

SOMMERFELD AND SUDAKOV CORRECTIONS IN HEAVY NEUTRALINO ANNIHILATION

Guillaume CHALONS

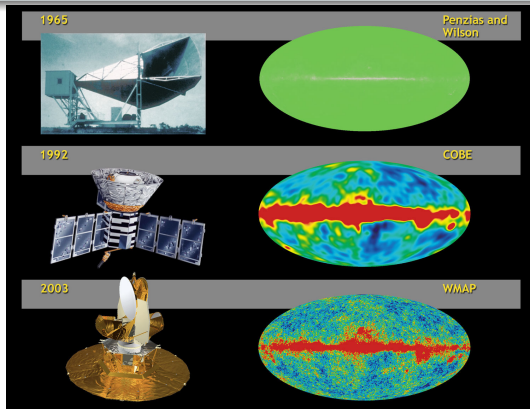
in collaboration with N. Baro, F. Boudjema, Sun Hao

August, 26th, 2010



RELIC DENSITY OF DARK MATTER

- WMAP : $0.0997 < \Omega_{DM} h^2 < 0.1221$ (10% precision)
- PLANCK : 2% precision



PRECISION MEASUREMENTS

RELIC DENSITY IN THE STANDARD SCENARIO

$$\Omega_{DM} h^2 \simeq \frac{3 \times 10^{-27} \text{ cm}^3 \text{ s}^{-1}}{\langle \sigma(\chi\chi \rightarrow \textcolor{red}{SM}) \nu \rangle}$$

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PRECISION

- Need for precise theoretical predictions w.r.t experimental measurements.
- Precision needed at the level of $\sigma \Rightarrow$ **One-loop** calculations (at least).
- If **SUSY** found \Rightarrow **Reconstruction** of fundamental underlying parameters.
- **Radiative corrections** must be under **control** to be able to **constrain** the cosmological **underlying** scenario.

EW + QCD corrections

- $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \gamma\gamma, Z\gamma, gg$: Boudjema, Semenov, Temes, *Phys. Rev.* **D72**, 055024 (2005)
- $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow ZZ, W^+ W^-$: Baro, Boudjema, Semenov, *Phys. Lett.* **B660** (2008) 550
Baro, Boudjema, G.C., Sun Hao, *Phys. Rev.* **D81** (2008) 015005
- $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \tau^+ \tau^-, b\bar{b}$: Baro, Boudjema, Semenov, *Phys. Lett.* **B660** (2008) 550
- Co-annihilation with $\tilde{\tau}$: Baro, Boudjema, Semenov, *Phys. Lett.* **B660** (2008) 550,

QCD corrections

- Co-annihilation with $\tilde{\tau}$ Freitas *Phys. Lett.* **B652** (2007) 280
- Annihilation into massive quarks Hermann, Klasen, Kovarik *Phys. Rev.* **D79** (2009)
Herrmann, Klasen, *Phys. Rev.* **D76** (2007) 117704
Herrmann, Klasen and Kovarik, *Phys. Rev.* **D80** (2009) 085025

- At **tree-level** we have for $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow WW$ 7 diagrams.
- Relic density **predictions** involve **many** annihilation (and coannihilation) channels.

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- DarkSUSY [Bergström *et al.* (2004)]
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- Mainly $2 \rightarrow 2$ processes are taken into account in the **computation**.

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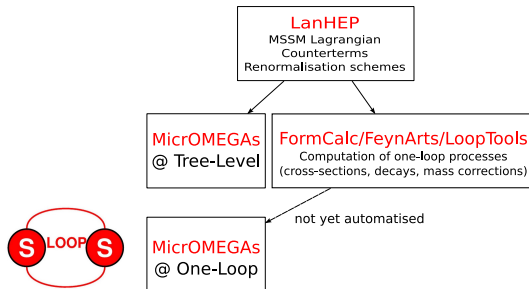
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- To deal with **IR** and **collinear divergencies** → include bremsstrahlung.
- To evaluate **many processes** entering $\langle \sigma v \rangle$.



- Evaluation of one-loop diagrams including a **complete** and **coherent** renormalisation of **each sector** of the MSSM with an **OS** scheme.
- Modularity between different renormalisation schemes.
- **Non-linear** gauge fixing.
- Handles a **large number** of Feynman diagrams.
- Checks : results **UV,IR** finite and **gauge** independent.

<http://code.sloops.free.fr/>

FERMION + GAUGE SECTOR

Input parameters as in the Standard Model $m_f, \alpha(0), M_W, M_Z$

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NEUTRALINOS/CHARGINOS SECTOR

Input parameters : 2 charginos $m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_2^\pm}$ and 1 neutralino $\tilde{\chi}_1^0$

COSMOLOGY (MicrOmegas)

THERMAL RELIC DENSITY

- Solve the Boltzmann equation

$$dn/dt = -3Hn - \langle \sigma v \rangle (n^2 - n_{\text{eq}}^2)$$

Processes with gauge boson production are the most difficult because gauge invariance plays a dominant role.

COSMOLOGY (MicroMegas)

THERMAL RELIC DENSITY

- Solve the Boltzmann equation
- Thermal average (MB approx.)

$$\langle \sigma v \rangle \propto \int_0^\infty (\sigma v) v^2 e^{-xv^2/4} dv$$

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COSMOLOGY (*MicrOmegas*)

THERMAL RELIC DENSITY

- Solve the *Boltzmann equation*
- Thermal average (MB approx.)

PARTICLE PHYSICS (*SloopS*)

GAUGE BOSON PRODUCTION

- *SU(2)_L* type couplings
- Channels contributions > 5% to Ωh^2 at TL corrected at one-loop
- Coannihilation channels
- *Large* number of diagrams to compute

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$\tilde{\chi}_1^0 \tilde{\chi}_1^\pm W^\pm$ and $\tilde{\chi}_1^\pm \tilde{\chi}_1^\pm Z^0$ vertices

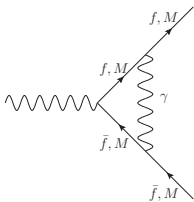
$\alpha(0) \rightarrow \alpha(M_Z^2)$: 13% corrections absorbed.

Also $q\bar{q}'$ production \rightarrow QCD corrections

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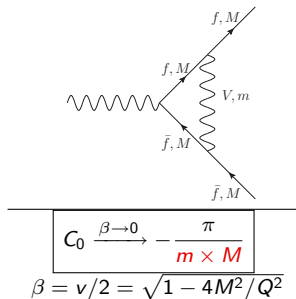
The diagram shows a scalar triangle loop. An incoming wavy line (photon) splits into two fermion lines (solid lines with arrows). The top fermion line is labeled f, M and the bottom fermion line is labeled \bar{f}, M . These two fermion lines meet at a vertex, from which a wavy line (photon) labeled γ emerges. The fermion lines are also labeled f, M and \bar{f}, M at the other vertex. Below the diagram, a box contains the expression for the scalar triangle loop integral C_0 in the limit $\beta \rightarrow 0$:

$$C_0 \xrightarrow{\beta \rightarrow 0} -\frac{\pi^2}{Q^2 \beta}$$

Below the box, the definition of β is given:

$$\beta = v/2 = \sqrt{1 - 4M^2/Q^2}$$

- **Singularities** arise in **scalar** triangle C_0 and box D_0 loop integrals when $\beta \rightarrow 0$.

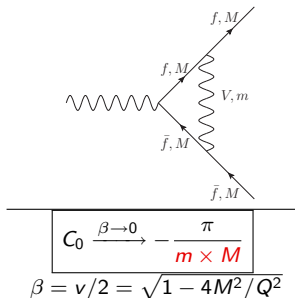


$$C_0 \xrightarrow{\beta \rightarrow 0} -\frac{\pi}{m \times M}$$

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- If two heavy masses M and one internal mass very small $m \ll M$
- See also study of these integrals in the non-relativistic limit with application to the relic density [Drees, Kim, Nagao *Phys. Rev. D* **81** (2010) 105004] and Kim's talk

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Results in **numerical instabilities** (vanishing **Gram** determinant).

Avoided using **Segmentation** of the loop integrals. [Boudjema-Semenov-Temes (2005)].

Idea : **split** 4 pt function \rightarrow 3 pt function when $\beta \rightarrow 0$.

Parameter	M_1	M_2	μ	t_β	M_3	$M_{L,\tilde{Q}}$	A_i	M_{A0}
Value(GeV)	3500	1800	4500	15	5000	5000	0	5000

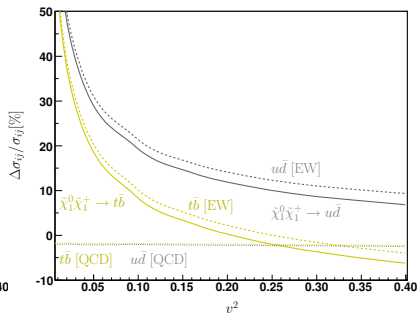
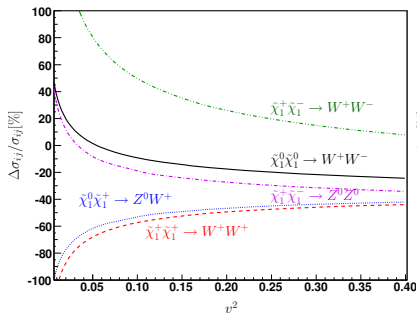
$$\tilde{\chi}_1^0 = 0.000\tilde{B} - 0.999\tilde{W} + 0.004\tilde{H}_1^0 + 0.032\tilde{H}_2^0$$

	Tree	
$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow W^+ W^-$ [10%]	a	+2.43
	b	+0.52
$\tilde{\chi}_1^+ \tilde{\chi}_1^+ \rightarrow W^+ W^+$ [10%]	a	+1.22
	b	+0.26
$\tilde{\chi}_1^0 \tilde{\chi}_1^+ \rightarrow Z^0 W^+$ [9%]	a	+0.51
	b	+0.12
$\tilde{\chi}_1^0 \tilde{\chi}_1^+ \rightarrow t\bar{b}$ [9%]	a	+0.54
	b	-0.23
$\tilde{\chi}_1^0 \tilde{\chi}_1^+ \rightarrow u\bar{d}$ [9%]	a	+0.54
	b	-0.23
$\tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow Z^0 Z^0$ [6%]	a	+0.73
	b	+0.16
$\tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow W^+ W^-$ [6%]	a	+0.65
	b	+0.17
$\Omega_\chi h^2$	0.0997	

- $m_{\tilde{\chi}_1^0} = 1799.1$ GeV
- $\delta(m_{\tilde{\chi}_1^+} - m_{\tilde{\chi}_1^0}) = 0.0003$ GeV

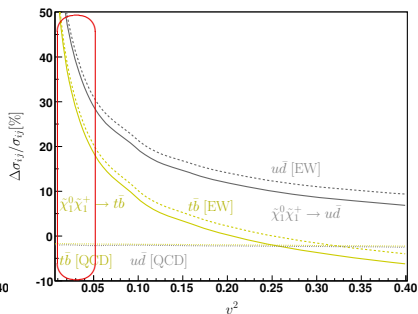
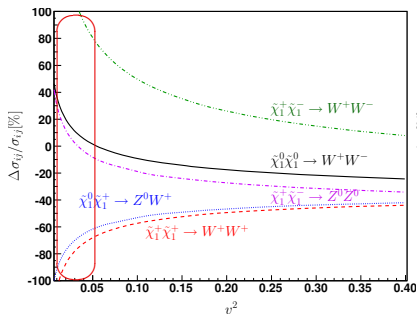
- $\sigma_0 v = a + bv^2$
- $m_{\tilde{\chi}_1^0}, m_{\tilde{\chi}_1^\pm}$ almost degenerate
- Coannihilation very important
- Degeneracy between processes $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow W^+ W^-$ and $\tilde{\chi}_1^+ \tilde{\chi}_1^+ \rightarrow W^+ W^+$
- A lot of processes contribute

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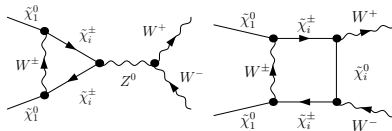


HEAVY-WINO NEUTRALINO

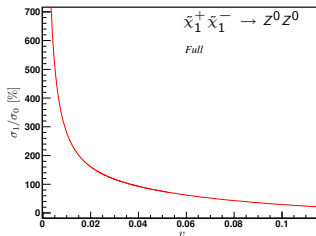
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- $M_W/m_{\tilde{\chi}_1^0} = 0.045 \Rightarrow W^\pm, Z^0$ bosons almost considered as massless.
- $v \rightarrow 0$: Large Sommerfeld (QED+EW) enhancement.



- The EW Sommerfeld effect is expected to be cut-off, as opposed to the QED one.
- To extract it, remove the QED Coulomb effect first.



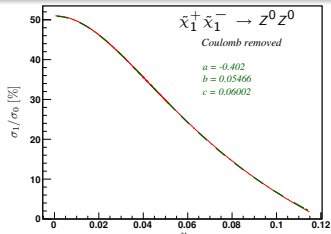
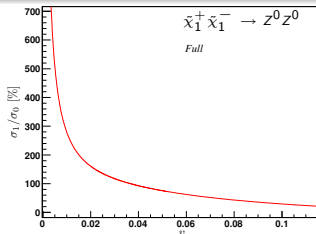
- $S_{nr} = X_{nr}/(1 - e^{-X_{nr}})$ $X_{nr} = 2\pi\alpha Q_i Q_j/v$
- $S_{1L} = \frac{\pi\alpha}{v} \times \sigma_0 Q_i Q_j$

EXTRACTING THE ONE-LOOP SOMMERFELD EFFECT

- The EW Sommerfeld effect is expected to be cut-off, as opposed to the QED one.
- To extract it, remove the QED Coulomb effect first.
- Then, as behavior expected to be cut-off, fit with,

$$\sigma_1/\sigma_0 = a + \frac{b}{\sqrt{v^2 + c^2}}$$

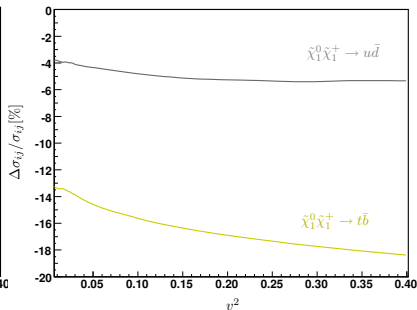
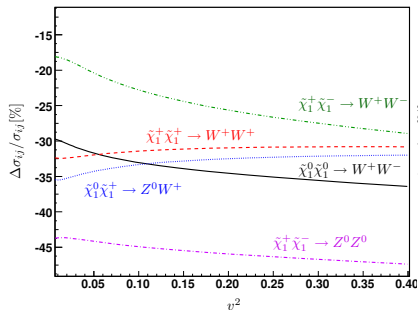
where c is supposed to be the cut-off, of order $M_W/m_{\tilde{\chi}_1^0}$.



- Large corrections but $<$ QED Sommerfeld for $v \rightarrow 0$.

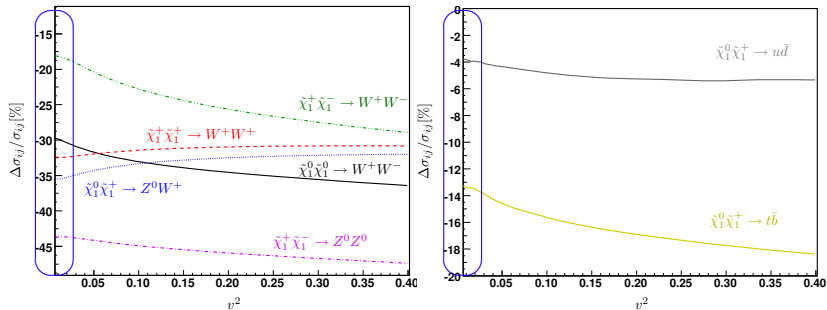
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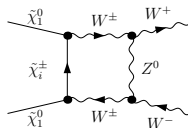


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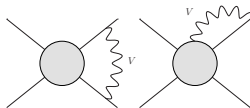


- $v \rightarrow 1$: Large negative corrections of Sudakov type.



- Originate from vertex and box diagrams involving virtual bosons.
- General form of one-loop Sudakov corrections

$$\alpha \left[\underbrace{C_2 \ln^2 \left(\frac{s}{M_V^2} \right)}_{\text{LL}} + \underbrace{C_1 \ln^1 \left(\frac{s}{M_V^2} \right)}_{\text{NLL}} + C_0 \right] + \mathcal{O} \left(\frac{M_V^2}{s} \right) \quad V = \gamma, W^\pm, Z^0$$



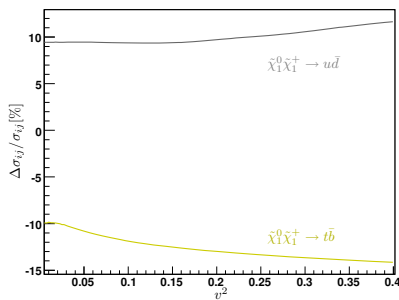
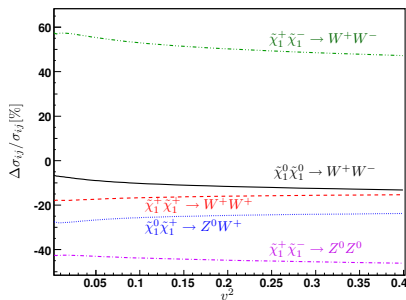
- The $\ln(s/M_V^2)$ represents **mass singularities** and originate from **soft** and **collinear** regions.
- For **QED** corrections always present ($M_\gamma \rightarrow 0$), for **EW** ones when $s \gg M_{W,Z}^2$.

- Originate from vertex and box diagrams involving virtual bosons.
- General form of one-loop Sudakov corrections

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- For **QED** corrections always present ($M_\gamma \rightarrow 0$), for **EW** ones when $s \gg M_{W,Z}^2$.
- Dependency on M_γ **unphysical** \Rightarrow removed by adding **real emission** as stated by the Bloch-Nordsieck theorem [Bloch,Nordsieck(1937)].
- For **EW** corrections, $M_{W,Z}$ **physical** and retained in the calculation.
- Adding real emission can **counterbalance virtual** corrections.



- Still large correction for some processes (-45% corrections for $\tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow Z^0 Z^0$).

RELIC DENSITY

	Tree	$A_{\tau\tau}$	$+ Z^0$ brems	$+ \text{Coul resum}$	EW Som removed
$\Omega_\chi h^2$	0.0993	0.104	0.0934	0.0934	0.103
$\frac{\delta\Omega_\chi h^2}{\Omega_\chi h^2}$		+4.7%	-5.9%	-5.9%	+3.7%

- Taking into account Z^0 real emission or not has an important effect.
- The QED Coulomb one-loop effect is enough, and doesn't change the result.
- One-loop EW Sommerfeld still relevant for relic density calculation.
- Large corrections for individual processes but final result not so affected, large compensation between processes.

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ORIGIN OF LARGE REMAINING CORRECTIONS

- **Non-compensation** between **real** and **virtual** corrections.
- Due to **Bloch-Norsieck violations** [Ciafaloni, Ciafaloni, Comelli (2000)] ?
- **Non cancellation** between **real** and **virtual** contributions due to W emission.
- One-loop expansion **not enough** \Rightarrow **Two-loop** effects important ?

- Importance of **radiative corrections** in the relic density calculations.
- Need to **control** them to be able to **extract** informations from it and to **constrain** the underlying **cosmological scenario**.
- For a heavy neutralino scenarios taking into account $2 \rightarrow 3$ processes is **necessary**.
- Large corrections due to soft/collinear logs and Sommerfeld enhancement.
- Study of the **dependency** of the results on the **chargino/neutralino** renormalisation scheme.
- Improve the **interface** with micrOMEGAs .

BACKUP

A WORD ABOUT LOOP INTEGRALS

- Loop **tensor** integrals **reduced** to a basis of **scalar** integrals [Passarino-Veltman (1979)]
- Reduction method rely on a kinematical ingredient : The **Gram Determinant**.
- For $2 \rightarrow 2$ processes, **Gram determinant** vanishes when relative velocity $v \rightarrow 0$
- In this case reduction method **inefficient** \Rightarrow **different** approach

A WORD ABOUT LOOP INTEGRALS

- Loop **tensor** integrals **reduced** to a basis of **scalar** integrals [Passarino-Veltman (1979)]
- Reduction method rely on a kinematical ingredient : The **Gram Determinant**.
- For $2 \rightarrow 2$ processes, **Gram determinant** vanishes when relative velocity $v \rightarrow 0$
- In this case reduction method **inefficient** \Rightarrow **different** approach
- **Segmentation** has been used to study the **analytical** and numerical behaviour for $v \rightarrow 0$ [Boudjema-Semenov-Temes (2005)].

$$\frac{1}{D_0 D_1 D_2 D_3} = \left(\frac{1}{D_0 D_1 D_2} - \alpha \frac{1}{D_0 D_2 D_3} - \beta \frac{1}{D_0 D_1 D_3} + (\alpha + \beta - 1) \frac{1}{D_1 D_2 D_3} \right) \times$$

$$\frac{1}{A + 2\ell \cdot (s_3 - \alpha s_1 - \beta s_2)}$$

$$A = (s_3^2 - M_3^2) - \alpha(s_1^2 - M_1^2) - \beta(s_2^2 - M_2^2) - (\alpha + \beta - 1)M_0^2.$$

$$D_i = (\ell + s_i)^2 - M_i^2, \quad s_i = \sum_{j=1}^i p_j$$

Relevant mostly for **indirect detection** : $\chi\chi \rightarrow W^+W^-, \gamma\gamma \dots$ in our galaxy ($v \simeq 10^{-3}c$).

- Segmentation has been used to study the analytical and numerical behaviour for $v \rightarrow 0$.

$$\frac{1}{D_0 D_1 D_2 D_3} = \left(\frac{1}{D_0 D_1 D_2} - \alpha \frac{1}{D_0 D_2 D_3} - \beta \frac{1}{D_0 D_1 D_3} + (\alpha + \beta - 1) \frac{1}{D_1 D_2 D_3} \right) \times \frac{1}{A + 2\ell \cdot (s_3 - \alpha s_1 - \beta s_2)}$$

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$$D_i = (\ell + s_i)^2 - M_i^2, \quad s_i = \sum_{j=1}^i p_j$$

- For any graph if $\det G(s_1, s_2, s_3) \simeq 0$, construct all 3 sub-determinants $\det G(s_1, s_2)$ and take the couple s_i, s_j (as independant basis) that corresponds to $\text{Max}|\det G(s_i, s_j)|$. Then

$$s_3 = \alpha s_1 + \beta s_2 + \varepsilon_T \quad \text{with } s_1 \cdot \varepsilon_T = s_2 \cdot \varepsilon_T = 0$$

$$\alpha = \frac{s_2^2(s_3 \cdot s_1) - (s_1 \cdot s_2)(s_2 \cdot s_3)}{\det G(s_1, s_2)}, \quad \beta = \alpha(s_1 \leftrightarrow s_2)$$

$$\det G(s_1, s_2, s_3) = \varepsilon_T^2 \det G(s_1, s_2)$$

Linear gauge fixing

$$\begin{aligned}
 \mathcal{L}_{GF} = & -\frac{1}{\xi_W} \left| \partial_\mu W^{\mu+} + i\xi_W \frac{g}{2} v G^+ \right|^2 \\
 & -\frac{1}{2\xi_Z} \left(\partial_\mu Z^\mu + \xi_Z \frac{g}{2c_W} v G^0 \right)^2 \\
 & -\frac{1}{2\xi_A} (\partial_\mu A^\mu)^2
 \end{aligned}$$

$$\Gamma^{\nu\nu} = \frac{-i}{q^2 - M_V^2 + i\epsilon} \left[g_{\mu\nu} + (\xi_V - 1) \frac{q_\mu q_\nu}{q^2 - \xi_V M_V^2} \right]$$

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$$\xi_{W,Z,A} = 1 \text{ (Feynman gauge)}$$

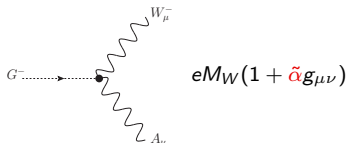
Non-Linear gauge fixing

$$\begin{aligned}
\mathcal{L}_{GF} = & -\frac{1}{\xi_W} |(\partial_\mu - ie\tilde{\alpha}A_\mu - ig_{c_W}\tilde{\beta}Z_\mu)W^\mu + \\
& + i\xi_W \frac{g}{2} (\nu + \tilde{\delta}h^0 + \tilde{\omega}H^0 + i\tilde{\kappa}G^0 + i\tilde{\rho}A^0)G^+|^2 \\
& - \frac{1}{2\xi_Z} (\partial_\mu Z^\mu + \xi_Z \frac{g}{2c_W} (\nu + \tilde{\epsilon}h^0 + \tilde{\gamma}_H^0)G^0)^2 \\
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Non-Linear gauge fixing

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$\xi_{W,Z,A} = 1$ (Feynman gauge)

- Gauge parameter dependence in gauge/Goldstone/ghost vertices.
- No "unphysical" threshold, no higher rank tensor.

SUDAKOV VIRTUAL+REAL CORRECTIONS : ABELIAN EXAMPLE

- However adding real emission of EW gauge boson can counterbalance virtual effects.
- Abelian $Z' \rightarrow \bar{\nu}\nu + Z^0$ (of mass \sqrt{s}) as an example (in the limit $s \gg M_Z^2$) :

$$\Gamma_{\nu\bar{\nu}}^V = - \Gamma_{\nu\bar{\nu}}^0 \frac{\alpha_Z}{4\pi} \left[2 \left(\ln^2 \left(\frac{m_Z^2}{s} \right) + 3 \ln \left(\frac{m_Z^2}{s} \right) \right) - \frac{2\pi^2}{3} + 7 \right]$$

$$\Gamma_{\nu\bar{\nu}}^R = + \Gamma_{\nu\bar{\nu}}^0 \frac{\alpha_Z}{4\pi} \left[2 \left(\ln^2 \left(\frac{m_Z^2}{s} \right) + 3 \ln \left(\frac{m_Z^2}{s} \right) \right) - \frac{2\pi^2}{3} + 10 \right]$$

- Complete compensation between virtual and real logarithmic corrections.

- Factorisation and universality : $\mathcal{M}_1^{Sud} = \mathcal{M}_0 \times \text{Sudakov form factor}$
- For our heavy-wino case Sudakov corrections important ($M_W^2/m_{\tilde{\chi}_1^0}^2 = 2.10^{-3}$).
- $2 \rightarrow 3$ to be taken into account for relic density.
- Real emission of a Z^0 boson added.