

TB 2019 and 2020 data analysis

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LUXE meeting
June 18, 2020

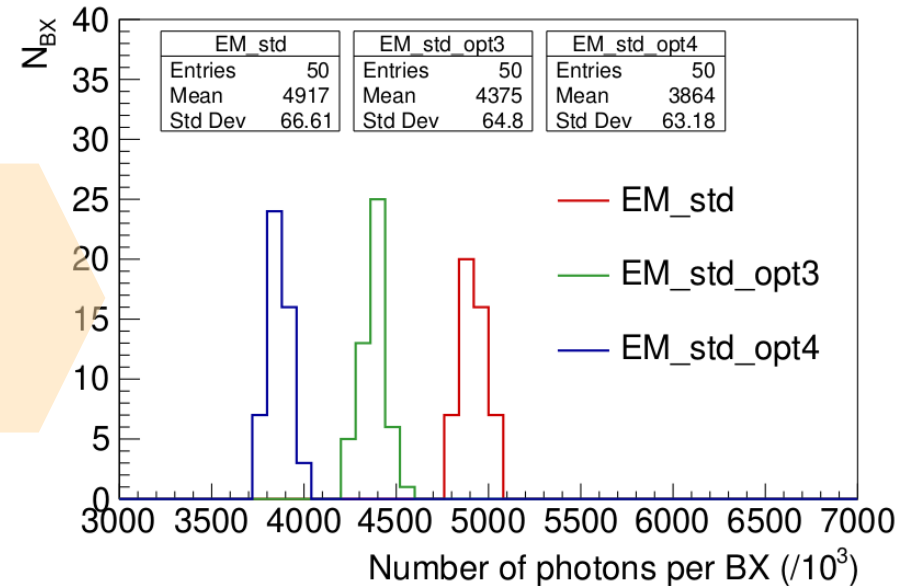
Geant4 simulation with different physics lists

- Gaussian beam, focused on IP;
- Tungsten target 1%X0 (35 μ m) thickness
- 5 m from IP;
- 6.25 M electrons (BX/1000);
- Production cut: 1 μ m.

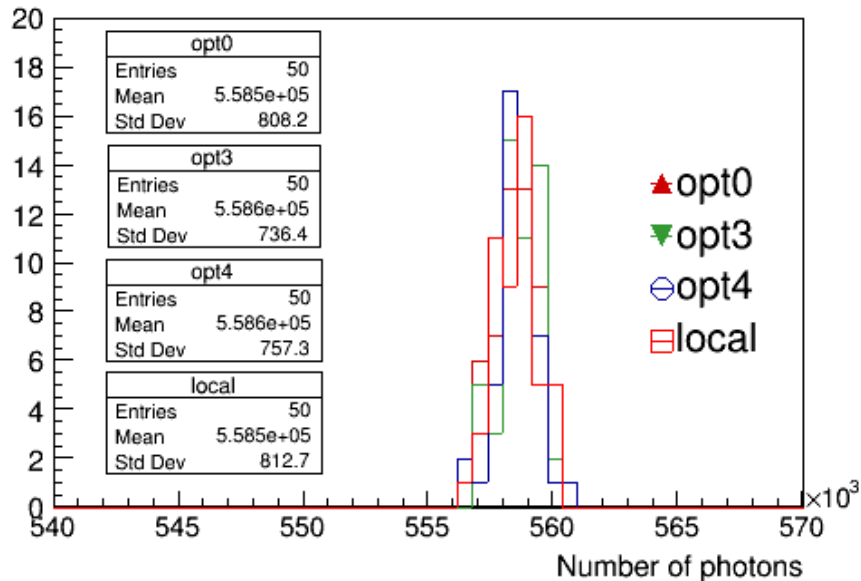
- Angular distribution is the widest for option_4 physics list and the narrowest for option_0.
- Total number of photons in forward region is identical for all physics lists.

Number of photons inside $|x| < 25\mu\text{m}$ and $|y| < 25\mu\text{m}$;

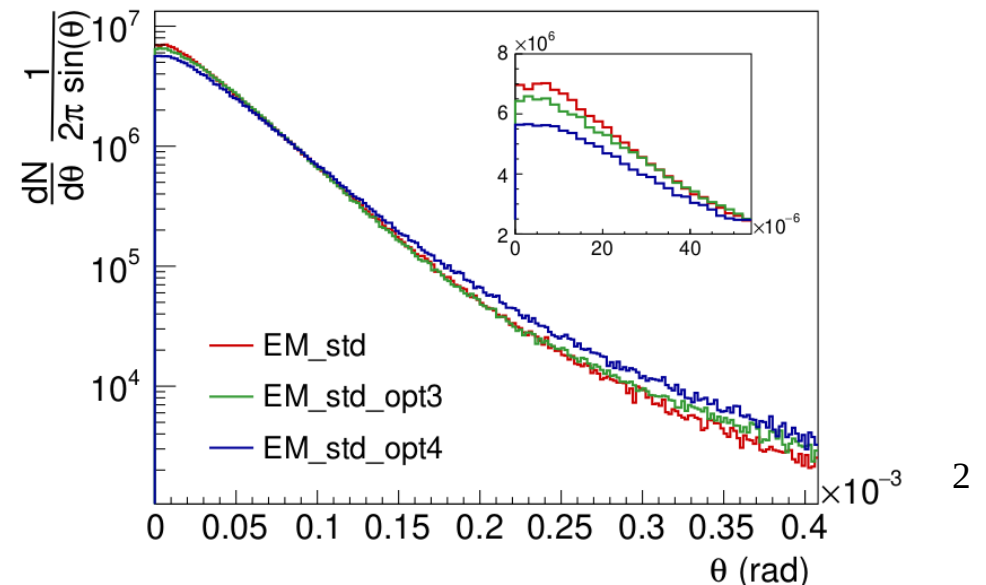
Different physics lists



Number of photons in wide range of θ



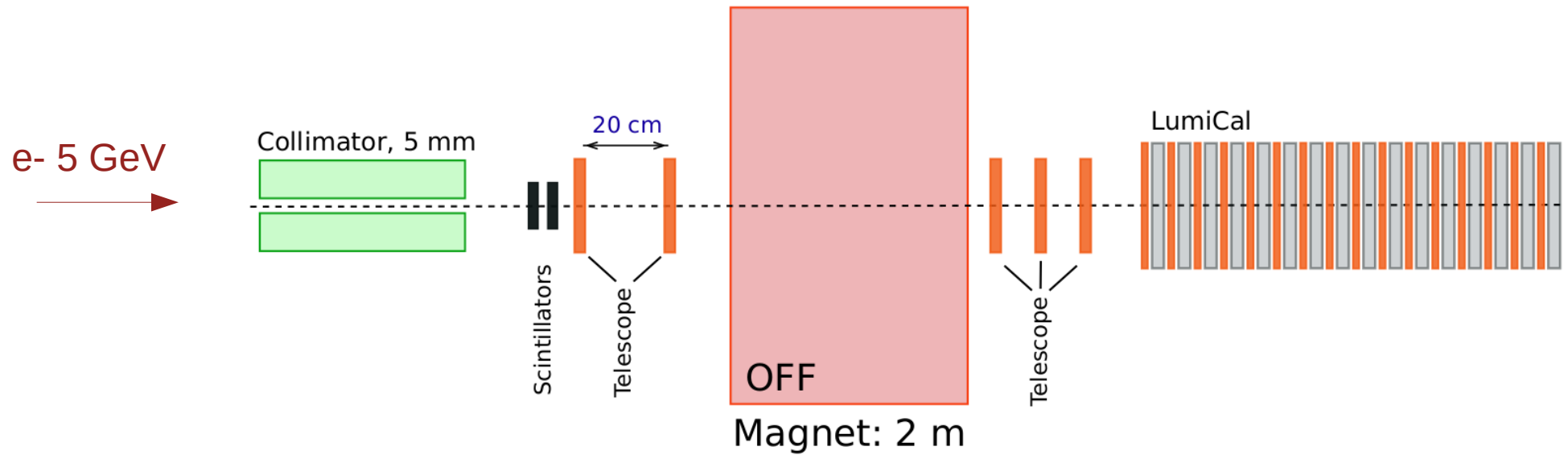
Angular distribution of photons



Outline

- Telescope planes alignment and track reconstruction performance;
 - 5 planes;
 - 3 planes;
 - 4 planes;

Setup 1



- Measure the effect of the air $\sim 2 \text{ m}$.
- Collimator with 5 mm square cross section?

Data processing

- Data converter from raw format to LCIO
- Eutelescope software. It uses ILC software:
 - for geometry settings (GEAR)
 - Marlin (Modular Analysis and Reconstruction for the LINear Collider) for data processing;
 - LCIO for input/output;
 - Converting data to root format;
 - Alignment and track reconstruction.

Noisy pixels (default settings for threshold)

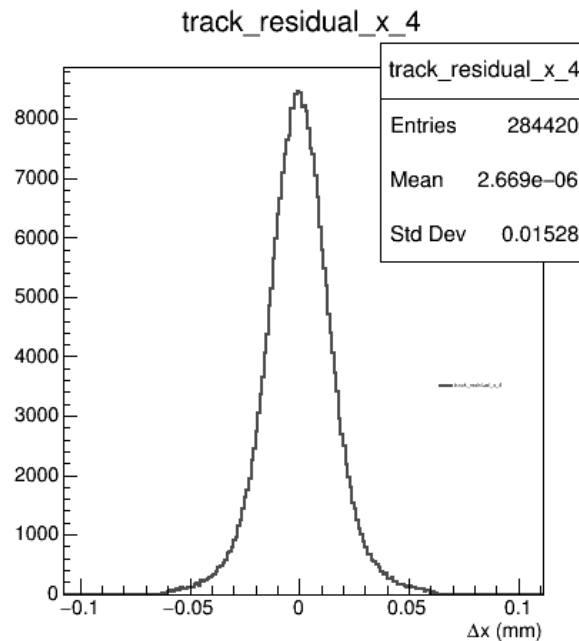
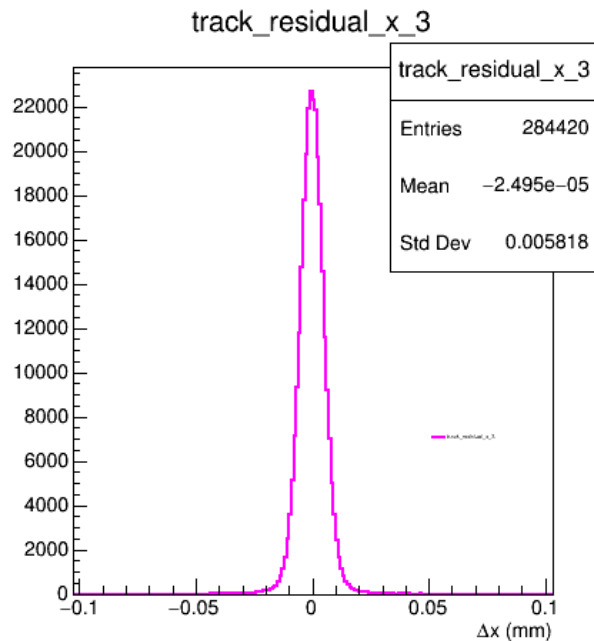
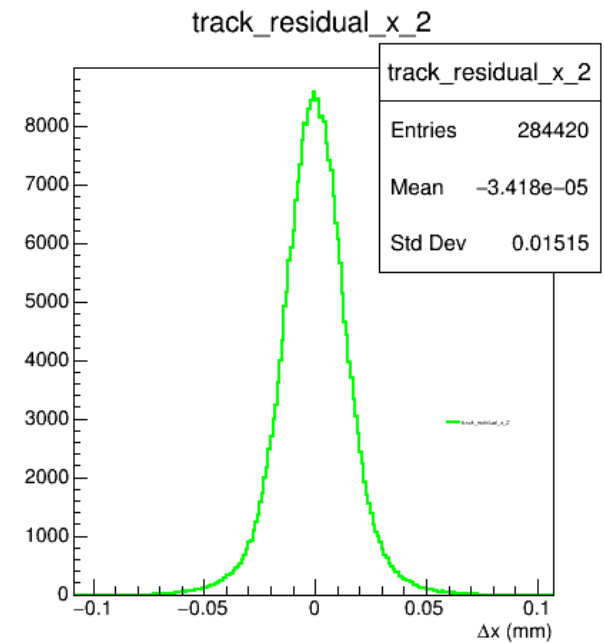
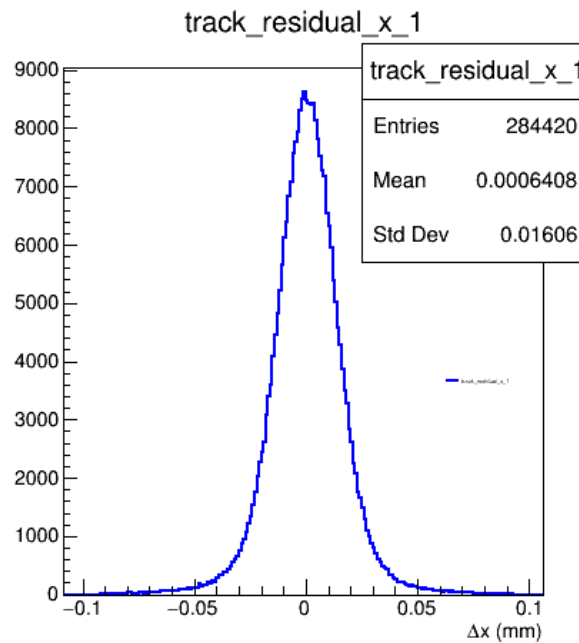
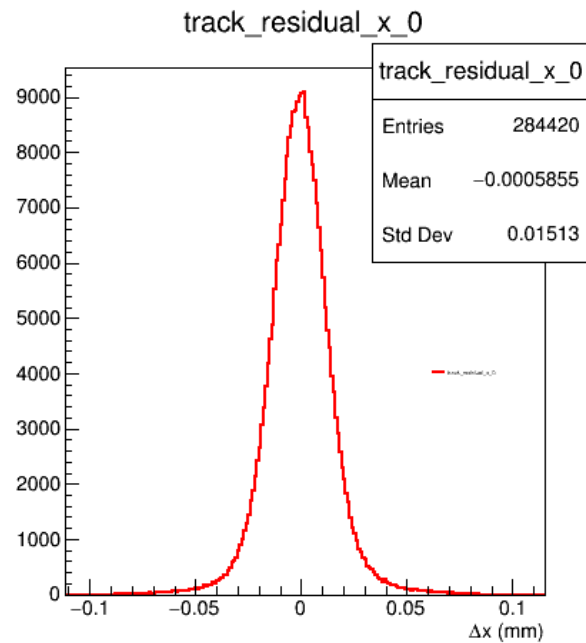
Run 49

```
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jobsub.noisypixel(INFO): [ MESSAGES "HotPixelMasker"] Found 0 noisy pixels on sensor: 2
jobsub.noisypixel(INFO): [ MESSAGES "HotPixelMasker"] Found 1 noisy pixels on sensor: 3
jobsub.noisypixel(INFO): [ MESSAGES "HotPixelMasker"] Found 0 noisy pixels on sensor: 4
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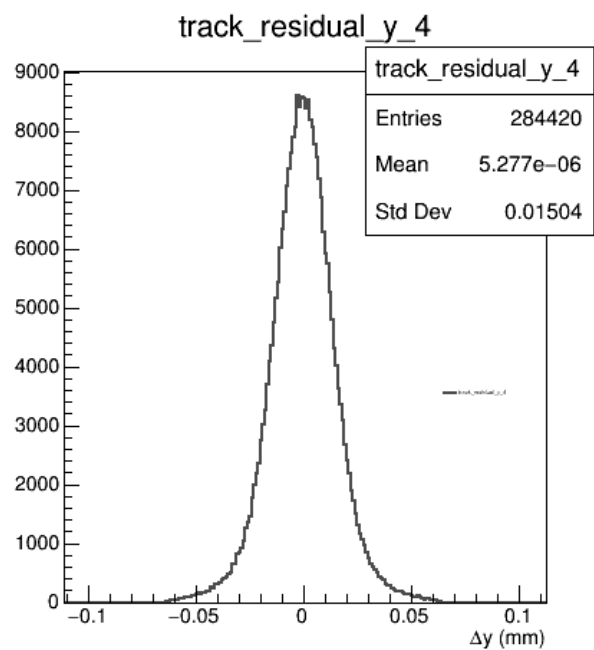
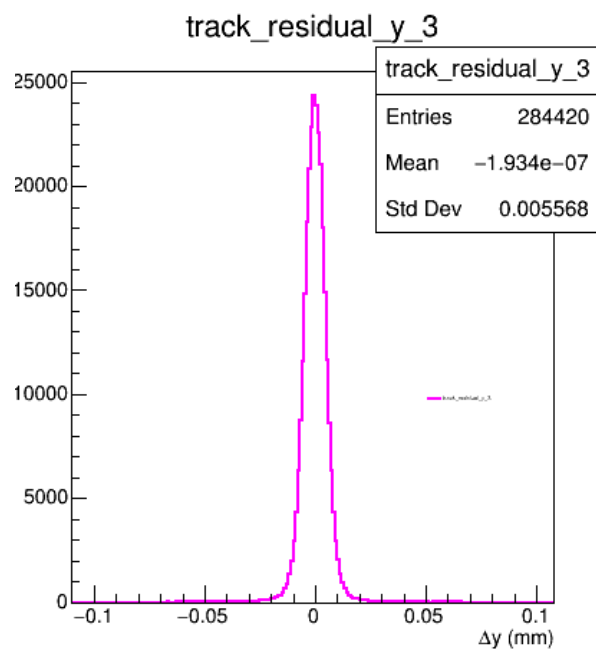
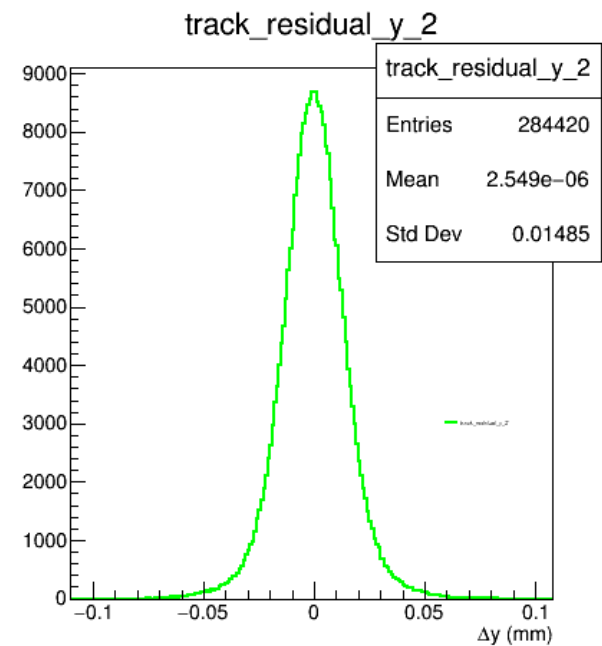
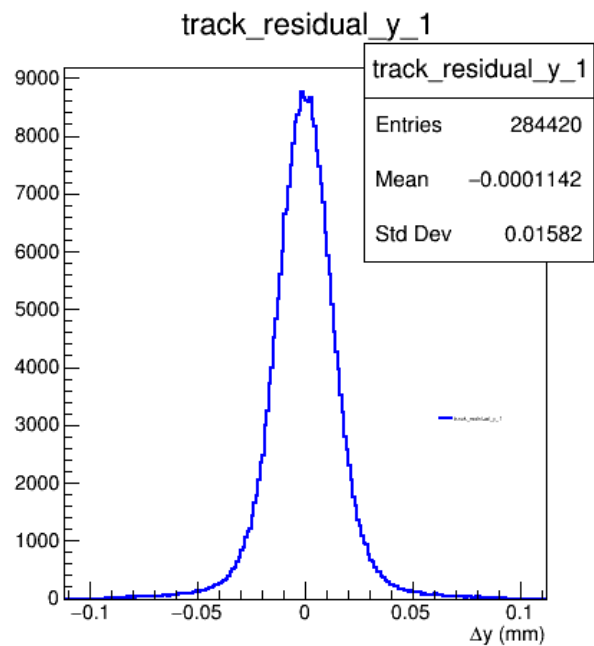
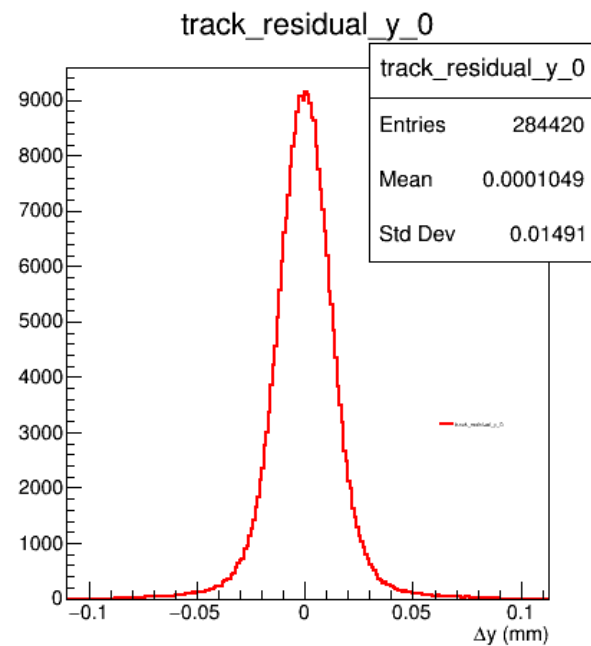
Run 60

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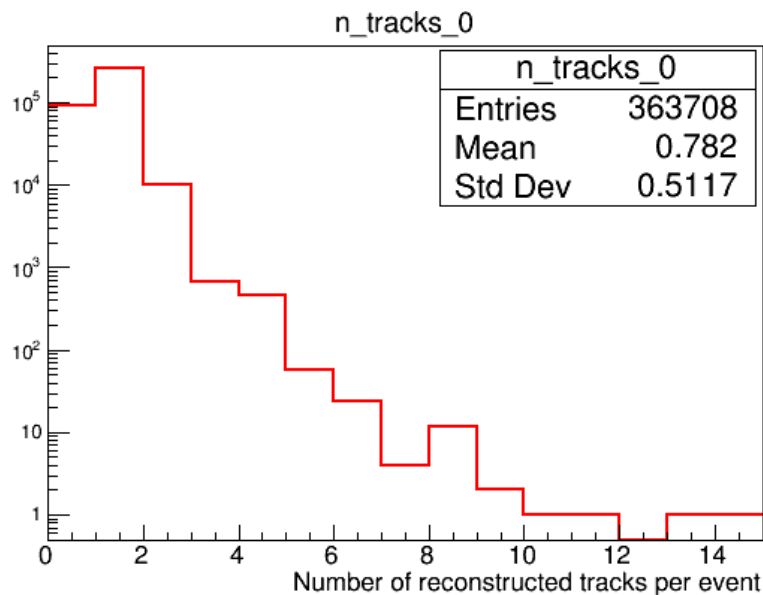
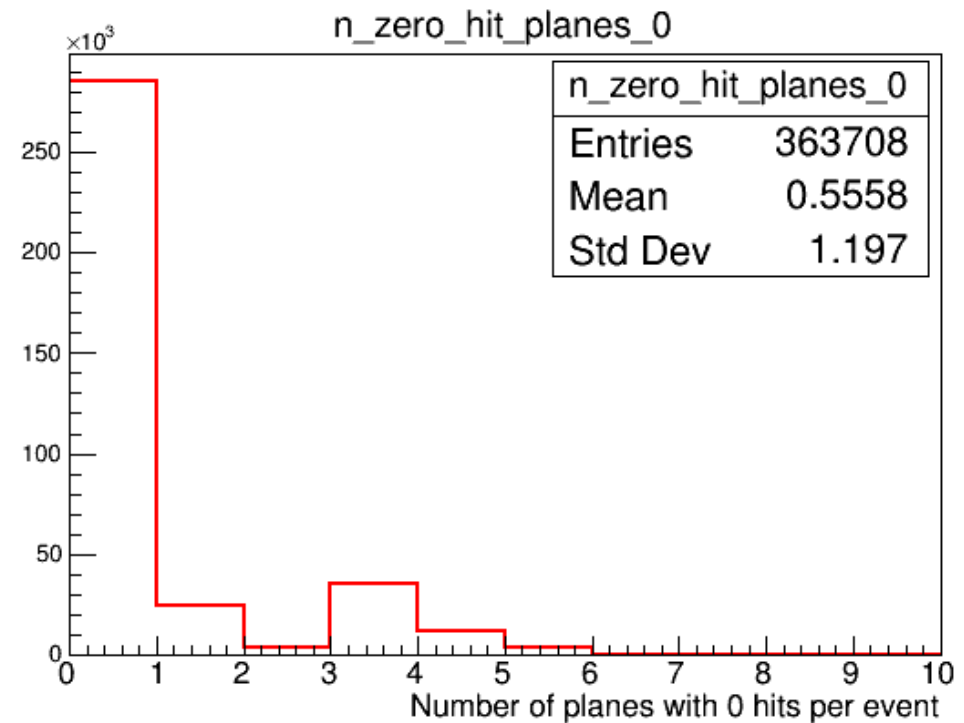
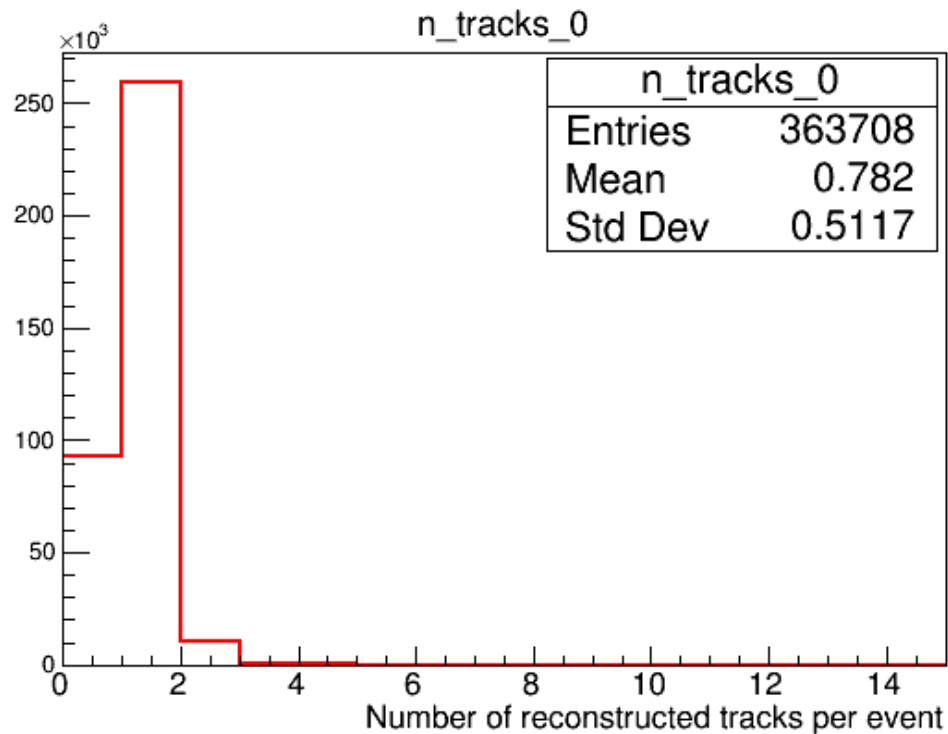
Run 60, Single track fit for all 5 planes



Run 60, Single track fit for all 5 planes

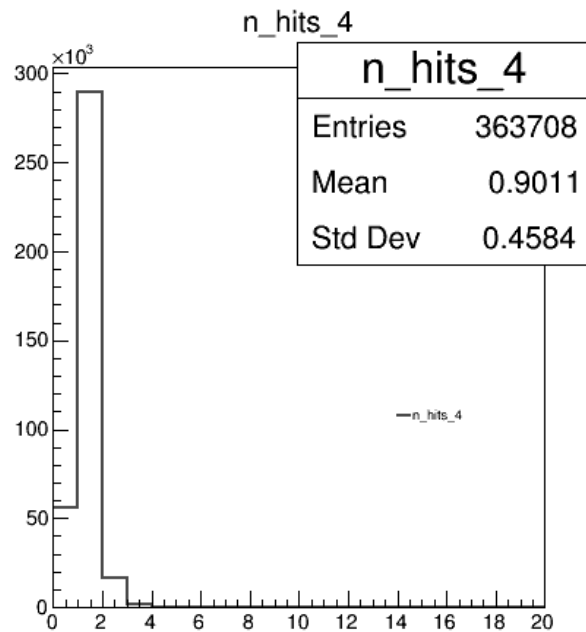
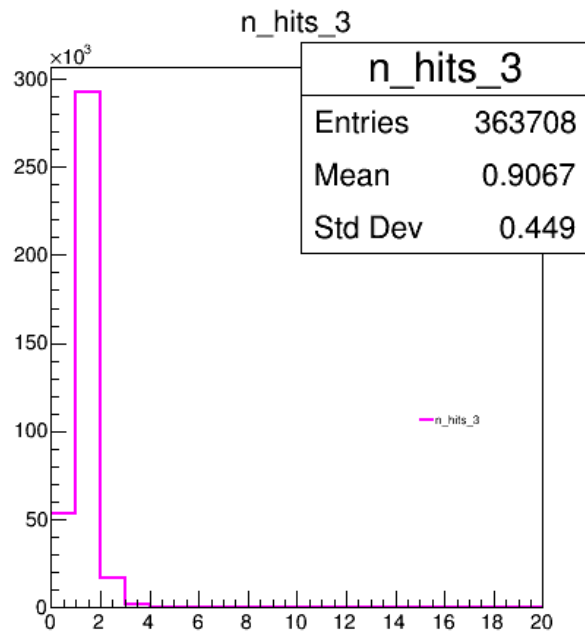
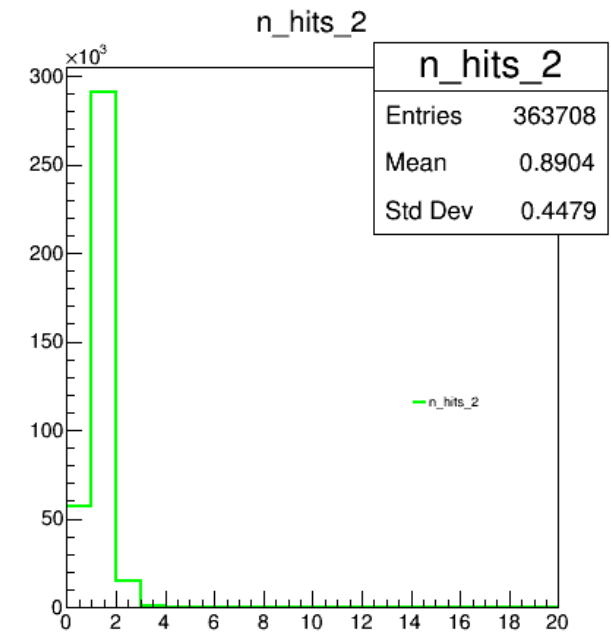
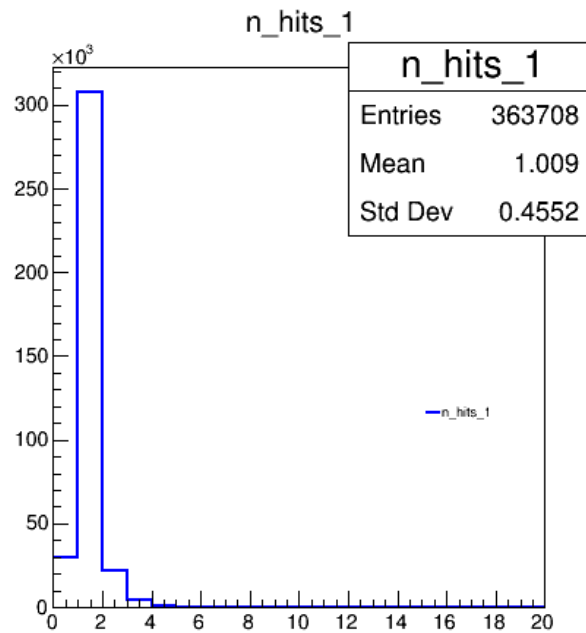
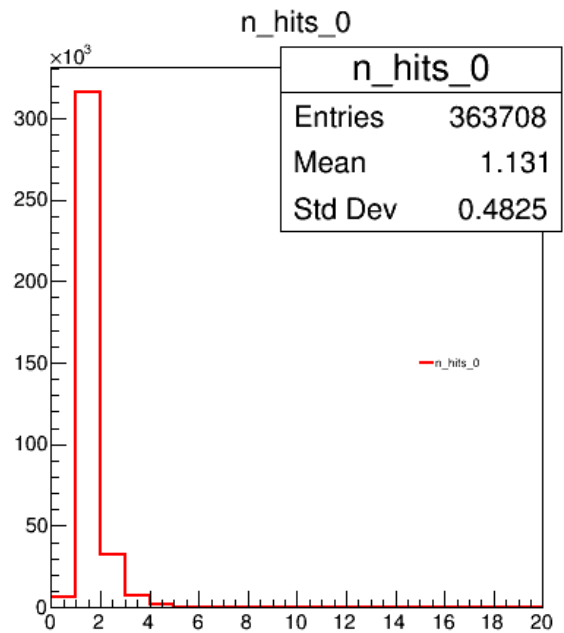


Number of reconstructed tracks

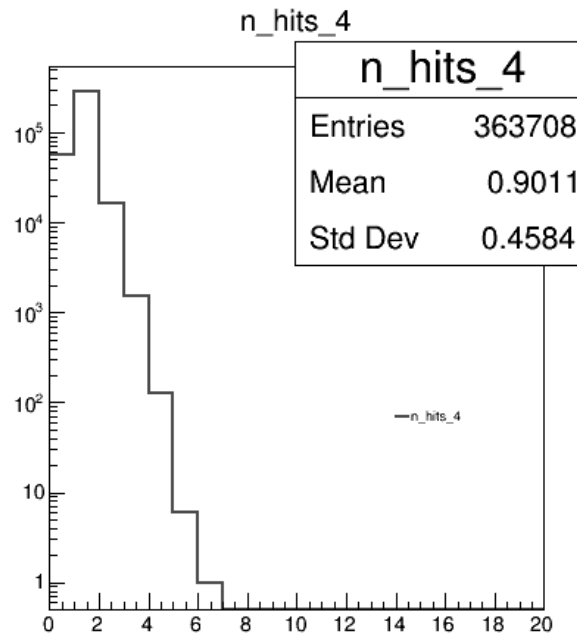
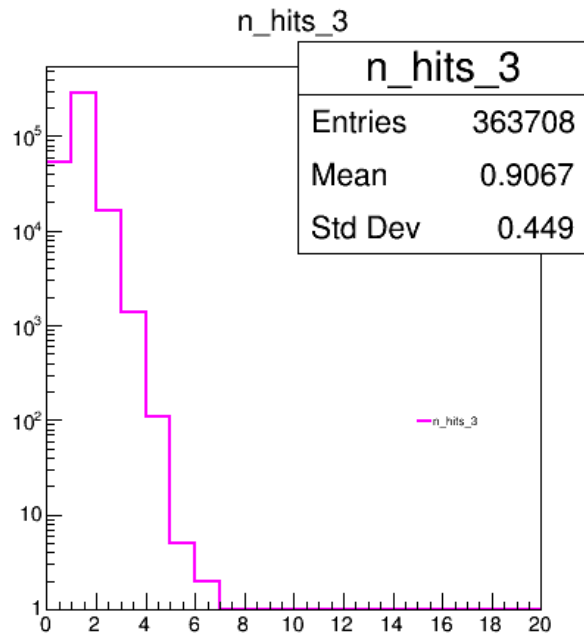
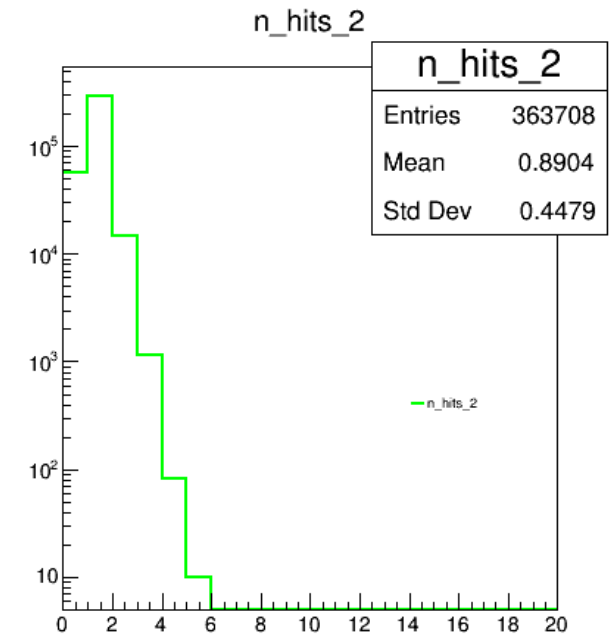
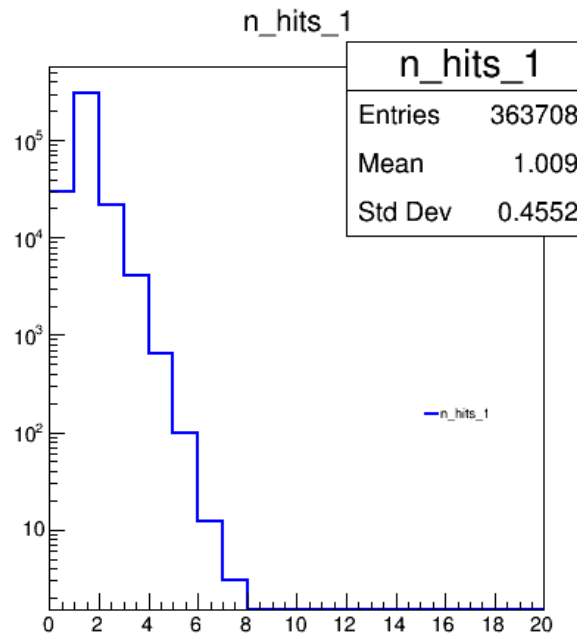
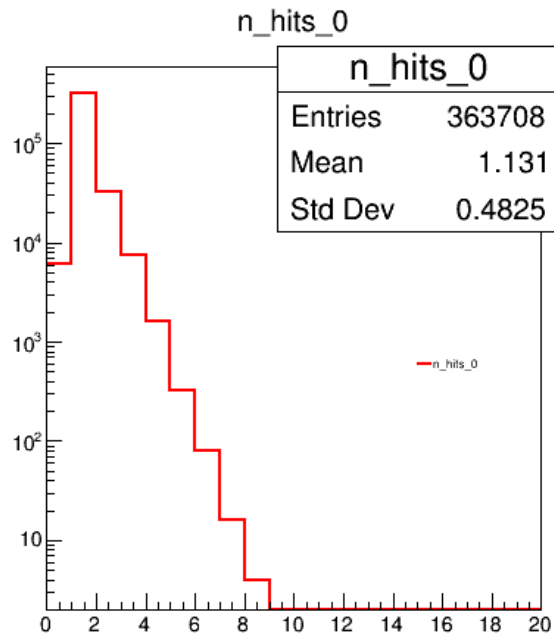


- Number of reconstructed tracks: 284420;
- Number of events with at least one plane with 0 hits: 78640;
- $363708 - 78640 = 285068$;
- $285068 - 284420 = 648$ - number of events without reconstructed track but with hits in all planes.

Number of clusters per plane

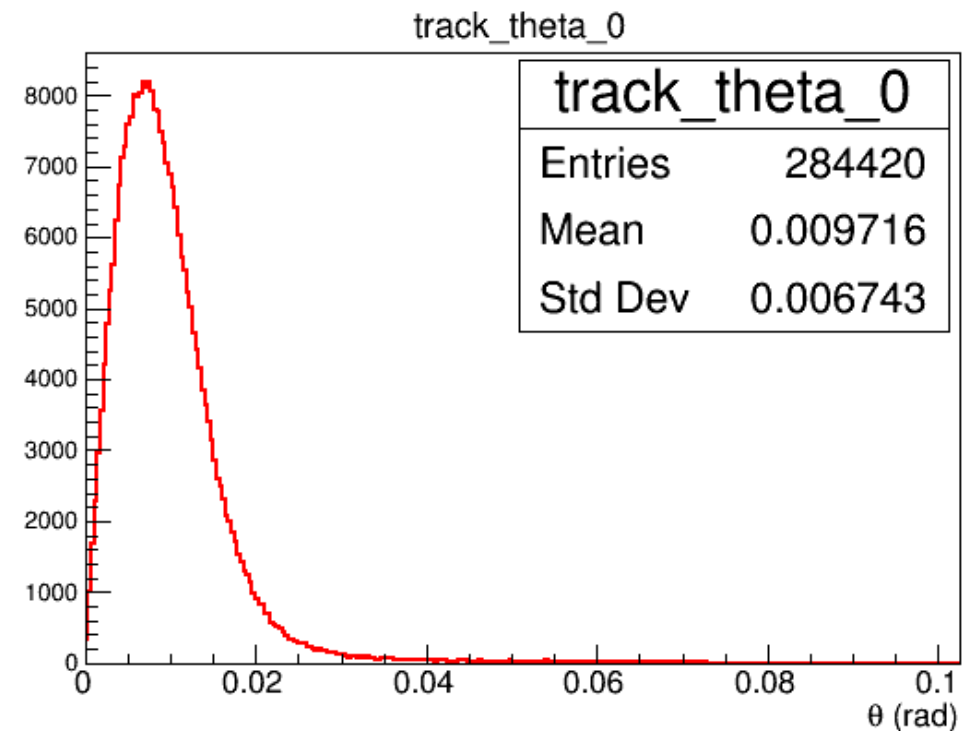
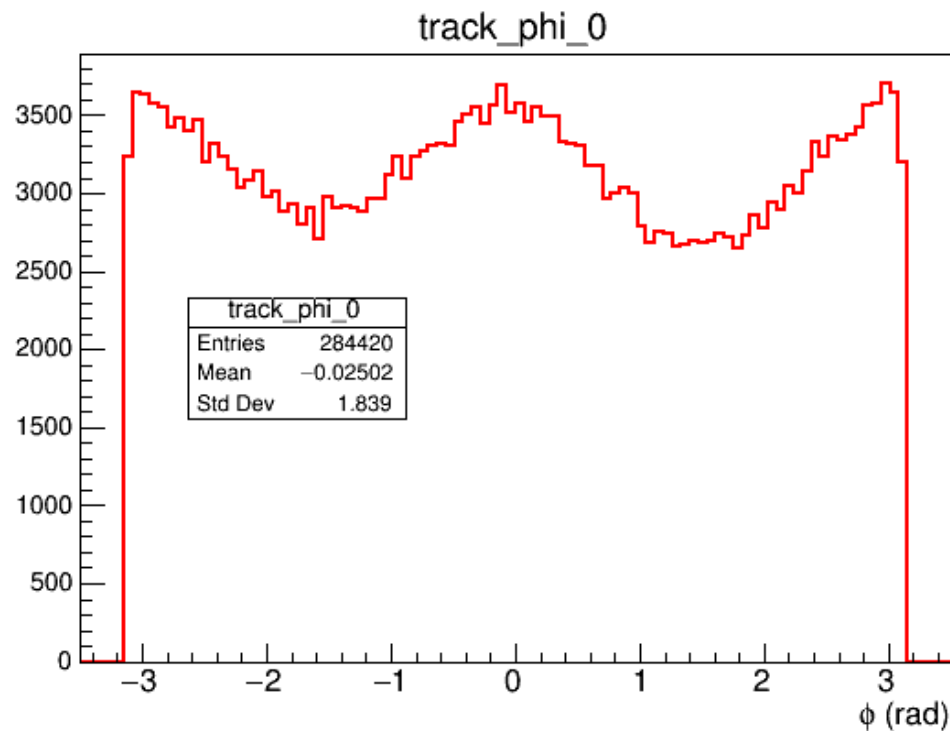


Number of clusters per plane

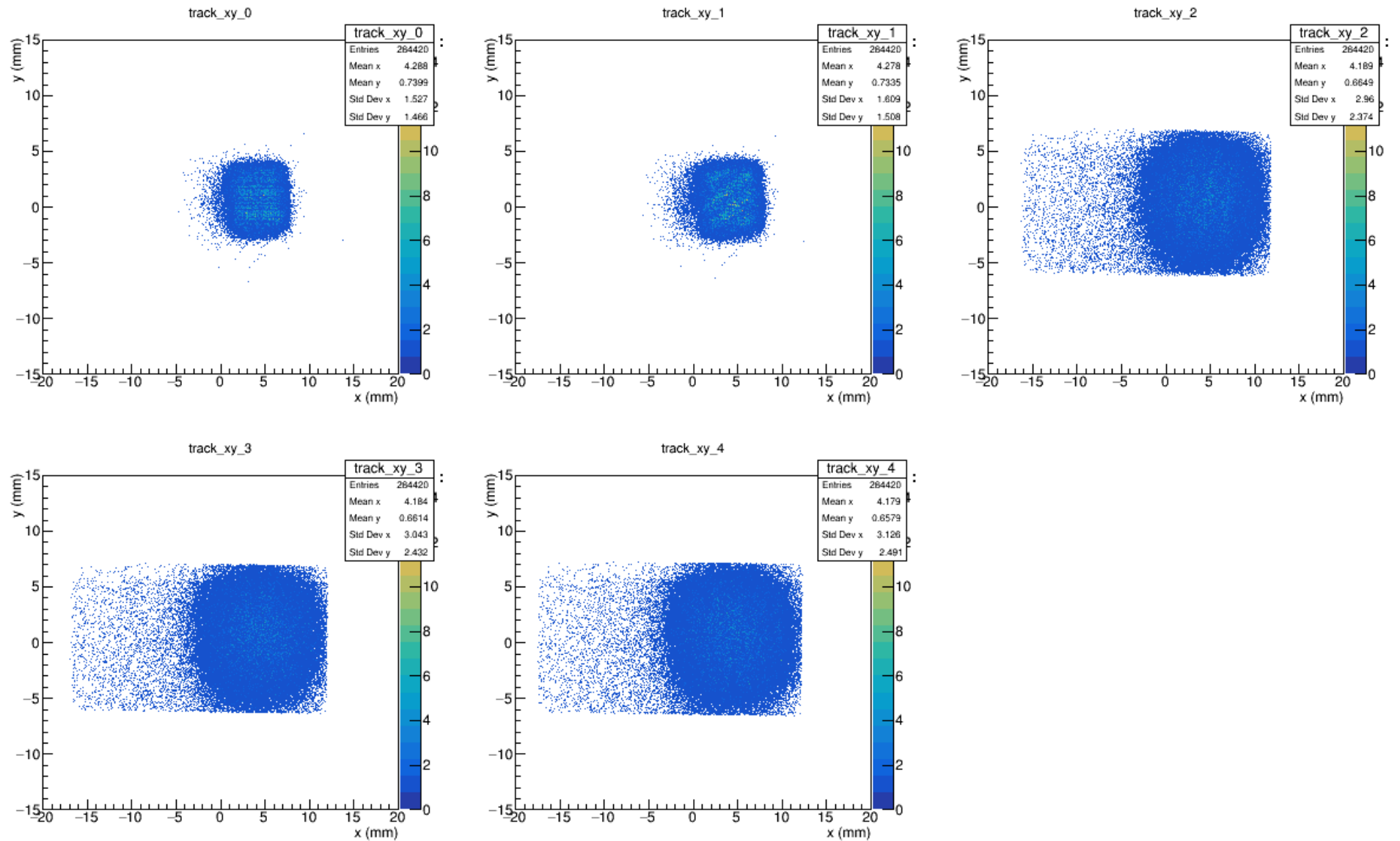


- Seems to be too many events with many (4+) tracks reconstructed;
- Need to check it.

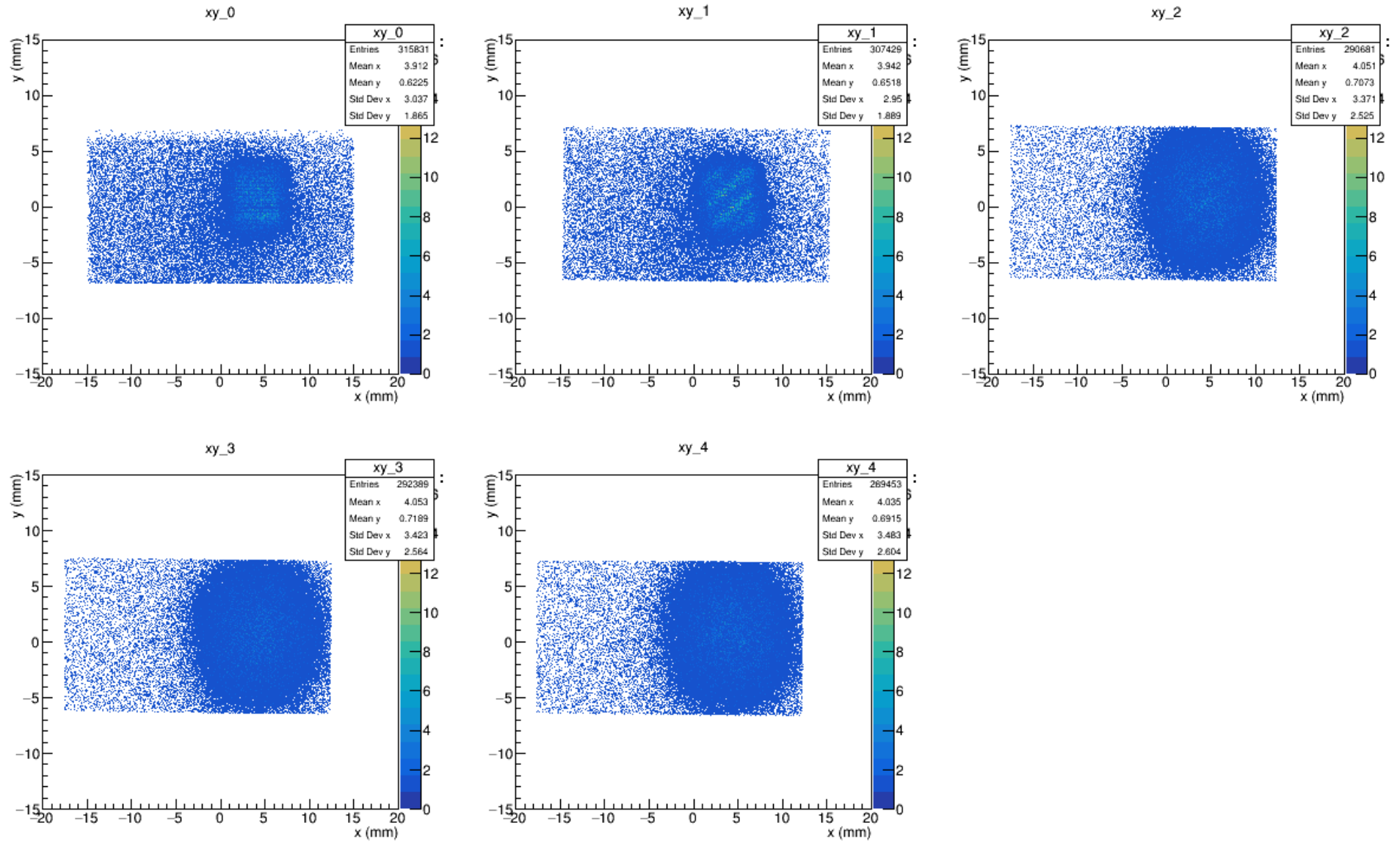
Angular distributions of the tracks



Hits assigned to tracks

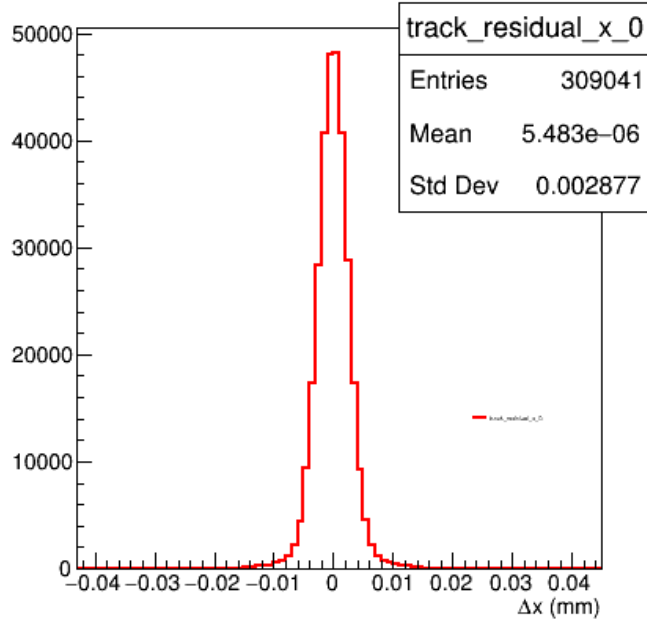


All hits

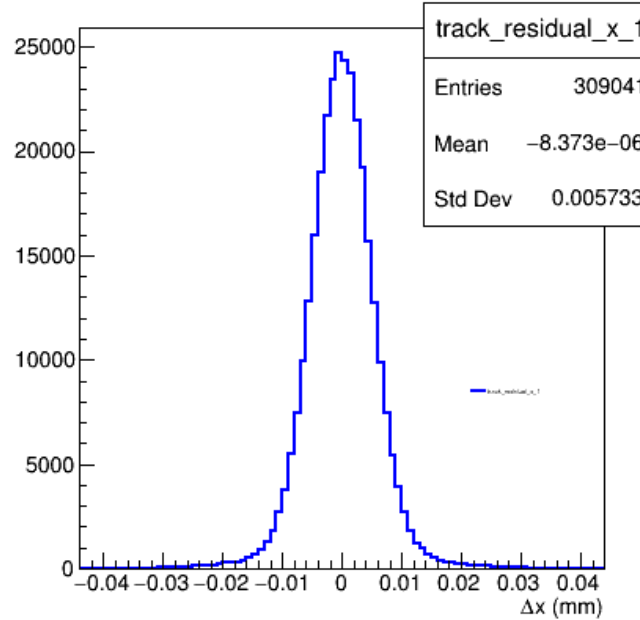


Track reconstruction in three planes after the magnet

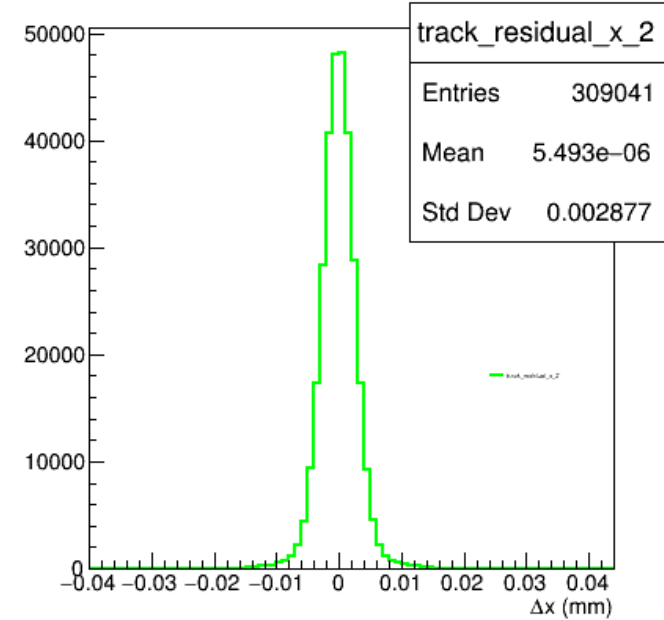
track_residual_x_0



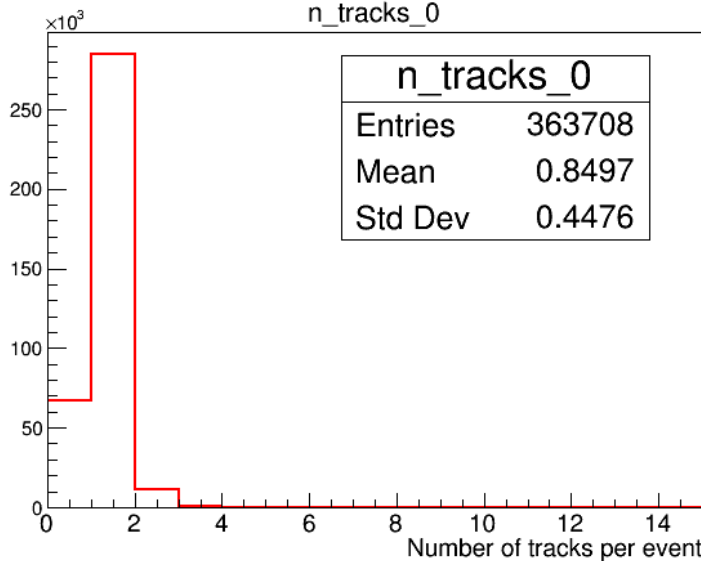
track_residual_x_1



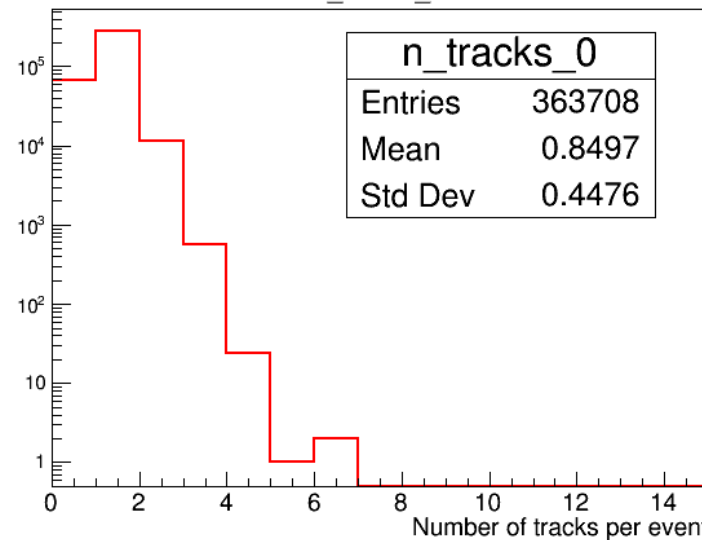
track_residual_x_2



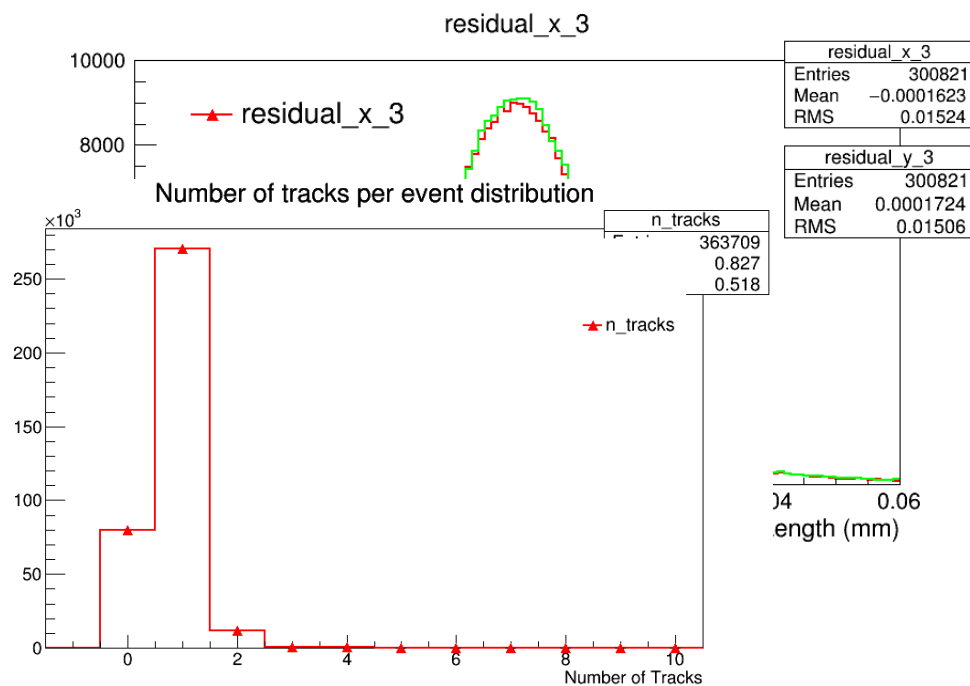
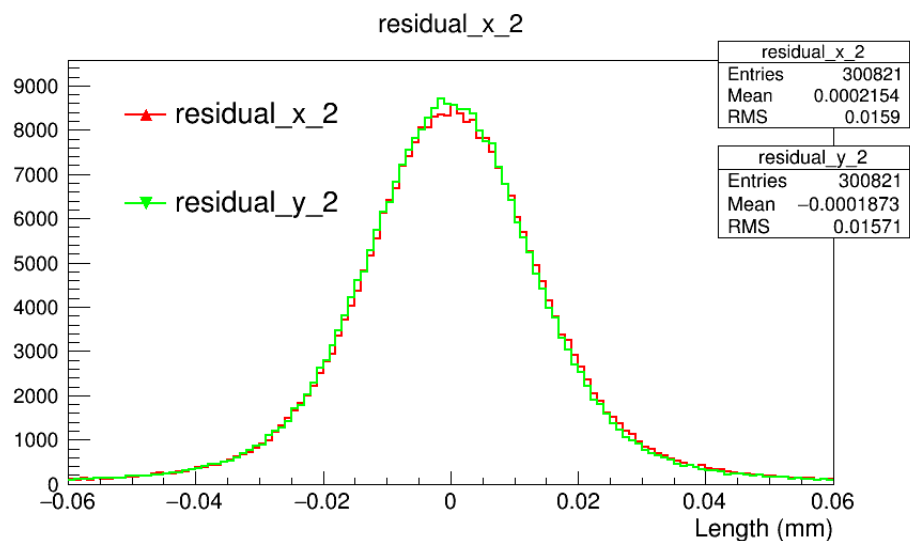
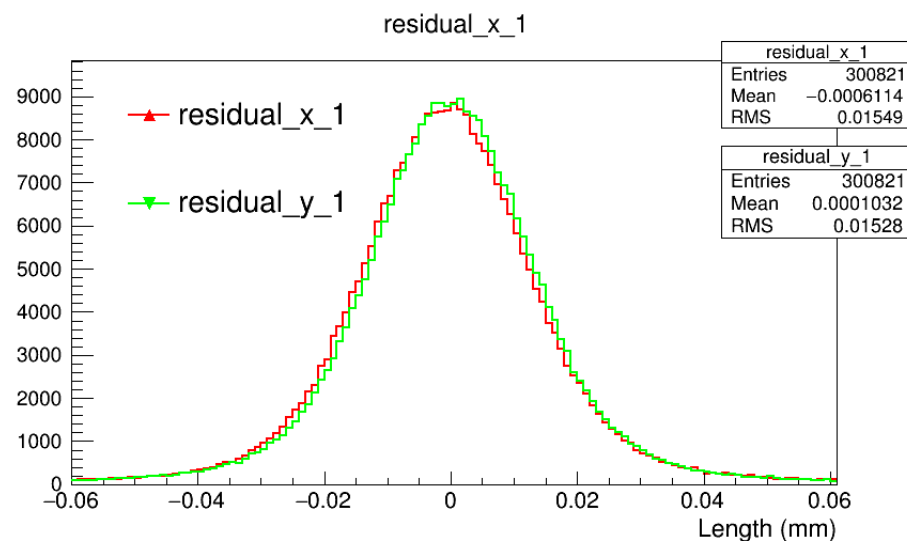
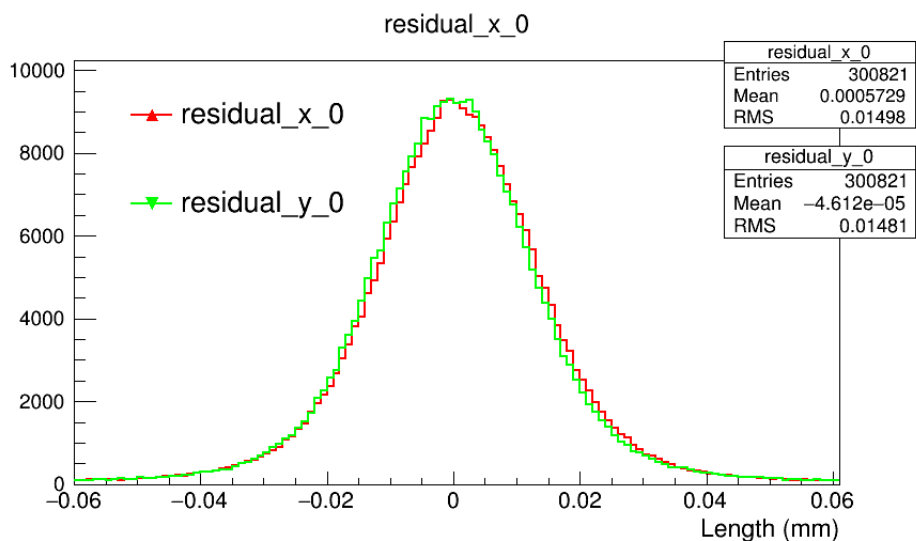
n_tracks_0



n_tracks_0



Test with 4 planes: 1,2 and 3,5



ALPIDE spacial resolution

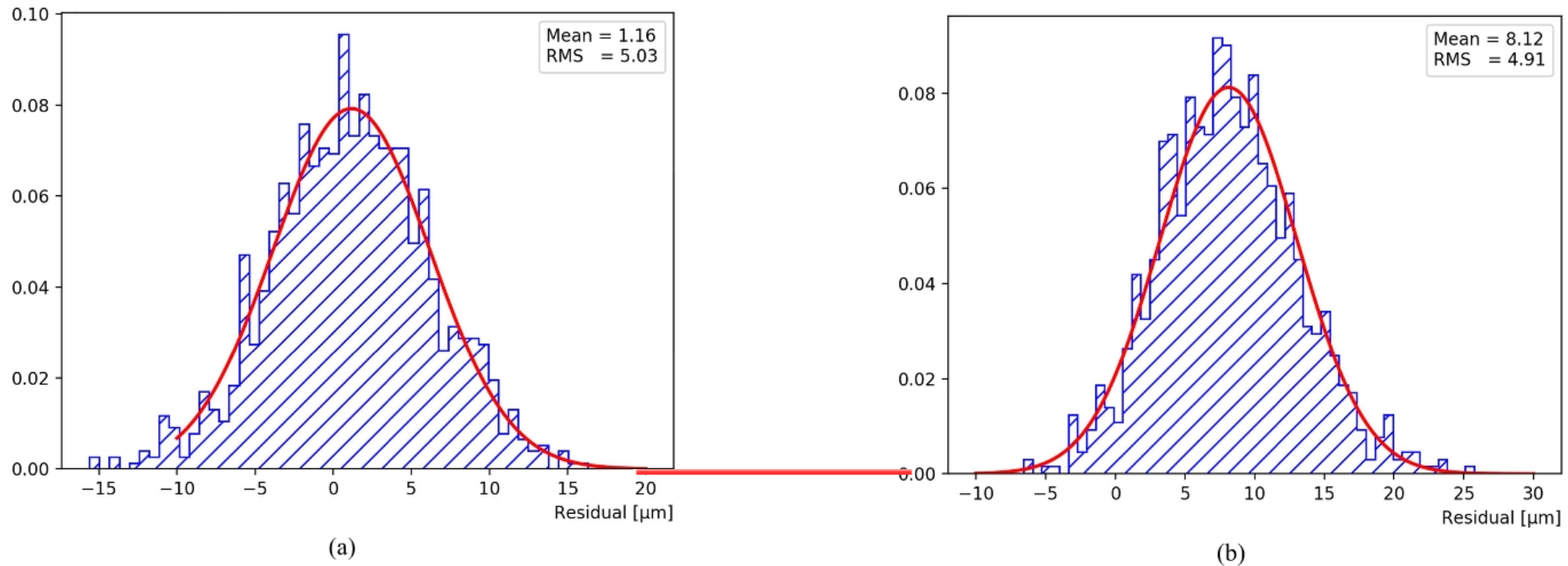


Fig. 17. Residual distribution for stage movements in (a) X-direction and (b) Y-direction. Red line: Gaussian fit results.

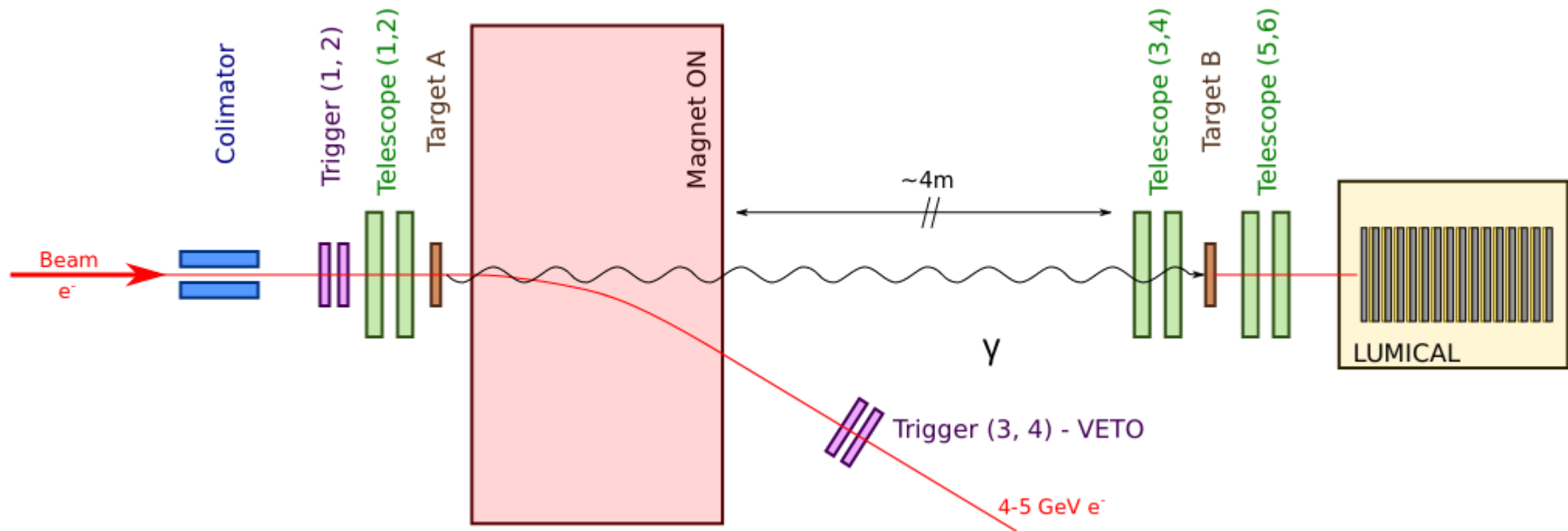
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in the X-direction of the XY stage, and 4.91 μm in the Y-direction. The influence of the XY stage movement resolution was rather small and ignored. The mean values being not zero were caused by the error in the measurement at the origin position. Taking the rotation angle between the ALPIDE chip and the XY stage into consideration, the resolution was calculated to be 5.18 μm in the X-direction of ALPIDE chip where the pixel size is 29.24 μm , and 4.75 μm in the Y-direction where the pixel size is 26.88 μm .

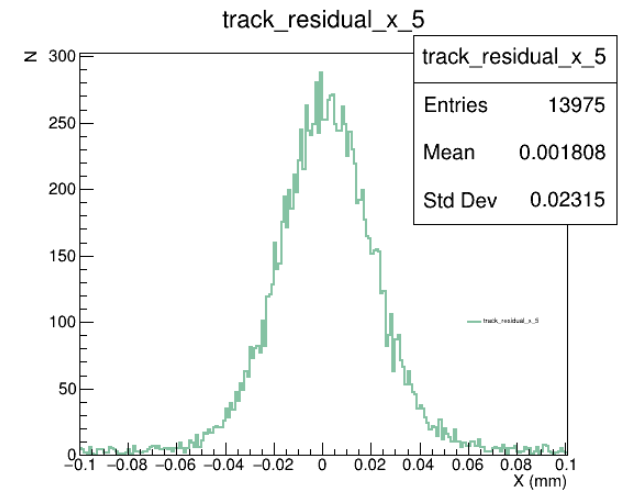
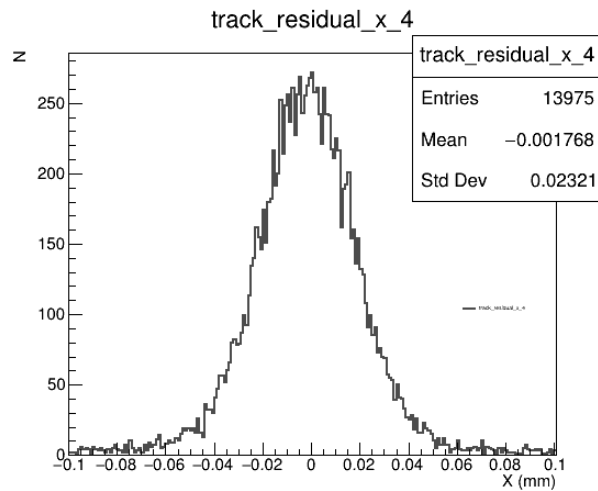
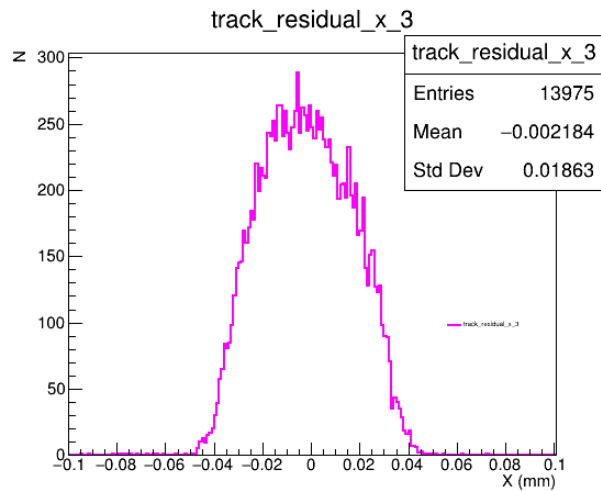
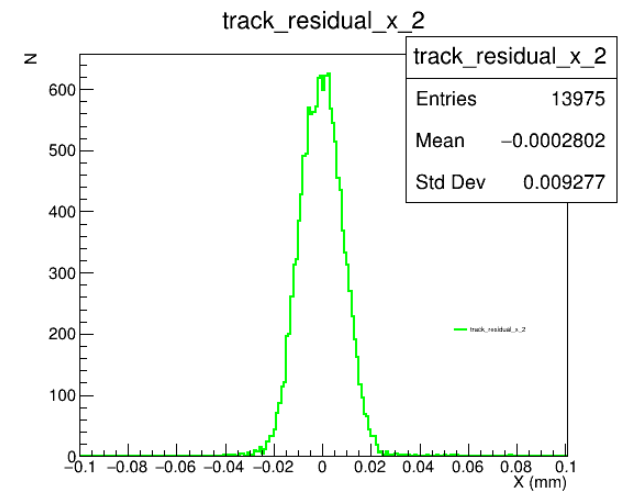
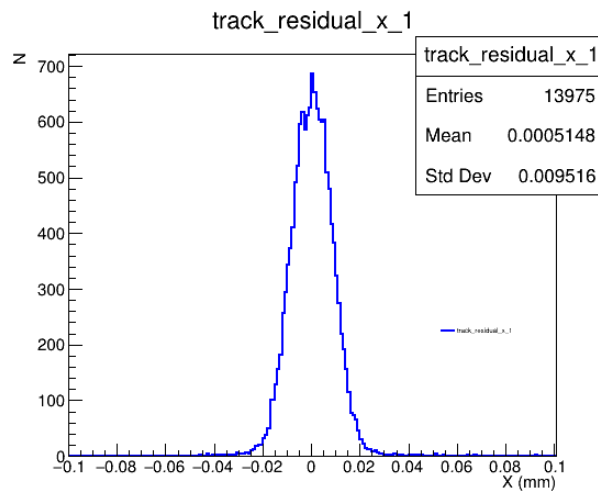
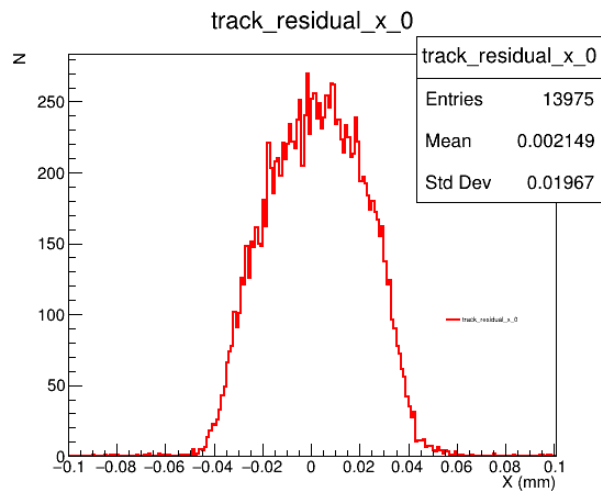
Summary

- Converter for ALPIDE raw data to LCIO works reasonably well.
- Noisy pixel analysis, clustering and hits reconstruction produces reasonable results.
- Alignment procedure converges reasonably well after good prealignment.
- Track reconstruction test for one run (run 60). Look reasonable, but some tuning of reconstruction algorithm parameters are needed.
- Continue with other runs and analysis of scattering angle.
- For TB2020 alignment code has been significantly modified to address ~7m distance between telescope arms.

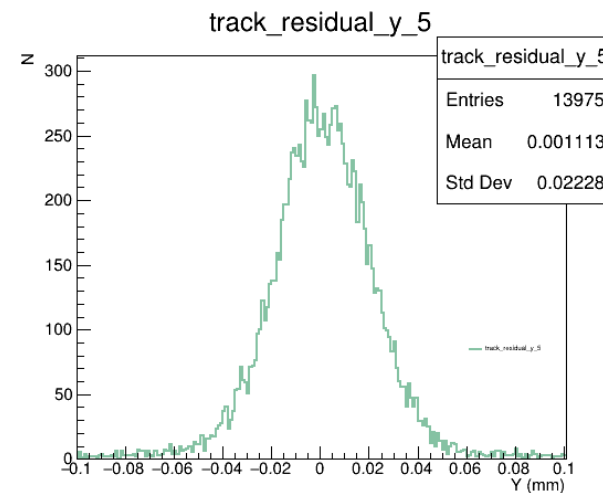
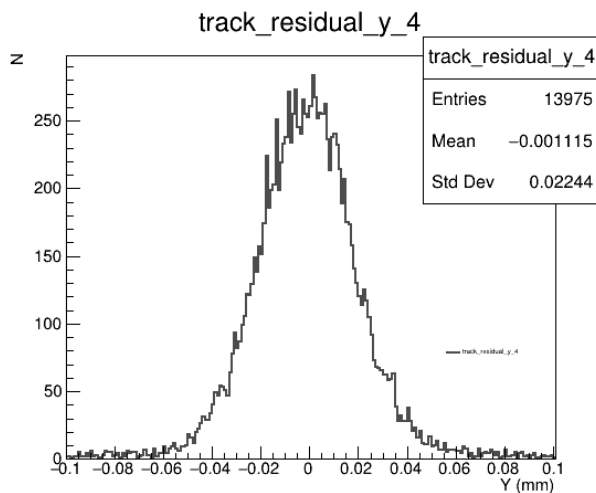
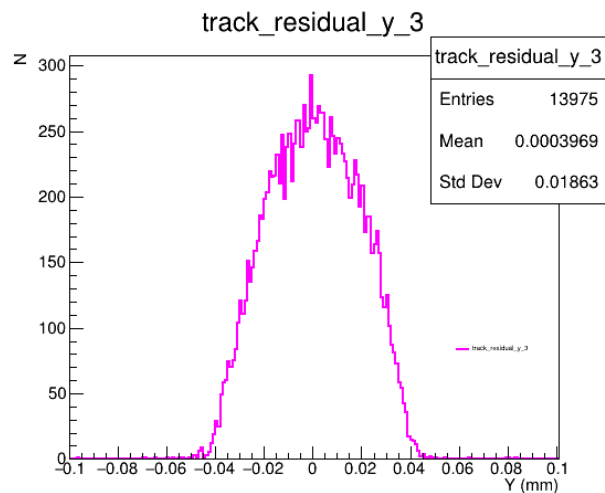
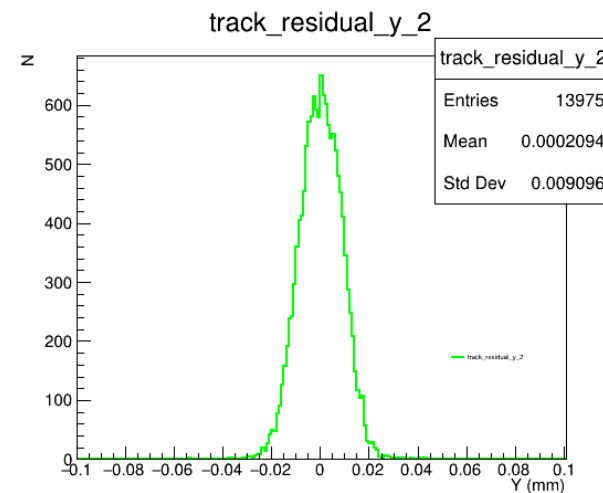
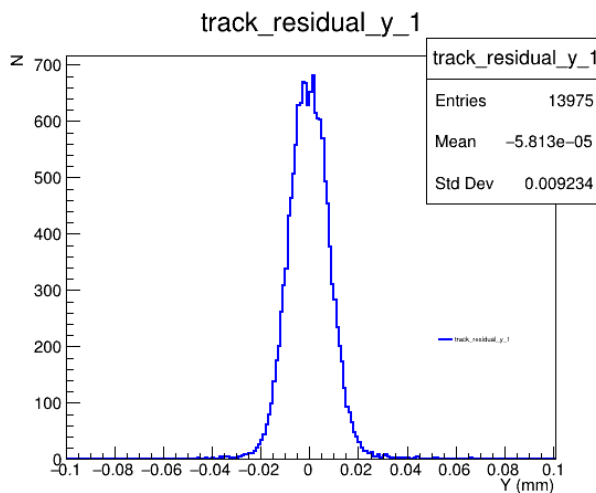
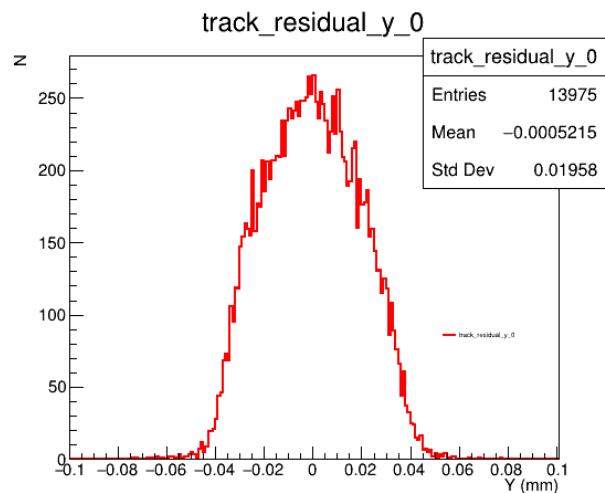
TB2020 Alpide Telescope Alignment



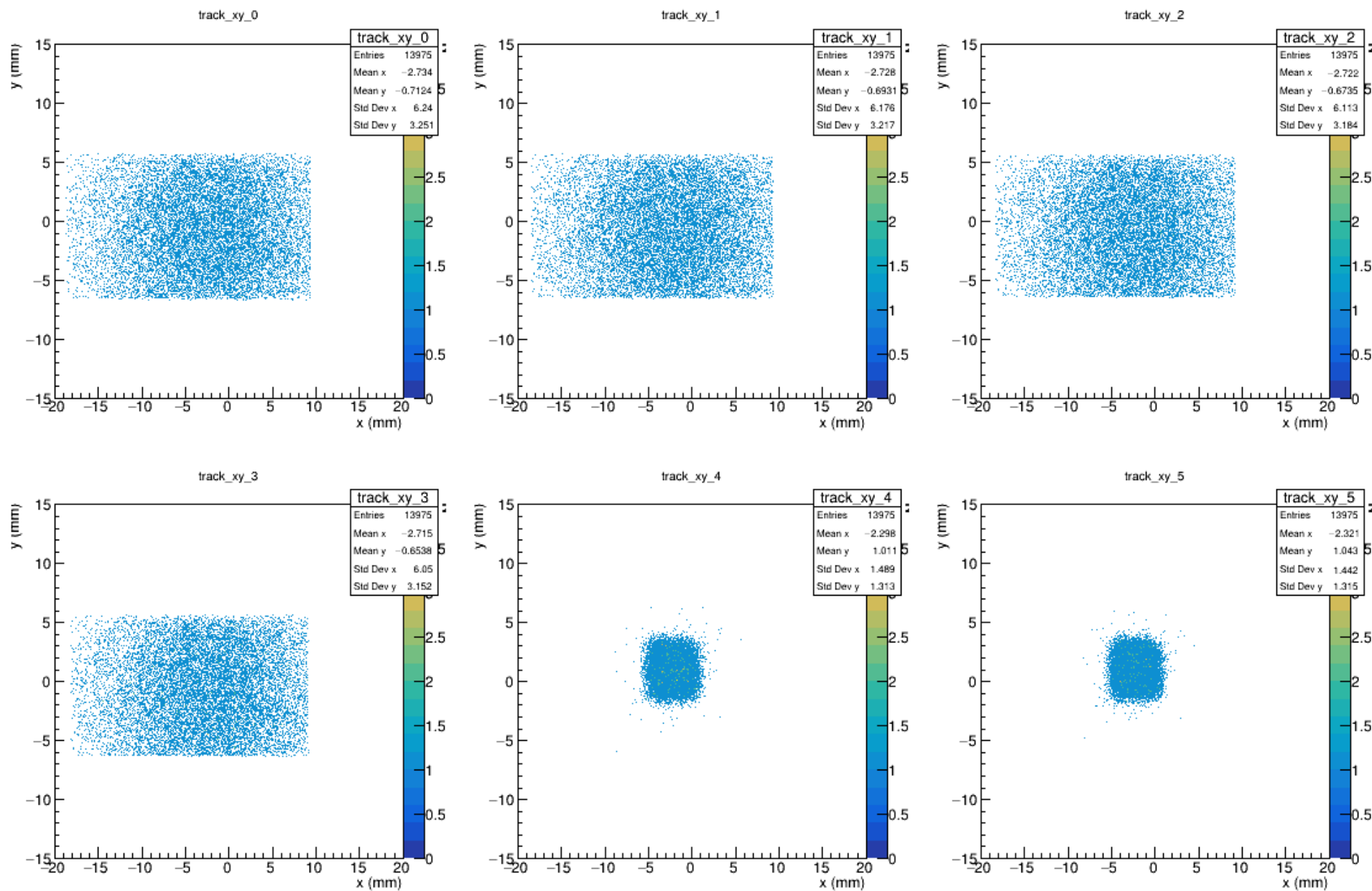
Residuals in planes 5-0



Residuals in planes 5-0

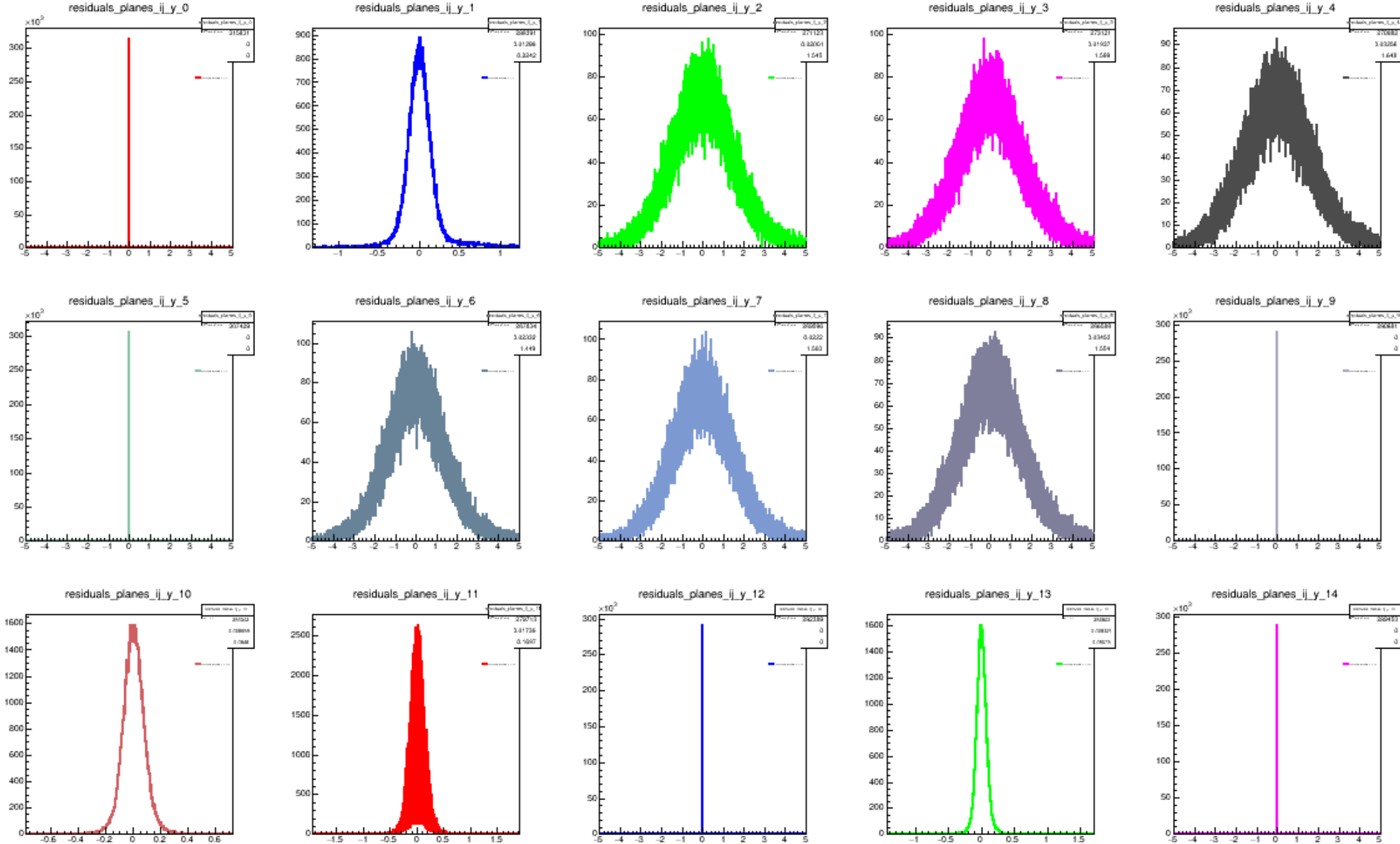


Track hits in planes 5-0



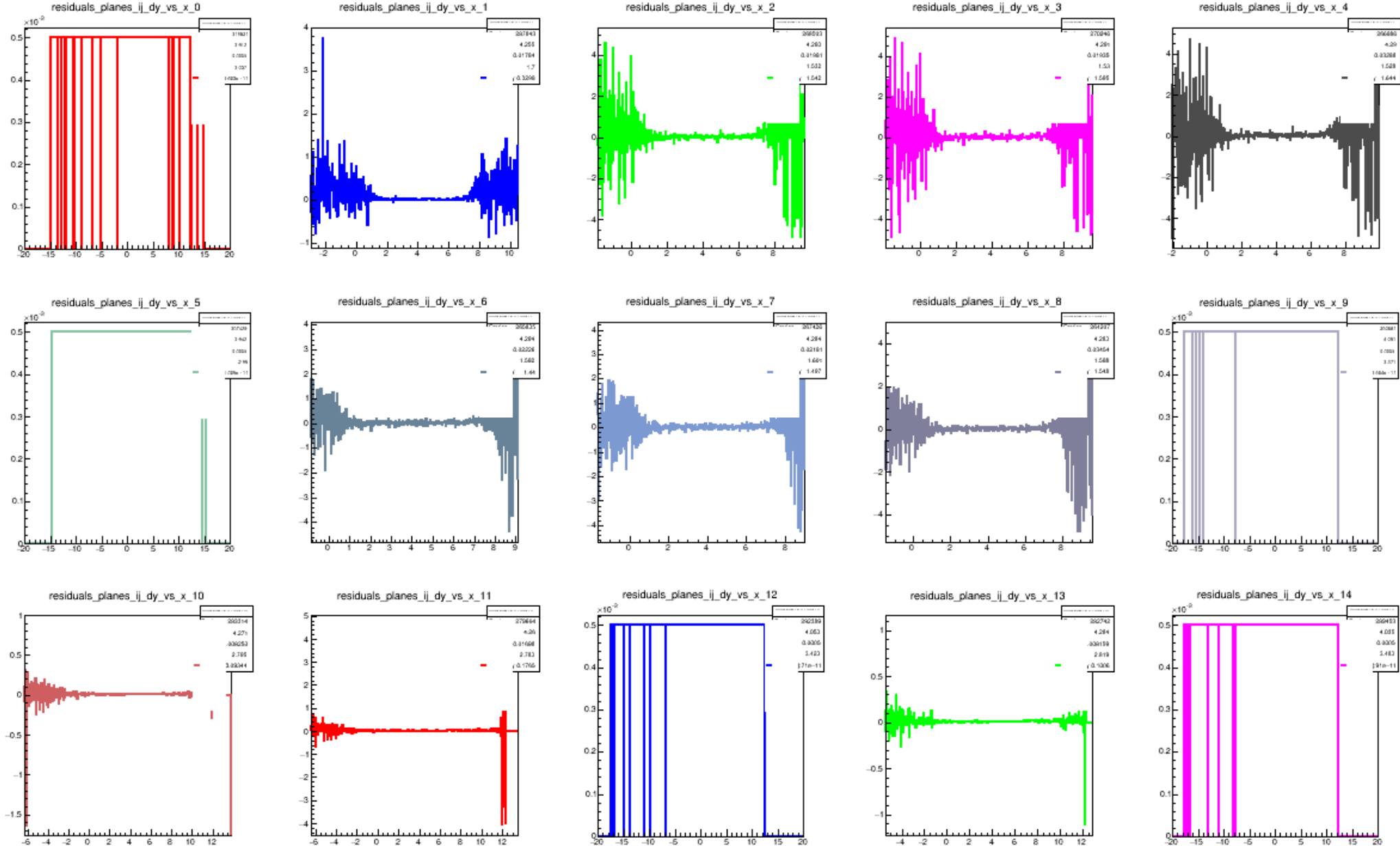
Back up

Prealignment. Y correlations between planes



Prealignment of rotation around Z.

Profiling plot of dy vs x distribution.



Least square fit of line to 3 points

$$\hat{\mathbf{x}} = (A^T A)^{-1} A^T \mathbf{b}$$

Slope is determined by the two outer points

The distance to outer points is twice smaller than to middle one

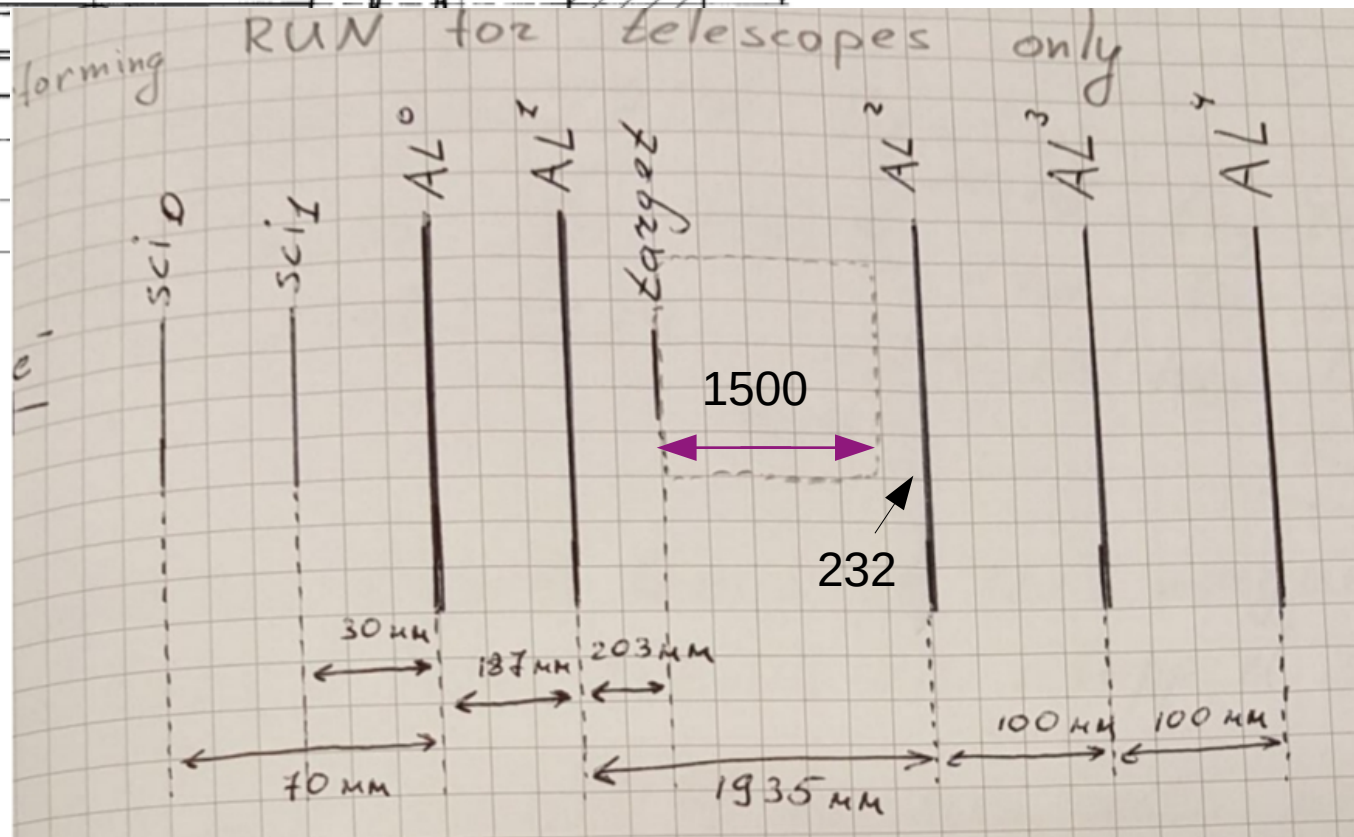
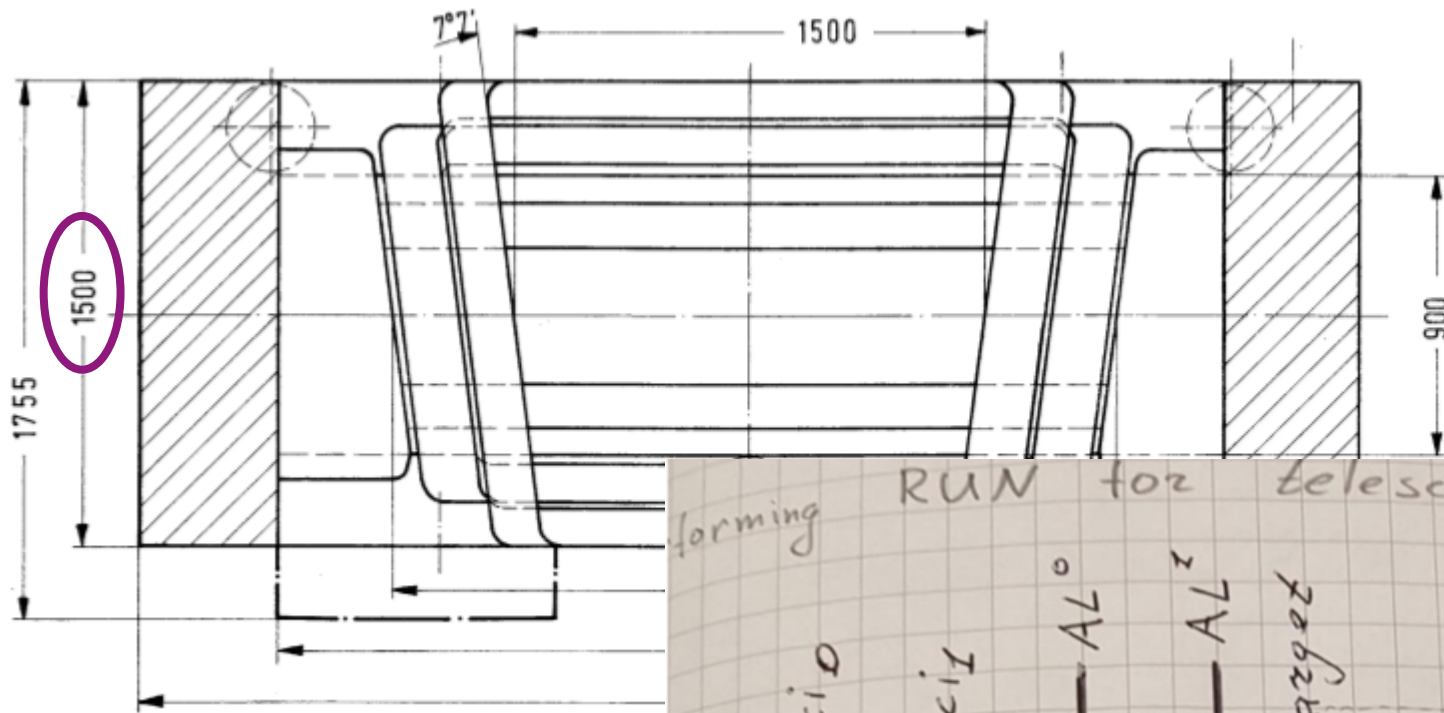
```
(%i1) A: matrix( [x0,1], [x0+d,1], [x0+2*d,1]);
(%o1)  $\begin{bmatrix} x_0 & 1 \\ x_0+d & 1 \\ x_0+2d & 1 \end{bmatrix}$ 

(%i2) Y: matrix( [y0], [y1], [y2]);
(%o2)  $\begin{bmatrix} y_0 \\ y_1 \\ y_2 \end{bmatrix}$ 

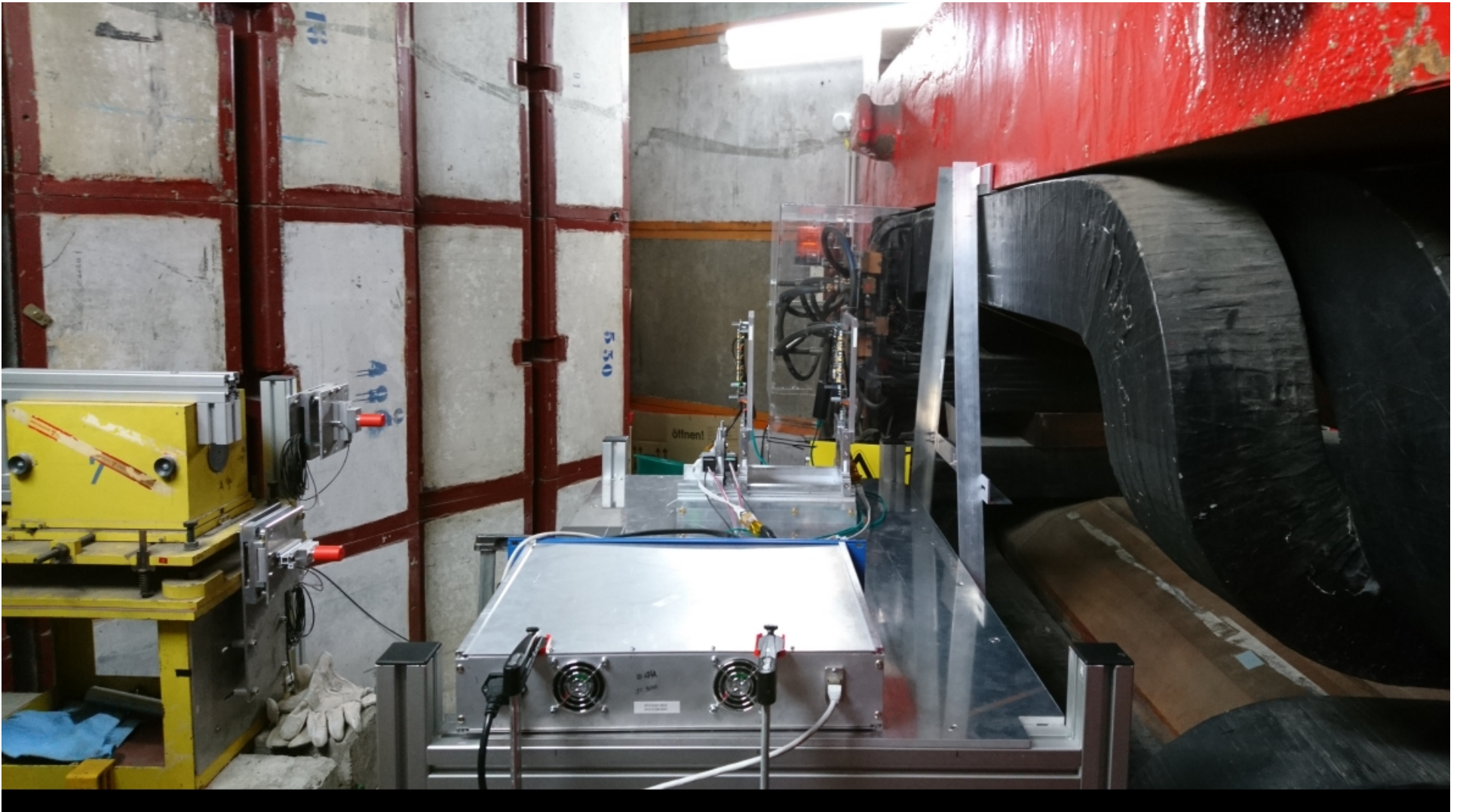
(%i3) B: ratsimp(invert(transpose(A) . A) . transpose(A) . Y);
(%o3)  $\begin{bmatrix} \frac{y_2 - y_0}{2d} \\ -\frac{(-3x_0 - 5d)y_0 - 2dy_1 + (3x_0 + d)y_2}{6d} \end{bmatrix}$ 

(%i4) ratsimp(A . B - Y);
(%o4)  $\begin{bmatrix} -\frac{y_0 - 2y_1 + y_2}{6} \\ \frac{y_0 - 2y_1 + y_2}{3} \\ -\frac{y_0 - 2y_1 + y_2}{6} \end{bmatrix}$ 
```

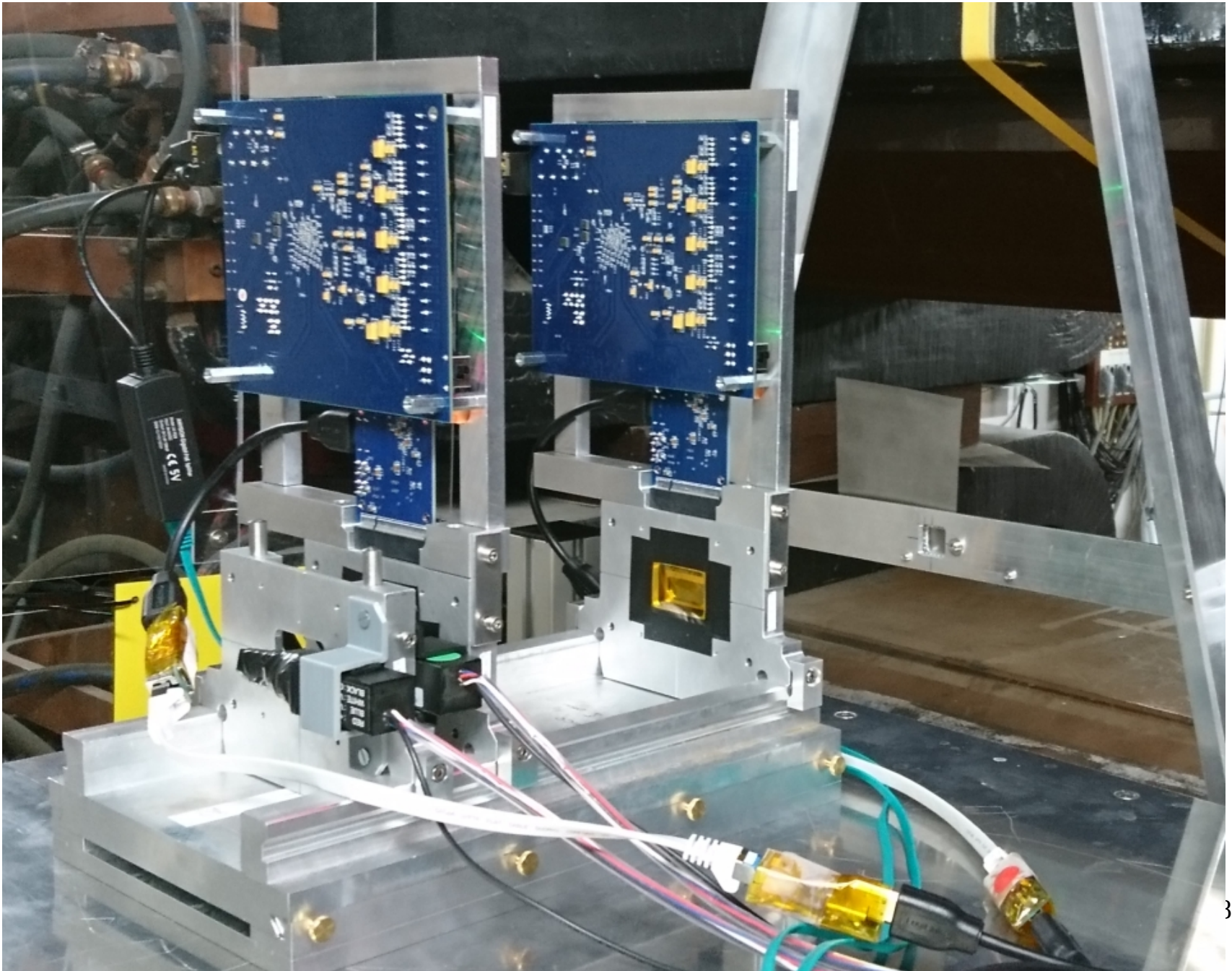
Magnet and TB setup geometry



Upstream of the target



Upstream of the target



Downstream of the target



Downstream of the target

