



## Update on PXD Offline Calibration

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## Outline

- New features
- Status of pixel masking
- Studies on gain calibration
- Summary & outlook

### New features

- Airflow support for automated calibration (release04-01)
  - Dedicated python module for PXD calibration
  - Script for Airflow framework, handling run chunks and configurations
- Issues/limits:
  - All data are grouped into run chunks
    - Cosmic, beam & physics runs require different treatment.
    - Annealing in gain calibration
  - No enough data or broken files in a run chunk will generate FAILURES
    - Will prevent the processing of calibration algorithms
    - Could be removed manually, but current <u>Airflow framework will get stuck</u>
- Workaround:
  - Always sending good runs for calibration
  - Different calibration strategies for sorted chunks
    - Physics/beam: Sequential run-by-run strategy, several runs/IOV, the last IOV of each chunk is open
    - Cosmic data: Simple run-by-run strategy, only one run/IOV
      No gaps, but with overlaps

## Status of pixel masking

- Local calibrations for Bucket8 (all runs in experiment 10) have been finished.
  - We also keep our pxd\_calibration GT up to date.
- However, the recent background study still showed strong noise component (-> Sally's talk).
  - Changing threshold (7→3) doesn't help
  - Gain study also showed the hint that the noise peak around 8 ADU may be coming from a few pixels near the masked gates.

### Gain calibration studies

- lacksquare Track cluster study with heta correction
  - Cluster charge will be normalised→ More landau-like
  - Low statistics for bins in high V
  - Some bugs for regions with no enough data.
- Pixel level study
  - New figure of merit for gain comparison in a local region:
    - Seed probability for clusters with a certain size

## Summary & outlook

#### Summary

- In the past three months, Airflow related development is of the highest priority.
  - The test is still ongoing and PXD scripts work well so far
- The current hot pixel masking algorithm works fine (efficiency and noise tradeoff), but a dedicated algorithm aiming at cleanness is better for background studies.
- Full validation of gain calibration code has been resumed and more detailed studies are also in progress.

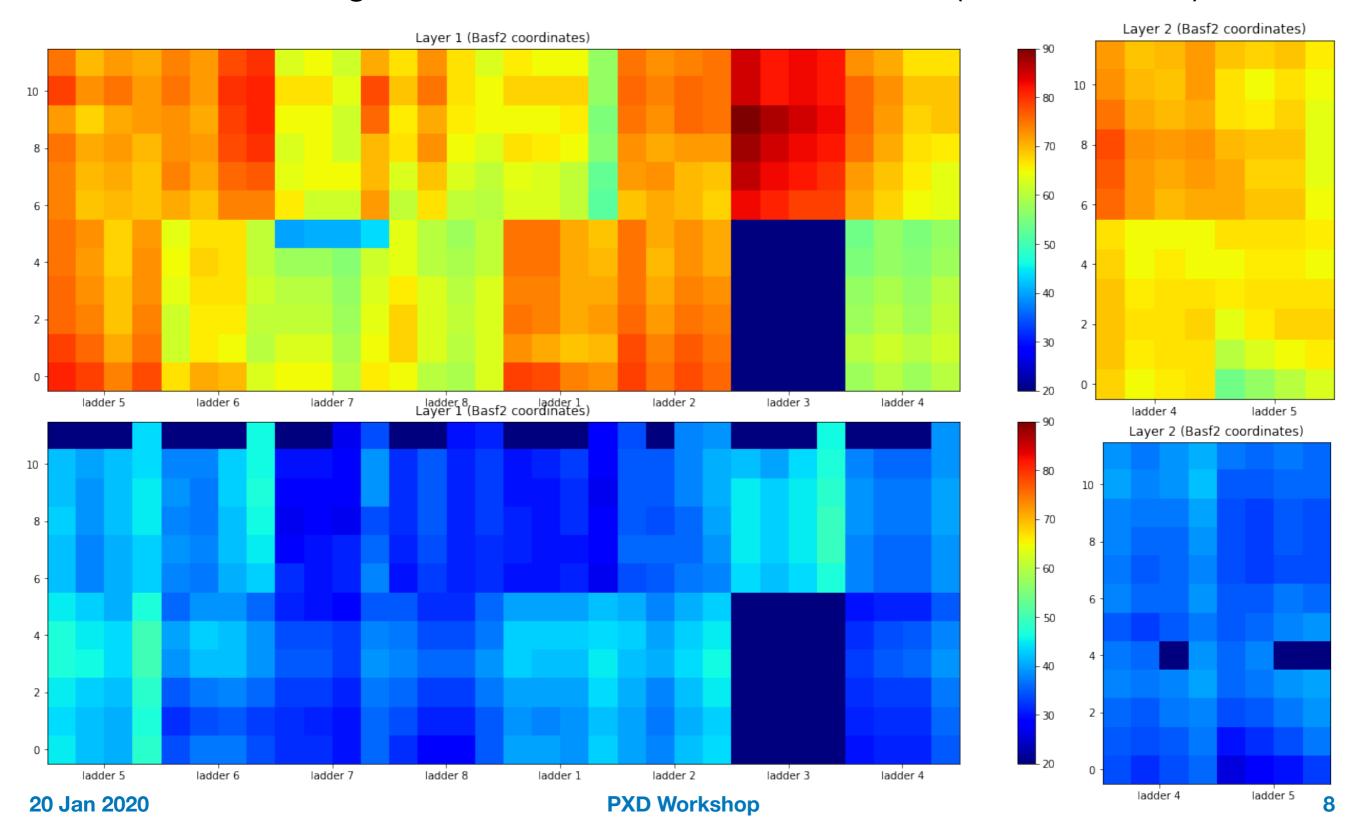
#### Outlook

- Track cluster charge distribution from MC
- Bug fixing in gain calibration code
- Request PXD clusters in CDST for post-tracking calibration

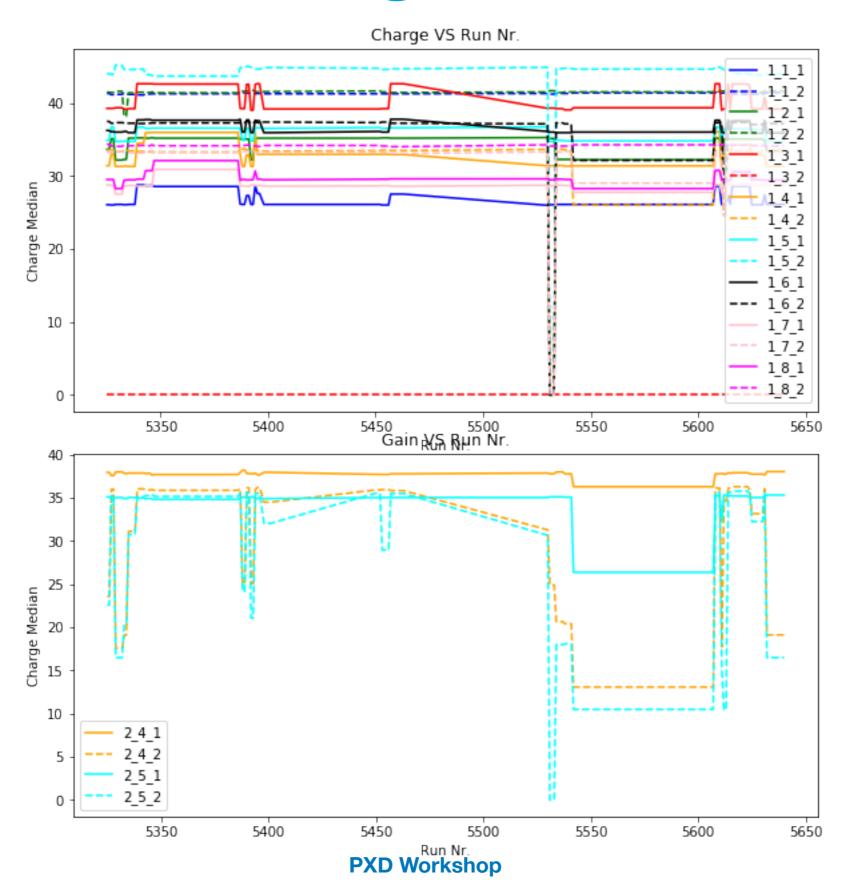
# Thank you!

## Backups

Cluster charge MPV before and after  $\theta$  correction (run 5401, e 10)

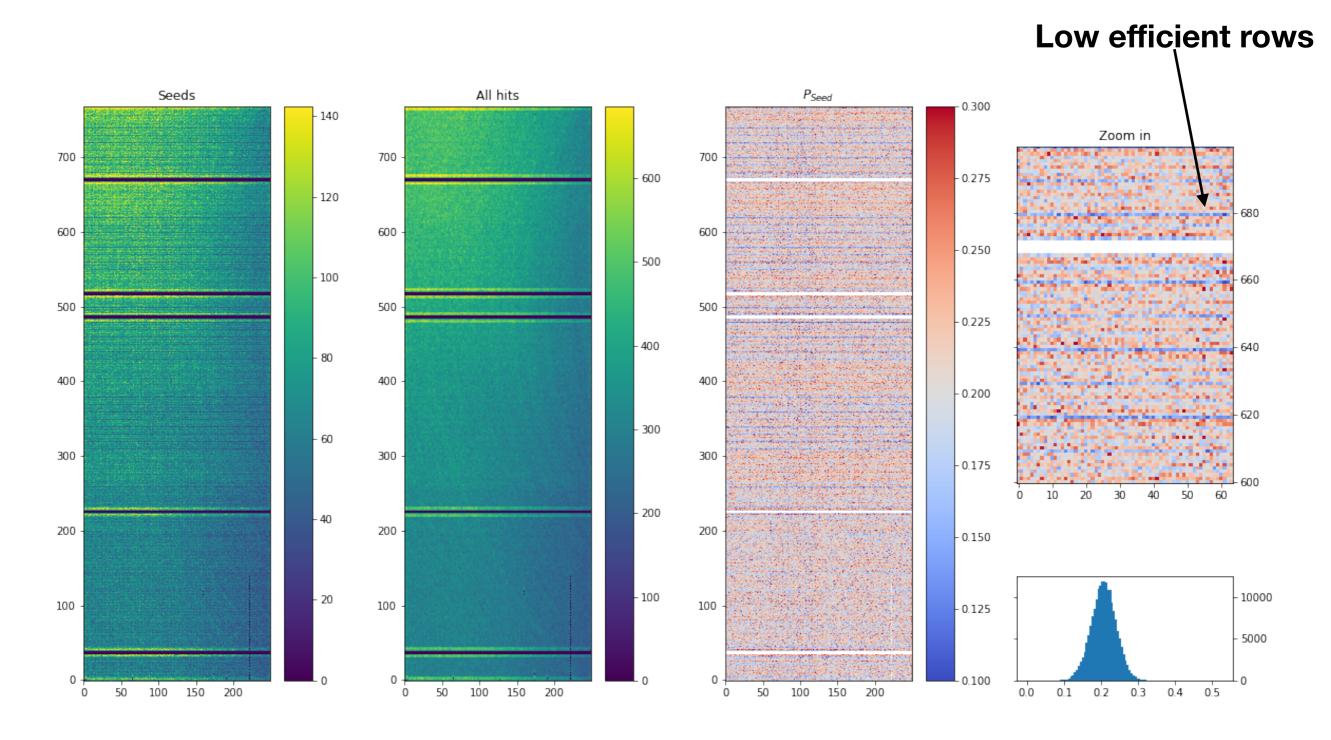


## Mean charge vs run No.



### Seed probability of clusters(size = 5)

No only track clusters (run 4700 - 4930)



### Seed charge distribution of single pixels

