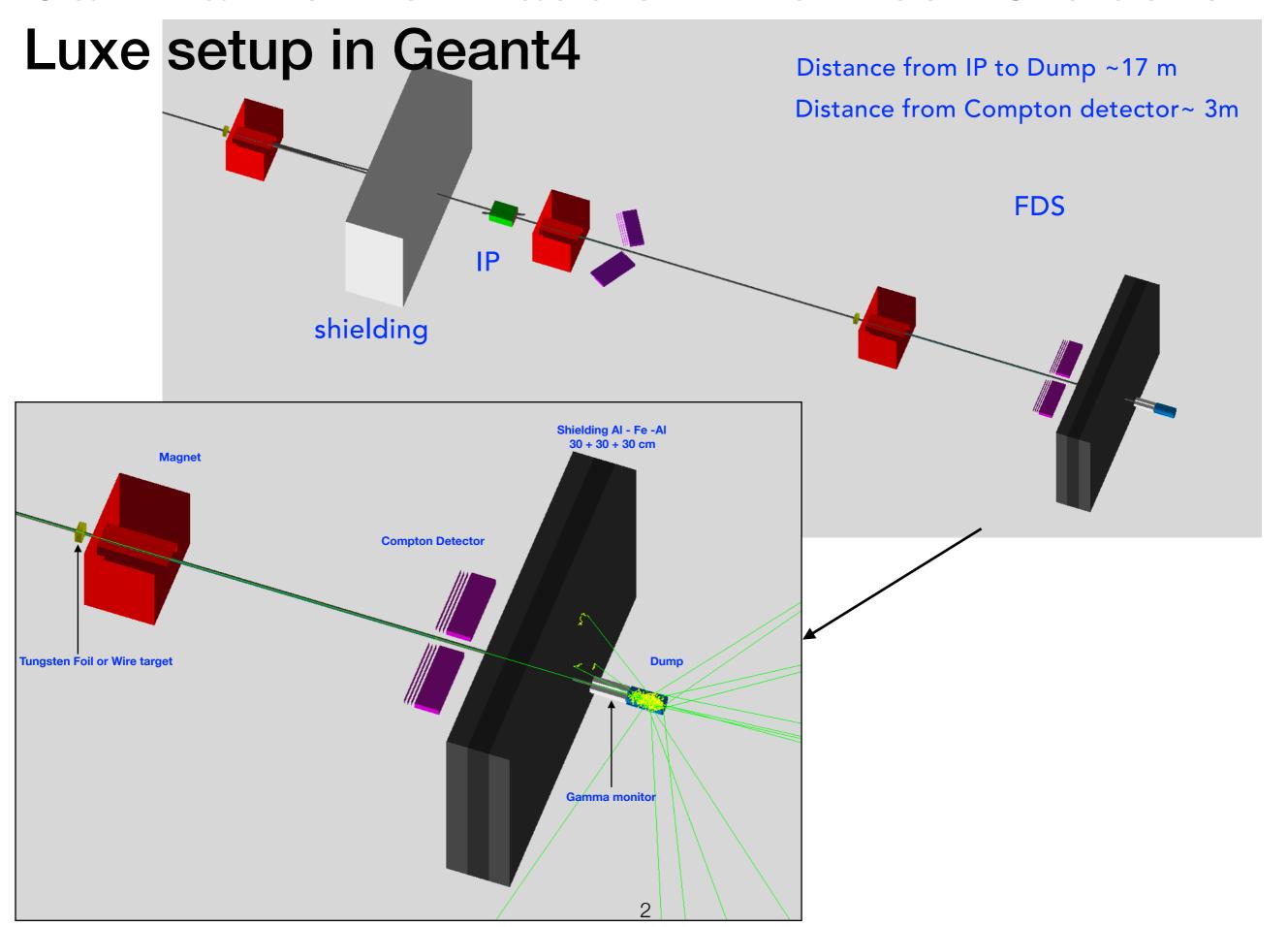
# Gamma Monitor using backscatters

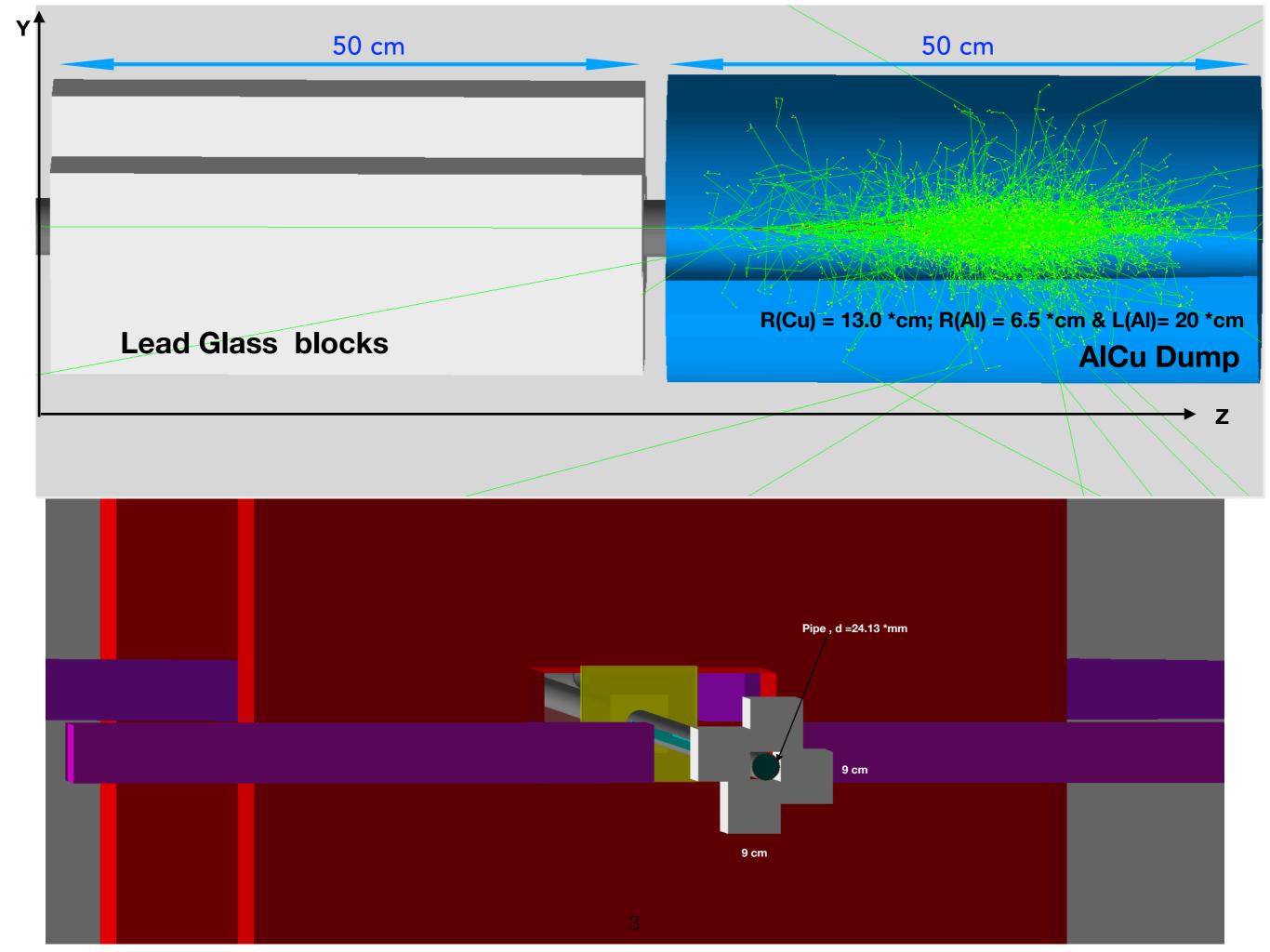
Borysova Maryna (KINR) 16/01/20

LUXE technical meeting

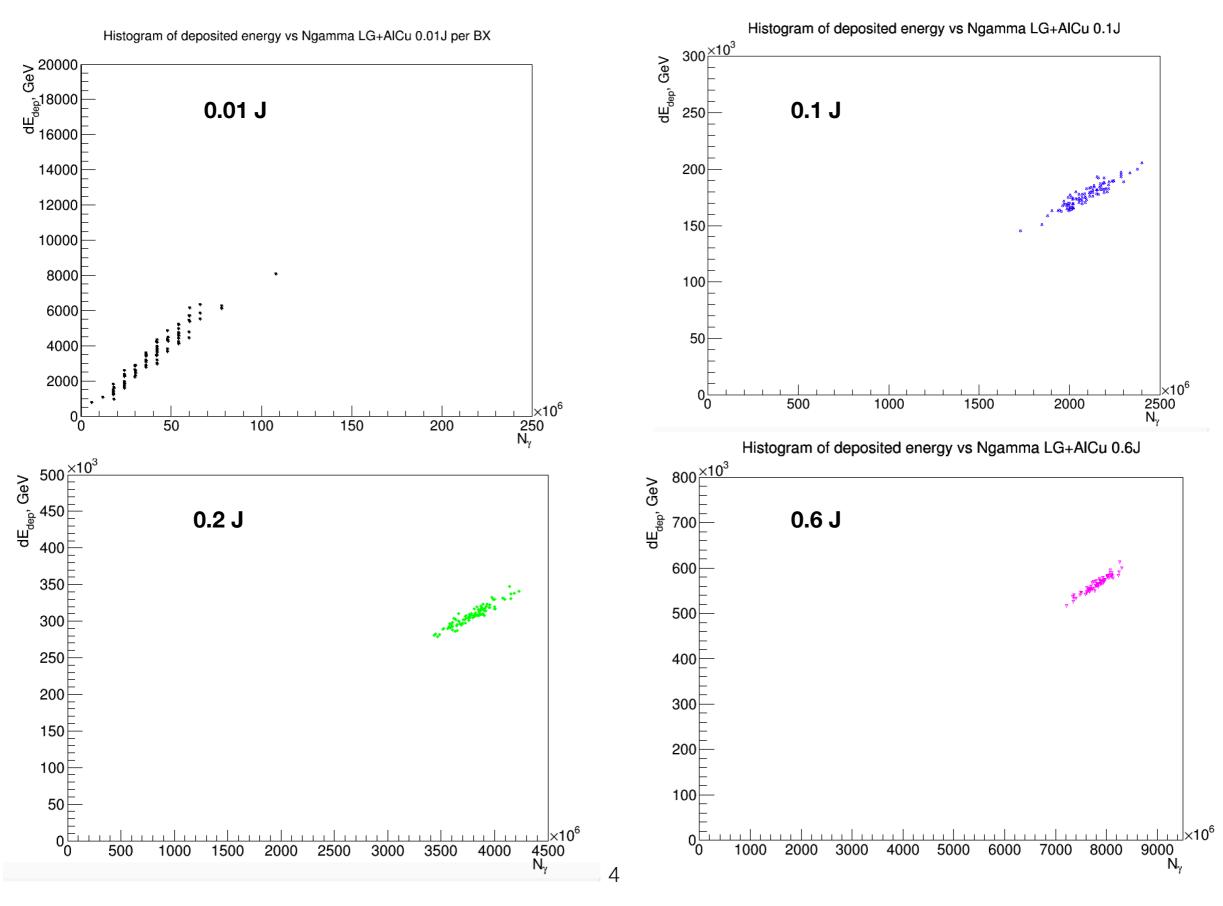


#### Gamma Monitor made of 4 Hermes LG blocks in

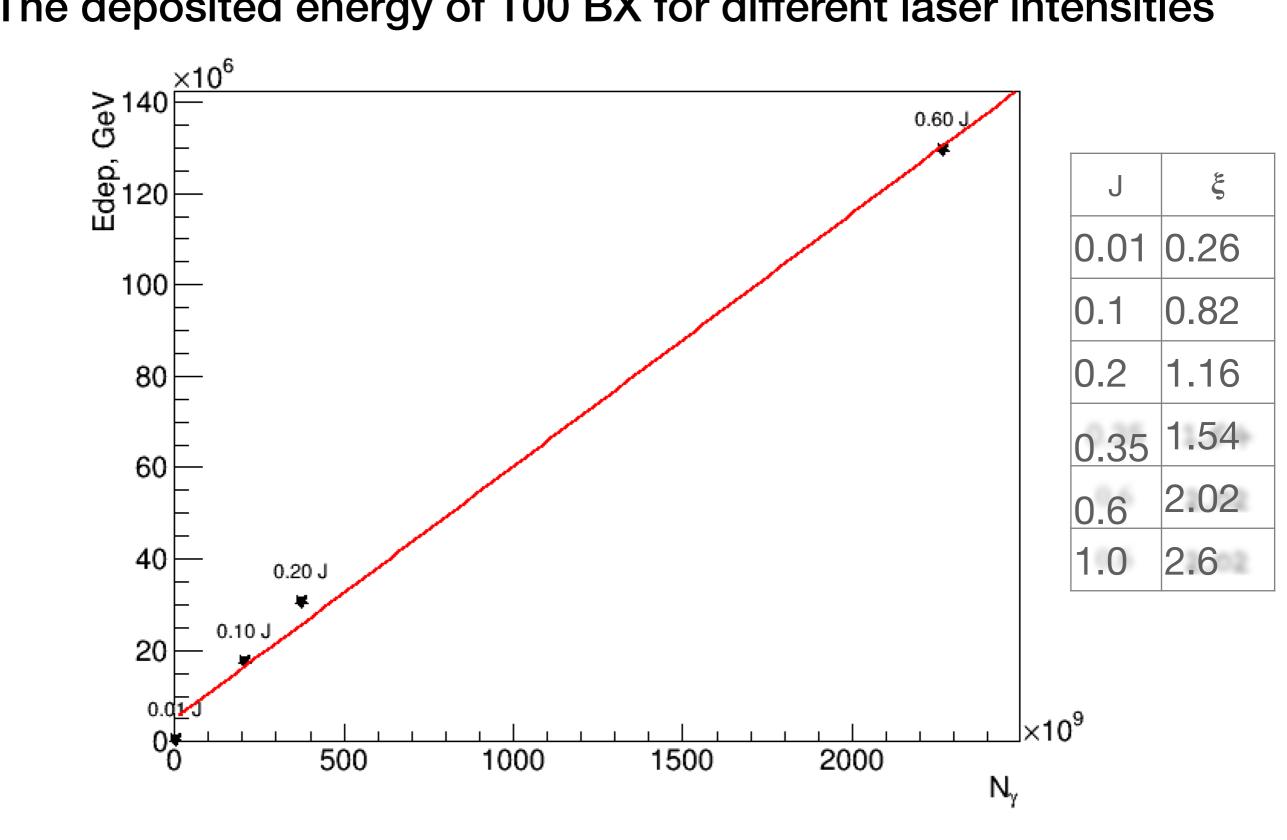




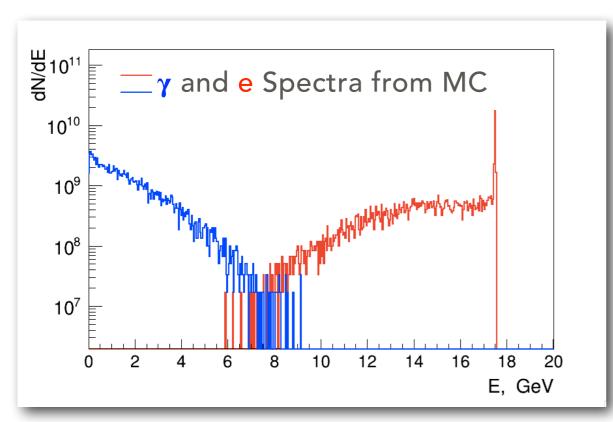
## The dependence of deposited energy on number of incoming photons per BX for LG Gamma monitor and AlCu dump



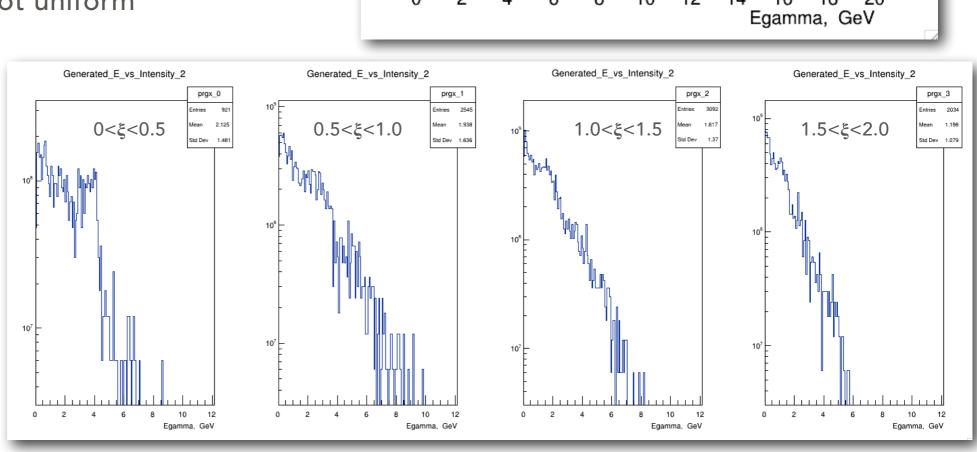
#### The deposited energy of 100 BX for different laser intensities

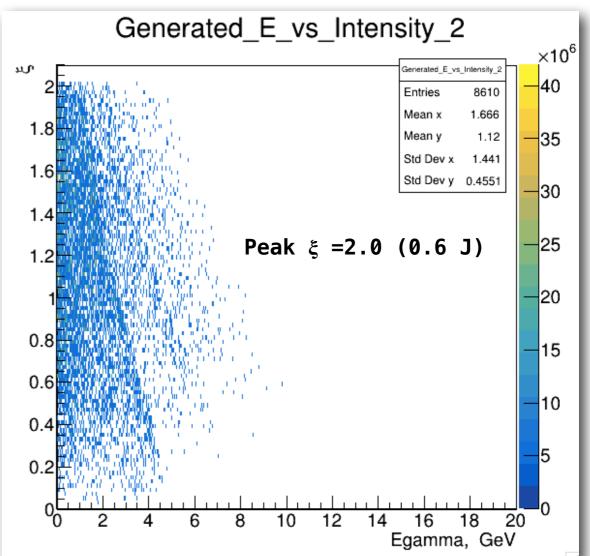


#### $\xi$ vs E $\gamma$ FROM MC



- Laser Intensity ( $\xi$ ) is not uniform
- This makes the kinematic edges from different n not visible
- ξ distribution might be reconstructed by fitting measured spectra w/ convolution of HICS xsection & ξ trial distribution

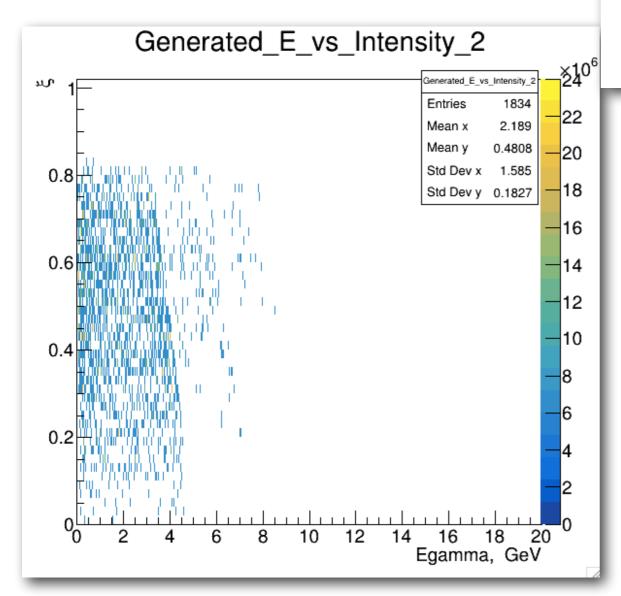


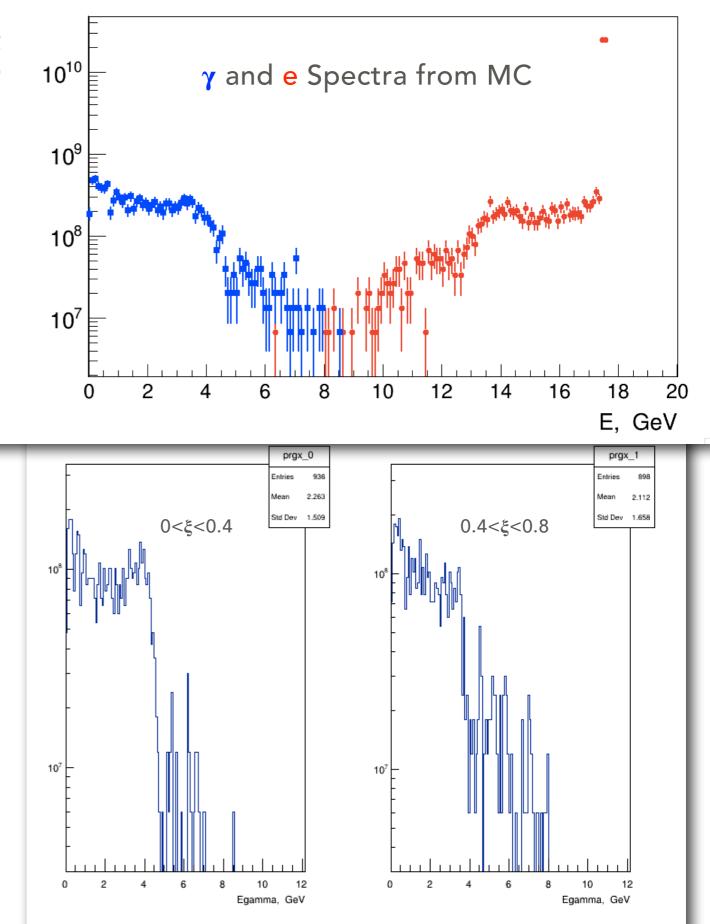


#### $\xi$ vs E $\gamma$ FROM MC

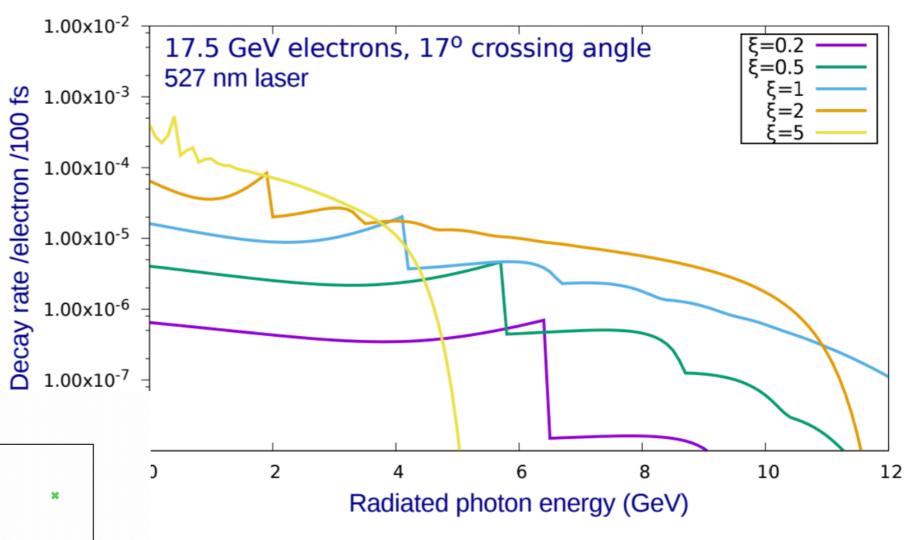
Peak  $\xi = 0.8 (0.1J)$ 

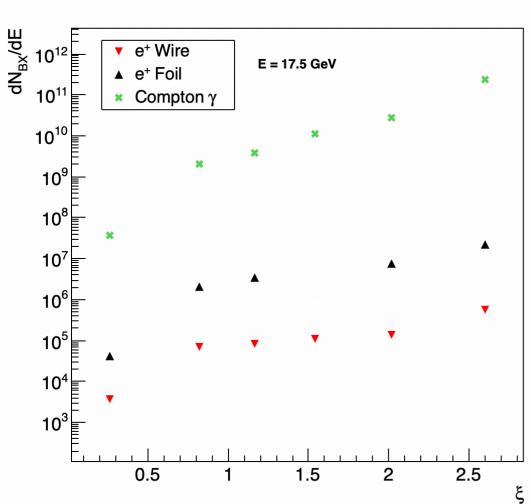
The kinematic edges can be seen at the low intensity.



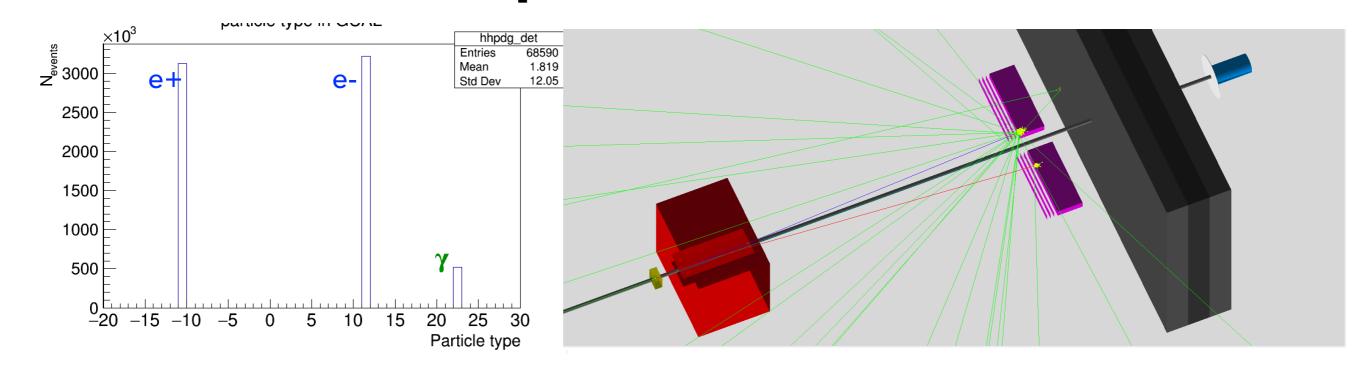


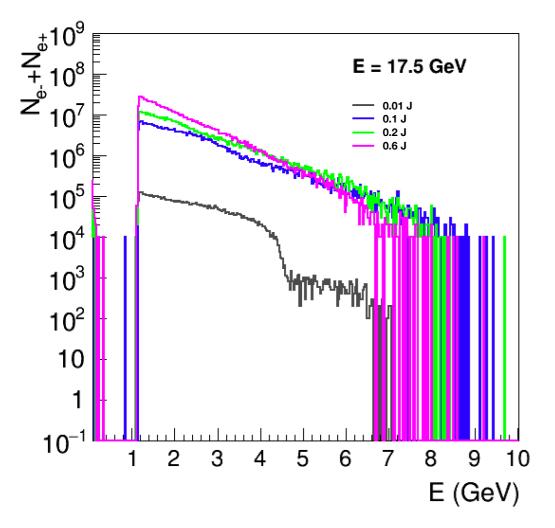
#### HICS for 17 GeV electrons, intensity sweep

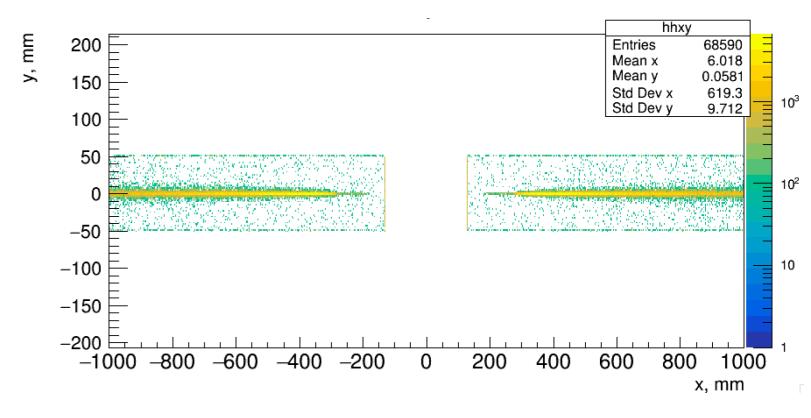




### Compton detector







## Outlook

- Gamma monitor studies:
- \*The implementation in Luxe geometry the LG Gamma Monitor made of Hermes LG blocks in front of Al-Cu Dump
- \*Gamma Monitor was studied in GEANT4 w/ LG Monitor in front of Al-Cu Dump for the lowest intensity
- \*The linear dependence of deposited energy on number of incoming photons allows the usage of backscatters for counting the photon flux for this configuration

#### **Further studies:**

To implement the simulation of the electromagnetic shower and the resulting Cherenkov photon production

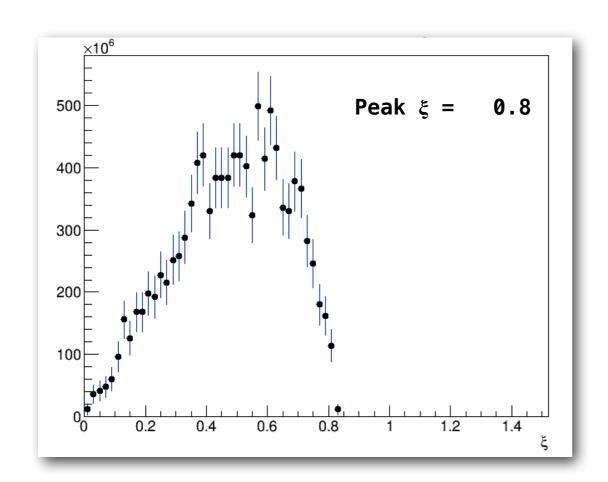
To implement mylar foil to reflect the Cherenkov photons, and a tedlar foil to block light To study background

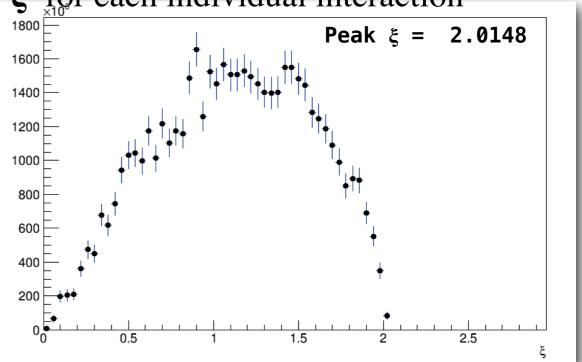
## Back up

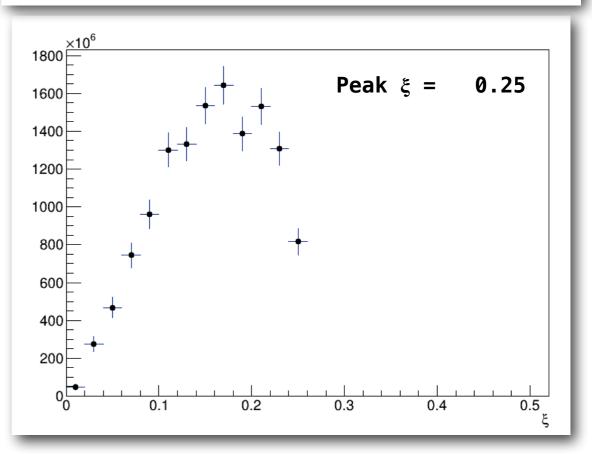
### **Laser Intensity**

MC simulation provides information for  $\xi$  for each individual interaction

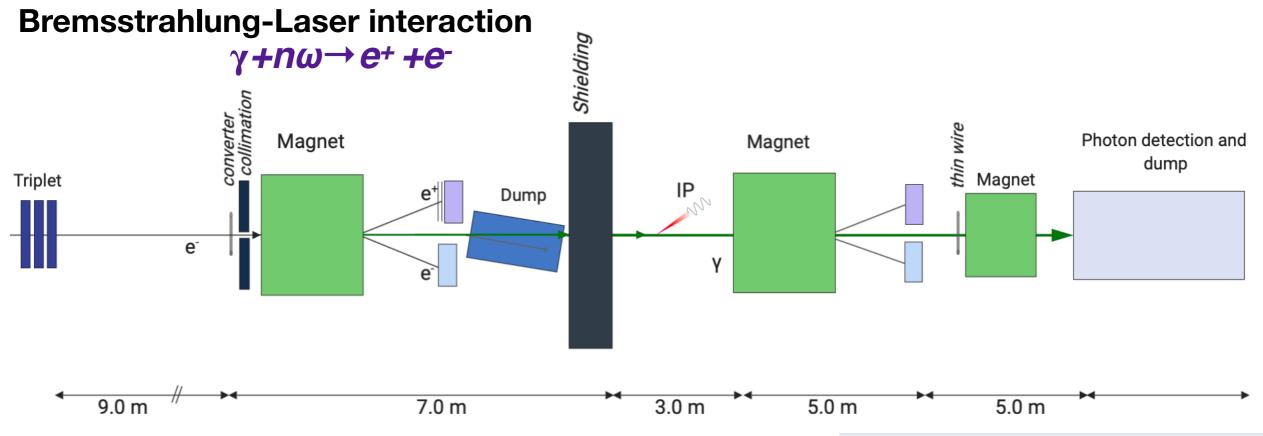
- •Realistic simulation of laser pulse intensity distribution.
- The field is not the same across the laser pulse.



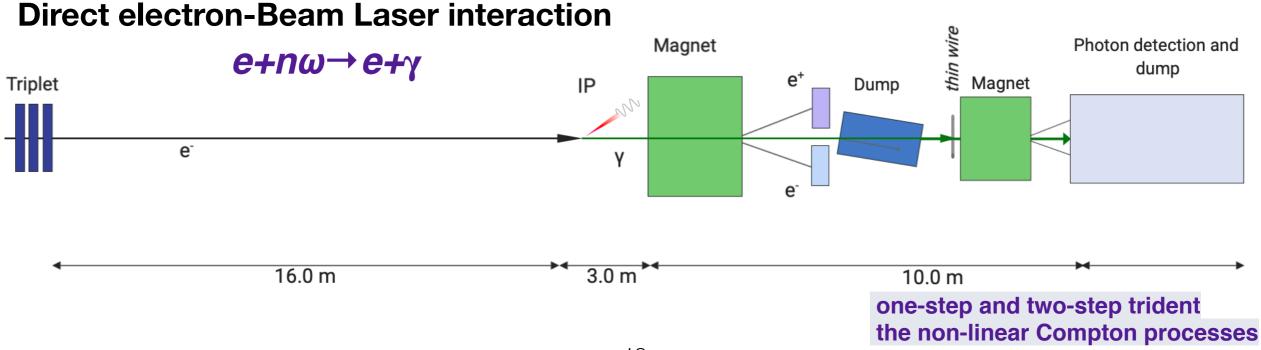




## LUXE Set Up



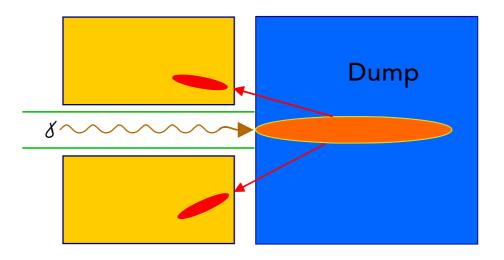




#### Rates from LOI For 6.0e9 electrons in BX $\omega$ / E= 17 GeV

Location	particle type	rate for $\xi$ =2.6	rate for ξ=0.26
e- detector	e-, E <16 GeV	5.9e+9	2.4e+07
e+ detector	e+	61.07	0.0
Photon detector	γ	2.4e+11	3.8e+07
Photon detector	e+ and e-	2.3e+07	4.2e+04
Photon detector	e+ and e-	5.8e+5	3.8e+03

#### The Idea:



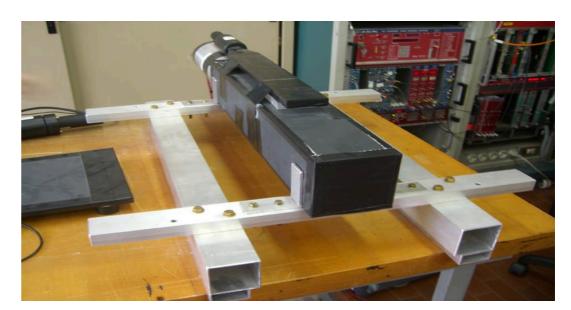
#### Lead glass blocks from Hermes Experiment

Available: 6 calorimeter blocks w/ measures  $9 \times 9$  cm<sup>2</sup>, length is 50 cm



Chemical composition (weight%)		
$Pb_3O_4$	51.23	
$SiO_2$	41.53	
$K_2O$	7.0	
Ce	0.2	
Radiation length (cm)	2.78	
Density $(g/cm^3)$	3.86	
Critical energy (MeV)	17.97	
Moliere radius (cm)	3.28	
Refraction index	1.65	
Thermal expansion coefficient $(C^{-1})$	$8.5\cdot 10^{-6}$	

Table 1: Chemical composition and physical properties of the F101 LG



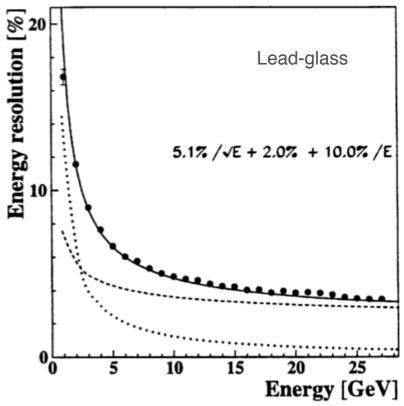
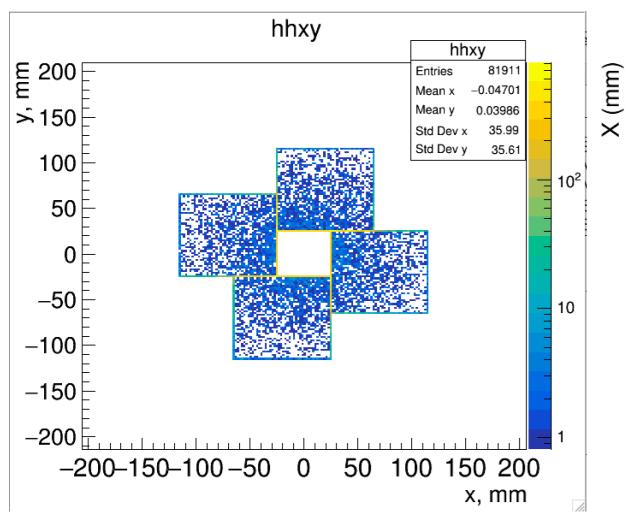
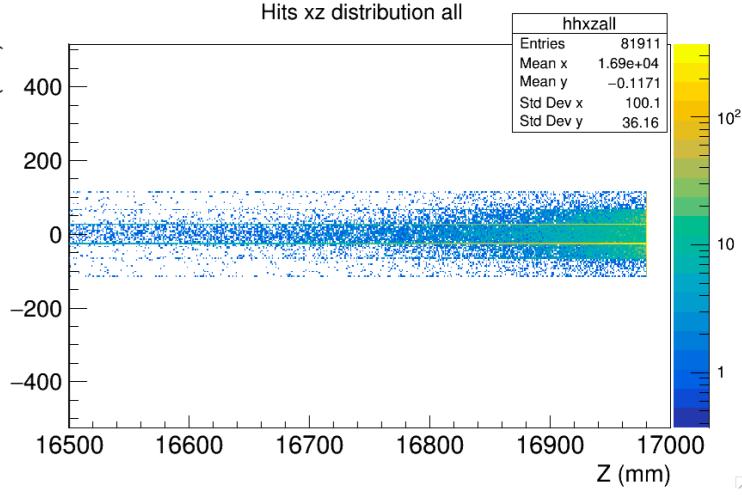
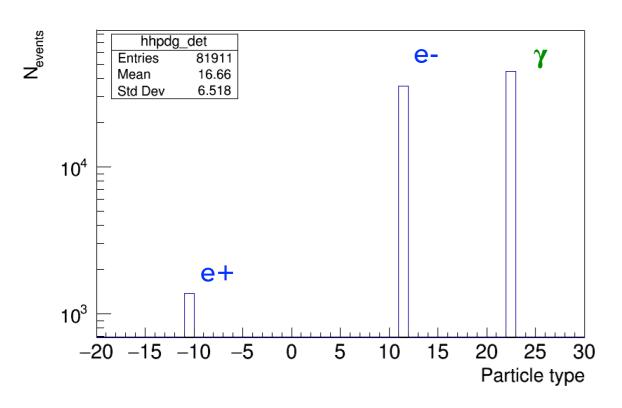


Fig. 5. Energy resolution of the calorimeter; the circles correspond to the 1996 data, the solid curve is the sum of the contributions from the lead-glass (dashed curve) and from the pre-shower (dotted curve) provided at test beam measurement [7].

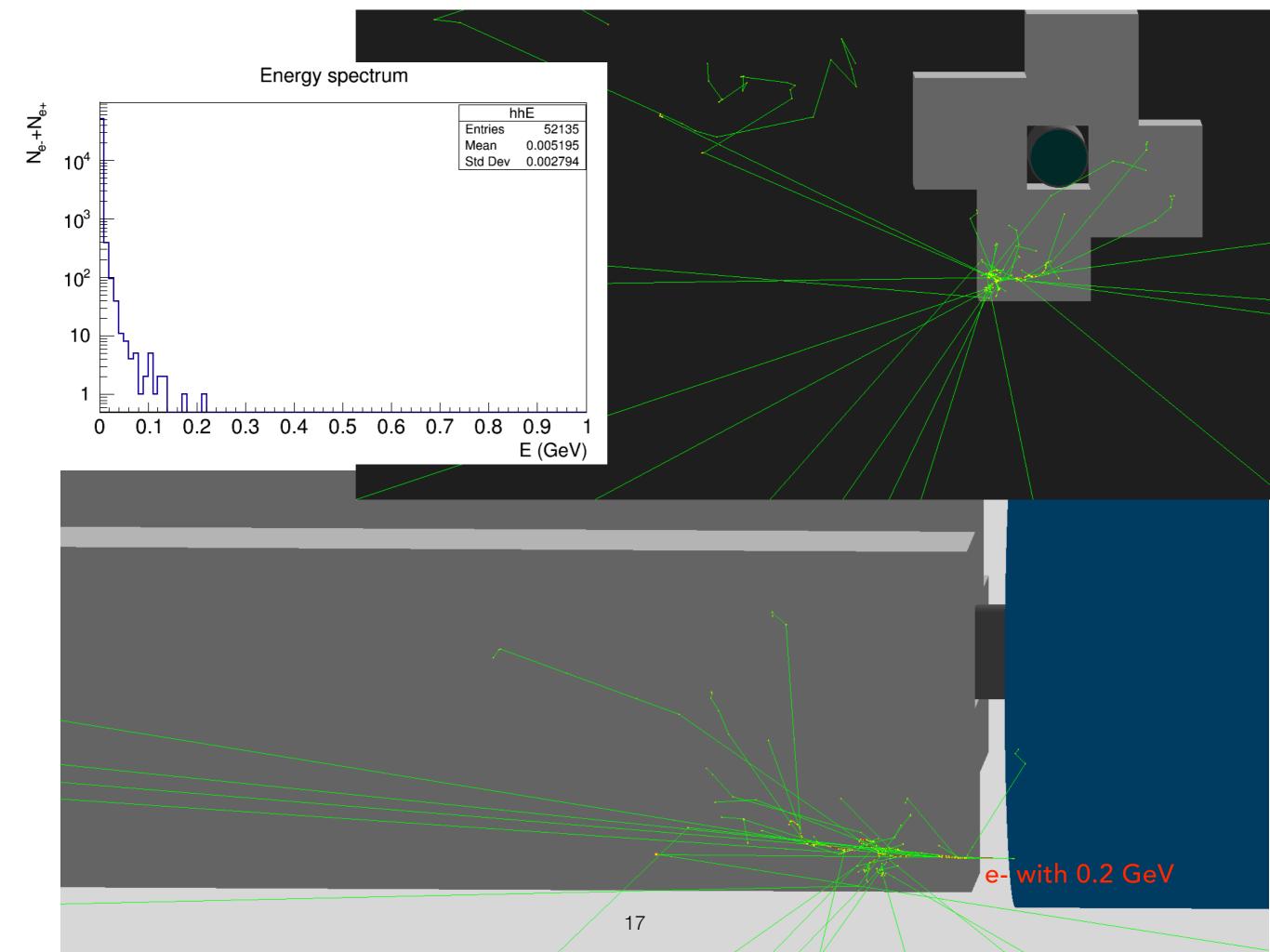
### Control plots for LG monitor



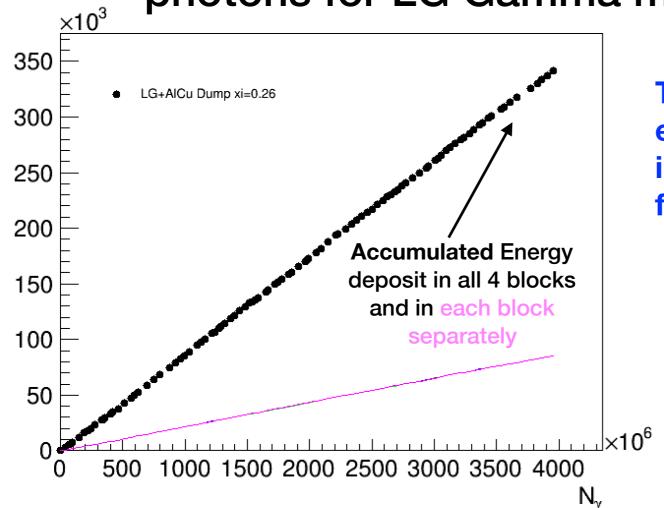




The distribution of hits of particles entering LG Gamma monitor in XY and XZ planes



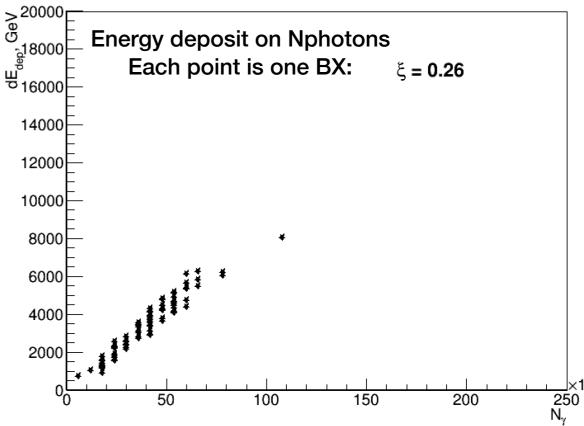
## The dependence of deposited energy on number of incoming photons for LG Gamma monitor and AlCu dump



Edep, GeV

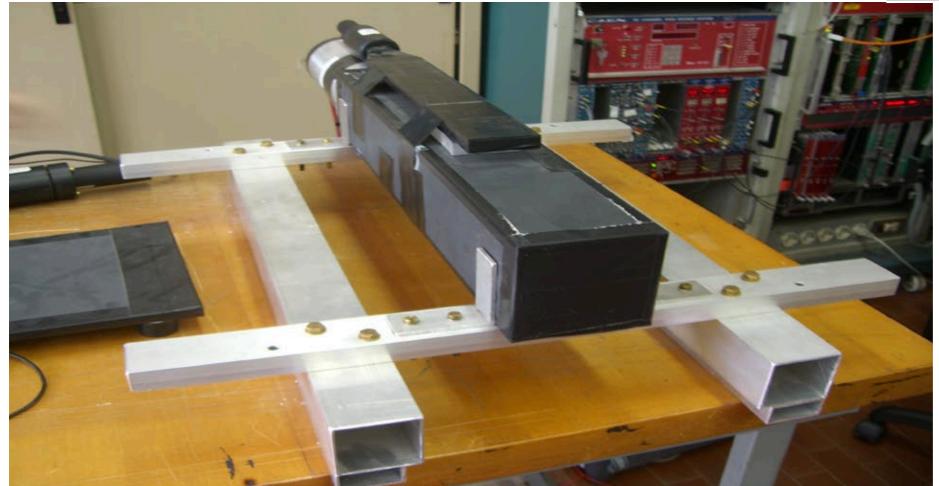
The linear dependence of deposited energy on number of incoming photons in LG allows the usage of backscatters for estimating the photon flux



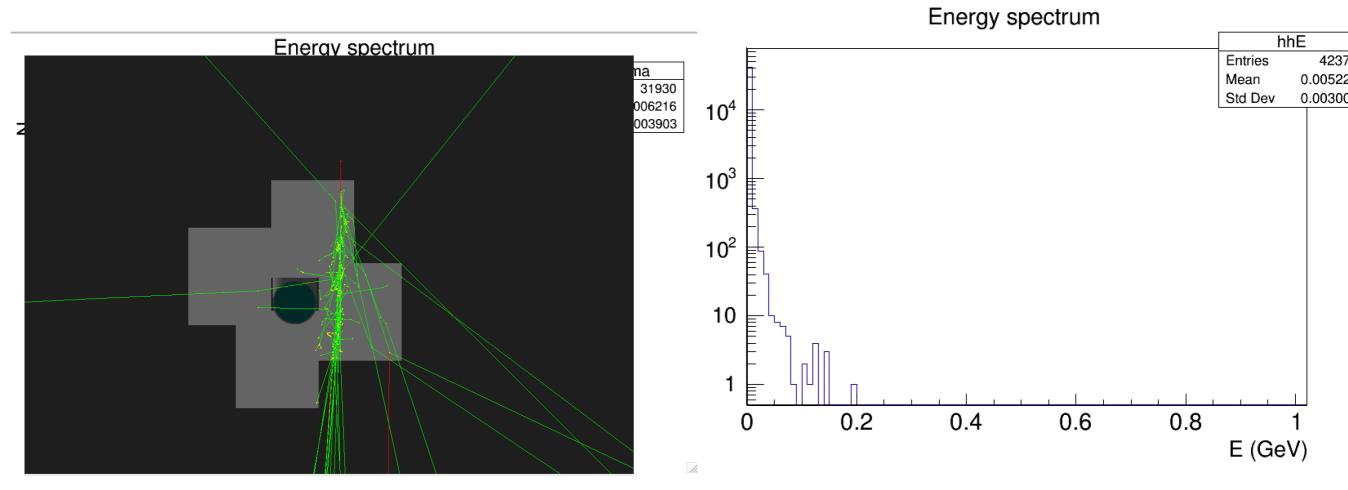


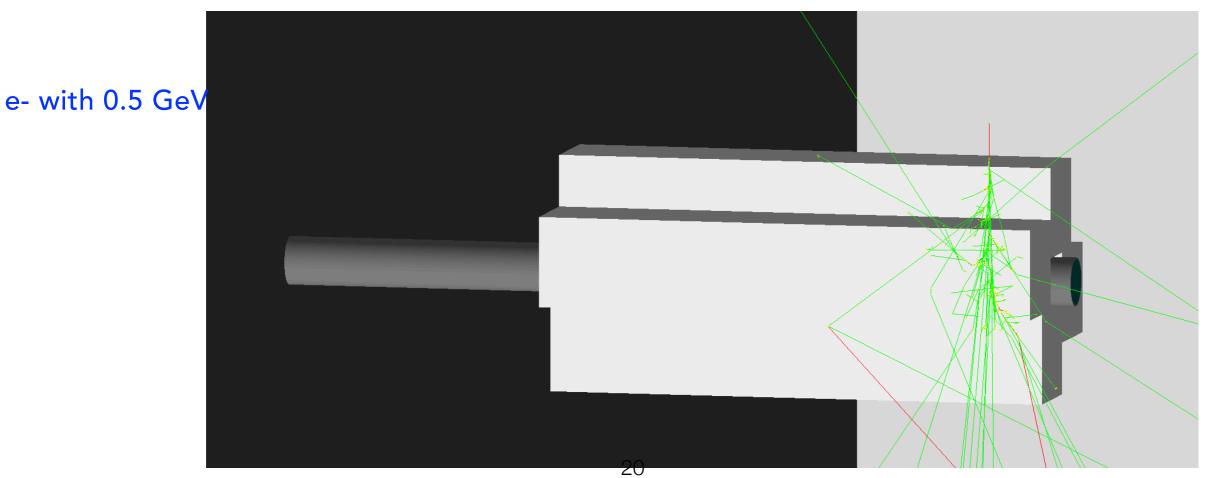
## Wrapped LG block

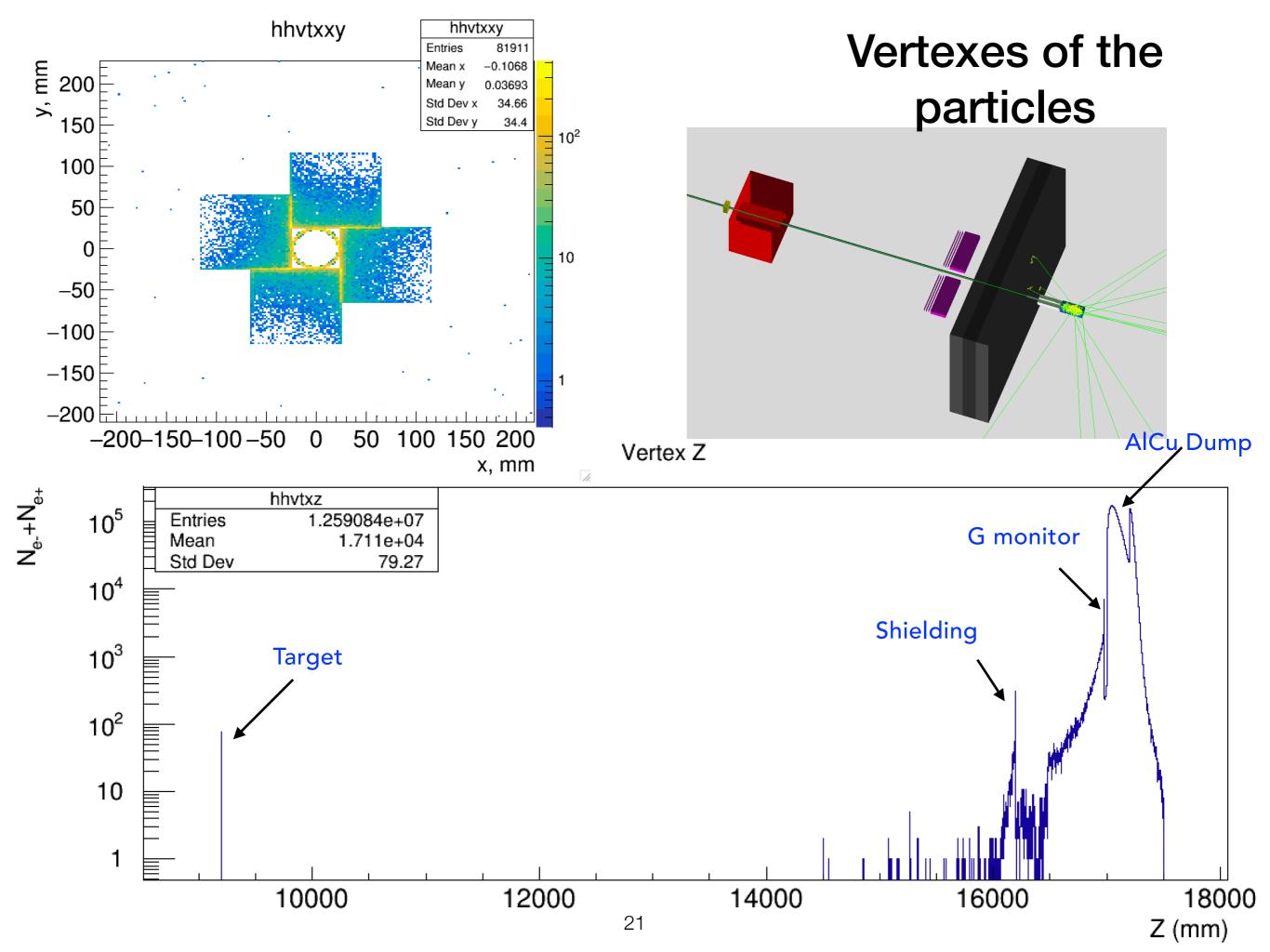


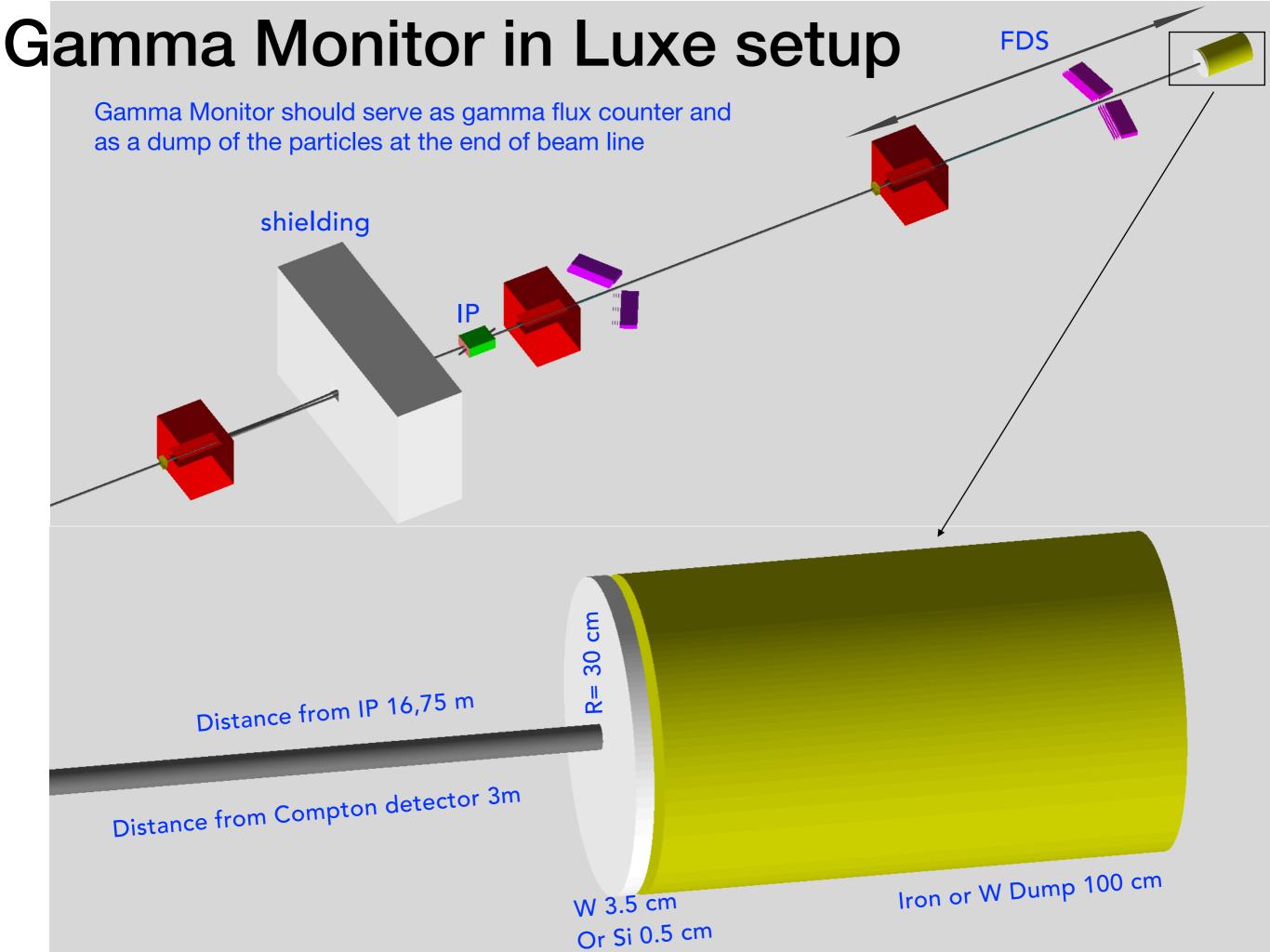


blocks are wrapped with an aluminized mylar foil to reflect the Cherenkov photons, and a tedlar foil to make the blocks light tight.



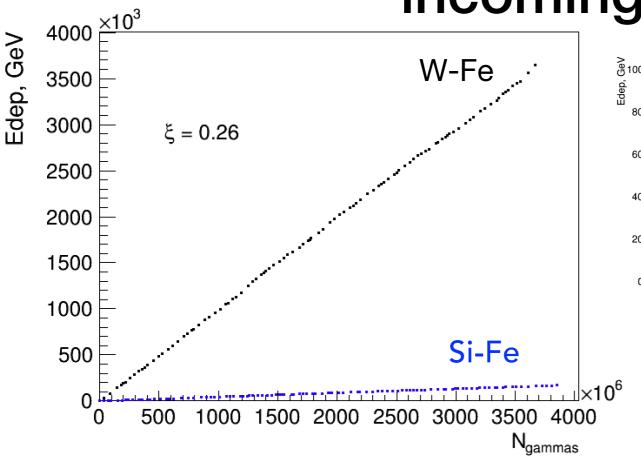


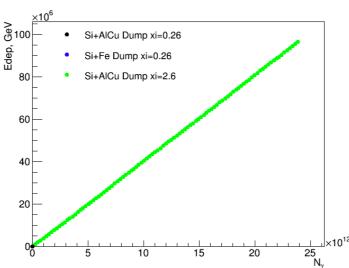




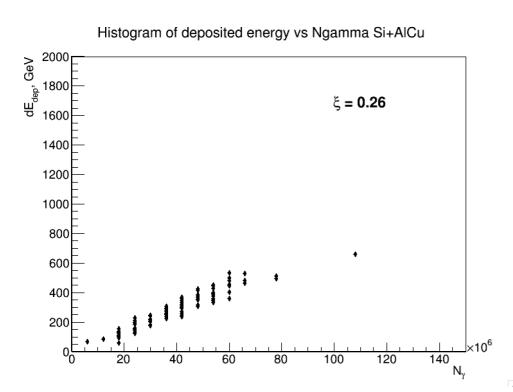
Energy dependence on number of incoming photons

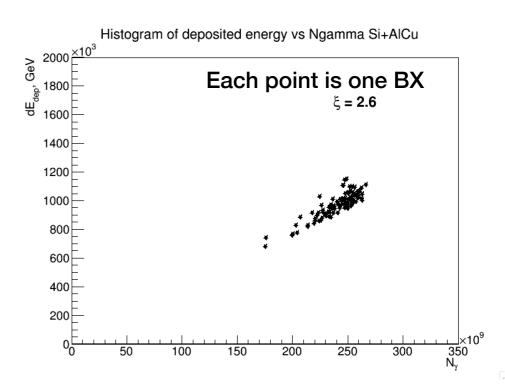
1



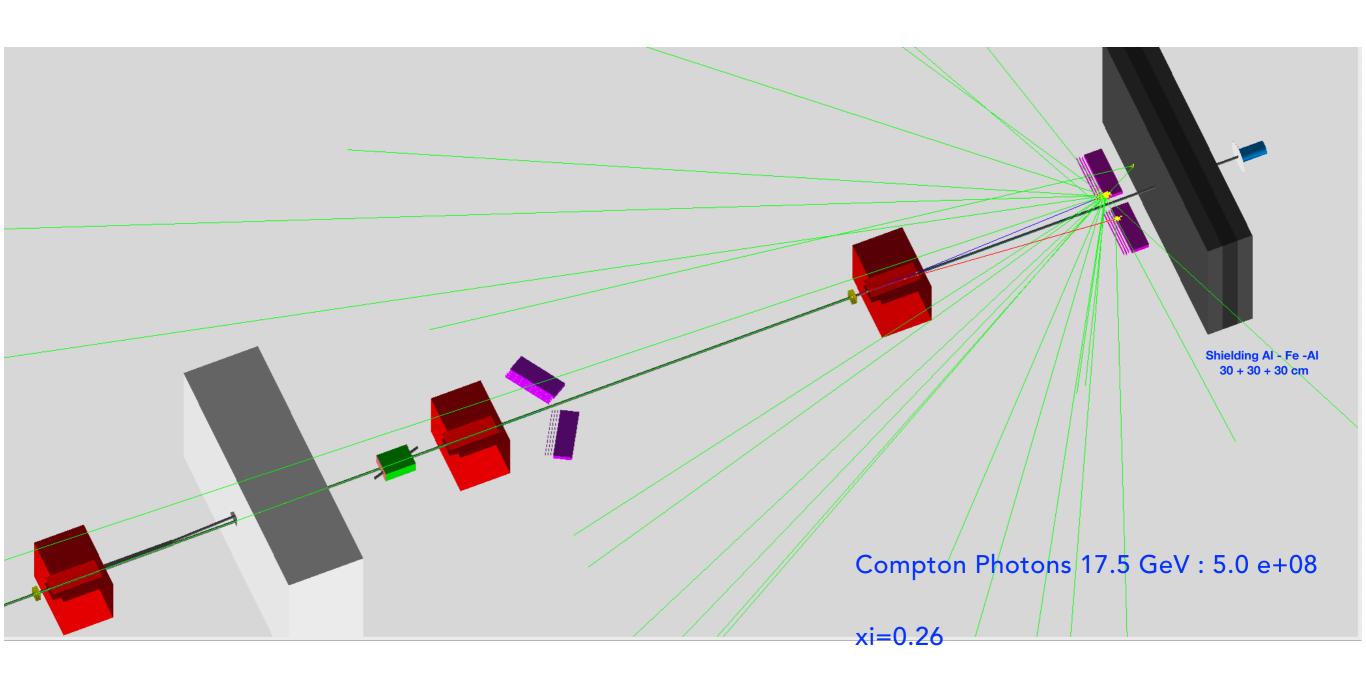


The linear dependence of deposited energy on number of incoming photons allows the usage of backscatters for estimating the photon flux





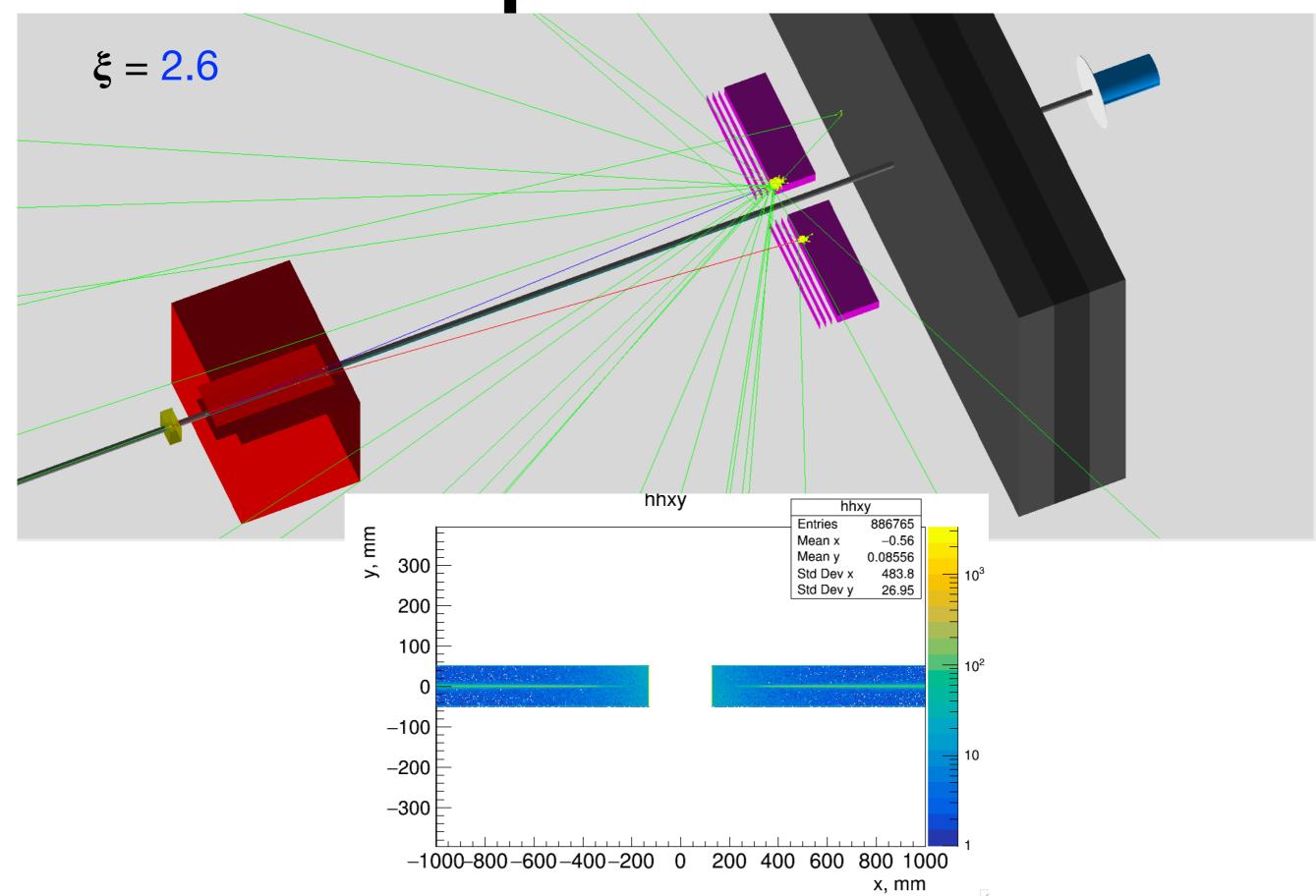
#### Luxe setup with non-tilted Compton Detector



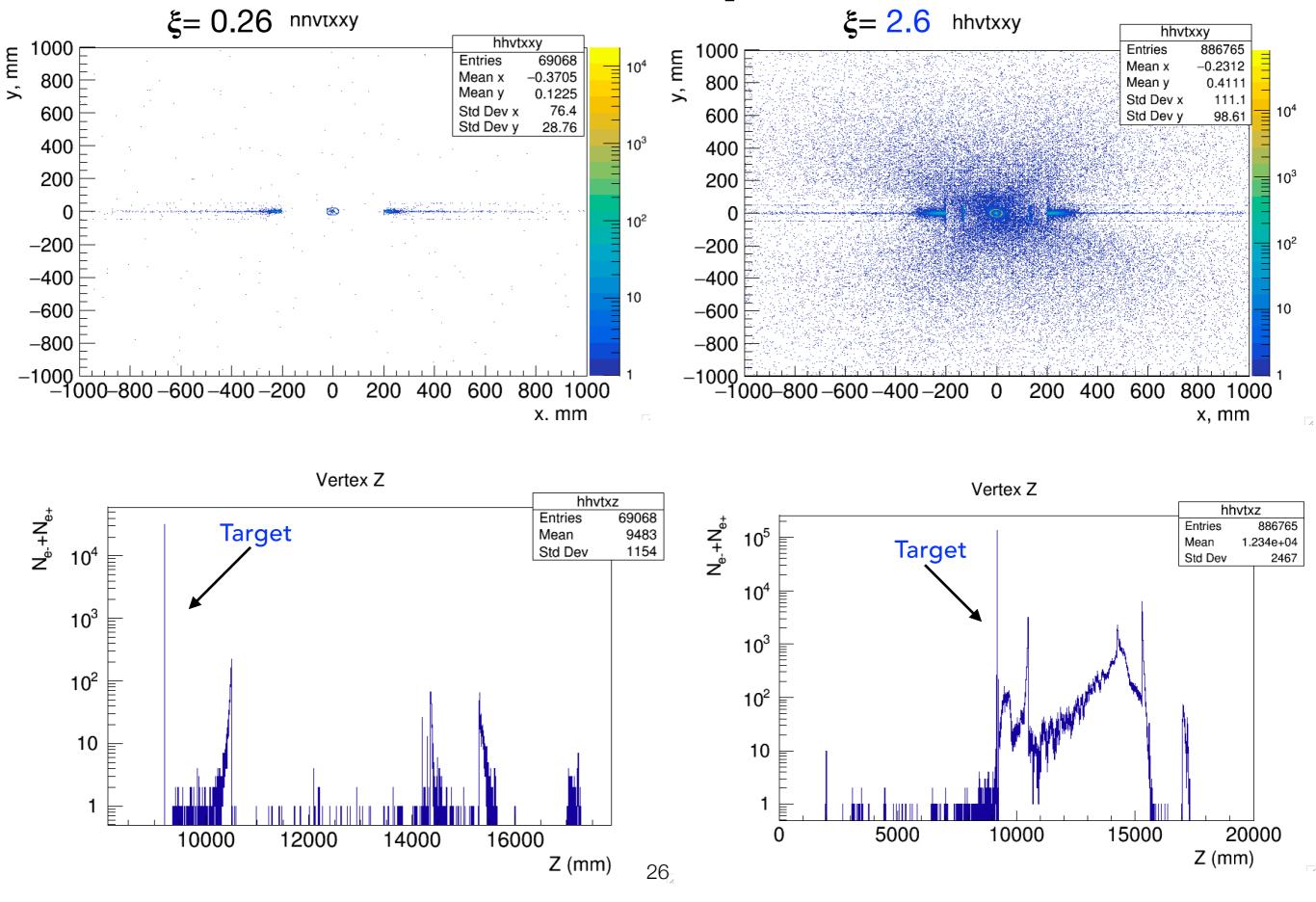
100 BX

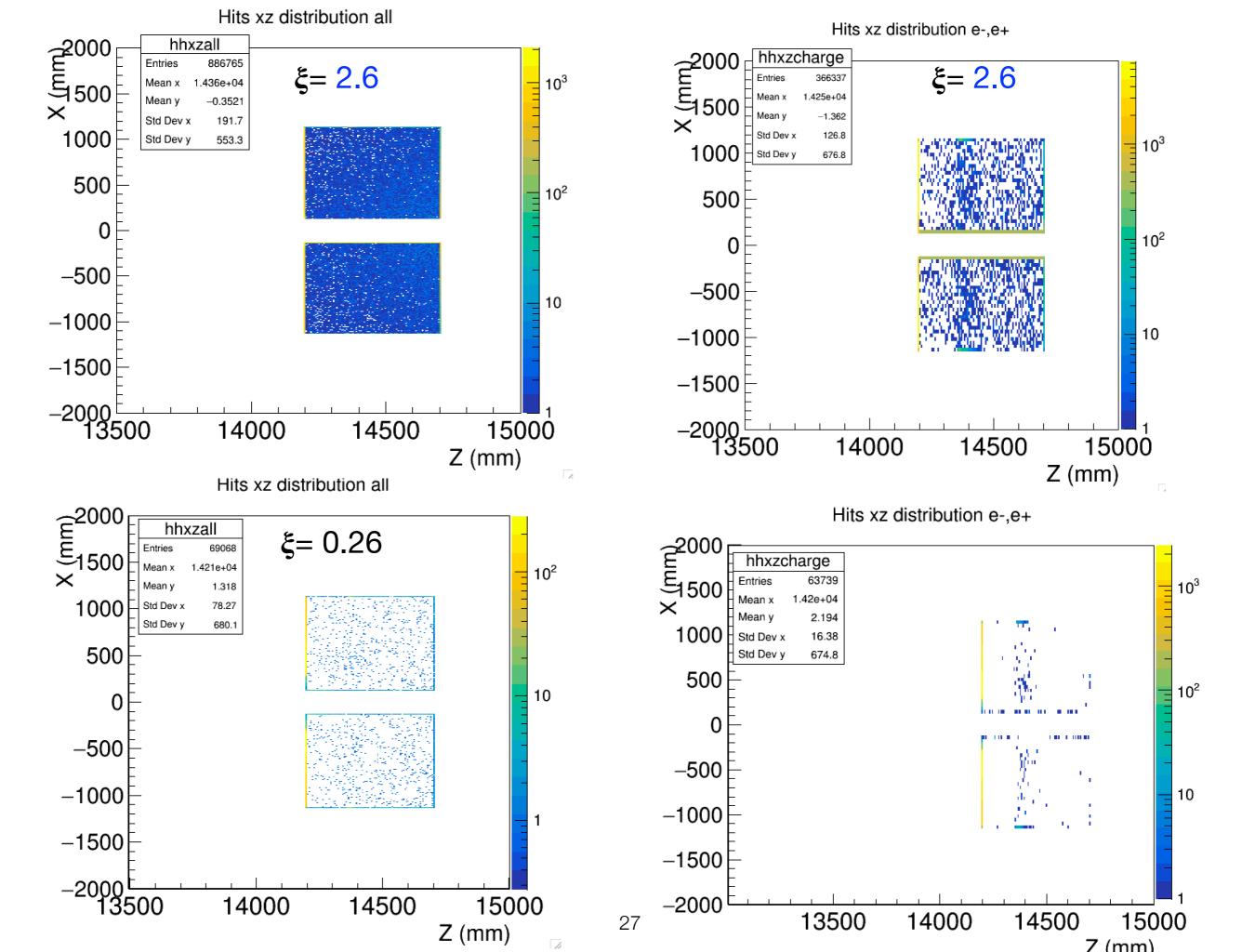
Target: W foil 10 um

Compton detector

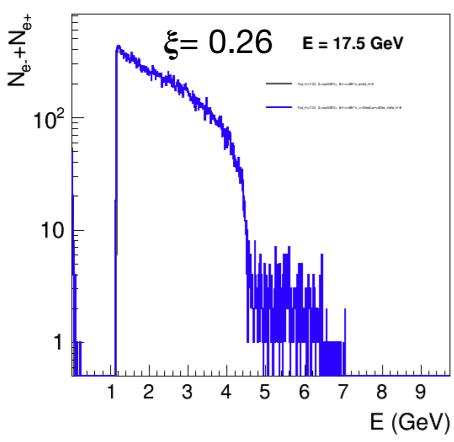


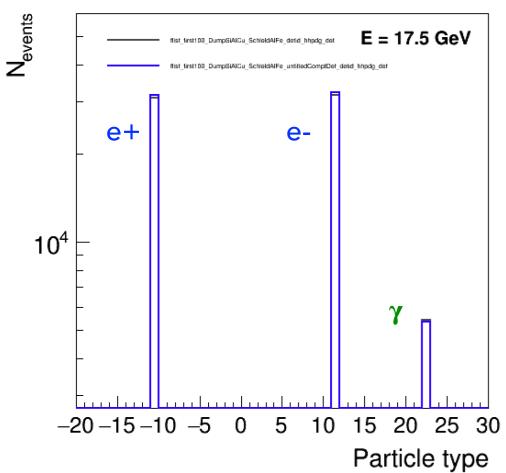
## Vertexes in Compton detector

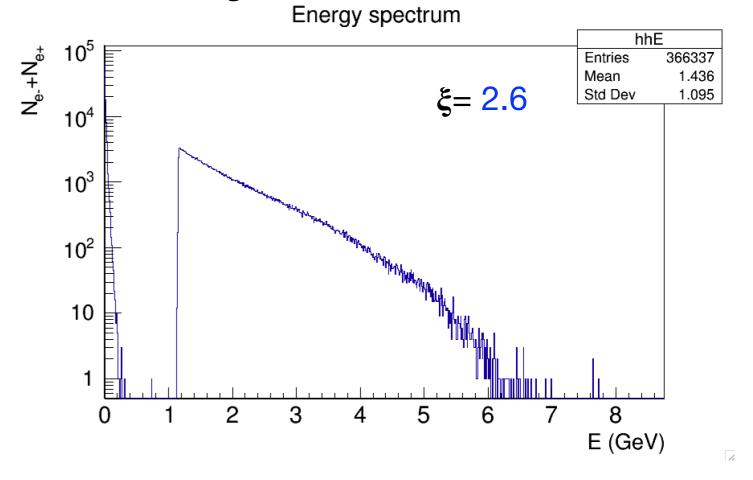




### Compton detector: $\xi$ = 2.6 vs 0.26

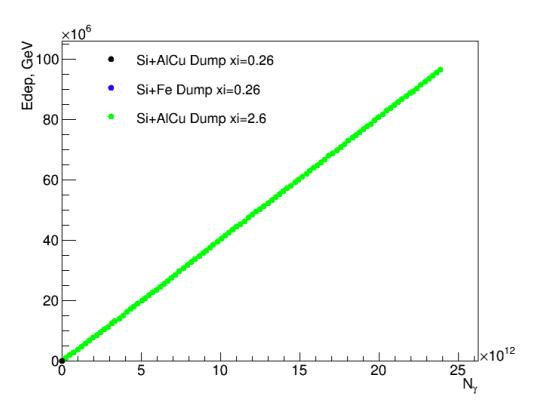




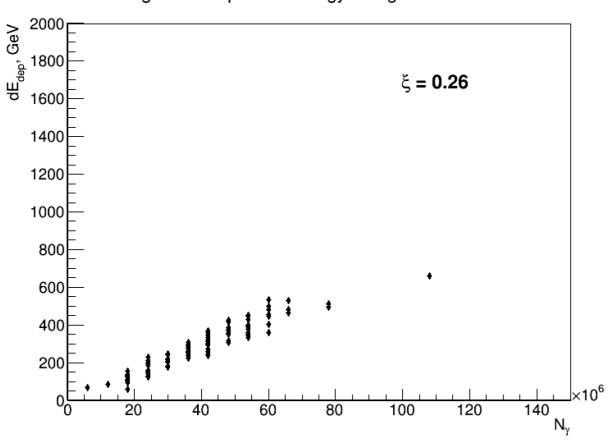


# The dependence of deposited energy on number of incoming photons for Si Gamma monitor and AlCu dump for different laser intensities

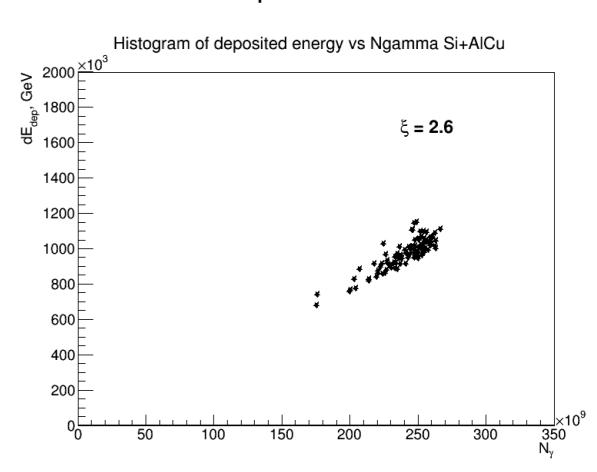
29



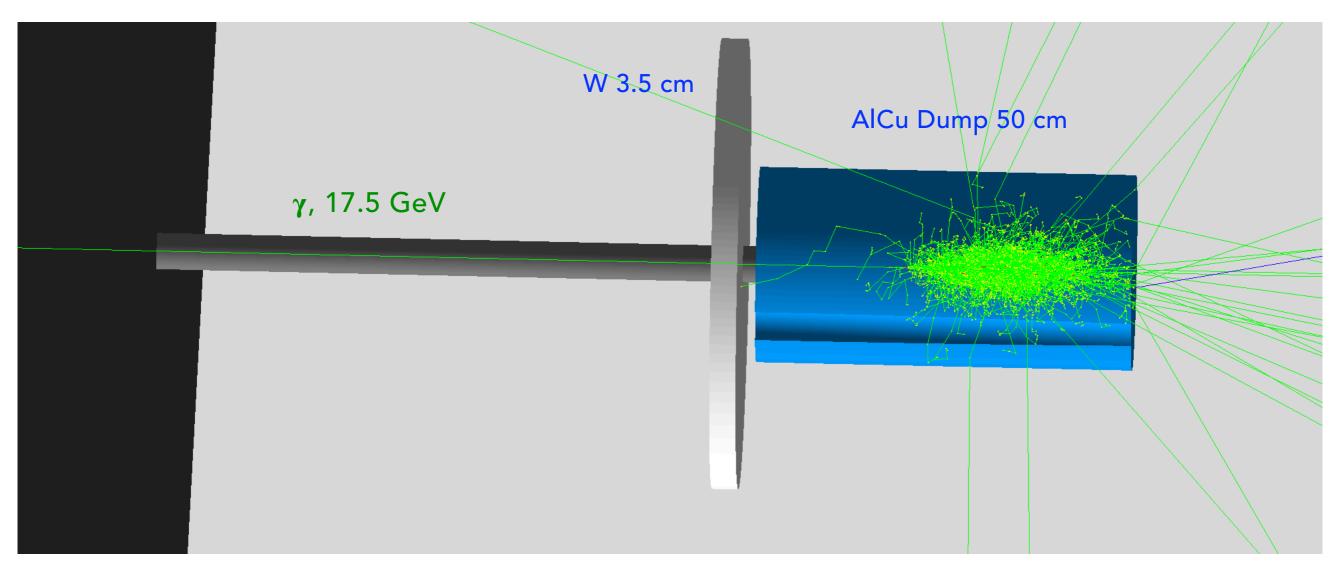




#### Energy deposit on Nphotons Each point is one BX



### Tungsten Gamma Monitor in Luxe setup



Compton Photons, 100 BX

 $\xi = 2.6$ 

Target: W foil 10 um

## The deposited energy on number of incoming photons for W Gamma monitor and AlCu dump

