Probing the Earth's Core using Atmospheric Neutrinos at INO

Sanjib Kumar Agarwalla sanjib@iopb.res.in https://twitter.com/sanjibneutrino

Institute of Physics (IOP), Bhubaneswar, India



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Neutrino Tomography of the Earth

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Validating the Earth's Core using Atmospheric Neutrinos with ICAL at INO

Anil Kumar,^{*a,b,c*} Sanjib Kumar Agarwalla^{*a,c,d*}

^a Institute of Physics, Sachivalaya Marg, Sainik School Post, Bhubaneswar 751005, India
 ^b Applied Nuclear Physics Division, Saha Institute of Nuclear Physics, Block AF, Sector 1, Bidhannagar, Kolkata 700064, India
 ^c Homi Bhabha National Institute, Anushakti Nagar, Mumbai 400094, India
 ^d International Centre for Theoretical Physics, Strada Costiera 11, 34151 Trieste, Italy
 E-mail: anil.k@iopb.res.in (ORCID: 0000-0002-8367-8401),
 sanjib@iopb.res.in (ORCID: 0000-0002-9714-8866)

Magnetized Iron Calorimeter Detector (ICAL) at INO

- ICAL@INO: 50 kton magnetized iron calorimeter detector at the proposed India-based Neutrino Observatory (INO)
- Location: Bodi West Hills, Theni District, Tamil Nadu, India
- Aim: To determine mass ordering and precision measurement of atmospheric oscillation parameters.
- Source: Atmospheric neutrinos and antineutrinos in the multi-GeV range of energies over a wide range of baselines.
- Uniqueness: Charge identification capability helps to distinguish μ^- and μ^+ and hence, ν_{μ} and $\bar{\nu}_{\mu}$
- Muon energy range: 1 25 GeV, Muon energy resolution: $\sim 10\%$
- Baselines: 15 12000 km, Muon zenith angle resolution: $\sim 1^\circ$



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Neutrino Oscillations in Matter: MSW Effect

 ν_e Neutrino propagation through matter modify the oscillations significantly Coherent forward elastic scattering of neutrinos with matter particles W^{\pm} Charged current interaction of v_e with electrons creates an extra potential for v_e ν_e $A(eV^2) = 0.76 \times 10^{-4} \rho \ (g/cc) E(GeV)$ $A = \pm 2\sqrt{2}G_F N_e E$ MSW matter term: or N_e = electron number density, + (-) for neutrinos (anti-neutrinos), ρ = matter density in Earth Matter term changes sign when we switch from neutrino mode to antineutrino mode even if $\delta_{CP} = 0$, causes fake CP asymmetry $P(\nu_{\alpha} \to \nu_{\beta}) - P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) \neq 0$ Matter term modifies oscillation probability differently depending on the sign of Δm^2 $\Delta m^2 \simeq A$ $= 6 - 8 \,\mathrm{GeV}$ \Leftrightarrow Resonant conversion – Matter effect V $\bar{\nu}$ **Resonance occurs for neutrinos (anti-neutrinos)** $\Delta m^2 > 0$ MSW if Δm^2 is positive (negative) $\Delta m^2 < 0$ MSW



Three-layered model of Earth

Crust: solid, rocks, brittle, lowest density

Mantle: hot, solid upper mantle, viscous plastic lower mantle

Core: solid inner core, liquid outer core, iron and nickel

Region	R_{\min} (km)	R_{\max} (km)	Density (g/cm^3)
Core	0	3480	11.37
Mantle	3480	5701	5
Crust	5701	6371	3.3

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D. E. Loper and T. Lay, The core-mantle boundary region, Journal of Geophysical Research: Solid Earth 100 (1995), no. B4 6397–6420.
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ICAL cannot distinguish between profiles having 81 and 25 layers



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Core-mantle	(0, 3480, 6371)	(11.37, 4.42)

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Profiles	Layer boundaries (km)	Layer densities (g/cm ³)
Uniform	(0, 6371)	(5.55)

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Profiles	Layer boundaries (km)	Layer densities (g/cm^3)
PREM	25 layers	25 densities
Core-mantle-crust	(0, 3480, 5701, 6371)	(11.37, 5, 3.3)
Mantle-crust	(0, 5701, 6371)	(6.45, 3.3)
Core-mantle	(0, 3480, 6371)	(11.37, 4.42)
Uniform	(0, 6371)	(5.55)

While constructing alternative profiles of Earth, the radius and mass of Earth remain invariant

Effect of Various Density Profiles on Muon *v*Survival Probability



MSW resonance: red yellow patch (- $0.8 < \cos\theta < -0.5$ and 6 GeV $< E_{\nu} < 10$ GeV)

Neutrino Oscillation Length/Parametric resonance: red and yellow patches ($\cos\theta < -0.8$ and 3 GeV $< E_{\nu} < 6$ GeV

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Effect of Various Density Profiles on Muon v Appearance Probability

$\sin^2 2\theta_{12}$	$\sin^2 \theta_{23}$	$\sin^2 2\theta_{13}$	$\Delta m_{\rm eff}^2 ~({\rm eV^2})$	$\Delta m^2_{21} \ ({\rm eV^2})$	δ_{CP}	Mass Ordering
0.855	0.5	0.0875	2.49×10^{-3}	$7.4 imes 10^{-5}$	0	Normal (NO)



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Reconstructed Muon Events Passing Through Different Layers of Earth



Neutrino flux (Honda) at INO location

500 kt•yr ICAL (magnetized) exposure

NUANCE event generator

Migration matrices using GEANT4

3-flavor **v** oscillations in matter with PREM profile

Reconstructed muon events

Regions	$\cos \theta_{\nu}$	L_{ν} (km)	μ^- Events	μ^+ Events
Crust-mantle-core	(-1.00, -0.84)	(10691, 12757)	331	146
Crust-mantle	(-0.84, -0.45)	(5721, 10691)	739	339
Crust	(-0.45, 0.00)	(437, 5721)	550	244
Downward	(0.00, 1.00)	(15, 437)	2994	1324
Total	(-1.00, 1.00)	(15, 12757)	4614	2053

Reconstructed Muon Event Distributions Through Various Earth Layers



Reconstructed muon event distributions at ICAL for 500 kt•yr exposure for neutrinos passing through various regions in the Earth

Binning Scheme and Systematic Uncertainties

Observable	Range	Bin width	Number	of bins
	[1, 4]	0.5	6	
Frec (CoV)	[4, 7]	1	3	12
L_{μ} (GeV)	[7, 11]	4	1	(12
	[11, 21]	5	2	
	[-1.0, -0.4]	0.05	12	
$\cos \theta_{\mu}^{ m rec}$	[-0.4, 0.0]	0.1	4	21
P	$[0.0, \ 1.0]$	0.2	5	
	[0, 2]	1	2	
$E'_{\rm had}^{\rm rec} ({\rm GeV})$	[2, 4]	2	1	4
	[4, 25]	21	1	

- 20% flux normalization error
- 10% cross section error
- 5% energy dependent tilt error in flux
- 5% error in zenith angle dependence of flux
- 5% overall systematics

Statistical Analysis

In this analysis, the χ^2 statistics is expected to give median sensitivity of the experiment in the frequentist approach.

$$\chi_{-}^{2} = \min_{\xi_{l}} \sum_{i=1}^{N_{E_{\mu}^{rec}}} \sum_{j=1}^{N_{exp}} \sum_{k=1}^{N_{exp}} \left[2(N_{ijk}^{\text{theory}} - N_{ijk}^{\text{data}}) - 2N_{ijk}^{\text{data}} \ln\left(\frac{N_{ijk}^{\text{theory}}}{N_{ijk}^{\text{data}}}\right) \right] + \sum_{l=1}^{5} \xi_{l}^{2}$$

where,

$$\mathcal{N}^{ ext{theory}}_{ijk} = \mathcal{N}^0_{ijk} \left(1 + \sum_{l=1}^5 \pi^l_{ijk} \xi_l
ight)$$

Similarly, χ^2_+ is defined for μ^+

$$\chi^2_{\rm ICAL} = \chi^2_- + \chi^2_+$$

 $\Delta \chi^2_{\text{ICAL-profile}} = \chi^2_{\text{ICAL}} \text{ (Mantle-Crust)} - \chi^2_{\text{ICAL}} \text{ (Core-Mantle-Crust)}$

• Marginalization over systematic uncertainties and $\sin^2 \theta_{23}$: (0.36, 0.66), Δm_{eff}^2 : (2.1, 2.6) ×10⁻³ eV², and mass ordering: (NO, IO)

		$\Delta \chi^2_{\rm ICAL-profile}$				
MC Data	Theory	NO(t	NO(true)		rue)	
		with CID	w/o CID	with CID	w/o CID	
Core-mantle-crust	Vacuum	4.65	2.96	3.53	1.43	
Core-mantle-crust	Mantle-crust	6.31	3.19	3.92	1.29	
Core-mantle-crust	Core-mantle	0.73	0.47	0.59	0.21	
Core-mantle-crust	Uniform	4.81	2.38	3.12	0.91	
PREM profile	Core-mantle-crust	0.36	0.24	0.30	0.11	
PREM profile	Vacuum	5.52	3.52	4.09	1.67	
PREM profile	Mantle-crust	7.45	3.76	4.83	1.59	
PREM profile	Core-mantle	0.27	0.18	0.21	0.07	
PREM profile	Uniform	6.10	3.08	3.92	1.18	

Impact of Marginalization Over Various Oscillation Parameters

			$\Delta \chi_{ m I}^2$	2 CAL-profile		
MC Data	Theory	Fixed	Marginalization over			
		parameter	$\sin^2 \theta_{23}$	$ \Delta m_{\rm eff}^2 $	$\pm \Delta m_{\rm eff}^2 $	All
Core-mantle-crust	Mantle-crust	6.90	6.36	6.84	6.84	6.31
Core-mantle-crust	Vacuum	6.80	6.44	5.16	4.94	4.65
PREM	Mantle-crust	7.88	7.47	7.81	7.81	7.45
PREM	Vacuum	7.71	7.28	6.10	5.89	5.52

Impact of Different True Choices of $\sin^2 \theta_{23}$



We explore the possibility of utilizing neutrino oscillations in the presence of matter to extract information about the internal structure of Earth complementary to seismic studies

Using good directional resolution, ICAL observe 331 μ ⁻ and 146 μ ⁺ core-passing events with 500 kt•yr exposure

The presence of Earth's core can be independently confirmed at ICAL with a median $\Delta \chi^2$ of 7.45 (4.83) assuming normal (inverted) mass ordering

Thank you!

Backup Slides

Effective Regions in $(E_{\mu}^{\text{rec}}, \cos \theta_{\mu}^{\text{rec}})$ Plane to Validate Earth's Core • MC Data: Core-mantle-crust • Theory: Mantle-crust • 500 kt-yr exposure at ICAL •

Systematic uncertainties are marginalized whereas oscillation parameters are kept fixed in theory



	Fixed-parameter $\Delta \chi^2$		
	NO	10	
Contribution from μ^-	6.85	0.02	
Contribution from μ^+	0.05	4.08	
Total	6.90	4.10	

Note: $\Delta \chi_{-}^2$ and $\Delta \chi_{+}^2$ are calculated without pull penalty $\sum_{l=1}^{5} \xi_l^2$ to explore contributions from each bin in $(E_{\mu}^{\text{rec}}, \cos \theta_{\mu}^{\text{rec}})$ plane for μ^- and μ^+ events, respectively.

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Atmospheric Neutrinos: Neutrinos from Cosmic Rays



$$\pi^{+} \rightarrow \mu^{+} + \nu_{\mu}$$

$$\mu^{+} \rightarrow e^{+} + \nu_{e} + \bar{\nu}_{\mu}$$

"\nu_{\mu}" flux = 2\times "\nu_{\mega"}" flux
"Down" flux = "Up" flux



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Schematic of Iron Calorimeter (ICAL) Detector at INO







51 kt world's largest electromagnet



Glass RPC for detecting charged particles ~30,000 RPCs required, ~ 3.8 M channels

[B] smoothed 1.71038 1.52464 Actions 1.28228 1.19722 1.11176 0.940721 0.895206 0.769691 0.684176 0.52666 0.612146 0.42783 0.342114 0.266699 0.171084 0.0855860 5.33811 p-004

B-field for 60 kA-turns, typical low C steel

Observe atmospheric neutrinos and antineutrinos separately over a wide range of energies & baselines using 50 kt magnetized Iron Calorimeter (ICAL) detector

ICAL Design and Specfications



Resistive plate chamber (RPC) (active element) sandwiched between iron plates (passive element)

Detector Response of ICAL

In CC events at ICAL:

• Muon \rightarrow track

 0.35_{1}

0.3

0.25

0.15

0.1

0.0

 $\sigma/P_{\rm in}$

• Hadron \rightarrow shower



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°P_{in} (GeV/c)

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Coordinates of INO



Located 115 km west of the Madurai city in the Theni district of Tamil Nadu

Come up with an underground lab, surface facilities, & build massive 50 kton magnetized Iron Calorimeter (ICAL) detector to study atmospheric neutrinos

Three Neutrino Mixing

Three neutrino mixing firmly established...

flavor states flavor states flavor states
$$\begin{pmatrix} v_e \\ v_\mu \\ v_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$
 neutrino mass states



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Some Things We Know and Don't Know

Three neutrino mixing firmly established...



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India-based Neutrino Observatory

India based Neutrino Observatory at Pottipuram (Theni)



Event Display Inside the ICAL Detector



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Identifying Neutrino Mass Hierarchy with ICAL

