

Track Reconstruction in Liquid Scintillators with Graph Neural Networks

Rosmarie Wirth

November 12, 2020

Short reminder on neutrino physics

Neutrino oscillation and masses

- Neutrinos only interact weakly
- Neutrinos have masses
- Neutrinos can change flavor
- Different mass and flavor eigenstates

Open Questions:

- Mass hierarchy
- CP violation
- Majorana particle

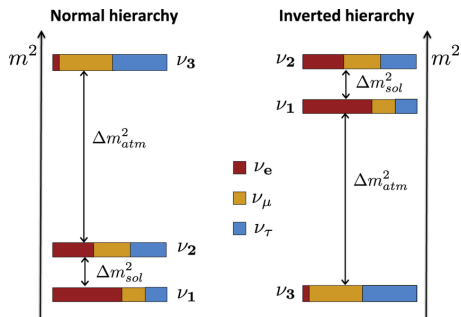


Figure: mass hierarchy

arXiv:1310.4340

Neutrino observation with scintillation detectors

Neutrino detection:

Inverse β Decay: $\bar{\nu}_e + p \rightarrow \bar{e} + n$

→ Positron immediately gives
signal

→ Neutron releases γ after

$\bar{t} \approx 200\mu s$

Background:

- Cosmic muons triggering cosmogenics

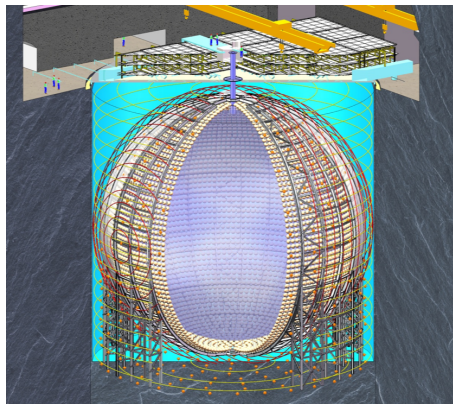


Figure: Jiangmen Underground Neutrino Observatory (JUNO),

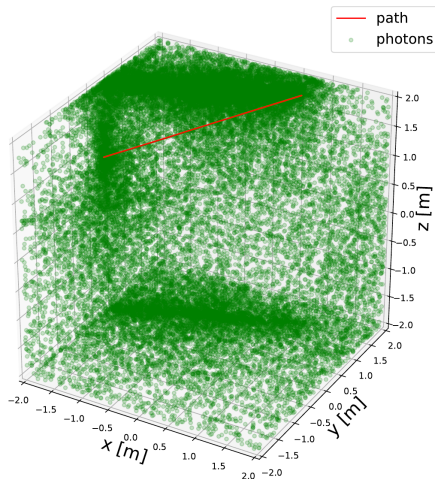
<https://www.etap.physik.uni-mainz.de/research-groups/ex-juno/> 30.07.2020

Muon path reconstruction in liquid scintillators with artificial neural networks

→ Reconstruct photon emission distribution from PMT response

- Reconstruction is crucial in neutrino detection
- Spatial muon tracking can reduce veto volume
- Shower identification could work as spatial veto

TOY - Monte Carlo



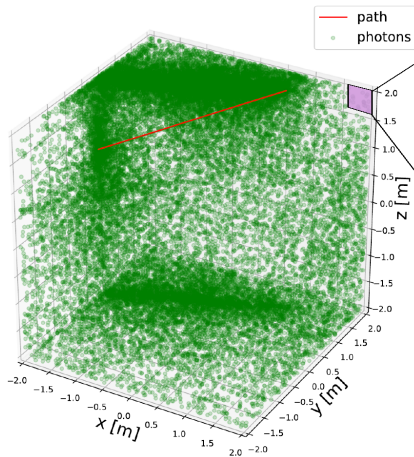
Simulation:

- 4 x 4 x 4 m
- Random, straight path
- 10 photons/ 0.001 m
- Peak with 5000 photons
- Absorption with $mfw = 80m$
- Scattering with $mfw = 25m$

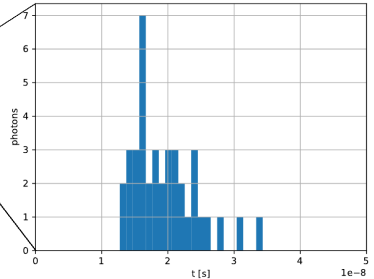
Not included:

- Cherenkov light
- Attenuation
- Other path types

Data set



PMT signal:



- $10 \times 10 \times 6$ PMTs
- time resolution 1 ns

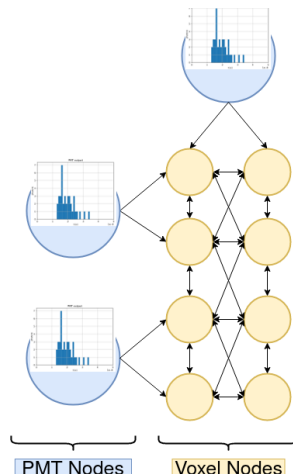
→ 600x50 data points per event

Architecture Idea

- Photon distribution representation by $[20,20,20]$ voxels
- Network should propagate photons back to their origin

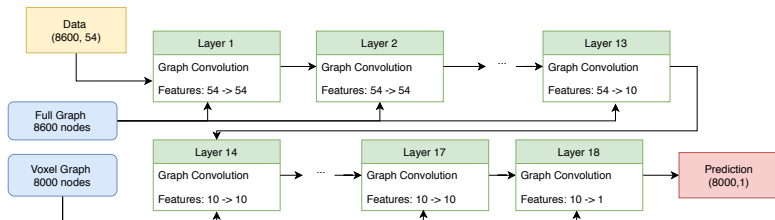
Full Graph

- 8600 nodes (8000 voxels + 600 PMTs)
- Each PMT node has edges to the closest voxel nodes
- Each voxel node has 26 edges to neighbors



→ Graph resembles geometrical structure of the setup

Architecture



→ Implemented in python with pytorch and dgl

Graph Convolution Layers:

$$\nu_i^{(l+1)} = f(b^{(l)} + \sum_{j \in \mathcal{N}(i)} \frac{1}{c_{ij}} \nu_j^{(l)} \cdot W^{(l)}) \quad (1)$$

ν_i - node, $\mathcal{N}(i)$ - neighboring nodes, b bias, W - weights, l - layer

Wasserstein Loss

Wasserstein distance: metric defined between probability distributions on a metric space

Sinkhorn distance:

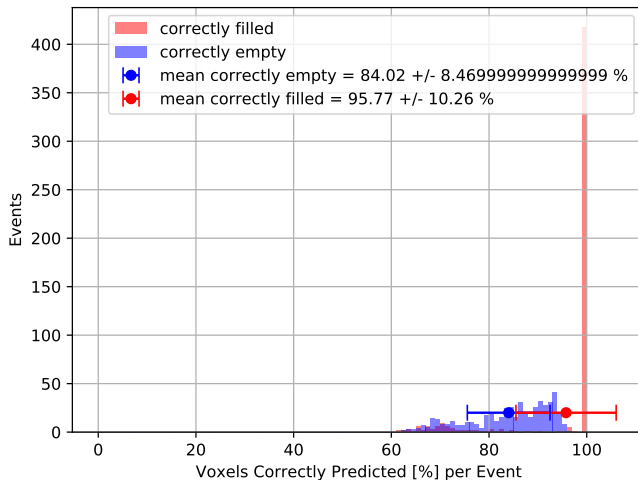
$$d_{M,\alpha}(r, c) = \min_{P \in U_\alpha(r, c)} \langle P, M \rangle \quad (2)$$

with M - cost matrix, P - transport matrix, r and c distributions, $U(r, c)$ joint probability matrices

- Reshape form [20,20,20] into 3D point cloud with photons as weights
- 3D point cloud uses coordinates of the filled voxels
- Supports autograd

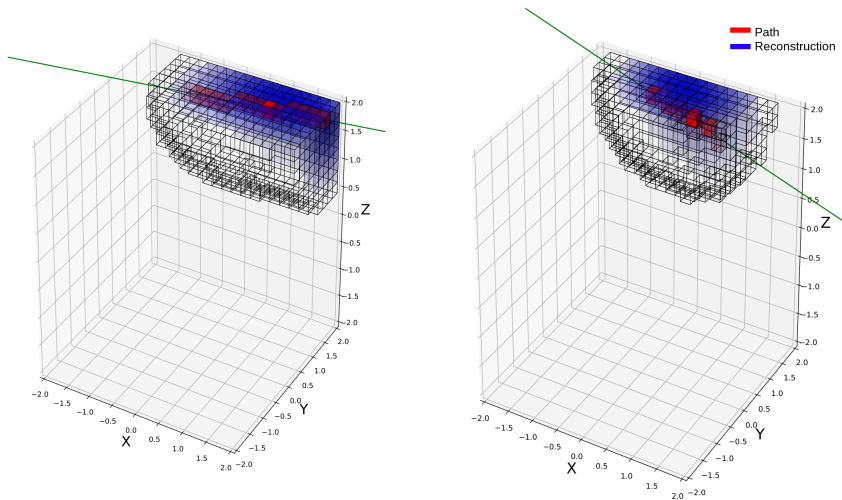
→ Including spacial gradients

Results - work in progress



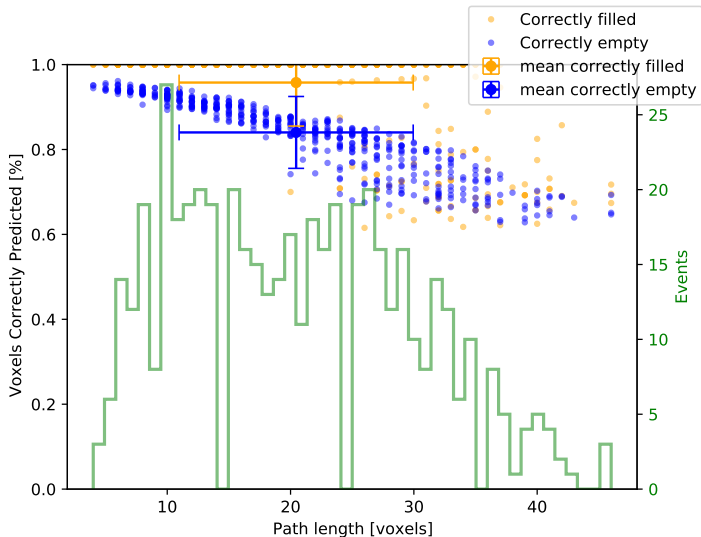
→ Only $\approx 16\%$ of the detector volume would need to be vetoed

Results - work in progress



→ Reconstruction surrounds path, but too many voxels are filled

Results - work in progress



→ Especially short paths are reconstructed nicely

→ **Graph Neural Networks can reconstruct muon tracks, but there is room for improvement**

Next Steps:

- Further improvement of the predictions
- Improve photon sum
- Include more complex paths
- More complex simulation

Outlook:

- Could be used to improve the muon veto by a spacial information
- Architecture could be adjusted for different detector setups

Thank you!