

Reconciling Higgs physics and pseudo-Nambu-Goldstone dark matter in the S2HDM using a genetic algorithm

[2108.10864]

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in collaboration with María Olalla Olea Romacho

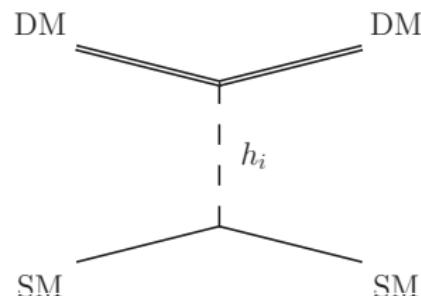
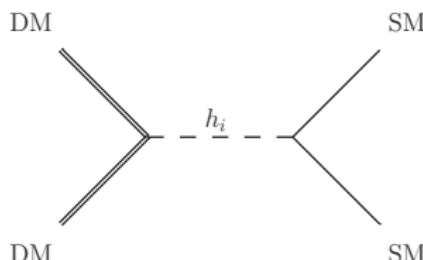
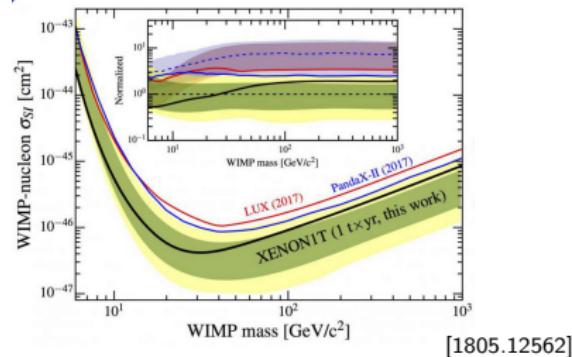
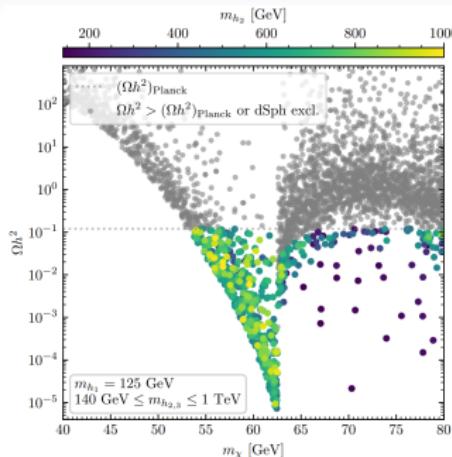
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Helmholtz Alliance

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QUANTUM UNIVERSE

Higgs portal dark matter



Can we have one without the other?

Pseudo-Nambu-Goldstone dark matter (pNG)

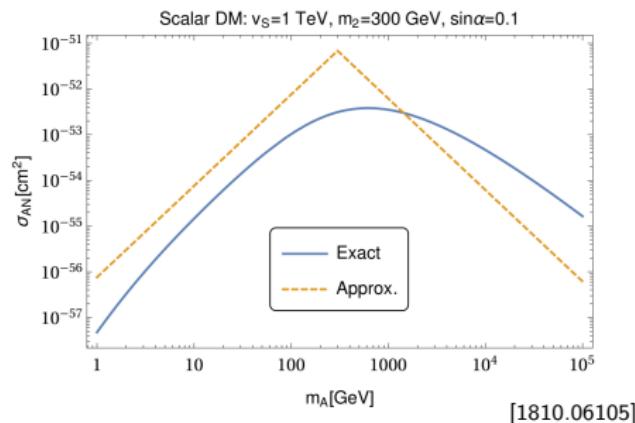
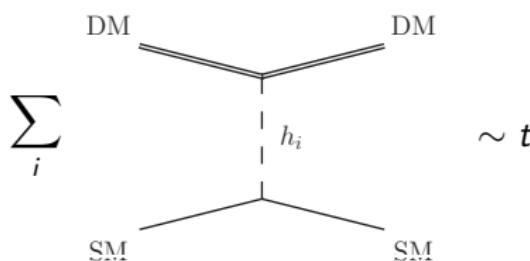
S : Complex field charged under a softly-broken global $U(1)$

$$\mathcal{L} = (\partial_\mu S)^* \partial^\mu S - V(\phi_i, S)|_{U(1)} - V(S)|_{\cancel{U(1)}} - \text{soft}$$

$$S = \frac{1}{\sqrt{2}} (v_s + s) e^{i \frac{\chi}{v_s}} \quad \Rightarrow \quad \mathcal{L}_{\chi \chi s} = \frac{1}{2v_s} (\partial^2 s) \chi \chi - \frac{s}{v_s} \chi (\partial^2 + m_\chi^2) \chi$$

[2109.11499]

On-shell χ interactions with Higgs sector proportional to momentum of s



Direct detection constraints largely irrelevant

Pseudo-Nambu-Goldstone dark matter (pNG)

Most studied case: $V = V_{\text{SM}}(H) + V(H, S)|_{U(1)} + \mu_\chi (S^2 + (S^*)^2)$

[0811.0393], [1609.07490], [1708.02253], [1812.05952], [1810.06105], [1810.08139], [1912.04008], [1906.02175], ...

- Predict DM relic abundance :)
 - DM constraints: ID important, almost no sensitivity with DD
 - Collider: $h_{125}-s$ mixing, $E_{T_{\text{missing}}}$ signatures
-

Here: $V = V_{\text{2HDM-II}}(\phi_1, \phi_2) + V(\phi_1, \phi_2, \phi_S)|_{U(1)} + \mu_\chi (\phi_S^2 + (\phi_S^*)^2)$

[2108.10864]

- 2HDM: First-order EW phase transitions, Susy, Axion models
- DM: Richer DM-Higgs portal interactions
- Collider: Richer mixing patterns and new states

Aim: Model exploration taking into account all constraints

We explore parameter region motivated by

- (i) Cosmic ray excesses (Fermi and AMS)
- (ii) collider excesses at 96 GeV

S2HDM: Singlet-extended 2 Higgs doublet model

ϕ_1, ϕ_2 : SU(2) doublets, ϕ_S : SM singlet, charged under global U(1)

Scalar potential:

$$V = V_{\text{2HDM-II}}(\phi_1, \phi_2) + V(\phi_1, \phi_2, \phi_S)|_{U(1)} + \mu_\chi (\phi_S^2 + (\phi_S^*)^2)$$

EW vacuum:

$$\langle \phi_1 \rangle = \begin{pmatrix} 0 \\ v_1/\sqrt{2} \end{pmatrix}, \quad \langle \phi_2 \rangle = \begin{pmatrix} 0 \\ v_2/\sqrt{2} \end{pmatrix}, \quad \langle \phi_S \rangle = v_S/\sqrt{2} \in \mathbb{R}$$

BSM particles:

$h_{1,2,3}$: CP-even Higgs bosons

H^\pm : Charged Higgs bosons

A : CP-odd Higgs boson

χ : pNG DM

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = R(\alpha_1, \alpha_2, \alpha_3) \cdot \begin{pmatrix} \text{Re}(\phi_1^0) \\ \text{Re}(\phi_2^0) \\ \text{Re}(\phi_S^0) \end{pmatrix}$$

Singlet components: $\Sigma_{h_i} = R_{i3}^2$

Free parameters (Yukawa type II):

$$m_{h_{1,2,3}}, \quad m_A, \quad m_{H^\pm}, \quad m_\chi, \quad \alpha_{1,2,3}, \quad \tan \beta = v_1/v_2, \quad M = \sqrt{\mu_{12}^2 / (s_\beta c_\beta)}, \quad v_S$$

μ_{12}^2 : Soft Z_2 -breaking parameter

Theoretical and experimental constraints

Theory:

Vacuum stability: EW minimum is global minimum (strict) [Hom4PS2]

Perturbativity: Upper limit on scalar $2 \rightarrow 2$ scattering amplitudes in large s limit
 $\rightarrow |\text{Eig}[\mathcal{M}(\lambda_i)]| < 8\pi$

RGE evolution: Check for boundedness and perturbativity until at least $\mu_\nu = 1$ TeV

Experiment:

Colliders: Searches, measurements of h_{125} [N2HDECAY, HiggsBounds, HiggsSignals]

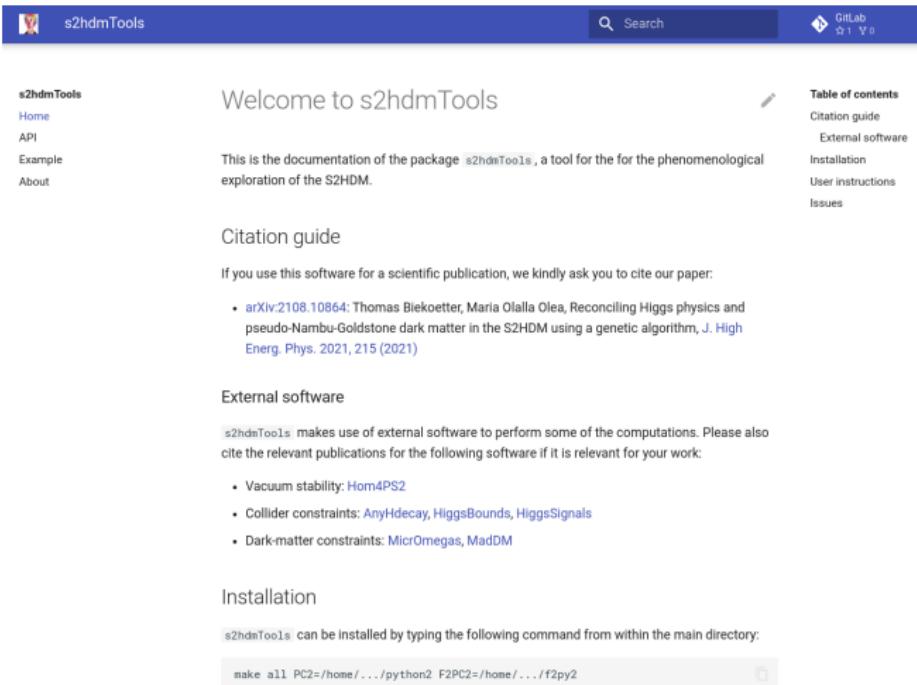
STU: EW precision observables (at one loop)

Flavour: $\tan \beta > 1.5$ and $m_{H^\pm} > 600$ GeV to avoid bounds (type II)

Dark matter: Relic abundance $h^2 \Omega < 0.12$ [Micromegas]

Indirect detection limits from Fermi dSph observations [MadDM]

s2hdmTools



The screenshot shows the homepage of the `s2hdmTools` documentation. At the top, there's a navigation bar with the logo, the project name, a search bar, and links to GitLab and GitHub. On the left, a sidebar lists "s2hdmTools" and several sub-links: Home, API, Example, and About. The main content area has a title "Welcome to s2hdmTools" with an edit icon. Below it, a text block states: "This is the documentation of the package `s2hdmTools`, a tool for the for the phenomenological exploration of the S2HDM." To the right of the main content is a "Table of contents" sidebar with links to Citation guide, External software, Installation, User instructions, and Issues.

Welcome to s2hdmTools

This is the documentation of the package `s2hdmTools`, a tool for the for the phenomenological exploration of the S2HDM.

Citation guide

If you use this software for a scientific publication, we kindly ask you to cite our paper:

- arXiv:2108.10864: Thomas Biekoetter, Maria Olalla Olea, Reconciling Higgs physics and pseudo-Nambu-Goldstone dark matter in the S2HDM using a genetic algorithm, *J. High Energ. Phys.* 2021, 215 (2021)

External software

`s2hdmTools` makes use of external software to perform some of the computations. Please also cite the relevant publications for the following software if it is relevant for your work:

- Vacuum stability: `Horn4PS2`
- Collider constraints: `AnyHdecay`, `HiggsBounds`, `HiggsSignals`
- Dark-matter constraints: `MicroOmegas`, `MadDM`

Installation

`s2hdmTools` can be installed by typing the following command from within the main directory:

```
make all PC2=/home/.../python2 F2PC2=/home/.../f2py2
```

The makefile uses `pip` to install the package to your python environment.

<https://gitlab.com/thomas.biekoetter/s2hdmtools/>

Parameter Scan

S2HDM Type II:

$$\begin{aligned}
 & 1.5 \leq \tan \beta \leq 10, \quad m_{h_1} = 125.09 \text{ GeV}, \quad 140 \text{ GeV} \leq m_{h_{2,3}} \leq 1 \text{ TeV}, \\
 & 40 \text{ GeV} \leq m_\chi \leq 80 \text{ GeV}, \quad 40 \text{ GeV} \leq v_S \leq 1 \text{ TeV}, \quad -\pi/2 \leq \alpha_{1,2,3} \leq \pi/2, \\
 & 400 \text{ GeV} \leq M \leq 1 \text{ TeV}, \quad 600 \text{ GeV} \leq m_{H^\pm} \leq 1 \text{ TeV}, \quad m_A \leq 1 \text{ TeV}, \\
 & \Delta M_{\max} = \max(|m_H - M|, |m_A - M|, |m_{H^\pm} - M|) < 200 \text{ GeV}, \\
 & \text{with } m_H = m_{h_2}(m_{h_3}) \text{ for } \Sigma_{h_2} > (<) \Sigma_{h_3}
 \end{aligned}$$

Second scan: $m_{h_1} = 96 \text{ GeV}$, $m_{h_2} = 125 \text{ GeV}$

Genetic algorithm:* Minimizing the loss function L

$$L_{1,2} = \chi^2_{125} (+\chi^2_{96}) + \max \left[0, (r_{\text{obs}}^{\text{HB}} - 1) \cdot 100 \right] + \begin{cases} C, & \chi^2_{ST} > 5.99 \text{ or theo. constr.} \\ 0, & \text{otherwise} \end{cases}$$

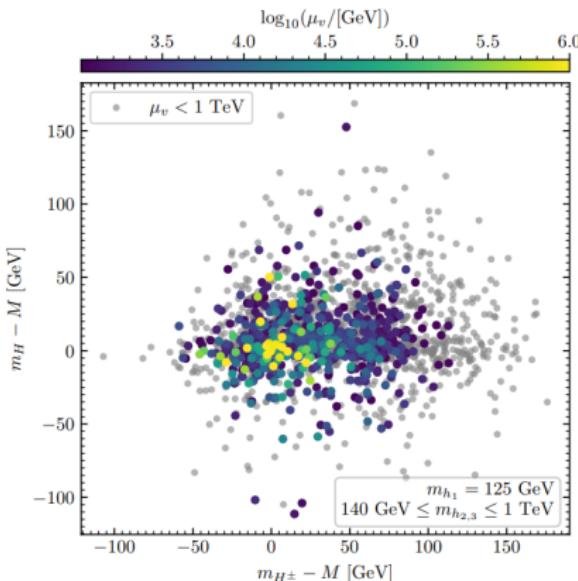
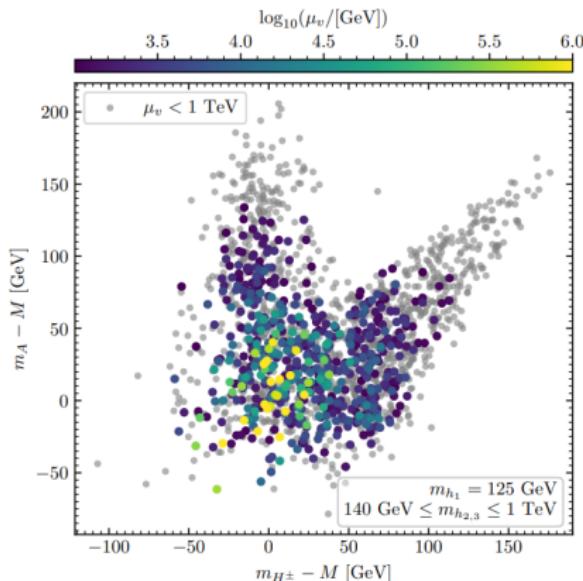
⇒ Points that fulfill theory constraints at $\mu = v_{\text{EW}}$ and are allowed by collider measurements

- Check theory constraints including **RGE evolution**
- Calculate **DM observables**

*Details appendix/questions

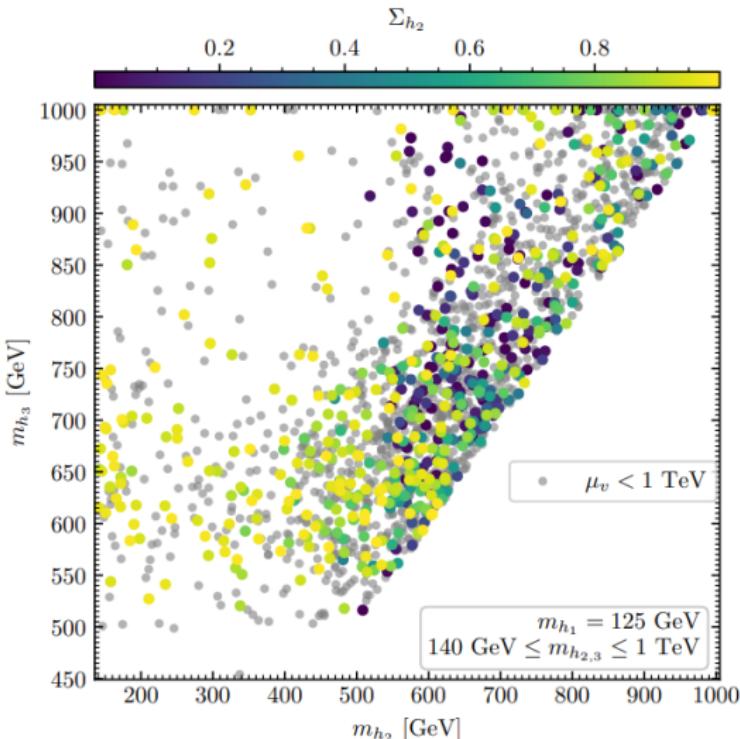
Theoretical constraints: Impact of RGE running

μ_v : Energy scale until the model is theoretically **viable**



⚠ First-order EW phase transitions (baryogenesis and GW) require sizable mass splittings

Scalar spectrum



Most common:

$$m_H \sim m_A \sim m_{H^\pm} \sim M \gtrsim 500 \text{ GeV}$$

$$m_{h_i} \lesssim v_S \text{ for } \Sigma_{h_i} \sim 1$$

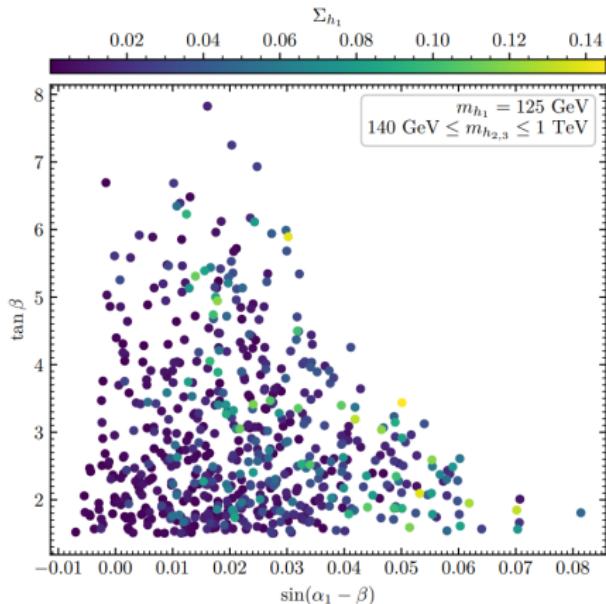
- Higgs cascade decays
- SM decay modes of H, A, H^\pm can be suppressed w.r.t. 2HDM
- Spectrum can be somewhat lighter

DM: Here $M \gg m_\chi$

→ Annihilation via $h_{1,2} = h_{125,S}$

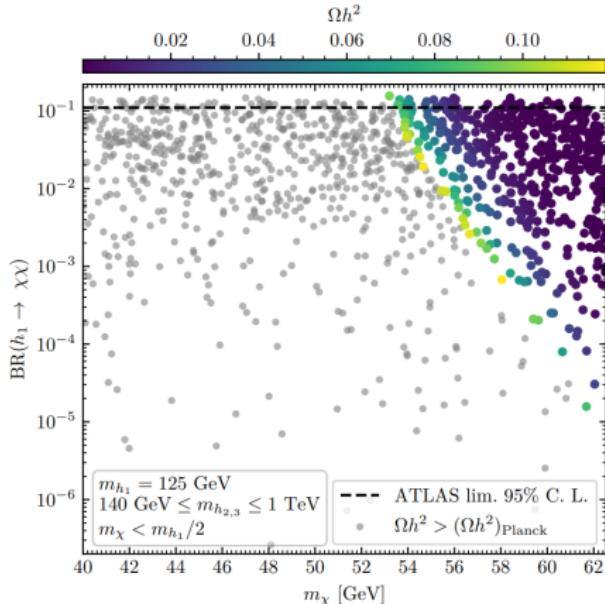
SM-like Higgs boson at 125 GeV

$$h_1 = h_{125}$$



$\max(\Sigma_{h_1}) \sim 0.14$ (in SM+S model ~ 0.07)

Possibility to distinguish the models

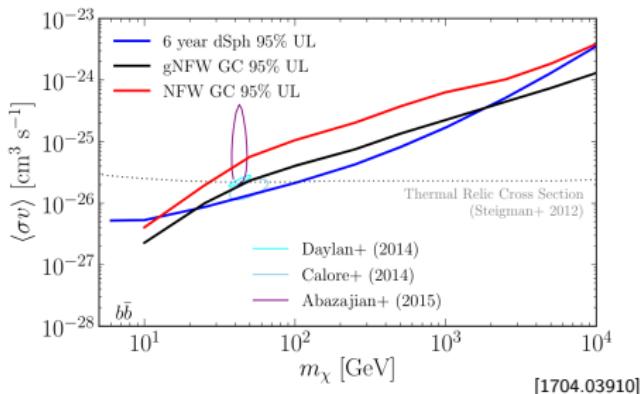


$m_\chi \gtrsim 54 \text{ GeV}$ (similar to SM+S limit)

But depends on $m_{h_5} > m_{h_{125}}$

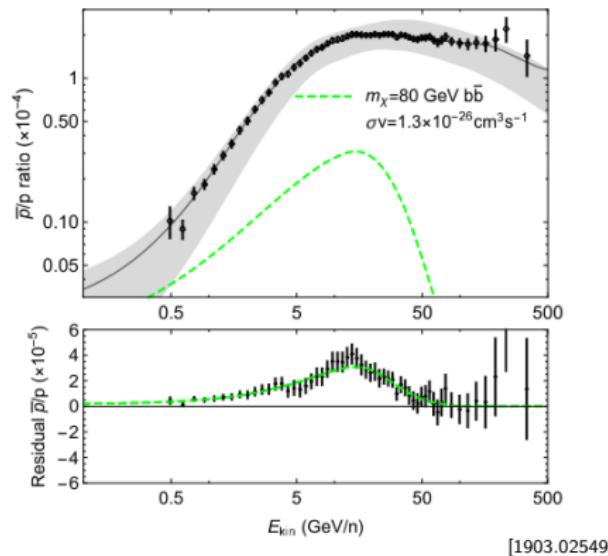
Galactic center excess and antiproton excess

Fermi γ excess



AMS \bar{p} excess

ISM Model I



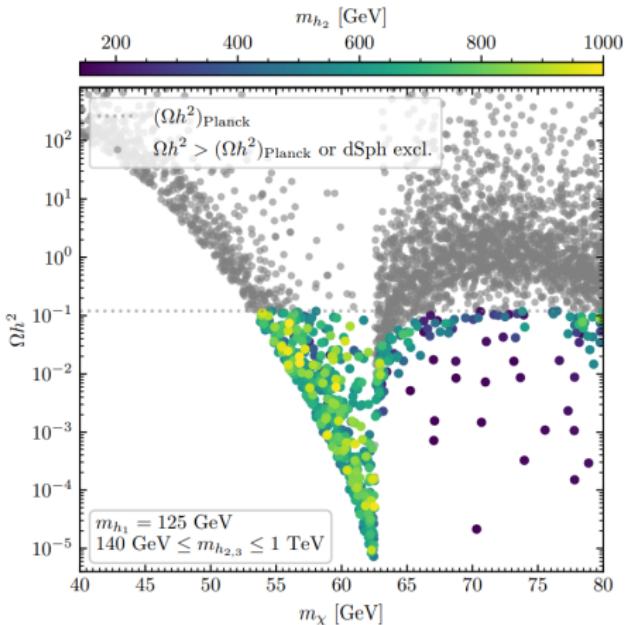
Assuming that origin is DM annihilation:

1. Both excesses are compatible
2. They require annihilation XS of the order of the thermal relic XS
3. Both consistent with $b\bar{b}$ annihilation \rightarrow Higgs portal DM
4. Currently probed by observation of dSph

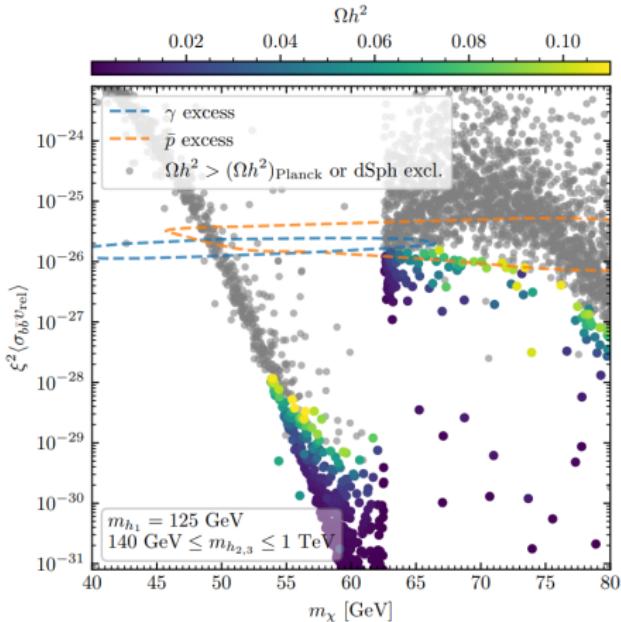
Exciting :)

Thermal DM relics in the Higgs funnel region

Relic abundance



Indirect detection ($\xi = \Omega h_{\text{pred}}^2 / \Omega h_{\text{Planck}}^2$)



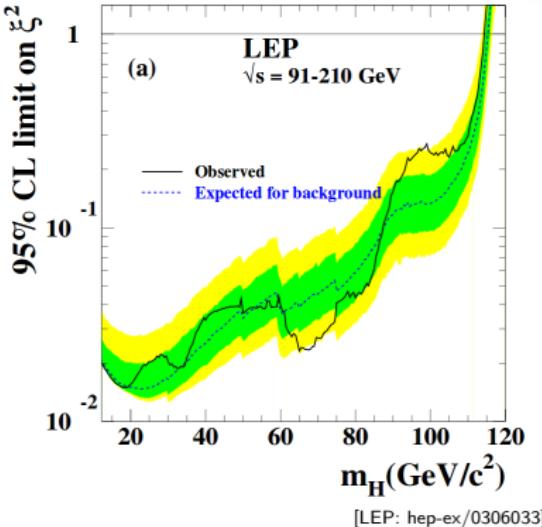
Very interesting region at $62 \text{ GeV} \lesssim m_\chi \lesssim 65 \text{ GeV}$:

✓ Relic abundance ✓ Fermi γ excess* ✓ AMS \bar{p} excess*

(✗) In tension with dSph exclusion limits (but there are large uncertainties)

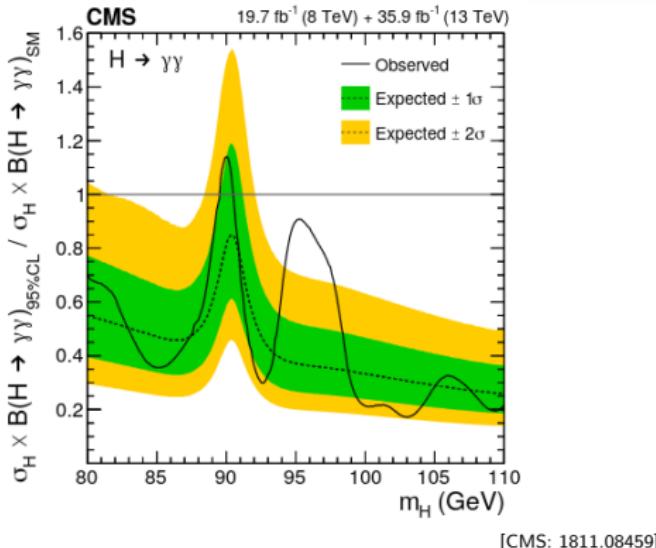
*Preferred regions from [1409.0042, 1903.02549]

“The 96GeV excesses” (LEP and CMS)



$$\mu_{\text{LEP}} (e^+ e^- \rightarrow Z h \rightarrow Z b\bar{b}) = 0.117 \pm 0.057$$

[1612.08522]



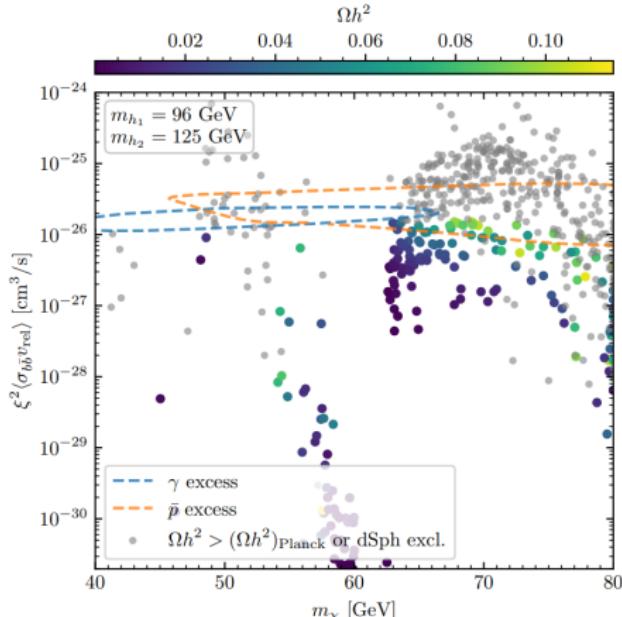
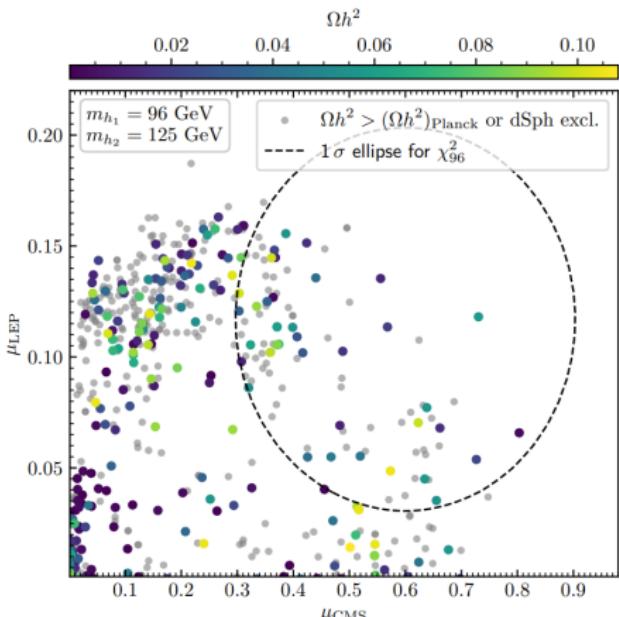
$$\mu_{\text{CMS}} (gg \rightarrow h \rightarrow \gamma\gamma) = 0.6 \pm 0.2$$

$\rightarrow \chi^2_{96}(\mu_{\text{LEP}}, \mu_{\text{CMS}})$ assuming no correlation between μ_{LEP} and μ_{CMS}

Many model interpretations with common origin of both excesses, including N2HDM and NMSSM
 see [2003.05422] for a list models

Higgs funnel DM and a Higgs boson at 96 GeV

Scan: As before, but with $m_{h_1} = 96$ GeV and $m_{h_2} = 125$ GeV



Very interesting region at $62 \text{ GeV} \lesssim m_\chi \lesssim 65 \text{ GeV}$:

- ✓ Relic abundance ✓ Fermi γ excess ✓ AMS \bar{p} excess ✓ CMS excess ✓ LEP excess
- (✗) In tension with dSph exclusion limits (but there are large uncertainties)

Summary and outlook

- The S2HDM is theoretically well motivated and has a rich Higgs and DM phenomenology
- We demonstrated that important (theoretical) constraints have been overlooked so far
- We provide the public code `s2hdmTools` for model explorations

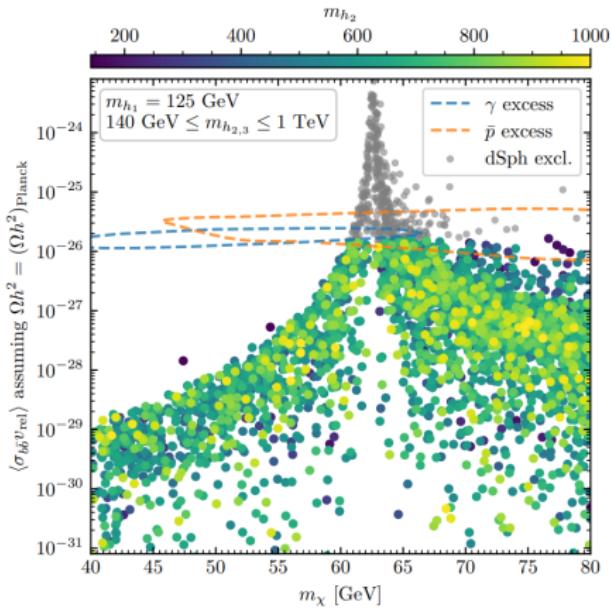
In the Higgs funnel region recent experimental anomalies can be explained:

- ✓ Relic abundance ✓ Fermi γ excess ✓ AMS \bar{p} excess ✓ CMS excess ✓ LEP excess
- (✗) In tension with dSph exclusion limits (but there are large uncertainties)

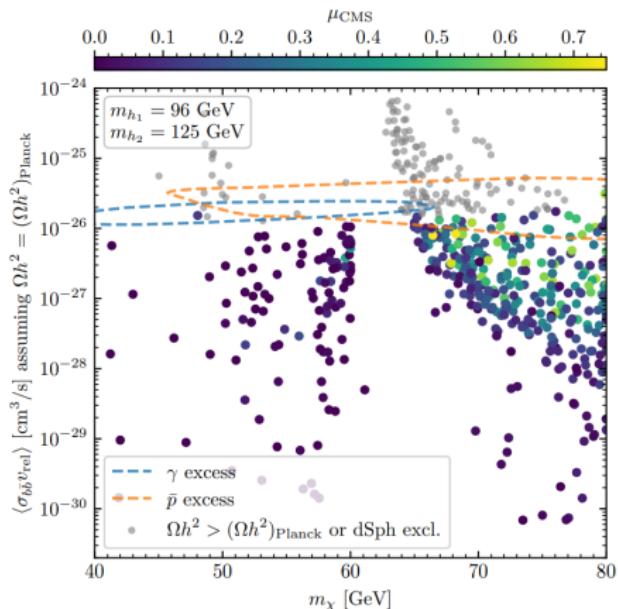
THANKS FOR YOUR ATTENTION!

Assuming $\Omega h^2 = \Omega h_{\text{Planck}}^2$ – Non-standard cosmological history

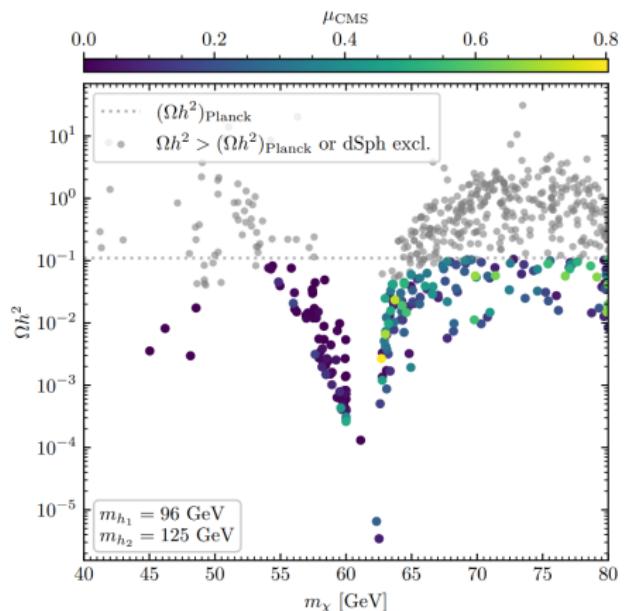
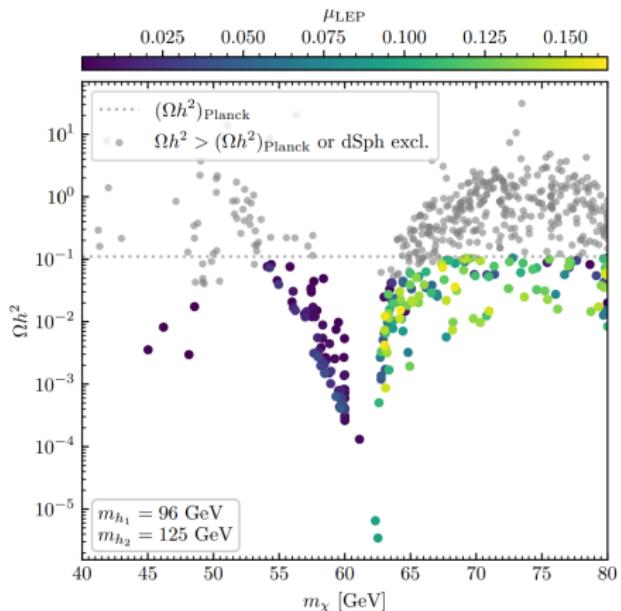
$m_{h_1} = 125 \text{ GeV}$, $m_{h_2} > 140 \text{ GeV}$



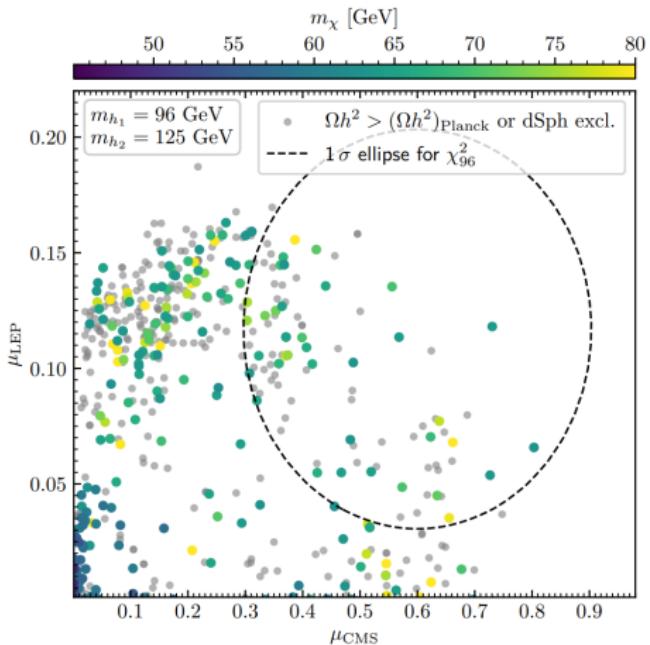
$m_{h_1} = 96 \text{ GeV}$, $m_{h_2} = 125 \text{ GeV}$



Higgs funnel DM and a Higgs boson at 96 GeV



Higgs funnel DM and a Higgs boson at 96 GeV



Parameter Scan

S2HDM Type II:

$$\begin{aligned}1.5 \leq \tan \beta \leq 10, \quad m_{h_1} = 125.09 \text{ GeV}, \quad 140 \text{ GeV} \leq m_{h_{2,3}} \leq 1 \text{ TeV}, \\40 \text{ GeV} \leq m_\chi \leq 80 \text{ GeV}, \quad 40 \text{ GeV} \leq v_S \leq 1 \text{ TeV}, \quad -\pi/2 \leq \alpha_{1,2,3} \leq \pi/2, \\400 \text{ GeV} \leq M \leq 1 \text{ TeV}, \quad 600 \text{ GeV} \leq m_{H^\pm} \leq 1 \text{ TeV}, \quad m_A \leq 1 \text{ TeV}, \\|\Delta M_{\max}| = \max(|m_H - M|, |m_A - M|, |m_{H^\pm} - M|) < 200 \text{ GeV}, \quad m_H = m_{h_2} \text{ or } m_{h_3}\end{aligned}$$

Genetic algorithm: Minimizing the loss function L

$$L = \chi^2_{125} (+\chi^2_{96}) + \max \left[0, (r_{\text{obs}}^{\text{HB}} - 1) \cdot 100 \right] + \begin{cases} C, & \chi^2_{ST} > 5.99 \text{ or theo. constr.} \\ 0, & \text{otherwise} \end{cases}$$

Individuals: $[n_1, n_2, \dots, n_{14}]$, $0 < n_i < 1 \Rightarrow \tan \beta(n_i)$, $m_{h_i}(n_i)$, ...

Population: 50 000 individuals, randomly generated

Evolution: Selection: Tournament selection with size 3

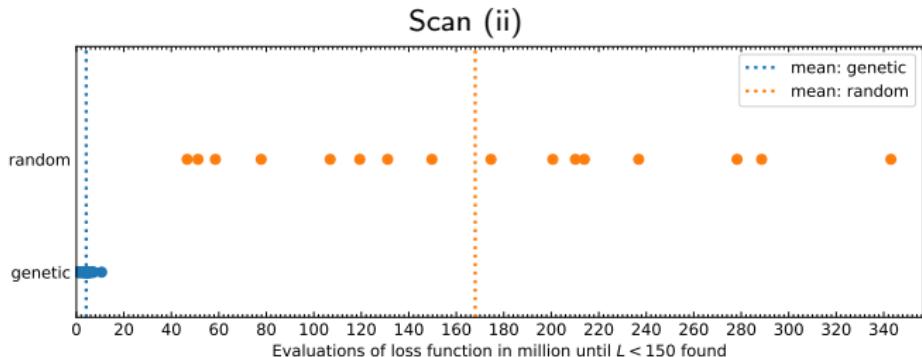
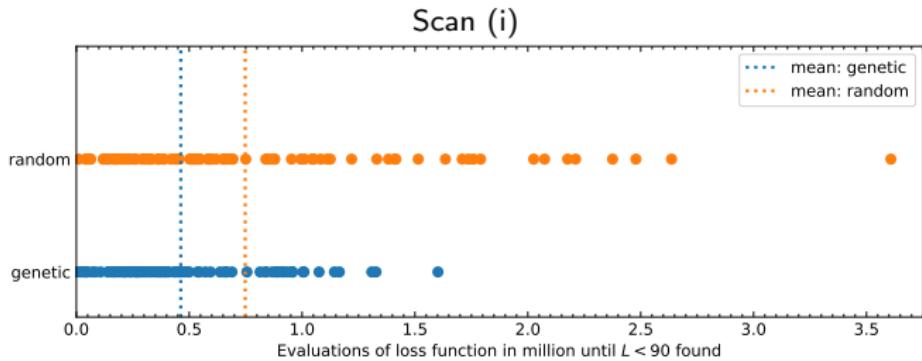
Mating: Uniform crossover of 2 individuals with $p = 20\%$, mating probability 80%

Mutation: Float uniform mutator with $p = 10\%$, mutation probability 20%

Generations: Maximum 40, or until individual with $L \leq L_{\text{threshold}}$ has been found

→ For the resulting points calculate **DM observables**

Performance of genetic algorithm



S2HDM: Singlet-extended 2 Higgs doublet model

ϕ_1, ϕ_2 : SU(2) doublets, ϕ_S : SM singlet, charged under global U(1)

Scalar potential:

$$\begin{aligned} V = & \mu_{11}^2 (\phi_1^\dagger \phi_1) + \mu_{22}^2 (\phi_2^\dagger \phi_2) - \mu_{12}^2 ((\phi_1^\dagger \phi_2) + (\phi_2^\dagger \phi_1)) + \frac{1}{2} \mu_S^2 |\phi_S|^2 - \frac{1}{4} \mu_\chi^2 (\phi_S^2 + (\phi_S^*)^2) \\ & + \frac{1}{2} \lambda_1 (\phi_1^\dagger \phi_1)^2 + \frac{1}{2} \lambda_2 (\phi_2^\dagger \phi_2)^2 + \lambda_3 (\phi_1^\dagger \phi_1) (\phi_2^\dagger \phi_2) + \lambda_4 (\phi_1^\dagger \phi_2) (\phi_2^\dagger \phi_1) \\ & + \frac{1}{2} \lambda_5 ((\phi_1^\dagger \phi_2)^2 + (\phi_2^\dagger \phi_1)^2) + \frac{1}{2} \lambda_6 (|\phi_S|^2)^2 + \lambda_7 (\phi_1^\dagger \phi_1) |\phi_S|^2 + \lambda_8 (\phi_2^\dagger \phi_2) |\phi_S|^2 \end{aligned}$$

EW vacuum: $\langle \phi_1 \rangle = \begin{pmatrix} 0 \\ v_1/\sqrt{2} \end{pmatrix}, \quad \langle \phi_2 \rangle = \begin{pmatrix} 0 \\ v_2/\sqrt{2} \end{pmatrix}, \quad \langle \phi_S \rangle = v_S/\sqrt{2} \in \mathbb{R}$

BSM particles:

$h_{1,2,3}$: CP-even Higgs bosons

H^\pm : Charged Higgs bosons

A : CP-odd Higgs boson

χ : pNG DM

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = R(\alpha_1, \alpha_2, \alpha_3) \cdot \begin{pmatrix} \text{Re}(\phi_1^0) \\ \text{Re}(\phi_2^0) \\ \text{Re}(\phi_S^0) \end{pmatrix}$$

Free parameters (Yukawa type II):

$$m_{h_{1,2,3}}, \quad m_A, \quad m_{H^\pm}, \quad m_\chi, \quad \alpha_{1,2,3}, \quad \tan \beta, \quad M = \sqrt{\mu_{12}^2 / (s_\beta c_\beta)}, \quad v_S$$