

Neutrino Self-interaction as a solution(?) to the Hubble Tension

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Based on 2011.12315
with Subhajit Ghosh



Outline

- **Cosmology of self-interacting neutrino**
- **Flavor-universal S_{Inu} & the Hubble tension**
- **Laboratory constraints on S_{Inu}**
- **Flavor-specific S_{Inu}**

A phenomenological model

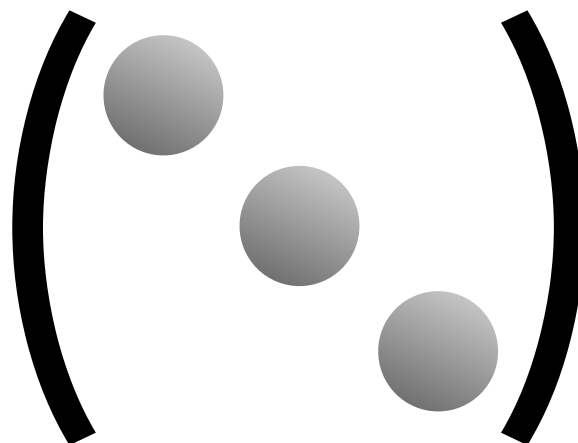
$$\mathcal{L} \supset \frac{1}{2} g_{ij} \bar{\nu}_i \nu_j \phi, \quad g_{ij} = g \delta_{ij}$$



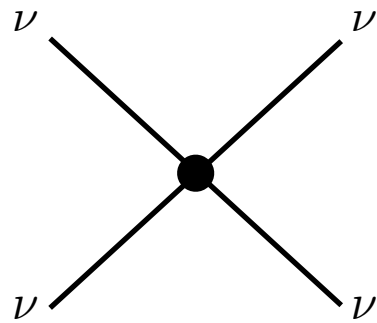
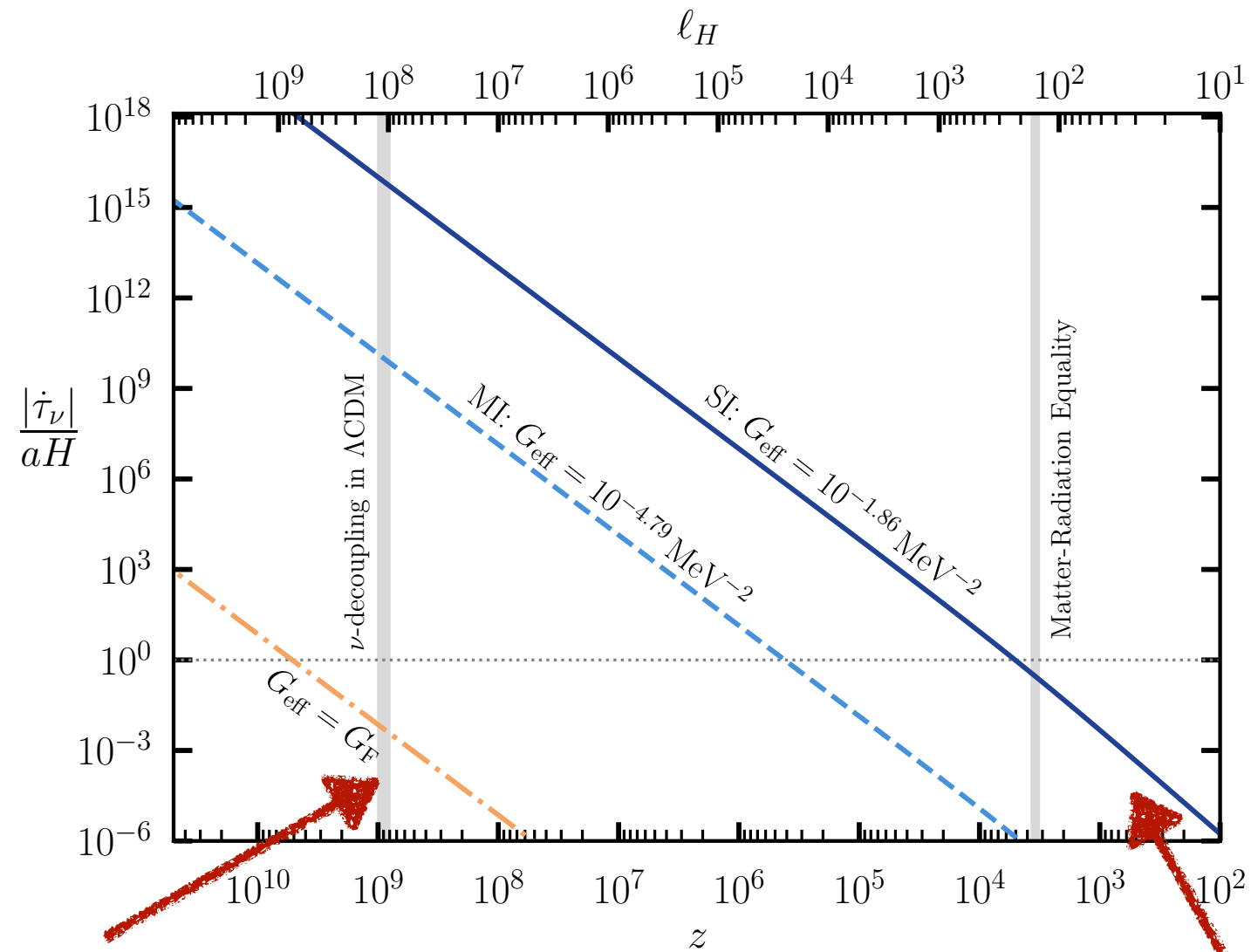
$$\mathcal{L} \supset G_{\text{eff}} \bar{\nu} \nu \bar{\nu} \nu, \quad G_{\text{eff}} = \frac{g^2}{M_\phi^2}$$

1306.1536
1704.06657
1902.00534

All neutrino flavors interact with the same strength

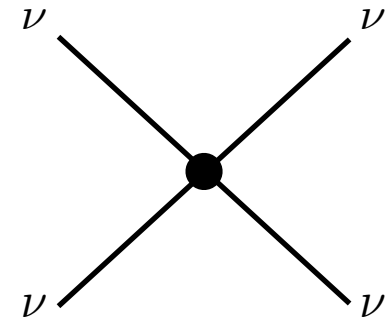


Cosmology of Slnu



$$\sigma \propto G_F^2$$

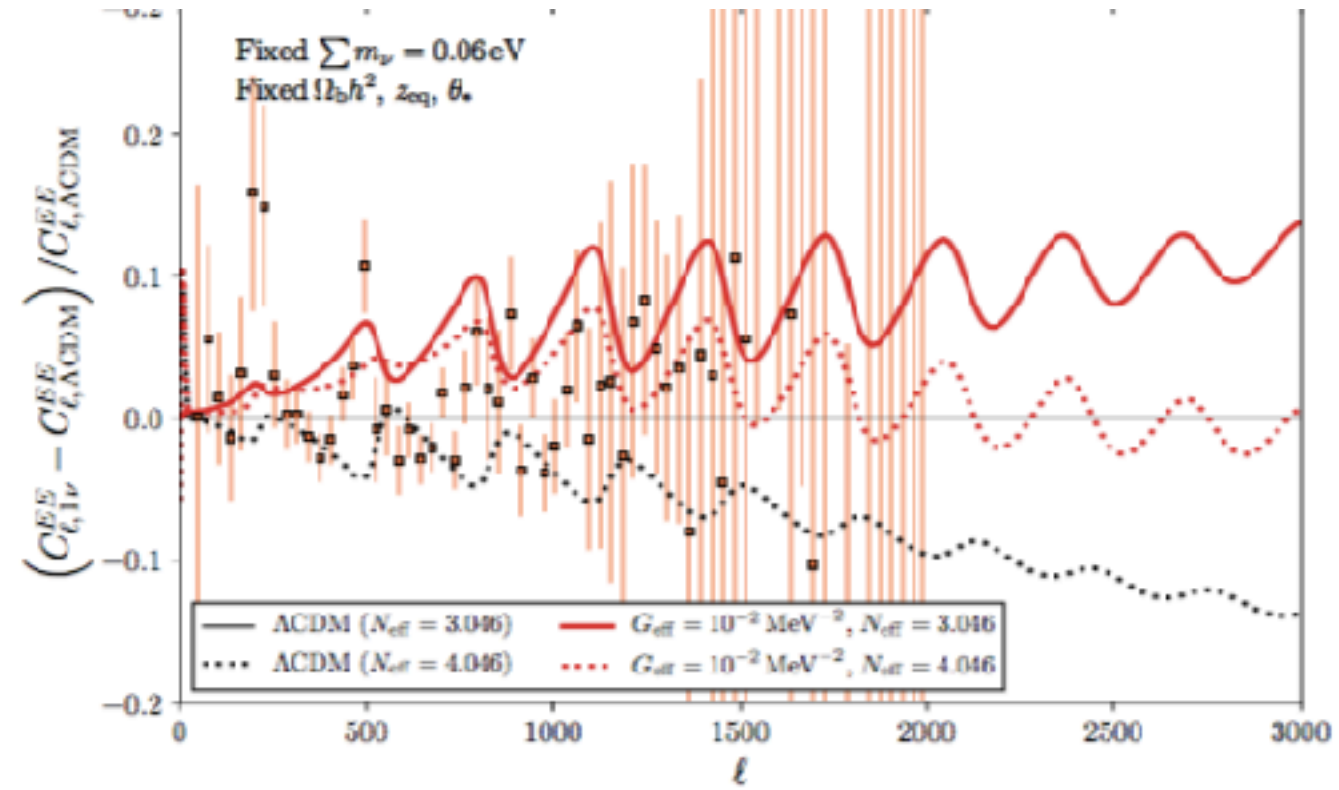
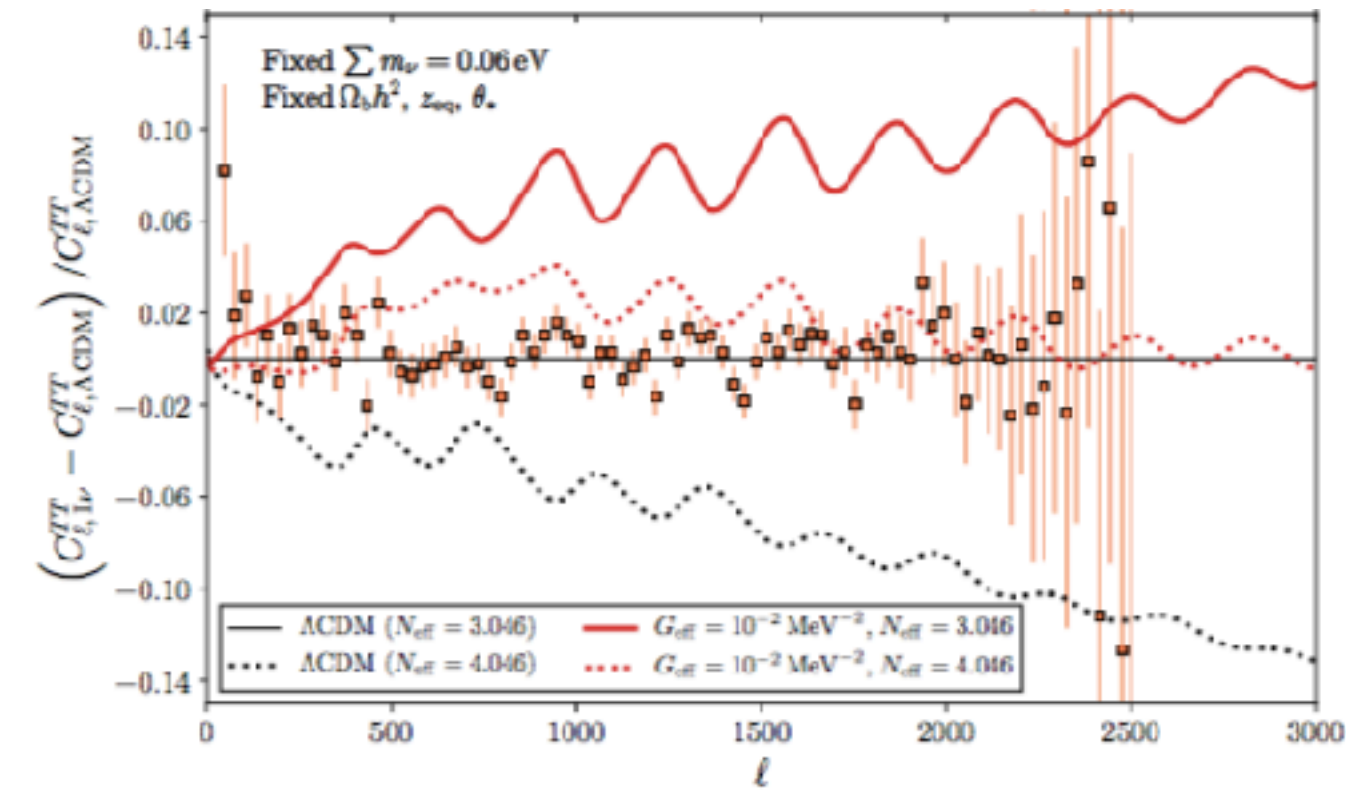
$$G_{\text{eff}} \simeq 10^9 G_F$$



$$\sigma \propto \frac{g_\phi^4}{M_\phi^4} \equiv G_{\text{eff}}^2$$

CMB Power Spectra

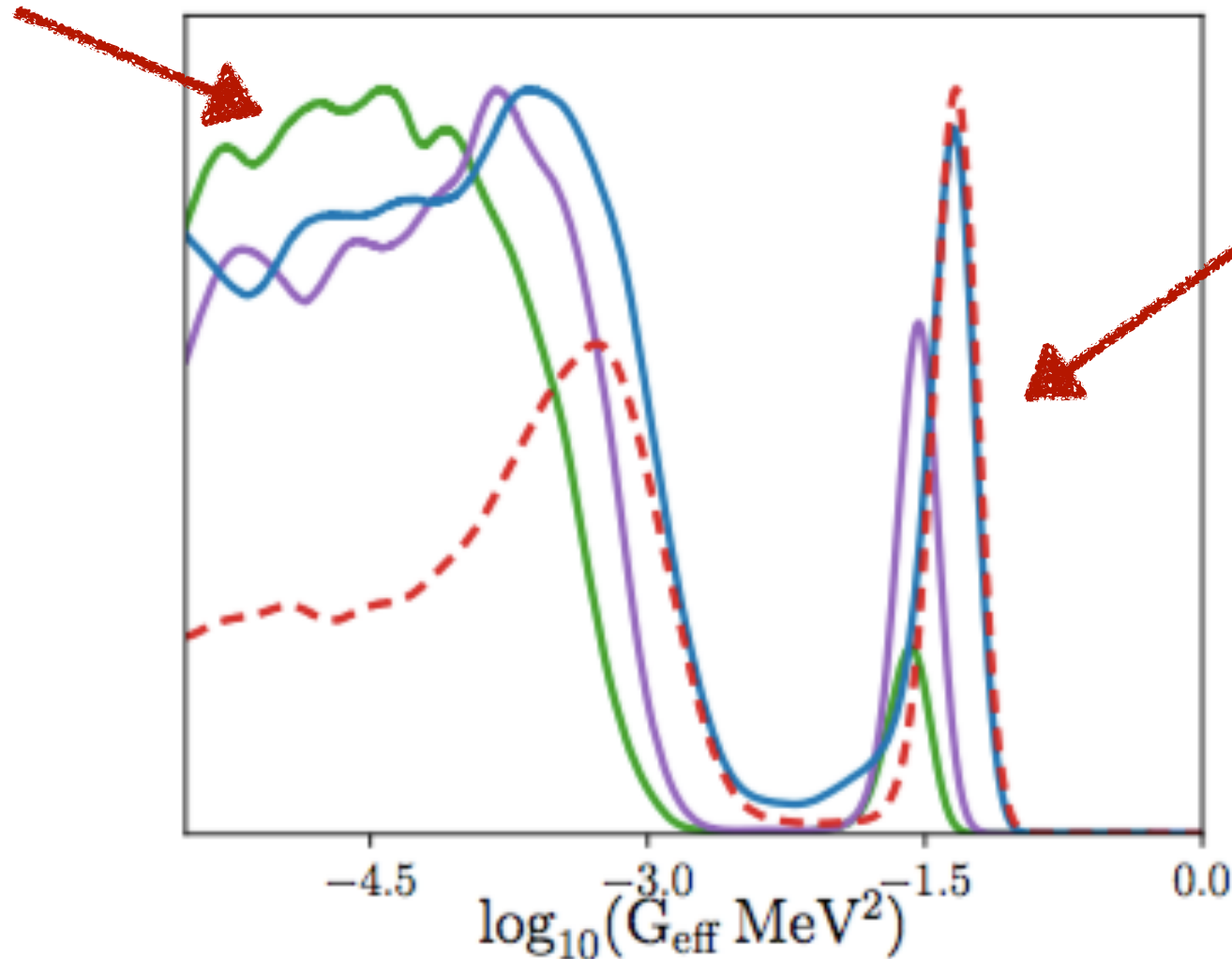
Gravitational potentials enhance photon perturbation \Rightarrow larger C_l



Kreisch et al 1902.00534

Cosmology of Slnu

Moderately Interacting
(MI)



Strongly Interacting
(SI)

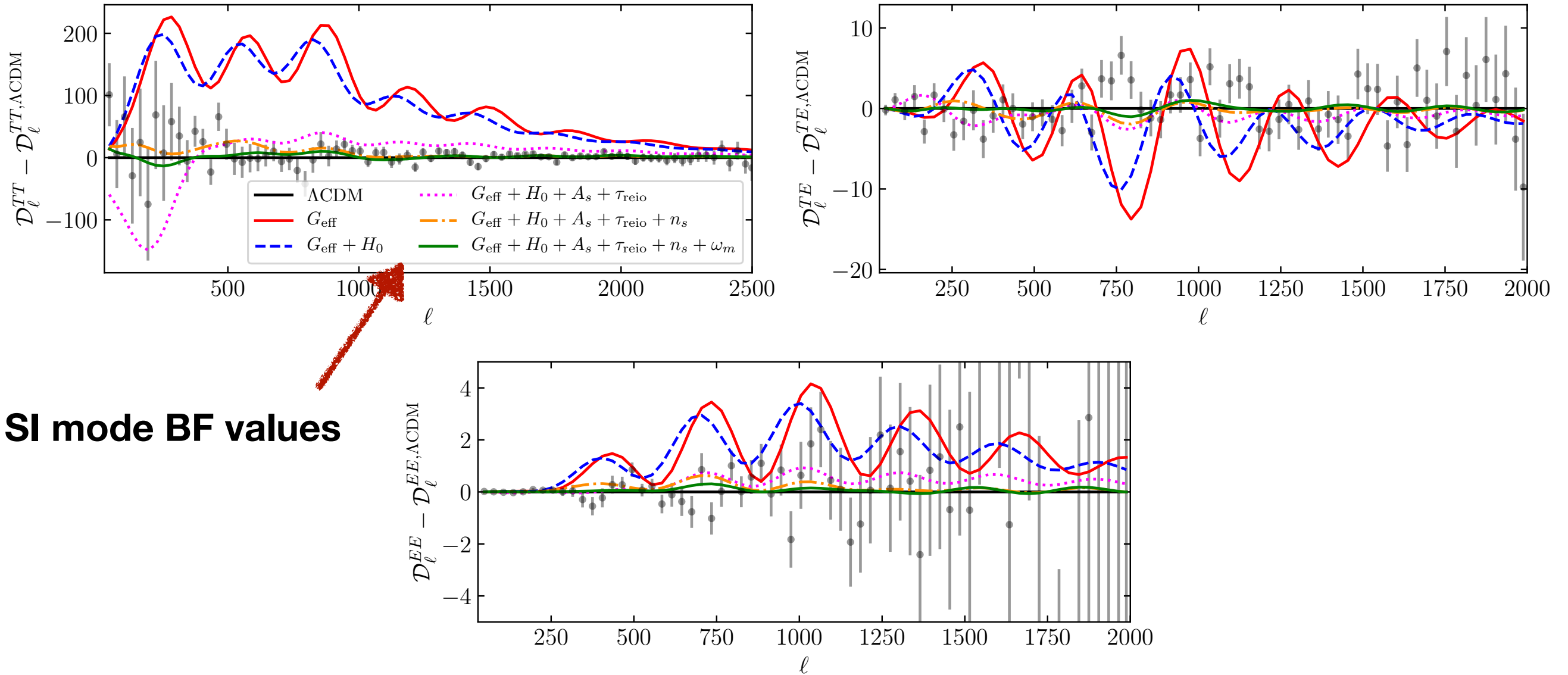
Kreisch et al 1902.00534

Questions:

- Origin of the SI mode?
- Why is there a valley in between?

Cosmology of Slnu: G_{eff} degeneracy w/ other params

P18 : TTTEEE + lowE | **3c + 0f**



SI mode BF values

SI mode coupling affects $l > 200$

Slnu & Hubble

Phase shift $\phi_\nu \simeq 0.19\pi R_\nu$ Free-streaming neutrino

$$\ell \approx (m\pi - \phi_\nu) \frac{D_A^*}{r_s^*}$$

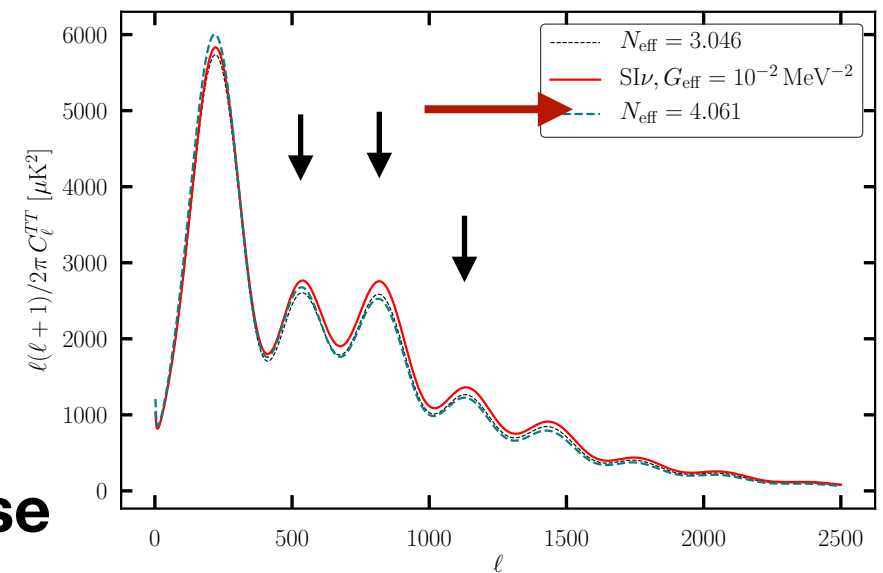
Slnu phase shift increases ℓ



To keep the spectrum fixed, $\theta_* = \frac{r_s^*}{D_A^*}$ needs to increase



Larger H_0 decreases D_A^* & increases θ_*



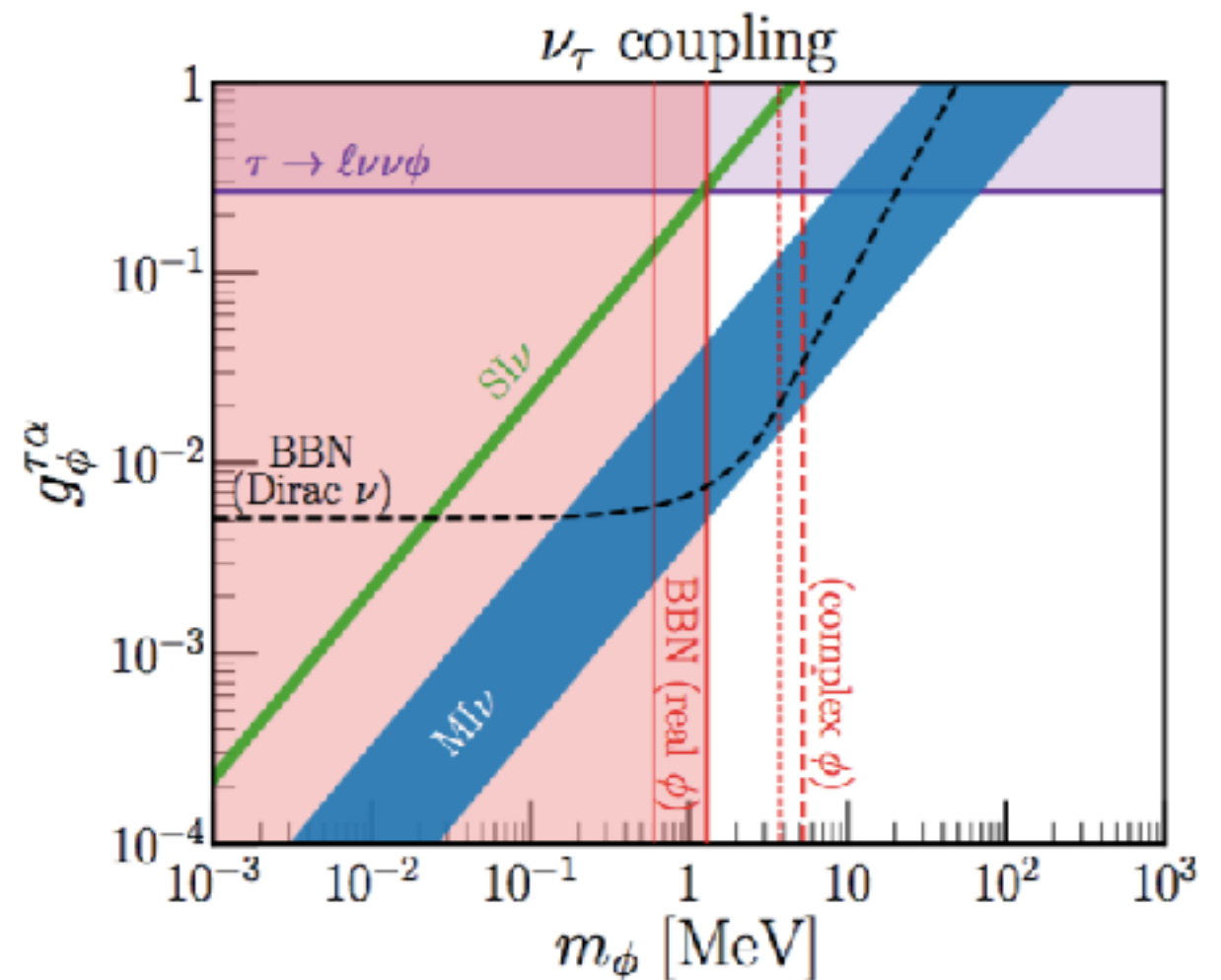
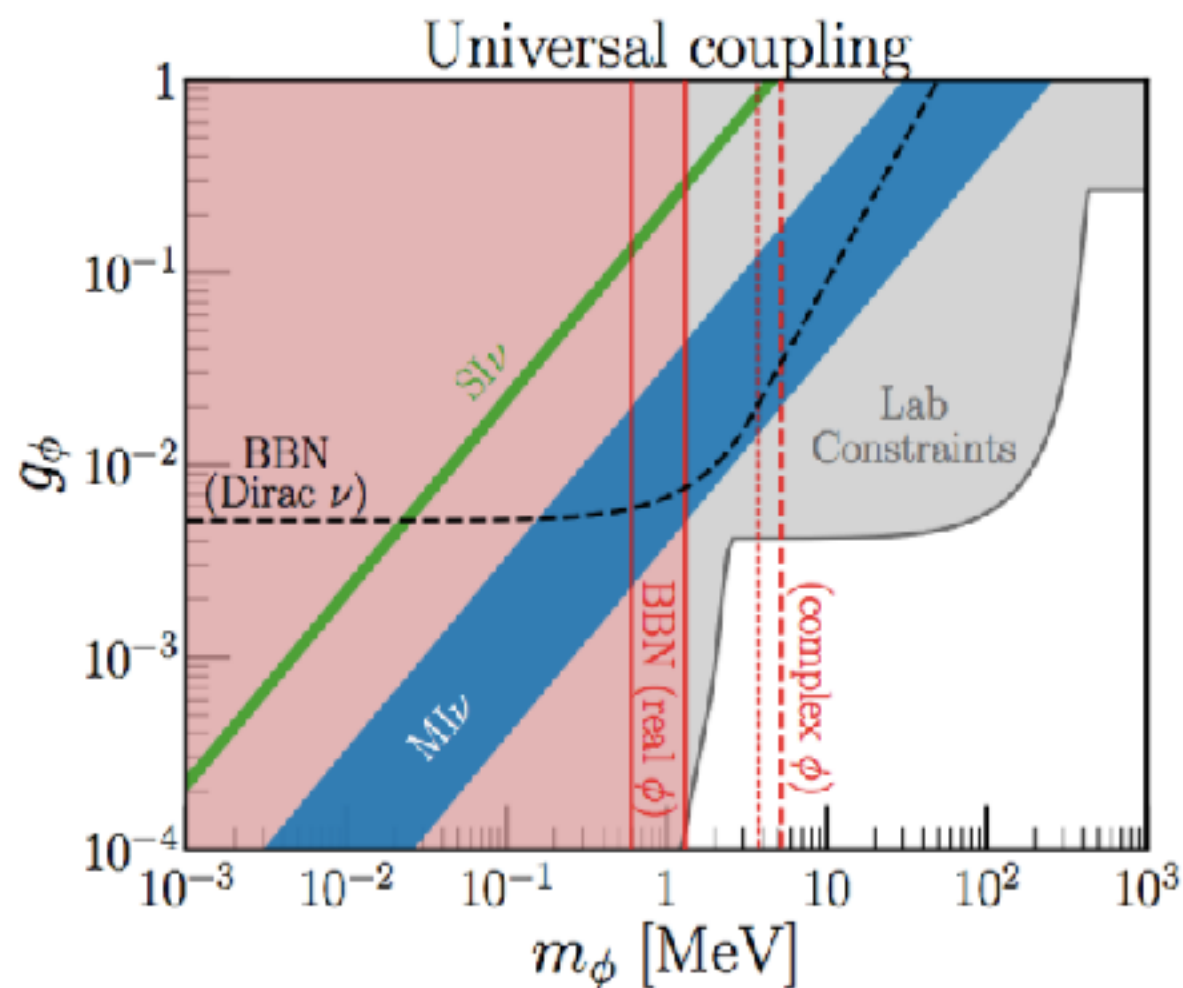
$$D_A^* = \int_0^{z^*} \frac{1}{H(z)} dz, \quad r_s^* = \int_{z^*}^{\infty} \frac{c_s(z)}{H(z)} dz$$

Laboratory Constraints

Example: $\nu\nu$ self-interaction could come from a scalar (1905.02727)

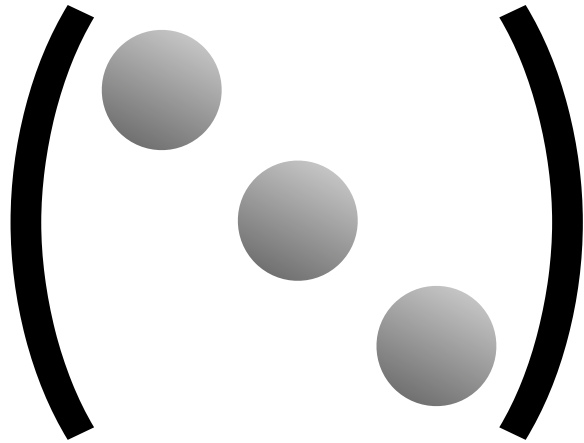
$$\mathcal{L} \supset -\frac{1}{2}m_\phi^2\phi^2 + \frac{1}{2}(g_\phi^{\alpha\beta}\nu_\alpha\nu_\beta\phi + \text{h.c.})$$

$$G_{\text{eff}} \equiv \frac{g_\phi^2}{m_\phi^2} = (10 \text{ MeV})^{-2} \left(\frac{g_\phi}{10^{-1}}\right)^2 \left(\frac{\text{MeV}}{m_\phi}\right)^2$$

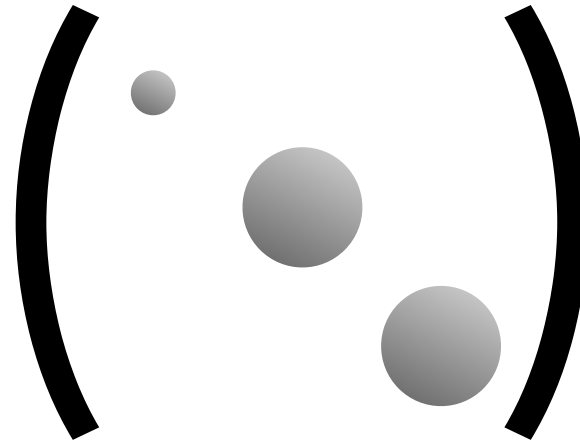


Flavor-specific Slnu

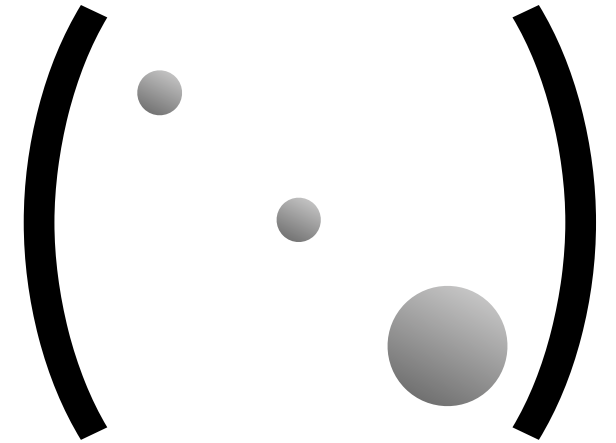
Flavor
Space



3c+0f



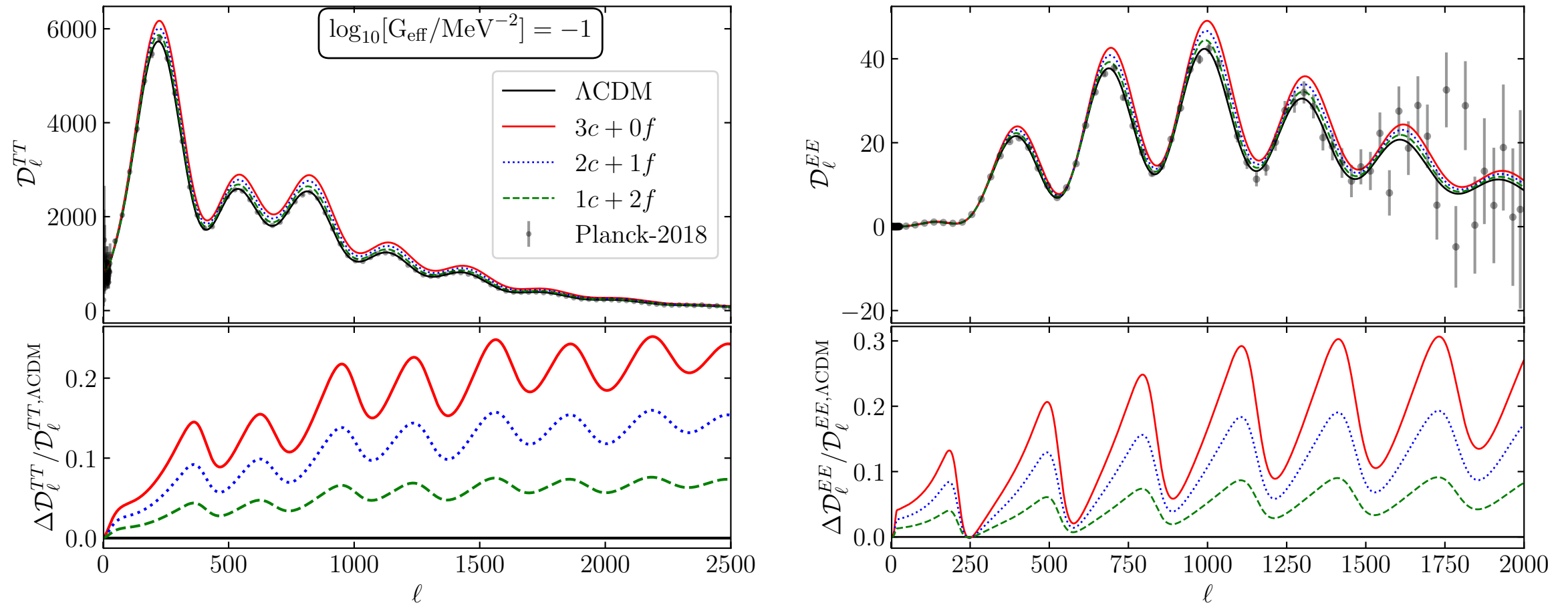
2c+1f



1c+2f

$$\mathcal{L} \supset G_{\text{eff}}^{(ijkl)} \bar{\nu}_i \nu_j \bar{\nu}_k \nu_l, \quad G_{\text{eff}}^{(ijkl)} \equiv \frac{g_{ij} g_{kl}}{M_\phi^2}$$

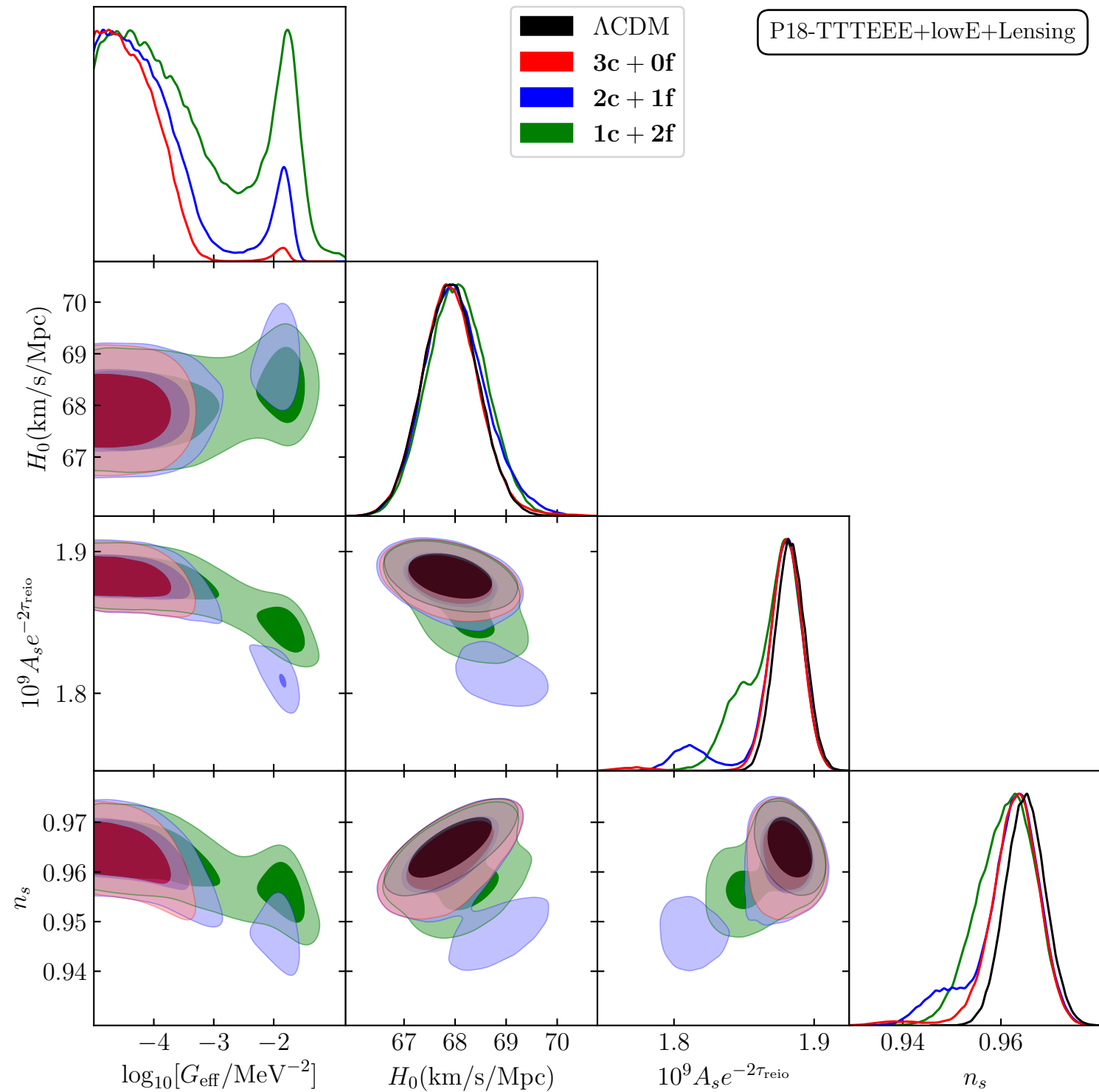
Flavor-specific Slnu



2011.12315

**Less number of interacting neutrinos reduces the change
in the CMB power spectra**

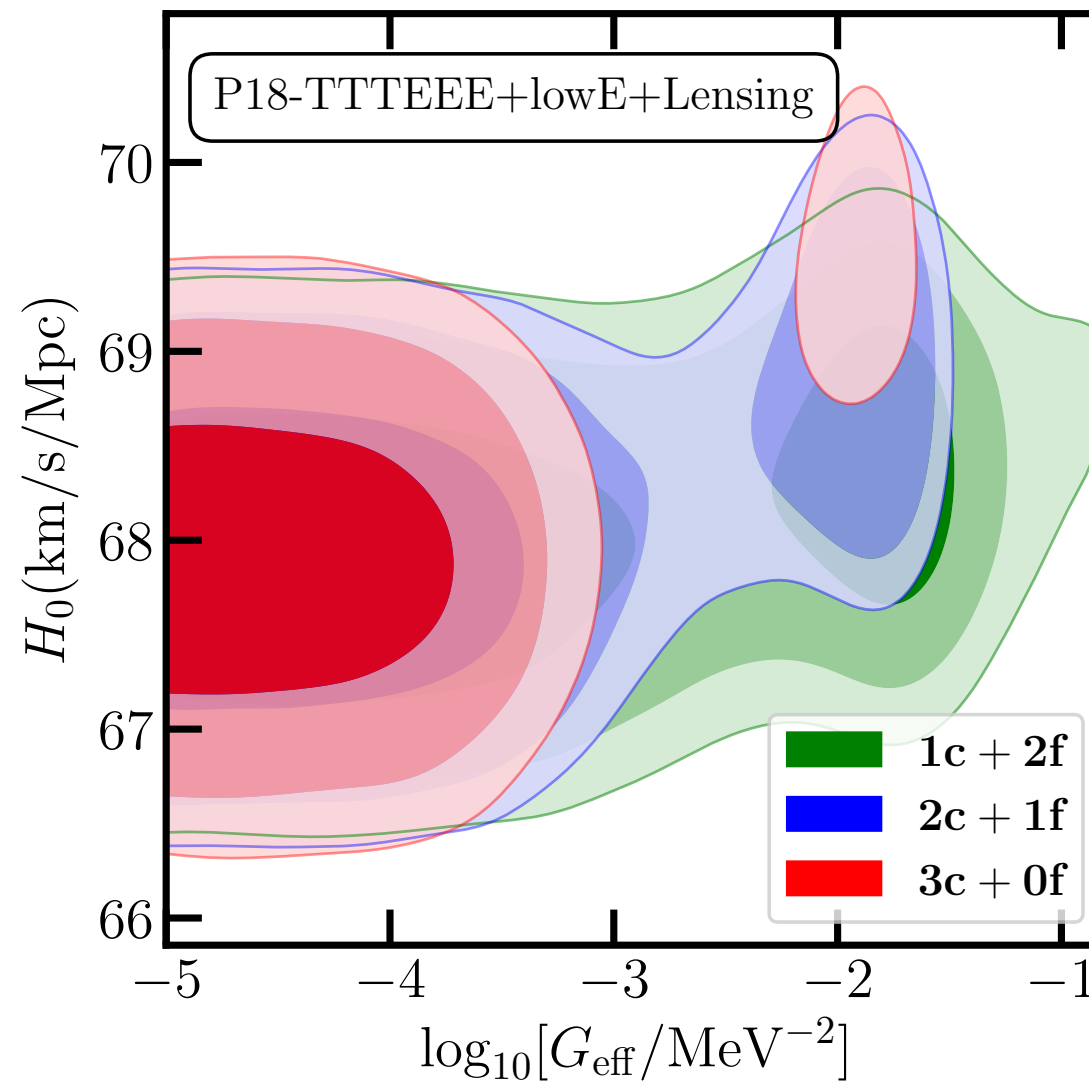
Flavor-specific Slnu



2011.12315

Flavor-specific S_{nu} : effect on H_0

G_{eff} has positive correlation with H_0



$H_0 = 69.5 \pm 0.6$ km/s/Mpc for TTTEEE+lowE+lens (2011.12315)

$H_0 = 69.6 \pm 2.5$ (TRGB, 2002.01550)

$H_0 = 73.2 \pm 1.3$ (Riess et al, 2012.08534)

Conclusions & Future Outlook

- Neutrino self-interaction is phenomenologically motivated
- Flavor-universal large $G_{\text{eff}} (\sim 10^{-2} \text{ MeV}^{-2})$ is incompatible with BBN & laboratory data. However, flavor-specific interaction may still be allowed.
- Changes in CMB spectrum is less with less #interacting nu, hence, more room to exploit the deg. with $A_s, n_s, \tau_{\text{reio}}, H_0 \implies$ **enhances the SI mode**
- The SI mode bestfit coupling in the flavor-specific case remains almost the same
- More work needed with massive neutrinos and off-diagonal couplings

Parameter values

Table 4: Parameter values and 68% confidence limits in 2c + 1f.

Parameters	TT+lowE		TTTEEE+lowE	
	SI	MI	SI	MI
$\Omega_b h^2$	0.022 ± 0.00027	0.022 ± 0.00021	0.022 ± 0.00016	0.022 ± 0.00015
$\Omega_c h^2$	0.1211 ± 0.0023	0.1203 ± 0.002	0.1205 ± 0.0014	0.1201 ± 0.0013
$100\theta_s$	1.0452 ± 0.00059	1.0419 ± 0.0005	1.045 ± 0.00076	1.0419 ± 0.00031
$\ln(10^{10} A_s)$	2.99 ± 0.0179	3.036 ± 0.01714	3 ± 0.0167	3.042 ± 0.0161
n_s	0.9407 ± 0.0079	0.9596 ± 0.0068	0.9473 ± 0.0046	0.9628 ± 0.005
τ_{reio}	0.0501 ± 0.008	0.0516 ± 0.0079	0.0538 ± 0.0077	0.0538 ± 0.0077
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	-1.69 ± 0.2	-4.03 ± 0.6	-1.93 ± 0.24	-4.24 ± 0.5
$H_0(\text{ km s}^{-1}\text{Mpc}^{-1})$	68.34 ± 1.00	67.57 ± 0.92	68.81 ± 0.63	67.83 ± 0.6
σ_8	0.823 ± 0.01	0.824 ± 0.009	0.829 ± 0.0079	0.824 ± 0.0075

Table 5: Parameter values and 68% confidence limits in 1c + 2f.

Parameters	TT+lowE		TTTEEE+lowE	
	SI	MI	SI	MI
$\Omega_b h^2$	0.022 ± 0.00023	0.022 ± 0.00021	0.022 ± 0.00015	0.022 ± 0.00015
$\Omega_c h^2$	0.1207 ± 0.0021	0.1203 ± 0.002	0.1203 ± 0.0014	0.1201 ± 0.0013
$100\theta_s$	1.0434 ± 0.00062	1.0419 ± 0.0004	1.043 ± 0.00058	1.0419 ± 0.0003
$\ln(10^{10} A_s)$	3.01 ± 0.0179	3.037 ± 0.01664	3.024 ± 0.0166	3.042 ± 0.016
n_s	0.9513 ± 0.0069	0.9609 ± 0.0059	0.9553 ± 0.0049	0.963 ± 0.005
τ_{reio}	0.051 ± 0.008	0.0519 ± 0.008	0.0539 ± 0.0076	0.0539 ± 0.0077
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	-1.75 ± 0.4	-3.94 ± 0.6	-1.9 ± 0.37	-4.06 ± 0.6
$H_0(\text{ km s}^{-1}\text{Mpc}^{-1})$	67.9 ± 1.00	67.56 ± 0.93	68.3 ± 0.62	67.83 ± 0.61
σ_8	0.821 ± 0.01	0.823 ± 0.009	0.825 ± 0.0083	0.824 ± 0.0075

Parameter values

Table 11: Parameter values and 68% confidence limits in **2c + 1f**.

Parameters	TTTEEE+lowE+lens		TTTEEE+lowE+lens+BAO+ H_0	
	SI	MI	SI	MI
$\Omega_b h^2$	0.022 ± 0.0001	0.022 ± 0.00014	0.022 ± 0.0001	0.022 ± 0.00013
$\Omega_c h^2$	0.1202 ± 0.0013	0.1199 ± 0.0012	0.12 ± 0.001	0.1188 ± 0.0009
$100\theta_s$	1.045 ± 0.0008	1.0419 ± 0.0003	1.045 ± 0.00068	1.042 ± 0.00029
$\ln(10^{10} A_s)$	3 ± 0.0158	3.041 ± 0.014	3 ± 0.0151	3.044 ± 0.0145
n_s	0.9476 ± 0.0043	0.9629 ± 0.0048	0.9483 ± 0.004	0.966 ± 0.0046
τ_{reio}	0.0541 ± 0.0074	0.0536 ± 0.0072	0.0544 ± 0.007	0.0565 ± 0.0071
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	-1.96 ± 0.26	-4.22 ± 0.51	-1.91 ± 0.22	-4.22 ± 0.52
$H_0(\text{km s}^{-1}\text{Mpc}^{-1})$	68.87 ± 0.58	67.9 ± 0.53	69.08 ± 0.42	68.47 ± 0.4
σ_8	0.829 ± 0.007	0.823 ± 0.006	0.827 ± 0.0065	0.821 ± 0.0059

Table 12: Parameter values and 68% confidence limits in **1c + 2f**.

Parameters	TTTEEE+lowE+lens		TTTEEE+lowE+lens+BAO+ H_0	
	SI	MI	SI	MI
$\Omega_b h^2$	0.022 ± 0.0001	0.022 ± 0.00014	0.022 ± 0.0001	0.022 ± 0.00013
$\Omega_c h^2$	0.1201 ± 0.0012	0.1199 ± 0.0012	0.12 ± 0.0009	0.1188 ± 0.0009
$100\theta_s$	1.043 ± 0.0006	1.0419 ± 0.0003	1.043 ± 0.00056	1.042 ± 0.00029
$\ln(10^{10} A_s)$	3.023 ± 0.0153	3.041 ± 0.015	3 ± 0.0151	3.045 ± 0.0142
n_s	0.9555 ± 0.0046	0.9633 ± 0.0045	0.9572 ± 0.004	0.966 ± 0.0042
τ_{reio}	0.0536 ± 0.0073	0.0536 ± 0.0074	0.0554 ± 0.007	0.0566 ± 0.0071
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	-1.91 ± 0.37	-4.04 ± 0.61	-1.86 ± 0.36	-4.03 ± 0.61
$H_0(\text{km s}^{-1}\text{Mpc}^{-1})$	68.35 ± 0.56	67.9 ± 0.54	68.75 ± 0.41	68.48 ± 0.41
σ_8	0.824 ± 0.007	0.823 ± 0.006	0.822 ± 0.0071	0.821 ± 0.006