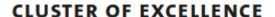
# Nonequilibrium scalar fields in the early universe

**Aleksandr Chatrchyan** 





QUANTUM UNIVERSE



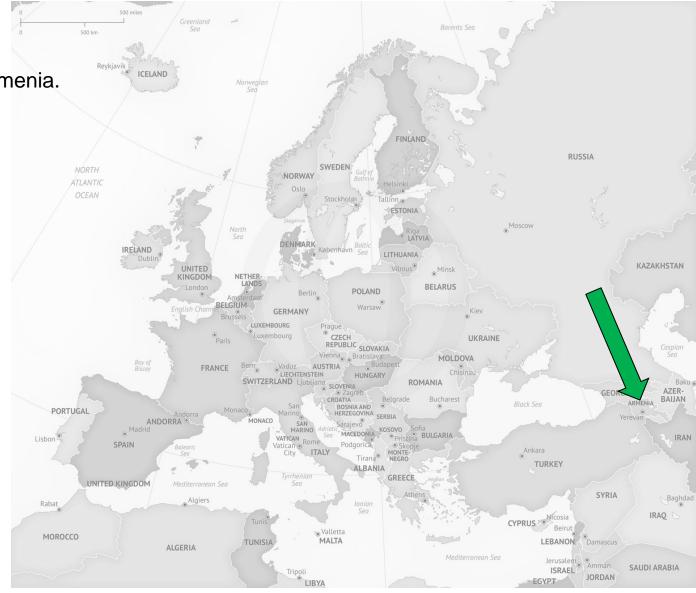
Email: aleksandr.chatrchyan@desy.de

## **About Me**

#### **Background, past activities**

2014: Bachelor's in physics, Yerevan State University, Armenia.





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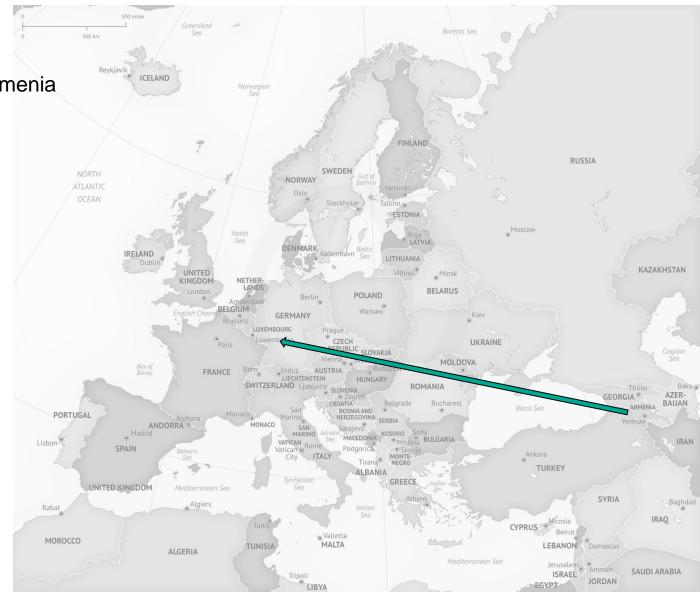
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• Thesis: <u>link</u>, supervisors: J. Berges, J. Jaeckel





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Main area of research: nonequilibrium quantum field theory

- Description in real-time
- Matrix elements for |in> and |out> states not so useful
- Applications: Heavy-ion collisions, ultracold atom, early universe physics

Approximate descriptions

Semiclassical (lattice) simulations, kinetic description, 2PI effective action framework, etc.

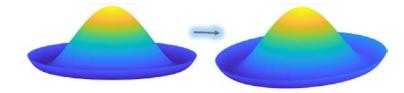


# My Research

## **Activities and challenges**

## Why scalars?

• Higgs, QCD axion, string theory compactifications.



## Significance for cosmology

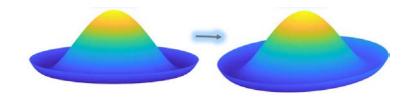
• Inflation & reheating, nonthermal scalar dark matter

# My Research

#### **Activities and challenges**

#### Why scalars?

Higgs, QCD axion, string theory compactifications.



## Significance for cosmology

Inflation & reheating, nonthermal scalar dark matter

## Can have very rich dynamics

Large occupation numbers: coherent field behavior (BEC)

• Axion-like particle dark matter <u>1903.03116</u>, <u>2004.07844</u>

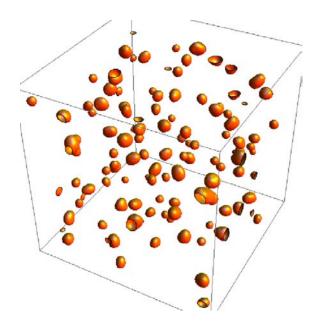
• Attractors and universal dynamics <u>1707.07696</u>, <u>2008.02290</u>

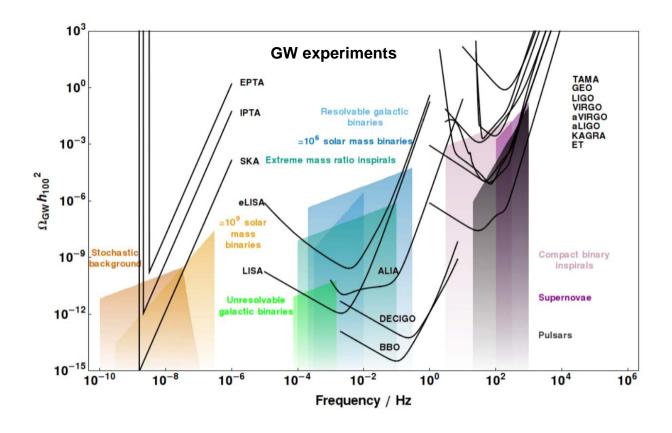
## My Research

#### **Activities and challenges**

 Gravitational waves signatures of scalar field dynamics in the early universe

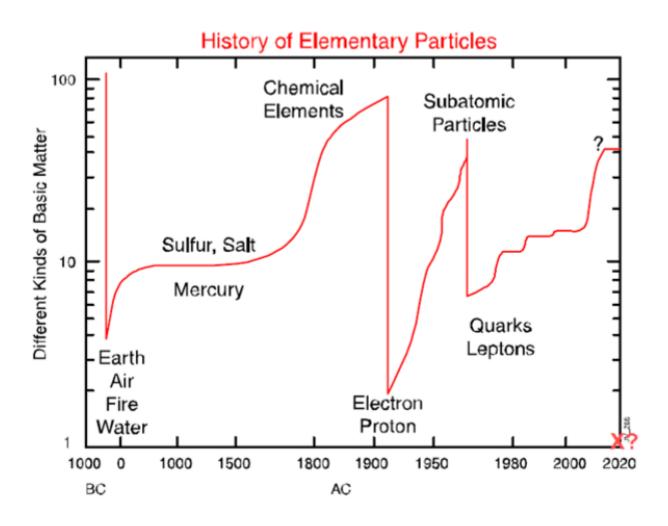
The role of oscillons (scalar field clumps),
power spectrum and implications for structure formation





## **My Favorite Plot**

Or the one question you always wanted to ask!



Taken from <u>1311.1769</u>

# Thanks for your attention!