



Cosmic Birefringence Triggered by Dark Matter Domination

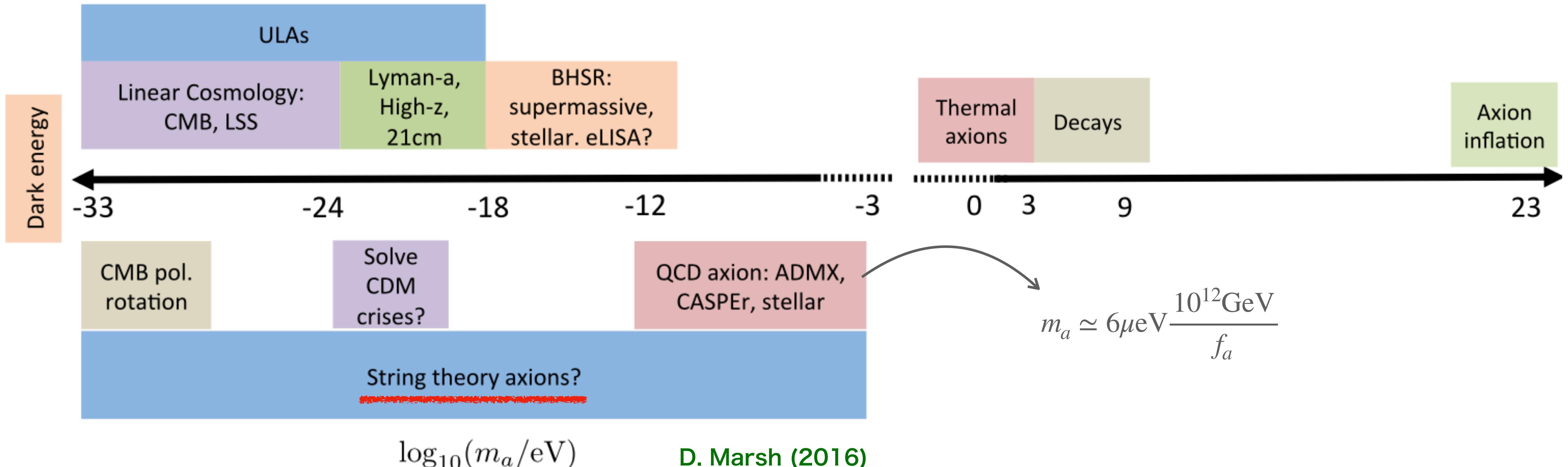
Shota Nakagawa (Tohoku U.)

In collaboration with F. Takahashi & M. Yamada

Based on arXiv:2103.08153

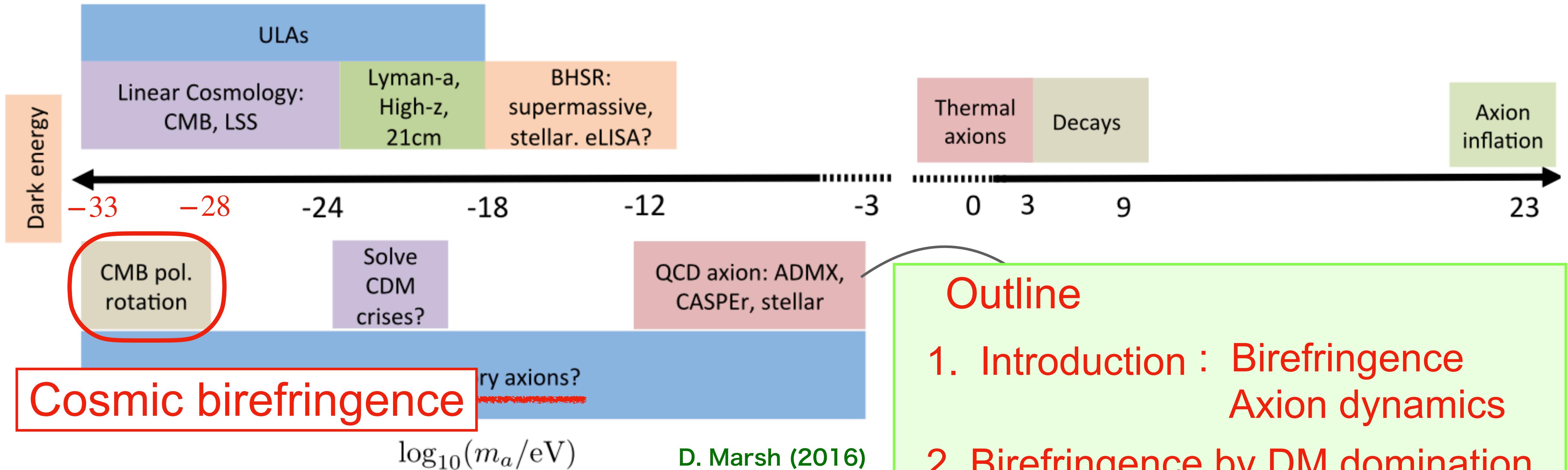
1. Introduction

Axion cosmology has various interesting topics.



1. Introduction

Axion cosmology has various interesting topics.



Outline

1. Introduction : Birefringence Axion dynamics
2. Birefringence by DM domination
3. UV theory
4. Summary

Cosmic birefringence

CMB polarization plane can be rotated after the axion starts to move.

$$\begin{aligned}\mathcal{L} &= -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - c_\gamma \frac{\alpha}{4\pi} \frac{\phi}{f_\phi} F_{\mu\nu} \tilde{F}^{\mu\nu} \\ &\simeq \frac{1}{2} \left[\left(\vec{E} + c_\gamma \frac{\alpha}{2\pi} \frac{\phi}{f_\phi} \vec{B} \right)^2 - \left(\vec{B} - c_\gamma \frac{\alpha}{2\pi} \frac{\phi}{f_\phi} \vec{E} \right)^2 \right] \\ &\quad \vec{D} \qquad \qquad \qquad \vec{H}\end{aligned}$$

S.M. Carroll, G.B.Field, R.Jackiw (1990)
D.Harari, P.Sikivie (1992)
S.M.Carroll (1998)

\vec{D} and \vec{H} obey the free field equation and corresponds to observable field.

Cosmic birefringence

CMB polarization plane can be rotated after the axion starts to move.

$$\begin{aligned}\mathcal{L} &= -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - c_\gamma \frac{\alpha}{4\pi} \frac{\phi}{f_\phi} F_{\mu\nu} \tilde{F}^{\mu\nu} \\ &\simeq \frac{1}{2} \left[\left(\vec{E} + c_\gamma \frac{\alpha}{2\pi} \frac{\phi}{f_\phi} \vec{B} \right)^2 - \left(\vec{B} - c_\gamma \frac{\alpha}{2\pi} \frac{\phi}{f_\phi} \vec{E} \right)^2 \right] \\ &\quad \vec{D} \qquad \qquad \qquad \vec{H}\end{aligned}$$

S.M. Carroll, G.B.Field, R.Jackiw (1990)
D.Harari, P.Sikivie (1992)
S.M.Carroll (1998)

\vec{D} and \vec{H} obey the free field equation and corresponds to observable field.

Rotation angle β

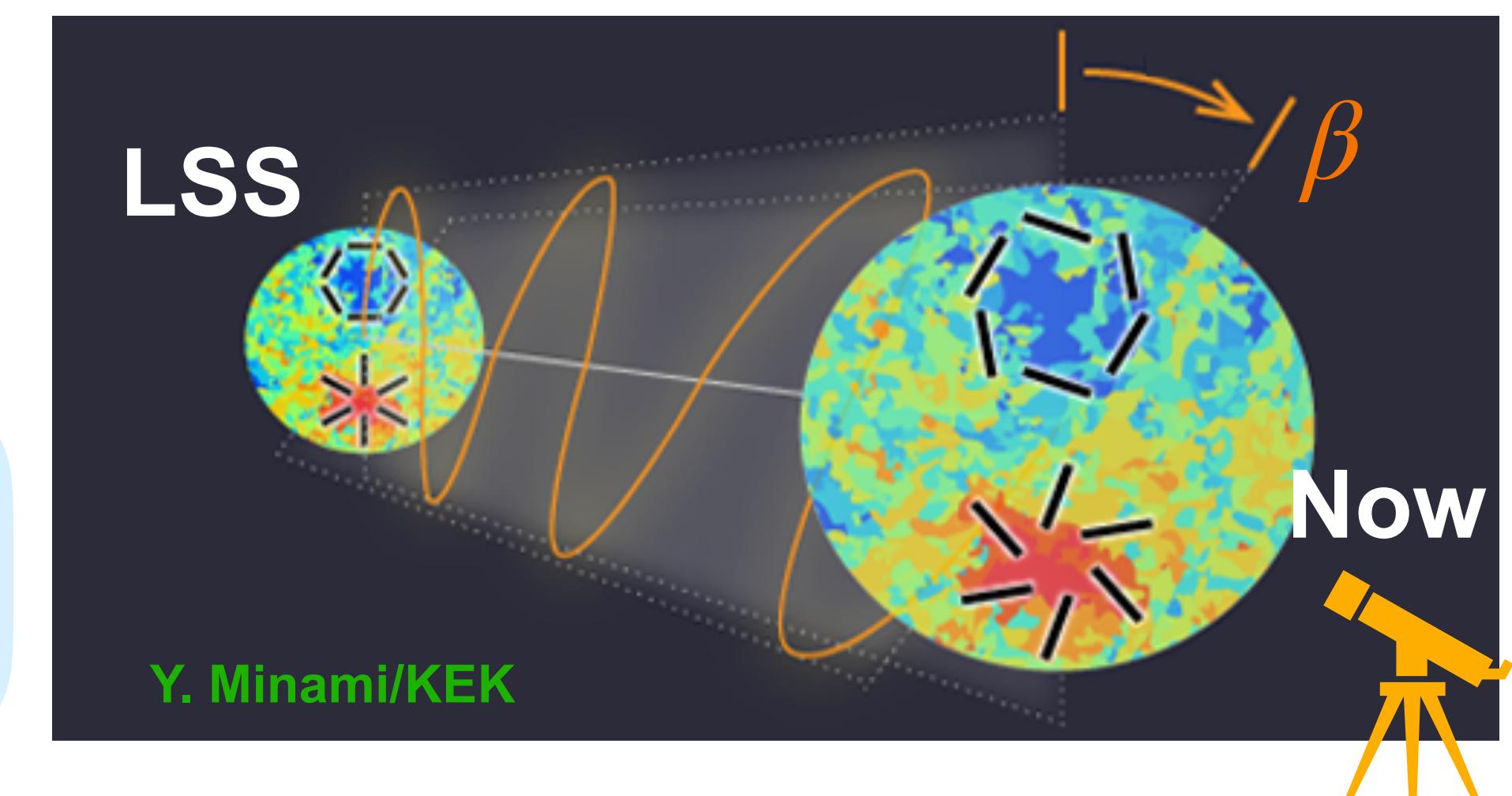
$$\beta = c_\gamma \frac{\alpha}{2\pi} \frac{\Delta\phi}{f_\phi} \simeq 0.42 \text{deg} \times \left(c_\gamma \frac{\phi_{\text{today}} - \phi_{\text{LSS}}}{2\pi f_\phi} \right)$$

$$\Delta\phi/f_\phi \sim O(1) \rightarrow$$

Expectation by Planck

$$\beta = 0.35 \pm 0.14 \text{deg}$$

Y.Minami & E.Komatsu
(2020)



Axion dynamics

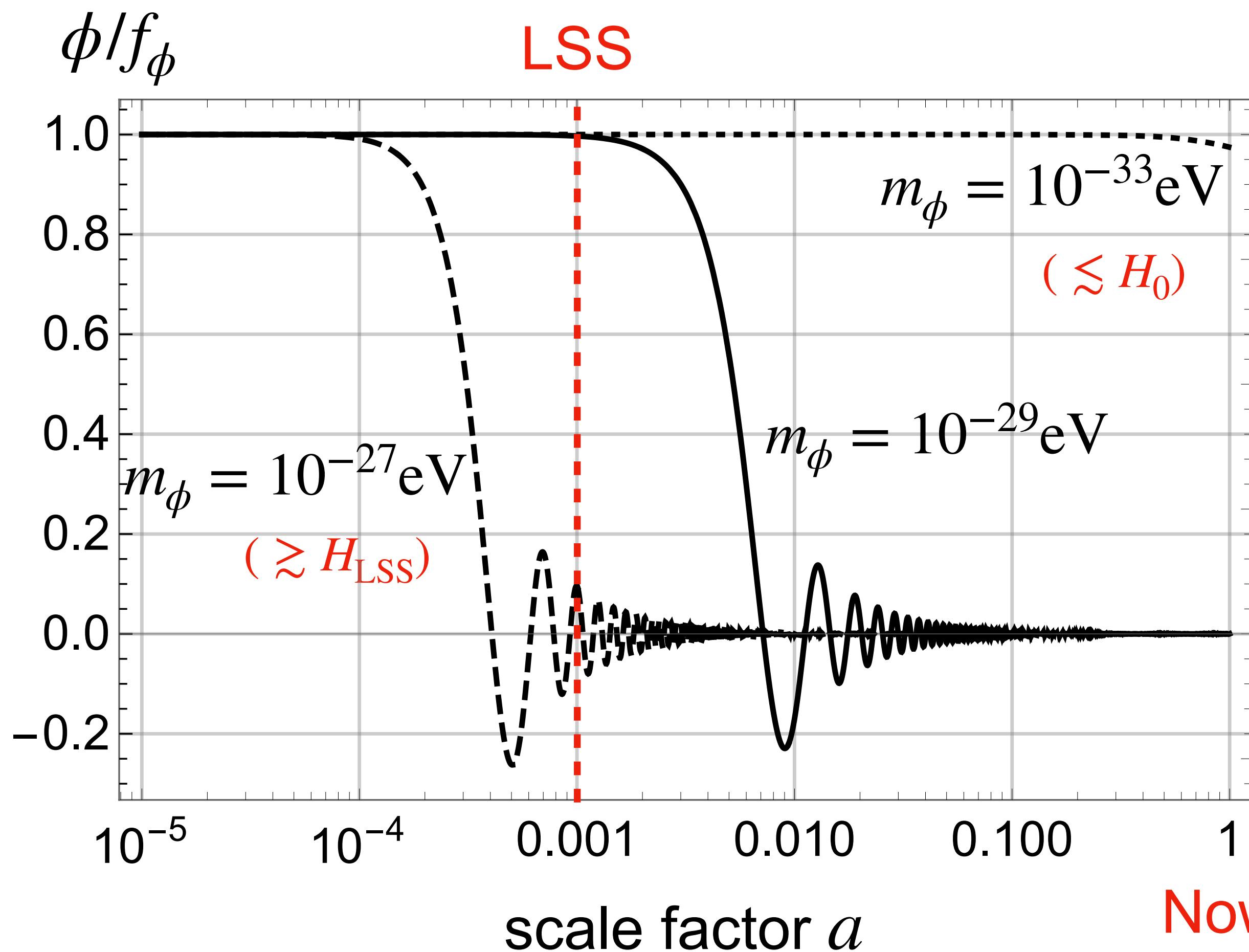
$$10^{-33}\text{eV} \lesssim m_\phi \lesssim 10^{-28}\text{eV} \rightarrow \Delta\phi/f_\phi \sim O(1) \rightarrow$$

$(\sim H_0)$ $(\sim H_{\text{LSS}})$

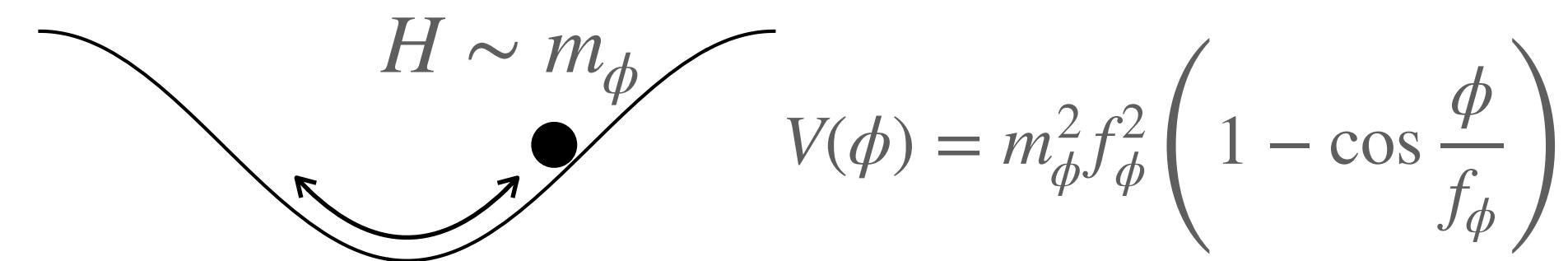
Expectation by Planck

$$\beta = 0.35 \pm 0.14\text{deg}$$

Y.Minami & E.Komatsu
(2020)



Now



Axion dynamics

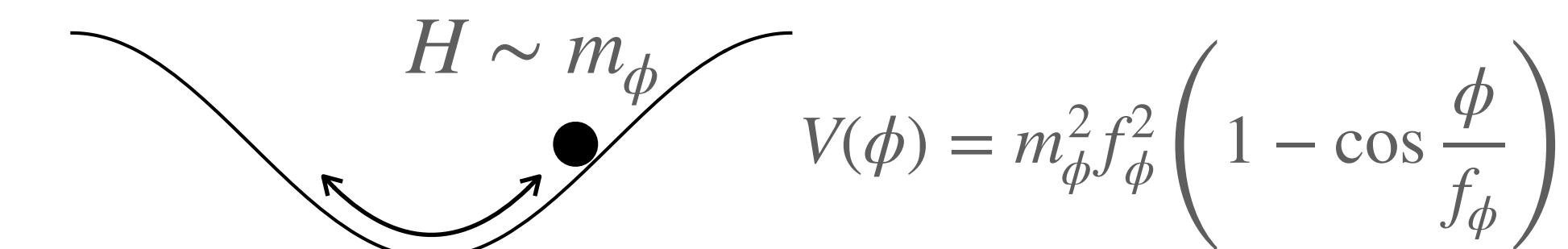
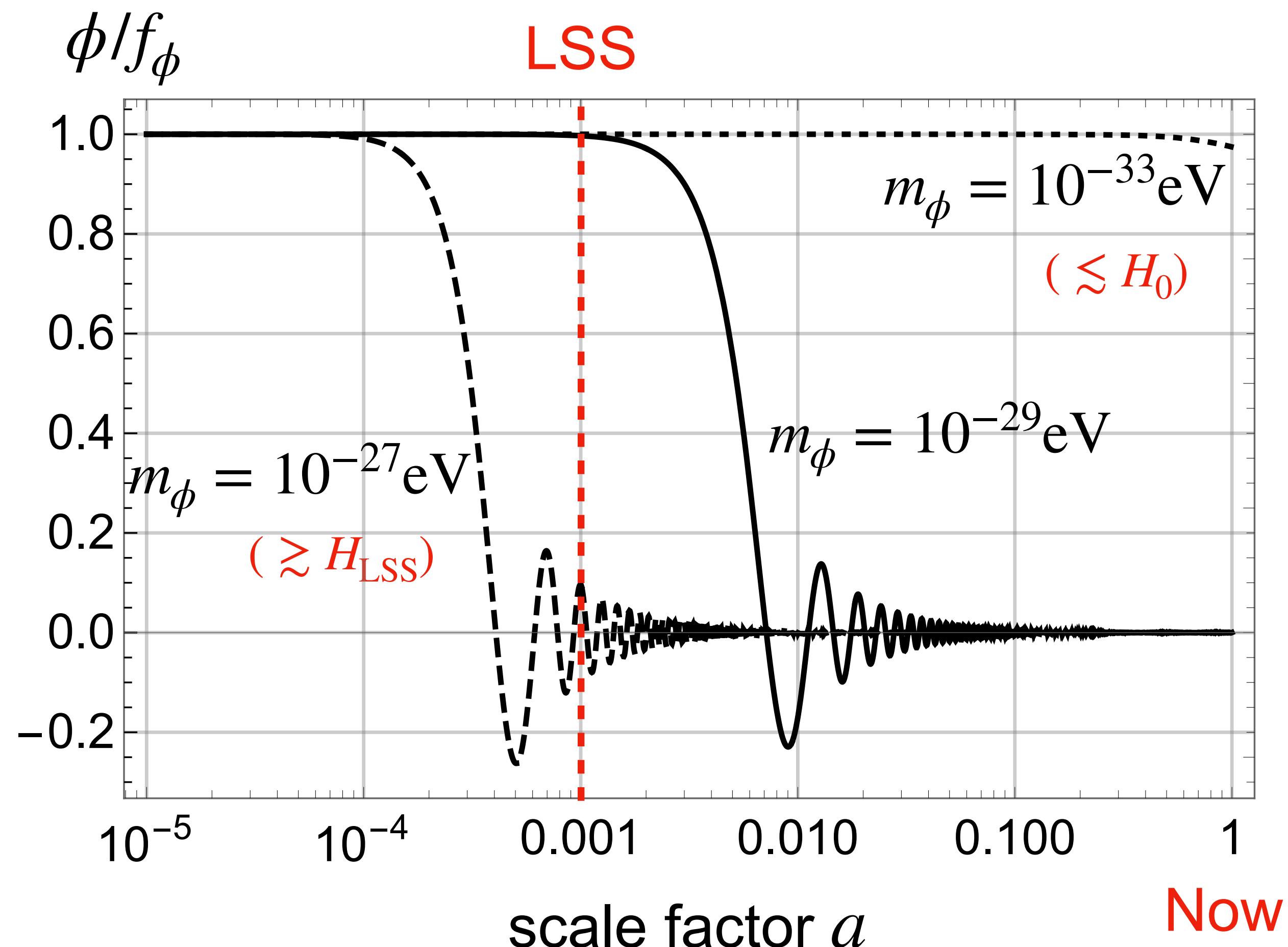
$$10^{-33}\text{eV} \lesssim m_\phi \lesssim 10^{-28}\text{eV} \rightarrow \Delta\phi/f_\phi \sim O(1) \rightarrow$$

$(\sim H_0)$ $(\sim H_{\text{LSS}})$

Expectation by Planck

$$\beta = 0.35 \pm 0.14\text{deg}$$

Y.Minami & E.Komatsu
(2020)



Why does the axion start to move between the recombination and the present?

↔ Why such a mass region?

“Why now?”
(coincidence problem)

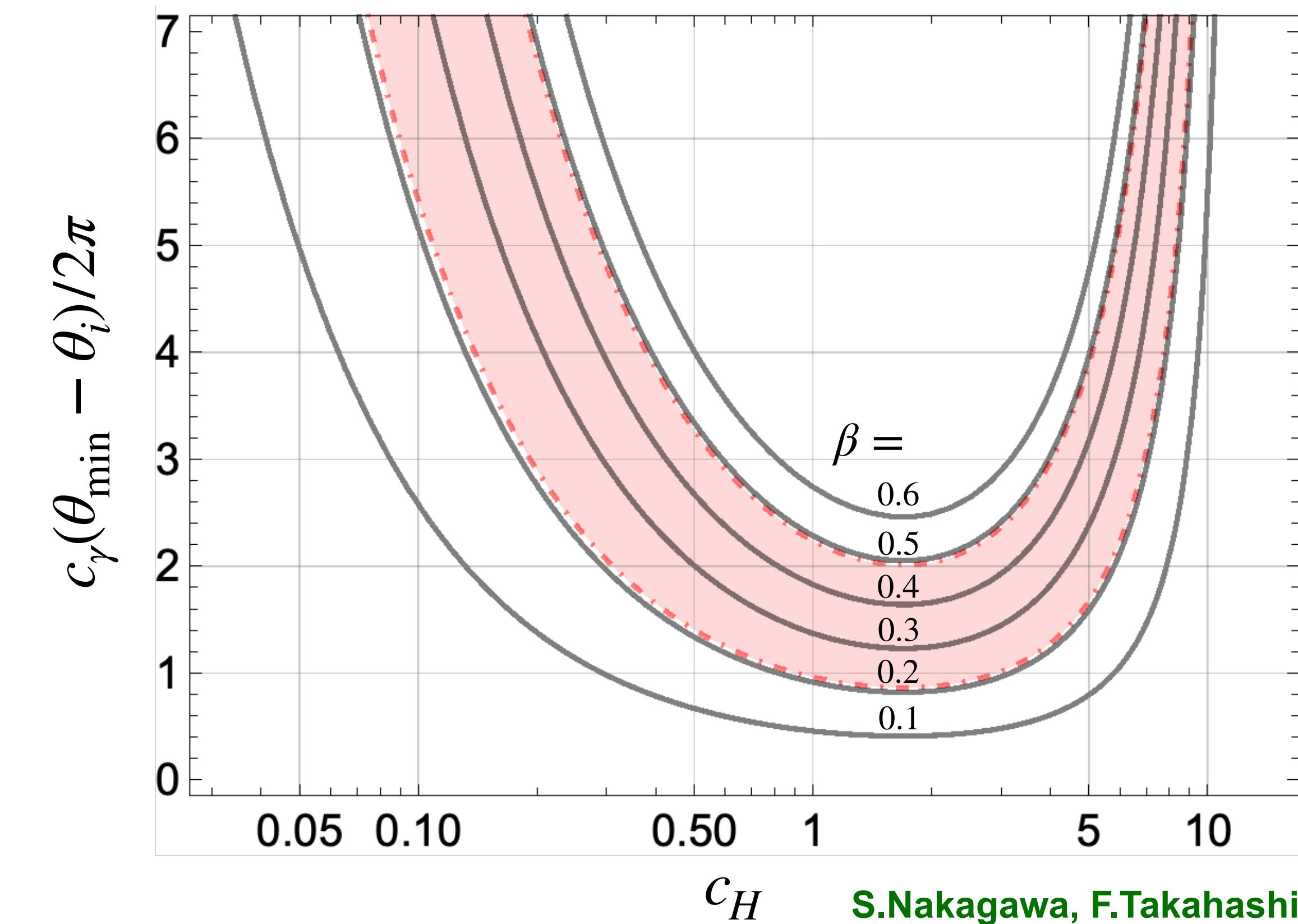
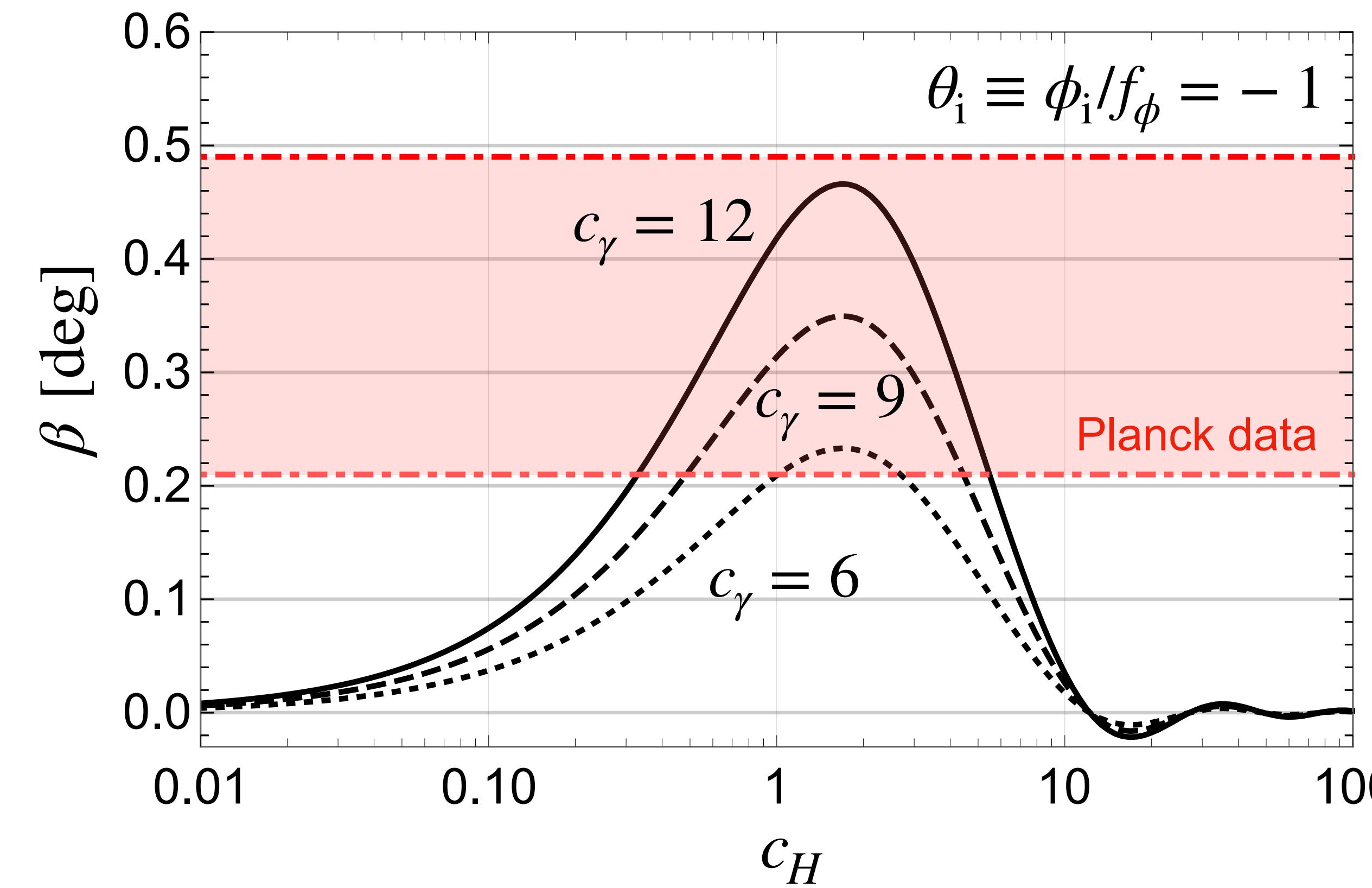
2. Birefringence by DM domination

$$V(\phi) = \frac{1}{2} c_H H_{\text{DM}}^2 \phi^2$$

$$H_{\text{DM}}^2 \equiv \frac{\rho_{\text{DM}}}{3M_{\text{Pl}}^2}$$

$$c_H \sim O(1)$$

The axion starts to move triggered by DM domination.
(just before the recombination)



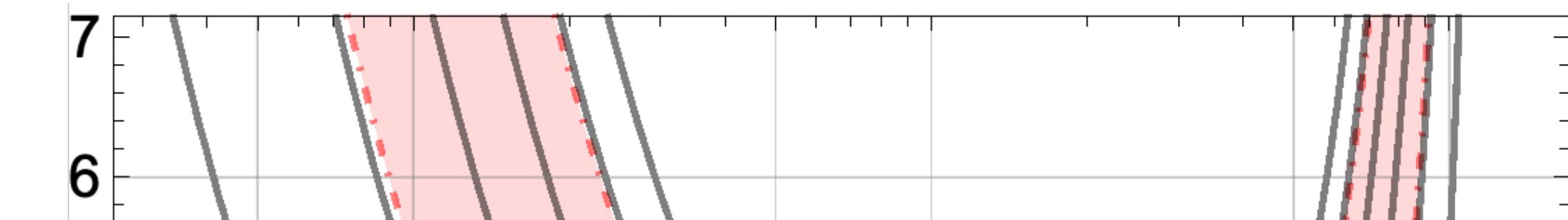
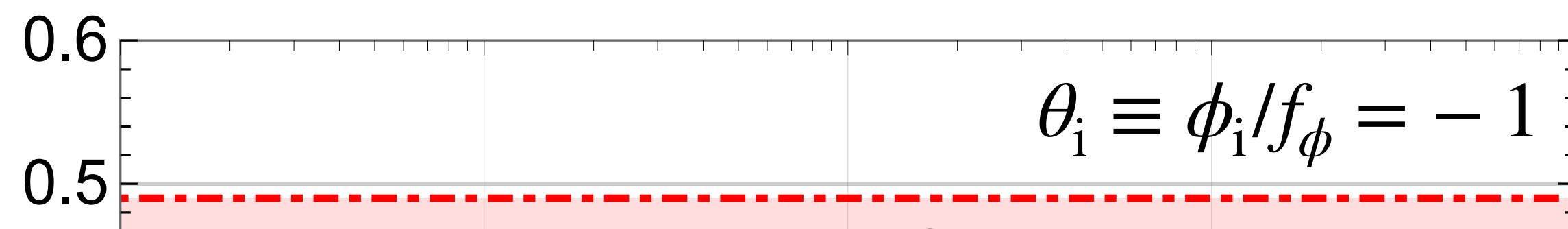
S.Nakagawa, F.Takahashi,
M.Yamada (2021)

2. Birefringence by DM domination

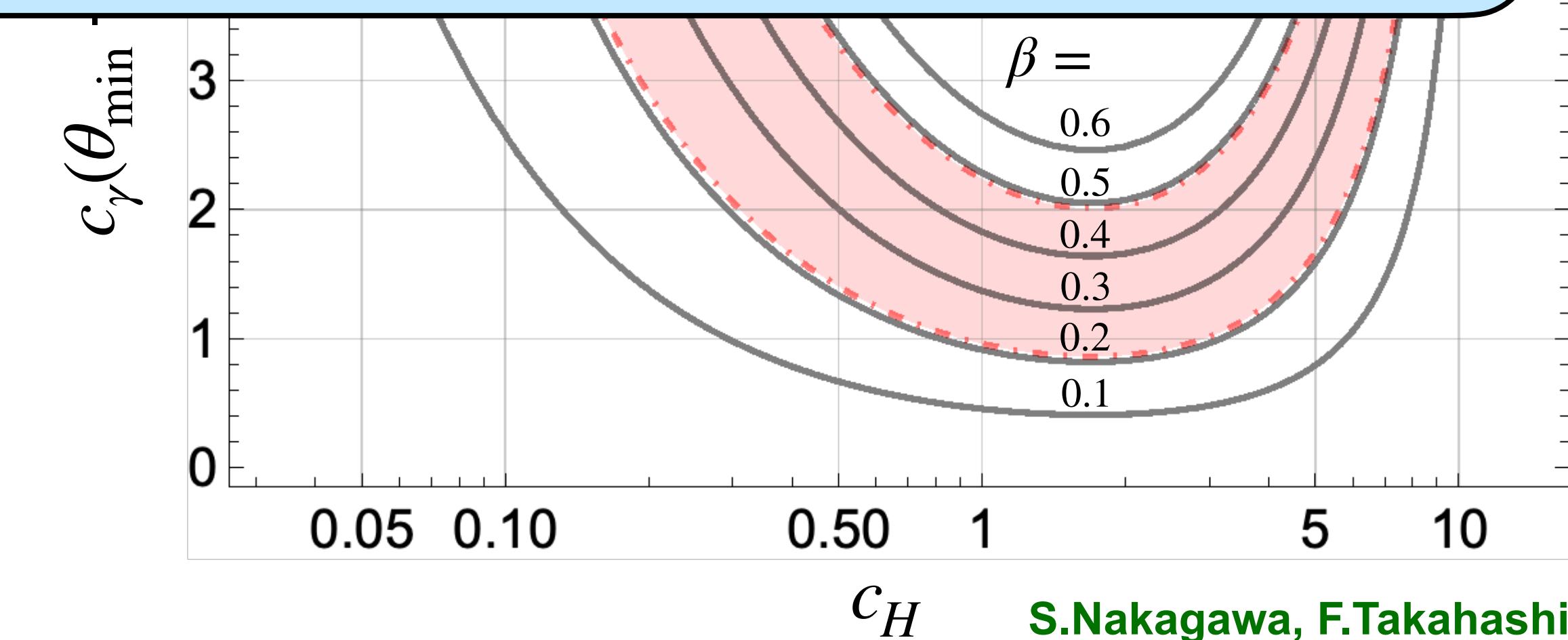
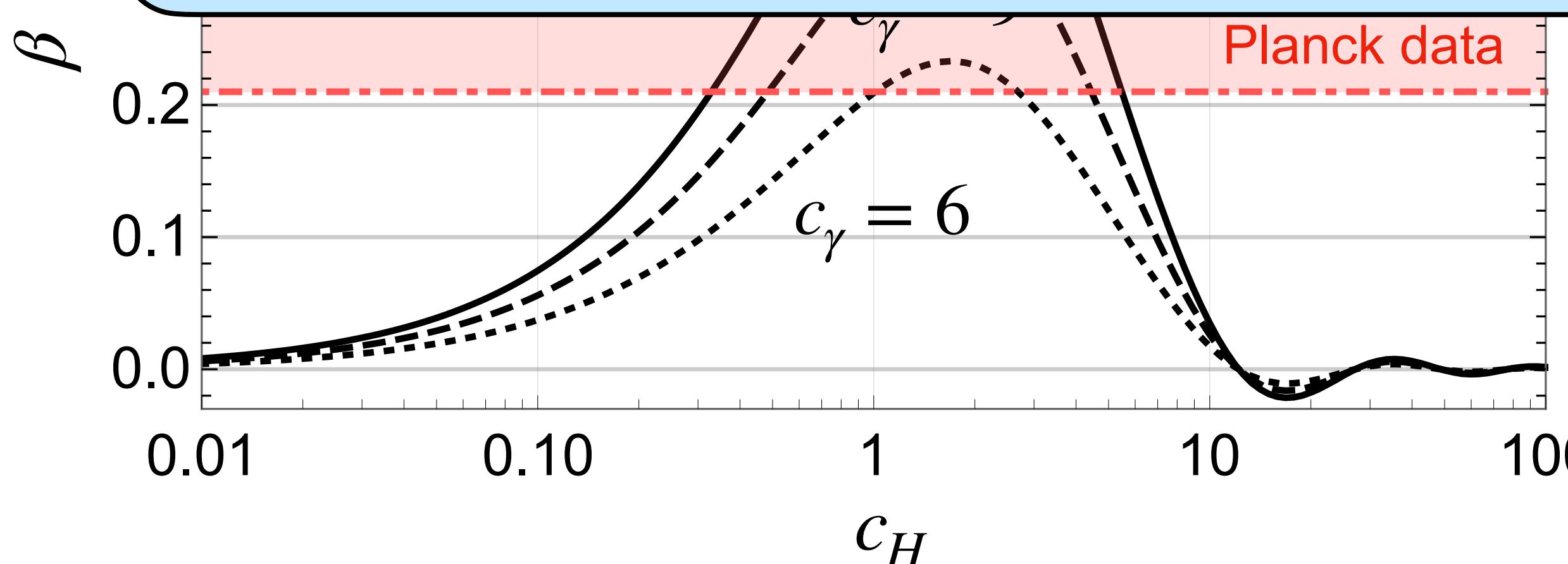
$$V(\phi) = \frac{1}{2} c_H H_{\text{DM}}^2 \phi^2$$

$$H_{\text{DM}}^2 \equiv \frac{\rho_{\text{DM}}}{3M_{\text{Pl}}^2}$$
$$c_H \sim O(1)$$

The axion starts to move triggered by DM domination.
(just before the recombination)



Is there a UV theory which produces the effective mass?



S.Nakagawa, F.Takahashi,
M.Yamada (2021)

3. UV theory

Low energy EFT $\mathcal{L}_\phi = -\frac{1}{2}(\partial\phi)^2 - \frac{1}{2}c_H H_{\text{DM}}^2 \phi^2 - c_\gamma \frac{\alpha}{4\pi} \frac{\phi}{f_\phi} F_{\mu\nu} \tilde{F}^{\mu\nu}$

1. non-minimal coupling to gravity

$$\mathcal{L} \supset -\xi R \phi^2 \simeq -3\xi H_{\text{DM}}^2 \phi^2 \quad (\text{Matter dominated era})$$

S.Nakagawa, F.Takahashi,
M.Yamada (2021)

$$\xi \sim O(1) \rightarrow c_H = 6\xi \sim O(1)$$

$\ddot{*}R \simeq 0$ during RD, so the axion was almost massless.

3. UV theory

Low energy EFT $\mathcal{L}_\phi = -\frac{1}{2}(\partial\phi)^2 - \frac{1}{2}c_H H_{\text{DM}}^2 \phi^2 - c_\gamma \frac{\alpha}{4\pi} \frac{\phi}{f_\phi} F_{\mu\nu} \tilde{F}^{\mu\nu}$

2. Witten effect on hidden monopole DM

- Considering a breaking $SU(2)_H \rightarrow U(1)_H$, hidden magnetic monopole is produced as dark matter
- If the axion has a $U(1)_H$ coupling, the monopole acquires a hidden electric charge (Witten effect). $\mathcal{L} \supset -\frac{1}{4}X_{\mu\nu}X^{\mu\nu} - \frac{\alpha_H \phi}{8\pi f_\phi} X_{\mu\nu} \tilde{X}^{\mu\nu} \rightarrow \nabla \cdot \vec{E}_H = -\frac{\alpha_H \phi}{2\pi f_\phi} \nabla \cdot \vec{B}_H$ E.Witten (1979)
- The axion acquires an effective mass induced from \vec{E}_H . W.Fischler, J.Preskill (1983)

$$m_{\text{eff}}^2 = \left(\frac{\alpha_H}{4\pi f_\phi} \right)^2 \rho_M \rightarrow c_H = 3 \left(\frac{\rho_M}{\rho_{\text{DM}}} \right) \left(\frac{\alpha_H M_{\text{Pl}}}{4\pi f_\phi} \right)^2 \sim O(1)$$

S.Nakagawa, F.Takahashi,
M.Yamada (2021)

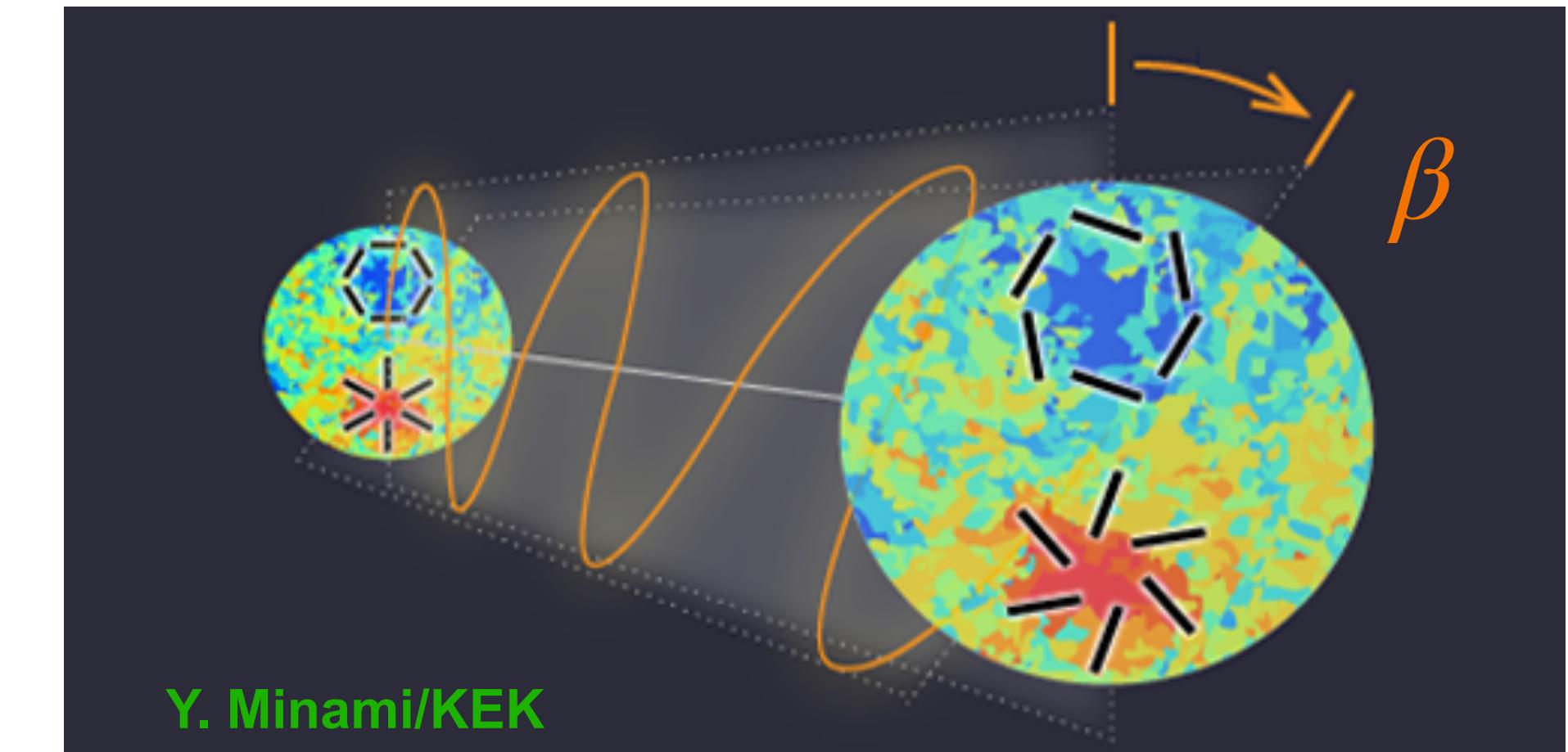
for $\rho_M \sim \rho_{\text{DM}}$, $\alpha_H \sim O(0.01)$, $f_\phi \sim 10^{16}\text{GeV}$

4. Summary

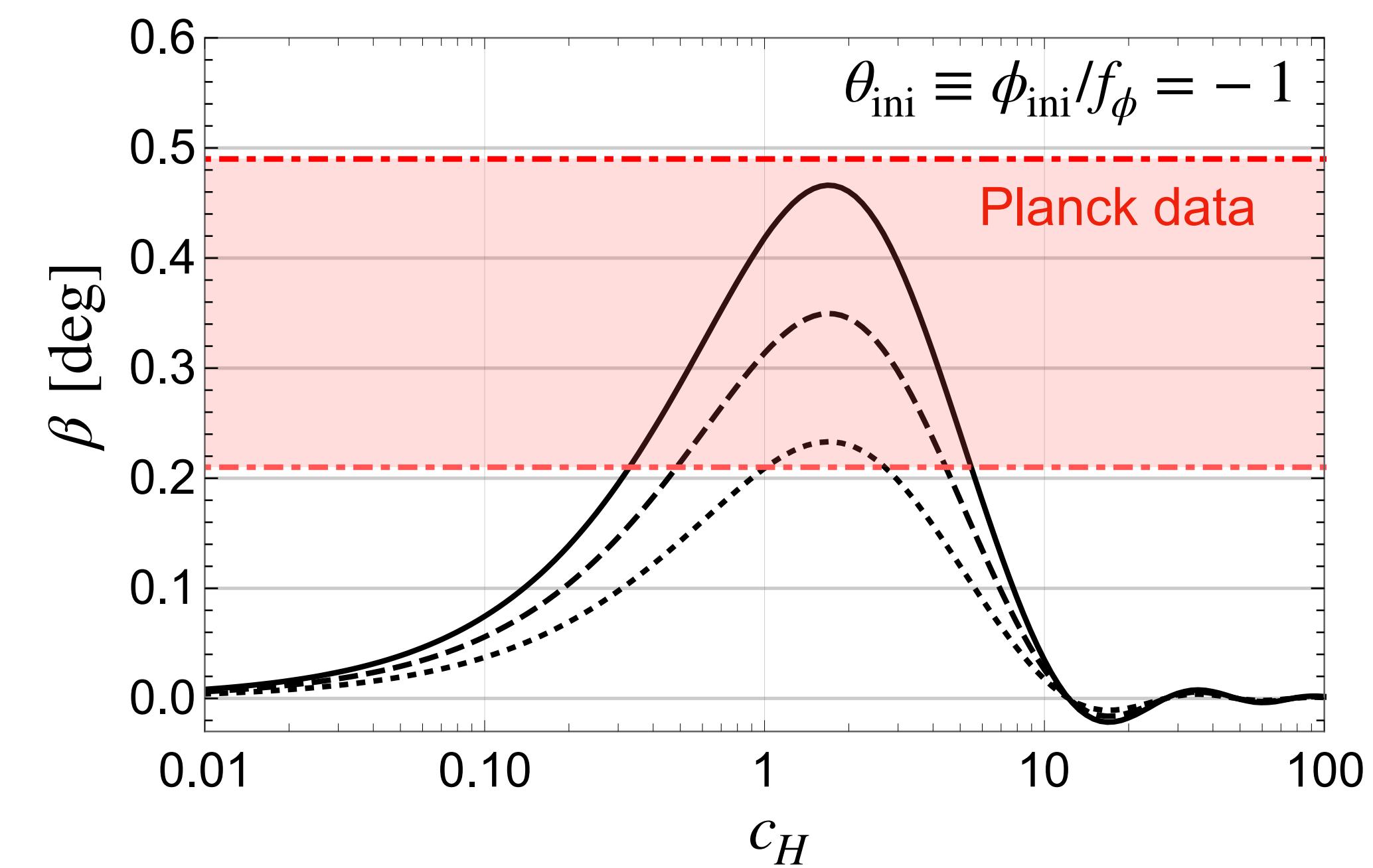
- On the basis of Planck2018 result, CMB polarization angle has been reported:

$$\beta = 0.35 \pm 0.14 \text{deg}$$

Y.Minami & E.Komatsu
(2020)



- There is no theoretical reason “why now” (coincidence problem).
- The effective mass proportional to DM density can explain the problem.
- We propose two specific model:
 - non-minimal coupling
 - Witten effect



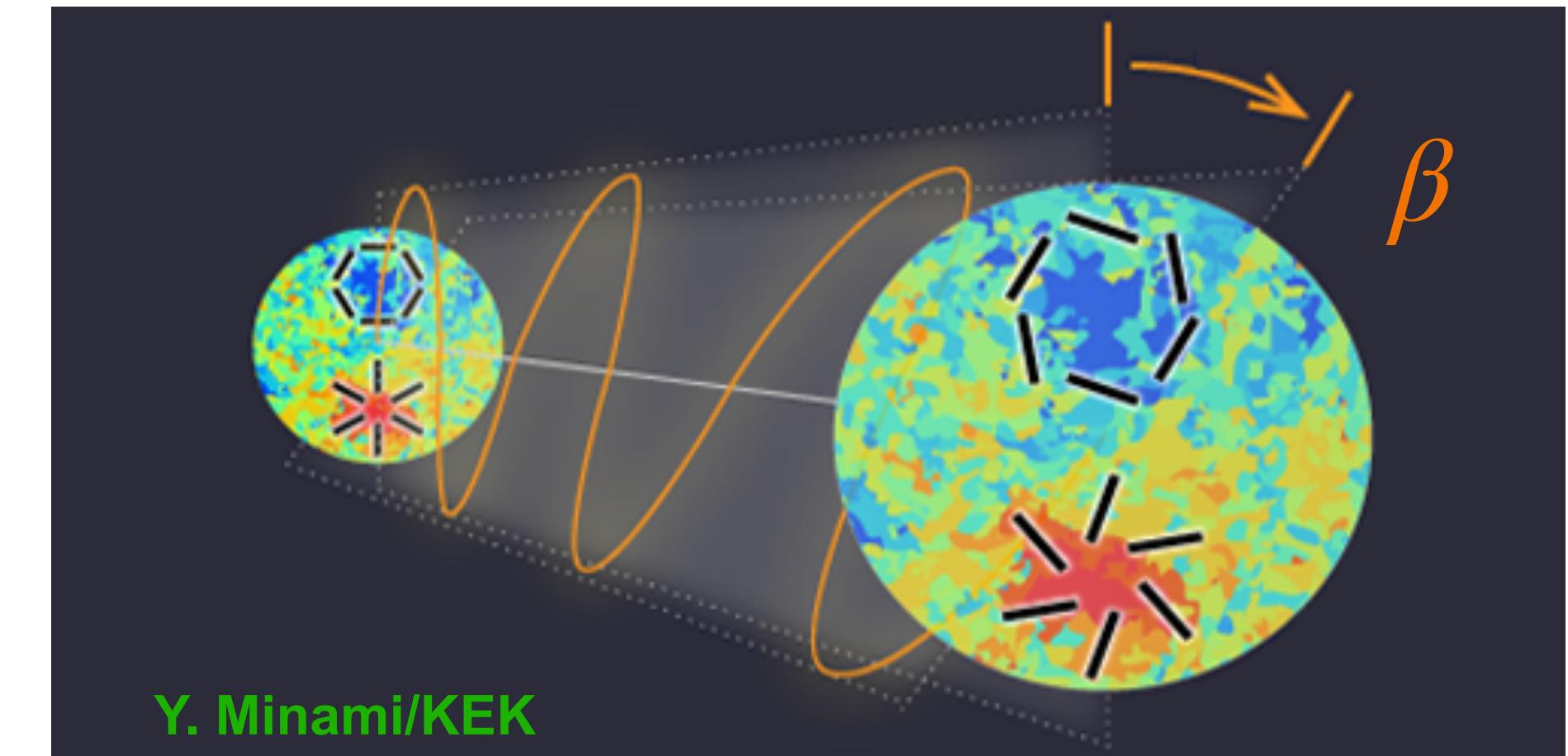
S.Nakagawa, F.Takahashi, M.Yamada (2021)

4. Summary

- On the basis of Planck2018 result, CMB polarization angle has been reported:

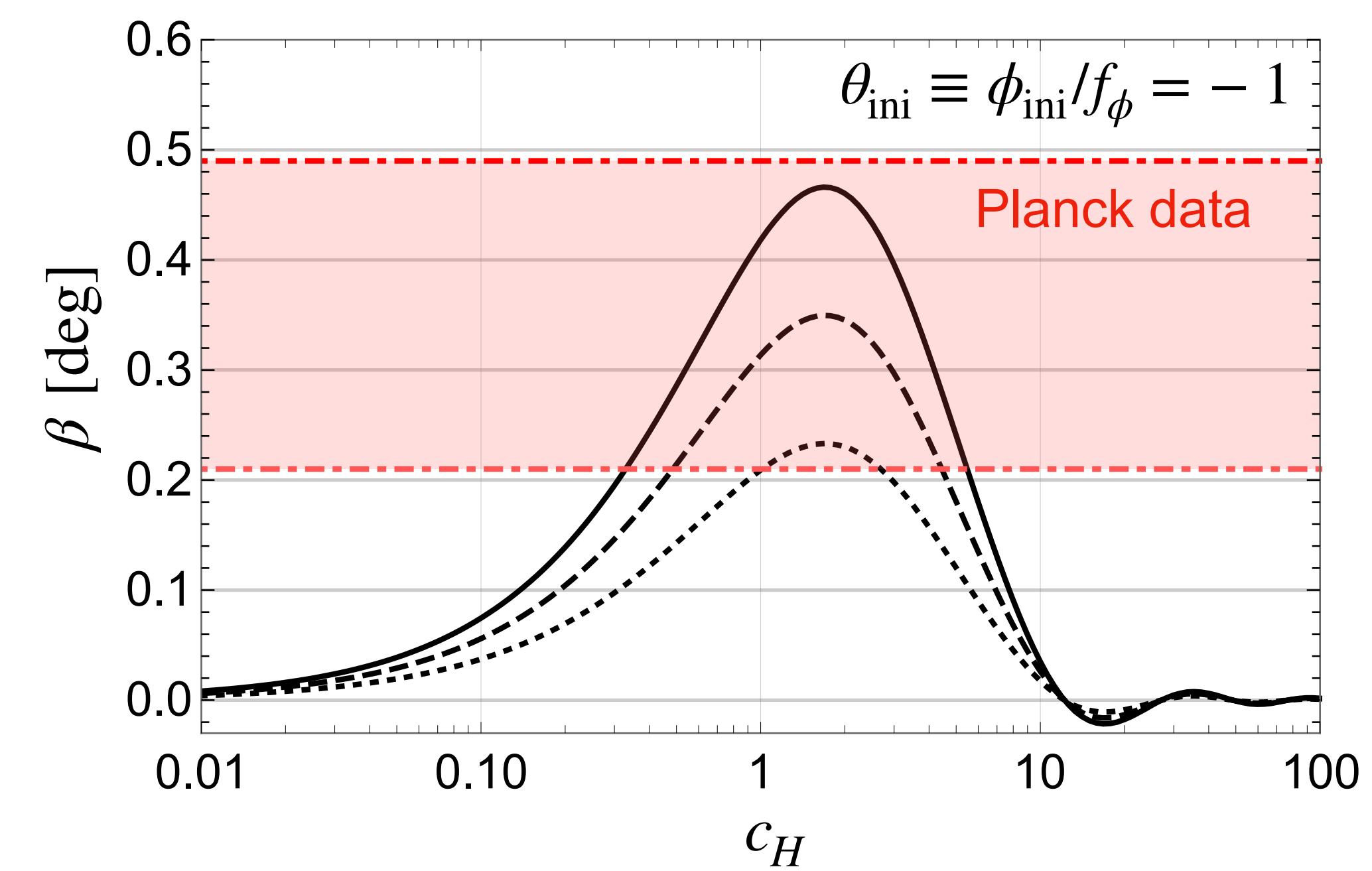
$$\beta = 0.35 \pm 0.14 \text{deg}$$

Y.Minami & E.Komatsu
(2020)



- There is no theoretical reason “why now” (coincidence problem).
- The effective mass proportional to DM density can explain the problem.
- We propose two specific model:
 - non-minimal coupling
 - Witten effect

Thank you.



S.Nakagawa, F.Takahashi, M.Yamada (2021)