Deep Learning meets Physics



Prof. Dr. Martin Erdmann, RWTH Aachen University, 14-Sep-2021



Deep Learning spectacular success

Image recognition challenge









quail



flamingo



cock



Persian cat Siamese cat



tabby

lynx

partridge

ImageNet: 1.2 million images in 1000 categories

ruffed grouse



O. Russakovsky et al, arXiv:1409.0575; K. He, X. Zhang, S. Ren, J. Sunar, arXiv:1512.03385 WMW Jie Hu, Li Shen (Oxford), Gang Sun, 2017

Generative Modeling



https://thispersondoesnotexist.com

Plan for today

- 1. What deep learning is precisely: neural networks
- 2. Deep learning & data symmetries
- 3. Autonomous model building
- 4. Experiments' operation reality and network insight

Data analysis \rightarrow deep learning



variable = *feature*



McCulloch, W.S., Pitts, W.: Bulletin of Mathematical Biophysics (1943) 5: 115. Frank Rosenblatt, Principles of Neurodynamics: Perceptrons and the Theory of Brain Mechanisms. Spartan Books, Washington DC, 1961

Neural Network Operations





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Neural Network Training



Data set

 $\{x_i, y_i\} \quad i = 1, ..., N$

Define model

 $y_m(x) = W x + b$

- Define objective function (=loss, cost) $\mathcal{L}(W, b) = \frac{1}{N} \sum_{i=1}^{N} [y_m(x_i) - y_i]^2$
- Train model by optimizing the parameters
 - $(\widehat{W}, \widehat{b}) = \arg \min \mathcal{L}(W, b)$

('supervised')



Automated parameterization of arbitrary function





7 hidden layers 200 nodes each **ReLU** activation function

original function (black symbols): fair description after 2800 training steps (purple)

- $\vec{x} \in \mathbb{R}^n \rightarrow \vec{z} \in \mathbb{R}^m$
- Function: training is million-parameter fit

Reality: function working in multi-dimensions

Deep Learning Progress

Concepts

- Fully connected
- Convolutional
- Graph
- Recurrent
- Lorentz Boost Network
- Autoencoder
- Adversarial
- Reinforcement
- Invertible



Improved set of tools



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Computing

- Graphics Processing Unit (GPU)
- Software Libraries
 - TensorFlow
 - keras...



Deep learning & data symmetries

...looking for better ways than 1 pixel = 1 network input node

Convolutional network to analyse image-like data



Convolutional network to identify electron neutrinos



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A. Aurisano et al., JINST 11 (2016) P09001



od	<pre>v_eefficiency (same purity)</pre>	
ists thm	35%	
learning I network	49%	

Convolution: *Classic* versus *Graph* network







T. Bister, M. Erdmann, J. Glombitza, N. Langner, J. Schulte, M. Wirtz, arXiv:2003.13038

Graph convolutions to detect cosmic magnetic fields





World's largest Calorimeter for Cosmic Rays





air Water Cherenkov detectors 55 km

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Pierre Auger Observatory



Cosmic ray arrival directions by physicist or network



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M. Erdmann, Jonas Glombitza, David Walz, 10.1016/j.astropartphys.2017.10.006

Deep Neural Network

No physics education

- No explicit information about
- locations of detectors
- speed of light

Needs data with true target θ

Deep Neural Network learns physics from data within 3h

M. Erdmann, Jonas Glombitza, David Walz, 10.1016/j.astropartphys.2017.10.006 Jonas Glombitza for the Pierre Auger Collaboration arXiv:2101.02946

Recurrent network to characterize signal traces



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Deep Neural Network

Emiel Hoogeboom, Jorn W.T. Peters, Taco S. Cohen, Max Welling, arXiv:1803.02108 Jonas Glombitza for the Pierre Auger Collaboration arXiv:2101.02946

Hexagonal convolutions to symmetrize azimuth $\boldsymbol{\varphi}$



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Autonomous model building

Assign functional target → training data optimize network (`unsupervised')

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)g pervised´)

Autoencoder networks to identify new physics

T. Heimel, G. Kasieczka, T. Plehn, J.M. Thompson, SciPost Phys. 6, 030 (2019)





Ian J. Goodfellow, Jean Pouget-Abadie, Mehdi Mirza, Bing Xu, David Warde-Farley, Sherjil Ozair, Aaron Courville, Yoshua BengioarXiv:1406.2661 M. Erdmann, J. Glombitza, T. Quast, Comput Softw Big Sci (2019) 3: 4

Generative Modeling to simulate particle showers



Invertible Networks to map probability distributions



S. Radev, U. Mertens, A. Voss, L. Ardizzone, U. Köthe, arxiv 2003.06281 J. Schulte, T. Bister, M. Erdmann, RWTH Aachen M. Bellagente, A. Butter, G. Kasieczka, T. Plehn, A. Rousselot, R. Winterhalder, L. Ardizzone, U. Köthe, SciPost Phys. 9, 074 (2020)



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Invertible Networks to map probability distributions





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Invertible Networks to unfold observed data



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Operation reality & network insight

Solved long-standing machine learning problem data ≠ simulation

M. Erdmann, Lukas Geiger, Jonas Glombitza, David Schmidt, Comput Softw Big Sci (2018) 2:4 M. Erdmann, Jonas Glombitza, David Walz, 10.1016/j.astropartphys.2017.10.006

Domain adaption to simulate data-like



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improved

resolution

Simulation: include operation reality for network training



\rightarrow Improved generalization capability of trained neural network

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Jonas Glombitza for the Pierre Auger Collaboration arXiv:2101.02946

Network training with simulated data including defects



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Causality: analysis of network predictions

Measure impact: input $x_i \rightarrow overall prediction$

G. Montavon, W. Samek, K.-R. Müller, Digital Signal Processing 73 (2018) 1





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Messages Deep Learning

- **Physicists** exploit more information from data: choose network according to data symmetries
- Standard network like fitting functions, except network is ultra-flexible physics model
- Advanced concepts assign **functional targets**, training data **autonomously** optimize network
- Tools exist for including experiment's **operation reality** and obtaining **insight** into networks

We ought to prepare for fundamental change to include machines in our daily work

backup

CMS jet flavor tagging



the 1-dim Convolution adds up all properties of a track. In the RNN the summed properties of the tracks are merged.

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Markus Stoye, ACAT2017, CMS DSP-2017-005/013/027



Denoising Gravitational Waves with **Recurrent Denoising Autoencoder**



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Hongyu Shen, Daniel George, E. A. Huerta, Zhizhen Zhao1, arXiv:1711.09919



Self Normalizing Networks

- Batch normalization adds perturbations for training fully connected networks
- Use activation function *selu* which ensures standard normalized output: $\mu = 0, \sigma = 1$
- Initialization:

Gauß with $\mu = 0$, $\sigma = 1/n$ with n=nodes in lower layer

• Alpha-dropout

(insert specified value instead of turning node off)

- Stabilizes the training
- Allows to build very deep networks!



$$\mu_j = \frac{1}{n} \sum_{i=1}^n x_i$$

$$\nu_j = \frac{1}{n} \sum_{i=1}^n (x_i - \mu_j)^2$$

$$(\mu_{i+1}) \qquad (\mu_i)$$

$$\begin{pmatrix} \mu_{j+1} \\ \nu_{j+1} \end{pmatrix} = K \begin{pmatrix} \mu_j \\ \nu_j \end{pmatrix}$$



$$\operatorname{selu}(x) =$$



Requirements: negative & positive values to control the mean Slope < 1 for damping the variance Slope > 1 to rise the variance

result activation function

Cosmic ray: shower maximum



correct for atmosphere





moonless nights

Fluorescence Light Detection





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 $\rho(h) = \rho_o e^{-\frac{\rho_o}{p_o}gh}$

Maximum of shower development directly visible in the camera during 35

Air shower maximum with particle detectors on Earth



electromagnetic component + muons: depending on height of 1st interaction shower particles differ when reaching earth







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Jonas Glombitza, PhD thesis, RWTH Aachen



Cosmic Rays: Energies & Nuclei



M. Erdmann, E. Geiser, Y. Rath, M. Rieger, JINST 14 (2019) P06006

Lorentz Boost Network to recover interaction



M. Erdmann, E. Geiser, Y. Rath, M. Rieger, JINST 14 (2019) P06006 Autonomous engineering of discriminating variables

