## Dark matter

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# Neutralino as dark matter candidate<sup>1</sup>

Motivation: Estimating the contribution of neutralinos to dark matter from LHC data

- R-parity conservation
  - ⇒ Pairwise production and annihilation of sparticles
  - $\implies$  Lightest supersymmetric particle (LSP) is stable
- In many models: Lightest neutralino  $\tilde{\chi}_1^0$  is LSP

 $\Longrightarrow \tilde{\chi}_1^0$  is good candidate for dark matter

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<sup>1 &</sup>quot;Supersymmetrie: Parameterbestimmung", T 52.3, T 52.4 ( ) > ( ) > ( ) > ( ) > ( )

#### Neutralino as dark matter candidate

- Cross-section for annihilation of WIMPs
  - proportional to  $\frac{1}{DM \text{ density}}$
  - well known
- Annihilation of neutralinos
  - $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \longrightarrow f \bar{f} via A^0 or Z$
  - Cross-section depends on  $m_{A^0}$  and  $\tan \beta$

 $\implies$  Measuring  $m_{A^0}$  and tan  $\beta$  to determine the contribution of neutralions to dark matter

or

 $\implies$  Assuming that dark matter is totally made up of  $\tilde{\chi}_1^0$  to introduce their annihilation cross-section as constraint for SUSY searches

#### Neutralino as dark matter candidate

#### Analysis:

- Vary  $\tan \beta$  and  $A_0$  for given  $m_0$  and  $m_{1/2}$  (and given sign  $\mu = +$  (?))
- Studie which scenarios match to the hypothesis that dark matter is totally made up of neutralinos

#### Result:

For many points in the  $m_0-m_{1/2}$  plane values for tan  $\beta$  and  $A^0$  (?) can be found such that the hypothesis may be true

# Kaluza-Klein-excitations as dark matter candidates<sup>2</sup>

- Models with infinite extra dimensions (ED)
  - States in ED are quantized
  - SM particles excited to these states are called Kaluza-Klein-particles (KK-particle)
  - ED not visible
    - $\implies$  Momentum of KK-particle in ED shows up as additional mass

⇒ Lightest KK-particle is good candidate for dark matter

→ Why is it stable?

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### Kaluza-Klein-excitations as dark matter candidates

- Significant excess in the cosmic-ray positron fraction observed by PAMELA experiment may be explained by
  - Extra dimensions:

Annihilation of lightest KK-particles (KK-photons)

- ⇒ BUT: Cross-section consistent with the dark matter observed relic abundance (?) is too small to explain excess and requires to introduce "boost" factors
- Many other explainations (e.g. decaying dark matter, pulsars)

⇒ Models are "tuned" to explain excess, but are not motivated by a fundamental principle like SUSY