Generative Models and Physics Applications Sascha Diefenbacher,



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Generative Models

CLUSTER OF EXCELLENCE QUANTUM UNIVERSE





Literature

- Further Reading:
 - Mehta et. al., A high-bias, low-variance introduction to Machine Learning for physicists, 1803.08823
 - Goodfellow et. al., Generative Adversarial Networks 1406.2661

 - Arjovsky et. al., Wasserstein GAN, 1701.07875 • Gulrajani et. al., Improved Training of Wasserstein GANs 1704.00028





Classification High Dim. Data





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Latent Space Noise

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Generation







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- Classification:
 - Easy to interpret model outputs
 - Model performance simple to measure
 - Accuracy, Area Under Curve, ...









- Classification:
 - Easy to interpret model outputs
 - Model performance simple to measure
 - Accuracy, Area Under Curve, ...
 - Wide range of available loss functions
 - Only needs to compare predictions with true labels
 - Common example: Cross-Entropy







- Generative:
 - How do you measure the model performance?







How is this expressed mathematically (and differentiable)





- Generative:
 - How do you measure the model performance?







• How is this expressed mathematically (and differentiable)



http://thesecatsdonotexist.com





Generation Difficulties

- Image Set:











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Does the new set have the same properties as the data?





Generation Difficulties

- Image Set:











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Does the new set have the same properties as the data?

Generated Data:









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AutoEncoder



Encoding function E(x)=z map high dimensional data X to low dimensional latent space Z









AutoEncoder



- Encoding function E(x)=z map high dimensional data X to low dimensional latent space Z
- Decoding function D(z)=x map latent space Z back to data X







AutoEncoder



- Encoding function E(x)=z map high dimensional data X to low dimensional latent space Z
- Decoding function D(z)=x map latent space Z back to data X Compare Input and Output pixel by pixel with mean squared error

Generative Models







• Sample for Z and pass it to $D(Z) \rightarrow C$ Generate new samples











• Sample for Z and pass it to $D(Z) \rightarrow C$ Generate new samples Problem: Need regularised later space to sample form Variational AutoEncoder





Variational AutoEncoder







• Latent space: Series of Gaussians, regularised match N(μ =0, σ =1)







Variational AutoEncoder



 Using Gaussians lets us use Kullback–Leibler divergence $\sum \sigma_i^2 + \mu_i^2 - \log(\sigma_i) - 1$ i=1



• Latent space: Series of Gaussians, regularised match N(μ =0, σ =1)







Variational AutoEncoder



- - Using Gaussians lets us use Kullback–Leibler divergence $\sum \sigma_i^2 + \mu_i^2 - \log(\sigma_i) - 1$ i=1
- Compare Input and Output again using MSE

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• Latent space: Series of Gaussians, regularised match N(μ =0, σ =1)







Generative Adversarial Network High Dim. Data

Latent Space Noise

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- Generator Network G(z)=x
 - Maps noise Z to Data X











- Generator Network G(z)=x
 - Maps noise Z to Data X
- Discriminator D(G(z)) and D(x)
 - Learns difference between real and fake





$$L = BCE(p(real), 1) + BCE(p(fake), 0)$$

Generative Models





- Generator Network G(z)=x
 - Maps noise Z to Data X
- Discriminator D(G(z)) and D(x)
 - Learns difference between real and fake
- D(G(z)) is differentiable function measuring performance



$$L = BCE(p(real), 1) + BCE(p(fake), 0)$$

Generative Models







- Generator Network G(z)=x
 - Maps noise Z to Data X
- Discriminator D(G(z)) and D(x)
 - Learns difference between real and fake
- D(G(z)) is differentiable function measuring performance
- Use D(G(z)) as loss to update G



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Code Example







(a)

(b)

- Guides Generated distribution to match real distribution

Generative Adversarial Network



Goodfellow et al.- arXiv:1406.2661

Discriminator of GAN approximates Jensen Shannon Divergence







- JSD only based on overlap
 - Problem for very separated distributions
 - Vanishing Gradient
- Alternative to JSE: Earth Mover Distance (aka Wasserstein Distance)



Arjovsky et al.- arXiv:1701.07875

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- Earth Mover Distance
- Optimal transport problem
- Most energy efficient way to match two distributions
 - Energy defined as (mass times distance)



https://lilianweng.github.io/lil-log/2017/08/20/from-GAN-to-WGAN.html









- Earth Mover Distance
- Optimal transport problem
- Most energy efficient way to match two distributions
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https://lilianweng.github.io/lil-log/2017/08/20/from-GAN-to-WGAN.html







- Earth Mover Distance
- Optimal transport problem $\frac{1}{2}$
- Most energy efficient way to match two distributions
 - Energy defined as (mass times distance)
- Sensitive to distance





https://lilianweng.github.io/lil-log/2017/08/20/from-GAN-to-WGAN.html

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- Earth Mover Distance Direct calculation not feasible $W(p_r, p_g) = \frac{1}{K} \sup_{\|f\|_{L} \leq K} \mathbb{E}_{x \sim p_r}[f(x)] - \mathbb{E}_{x \sim p_g}[f(x)]$
 - Gives upper bound, but needs K-Lipschitz continuous function f

 - Use Critic to estimate EMD

 $W(p_r, p_g) = \max_{x \sim p_r} \mathbb{E}_{x \sim p_r}[f_w(x)]$ w∈W

$$-\mathbb{E}_{z\sim p_r(z)}[f_w(g_\theta(z))]$$







- JSD only based on overlap
 - Problem for very separated distributions
 - Vanishing Gradient
- Alternative to JSE: Earth Mover Distance (aka Wasserstein Distance)
 - Gradient no longer vanishes



Arjovsky et al.- arXiv:1701.07875

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Wasserstein GAN

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Code Example





- - simulation takes significant time and resources





Monte Carlo simulations are invaluable part of particle physics

Detailed, full (matrix element, hadronisation, detector effects)





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- Monte Carlo simulations are invaluable part of particle physics
 - Detailed, full (matrix element, hadronisation, detector effects) simulation takes significant time and resources
- Increasing collider luminosities calls for more Monte Carlo data











- Monte Carlo simulations are invaluable part of particle physics
 - Detailed, full (matrix element, hadronisation, detector effects) simulation takes significant time and resources
- Increasing collider luminosities calls for more Monte Carlo data Need a way to speed up simulations











- Monte Carlo simulations are invaluable part of particle physics
 - Detailed, full (matrix element, hadronisation, detector effects) simulation takes significant time and resources
- Increasing collider luminosities calls for more Monte Carlo data
 - Need a way to speed up simulations
- Train Generative Network to emulate simulators
 - Networks are significantly faster than classical methods







Generative Models in Physics Training Real data, Data **Discr. N** points Noise-Generator













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Generative Models



Potential Problem Info(N real points) Info(M GANed points)

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Potential Problem Info(N real points) = Info(M GANed points) Little advantage to be gained from GAN











Potential Problem Info(N real points) = Info(M GANed points) Little advantage to be gained from GAN Info(N real points) < Info(M GANed points)













Potential Problem Little advantage to be gained from GAN GAN can speed up simulations

Info(N real points) = Info(M GANed points) Info(N real points) < Info(M GANed points)













Potential Problem Little advantage to be gained from GAN GAN can speed up simulations Check with small scale test

Info(N real points) = Info(M GANed points) Info(N real points) < Info(M GANed points)



Butter et al.: GANplifying Event Samples: 2008.06545









1-D Toy Model

 Camel back function: double peak Gaussian $p(X) = \frac{1}{2}(N_{-4,1}(x) + N_{4,1}(x))$ 0.09 0.08 0.07 0.06 0.05 (<u>×</u> 0.04 0.03 0.02 0.01

0.00

-6

-8

-4

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Quantiles

 Need measurement how well function is described







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Quantiles

- Need measurement how well function is described
- Define 10 quantiles on true distribution







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Quantiles

- Need measurement how well function is described
- Define 10 quantiles on true distribution
- Each quantile contains equal probability







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Parameter Fit

• Gives upper performance benchmark







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Parameter Fit

- Gives upper performance benchmark
- Fit 5 parameter camel back function to training samples

$$p(X) = a N_{\mu_1,\sigma_1}(x) + (1-a)N_{\mu_2,\sigma_2}(x)$$





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Parameter Fit

- Gives upper performance benchmark
- Fit 5 parameter camel back function to training samples

$$p(X) = a N_{\mu_1,\sigma_1}(x) + (1-a)N_{\mu_2,\sigma_2}(x)$$

 Analytically calculate integral for each quantile



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Train GAN on 100 data points from training sample











Train GAN on 100 data points from training sample





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- Train GAN on 100 data points from training sample
- Regularisation through
 - Dropout
 - Added training noise
 - Batch-statistics











- Generate $O(10^7)$ data points using GAN
- Calculate fraction of points in each quantile
- Calculate Quantile MSE







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 Calculate MSE for 100 independent training samples







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- Calculate MSE for 100 independent training samples
- Train GANs on samples and calculate MSE on for these







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- Calculate MSE for 100 independent training samples
- Train GANs on samples and calculate MSE for these
- Reference lines for 200 and 300 samples







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- Calculate MSE for 100 independent training samples
- Train GANs on samples and calculate MSE for these
- Reference lines for 200 and 300 samples
- Calculate MSE for fits on training samples





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- GAN describes distribution better than training data
- Needs 10,000 GANed points to match 150 true points







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How is this possible?







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- How is this possible?
- In terms of information:
 - sample: only data points
 - fit: data + true function
 - GAN: data + smooth, continuous function





Generative Models



- How is this possible?
- In terms of information:
 - sample: only data points
 - fit: data + true function
 - GAN: data + smooth, continuous function
- This allows the GAN to interpolate



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Statistical Properties

- For interpolatable dataset:
- Info(N real points) < Info(M GANed points) GANs have potential to amplify dataset Highly promising for application as simulation
- accelerators



Butter et al.: GANplifying Event Samples: 2008.06545







Simulator	Hardware	Batchsize	Time/shower*
GEANT4	CPU	N/A	4082 ± 170 ms



* average time for 10-100 GeV showers

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Shower Simulation

- Classically done using Geant4
- First principle simulation modelling individual particle interactions
- Very computationally expensive
- Timing even more significant for higher luminosities
- Significant resources needed





Calorimeters



http://ihp-lx.ethz.ch/CompMethPP/lhc/pictures/

- Homogeneous Calorimeter
 - Made up of single active block
 - Great energy resolution
 - No spatial information









http://ihp-lx.ethz.ch/CompMethPP/lhc/pictures/

- Homogeneous Calorimeter
 - Made up of single active block
 - Great energy resolution
 - No spatial information





- Sampling Calorimeter
 - Alternating passive and active layers
 - Only part of energy directly observed
 - Maintains spacial structure





ILD Calorimeter



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 International Large Detector (ILD) Detector for International Linear Collider (ILC)





ILD Calorimeter



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- International Large Detector (ILD)
 - Detector for International Linear Collider (ILC)
- ILD electromagnetic calorimeter
 - Highly granular sampling calorimeter
 - Active silicon, passive tungsten
 - 30 layers, 5mm x 5mm cells







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Buhmann et al.: Getting High: High Fidelity Simulation of High Granularity Calorimeters with High Speed: 2005.05334









 Initial VAE tests on 2D version of Data Mean Squared Error to compare input and reconstruction





- Initial VAE tests on 2D version of Data

 - Mean Squared Error to compare input and reconstruction Unable to reproduce outer hits







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MSE between two random sparse images is larger than between random and empty













- random and empty
- For outer shower hits:
 - Presence more important than exact position
- Needs more fitting reconstruction loss than MSE

MSE between two random sparse images is larger than between











Combines VAE and GAN ideas











- Combines VAE and GAN ideas
 - Enhances pixel wise Mean Squared Error loss with GAN-like adversarial network







- Combines VAE and GAN ideas
 - Enhances pixel wise Mean Squared Error loss with GAN-like adversarial network
 - Significant improvement in generated shower quality



Bounded Information Bottleneck AutoEncoder



- Further expansion of the VAE-GAN structure
- Critic network that judges shower quality
- Second critic that judges reconstruction

Slava Voloshynovskiy et al.: Information bottleneck through variational glasses: 1912.00830









Final Post Processor Network for fine tuning

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GAN



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Results

WGAN

BIB-AE

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Buhmann et al.: Getting High: High Fidelity Simulation of High Granularity Calorimeters with High Speed: 2005.05334

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Differential Distributions









Buhmann et al.: Getting High: High Fidelity Simulation of High Granularity Calorimeters with High Speed: 2005.05334

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Differential Distributions









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Differential Distributions

Buhmann et al.: Getting High: High Fidelity Simulation of High Granularity Calorimeters with High Speed: 2005.05334





Linearity and Resolution*



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* actually not the ECAL resolution as not correction for sampling fraction variation performed







Correlations

$m_{1,x}$	1.00			
$m_{1,y}$	0.26	1.00		
$m_{1,z}$	-0.11	-0.00	1.00	
$m_{2,x}$	-0.12	-0.02	-0.12	1.
$m_{2,y}$	0.05	0.16	-0.05	0.
$m_{2,z}$	0.07	0.00	-0.23	-0
$E_{\rm vis}$	0.20	0.30	0.28	-0
$E_{\rm inc}$	0.19	0.28	0.38	-0
$n_{ m hit}$	0.22	0.31	0.24	-0
$E_1/E_{\rm vis}$	0.11	-0.03	-0.95	0.
$E_2/E_{\rm vis}$	-0.04	0.10	0.45	0.
$E_3/E_{\rm vis}$	-0.11	-0.05	0.92	-0
	$m_{1,x}$	$m_{1,y}$	$m_{1,z}$	m_{o}

Geant4



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BIB-AE PP







Correlations

Geant4 - GAN



	$m_{1,x}$	$m_{1, y}$	$m_{1,z}$	200
$E_3/E_{\rm vis}$	0.18	0.13	0.00	-0
$E_2/E_{\rm vis}$	0.17	0.25	-0.01	0
$E_1/E_{\rm vis}$	-0.28	-0.30	0.00	0
$n_{ m hit}$	0.25	0.19	-0.14	0
$E_{\rm inc}$	0.25	0.20	-0.05	0
$E_{\rm vis}$	0.23	0.19	-0.06	0
$m_{2,z}$	-0.02	-0.15	-0.33	-0
$m_{2,y}$	-0.32	-0.56	-0.17	-0
$m_{2,x}$	-0.26	-0.47	-0.44	0
$m_{1,z}$	0.25	0.24	0.00	
$m_{1,y}$	-0.33	0.00		
$m_{1,x}$	0.00			

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Geant4 - WGAN



Geant4 - BIB-AE PP

	$m_{1,x}$	$m_{1,y}$	$m_{1,z}$	$m_{2,x}$	$m_{2,y}$	$m_{2,z}$	$E_{ m vis}$	$E_{ m inc}$	$n_{ m hit}$	$E_1/E_{ m vis}$	$E_2/E_{ m vis}$	F_{0}/F_{min}
$E_3/E_{\rm vis}$	0.04	0.03	0.01	-0.21	-0.21	-0.26	-0.14	-0.10	-0.16	0.02	-0.32	0.0
$E_2/E_{\rm vis}$	-0.09	-0.09	-0.33	-0.08	-0.03	0.37	0.11	0.07	0.12	0.23	0.00	
$E_1/E_{\rm vis}$	0.05	0.04	0.02	0.27	0.23	0.00	-0.01	-0.04	0.02	0.00		
$n_{ m hit}$	-0.28	-0.29	-0.09	0.21	0.15	-0.07	-0.00	-0.01	0.00			
$E_{\rm inc}$	-0.26	-0.26	-0.03	0.09	0.03	-0.17	-0.01	0.00				
$E_{\rm vis}$	-0.27	-0.28	-0.07	0.13	0.07	-0.15	0.00					
$m_{2,z}$	-0.01	0.01	-0.12	0.26	0.20	0.00						
$m_{2,y}$	-0.02	-0.17	-0.24	-0.16	0.00							
$m_{2,x}$	0.13	0.03	-0.26	0.00								
$m_{1,z}$	-0.01	-0.01	0.00									
$m_{1,y}$	-0.26	0.00										
$m_{1,x}$	0.00											

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Computation Time

- For 10-100 GeV showers:
 - 3 Orders of magnitude speedup compared to GEANT4

Simulator	Hardware	Batchsize	Time/shower	Speedup
GEANT4	CPU	N/A	4082 ± 170 ms	_
BIB-AE	CPU	1	426.3 ± 3.6 ms	x10
BIB-AE	GPU V100	1	3.19 ± 0.01 ms	x1279
BIB-AE	GPU V100	100	1.42 ± 0.01 ms	x2874

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