



Stefan Liebler

## Correlating neutralino decay modes with neutrino mixing angles

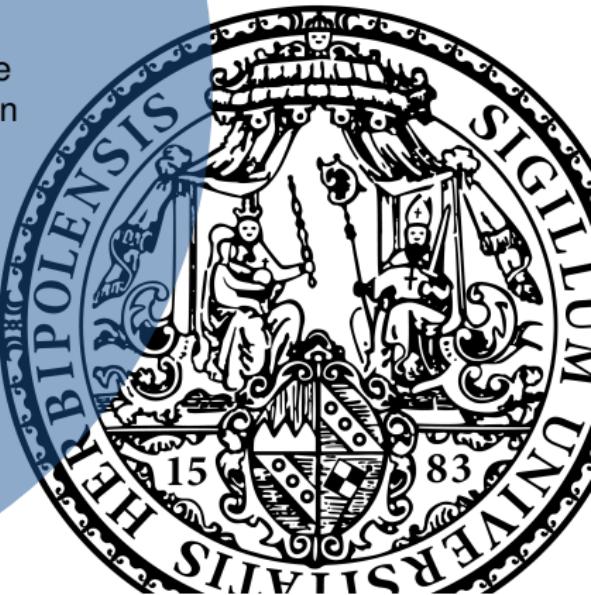
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Neutrino masses and lepton flavour violation

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## Outline

- 1 Supersymmetry with and without  $R$ -parity
- 2 Neutrino physics in models with broken  $R$ -parity
- 3 LSP properties in models with broken  $R$ -parity
- 4 Summary



## The Minimal Supersymmetric Standard Model (MSSM)

### 1. Particle content:

- Each SM fermion has a scalar as superpartner.
- There are two Higgs and two fermionic Higgsino doublets.
- Each gauge boson has a fermionic superpartner, called gaugino.

2. Particle mixing: Particles like the neutral Higgsinos and the neutral gauginos mix to e.g. four Neutralinos:

$$(\psi^0)^T = (\tilde{B}, \tilde{W}_3^0, \tilde{H}_d^0, \tilde{H}_u^0) \quad \longrightarrow \quad (F^0)^T = (\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0)$$

3. Particle interactions: Interactions in supersymmetric theories are described by the superpotential  $W$ , which is in the MSSM given by:

$$W_{\text{MSSM}} = (Y_u)_{ij} \hat{Q}_i \hat{H}_u \hat{u}_j^c + (Y_d)_{ij} \hat{H}_d \hat{Q}_i \hat{d}_j^c + (Y_e)_{ij} \hat{H}_d \hat{L}_i \hat{e}_j^c - \mu \hat{H}_d \hat{H}_u$$

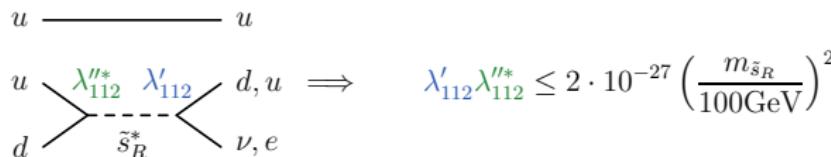
Remark:  $\hat{L}_i = (\hat{\nu}_i, \hat{e}_i)^T$ ,  $\hat{Q}_i = (\hat{u}_i, \hat{d}_i)^T$ ,  $\hat{H}_u = (\hat{H}_u^+, \hat{H}_u^0)^T$ ,  $\hat{H}_d = (\hat{H}_d^0, \hat{H}_d^-)^T$



Similar to the SM, the MSSM does not provide an explanation for neutrino masses. Nevertheless there are more terms  $W_R$ , which are invariant under SUSY- and gauge-transformations:

$$W_R = \frac{1}{2} \lambda_{ijk} \hat{L}_i \hat{L}_j \hat{e}_k^c + \lambda'_{ijk} \hat{L}_i \hat{Q}_j \hat{d}_k^c - \epsilon_i \hat{L}_i \hat{H}_u + \frac{1}{2} \lambda''_{ijk} \hat{u}_i^c \hat{d}_j^c \hat{d}_k^c$$

⇒ Strong restrictions from i.e.  $p \rightarrow \pi^+ \nu / \pi^0 e$ :



↔ All terms in  $W_R$  are forbidden by  $R$ -parity  $P_R = (-1)^{3(B-L)+2s}$ , which is  $P_R = +1$  for SM particles and  $P_R = -1$  for SUSY partners.

Idea:

**But** allowing only *L-violating* terms explains neutrino physics and leads to a phenomenology, which is testable at the LHC.

We will now focus on the phenomenology of models with (effective)  $\epsilon_i$ -terms.



## Generation of a Majorana mass term for $\nu_i$ via the $L$ -violating terms

In the basis

$$(\psi^0)^T = \left( \tilde{B}, \tilde{W}_3^0, \tilde{H}_d^0, \tilde{H}_u^0, \nu_1, \nu_2, \nu_3 \right)$$

one can write  $\mathcal{L}_{\text{neutral}}^{\text{mass}} = -\frac{1}{2} (\psi^0)^T \mathcal{M}_n \psi^0 + h.c.$  with

$$\mathcal{M}_n = \begin{pmatrix} M_n & \textcolor{red}{m} \\ \textcolor{red}{m}^T & 0 \end{pmatrix}.$$

- $M_n$  mixes the 4 heavy states
- $\textcolor{red}{m}$  mixes the heavy states with the neutrinos

This leads to an effective neutrino mass matrix  $m_{\text{eff}}$ , which is at NLO given by

$$(m_{\text{eff}})_{ij} = - \left( \textcolor{red}{m}^T M_n^{-1} \textcolor{red}{m} \right)_{ij} = a \Lambda_i \Lambda_j + b (\textcolor{green}{\Lambda}_i \epsilon_j + \epsilon_i \textcolor{green}{\Lambda}_j) + c \epsilon_i \epsilon_j$$

with  $\textcolor{green}{\Lambda}_i = \mu v_i + v_d \epsilon_i$  and  $\langle \tilde{\nu}_i \rangle = \frac{1}{\sqrt{2}} v_i$



## Remark: Present neutrino data

From neutrino oscillations: mass differences  $\Delta m_{ij}^2 = m_i^2 - m_j^2$  between the different mass eigenstates and mixing angles:

parameter	best fit	$2\sigma$
$\Delta m_{21}^2 [10^{-5} \text{eV}^2]$	$7.59^{+0.23}_{-0.18}$	$7.22 - 8.03$
$ \Delta m_{31}^2  [10^{-3} \text{eV}^2]$	$2.40^{+0.12}_{-0.11}$	$2.18 - 2.64$
$\tan^2 \theta_{12} = \tan^2 \theta_{sol}$	$0.466^{+0.033}_{-0.042}$	$0.41 - 0.56$
$\tan^2 \theta_{23} = \tan^2 \theta_{atm}$	$1.00^{+0.33}_{-0.21}$	$0.64 - 1.70$
$\tan^2 \theta_{13} = \tan^2 \theta_R$	$0.013^{+0.013}_{-0.009}$	$\leq 0.041$

[Schwetz et al., 2010, arXiv:0808.2016]

Example for fitting the neutrino data with the parameters of ( $m_{\text{eff}}$ ):

$$\tan^2 \theta_{atm} \approx \left( \frac{\Lambda_2}{\Lambda_3} \right)^2, \quad \tan^2 \theta_{sol} \approx \left( \frac{\tilde{\epsilon}_1}{\tilde{\epsilon}_2} \right)^2, \quad \tan^2 \theta_R \approx \left( \frac{\Lambda_1}{\sqrt{\Lambda_2^2 + \Lambda_3^2}} \right)^2,$$

⇒ Alignment parameter  $\Lambda_i$  fits atmospheric and  $\epsilon_i$  solar scale.



What features do these  $R$ -parity violating models have?

- Neutrino data is explainable ( $\cancel{R}$  via **L-violating** terms).
- Lightest supersymmetric particle (LSP) is not stable any more.  
 $\Rightarrow$  Decay length of **mm** up to several **km**  $\iff$  Displaced vertices
- Decay modes of LSP are **correlated** with the neutrino mixing angles.
- Special phenomenology in some of these models due to singlets states

$\Rightarrow$  Testable at the LHC!

All the properties above are for example given in the  $\mu\nu$ SSM, which will be discussed in the following:

[D.E. Lopéz-Fogliani and C. Muñoz, 2005]

$$\begin{aligned}
 W_{\mu\nu\text{SSM}} = & (Y_u)_{ij} \hat{Q}_i \hat{H}_u \hat{u}_j^c + (Y_d)_{ij} \hat{H}_d \hat{Q}_i \hat{d}_j^c + (Y_e)_{ij} \hat{H}_d \hat{L}_i \hat{e}_j^c \\
 & - \lambda \hat{\nu}^c \hat{H}_d \hat{H}_u + (Y_\nu)_i \hat{L}_i \hat{H}_u \hat{\nu}^c + \frac{1}{3!} \kappa \hat{\nu}^c \hat{\nu}^c \hat{\nu}^c \\
 \Leftrightarrow & \textcolor{red}{\mu} = \lambda \langle \tilde{\nu}^c \rangle, \quad \epsilon_i = (Y_\nu)_i \langle \tilde{\nu}^c \rangle \text{ after EWSB}
 \end{aligned}$$



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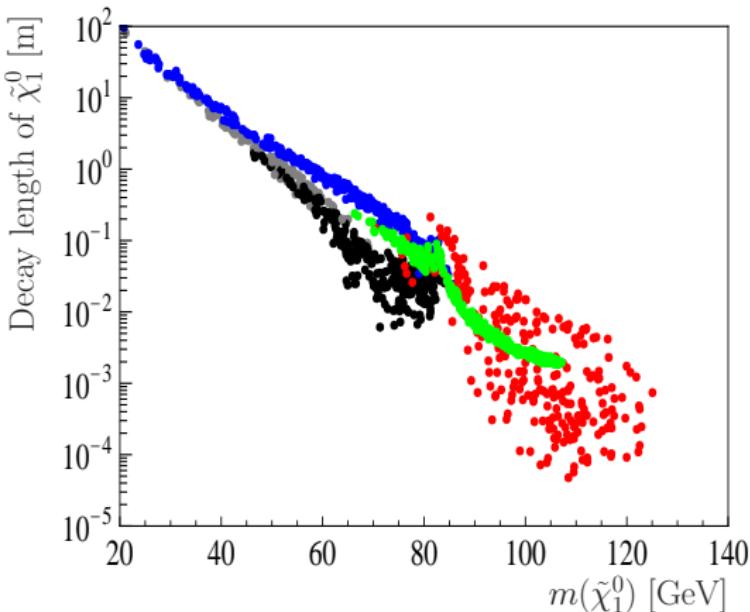
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 \end{aligned}$$

Total decay length for a singlino-type LSP  $\tilde{\chi}_1^0 = \nu^c$ 

- $\lambda \in [0.2, 0.5]$   
 $\kappa \in [0.025, 0.2]$   
 $\mu \in [110, 170]\text{GeV}$
- SPS1a' ( $\nu^c$ )  
SPS1a' (mixture)  
SPS3 ( $\nu^c$ )  
SPS3 (mixture)  
SPS4 (mixture)
- $\{m(S_1^0), m(P_1^0)\} > m(\tilde{\chi}_1^0)$   
by appropriate choice of  
 $T_\kappa \in [-20, -0.05]\text{GeV}$

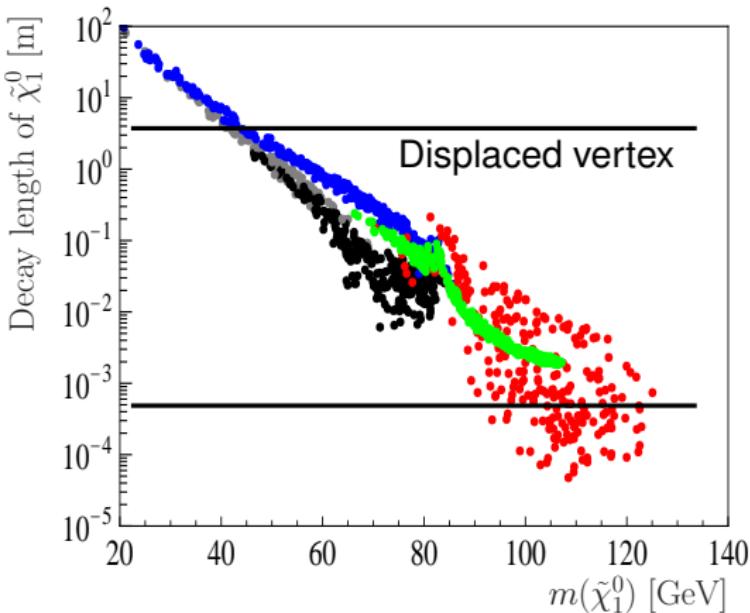
[Bartl, Hirsch, S.L, Porod, Vicente, arXiv:0903.3596]



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Decay modes of the lightest neutralino  $\tilde{\chi}_1^0$  as LSP

Most important decay modes:

decay	$m_{\tilde{\chi}_1^0} < m_W$	$m_W < m_{\tilde{\chi}_1^0} < m_Z$	$m_Z < m_{\tilde{\chi}_1^0}$
$\tilde{\chi}_1^0 \rightarrow Z \nu_i$			•
$\tilde{\chi}_1^0 \rightarrow W^\pm l^\mp$		•	•
$\tilde{\chi}_1^0 \rightarrow S_i^0 \nu_j / P_i^0 \nu_j$	○	○	○
$\tilde{\chi}_1^0 \rightarrow l_i^\pm l_j^\mp \nu_k$	•	•	•
$\tilde{\chi}_1^0 \rightarrow q_i \bar{q}_j l_k$	•	•	•

Also present:  $\tilde{\chi}_1^0 \rightarrow 3\nu$  or  $\tilde{\chi}_1^0 \rightarrow q_i \bar{q}_j \nu_k$ 

Notation:

- $\iff$  Link to neutrino physics
- $\iff$  Masses of scalars and pseudoscalars crucial



Where does the **correlation** come from?

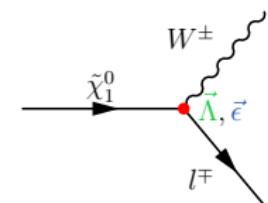
Consider the lightest neutralino  $\tilde{\chi}_1^0 = \tilde{W}_3^0$  as LSP in the  $\mu\nu$ SSM.

Two-body decay: At tree level the left-handed  $W$ - $\tilde{\chi}_1^0$ - $l_i$ -coupling reads:

$$\mathcal{L} = \overline{l_i^-} \gamma^\mu (O_{Li} P_L + O_{Ri} P_R) \tilde{\chi}_1^0 W_\mu^- + h.c.$$

$$O_{Li} \approx \frac{g}{\sqrt{2}} \left[ \frac{g \Lambda_i}{\det_+} N_{12} - \left( \frac{\epsilon_i}{\mu} + \frac{g^2 v_u \Lambda_i}{2\mu \det_+} \right) N_{13} - \sum_{j=1}^5 N_{1j} \xi_{ij} \right]$$

$$\implies \frac{Br(\tilde{\chi}_1^0 \rightarrow W^- \mu^+)}{Br(\tilde{\chi}_1^0 \rightarrow W^- \tau^+)} \propto \left| \frac{O_{L2}}{O_{L3}} \right|^2 = \left( \frac{\Lambda_2}{\Lambda_3} \right)^2 \approx \tan^2 \theta_{atm}$$

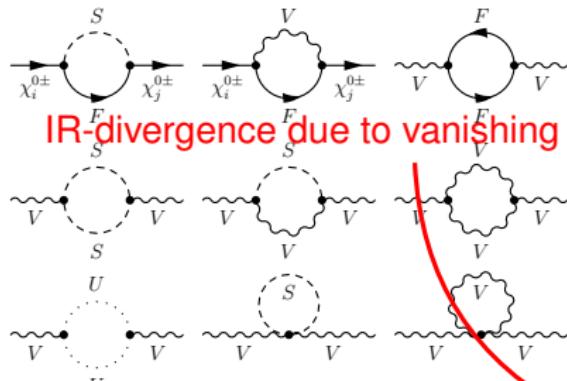


To maintain this correlation after calculating the one-loop Neutralino mass matrix one has to perform a complete NLO correction to the decay  $\tilde{\chi}_1^0 \rightarrow W^\pm l^\mp$ .



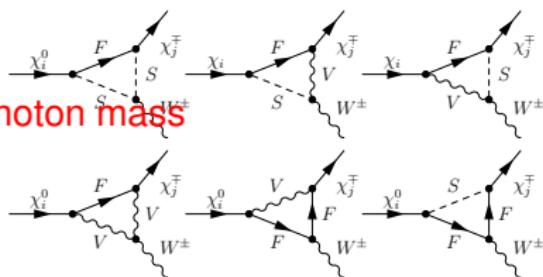
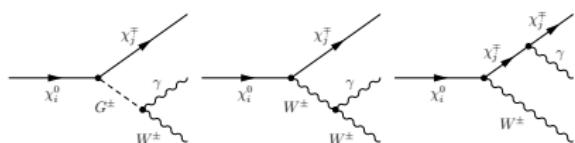
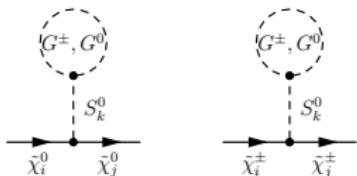
## Electroweak contributions

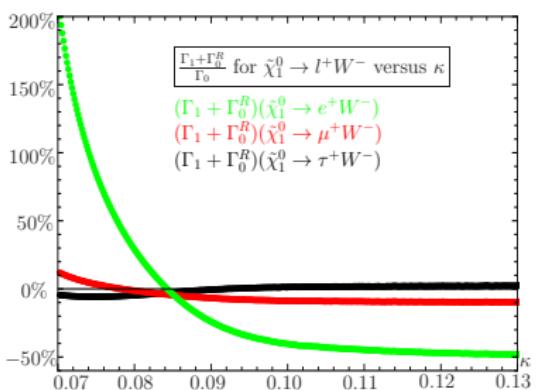
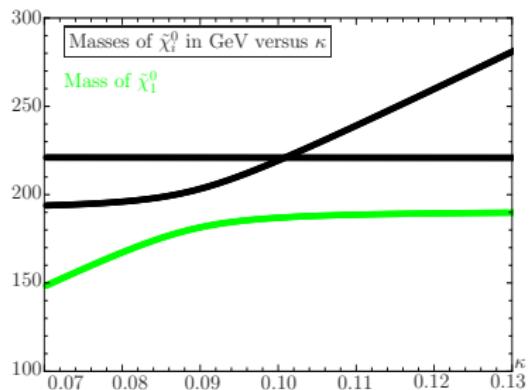
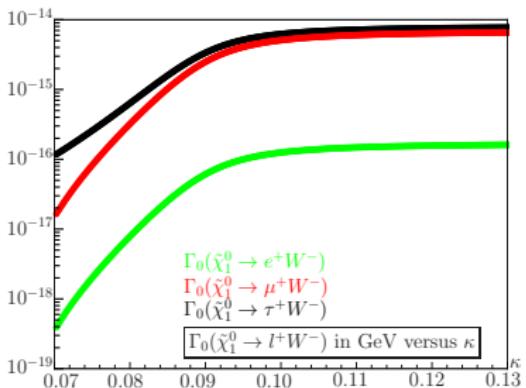
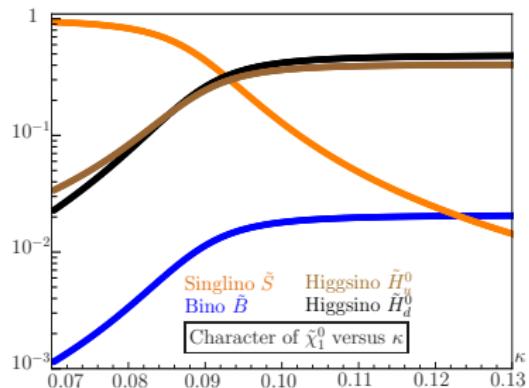
Self energies:



IR-divergence due to vanishing photon mass

Vertex corrections:

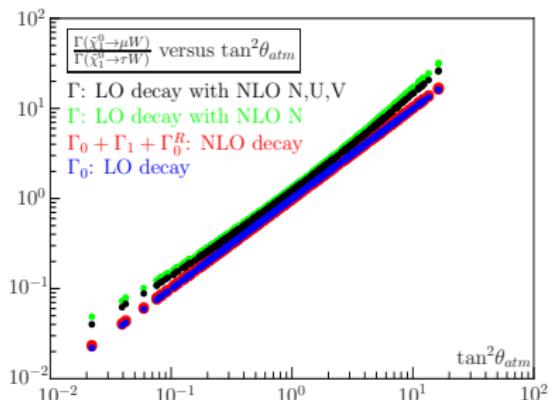
Pinch technique (for  $\xi_V \neq 1$  relevant): → Real photon emission:

Absolute corrections for  $\tilde{\chi}_1^0 \rightarrow l^+ W^-$ 

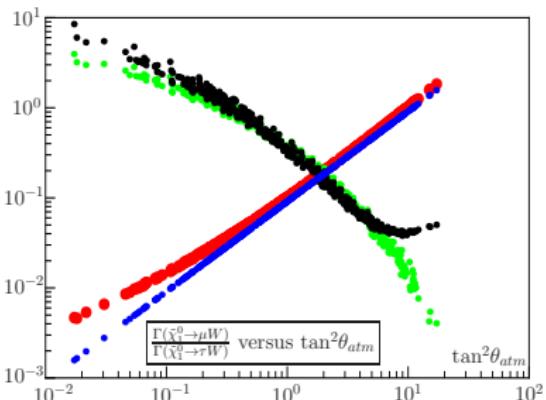


Finally we can compare the ratios of decay widths  $\tilde{\chi}_1^0 \rightarrow l^+ W^-$  with the neutrino mixing angles:

Bino  $\tilde{\chi}_1^0 = \tilde{B}$



Singlino  $\tilde{\chi}_1^0 = \nu^c$

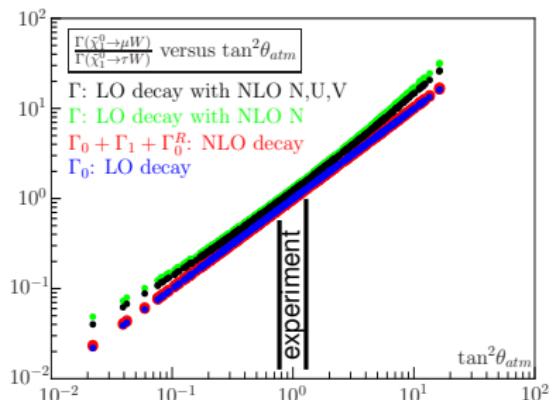


⇒ The full NLO corrections show the behaviour predicted on tree-level!

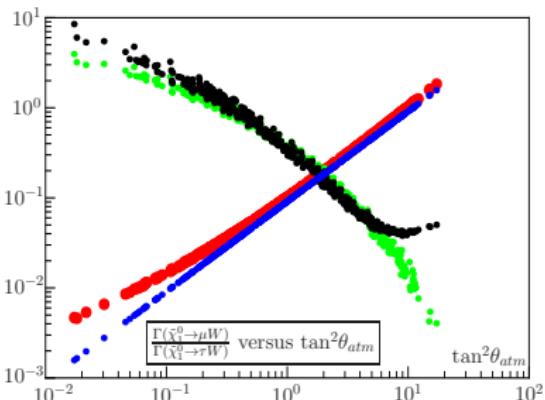


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## Models with broken $R$ -parity via (effective) $\epsilon_i$ -terms in SUSY theories ...

- ... provide an alternative ansatz to explain neutrino masses and mixings.
- ... show interesting correlations between decay modes and neutrino mixing angles. To maintain these tree-level relations a full NLO calculation is necessary. If some singlet neutralinos are light, interesting decay chains at the LHC could occur.
- ... are in complete agreement with experimental bounds on lepton flavor violating decays due to the smallness of *L-violating* parameters.
- ... can show interesting decay chains for e.g. the lightest SUSY Higgs.

**Thank you for your attention!**



## Particle content of the $\mu\nu$ SSM

MSSM + right-handed neutrino superfield  $\widehat{\nu}^c$  with  $L_{\widehat{\nu}^c} = -1$

The new,  $R$ -parity violating terms induce a mixing of the following flavor eigenstates to mass eigenstates:

- Neutralinos  $(\tilde{B}, \tilde{W}_3^0, \tilde{H}_d^0, \tilde{H}_u^0, \nu^c, \nu_1, \nu_2, \nu_3)$   
 $\implies$  5 heavy states including Singlino  $\nu^c$  and three light states
- Scalars/Pseudoscalars  $(H_d^0, H_u^0, \tilde{\nu}^c, \tilde{\nu}_1, \tilde{\nu}_2, \tilde{\nu}_3)$   
 $\implies$  Singlet scalar/pseudoscalar state  $\tilde{\nu}^c$

Similar to BRPV and SRPV:

- Charginos  $(\tilde{W}^-, \tilde{H}_d^-, e, \mu, \tau)$
- Charged Scalars  $(H_d^-, H_u^+, \tilde{e}, \tilde{\mu}, \tilde{\tau}, \tilde{e}^c, \tilde{\mu}^c, \tilde{\tau}^c)$



Definition of the matrix  $\xi = m^T M_n^{-1}$

The complete diagonalization of the neutral fermion mass matrix reads

$$\widehat{\mathcal{M}}_n = \mathcal{N}^* \mathcal{M}_n \mathcal{N}^\dagger \quad \text{with} \quad \mathcal{M}_n = \begin{pmatrix} M_n & m \\ m^T & 0 \end{pmatrix}$$

and can be done approximately with  $\xi = m^T M_n^{-1}$  in the form

$$\mathcal{N}^* = \begin{pmatrix} N^* & 0 \\ 0 & \mathcal{V}^T \end{pmatrix} \begin{pmatrix} 1 - \frac{1}{2}\xi^\dagger \xi & \xi^\dagger \\ -\xi & 1 - \frac{1}{2}\xi \xi^\dagger \end{pmatrix}$$

resulting in

$$\mathcal{N}^* \mathcal{M}_n \mathcal{N}^\dagger \approx \begin{pmatrix} N^* & 0 \\ 0 & \mathcal{V}^T \end{pmatrix} \begin{pmatrix} M_n & 0 \\ 0 & -m^T M_n^{-1} m \end{pmatrix} \begin{pmatrix} N^\dagger & 0 \\ 0 & \mathcal{V} \end{pmatrix} ,$$

where in the  $1-\mu\nu$ SSM the matrix  $\xi$  is given by

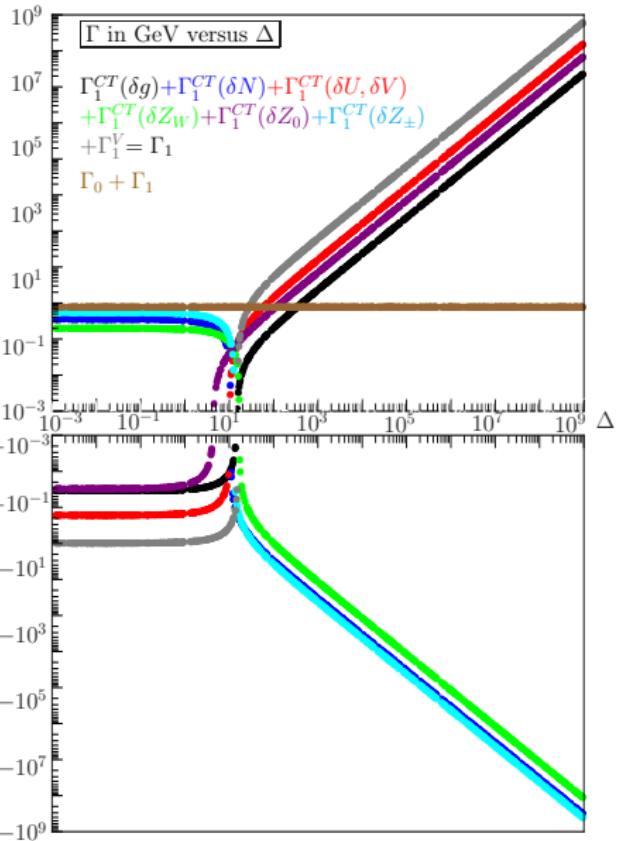
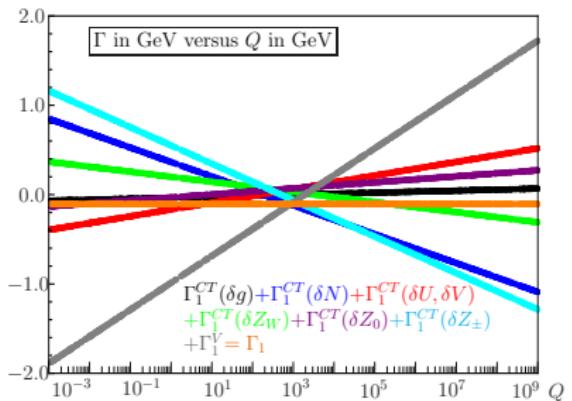
$$\xi_{ij} = K_\Lambda^j \Lambda_i - \frac{1}{\mu} \delta_{j3} \epsilon_i$$

with  $K_\Lambda^j$  neither proportional to  $v_j$  nor  $(Y_\nu)_j$ .



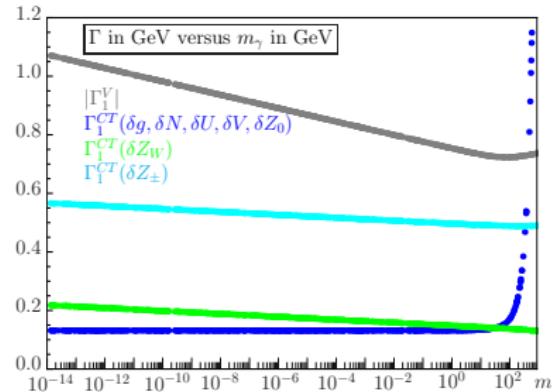
## UV divergence and Renormalization scale

Let's have a look at the dependence of the UV parameter  $\Delta$  and the renormalization scale  $Q$  for a NMSSM Neutralino decay.



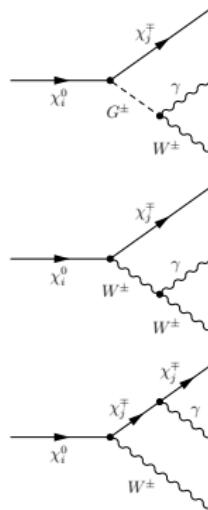


A massless photon produces IR divergences. Therefore one calculates with a photon with mass  $m_\gamma$ . But how to get rid of this dependence on an unphysical mass scale  $m_\gamma$ ?



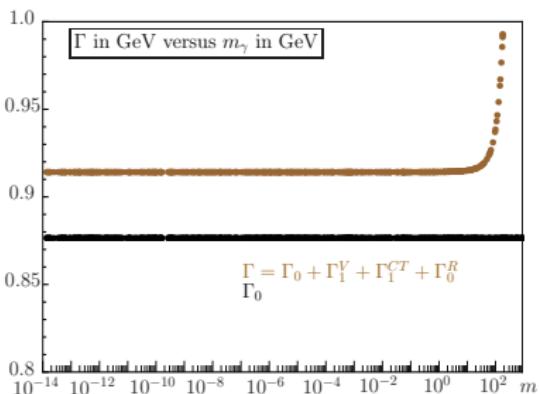
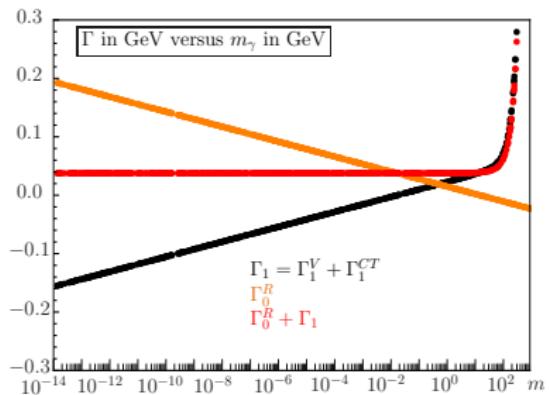
What is missing?

We have to take into account the real corrections coming from the emission of a photon with mass  $m_\gamma$ !





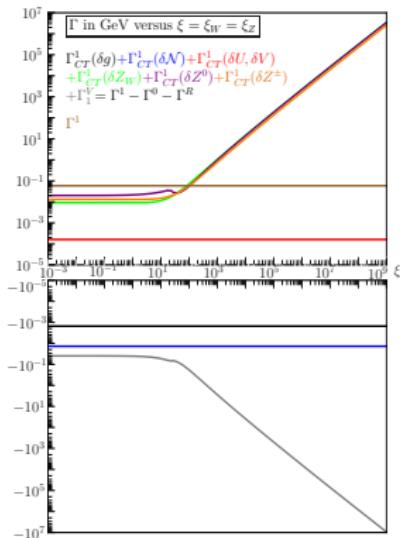
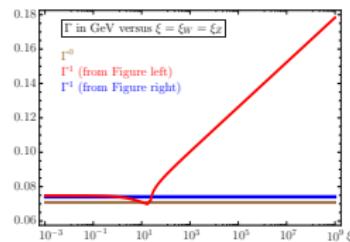
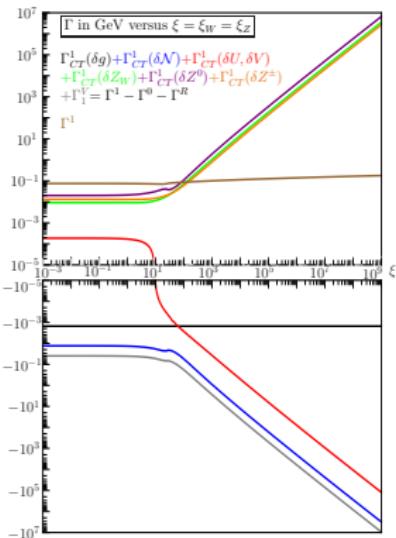
This can either be done by considering a cut-off energy for the real correction or by calculating the full hard photon emission, what was done here:



⇒ The final decay width is IR finite.



# Gauge dependence





Considering the possible extension of bilinear  $R$ -parity breaking, one can compare displaced vertex signals, completely invisible final state branching ratios for LSP decays and lightest Higgs decays:

	Displaced vertex	Comment	BR(Invis.)	Higgs decays
BRPV	Yes	Visible	$\leq 10\%$	standard
SRPV	Yes/No	anti-correlates with invisible	any	non-standard (invisible)
$\mu\nu$ SSM	Yes/No	anti-correlates with non-standard Higgs	$\leq 10\%$	non-standard

Combining the NMSSM with BRPV results in the superpotential:

$$W_{\text{NMSSM+BRPV}} = W_{\text{all}} - \lambda \hat{S} \hat{H}_d \hat{H}_u - \epsilon_i \hat{L}_i \hat{H}_u + \frac{1}{3!} \kappa \hat{S} \hat{S} \hat{S}$$

This model is very hard to distinguish from the  $\mu\nu$ SSM, since in both a light singlet scalar/neutralino can be present.