



Probing New Physics with Heavy Hadron decays

Fulvia De Fazio
INFN Bari

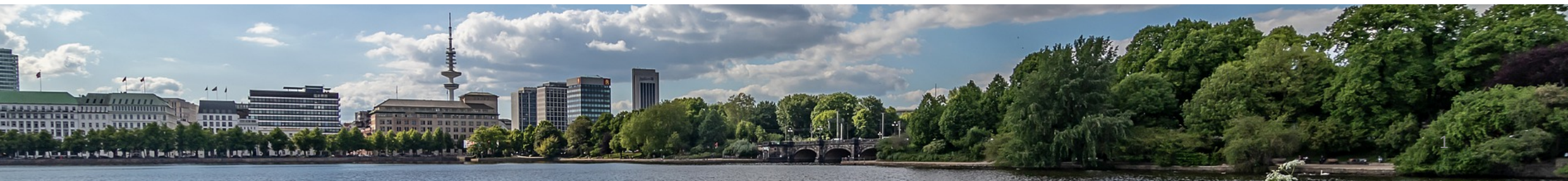
EPS 2021
July 26-30 Hamburg

based on:

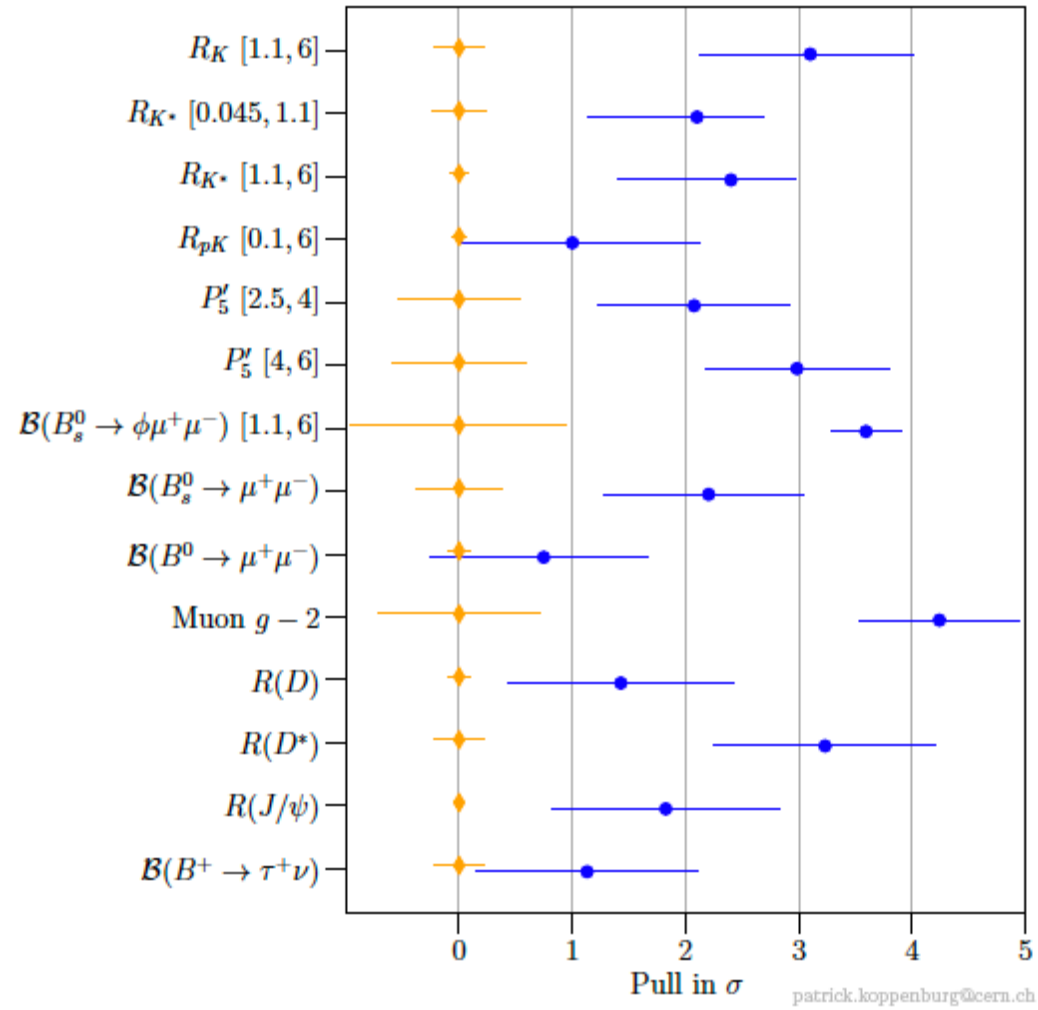
Inclusive semileptonic Λ_b decays in the Standard Model and beyond

P. Colangelo, F. Loparco, FDF

JHEP 11 (2020) 032 , arXiv:2006.13759



New & Old Anomalies



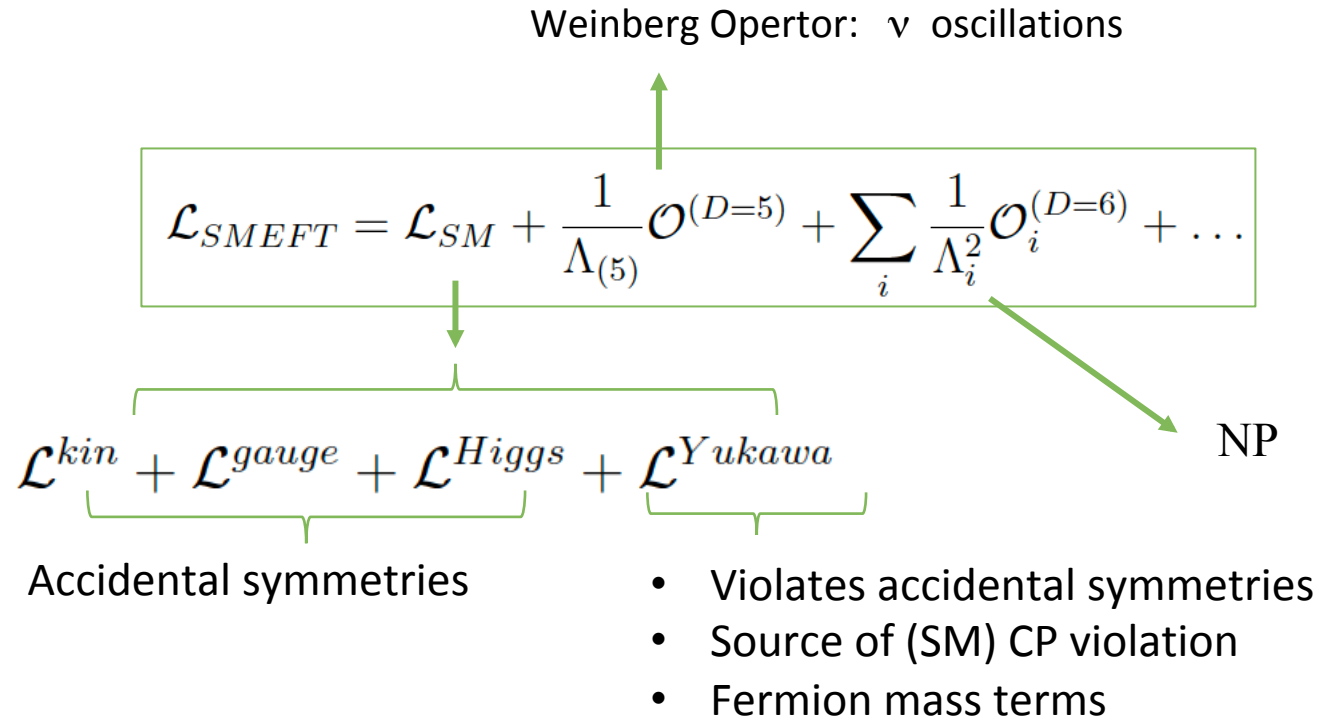
New & Old Anomalies

- 1) SM allowed processes: deviations from expectations → NP
 - tree level decays ($R(D^{(*)})$ & co.)
 - loop (rare) decays ($R(K^{(*)})$, P'_5 , ...)
 - puzzling quantities (V_{cb} , V_{ub} , ε'/ε , $(g-2)_\mu$...).
- 2) SM forbidden processes: observation → NP
 - LFV decays $\tau \rightarrow 3\mu$, $\mu \rightarrow e \gamma$...

- many hints of LFU violation \rightarrow relation between 1) and 2) expected
- Common origin of the anomalies?
ex: Problems with V_{cb} and V_{ub} determinations might be correlated
with observed anomalies in tree level modes

P. Colangelo, FDF
PRD 95 (17) 011701

Systematic extension of the SM



Exclusive $b \rightarrow c, u$ modes

- Identifications of suitable observables
 - Issue of FF uncertainties
 - $b \rightarrow c$ unsolved debate: role of the parametrization: BGL vs CLN in heavy-to-heavy FF
- Proposed to reconcile inclusive vs exclusive V_{cb} determinations



what about V_{ub} ?

Inclusive $b \rightarrow c, u$ modes (exploit HQE)

- Identifications of suitable observables
- Higher orders in $1/m_Q$ and α_s
- role of shape function

Common starting point (U=c,u)

$$\begin{aligned}
H_{\text{eff}}^{b \rightarrow U \ell \nu} = \frac{G_F}{\sqrt{2}} V_{Ub} \bigg[& (1 + \epsilon_V^\ell) (\bar{U} \gamma_\mu (1 - \gamma_5) b) (\bar{\ell} \gamma^\mu (1 - \gamma_5) \nu_\ell) \\
& + \epsilon_S^\ell (\bar{U} b) (\bar{\ell} (1 - \gamma_5) \nu_\ell) + \epsilon_P^\ell (\bar{U} \gamma_5 b) (\bar{\ell} (1 - \gamma_5) \nu_\ell) \\
& + \epsilon_T^\ell (\bar{U} \sigma_{\mu\nu} (1 - \gamma_5) b) (\bar{\ell} \sigma^{\mu\nu} (1 - \gamma_5) \nu_\ell) \\
& + \epsilon_R^\ell (\bar{U} \gamma_\mu (1 + \gamma_5) b) (\bar{\ell} \gamma^\mu (1 - \gamma_5) \nu_\ell) \bigg] + h.c. \quad .
\end{aligned}$$

Extends the SM effective Hamiltonian

Common starting point (U=c,u)

$$H_{\text{eff}}^{b \rightarrow U \ell \nu} = \frac{G_F}{\sqrt{2}} V_{Ub} \left[\begin{aligned} & (1 + \epsilon_V^\ell) (\bar{U} \gamma_\mu (1 - \gamma_5) b) (\bar{\ell} \gamma^\mu (1 - \gamma_5) \nu_\ell) \\ & + \epsilon_S^\ell (\bar{U} b) (\bar{\ell} (1 - \gamma_5) \nu_\ell) + \epsilon_P^\ell (\bar{U} \gamma_5 b) (\bar{\ell} (1 - \gamma_5) \nu_\ell) \\ & + \epsilon_T^\ell (\bar{U} \sigma_{\mu\nu} (1 - \gamma_5) b) (\bar{\ell} \sigma^{\mu\nu} (1 - \gamma_5) \nu_\ell) \\ & + \epsilon_R^\ell (\bar{U} \gamma_\mu (1 + \gamma_5) b) (\bar{\ell} \gamma^\mu (1 - \gamma_5) \nu_\ell) \end{aligned} \right] + h.c. \dots$$

SM

complex
LF-dependent couplings

Extends the SM effective Hamiltonian

- Investigate correlations to similar processes in meson (B_s , B_c) decays and in b-baryon (Λ_b , Ξ_b , Ω_b) decays
- inclusive widths : use
 - Optical theorem
 - Heavy Quark expansion (HQE)



Pioneering works: J. Chay, H. Georgi and B. Grinstein, PLB 247 (1990) 399
I.I.Y. Bigi, M.A. Shifman, N.G. Uraltsev and A.I. Vainshtein, PRL 71 (1993) 496

Inclusive decay width

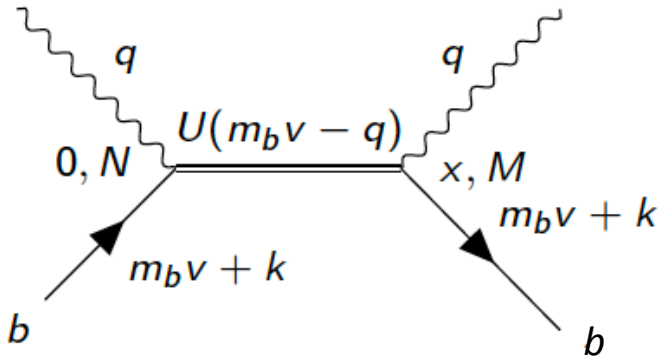
$$H_{\text{eff}}^{b \rightarrow U \ell \nu} = \frac{G_F}{\sqrt{2}} V_{Ub} \sum_{i=1}^5 C_i^\ell J_M^{(i)} L^{(i)M} + h.c.$$

Leptonic currents

$$C_1^\ell = (1 + \epsilon_V^\ell)$$

$$C_{2,3,4,5}^\ell = \epsilon_{S,P,T,R}^\ell$$

Hadronic currents



SM: $\epsilon_{V,S,P,T,R}^\ell = 0$

$$J_\mu^{(1)} = \bar{U} \gamma_\mu (1 - \gamma_5) b$$

$$L^{\mu(1)} = \bar{\ell} \gamma_\mu (1 - \gamma_5) \nu_\ell$$

$$d\Gamma = d\Sigma \frac{G_F^2 |V_{Ub}|^2}{4m_H} \sum_{i,j} C_i^* C_j \underbrace{(W^{ij})_{MN}}_{\text{Hadronic tensor}} \underbrace{(L^{ij})^{MN}}_{\text{Leptonic tensor}}$$

Hadronic tensor

Leptonic tensor

$$(W^{ij})_{MN} = \frac{1}{\pi} \text{Im}(T^{ij})_{MN}$$

$$(T^{ij})_{MN} = i \int d^4x e^{i(m_b v - q) \cdot x} \langle H_b(v, s) | T[\hat{J}_M^{(i)\dagger}(x) \hat{J}_N^{(j)}(0)] | H_b(v, s) \rangle$$

$$= \langle H_b(v, s) | \bar{b}_v(0) \Gamma_M^{(i)\dagger} S_U(p_X) \Gamma_N^{(j)} b_v(0) | H_b(v, s) \rangle$$

Intermediate U quark propagator

Expansion:

$$S_U(p_X) = S_U^{(0)} - S_U^{(0)}(i\mathcal{D})S_U^{(0)} + S_U^{(0)}(i\mathcal{D})S_U^{(0)}(i\mathcal{D})S_U^{(0)} + \dots$$

$$\frac{1}{m_b v - q - m_U}$$

$$\begin{aligned} \frac{1}{\pi} \text{Im}(T^{ij})_{MN} = & \frac{1}{\pi} \text{Im} \frac{1}{\Delta_0} \langle H_b(v, s) | \bar{b}_v [\Gamma_M^{(i)\dagger} \mathcal{P} \Gamma_N^{(j)}] b_v | H_b(v, s) \rangle \\ & - \frac{1}{\pi} \text{Im} \frac{1}{\Delta_0^2} \langle H_b(v, s) | \bar{b}_v [\Gamma_M^{(i)\dagger} \mathcal{P} \gamma^{\mu_1} \mathcal{P} \Gamma_N^{(j)}] (iD_{\mu_1}) b_v | H_b(v, s) \rangle \\ & + \frac{1}{\pi} \text{Im} \frac{1}{\Delta_0^3} \langle H_b(v, s) | \bar{b}_v [\Gamma_M^{(i)\dagger} \mathcal{P} \gamma^{\mu_1} \mathcal{P} \gamma^{\mu_2} \mathcal{P} \Gamma_N^{(j)}] (iD_{\mu_1}) (iD_{\mu_2}) b_v | H_b(v, s) \rangle \\ & - \frac{1}{\pi} \text{Im} \frac{1}{\Delta_0^4} \langle H_b(v, s) | \bar{b}_v [\Gamma_M^{(i)\dagger} \mathcal{P} \gamma^{\mu_1} \mathcal{P} \gamma^{\mu_2} \mathcal{P} \gamma^{\mu_3} \mathcal{P} \Gamma_N^{(j)}] (iD_{\mu_1}) (iD_{\mu_2}) (iD_{\mu_3}) b_v | H_b(v, s) \rangle . \end{aligned}$$

$$\left\{ \begin{array}{l} \Delta_0 = p_U^2 - m_U^2 \\ p_U = m_b v - q \end{array} \right.$$

HQE

Requires hadronic matrix elements with increasing number of derivatives:

$$\mathcal{M}_{\mu_1 \dots \mu_n} = \langle H_b(v, s) | (\bar{b}_v)_a (iD_{\mu_1}) \dots (iD_{\mu_n}) (b_v)_b | H_b(v, s) \rangle$$

can be expressed in terms of non perturbative parameters

$$\mathcal{O}\left(\frac{1}{m_b^n}\right) \dots \left\{ \mathcal{O}\left(\frac{1}{m_b^3}\right) \left\{ \begin{array}{l} \mathcal{O}\left(\frac{1}{m_b^2}\right) \left\{ \begin{array}{l} -2 M_H \hat{\mu}_\pi^2 = \langle H_b | \bar{b}_v i D^\mu i D_\mu b_v | H_b \rangle \\ 2 M_H \hat{\mu}_G^2 = \langle H_b | \bar{b}_v (-i \sigma_{\mu\nu}) i D^\mu i D^\nu b_v | H_b \rangle \end{array} \right. \\ 2 M_H \hat{\rho}_D^3 = \langle H_b | \bar{b}_v i D^\mu (i v \cdot D) i D_\mu b_v | H_b \rangle \\ 2 M_H \hat{\rho}_{LS}^3 = \langle H_b | \bar{b}_v (-i \sigma_{\mu\nu}) i D^\mu (i v \cdot D) i D^\nu b_v | H_b \rangle \end{array} \right. \right. \\ \dots \end{array}$$

$$\mathcal{O}\left(\frac{1}{m_b^n}\right) \dots \left\{ \mathcal{O}\left(\frac{1}{m_b^3}\right) \left\{ \begin{array}{l} \mathcal{O}\left(\frac{1}{m_b^2}\right) \left\{ \begin{array}{l} -2 M_H \hat{\mu}_\pi^2 = \langle H_b | \bar{b}_\nu i D^\mu i D_\mu b_\nu | H_b \rangle \\ 2 M_H \hat{\mu}_G^2 = \langle H_b | \bar{b}_\nu (-i \sigma_{\mu\nu}) i D^\mu i D^\nu b_\nu | H_b \rangle \end{array} \right. \\ 2 M_H \hat{\rho}_D^3 = \langle H_b | \bar{b}_\nu i D^\mu (i \not{v} \cdot D) i D_\mu b_\nu | H_b \rangle \\ 2 M_H \hat{\rho}_{LS}^3 = \langle H_b | \bar{b}_\nu (-i \sigma_{\mu\nu}) i D^\mu (i \not{v} \cdot D) i D^\nu b_\nu | H_b \rangle \end{array} \right. \\ \dots \end{array} \right.$$

$\hat{\mu}_\pi^2$ matrix element of the kinetic energy operator

is different for different hadrons (i.e. B and Λ_b)

$$\mu_\pi^2(B) - \mu_\pi^2(\Lambda_b) = \frac{2m_b m_c}{m_b - m_c} [(m_{\Lambda_b} - m_{\Lambda_c}) - (\overline{m}_B - \overline{m}_D)] (1 + \mathcal{O}(1/m_{b,c}^2))$$

$$\hat{\mu}_\pi^2(\Lambda_b) = (0.50 \pm 0.1) \text{ GeV}^2$$

$$\mathcal{O}\left(\frac{1}{m_b^n}\right) \dots \left\{ \mathcal{O}\left(\frac{1}{m_b^3}\right) \left\{ \begin{array}{l} \mathcal{O}\left(\frac{1}{m_b^2}\right) \left\{ \begin{array}{l} -2 M_H \hat{\mu}_V^2 = \langle H_b | \bar{b}_v i D^\mu i D_\mu b_v | H_b \rangle \\ 2 M_H \hat{\mu}_G^2 = \langle H_b | \bar{b}_v (-i \sigma_{\mu\nu}) i D^\mu i D^\nu b_v | H_b \rangle \end{array} \right. \\ 2 M_H \hat{\rho}_D^3 = \langle H_b | \bar{b}_v i D^\mu (i v \cdot D) i D_\mu b_v | H_b \rangle \\ 2 M_H \hat{\rho}_{LS}^3 = \langle H_b | \bar{b}_v (-i \sigma_{\mu\nu}) i D^\mu (i v \cdot D) i D^\nu b_v | H_b \rangle \end{array} \right. \dots \right.$$

$\hat{\mu}_G^2$ matrix element of the chromomagnetic operator

depends on the spin of the hadron and can be fixed from data (mass splittings)

$$\hat{\mu}_G^2(\Lambda_b) = 0$$

$$\mathcal{O}\left(\frac{1}{m_b^n}\right) \dots \left\{ \mathcal{O}\left(\frac{1}{m_b^3}\right) \left\{ \begin{array}{l} \mathcal{O}\left(\frac{1}{m_b^2}\right) \left\{ \begin{array}{l} -2 M_H \hat{\mu}_\pi^2 = \langle H_b | \bar{b}_v i D^\mu i D_\mu b_v | H_b \rangle \\ 2 M_H \hat{\mu}_G^2 = \langle H_b | \bar{b}_v (-i \sigma_{\mu\nu}) i D^\mu i D^\nu b_v | H_b \rangle \end{array} \right. \right. \\ 2 M_H \hat{\rho}_D^3 = \langle H_b | \bar{b}_v i D^\mu (i v \cdot D) i D_\mu b_v | H_b \rangle \\ 2 M_H \hat{\rho}_{LS}^3 = \langle H_b | \bar{b}_v (-i \sigma_{\mu\nu}) i D^\mu (i v \cdot D) i D^\nu b_v | H_b \rangle \end{array} \right. \dots$$

$\hat{\rho}_D^3$

Darwin term approx:

$$\rho_D^3(\Lambda_b) \simeq \rho_D^3(B)$$

$$\rho_D^3(\Lambda_b) = (0.17 \pm 0.08) \text{ GeV}^3$$

$\hat{\rho}_{LS}^3$

spin-orbit term

$$\hat{\rho}_{LS}^3(\Lambda_b) = 0$$

Results

$$\mathcal{M}_{\mu_1 \dots \mu_n} = \langle H_b(v, s) | (\bar{b}_v)_a (iD_{\mu_1}) \dots (iD_{\mu_n}) (b_v)_b | H_b(v, s) \rangle$$

General parametrization including **dependence on the spin** up to $O(m_b^{-3})$ not known before
Method derived for B mesons in B.M. Dassinger, T. Mannel and S. Turczyk, JHEP 03 (2007) 087

New terms depending on the spin appear for a **polarized baryon**

- Analytic results
- Main outcome of our study
- Applies to all spin ½ baryons

Results at $O(m_b^{-2})$: A.V. Manohar and M.B. Wise, PRD 49 (1994) 1310
S. Balk, J.G. Korner and D. Pirjol, EPJC 1 (1998) 221

$$\mathcal{M}^{\rho\sigma\lambda} = M_H \left[\left(\frac{\hat{\rho}_D^3}{3} \Pi^{\rho\lambda} v^\sigma P_+ + \frac{\hat{\rho}_{LS}^3}{6} v^\sigma i \epsilon^{\rho\lambda\alpha\beta} v_\alpha S_\beta \right) - \left(\frac{\hat{\rho}_D^3}{3} \Pi^{\rho\lambda} v^\sigma s^\mu S_\mu - \frac{\hat{\rho}_{LS}^3}{2} v^\sigma i \epsilon^{\rho\lambda\alpha\beta} v_\alpha s_\beta P_+ \right) \right]$$

$$\begin{aligned} \mathcal{M}^{\rho\sigma} = M_H \left[\left(\frac{\hat{\mu}_\pi^2}{3} \Pi^{\rho\sigma} P_+ + \frac{\hat{\mu}_G^2}{6} i \epsilon^{\rho\sigma\alpha\beta} v_\alpha S_\beta + \right. \right. \\ \left. \left. + \frac{\hat{\rho}_D^3 + \hat{\rho}_{LS}^3}{24m_b} (4 (i \epsilon^{\rho\sigma\alpha\beta} v_\alpha S_\beta - v^\rho v^\sigma \not{v}) + \right. \right. \\ \left. \left. + v^\rho (2 \gamma^\sigma + \not{v} \gamma^\sigma - \gamma^\sigma \not{v}) + v^\sigma (2 \gamma^\rho + \not{v} \gamma^\rho - \gamma^\rho \not{v})) \right) + \right. \\ \left. + \left(- \frac{\hat{\mu}_\pi^2}{3} \Pi^{\rho\sigma} P_+ \not{5} + \frac{\hat{\mu}_G^2}{2} i \epsilon^{\rho\sigma\alpha\beta} v_\alpha s_\beta P_+ + \right. \right. \\ \left. \left. + \frac{\hat{\rho}_D^3}{12m_b} (6 i \epsilon^{\rho\sigma\alpha\beta} v_\alpha s_\beta + i (v^\rho \epsilon^{\sigma\mu\alpha\beta} - v^\sigma \epsilon^{\rho\mu\alpha\beta}) v_\alpha s_\beta \gamma_\mu + \right. \right. \\ \left. \left. + s^\rho v^\sigma \not{v} \gamma_5 + v^\rho s^\sigma (2 \gamma_5 + \not{v} \gamma_5) + (2 v^\rho v^\sigma \not{v} - v^\rho \gamma^\sigma - v^\sigma \gamma^\rho) \not{5} \gamma_5 \right) + \right. \\ \left. \left. + \frac{\hat{\rho}_{LS}^3}{8m_b} (4 i \epsilon^{\rho\sigma\alpha\beta} v_\alpha s_\beta + i (v^\rho \epsilon^{\sigma\mu\alpha\beta} - v^\sigma \epsilon^{\rho\mu\alpha\beta}) v_\alpha s_\beta \gamma_\mu + \right. \right. \\ \left. \left. + (s^\rho v^\sigma + v^\rho s^\sigma) \gamma_5 + (2 v^\rho v^\sigma \not{v} - v^\rho \gamma^\sigma - v^\sigma \gamma^\rho) \not{5} \gamma_5 \right) \right] \end{aligned}$$

$$\begin{aligned} \mathcal{M}^\rho = M_H \bigg[& \left(\frac{\hat{\mu}_\pi^2 - \hat{\mu}_G^2}{12m_b} (v^\rho (3 + 5 \not{v}) - 2 \gamma^\rho) - \frac{\hat{\rho}_D^3 + \hat{\rho}_{LS}^3}{12m_b^2} (4 v^\rho \not{v} - \gamma^\rho) \right) + \\ & + \left(- \frac{\hat{\mu}_\pi^2}{12m_b} ((v^\rho (3 + 5 \not{v}) - 2 \gamma^\rho) \not{5} \gamma_5 + 4 s^\rho P_+ \gamma_5) + \right. \\ & + \frac{\hat{\mu}_G^2}{4m_b} ((v^\rho (1 + 2 \not{v}) - \gamma^\rho) \not{5} \gamma_5 + s^\rho \gamma_5) + \\ & + \frac{\hat{\rho}_D^3}{12m_b^2} ((v^\rho (1 + 4 \not{v}) - 2 \gamma^\rho) \not{5} \gamma_5 + s^\rho (2 - \not{v}) \gamma_5) + \\ & \left. \left. + \frac{\hat{\rho}_{LS}^3}{8m_b^2} ((3 v^\rho \not{v} - \gamma^\rho) \not{5} \gamma_5 + s^\rho \gamma_5) \right) \right] \end{aligned}$$

$$\mathcal{M} = M_H \left[\left(P_+ - \frac{\hat{\mu}_\pi^2 - \hat{\mu}_G^2}{4m_b^2} \right) + \left(P_+ + \frac{\hat{\mu}_\pi^2}{24m_b^2} (7 + 5 \not{v}) - \frac{\hat{\mu}_G^2}{8m_b^2} (3 + \not{v}) - \frac{\hat{\rho}_D^3}{6m_b^3} P_- \right) \not{5} \gamma_5 \right]$$

Fully differential decay width

$$\frac{d^4\Gamma}{dE_\ell dq^2 dq_0 d\cos\theta_P}$$

- E_ℓ : lepton energy, with $p_\ell = (E_\ell, \mathbf{p}_\ell)$;
- q^2 : dilepton invariant mass, with $q = (q_0, \mathbf{q})$;
- θ_P : angle between the hadron spin \mathbf{s} and the lepton 3-momentum \mathbf{p}_ℓ , i.e. $\cos\theta_P = \frac{\mathbf{s} \cdot \mathbf{p}_\ell}{|\mathbf{s}||\mathbf{p}_\ell|}$.

Computed with the inclusion of all NP operators

$$\Gamma(H_b \rightarrow X_U \ell^- \bar{\nu}_\ell) = \Gamma_0 \sum_{i,j} g_i^* g_j \left[c_0^{(i,j)} + \frac{\hat{\mu}_\pi^2}{m_b^2} c_{\hat{\mu}_\pi^2}^{(i,j)} + \frac{\hat{\mu}_G^2}{m_b^2} c_{\hat{\mu}_G^2}^{(i,j)} + \frac{\hat{\rho}_D^3}{m_b^3} c_{\hat{\rho}_D^3}^{(i,j)} + \frac{\hat{\rho}_{LS}^3}{m_b^3} c_{\hat{\rho}_{LS}^3}^{(i,j)} \right]$$

$$\Gamma_0 = \frac{G_F^2 |V_{Ub}|^2 m_b^5}{192\pi^3}$$

Include the effects of NP through the ε parameters

Complete expression at $O(1/m_b^3)$ in SM, including all NP operators and for non vanishing lepton mass
Rather lengthy [P. Colangelo, F. Loparco, FDF JHEP 11 \(2020\) 032](#), [arXiv:2006.13759](#)

Numerical Results

Identification of observables sensitive to Λ_b polarization and to BSM effects
(Longitudinal polarization expected for Λ_b resulting from b quark produced in top or Z^0 decays)

P. Colangelo, F. Loparco, FDF
JHEP 11 (2020) 032

$$\frac{d\Gamma(\Lambda_b \rightarrow X_U \ell \bar{\nu}_\ell)}{d\cos\theta_P} = A_\ell^U + B_\ell^U \cos\theta_P$$

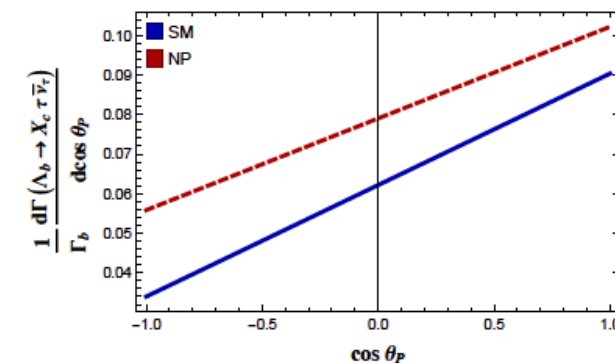
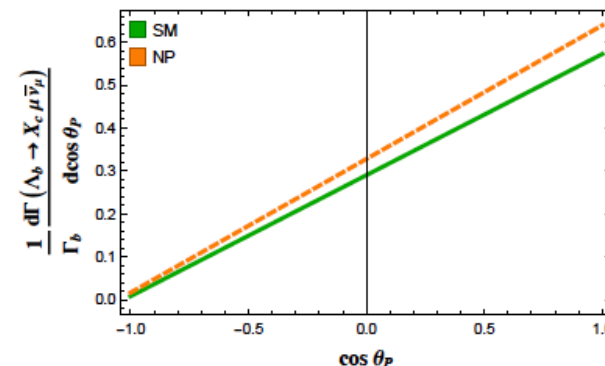
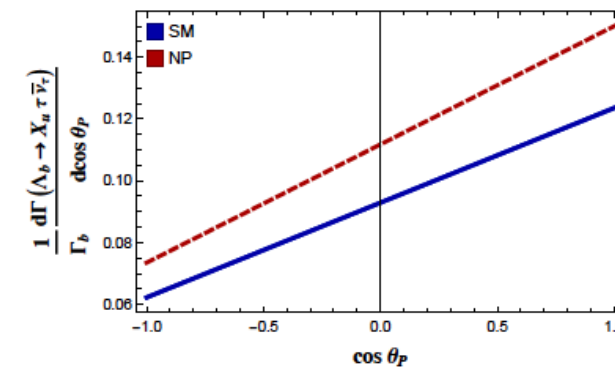
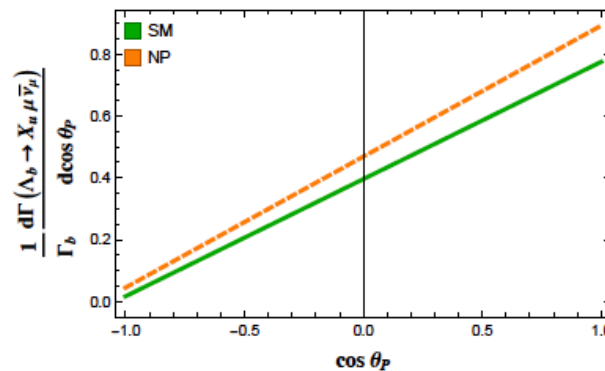
angle between
 ℓ direction and Λ_b spin

$$R_{\Lambda_b}(X_U) = \frac{A_\tau^U}{A_\mu^U}$$

$$R_S^U = \frac{B_\tau^U}{B_\mu^U}$$

slope ratios for $\ell=\tau$ and $\ell=\mu$

analogous to $R(D^{(*)})$
does not require
considering
polarization



NP: benchmark points

b -> u P. Colangelo, F. Loparco, FDF, PRD 100 (19) 075037

b -> c P. Colangelo, FDF, JHEP 06 (18) 082 (BP1)

R.X. Shi, L.S. Geng, B. Grinstein, S. Jager, J.M. Camalich JHEP 12 (19) 065. (BP2)

Identification of observables sensitive to Λ_b polarization and to BSM effects
(Longitudinal polarization expected for Λ_b resulting from b quark produced in top or Z^0 decays)

P. Colangelo, F. Loparco, FDF
JHEP 11 (2020) 032

$$\frac{d\Gamma(\Lambda_b \rightarrow X_U \ell \bar{\nu}_\ell)}{d\cos\theta_P} = A_\ell^U + B_\ell^U \cos\theta_P$$

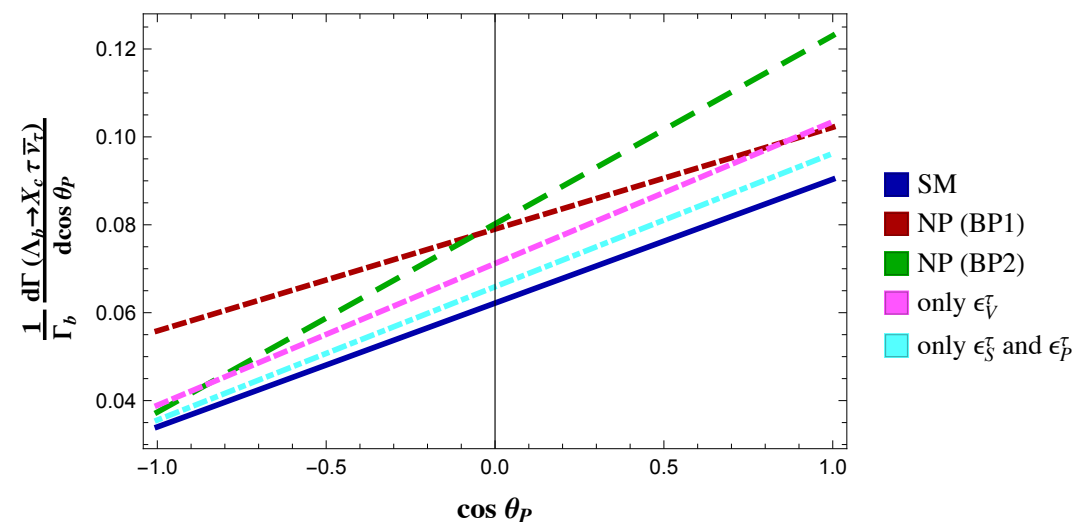
angle between ℓ direction and Λ_b spin

$$R_{\Lambda_b}(X_U) = \frac{A_\tau^U}{A_\mu^U}$$

$$R_S^U = \frac{B_\tau^U}{B_\mu^U}$$

slope ratios for $\ell=\tau$ and $\ell=\mu$

analogous to $R(D^{(*)})$
does not require
considering
polarization



NP: benchmark points

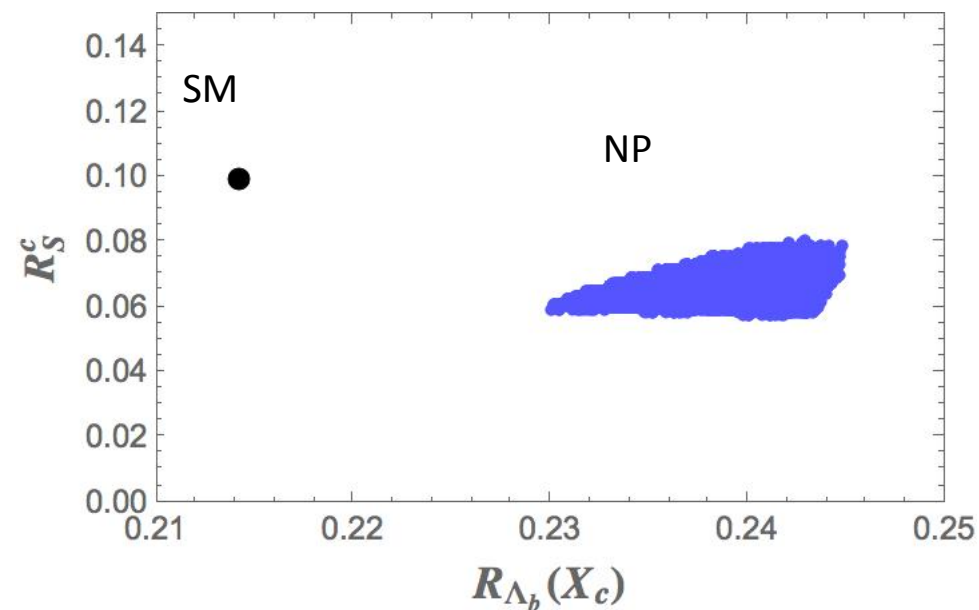
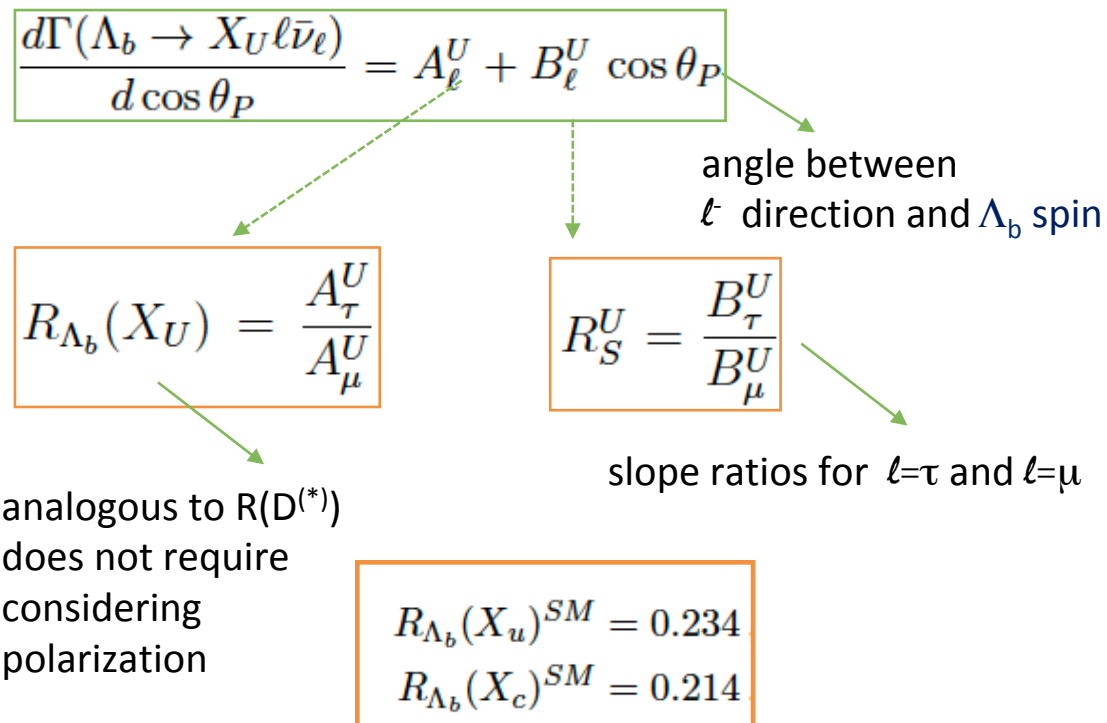
b \rightarrow u P. Colangelo, F. Loparco, FDF, PRD 100 (19) 075037

b \rightarrow c P. Colangelo, FDF, JHEP 06 (18) 082 (BP1)

R.X. Shi, L.S. Geng, B. Grinstein, S. Jager, J.M. Camalich JHEP 12 (19) 065. (BP2)

Identification of observables sensitive to Λ_b polarization and to BSM effects
(Longitudinal polarization expected for Λ_b resulting from b quark produced in top or Z^0 decays)

P. Colangelo, F. Loparco, FDF
JHEP 11 (2020) 032



- clean observable but hard to measure at LHC
- possible using polarized Λ_b from Z decays at FCC-e or muon collider

NP: benchmark points

b \rightarrow u P. Colangelo, F. Loparco, FDF, PRD 100 (19) 075037

b \rightarrow c P. Colangelo, FDF, JHEP 06 (18) 082 (BP1)

R.X. Shi, L.S. Geng, B. Grinstein, S. Jager, J.M. Camalich JHEP 12 (19) 065. (BP2)

Inclusive heavy hadron decays

- Further testing ground for NP
- Essential to understand the inclusive/exclusive discrepancy in the V_{cb} and V_{ub} determinations

Theoretical improvement:

calculation of the necessary hadronic matrix elements in the case of a polarized baryon at $O(1/m_b^3)$ and including all possible NP operators

Results: distributions in $\cos(\theta_p)$ sensitive to NP

- Ratio $R(\Lambda_b)$ can deviate from SM
- Ratio of slope parameter to be considered together with $R(\Lambda_b)$
- correlation between the two ratio shows the predicted pattern of deviation from SM
- measurable at FCC