

# Massive Quark Form Factors

Loops and Legs in Quantum Field Theory, 2022 (LL2022)

Kay Schönwald | Ettal, April 26, 2022

TTP KARLSRUHE

[based on: Fael, Lange, Schönwald, Steinhauser [arxiv:2202.05276](https://arxiv.org/abs/2202.05276)

Egner, Fael, Lange, Schönwald, Steinhauser [arxiv:2203.11231](https://arxiv.org/abs/2203.11231)]



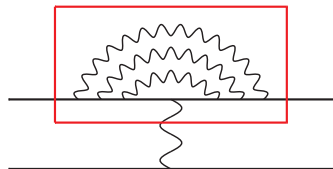
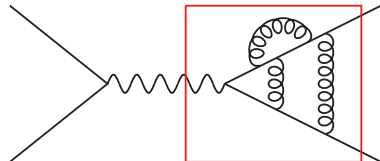
Institute for Theoretical Particle Physics



TRR 257 - Particle Physics Phenomenology  
after the Higgs Discovery

- 1 Motivation
- 2 Definition and Previous Calculations
- 3 Technical Details
- 4 Results
- 5 Conclusions and Outlook

- Form factors are basic building blocks for many physical observables:
  - $t\bar{t}$  production at hadron and  $e^+e^-$  colliders
  - $\mu e$  scattering
  - Higgs production and decay
  - ...
- Form factors exhibit an universal infrared behavior.



$$X(q) \rightarrow Q(q_1) + Q(q_2)$$

$$q_1^2 = q_2^2 = m^2, \quad q^2 = s = \hat{s} \cdot m^2$$

vector :

$$j_\mu^v = \bar{\psi} \gamma_\mu \psi \quad \Gamma_\mu^v = F_1^v(s) \gamma_\mu - \frac{i}{2m} F_2^v(s) \sigma_{\mu\nu} q^\nu$$

axial-vector :

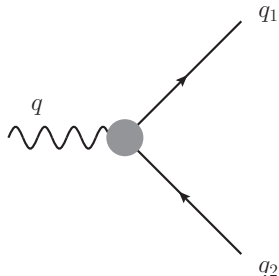
$$j_\mu^a = \bar{\psi} \gamma_\mu \gamma_5 \psi \quad \Gamma_\mu^a = F_1^a(s) \gamma_\mu \gamma_5 - \frac{1}{2m} F_2^a(s) q_\mu \gamma_5$$

scalar :

$$j^s = m \bar{\psi} \psi \quad \Gamma^s = m F^s(s)$$

pseudo-scalar :

$$j^p = im \bar{\psi} \gamma_5 \psi \quad \Gamma^p = im F^p(s) \gamma_5$$



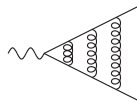
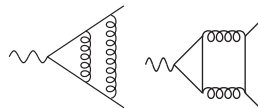
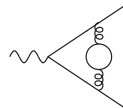
## NNLO

$F_I^{(2)}$  fermionic corrections [Hoang, Teubner '97]

$F_I^{(2)}$  [Bernreuther, Bonciani, Gehrmann, Heinesch, Leineweber, Mastrolia, Remiddi '04-'06]

+ $\mathcal{O}(\epsilon)$  [Gluza, Mitov, Moch, Riemann '09]

+ $\mathcal{O}(\epsilon^2)$  [Ahmed, Henn, Steinhauser '17; Ablinger, Behring, Blümlein, Falcioni, Freitas, Marquard, Rana, Schneider '17]



## NNNLO

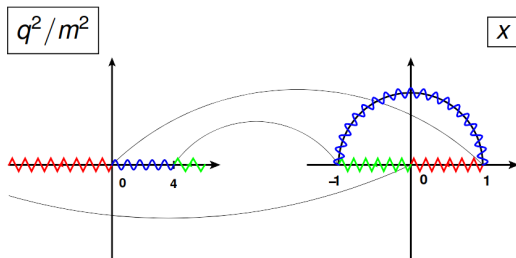
$F_I^{(3)}$  large- $N_c$  [Henn, Smirnov, Smirnov, Steinhauser '16-'18; Ablinger, Marquard, Rana, Schneider '18]

$n_l$  [Lee, Smirnov, Smirnov, Steinhauser '18]

$n_h$  (partially) [Blümlein, Marquard, Rana, Schneider '19]

this talk: full (numerical) results for non-singlet diagrams at NNNLO

$$q^2 = s = -\frac{(1-x)^2}{x}$$

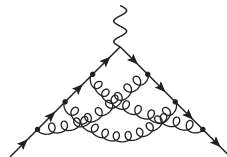
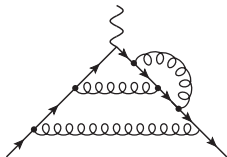
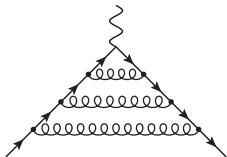


- The large- $N_c$  and  $n_l$  contributions at NNNLO can be written as iterated integrals over the letters:

$$\frac{1}{x}, \frac{1}{1+x}, \frac{1}{1-x}, \frac{1}{1-x+x^2}, \frac{x}{1-x+x^2}$$

- The  $n_h$  terms already contain structures which go beyond iterated integrals.

⇒ We aim at the full solution through analytic series expansions and numerical matching.



- Generate diagrams with QGRAF. [Nogueira '93]
- Use FORM [Ruij, Ueda, Vermaseren '17] for Lorentz Dirac and color algebra. [Ritbergen, Schellekens, Vermaseren '98]
- Map the output to predefined integral families with q2e/exp. [Harlander, Seidensticker, Steinhauser '97-'99]
- Reduce the scalar integrals to masters with Kira. [Klappert, Lange, Maierhöfer, Usovitsch, Uwer '17,20]
  - We ensure a good basis where denominators factorize in  $\epsilon$  and  $\hat{s}$  with `ImproveMasters.m`. [Smirnov, Smirnov '20]
- Establish differential equations in variable  $\hat{s}$  using LiteRed. [Lee '12,'14]

	non-singlet
diagrams	271
families	34
masters	422

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{l=-3}^{\infty} \sum_{j=0}^{l_{\max}} c_j^{(n)} \epsilon^l \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_j^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using *Kira with FireFly*.  
(Klappert, Klein, Lange, 10.21)
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.



- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using Kira with FireFly.  
[Klappert, Klein, Lange '19,20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using Kira with FireFly. [Klappert, Klein, Lange '19,20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
  - Solve the linear system in terms of a small number of boundary constants using *Kira with FireFly*.  
[Klappert, Klein, Lange '19,20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using **Kira with FireFly**.  
[Klappert, Klein, Lange '19,'20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using **Kira with FireFly**.  
[Klappert, Klein, Lange '19,'20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using Kira with FireFly. [Klappert, Klein, Lange '19,'20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
  - Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
  - Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

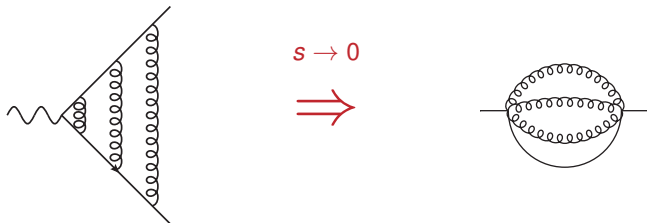
- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using Kira with FireFly. [Klappert, Klein, Lange '19,'20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

- Establish a system of differential equations for the master integrals in the variable  $\hat{s}$ .
- Compute an expansion around  $\hat{s} = 0$  by:
  - Inserting an ansatz for the master integrals into the differential equation.

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Compare coefficients in  $\epsilon$  and  $\hat{s}$  to establish a linear system of equations for the  $c_{ij}^{(n)}$ .
- Solve the linear system in terms of a small number of boundary constants using **Kira with FireFly**.  
[Klappert, Klein, Lange '19,'20]
- Compute boundary values for  $\hat{s} = 0$  and obtain an analytic expansion.
- Build a general expansion around a new point, e.g.  $\hat{s} = \hat{s}_0$ , by modifying the ansatz and repeating the steps above.
- Match both expansions numerically at a point where both expansions converge, e.g.  $\hat{s}_0/2$ .
- Repeat the procedure for the next point.

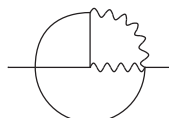




- For  $s = 0$  the master integrals reduce to 3-loop on-shell propagators:
  - These integrals are well studied in the literature. [Laporta, Remiddi '96; Meinel, Ritbergen '00; Lee, Smirnov '10]
- The reduction introduces high inverse powers in  $\epsilon$ , which require some integrals up to weight 9.
- We calculate the needed terms with `SummerTime.m` [Lee, Mingulov '15] and `PSLQ` [Ferguson, Bailey '92].

# Calculation of Boundary Conditions

E.g. extension of  $G_{66}$  (given up to and including  $\mathcal{O}(\epsilon^3)$  in [Lee, Smirnov '10]):



$$\begin{aligned}
 &= \dots + \epsilon^4 \left( -4704s_6 - 9120s_{7a} - 9120s_{7b} - 547s_{8a} + 9120s_6 \ln(2) + 28 \ln^4(2) + \frac{112 \ln^5(2)}{3} - \frac{808}{45} \ln^6(2) \right. \\
 &\quad \left. - \frac{347}{9} \ln^8(2) + 672\text{Li}_4\left(\frac{1}{2}\right) - \frac{5552}{3} \ln^4(2)\text{Li}_4\left(\frac{1}{2}\right) - 22208\text{Li}_4\left(\frac{1}{2}\right)^2 - 4480\text{Li}_5\left(\frac{1}{2}\right) - 12928\text{Li}_6\left(\frac{1}{2}\right) + \dots \right) \\
 &\quad + \epsilon^5 \left( 14400s_6 - \frac{377568s_{7a}}{7} - \frac{93984s_{7b}}{7} - 2735s_{8a} + 7572912s_{9a} - 3804464s_{9b} - \frac{5092568s_{9c}}{3} - 136256s_{9d} \right. \\
 &\quad \left. + 681280s_{9e} + 272512s_{9f} + \frac{377568}{7} s_6 \ln(2) - \frac{32465121}{20} s_{8a} \ln(2) - 10185136s_{8b} \ln(2) + 136256s_{7b} \ln^2(2) + \dots \right) \\
 &\quad + \mathcal{O}(\epsilon^6)
 \end{aligned}$$

- Special points:

$s = 0$	$s = 4m^2$	$s = \pm\infty$
$x = 1$	$x = -1$	$x = 0$
static limit	2-particle threshold	high energy limit

- Every expansion point needs a different ansatz:

$$M_n(\epsilon, \hat{s} = 0) = \sum_{i=-3}^{\infty} \sum_{j=0}^{j_{\max}} c_{ij}^{(n)} \epsilon^i \hat{s}^j$$

- Special points:

$s = 0$	$s = 4m^2$	$s = \pm\infty$
$x = 1$	$x = -1$	$x = 0$
static limit	2-particle threshold	high energy limit

- Every expansion point needs a different ansatz:

$$M_n(\epsilon, \hat{s} = 4) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[ \sqrt{4 - \hat{s}} \right]^j \ln^k \left( \sqrt{4 - \hat{s}} \right)$$

- Special points:

$s = 0$	$s = 4m^2$	$s = \pm\infty$
$x = 1$	$x = -1$	$x = 0$
static limit	2-particle threshold	high energy limit

- Every expansion point needs a different ansatz:

$$M_n(\epsilon, \hat{s} \rightarrow \pm\infty) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+6} c_{ijk}^{(n)} \epsilon^i \hat{s}^{-j} \ln^k(\hat{s})$$

- Special points:

$s = 0$	$s = 4m^2$	$s = \pm\infty$	$s = 16m^2$
$x = 1$	$x = -1$	$x = 0$	$x = 4\sqrt{3} - 7$
static limit	2-particle threshold	high energy limit	4-particle threshold

- Every expansion point needs a different ansatz:

$$M_n(\epsilon, \hat{s} = 16) = \sum_{i=-3}^{\infty} \sum_{j=-s_{\min}}^{j_{\max}} \sum_{k=0}^{i+3} c_{ijk}^{(n)} \epsilon^i \left[ \sqrt{16 - \hat{s}} \right]^j \ln^k \left( \sqrt{16 - \hat{s}} \right)$$

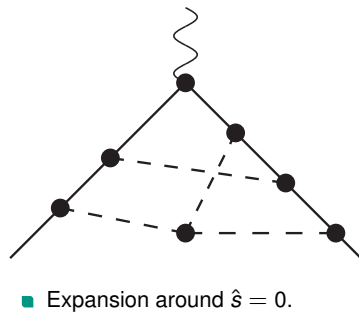
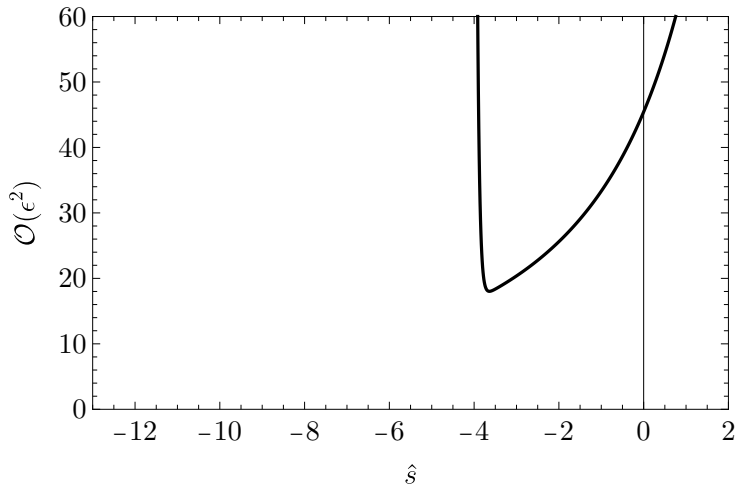
- Special points:

$s = 0$	$s = 4m^2$	$s = \pm\infty$	$s = 16m^2$
$x = 1$	$x = -1$	$x = 0$	$x = 4\sqrt{3} - 7$
static limit	2-particle threshold	high energy limit	4-particle threshold

- Every expansion point needs a different ansatz.
- We construct expansions with  $j_{\max} = 50$  around:

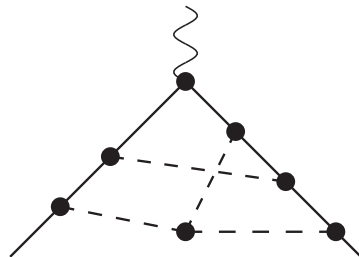
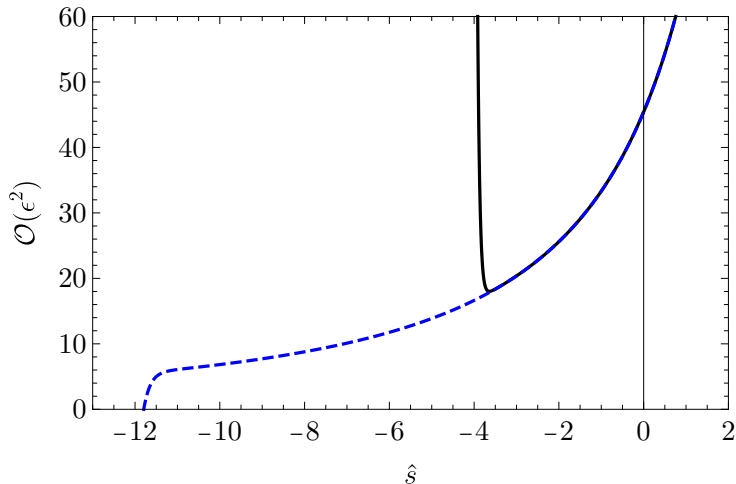
$$\hat{s} = \{ -\infty, -32, -28, -24, -16, -12, -8, -4, 0, 1, 2, 5/2, 3, 7/2, 4, 9/2, 5, 6, 7, 8, 10, 12, 14, 15, 16, 17, 19, 22, 28, 40 \}$$

# Example



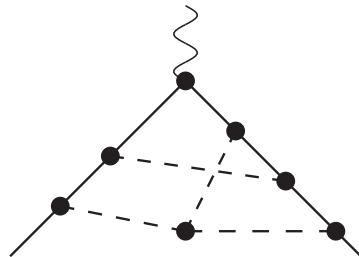
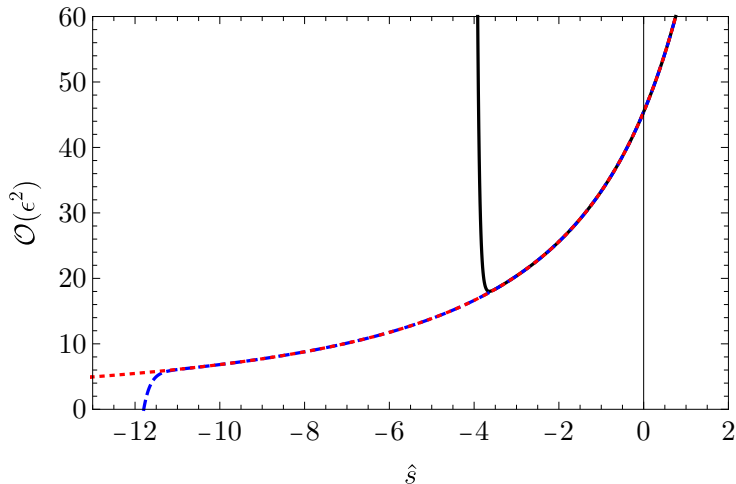


# Example



- Expansion around  $\hat{s} = 0$ .
- Expansion around  $\hat{s} = -4$ ,  
matched at  $\hat{s} = -2$ .

# Example



- Expansion around  $\hat{s} = 0$ .
- Expansion around  $\hat{s} = -4$ ,  
matched at  $\hat{s} = -2$ .
- Expansion around  $\hat{s} = -8$ ,  
matched at  $\hat{s} = -6$ .

There are other approaches based on expansions:

- `SolveCoupledSystems.m` [Blümlein, Schneider '17]
- `DESS.m` [Lee, Smirnov, Smirnov '18]
- `DiffExp.m` [Hidding '20]
- ...

Our approach ...

- ... does not require a special form of differential equation.
- ... provides approximation in whole kinematic range.
- ... is applied to physical quantity. [Fael, Lange, KS, Steinhauser '21]

## UV renormalization

- On-shell renormalization of mass  $Z_m^{\text{OS}}$ , wave function  $Z_2^{\text{OS}}$ , and (if needed) the currents.  
[Chetyrkin, Steinhauser '99; Melnikov, Ritbergen '00]

## IR subtraction

- Structure of the infrared poles is given by the cusp anomalous dimension  $\Gamma_{\text{cusp}}$ .  
[Grozin, Henn, Korchemski, Marquard '14]
- Define finite form factors  $F = Z_{\text{IR}} F^{\text{finite}}$  with the UV renormalized form factor  $F$  and

$$Z_{\text{IR}} = 1 - \frac{\alpha_s}{\pi} \frac{1}{2\epsilon} \Gamma_{\text{cusp}}^{(1)} - \left(\frac{\alpha_s}{\pi}\right)^2 \left( \frac{\dots}{\epsilon^2} + \frac{1}{4\epsilon} \Gamma_{\text{cusp}}^{(2)} \right) - \left(\frac{\alpha_s}{\pi}\right)^3 \left( \frac{\dots}{\epsilon^3} + \frac{\dots}{\epsilon^2} + \frac{1}{6\epsilon} \Gamma_{\text{cusp}}^{(3)} \right)$$

- $\Gamma_{\text{cusp}} = \Gamma_{\text{cusp}}(x)$  depends on kinematics.
- $\Gamma_{\text{cusp}}$  is universal for all currents.

Analytic expansion for  $\hat{s} = 0$ :

$$\begin{aligned}
 F_1^V(\hat{s} = 0) = & \left\{ C_F^3 \left( -15a_4 - \frac{17\pi^2\zeta_3}{24} - \frac{18367\zeta_3}{1728} + \frac{25\zeta_5}{8} - \frac{5l_2^4}{8} - \frac{19}{40}\pi^2 l_2^2 + \frac{4957\pi^2 l_2}{720} + \frac{3037\pi^4}{25920} \right. \right. \\
 & - \frac{24463\pi^2}{7776} + \frac{13135}{20736} \left. \right) + C_A C_F^2 \left( \frac{19a_4}{2} - \frac{\pi^2\zeta_3}{9} + \frac{17725\zeta_3}{3456} - \frac{55\zeta_5}{32} + \frac{19l_2^4}{48} - \frac{97}{720}\pi^2 l_2^2 \right. \\
 & + \frac{29\pi^2 l_2}{240} - \frac{347\pi^4}{17280} - \frac{4829\pi^2}{10368} + \frac{707}{288} \left. \right) + C_A^2 C_F \left( -a_4 + \frac{7\pi^2\zeta_3}{96} + \frac{4045\zeta_3}{5184} - \frac{5\zeta_5}{64} - \frac{l_2^4}{24} \right. \\
 & \left. \left. + \frac{67}{360}\pi^2 l_2^2 - \frac{5131\pi^2 l_2}{2880} + \frac{67\pi^4}{8640} + \frac{172285\pi^2}{186624} - \frac{7876}{2187} \right) \right\} \hat{s} + \text{fermionic corrections} + \mathcal{O}(\hat{s}^2)
 \end{aligned}$$

with  $l_2 = \ln(2)$ ,  $a_4 = \text{Li}_4(1/2)$  and  $C_A = 3$ ,  $C_F = 4/3$  for QCD.

- The expansions for all currents are available.
- We have calculated the expansion up to  $\mathcal{O}(s^{67})$ .

Analytic expansion for  $\hat{s} = 0$ :

$$\begin{aligned}
 F_1^V(\hat{s} = 0) = & \left\{ C_F^3 \left( -15a_4 - \frac{17\pi^2\zeta_3}{24} - \frac{18367\zeta_3}{1728} + \frac{25\zeta_5}{8} - \frac{5l_2^4}{8} - \frac{19}{40}\pi^2 l_2^2 + \frac{4957\pi^2 l_2}{720} + \frac{3037\pi^4}{25920} \right. \right. \\
 & - \frac{24463\pi^2}{7776} + \frac{13135}{20736} \left. \right) + C_A C_F^2 \left( \frac{19a_4}{2} - \frac{\pi^2\zeta_3}{9} + \frac{17725\zeta_3}{3456} - \frac{55\zeta_5}{32} + \frac{19l_2^4}{48} - \frac{97}{720}\pi^2 l_2^2 \right. \\
 & + \frac{29\pi^2 l_2}{240} - \frac{347\pi^4}{17280} - \frac{4829\pi^2}{10368} + \frac{707}{288} \left. \right) + C_A^2 C_F \left( -a_4 + \frac{7\pi^2\zeta_3}{96} + \frac{4045\zeta_3}{5184} - \frac{5\zeta_5}{64} - \frac{l_2^4}{24} \right. \\
 & \left. \left. + \frac{67}{360}\pi^2 l_2^2 - \frac{5131\pi^2 l_2}{2880} + \frac{67\pi^4}{8640} + \frac{172285\pi^2}{186624} - \frac{7876}{2187} \right) \right\} \hat{s} + \text{fermionic corrections} + \mathcal{O}(\hat{s}^2)
 \end{aligned}$$

with  $l_2 = \ln(2)$ ,  $a_4 = \text{Li}_4(1/2)$  and  $C_A = 3$ ,  $C_F = 4/3$  for QCD.

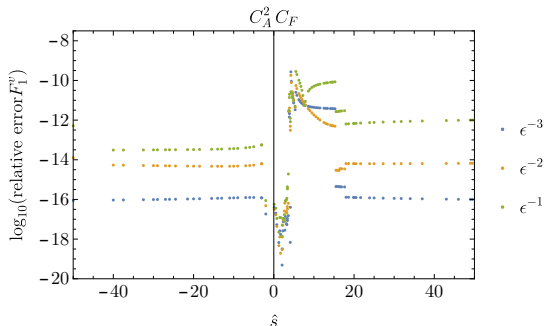
- The expansions for all currents are available.
- We have calculated the expansion up to  $\mathcal{O}(s^{67})$ .

- Except for  $s = 0$  the results of the expansions are not analytic.
  - We can use the pole cancellation to estimate the precision.
- ⇒ We find at least 8 significant digits, although some regions are much more precise.

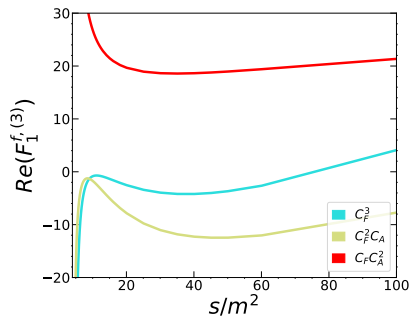
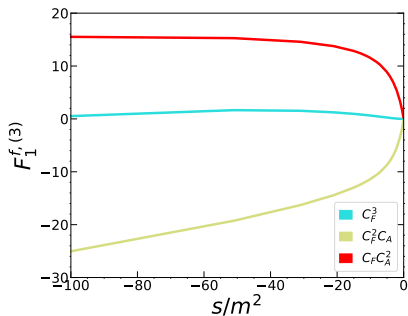
- To estimate the number of significant digits we use:

$$\log_{10} \left( \left| \frac{\text{expansion} - \text{analytic}}{\text{analytic}} \right| \right)$$

- The analytic expressions for the poles are expressed by Harmonic Polylogarithms which can be evaluated with `ginac`. [Vollinga, Weinzierl '05]



- Except for  $s = 0$  the results of the expansions are not analytic.
  - We can use the pole cancellation to estimate the precision.
- ⇒ We find at least 8 significant digits, although some regions are much more precise.





# Results – High Energy Limit

- For  $s \rightarrow \infty$  there is the prediction: [Liu, Penin, Zerf '17]

$$F_1^{f,(3)} = -\frac{C_F^3}{384} l_s^6 - \frac{m^2}{s} \left( \frac{C_F^3}{240} - \frac{C_F^2 C_A}{960} - \frac{C_F C_A^2}{1920} \right) l_s^6 + \dots, \quad \text{with } l_s = \ln \left( \frac{m^2}{-s} \right)$$

- We obtain:

$$\begin{aligned} F_1^{f,(3)} \Big|_{s \rightarrow -\infty} &= 4.7318 C_F^3 - 20.762 C_F^2 C_A + 8.3501 C_F C_A^2 + \left[ 3.4586 C_F^3 - 4.0082 C_F^2 C_A - 6.3561 C_F C_A^2 \right] l_s \\ &+ \left[ 1.4025 C_F^3 + 0.51078 C_F^2 C_A - 2.2488 C_F C_A^2 \right] l_s^2 + \left[ 0.062184 C_F^3 + 0.90267 C_F^2 C_A - 0.42778 C_F C_A^2 \right] l_s^3 \\ &+ \left[ -0.075860 C_F^3 + 0.20814 C_F^2 C_A - 0.035011 C_F C_A^2 \right] l_s^4 + \left[ -0.023438 C_F^3 + 0.019097 C_F^2 C_A \right] l_s^5 \\ &+ \left[ -0.0026042 C_F^3 \right] l_s^6 - \left\{ -92.918 C_F^3 + 123.65 C_F^2 C_A - 47.821 C_F C_A^2 + \left[ -10.381 C_F^3 + 2.3223 C_F^2 C_A \right. \right. \\ &+ 17.305 C_F C_A^2 \left. \right] l_s + \left[ 4.9856 C_F^3 - 19.097 C_F^2 C_A + 8.0183 C_F C_A^2 \right] l_s^2 + \left[ 3.0499 C_F^3 - 6.8519 C_F^2 C_A + 1.9149 C_F C_A^2 \right] l_s^3 \\ &+ \left[ 0.67172 C_F^3 - 0.91213 C_F^2 C_A + 0.24069 C_F C_A^2 \right] l_s^4 + \left[ 0.13229 C_F^3 - 0.051389 C_F^2 C_A + 0.0043403 C_F C_A^2 \right] l_s^5 \\ &+ \left. \left[ 0.0041667 C_F^3 - 0.0010417 C_F^2 C_A - 0.00052083 C_F C_A^2 \right] l_s^6 \right\} \frac{m^2}{s} + \mathcal{O} \left( \frac{m^4}{s^2} \right) + \text{fermionic contributions}, \end{aligned}$$

# Results – High Energy Limit

- For  $s \rightarrow \infty$  there is the prediction: [Liu, Penin, Zerf '17]

$$F_1^{f,(3)} = -\frac{C_F^3}{384} l_s^6 - \frac{m^2}{s} \left( \frac{C_F^3}{240} - \frac{C_F^2 C_A}{960} - \frac{C_F C_A^2}{1920} \right) l_s^6 + \dots, \quad \text{with } l_s = \ln \left( \frac{m^2}{-s} \right)$$

- We obtain:

$$\begin{aligned}
 F_1^{f,(3)} \Big|_{s \rightarrow -\infty} &= 4.7318 C_F^3 - 20.762 C_F^2 C_A + 8.3501 C_F C_A^2 + \left[ 3.4586 C_F^3 - 4.0082 C_F^2 C_A - 6.3561 C_F C_A^2 \right] l_s \\
 &+ \left[ 1.4025 C_F^3 + 0.51078 C_F^2 C_A - 2.2488 C_F C_A^2 \right] l_s^2 + \left[ 0.062184 C_F^3 + 0.90267 C_F^2 C_A - 0.42778 C_F C_A^2 \right] l_s^3 \\
 &+ \left[ -0.075860 C_F^3 + 0.20814 C_F^2 C_A - 0.035011 C_F C_A^2 \right] l_s^4 + \left[ -0.023438 C_F^3 + 0.019097 C_F^2 C_A \right] l_s^5 \\
 &+ \left[ -0.0026042 C_F^3 \right] l_s^6 - \left\{ -92.918 C_F^3 + 123.65 C_F^2 C_A - 47.821 C_F C_A^2 + \left[ -10.381 C_F^3 + 2.3223 C_F^2 C_A \right. \right. \\
 &+ 17.305 C_F C_A^2 \left. \right] l_s + \left[ 4.9856 C_F^3 - 19.097 C_F^2 C_A + 8.0183 C_F C_A^2 \right] l_s^2 + \left[ 3.0499 C_F^3 - 6.8519 C_F^2 C_A + 1.9149 C_F C_A^2 \right] l_s^3 \\
 &+ \left[ 0.67172 C_F^3 - 0.91213 C_F^2 C_A + 0.24069 C_F C_A^2 \right] l_s^4 + \left[ 0.13229 C_F^3 - 0.051389 C_F^2 C_A + 0.0043403 C_F C_A^2 \right] l_s^5 \\
 &+ \left. \left[ 0.0041667 C_F^3 - 0.0010417 C_F^2 C_A - 0.00052083 C_F C_A^2 \right] l_s^6 \right\} \frac{m^2}{s} + \mathcal{O} \left( \frac{m^4}{s^2} \right) + \text{fermionic contributions},
 \end{aligned}$$

- For  $Q\bar{Q}$  production close to threshold it is advantageous to calculate the cross section in non-relativistic QCD (NRQCD).
- The naive expansion around the threshold of the form factor

$$x = \sqrt{4 - \hat{s}} = 0$$

defines the matching coefficients between QCD and NRQCD. [Pineda '11]

- At threshold the momenta can have different scalings: [Beneke, Smirnov '98]
  - hard (h):  $k_0 \sim m, k_i \sim m$
  - potential (p):  $k_0 \sim x^2 \cdot m, k_i \sim x \cdot m$
  - soft (s):  $k_0 \sim x \cdot m, k_i \sim x \cdot m$
  - ultrasoft (u):  $k_0 \sim x^2 \cdot m, k_i \sim x^2 \cdot m$

- We can obtain the matching coefficient by modifying the ansatz:

$$\begin{array}{ll}
 l_a \sim x^{-0\epsilon} \cdot \text{Taylor expansion} & (h - h - h) \\
 l_b \sim x^{-2\epsilon} \cdot \text{Taylor expansion} & (h - h - p), (h - h - s), \dots \\
 l_c \sim x^{-4\epsilon} \cdot \text{Taylor expansion} & (h - h - u), (h - p - p), \dots \\
 l_d \sim x^{-6\epsilon} \cdot \text{Taylor expansion} & (h - p - u), (h - s - u), \dots \\
 l_e \sim x^{-8\epsilon} \cdot \text{Taylor expansion} & (h - u - u), (u - p - p), \dots \\
 l_f \sim x^{-10\epsilon} \cdot \text{Taylor expansion} & (u - u - p), (u - u - s) \\
 l_g \sim x^{-12\epsilon} \cdot \text{Taylor expansion} & (u - u - u)
 \end{array}$$

- For  $Q\bar{Q}$  production close to threshold it is advantageous to calculate the cross section in non-relativistic QCD (NRQCD).
- The naive expansion around the threshold of the form factor

$$x = \sqrt{4 - \hat{s}} = 0$$

defines the matching coefficients between QCD and NRQCD. [Pineda '11]

- At threshold the momenta can have different scalings: [Beneke, Smirnov '98]

- hard (h):  $k_0 \sim m, k_i \sim m$
- potential (p):  $k_0 \sim x^2 \cdot m, k_i \sim x \cdot m$
- soft (s):  $k_0 \sim x \cdot m, k_i \sim x \cdot m$
- ultrasoft (u):  $k_0 \sim x^2 \cdot m, k_i \sim x^2 \cdot m$

- We can obtain the matching coefficient by modifying the ansatz:

$$\begin{array}{ll}
 l_a \sim x^{-0\epsilon} \cdot \text{Taylor expansion} & (h - h - h) & l_e \sim x^{-8\epsilon} \cdot \text{Taylor expansion} & (h - u - u), (u - p - p), \dots \\
 l_b \sim x^{-2\epsilon} \cdot \text{Taylor expansion} & (h - h - p), (h - h - s), \dots & l_f \sim x^{-10\epsilon} \cdot \text{Taylor expansion} & (u - u - p), (u - u - s) \\
 l_c \sim x^{-4\epsilon} \cdot \text{Taylor expansion} & (h - h - u), (h - p - p), \dots & l_g \sim x^{-12\epsilon} \cdot \text{Taylor expansion} & (u - u - u) \\
 l_d \sim x^{-6\epsilon} \cdot \text{Taylor expansion} & (h - p - u), (h - s - u), \dots & & 
 \end{array}$$

- The previous calculation relied heavily on sector decomposition and numerical integration with FIESTA. [Marquard, Piclum, Seidel, Steinhauser '14]
- We can improve the precision significantly:

$$c_V^{(3)} = C_F^3 C_{FFF} + C_F C_A^2 C_{FFA} + C_F C_A^2 C_{FAA} + \text{fermionic and singlet contributions}$$

$$c_{FFF}^V = 36.55(0.53) \quad \rightarrow \quad 36.49486246$$

$$c_{FFA}^V = -188.10(0.83) \quad \rightarrow \quad -188.0778417$$

$$c_{FAA}^V = -97.81(0.38) \quad \rightarrow \quad -97.73497327$$

- We calculated the matching coefficients for all four currents.

## Conclusions

- We have calculated the non-singlet contributions to the massive quark form factors at NNNLO.
- We applied a semianalytic method by constructing series expansions and numerical matching.
- We can reproduce known results in the literature, e.g.
  - the large  $N_c$  limit.
  - expansion terms in the static, high energy and threshold expansion.
- We estimate the precision to 8 significant digits over the whole real axis.
- The method is promising to tackle other one-scale problems.

## Outlook

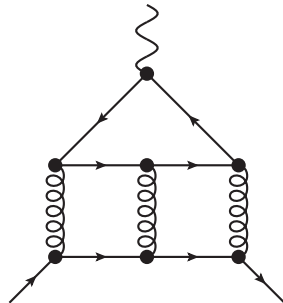
- Calculate the contributions of the singlet diagrams.
- Calculate the singlet contributions to the NRQCD matching coefficient.

## Conclusions

- We have calculated the non-singlet contributions to the massive quark form factors at NNNLO.
- We applied a semianalytic method by constructing series expansions and numerical matching.
- We can reproduce known results in the literature, e.g.
  - the large  $N_c$  limit.
  - expansion terms in the static, high energy and threshold expansion.
- We estimate the precision to 8 significant digits over the whole real axis.
- The method is promising to tackle other one-scale problems.

## Outlook

- Calculate the contributions of the singlet diagrams.
- Calculate the singlet contributions to the NRQCD matching coefficient.



# Backup



- The radius of convergence is at most the distance to the closest singularity.
- We can extend the radius of convergence by changing to a new expansion variable.
- If we want to expand around the point  $x_k$  with the closest singularities at  $x_{k-1}$  and  $x_{k+1}$ , we can use:

$$y_k = \frac{(x - x_k)(x_{k+1} - x_{k-1})}{(x - x_{k+1})(x_{k-1} - x_k) + (x - x_{k-1})(x_{k+1} - x_k)}$$

- The variable change maps  $\{x_{k-1}, x_k, x_{k+1}\} \rightarrow \{-1, 0, 1\}$ .

- Close to threshold it is interesting to consider:

$$\sigma(e^+e^- \rightarrow Q\bar{Q}) = \sigma_0\beta \underbrace{\left( |F_1^V + F_2^V|^2 + \frac{|(1-\beta^2)F_1^V + F_2^V|^2}{2(1-\beta^2)} \right)}_{=3/2 \Delta}$$

with  $\beta = \sqrt{1 - 4m^2/s}$ .

- Real radiation is suppressed by  $\beta^3$ .
- We find (with  $l_{2\beta} = \ln(2\beta)$ ):

$$\begin{aligned} \Delta^{(3)} = & C_F^3 \left[ -\frac{32.470}{\beta^2} + \frac{1}{\beta} (14.998 - 32.470 l_{2\beta}) \right] + C_A^2 C_F \frac{1}{\beta} [16.586 l_{2\beta}^2 - 22.572 l_{2\beta} + 42.936] \\ & + C_A C_F^2 \left[ \frac{1}{\beta^2} (-29.764 l_{2\beta} - 7.770339) + \frac{1}{\beta} (-12.516 l_{2\beta} - 11.435) \right] \\ & + \mathcal{O}(\beta^0) + \text{fermionic contributions} \end{aligned}$$