

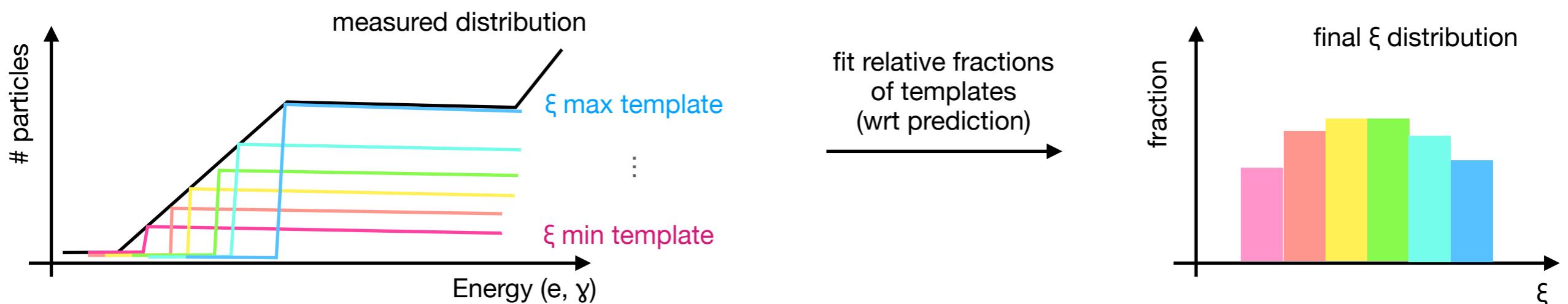
Finite Impulses Response Filters for Compton Edge reconstruction & Background G4 studies

Ruth Jacobs

LUXE analysis TF meeting
26th October 2020

Reminder: Why we are interested in the kink

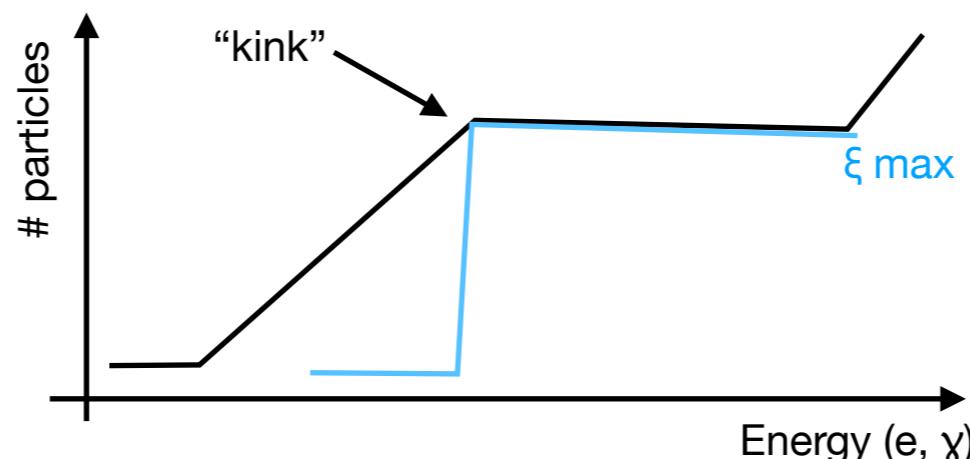
- Gaussian pulse: overlay of different true ξ leads to dramatic washing-out of edges
- final analysis should be a template fit (template of different ξ bins) fit to the spectrum



For the CDR propose simple approach:

- instead of differentiation, try to find the “kink” of the edge
- for low enough ξ (high w_0), this position corresponds to ξ_{max}

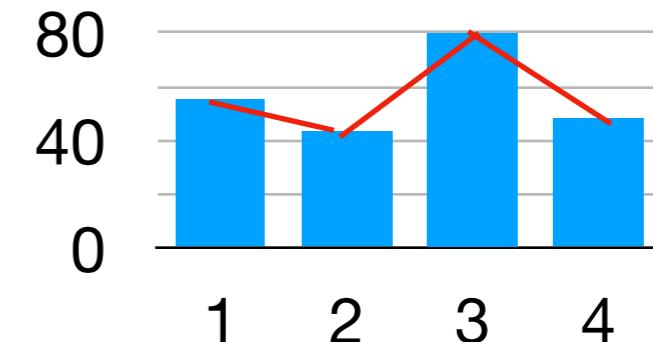
← **Today!**



Finite Impulses Response Filter (FIR)

Before: Simple Differentiation for Edge finding

- get electron x distribution
- calculate slope bin-by-bin
→ bin with max. slope = edge
- susceptible to statistical fluctuations!



Finite Impulses Response Filter

method used by J. List et. al.

- edge-like features in function $g(x)$ can be identified by maxima in the convolution $R(x)=h(x)*g(x)$ where $h(x)$ is a matched filter
- $R(x)$ is called the **Response**
- we have discrete data points $x=(x_0, \dots, x_i)$, need discretized Response $R_d(i)$

$$R_d(i) = \sum_{k=-N}^N h_d(k) \cdot g_d(i - k)$$

- different filters h_d available, optimal choice depends on the function $g(x)$
- Used here: **First derivative of a Gaussian (FDG)**

$$h_d(k) = -k \exp\left(-\frac{k^2}{2\sigma^2}\right) \text{ for } -N \leq k \leq N$$

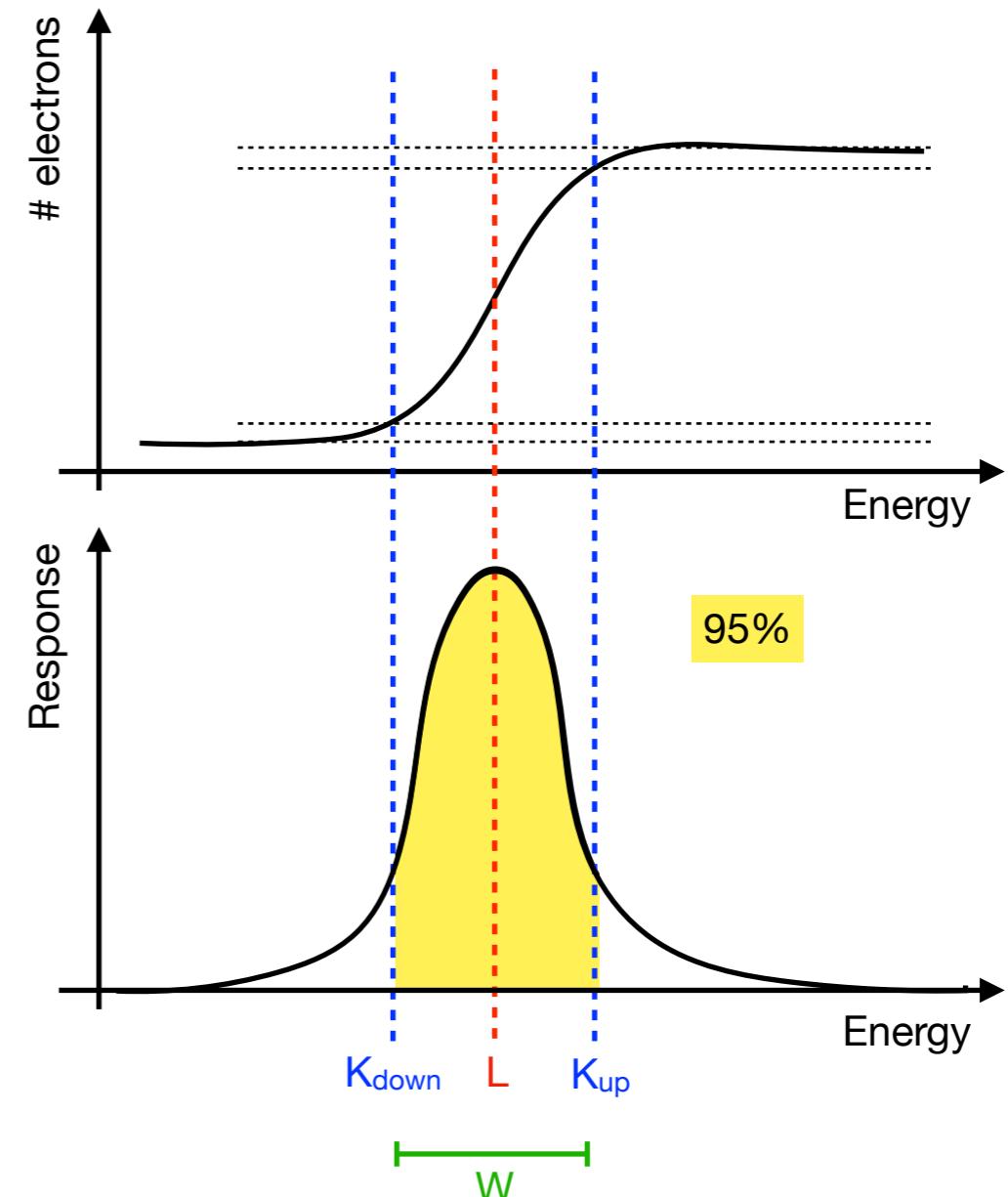
FIR approximates first derivative

DESY. — thanks to filters more robust against statistical fluctuations!

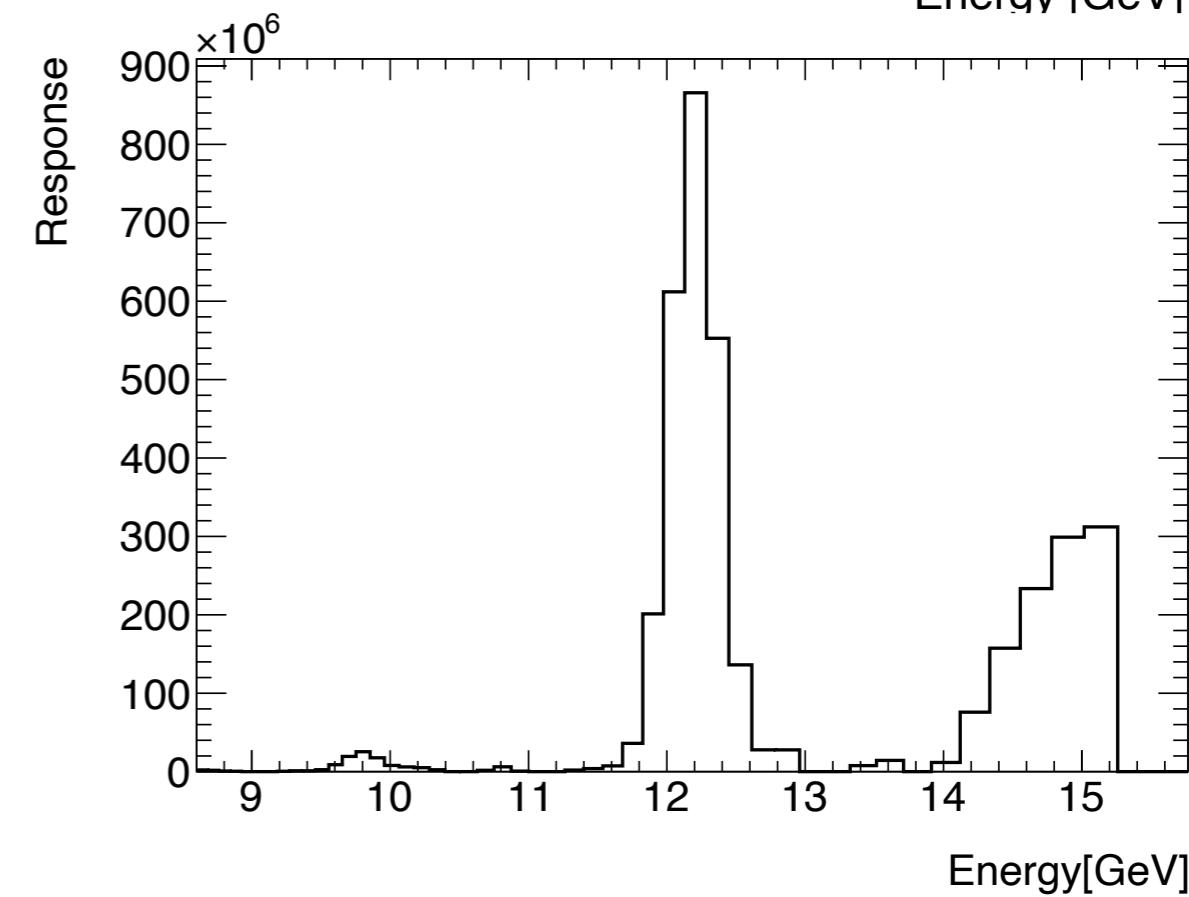
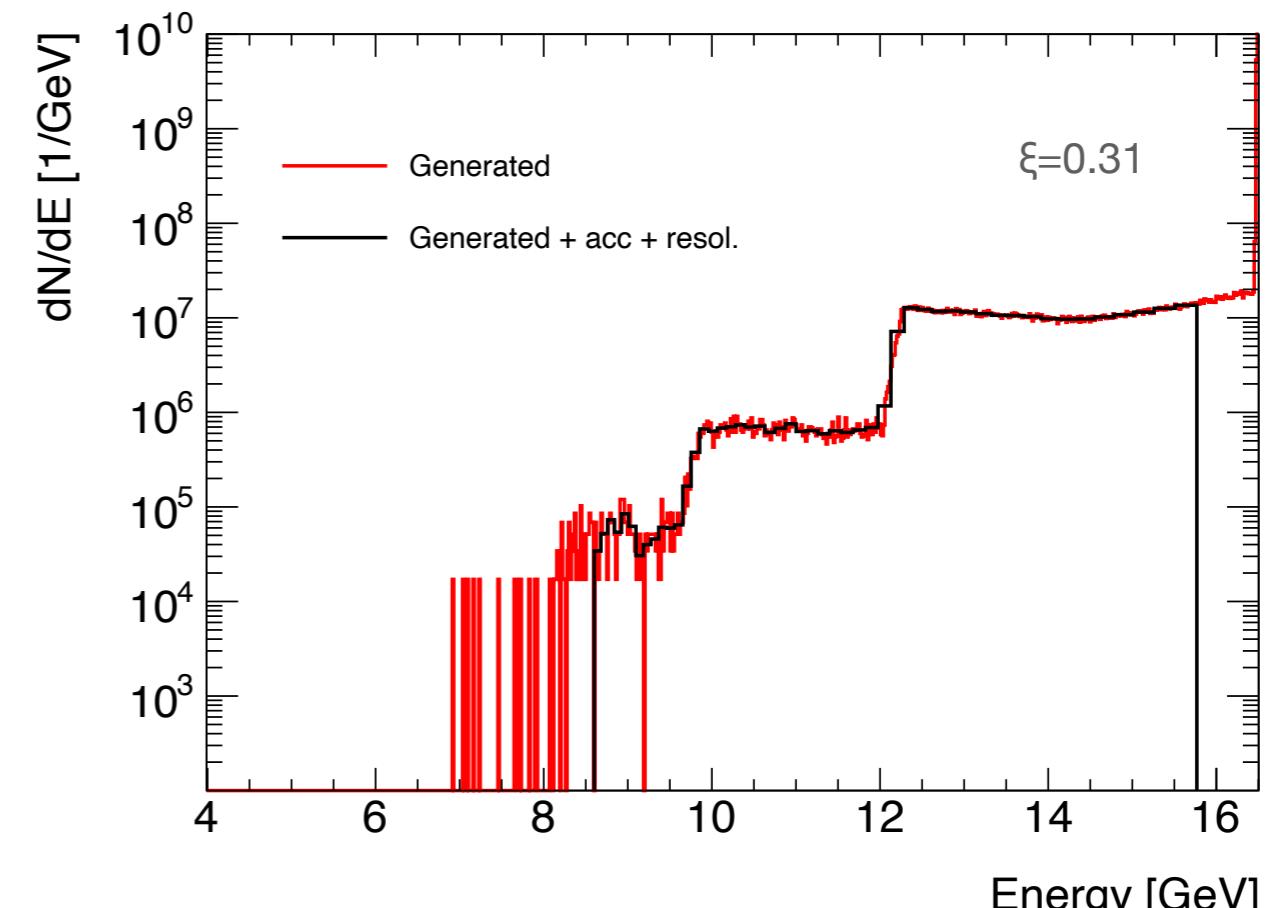
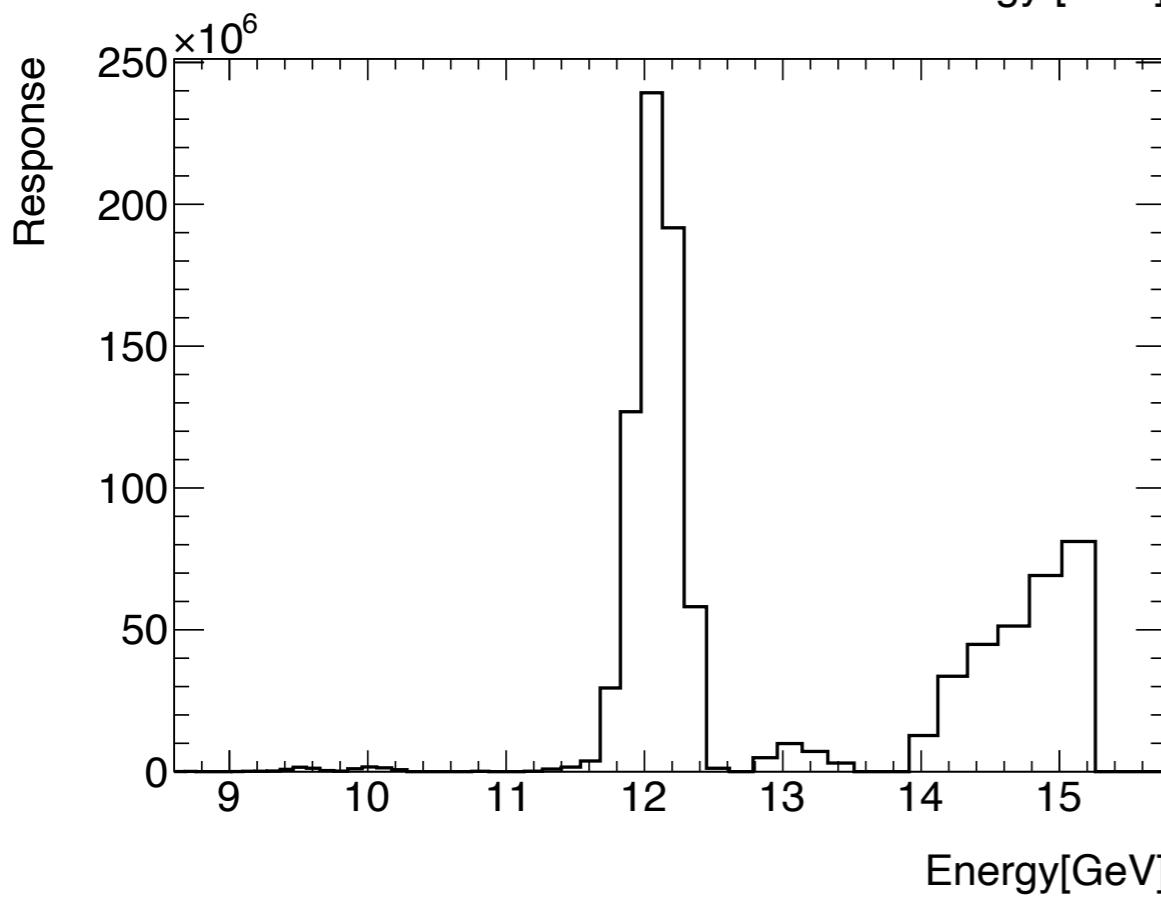
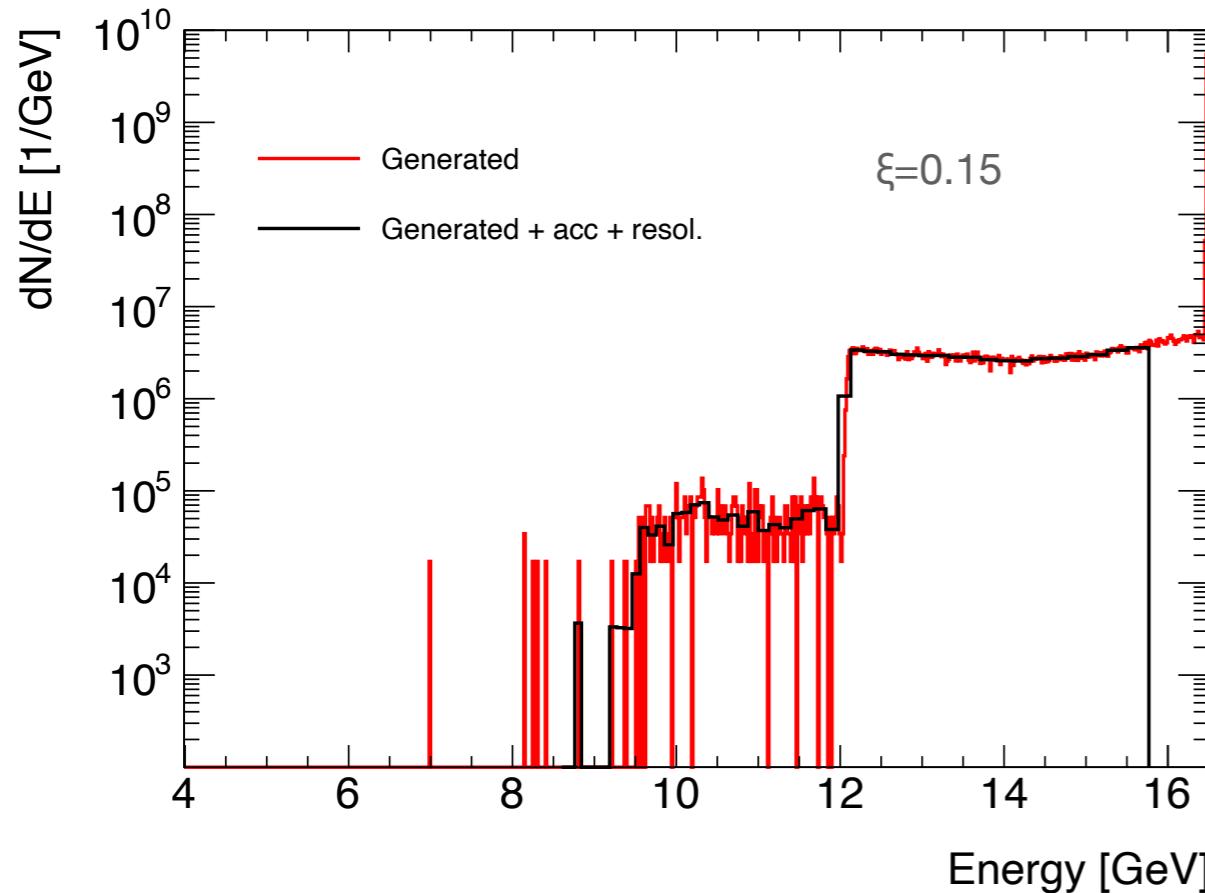
Finite Impulses Response Filter (FIR)

Features of interests:

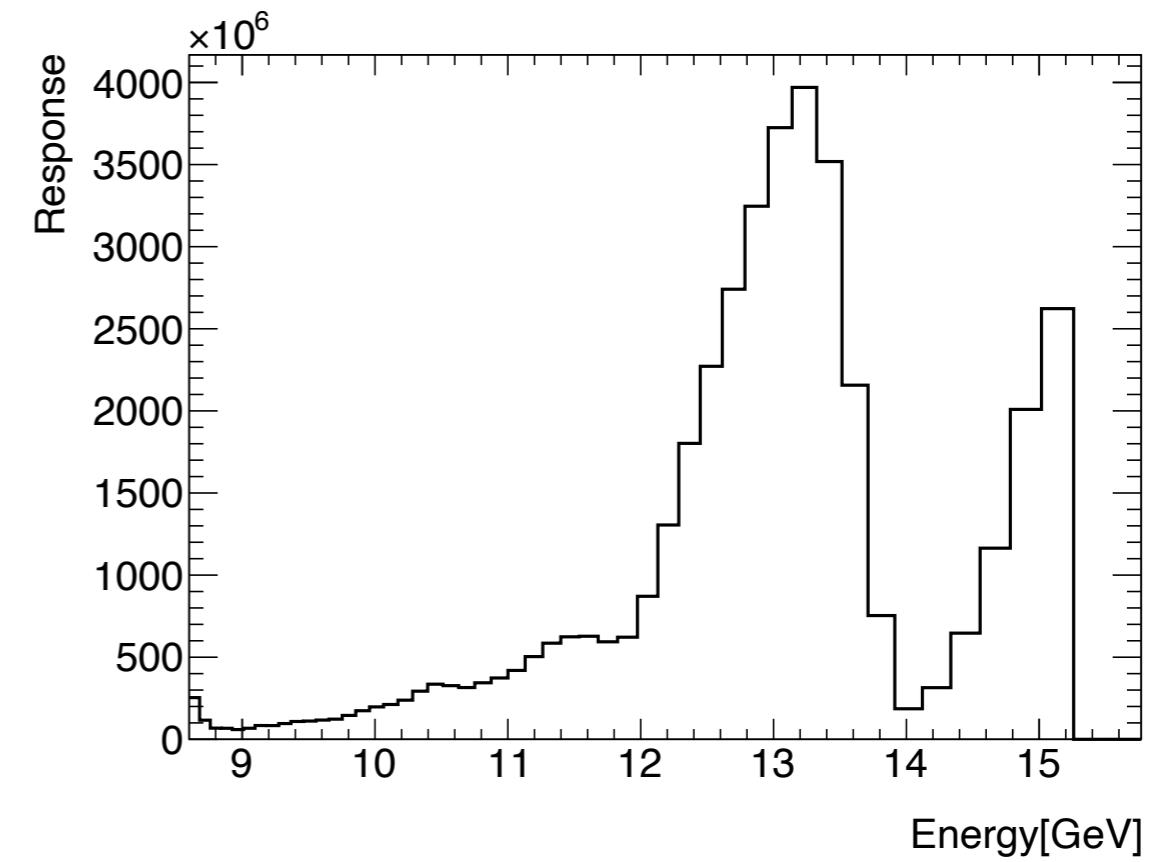
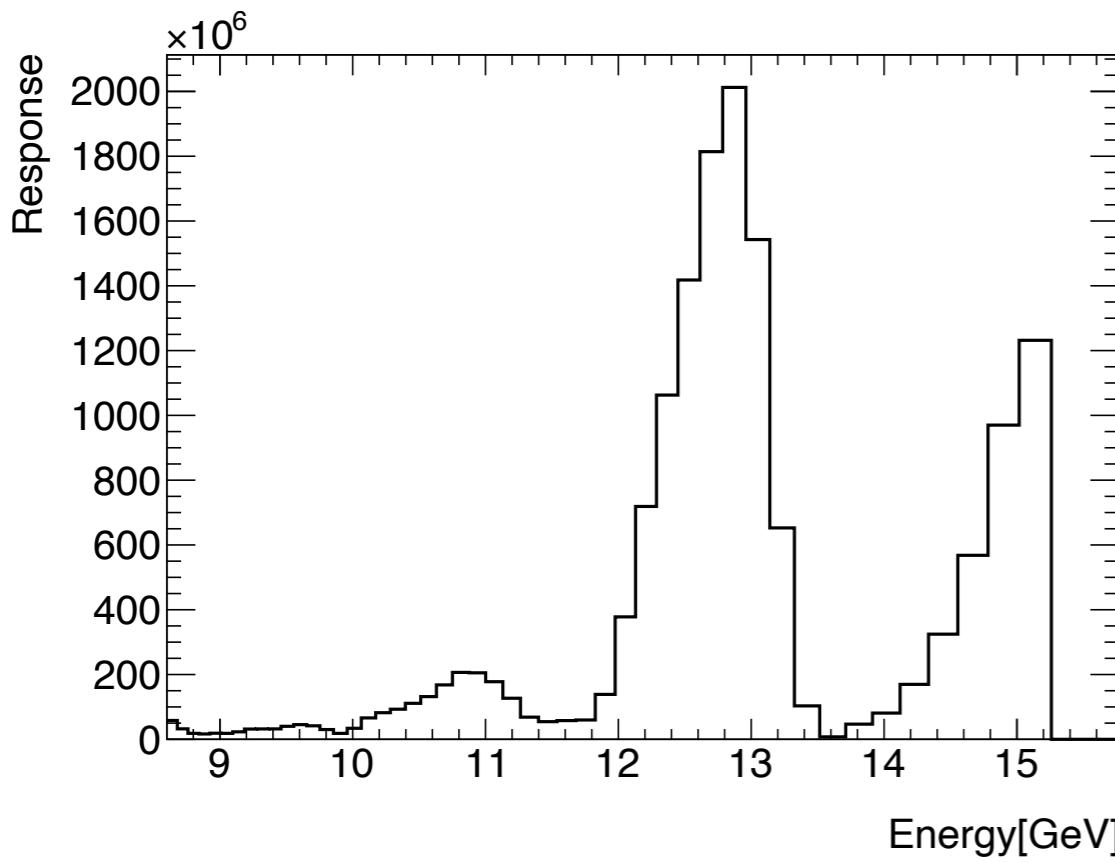
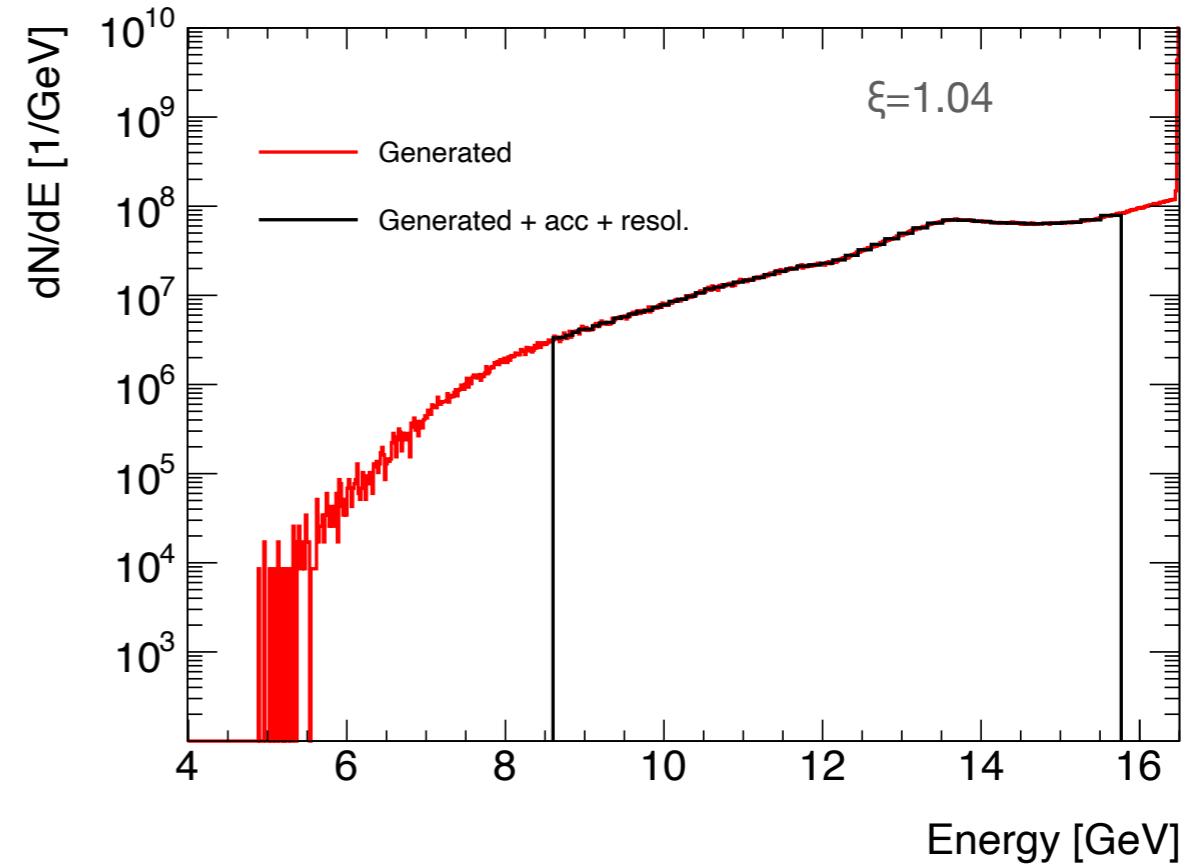
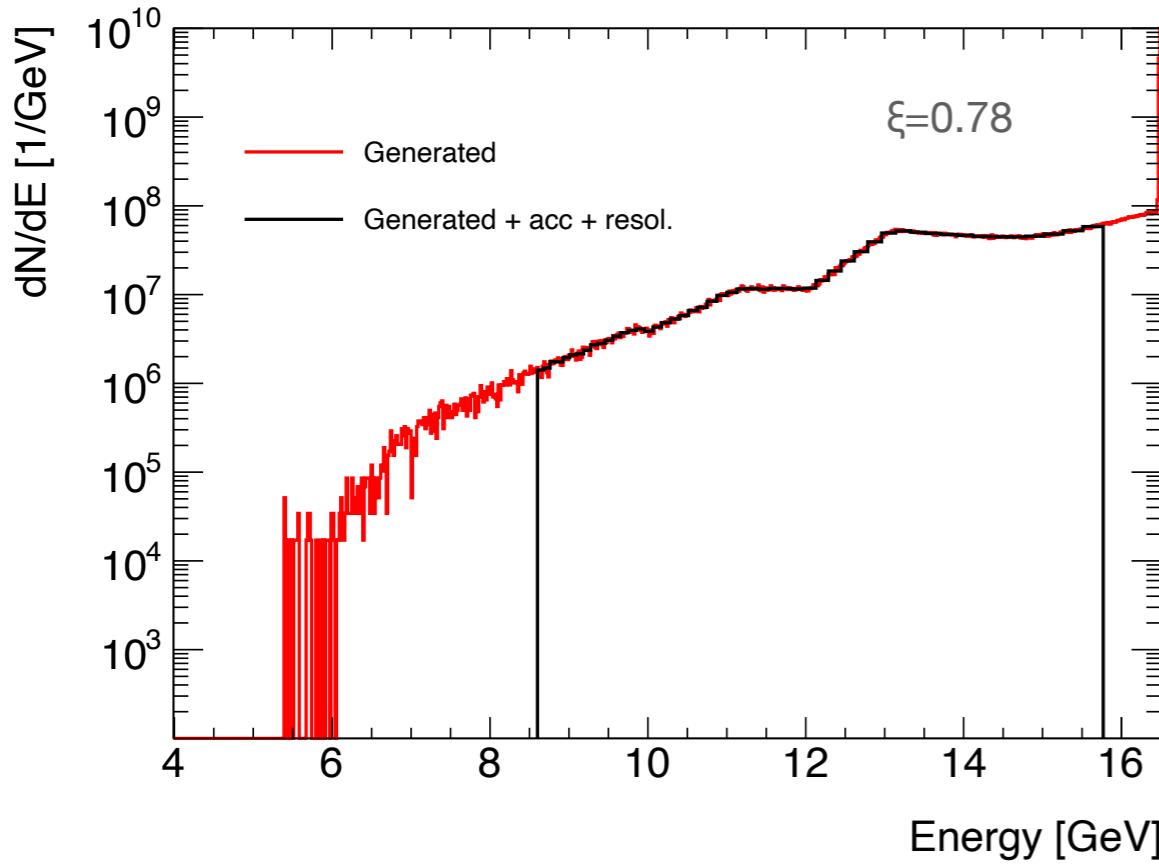
- location **L**: edge position, maximum of Response
- kinks **K_{up/down}**: edge start/end points, determined by finding 95% interval of Response function
- alternative **K_{up/down}**: zero-crossing of Response
- width **W = K_{up}-K_{down}**
- Note: Response is not necessarily Gaussian!
→ K_{up} and K_{down} can be very asymmetric



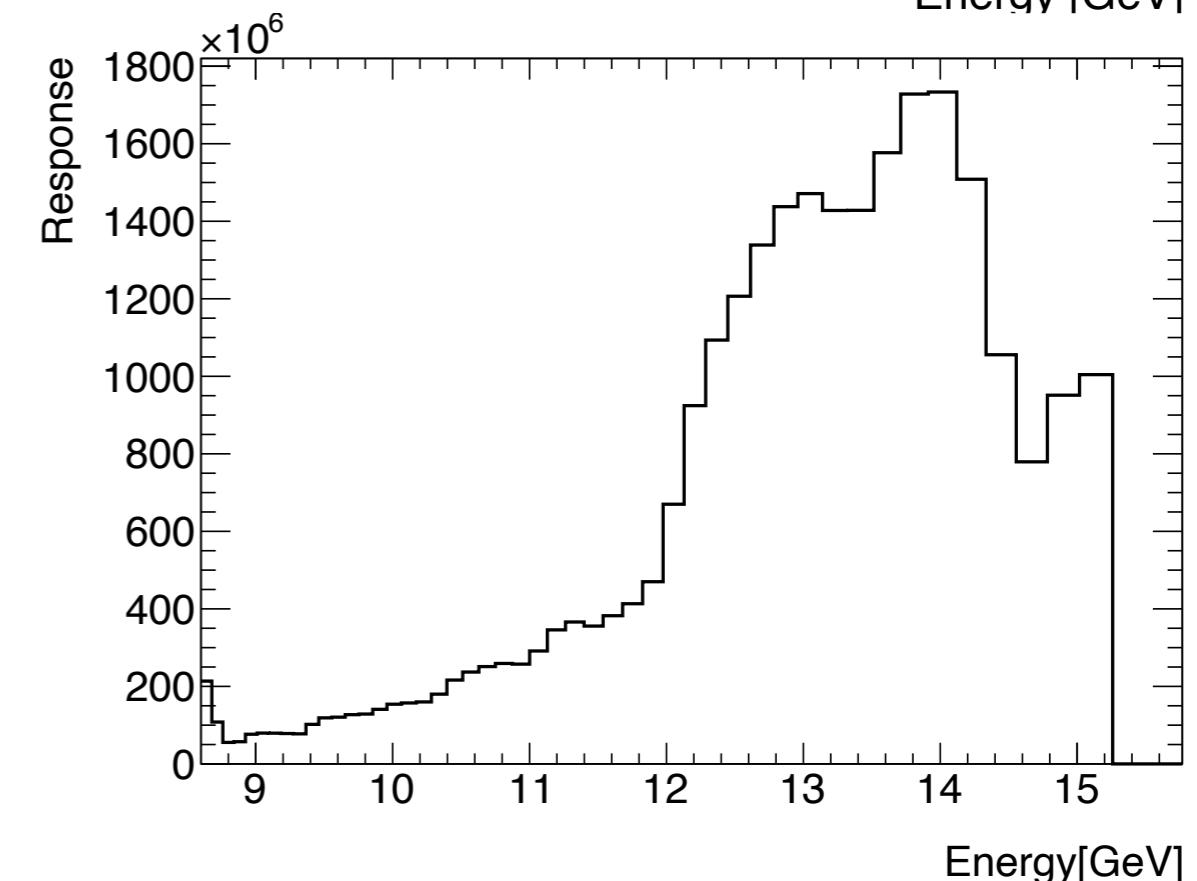
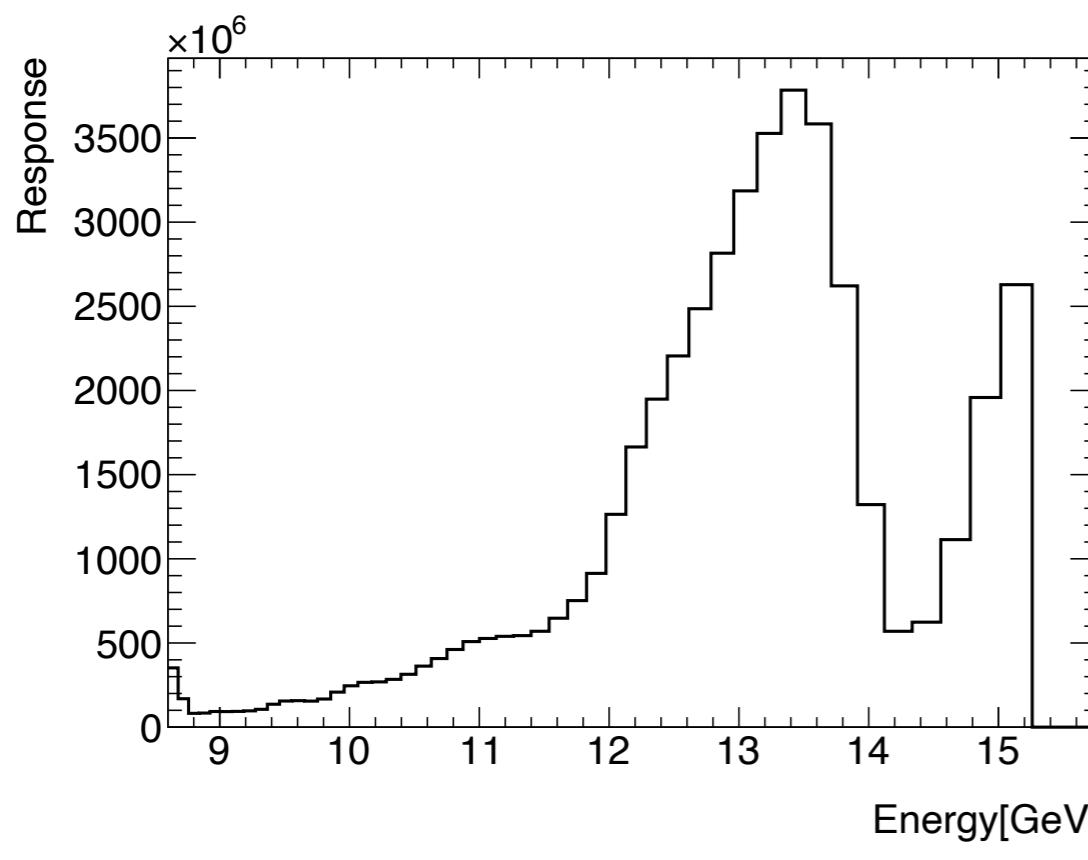
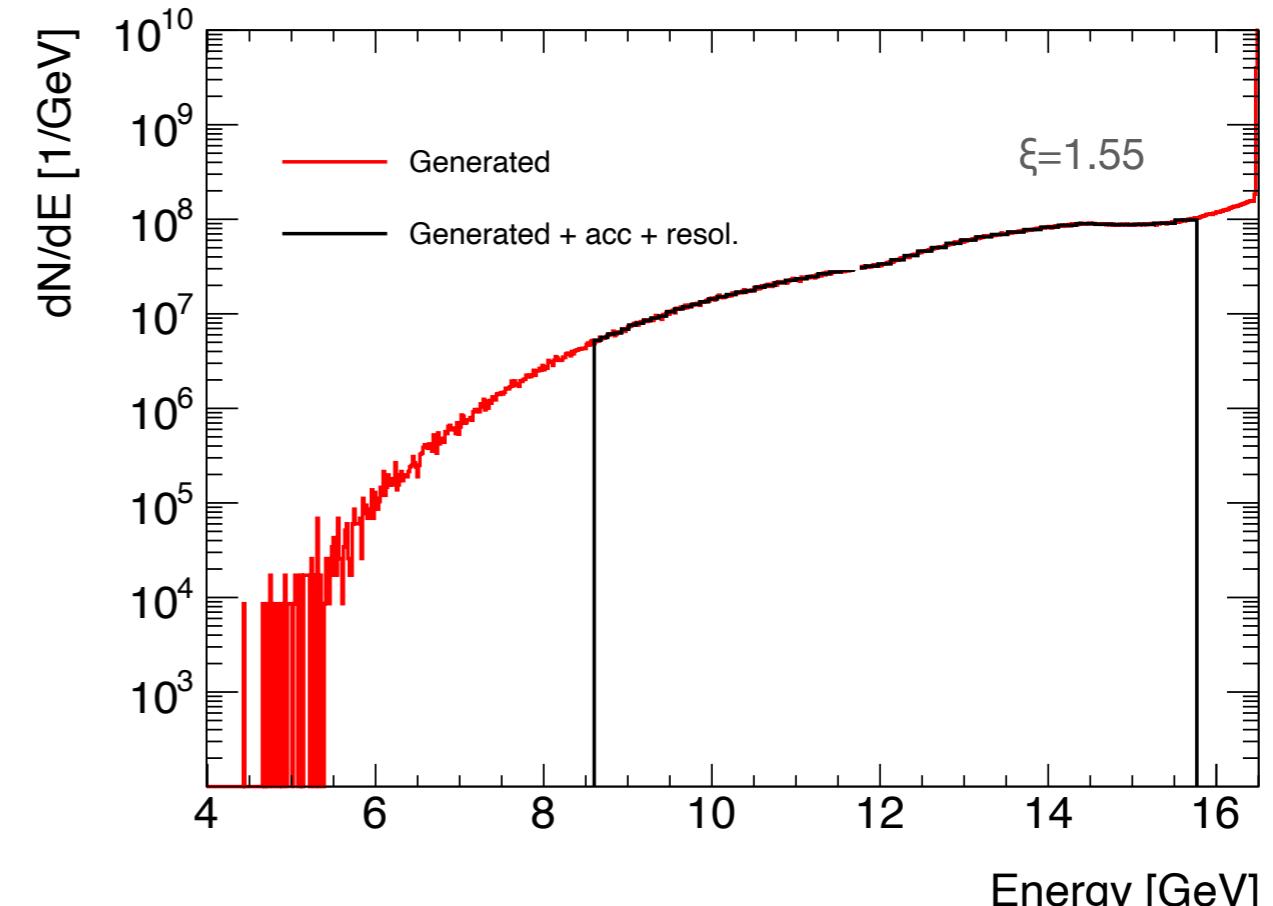
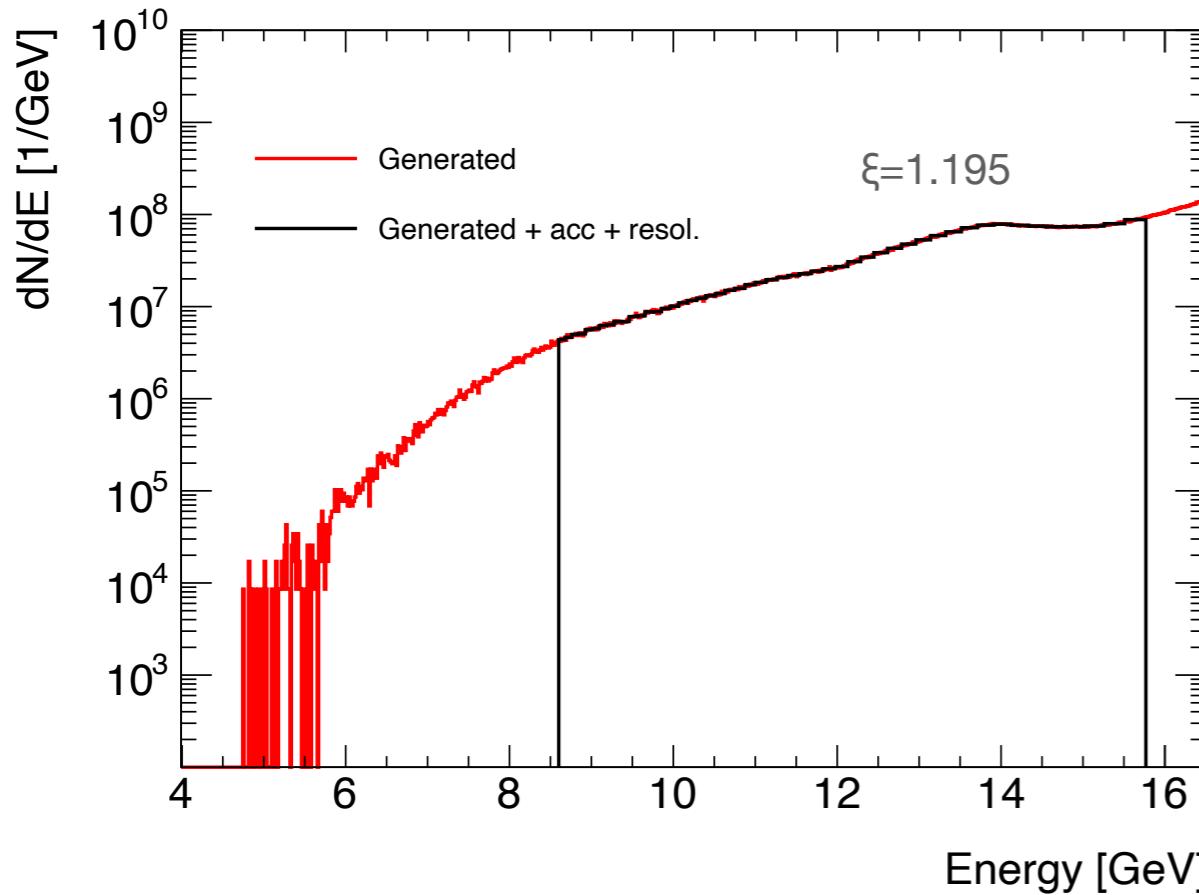
Spectra & Response



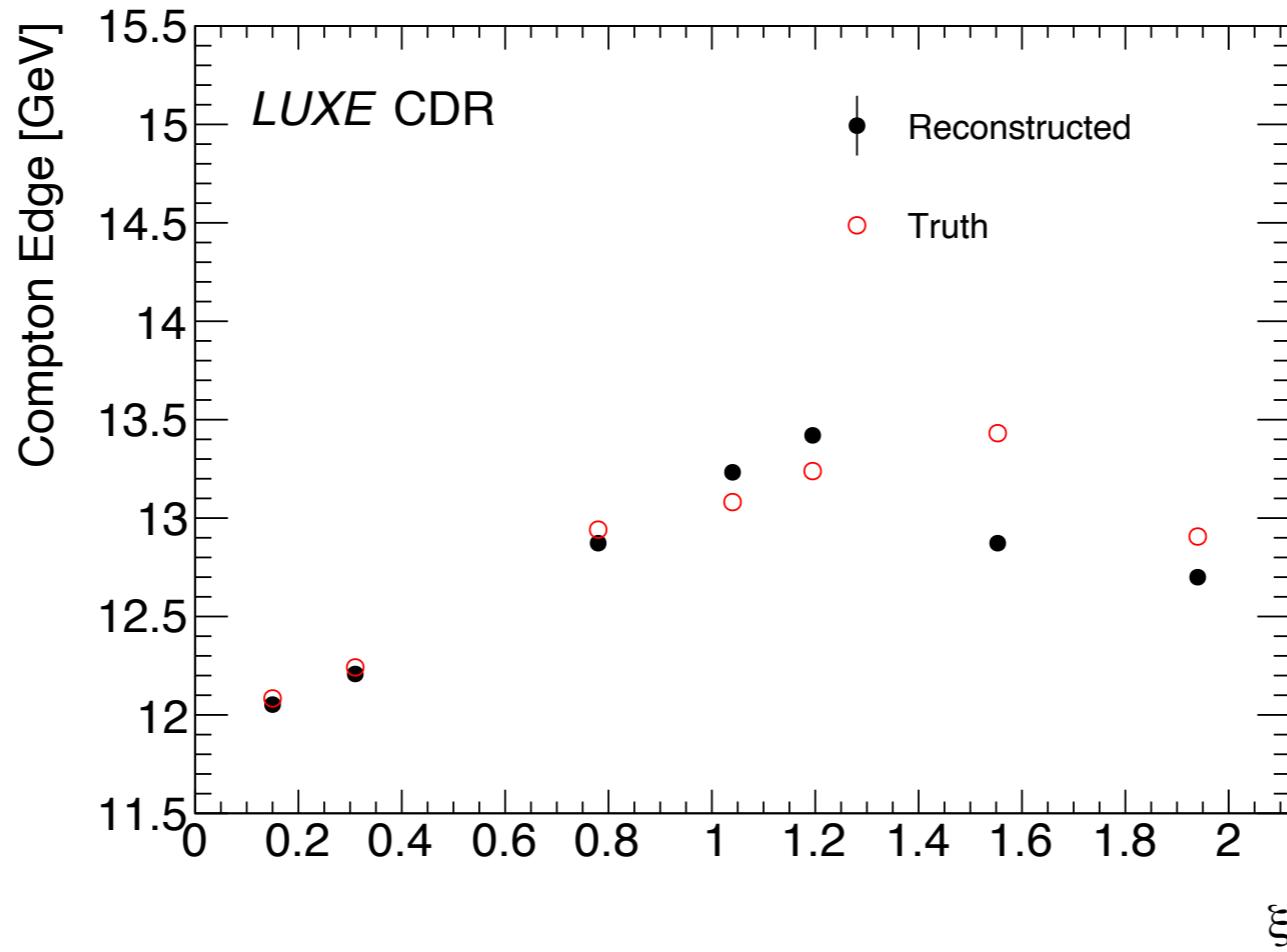
Spectra & Response



Spectra & Response

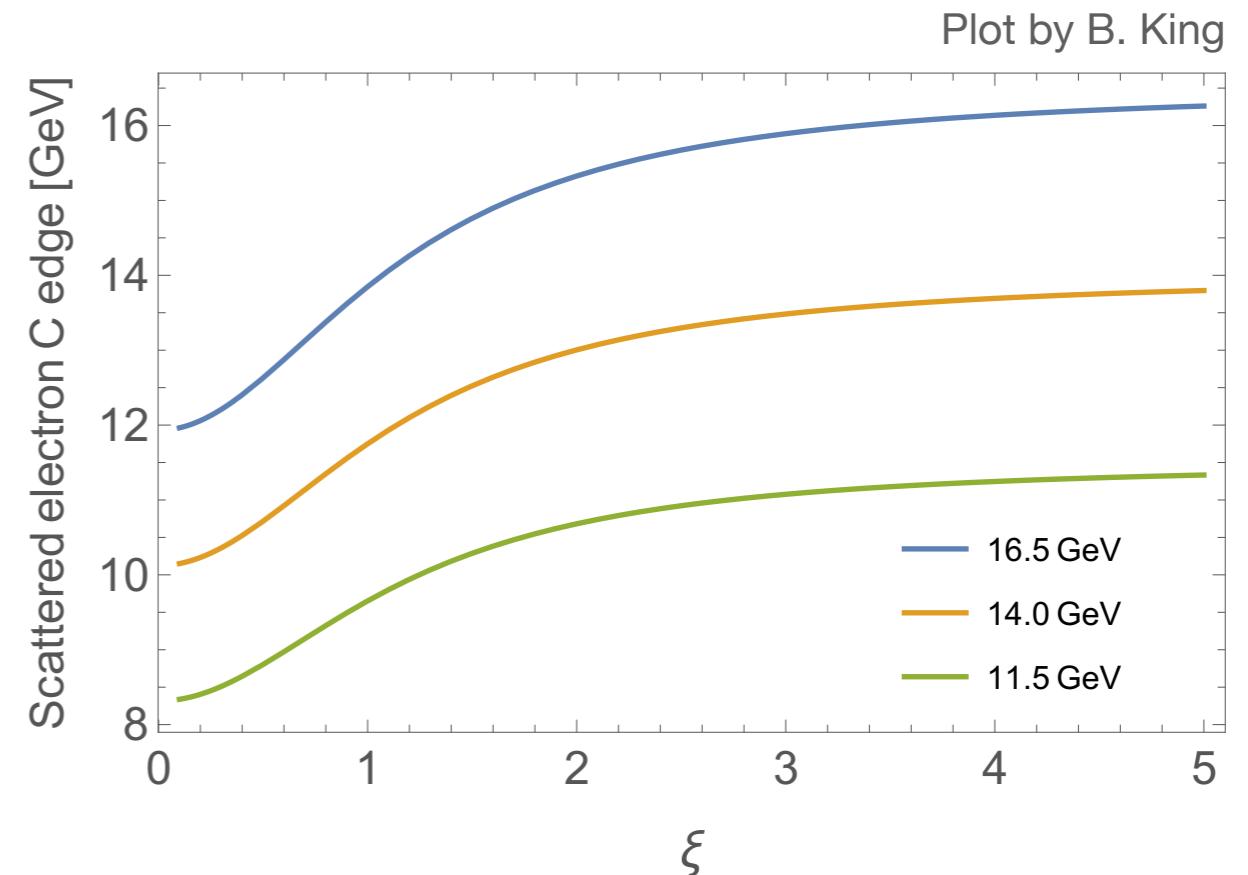
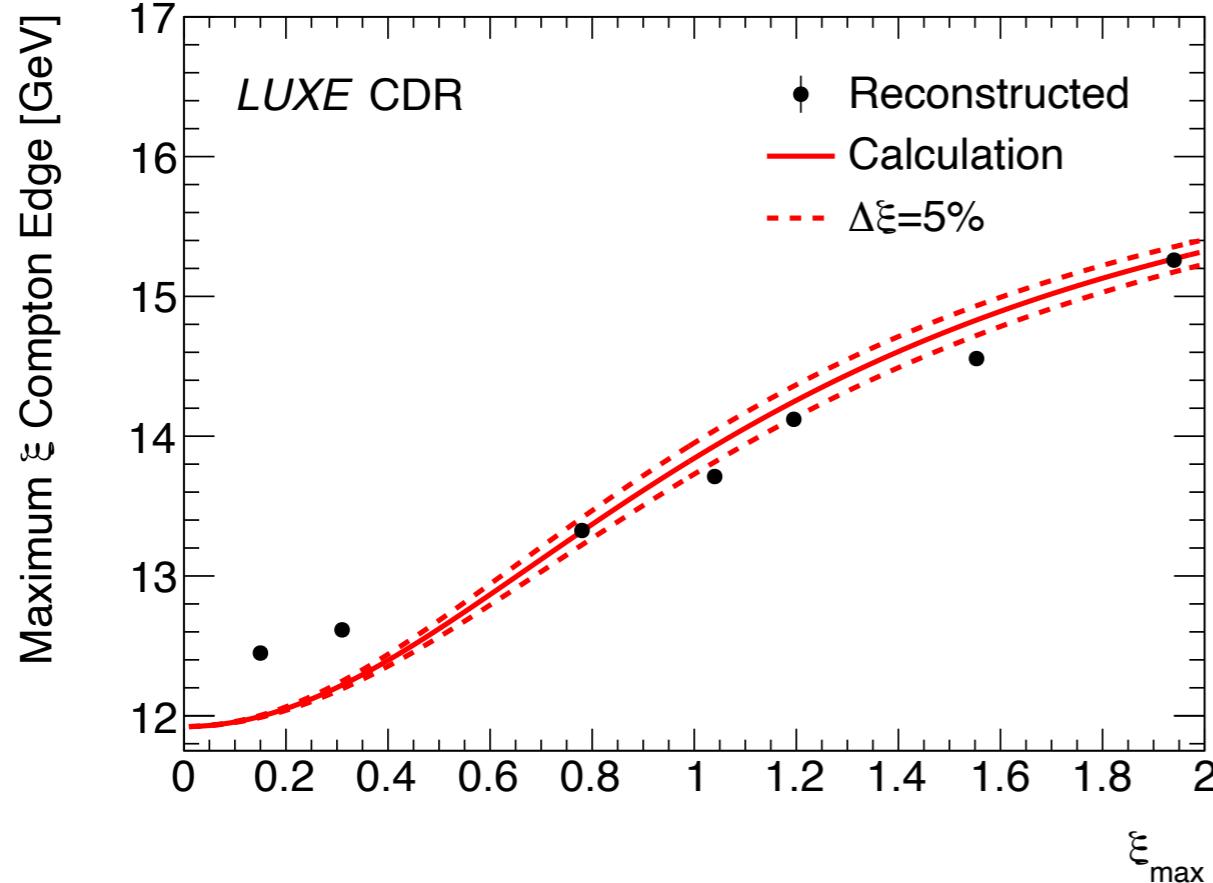


Edge Position (L) Reco vs. Generated



- Closure test: difference between generated and reco here is only due to binning

Kink Position (K_{up})



- the curve shows:

$$f(E)/\text{GeV} = 16.5 \left(1 - \frac{2\eta}{2\eta + 1 + \xi^2}\right) \quad \text{where } \eta = 0.192$$

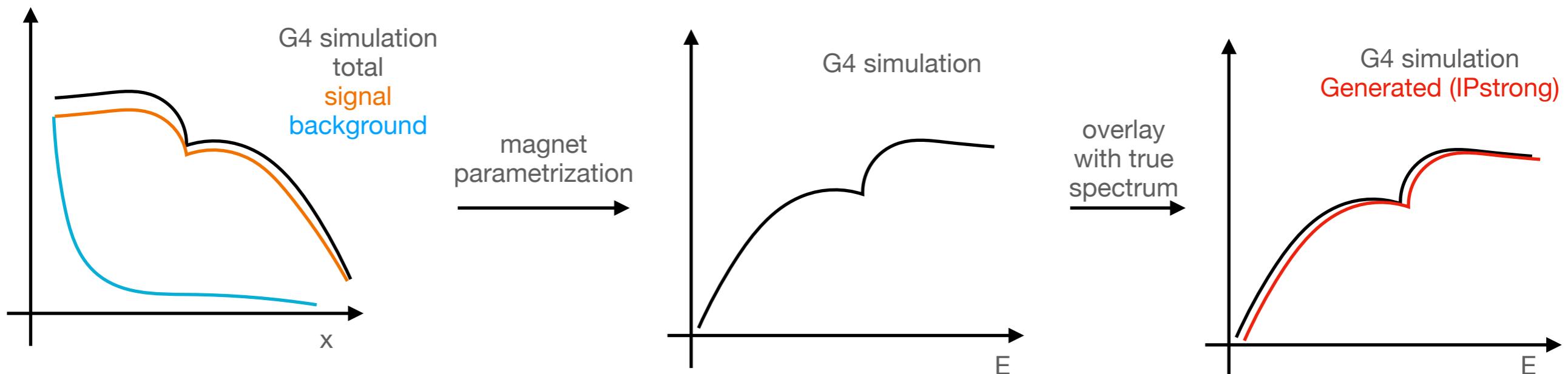
- uncertainty of points not final, slight bias at low ξ

A first look at Cerenkov IP backgrounds

- Sasha ran a special e+LASER G4 sim with 2T magnetic field at $\xi_{\max}=0.78$ (thanks!)

How to extract the "backgrounds":

- 1.) Select G4 events:
 - detid==6300 (hit Cerenkov sensitive volume)
 - pdg==11 (electrons)
 - E>20MeV (Cerenkov threshold)
- 2.) Get x-distribution, convert back to energy using dipole parametrization
- 3.) Overlay with generated (signal-only!) energy spectrum from Tony's MC
→ difference is background

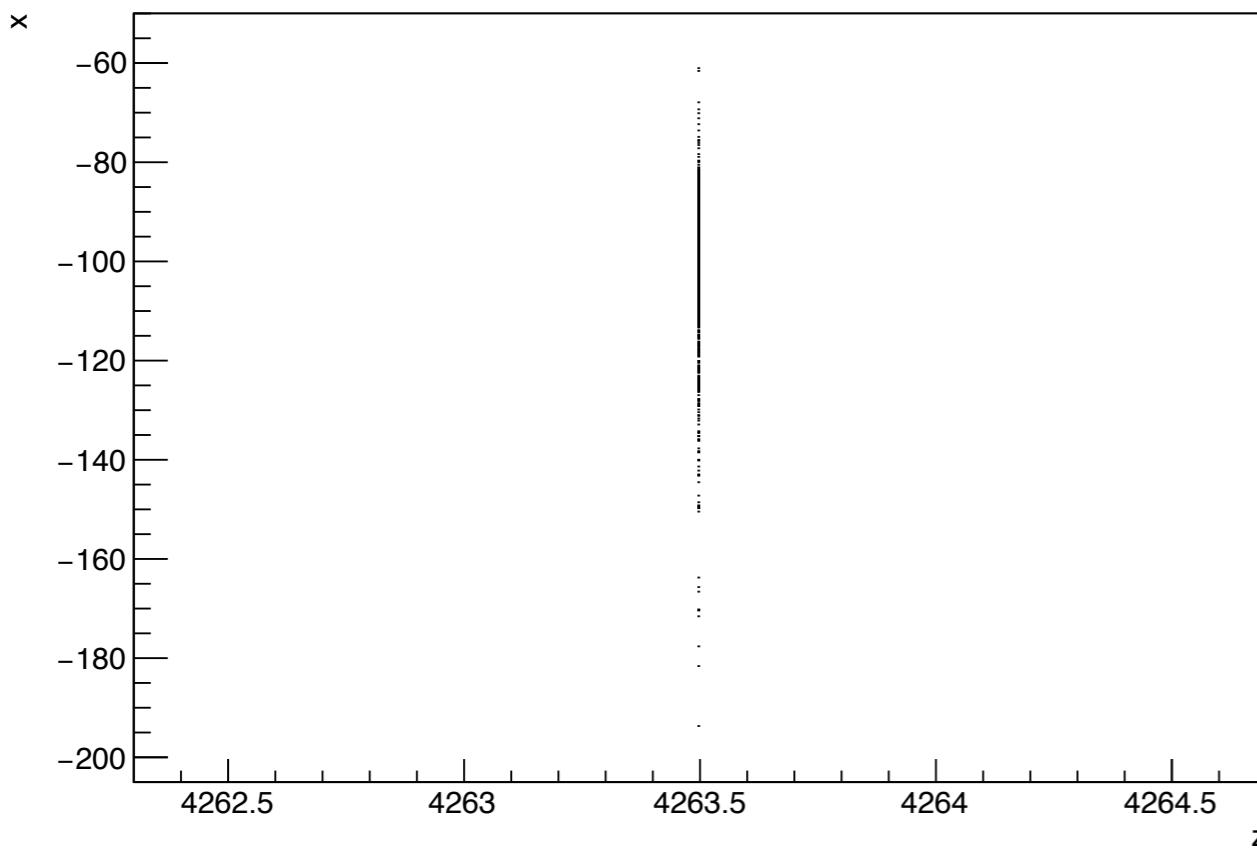


What is a good background estimate?

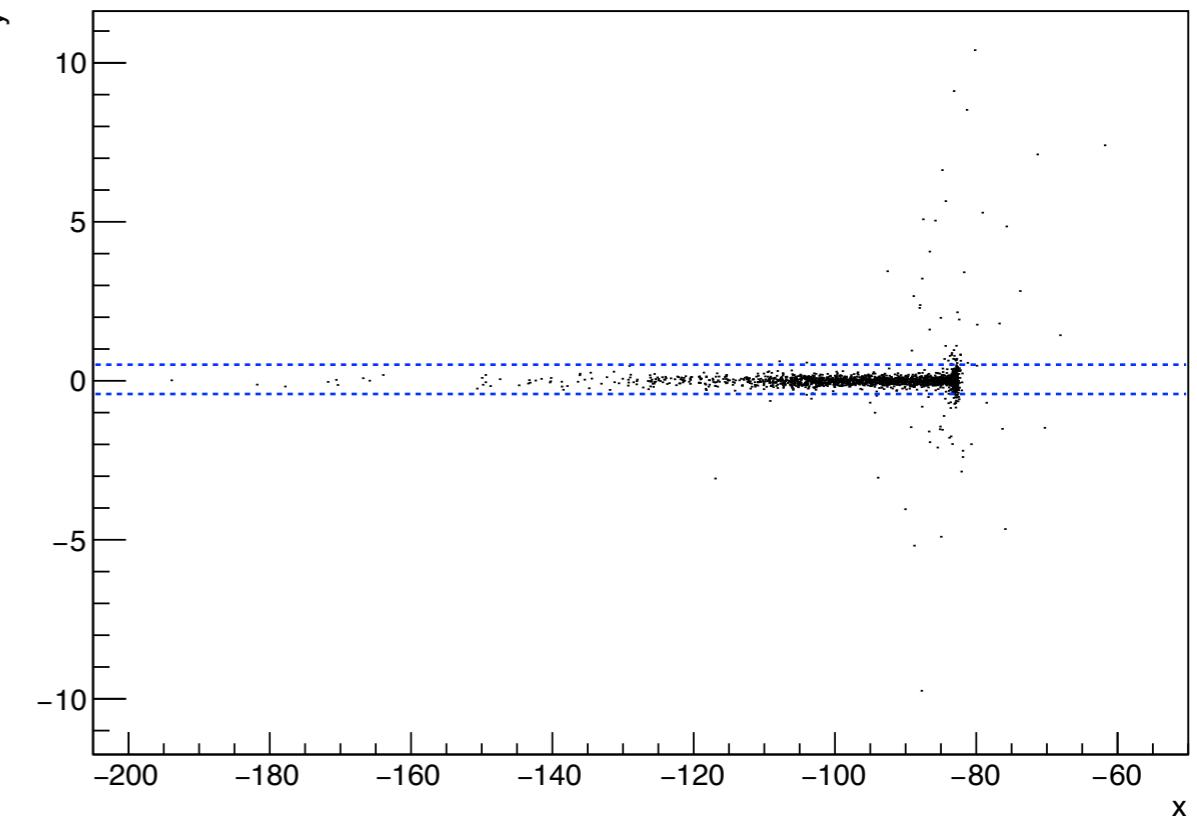
G4Sim-Generated?, Sim(trackid!=1)?

Some Facts about Particles at the IP Cerenkov

x:z {detid==6300 && abs(pdg)==11 && E>0.02}

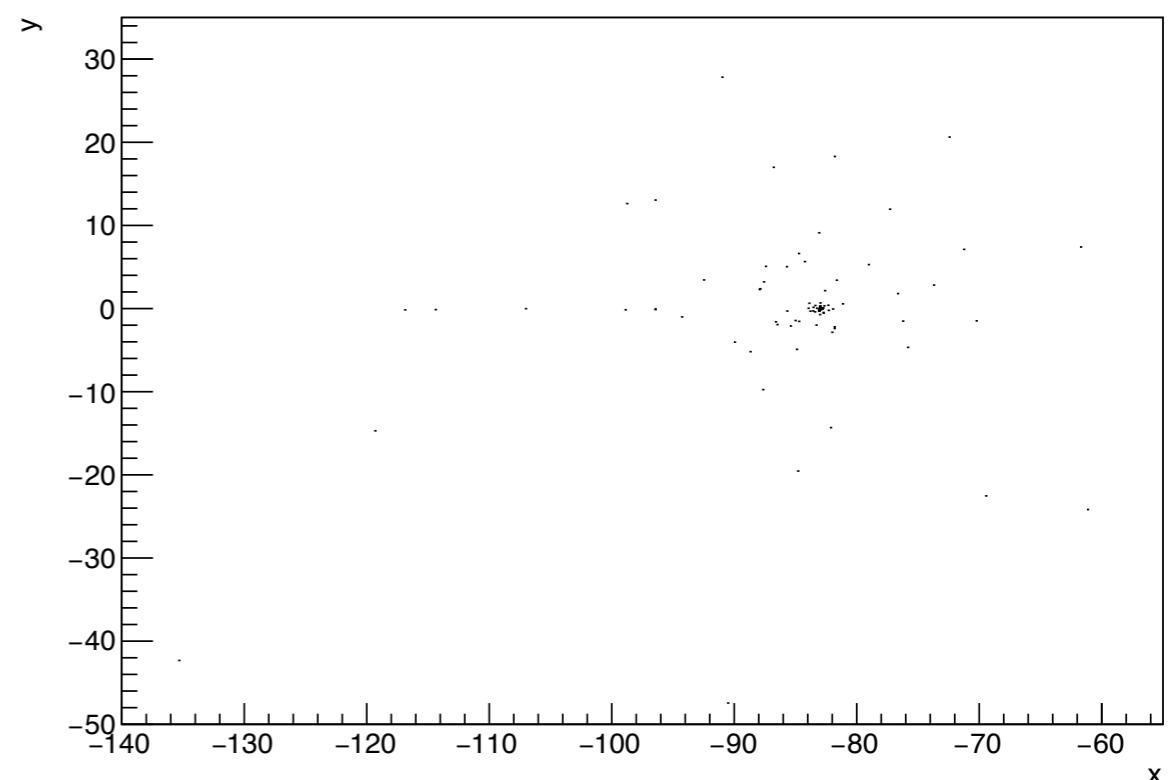


y:x {detid==6300 && abs(pdg)==11 && E>0.02}

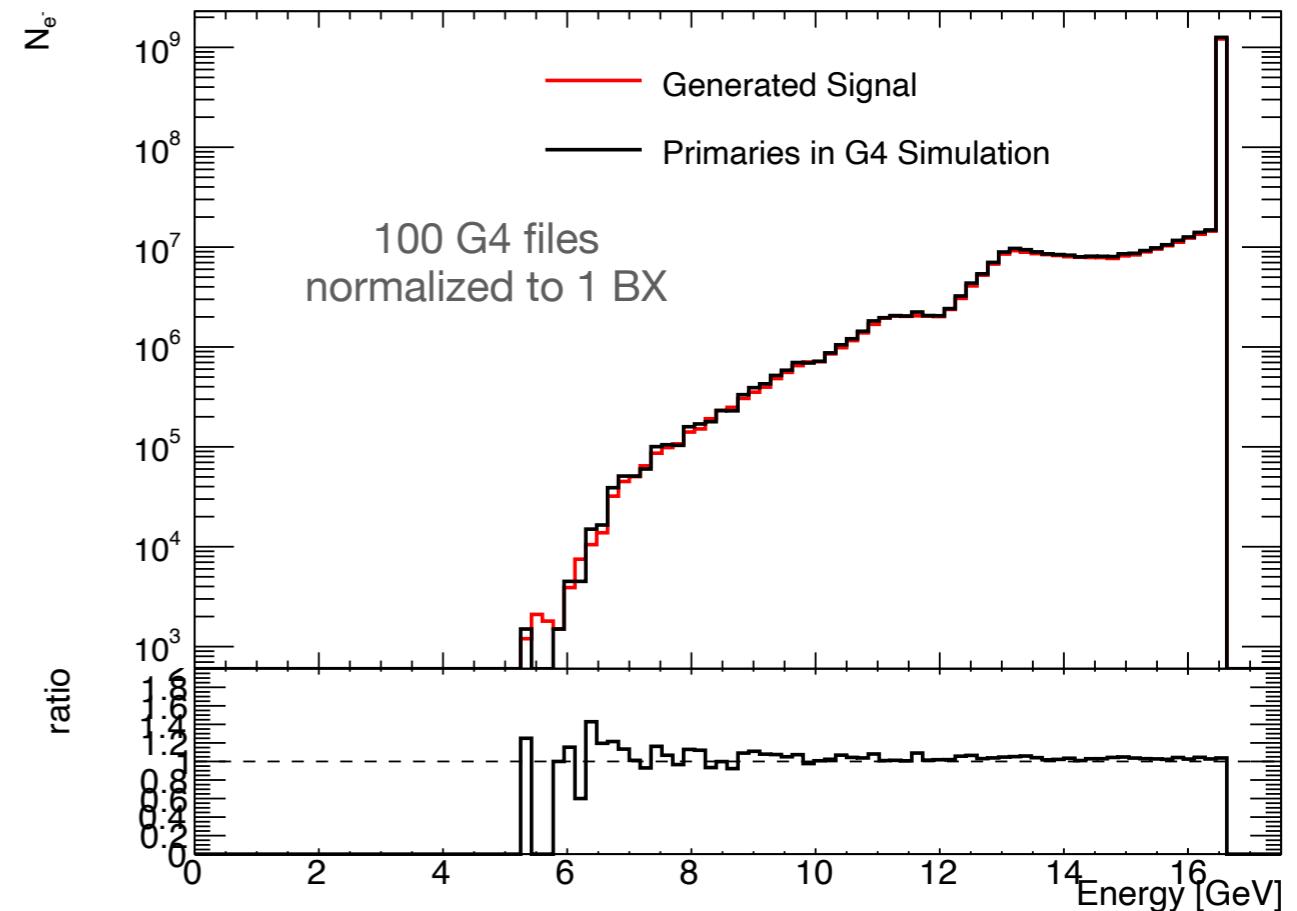
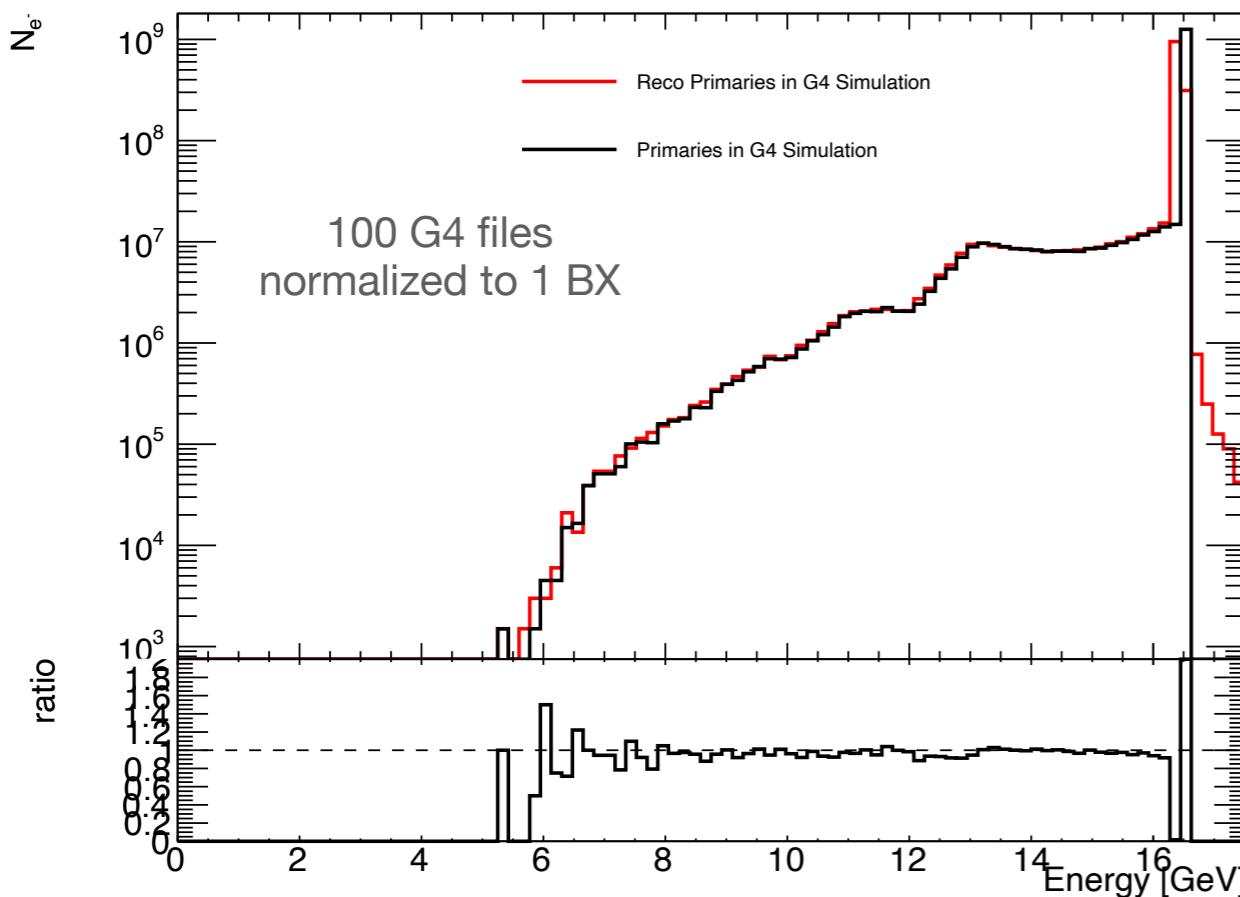


- all particles charged above Cerenkov threshold hit the detector from the front
- particles are very collimated
- most “non-primeries“ close to the beam

y:x {detid==6300 && abs(pdg)==11 && E>0.02 && trackid!=1}



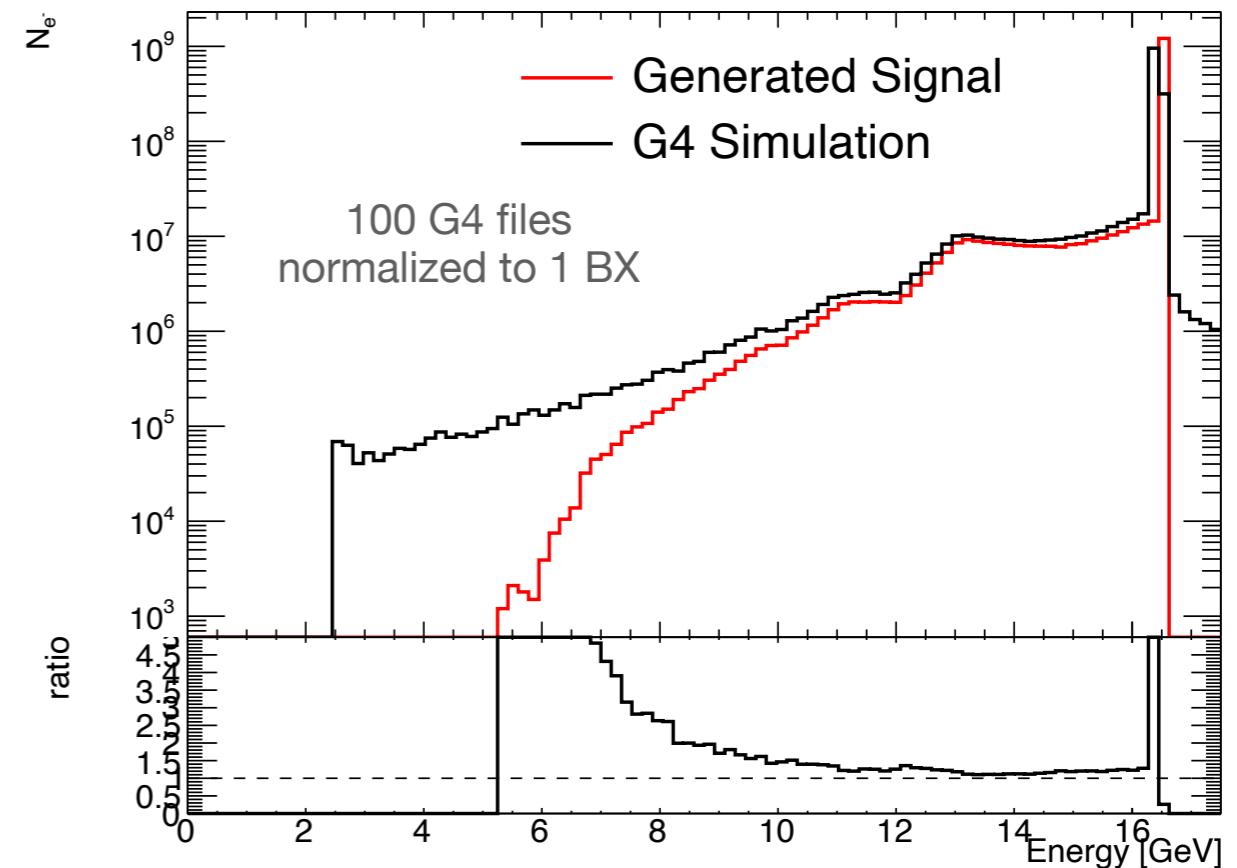
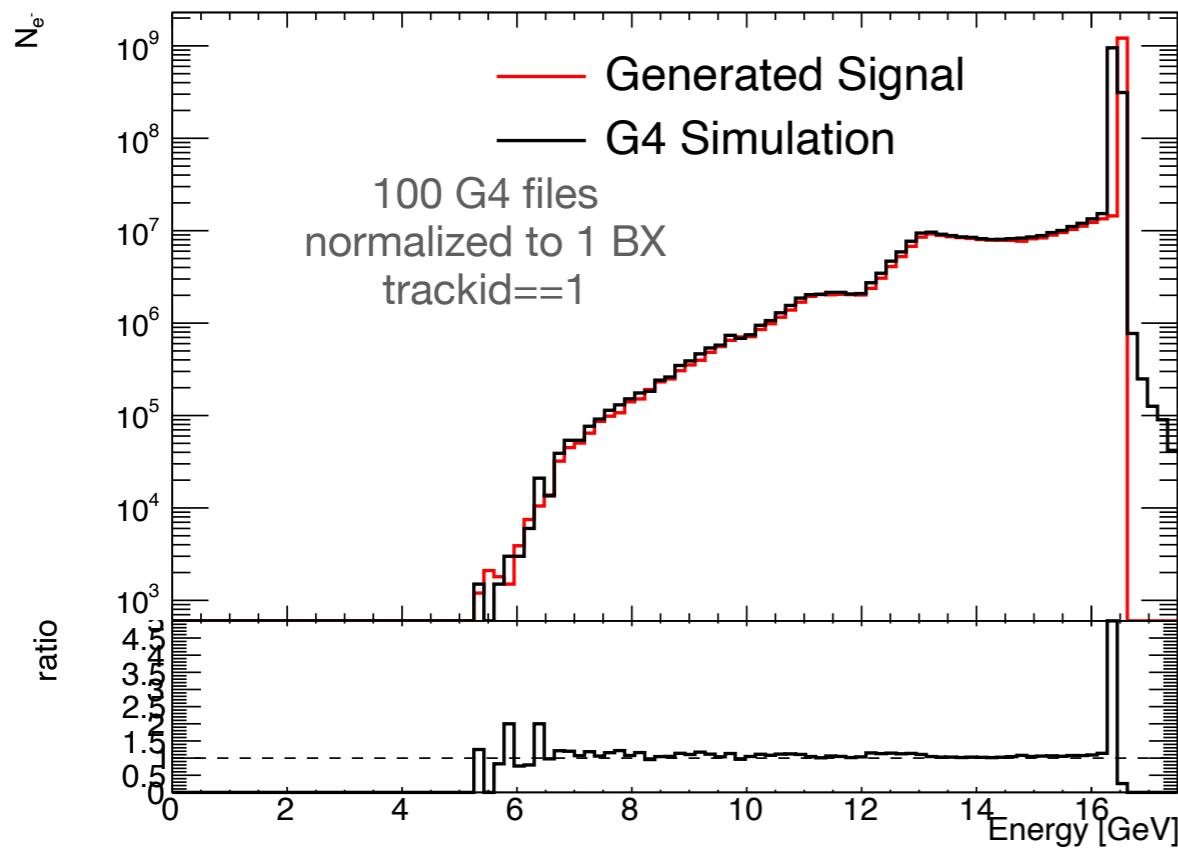
Sanity Cross-Checks



- “reco primaries”: reconstructed from x-dist. of particles with trackid==1
- primaries: particles with detid==−1
- difference: difference between simple B-field parametrization & actual field?

- generated: particles from Tonys MC
- primaries: particles with detid==−1

At the IP Cerenkov



- difference between generated