# Path to an ALPS II Upgrade

# Plans on improving the performance of the optical system in summer/fall 2021

Aaron Spector ALPS Collaboration Meeting, June 8-9, 2021

## **Characterizing performance**

**Duty Cycle vs. Sensitivity** 

### Improving Duty Cycle

- Actuators (or lack thereof) likely limiting duty cycle
  - Proposal: Floating tables
- Very low duty cycle could limit sensitivity if measurement time is limited

### **Improving Sensitivity**

- Increasing power circulating in PC
  - What needs to change for this?
- Increasing power build up in the RC
  - Plans to procure lower loss mirrors
- Other ideas (Guido)

## **Actuator limited by pointing noise**

### Improving Duty Cycle

- New long range actuator design could increase dynamic range
  - Same wave washer 'sandwich' concept as previous actuator, but with short range and long range piezo
    - Long range Piezo 100um with 1000V commercial HV
    - Short range Piezo 2.5um with 400V AEI HV
- Newport stage is the suspected cause of the pointing coupling
- Not expected to be limited by long term pointing of tables
- Even with long range actuator PZT actuator pointing coupling is still expected to be a problem

### **Other Actuator Options**

#### Floating tables at the end stations as long range actuators

- Would potentially provide long range angular and longitudinal actuators
  - No active actuation of pointing with current system
- Likely improved performance in length lock PLL
  - Newport stage not needed, mount could be rigidly fixed to table
    - Increases the frequency of resonances (BW increases)
    - Isolation at critical frequencies (10-100Hz) of end stations
      - No isolation at central table, but potential for feed forward from a single sensor on COB
    - Likely to reduce pointing coupling of length actuator

### **Other Actuator Options**

#### **Drawbacks on floating tables**

- From correspondence with BILZ
  - Step size 10um
  - Company doesn't recommend using them for active horizontal positioning as it will short the isolation
    - Does that matter for such small actuation range?
  - Cost ~100k euro total for two tables
- Noise perfomance at resonances?
- Stability during commissioning?

# **Operating PC at higher power**

#### What needs to change

- COB QPDs gain (damage threshold?)
- Straylight?
- Other components that may have longer lead times?

# **Cavity Scattering loss**

Challenges to achieving high power buildup

#### Scattering losses due to surface roughness of substrates

- Deviations from a perfect sphere scatters light into higher order modes
- Integrated over the area of the beam
  - Larger beams → more scattering losses
- To achieve 40,000 power build requires losses < 25 ppm
  - ALPS IIc substrates: 40-65 ppm losses with 6 9 mm beam radius
  - aLIGO test masses: 40-65 ppm losses with 50 60 mm beam radius

## **Cavity Scattering loss**

#### For RC mirror $T_i = 100$ ppm, $T_o = 5$ ppm (with 40-60 ppm losses)

• Power build up in the range of 13,800 - 19,000

#### For RC mirror $T_i = 120$ ppm, $T_o = 5$ ppm (with 40-60 ppm losses)

• Power build up in the range of 13,300 - 17,600

### For RC mirror $T_i = 45$ ppm, $T_o = 5$ ppm (with 40 ppm losses)

• Power build up of 22,200

For RC mirror  $T_i = 25$  ppm,  $T_o = 5$  ppm (with 20 ppm losses)

• Power build up of 40,000

# **Cavity Scattering loss**

#### How do we acquire substrates with improved scattering losses???

- Coastline claimed they would not have bid had they fully understood the specifications
- All others have also no-bid: Coherent, Gooch and Housego, REO, and ATF
- Need LIGO Test mass quality substrates in terms of surface quality
  - LMA???