## LABORATORY PROBES OF ULTRALIGHT DARK MATTER

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#### **ULTRALIGHT SCALAR DARK MATTER**

Ultralight DM (m<0.1 eV) - large occupation number in the galaxy</p>

$$N = n \cdot \lambda_{dB}^3 = \frac{\rho_{DM}}{m^4 v^3} \approx \left(\frac{\rho_{DM}}{0.3 GeV/cm^3}\right) \left(\frac{0.1 eV}{m}\right)^2$$

► Can be described as a classical field

$$\phi(t, \vec{x}) = A\cos(m_{\phi}\left(t - \vec{v} \cdot \vec{x} + \dots\right))$$

Energy density

$$\rho_{DM} = T_{00} \approx \frac{1}{2} m_{\phi}^2 A^2$$

► Ultralight scalar DM is the coherently oscillating as

$$\left\langle \phi\left(t,\overrightarrow{x}\right)\right\rangle \simeq \frac{\sqrt{2\rho_{DM}}}{m_{\phi}}\cos(m_{\phi}\left(t-\overrightarrow{v}\cdot\overrightarrow{x}+\ldots\right))\,,$$

#### A. Khmelnitskya and V. Rubakov, 1309.5888

#### **ULTRALIGHT DM COUPLING TO THE SM**

- ► If the DM field is coupled to photons and/or electrons  $\mathscr{L}_{int} \supset \frac{g_{\phi\gamma}}{4} \phi F^{\mu\nu} F_{\mu\nu} - g_{\phi e} m_e \phi \overline{e} e$
- ► e.g. Higgs-portal models, in which the scalar couples to the SM through its mixing with the Higgs  $g_{\phi i} = g_{hi} \sin \theta$

$$g_{\phi e} = \frac{\sqrt{2}m_e}{v}\sin\theta$$
,  $g_{\phi\gamma} = -\frac{\alpha_0\sin\theta}{2\pi v}\left|A_W(\tau_W) + \sum_{\text{fermions}}N_{c,f}Q_f^2A_F(\tau_f)\right|$ 

- ► Temporal oscillations of  $\alpha$ ,  $m_e$  at frequency determined by the mass of the field.
- ► A mediator of a new fifth-force.

### **TEMPORAL VARIATIONS OF FUNDAMENTAL CONSTANTS**

Solutions of the DM field induce oscillations of  $m_e, \alpha$ 

$$\delta m_e = g_{\phi e} m_e \langle \phi(t, \vec{x}) \rangle, \qquad \delta \alpha = g_{\phi \gamma} \langle \phi(t, \vec{x}) \rangle \alpha^{SM} \qquad \langle \phi(t, \vec{x}) \rangle \simeq \frac{\sqrt{2\rho_{DM}}}{m_{\phi}} \cos(m_{\phi} t),$$

Temporal oscillations of atomic frequencies

- ► Rydberg energy levels  $R_{\infty} \propto \alpha^2 m_e$
- ► cavity modes frequencies  $a_0 \propto (\alpha m_e)^{-1}$
- ► H hyperfine transition  $\omega_{HF} \propto \alpha^4 m_e^2$
- > Detectable by comparing two atomic probes that depend differently on  $\alpha, m_e$

$$\Delta \Phi = \int \delta \omega_1(t) - \delta \omega_2(t) dt \approx \omega \int (n_{\alpha_2} - n_{\alpha_1}) \frac{\delta \alpha}{\alpha} + (n_{m_{e_2}} - n_{m_{e_1}}) \frac{\delta m_e}{m_e} dt$$

Y. V. Stadnik1 and V. V. Flambaum, PHYSICAL REVIEW A 93, 063630 (2016)

#### **RAPIDALLY OSCILLATING DARK MATTER**

- ► "Heavy" ultralight DM  $m_{\phi} \gtrsim 10^{-12} \text{ eV}, \nu_{\phi} \gtrsim 100 \text{ Hz}$ 
  - ► Theoretically motivated relaxion DM.
  - Blind spot for atomic probes of temporal variations of fundamental constants due to rapid oscillations.
- Dynamical Decoupling modulation - accumulating effects at a probe frequency, allows for long interrogation times.



#### **PROBING RAPIDLY OSCILLATING DM USING DD**

$$\frac{\Delta f(\nu)}{f_0} = \frac{\delta f_{\text{ion}}(\nu) - \delta f_{\text{laser}}(\nu)}{f_0} = \left(2 - F(\nu)\right) \frac{\delta \alpha}{\alpha} + \left(1 - F(\nu)\right) \frac{\delta m_e}{m_e}$$



First model independent bounds on variations of  $\alpha$ ,  $m_e$  up to MHz frequencies. S. Aharony et al., [1902.02788].

#### **TESTS OF FIFTH-FORCE AND EP**

- New scalar mediates a force (interaction) between neutral masses
- Searches for deviations from gravity
  - massive force carrier Yukawa potential  $V_{\phi} \propto \frac{e}{-}$
  - Non-universal couplings to masses of different compositions - violation of the universality of free-fall (Einstein's Equivalence Principle)

$$\eta = \frac{a_A - a_B}{a_A + a_B} \propto \frac{1}{G} \left( g_{\phi e} Q_{m_e}^C + g_{\phi \gamma} Q_e^C \right) \left( g_{\phi e} \left( Q_{m_e}^A - Q_{m_e}^B \right) + g_{\phi \gamma} \left( Q_e^A - Q_e^B \right) \right)$$





#### **COMPARISON – FIFTH FORCE VS. FUND. CONSTANTS**



#### **QUADRATICALLY COUPLED DM**

► Higher order interactions of the DM candidate with the SM

$$\mathscr{L} = g_{\phi^2 e} \phi^2 m_e \bar{e} e$$

> Variations of  $m_e$ 

$$\frac{\delta m_e}{m_e} = g_{\phi^2 e} \frac{\rho_{DM}}{m_{\phi}^2} \cos(2m_{\phi}t)$$

► 5F/EP - gains DM density sensitivity, different r-dependance

$$V(r) \propto \phi^2 \sim \frac{\rho_{DM}}{m_{\phi}^2} \left(1 - s_A \frac{GM_A}{r}\right)^2$$

$$\eta = \frac{a_A - a_B}{a_A + a_B} \propto \frac{\rho_{DM}}{Gm_{\phi}^2} \left( g_{\phi^2 e} Q_{m_e}^C + g_{\phi^2 \gamma} Q_e^C \right) \left( g_{\phi^2 e} \left( Q_{m_e}^A - Q_{m_e}^B \right) + g_{\phi^2 \gamma} \left( Q_e^A - Q_e^B \right) \right)$$

#### A. Hees et al, 1807.04512

#### **COMPARISON – QUADRATICALLY COUPLED DM**



### **OUTLOOK AND FUTURE WORK**

- Ultralight DM could result in temporal oscillations of fundamental constants, and non-universal acceleration.
- Different couplings/density enhancement can make atomic probes more sensitive.
- Ongoing work Model building for Natural theories with quadratically coupled DM.
- Future work using tools from signal processing to increase the sensitivity to rapidly oscillating DM.

## THANK YOU!

# **BACKUP SLIDES**

### **DM HALOS CENTERED AROUND EARTH?**

- Temporal variation probes are sensitive to the local DM density at the position of the experiment.
- ► What if the DM density is locally enhanced?
  - "SM" enhancement (?)
  - DM halos earth's gravity as a source binding the star together.

#### **RESULTS – DM HALOS AROUND EARTH**

