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Application Reference Summary & Back-up

> UH Ĥ

## **STEP:** a tool to perform tests of smoothness on differential distributions

**Terascale alliance meeting** 

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## Introduction

#### Step P.L.S. Connor

Introduction Motivation Method Application Reference Summary &

Conclusior

Back-up

## UH 2/18

### STEP library

# Smoothness Test with Expansion of Polynomials $\rightarrow$ automated 1D fits with correlations (optional) and early stopping (optional)

### Motivation

- Test the quality of a differential distribution in terms of the "alignment" of the points around a smooth function.
- Here we will discuss it in the context of the QCD interpretation of inclusive jet production<sup>a</sup>.
- But applications beyond QCD interpretation or even HEP are very likely.

 $^{a}\mbox{e.g.}$  Toni's presentation in which context it was developed.

### Applications in this presentation

Tevatron DØ & CDF ( $p\bar{p}$  at  $\sqrt{s} = 1.96 \text{ TeV}$ )

LHC CMS & ATLAS (pp at  $\sqrt{s} = 8 \text{ TeV}$ )

 $\longrightarrow$  These data have been used in QCD fits by PDF collaborations [1, 2, 3].



UH #

 $\longrightarrow$  Relevant for any differential measurement with "large" number of bins.

« Physics is smooth. » « Logarithmic scales can hide monsters. »



### Test of smoothness

Find an adequate analytical function to fit the spectrum in order to find deviations beyond bin-to-bin fluctuations.

- $\longrightarrow$  Applicable in different circumstances, as for instance:
  - Impact from bin-to-bin correlations.
  - Smooth systematic uncertainties.

- Check combination of different triggers.
- Smooth binned corrections.

UH H. 4/18

Motivation

Back-up

## Method

Chebyshev polynomials General form



ntroduction

- Method Chebyshev polynomials General form
- Application
- Reference
- Summary & Conclusions
- Back-up

UH # 5/18





## **Chebyshev polynomials**

### Difficulties with standard polynomials

1 Runge phenomenon

Ø Many parameters

ightarrow fits based on standard polynomials  $\sum a_i x^i$  difficult...

### Definition (first kind)

$$P_n(x) = \sum_{i=0}^n a_i T_i(x) \tag{1}$$

where 
$$T_0(x) = 1$$
,  $T_1(x) = x$  (2)

and 
$$T_{i+1}(x) = 2xT_i(x) - T_{i-1}(x)$$
 (3)

### Interesting properties

- Fits with Chebyshev polynomials more robust against Runge phenomenon.
- **2**  $P_n$  is a good approximation of  $P_{n+1}$ .

## **General form**

### Chebyshev fit of inclusive jet $p_T$ spectrum

$$f_n(p_{\rm T}) = \exp\left(\sum_{i=0}^n b_i T_i \left(2\frac{\log p_{\rm T}/\log p_{\rm T}^{\rm min}}{\log p_{\rm T}^{\rm max}/\log p_{\rm T}^{\rm min}} - 1\right)\right)$$
(4)

- The exponential and the logarithm are used because of the steeply falling nature of spectrum<sup>a</sup>
- Then rescaling is necessary as Chebyshev polynomials only cover [-1, 1].
- *n* is the **number of parameters** (one unity larger than the degree).

<sup>a</sup>They may of course be removed in case of other differential distributions.

### Remark

The **only assumption** is a smooth behaviour...  $\rightarrow$  No assumption on the physical nature of the fitted data!

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Introduction

Method Chebyshev polynomials General form

Application

Reference

Summary & Conclusions

Back-up

UH 6/18

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General form

Application

Summary &

UH Ĥ

### **Objective function**

$$\chi_n^2 = \min_{b_i \leq n} \left[ \left( \mathbf{x} - \mathbf{y}_{b_i \leq n} \right)^{\mathsf{T}} \mathbf{V}^{-1} \left( \mathbf{x} - \mathbf{y}_{b_i \leq n} \right) \right]$$
(5)

General form

- x corresponds to the binned differential distribution;
- $\mathbf{y}_{b_{i < n}}$  corresponds to the values of the smooth fit evaluated at, for instance, the centre of the bins for a given set of parameters  $b_{i < n}$ ;
- V is the covariance matrix of the binned differential distribution.

### Algorithm

- **()** The two first parameters are initialised from the first and last points of the spectrum whereas all other parameters are fixed to 0.
- Fit with all released parameters.
- **Stop** if one of the following statements is satisfied<sup>a</sup>:
  - 1 the  $\chi^2$  is compatible with the number of degrees of freedom (ndf);
  - f) the  $\chi^2$ /ndf is no longer decreasing;
  - f or the number of released parameters has reached the maximum allowed.
- **3** Release the next parameter and go back to item 1.

<sup>&</sup>lt;sup>a</sup>The two first criteria can be deactivated

### Application CDF DØ CMS ATLAS Comparison

## **Application**

### Datasets

Tevatron Run 2 DØ & CDF ( $p\bar{p}$  at  $\sqrt{s} = 1.96 \text{ TeV}$ )

LHC Run 1 CMS & ATLAS (pp at  $\sqrt{s} = 8 \text{ TeV}$ )

 $\rightarrow$  These data have been used in QCD fits by PDF collaborations [1, 2, 3].

Bin-to-bin uncertainties						
detector	stat. corr.	syst. unc.				
CDF [4]			-			
DØ [5]		×				
CMS [6]	×	×				
ATLAS [7]	×					

UH #18

Step PLS

CONNOR

Method

ATLAS.

Reference Summary &

Back-up

Application CDF

### CDF

Method Application CDF DØ CMS ATLAS Comparison

Reference

Summary & Conclusions Back-up



### Figure

Steps due to the trigger strategy visible at 146 GeV (96 GeV) in 2nd (3rd) y bin.  $\rightarrow$  Seems to be negligible in front of the statistical uncertainties.

	y  < 0.1	0.1 <  y  < 0.7	0.7 <  y  < 1.1	1.1 <  y  < 1.6	1.6 <  y  < 2.1
2	$163.16 \pm 0.38$	$1896.56 \pm 0.38$	$1165.27 \pm 0.39$	$1441.15 \pm 0.41$	$1224.13 \pm 0.47$
3	$7.03 \pm 0.39$	$97.04 \pm 0.39$	$67.30 \pm 0.41$	$62.69 \pm 0.43$	$41.67 \pm 0.50$
4	$0.70 \pm 0.41$	$8.82 \pm 0.41$	$9.89 \pm 0.43$	$1.69 \pm 0.45$	$1.97 \pm 0.53$
5	$0.56 \pm 0.43$	$1.17 \pm 0.43$	$0.54 \pm 0.45$	$0.43 \pm 0.47$	$0.72 \pm 0.58$
6	$0.55 \pm 0.45$	$0.65 \pm 0.45$	$0.60 \pm 0.47$	$0.42 \pm 0.50$	$0.42 \pm 0.63$

UH #



Method Application CDF DØ CMS ATLAS Comparison

Reference

Summary & Conclusions Back-up



### Figure

Additional bin-to-bin uncorrelated systematic uncertainty seems slightly overestimated at |y| > 0.8.

	y  < 0.4	0.4 <  y  < 0.8	0.8 <  y  < 1.2	1.2 <  y  < 1.6	1.6 <  y  < 2.0	2.0 <  y  < 2.4
2	$811.57 \pm 0.31$	$1016.79 \pm 0.32$	$795.84 \pm 0.33$	$432.24 \pm 0.37$	$386.47 \pm 0.39$	$329.43 \pm 0.43$
3	$59.47 \pm 0.32$	$76.89 \pm 0.32$	$54.23 \pm 0.34$	$35.55 \pm 0.38$	$27.01 \pm 0.41$	$14.65 \pm 0.45$
4	$6.45 \pm 0.32$	$7.31 \pm 0.33$	$3.35 \pm 0.35$	$1.47 \pm 0.39$	$2.04 \pm 0.43$	$1.02 \pm 0.47$
5	$0.93 \pm 0.33$	$1.44 \pm 0.34$	$0.50 \pm 0.37$	$0.33 \pm 0.41$	$0.31 \pm 0.45$	$0.42 \pm 0.50$
6	$0.88 \pm 0.34$	$1.29 \pm 0.35$	$0.45 \pm 0.38$	$0.35 \pm 0.43$	$0.32 \pm 0.47$	$0.42 \pm 0.53$

UH 10/18

#### Method Application CDF DØ CMS ATLAS Comparison Reference

Summary & Conclusions Back-up

> UH 11/18



**CMS** 



### CMS

#### Step P.L.S. Connor

#### Introduction

Vethod

Application CDF DØ CMS ATLAS

Reference

Summary & Conclusions

Back-up

## UH #

### Figure

- Only the **phase space** relevant for QCD interpretation has been considered.
- Statistical correlations are provided directly by CMS and accounted for in each fit.
- Steps related to the **unfolding** may still be seen but are more difficult to find due to correlations and to the unfolding.
- The additional 1% bin-to-bin uncorrelated systematic uncertainty is necessary but likely too conservative (wavy shape).

- 1	y  < 0.5	0.5 <  y  < 1.0	1.0 <  y  < 1.5	1.5 <  y  < 2.0	2.0 <  y  < 2.5
2	$841.51 \pm 0.25$	$1019.44 \pm 0.25$	$1442.31 \pm 0.25$	$1965.81 \pm 0.26$	$1943.62 \pm 0.29$
3	$47.77 \pm 0.25$	$66.72 \pm 0.25$	$115.51 \pm 0.25$	$185.72 \pm 0.26$	$187.95 \pm 0.30$
4	$7.52 \pm 0.25$	$7.55 \pm 0.25$	$14.07 \pm 0.25$	$21.57 \pm 0.27$	$17.79 \pm 0.31$
5	$0.78 \pm 0.26$	$0.84 \pm 0.26$	$0.93 \pm 0.26$	$0.90 \pm 0.27$	$0.70 \pm 0.32$
6	$0.76 \pm 0.26$	$0.65 \pm 0.26$	$0.50 \pm 0.26$	$0.47 \pm 0.28$	$0.41 \pm 0.32$

#### Method Application CDF DØ CMS ATLAS Comparison Reference

Summary & Conclusions Back-up

> UH 13/18





## ATLAS

## **ATLAS**

### Figure

- **Statistical correlations** are accounted for using replica method.
- The statistical uncertainties seem insufficient to cover the deviations from a smooth behaviour in the two first rapidity bins, and no additional bin-to-bin uncorrelated systematic uncertainties are provided.
- This might be related to the issues observed by several PDF collaborations [1, 2, 3].

	y  < 0.5	0.5 <  y  < 1.0	1.0 <  y  < 1.5	1.5 <  y  < 2.0	2.0 <  y  < 2.5	2.5 <  y  < 3.0
2	$2188.98 \pm 0.25$	$3033.97 \pm 0.25$	$4115.24 \pm 0.26$	$3860.63 \pm 0.27$	$2722.43 \pm 0.30$	$1142.21 \pm 0.35$
3	$99.77 \pm 0.25$	$178.11 \pm 0.26$	$325.30 \pm 0.26$	$322.57 \pm 0.27$	$261.13 \pm 0.31$	$92.98 \pm 0.37$
4	$17.57 \pm 0.26$	$27.64 \pm 0.26$	$49.63 \pm 0.27$	$37.35 \pm 0.28$	$17.49 \pm 0.32$	$9.03 \pm 0.38$
5	$1.61 \pm 0.26$	$4.62 \pm 0.27$	$4.62 \pm 0.27$	$3.67 \pm 0.28$	$2.70 \pm 0.32$	$1.92 \pm 0.39$
6	$1.46 \pm 0.27$	$1.84 \pm 0.27$	$0.91 \pm 0.28$	$1.15 \pm 0.29$	$0.93 \pm 0.33$	$0.80 \pm 0.41$
7	$1.46 \pm 0.27$	$1.84 \pm 0.28$	$0.92 \pm 0.28$	$1.18 \pm 0.29$	$0.73 \pm 0.34$	$0.86 \pm 0.43$

UH 14/18

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CONNOR

Method

ATLAS

Reference

Back-up

Summary &

Application CDF

## Comparison



### Interpretation

- Steps due to the **triggers** are visible (e.g. CDF & CMS).
- Certain bin-to-bin decorrelated uncertainties are likely overestimated (DØ in forward region; CMS in all regions) or underestimated (ATLAS).
- Need for higher number of parameters in central region of ATLAS measurement might be related to difficulties encountered in QCD interpretation.

Step P.L.S. Connor

Application

CDF

ATLAS Comparison

Reference

Back-up

Summary &

UH 15/18

## Reference

Article Source code

## Article

 $\ensuremath{\mathsf{STEP:}}$  a tool to perform tests of smoothness on differential distributions based on Chebyshev polynomials of the first kind

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Nov

[hep-

Aarmser: We motivate and docribe a method based on fits with Chebyshev polynomials to tee the smeetheness of differential distributions. We also provide a header-only tool in C++ called Srav to perfere we have has a A a demonstration we use the tool in the constraint of the measurement of inclusive deduke-differential cross section in the jet transverse momentum and rapidity at the Transron and LHC.

Krywoatse: smoothness, step, Chebyshev polynomials, inclusive jet, correlations, D0, CDF, CMS ATLAS, library, C++

### Outline

- Motivation and description of the method
- Application on real data
- $\rightarrow$  Essentially the content of this presentation.

### Status

- Posted on arXiv [8].
- Plan to submit it to a journal in the coming hours/days.

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P.L.S.

Application

Summary &

Article

## Reference

### Source code

### How-to

- Just download the header Step.h from the GitLab repository [9], then the following call is enough to get a ROOT TF1 object:
  - TF1 \* f = Step:::GetSmoothFit<log,exp>(hist,n);
- Here is an example of a more advanced call (but more complex isn't possible):

TF1 \* f = Step:::GetSmoothFit (h,cov,im,iM,n,1,true,cout);

Most general prototype:

 $\longrightarrow$  Additional information on the fit performance is available using the static STL container Step::chi2s.

### Examples

All results shown in the presentation and in the article may be reproduced.

UH 17/18

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Application

Summary &

Back-up

Article Source code

## **Summary & Conclusions**

### Step P.L.S. CONNOR ntroduction Method Application

### Application Reference Summary &

Conclusions Back-up **Summary & Conclusions** 

- We have presented an original method to test the quality of a differential distribution.
- We have applied the method to several measurements used by PDF collaborations.
- We have provided a package and briefly instructed how to use it.

# Thank you for your attention!

## Back-up

## Acronyms I

#### ATLAS A Toroidal LHC ApparatuS. 3, 11, 18

- CDF Collider Detector at Fermilab. 3, 11, 18
- CMS Compact Muon Solenoid. 3, 11, 15, 18
- DØ after the location of the detector. 3, 11, 18

- HEP High-Energy Physics. 3
- PDF Parton Distribution Function. 3, 11, 17, 23
- QCD Quantum Chromodynamics. 3, 11, 15, 18
- STL Standard Template Library. 21

UH 19/18

Step

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Acronyms References Visiting card

## **References I**

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UH 21/18

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UH 22/18