## Supersymmetric gauged U(1)B×U(1)L Model

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## **1. Introduction**

- We discuss supersymmetric model with  $U(1)_B$  and  $U(1)_L$  gauge symmetries.
- Baryon symmetry and Lepton symmetry forbid the operator to cause proton decay. However, they must be broken to realize Baryon asymmetry.
- Our scenario is ...

The both symmetries are broken around TeV scale.

(We do not discuss U(1)L sector in this talk.)

T.R. Dulaney, P. F. Perez, and M.B. Wise suggested the non-SUSY model with U(1)<sub>B</sub> × U(1)<sub>L</sub>. (1002.1754;1005.0617[hep-ph])
 They discussed how to realize Baryon asymmetry, and found that the extension of U(1)<sub>B</sub> symmetry naturally gives a cold dark matter (CDM) candidate.

- In the experiments concerned with DM, there are many signals which we can expect to relate to DM physics. The direct searches (DAMA, CoGeNT etc.) suggest light DM, and the indirect (PAMELA etc.) heavy DM.
- The extension of MSSM to U(1)<sup>B</sup> and U(1)<sup>L</sup>, we discuss here, provides several DM candidates. Our model provides DM physics and experiments with interesting observations.

Furthermore, this model suggests a lot of interesting aspects, such as higgs physics, flavor physics and so on.

In this talk, I introduce DM candidates and focus on ~7GeV CDM, discussed in 1007.1005 [hep-ph] by Hooper, Collar, Hall and McKinsey . Then we discuss several experimental constraints.

## 2. U(1)B ×U(1)L Model

In MSSM, Baryon symmetry and Lepton symmetry are good symmetries at classical level.

$$U(1)_{B}: Q^{i} \to e^{\frac{ig_{B}}{3}}Q^{i}, U^{i} \to e^{\frac{-ig_{B}}{3}}U^{i}$$
$$U(1)_{L}: L^{i} \to e^{ig_{L}}L^{i}, E^{i} \to e^{-ig_{L}}E^{i}$$

However, these symmetries are anomalous,

$$SU(2)^{2}U(1)_{B} = \frac{3}{2}, \quad U(1)_{Y}^{2}U(1)_{B} = -\frac{3}{2}$$
$$SU(2)^{2}U(1)_{L} = \frac{3}{2}, \quad U(1)_{Y}^{2}U(1)_{L} = -\frac{3}{2}$$

where right-handed neutrino, Ni, are added.

Extra chiral superfields must be added to built gauged B and L model.

### field contents

 $SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_B \times U(1)_L$ 

Field	Charges	Fields	Charges	
$Q^i$	(3, 2, 1/6; 1/3, 0)	$Q^{'}$	$(\bar{3}, 2, -1/6; -1, 0)$	
$U^i$	$(\bar{3}, 1, -2/3; -1/3, 0)$	$U^{\prime}$	(3, 1, 2/3; 1, 0)	Extra Quarks (chirality is
$D^i$	$(\bar{3}, 1, +1/3; -1/3, 0)$	$D^{\prime}$	(3, 1, -1/3; 1, 0)	opposite.)
$L^i$	(1, 2, -1/2; 0, 1)	$L'_{\downarrow}$	(1, 2, 1/2; 0, -3)	
$E^i$	(1, 1, 1; 0, -1)	$E'_{.}$	(1, 1, -1; 0, 3)	
$N^i$	(1, 1, 0; 0, -1)	$N^{\prime}$	(1, 1, 0; 0, 3)	
$H_u$	(1, 2, 1/2; 0, 0)	$H_d$	(1, 2, -1/2; 0, 0)	
$X_B$	(1, 1, 0; 2/3, 0)	$\overline{X_B}$	(1, 1, 0; -2/3, 0)	
$S_L$	(1, 1, 0; 0, 2)	$\overline{S_L}$	(1, 1, 0; 0, -2)	They break
$S_B$	$(1, 1, 0; n_B, 0)$	$\overline{S_B}$	$(1, 1, 0; -n_B, 0)$	U(1)L and U(1)в.
$X_L$	$(1, 1, 0; n_L, 0)$	$\overline{X_L}$	$(1, 1, 0; -n_L, 0)$	
$\overline{n_B \neq \pm \frac{2}{3}, n_L \neq \pm 2}$ to forbid couplings between SM particles and $S_B(\overline{S}_B), X_L(\overline{X}_L)$ .				

## **Superpotential for extra superfields**

The difference is only chirality. Extra quark and lepton mass terms are

In order to avoid stable charged particles and generate neutrino masses, the couplings between extra B and L symmetric superfields and MSSM fields are given by

$$\lambda_{Qi}X_BQ'Q^i + \lambda_{Ui}\overline{X_B}U'U^i + \lambda_{Di}\overline{X_B}D'D^i$$
$$\lambda_{Ei}\overline{S_L}E'E_i + \lambda_iS_LL'L_i + \lambda_{ij}N^iN^jS_L + \lambda_{Ni}N_i\overline{S_L}N'$$

Assumptions to avoid large FCNC are

$$\begin{array}{l} \left\langle X_{B} \right\rangle = \left\langle \overline{X}_{B} \right\rangle = 0 \\ \left\langle S_{L} \right\rangle, \left\langle \overline{S}_{L} \right\rangle \neq 0, \quad \lambda_{Ei}, \lambda_{i} \approx 0 \end{array}$$

### 3. CDM in the U(1)B ×U(1)L Model

#### The symmetries of cold dark matter (CDM)

1) The B and L charges of SB and SL are nB=2k/3 (k=2,3,...) and nL=2k (k=2,3,...),

$$U(1)_{B} \text{ breaking: } \langle S_{B} \rangle, \langle \overline{S}_{B} \rangle \neq 0 \longrightarrow Z_{2}^{B}$$

$$U(1)_{L} \text{ breaking: } \langle S_{L} \rangle, \langle \overline{S}_{L} \rangle \neq 0 \longrightarrow Z_{2}^{L}$$

$$R\text{-parity } (-1)^{3B+L+2j}$$

***R*-parity** even: 
$$q, X_B, S_B, S_L$$
  
odd:  $\tilde{q}, \tilde{X}_B, \tilde{S}_B, \tilde{S}_L$ 

2) Global U(1) symmetries can be assigned, because of

$$\langle X_B \rangle = \langle \overline{X}_B \rangle = 0, \langle X_L \rangle = \langle \overline{X}_L \rangle = 0.$$

$$U(1)_{XB}: X_B \to e^{i\alpha} X_B, Q' \to e^{-i\alpha} Q', U' \to e^{-i\alpha} U'$$

$$U(1)_{XL}: \quad X_L \to e^{i\alpha} X_L, \overline{X}_L \to e^{-i\alpha} \overline{X}_L$$

#### five candidates for DM



 3 particles of them can be stable, because of 3 global symmetries, U(1)xB, U(1)xL, and R-parity.







2 of 5 decay to 2 CDM.

- $\succ$  q' limits the masses of CDM.
- $\succ$  q' is charged under U(1)xB, so at least one of  $X_B^l$  and  $\tilde{X}_B$  must be smaller than q'.
- $\succ$  R-parity of q' is even, so it cannot decay to only  $\widetilde{X}_{R}$  and quarks.

<u>I introduce the scenario that  $\tilde{X}_{B}$  is the lightest.</u> We assume that SUSY particles, such as squarks and sleptons, are very heavy, around 1 TeV.

### Example



## <u>4. ~ 7 GeV Dark Matter</u>

• D. Hooper, J.I. Collar, J. Hall and D. McKinsey (1007.1005 [hep-ph]) suggest

$$m_{DM} \cong 7 GeV, \ \sigma_{SI} \cong 2 \times 10^{-40} cm^2.$$

Direct detection vs Relic density due to vector current coupling
 <u>Direct detection</u>

U(1)B charged particles scatters with Nuclei through ZB boson (squark'),



Roughly,  $\frac{M_{Z_B}}{g_B} \sim O(TeV)$ . It must explain relic density,  $\Omega h_{CDM}^2 \cong 0.11$ .

# **<u>Relic density</u>** $\frac{dn_x}{dt} + 3Hn_x = -\langle \sigma v \rangle (n_x^2 - n_{eq}^2)$

As well known, the Boltzmann equation is approximatelly estimated as

$$\Omega h^2 \cong 0.1 \times \left(\frac{1pb}{\langle \sigma v \rangle}\right).$$

Annihilation cross section of  $\,\widetilde{X}_{\scriptscriptstyle B}\,$  through ZB boson is

$$\sigma v \cong \frac{A_{an} \left( m_{DM}, M_{Z_B} \right) Q_X^2}{\pi} \left( \frac{g_B}{M_{Z_B}} \right)^4 \left( m_{DM}^2 \sum_q Q_q^2 + \cdots \right)$$
Boson case is p-wave, no S-wave.

#### **Direct scattering leads**

$$\Omega h^{2} \cong O(10) \times \left( \frac{BQ_{N}^{2}}{A_{an} \sum Q_{q}^{2}} \right)$$
  
for  $m_{DM} \cong 7 GeV.$ 

B = 1, for dirac fermion and boson

 $A_{an}(m_{DM}, M_{Z_B})$  must be large to explain the relic density, around 0.11.

#### We use resonance enhancement.



Very small mass and coupling are required,

$$\alpha_B \sim 2 \times 10^{-5}, M_{Z_B} \sim 2m_X.$$

## **5.** Constraints

- Constraints on very light *ZB* boson mass have been discussed.
- For example, the hadoronic decay width of Z boson require  $\alpha_B < 0.2$ . Upsiron decay to hadrons gives the strongest constraint.

(PRL74,3122(1995) by C.D.Carone and H. Murayama, PLB443, 352(1998) by A.Aranda and C.D.Carone. )

### Our model has enough small U(1)<sup>B</sup> gauge coupling.

- On the other hand, kinematic mixing must be enough small to avoid conflicts with experiments.  $Tr(BY) \neq 0$
- Other stable particles also contribute to the relic density. There are a lot of arguments about heavy DM and scalar DM,
   Scalar DM with yukawa:PLB670,37(2008) by J.L.Feng, J.Kumar, L.E.Strigari
   MSSM DM: PR267,195(1996), by G.Jungman, M.Kamionkowski, K.Griest, etc.

Our model predicts very light scalar.

 $(S_B, \overline{S}_B)$ , which correspond to Higgs, give this lightest scalar.  $m_{S1}^2 \le M_{Z_R}^2 \cos^2 2\beta_B + \alpha_B M_{Z_R}^2 \cos^2 2\beta_B f(m_{\tilde{t}_1}^2, m_{\tilde{t}_2}^2)$ 

m<sub>s2</sub>(GeV)

 $O(10^{-3})GeV^2$ 

 They do not couple with SM 500 particles at tree level. 450 The one-loop corrections are also strongly suppressed by small  $U(1)_B$  coupling.

ex)





- We discussed supersymmetric model with U(1)<sub>B</sub> × U(1)<sub>L</sub> gauge symmetries.
- This model provides a lot of interesting topics, such as Flavor physics, Higgs physics, dark matter, and so on. I introduced DM physics related to direct detections.
- There are several candidates for CDM and several scenarios. In any cases, 3 stable particles appear.
- In order to explain ~7GeV CDM of CoGeNT/DAMA, we considered very light ZB mass and the resonance effect. U(1)B gauge coupling is also very small,

$$\alpha_{B} \sim 2 \times 10^{-5}, M_{Z_{B}} \sim 2m_{DM}.$$

• I introduced several constraints, especially the predicted light scalar.