

SUSY at future colliders: Will e^+e^- -colliders be relevant? ¹

Mikael Berggren¹

¹DESY, Hamburg

First ECFA Workshop on e^+e^- Higgs/EW/Top Factories,
October 5 to 7, 2022, DESY



CLUSTER OF EXCELLENCE
QUANTUM UNIVERSE

¹Largely based on [arXiv:2003.12391](https://arxiv.org/abs/2003.12391)

SUSY: What *do* we know ?

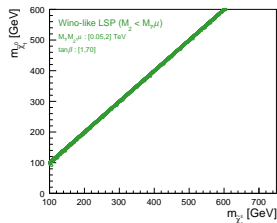
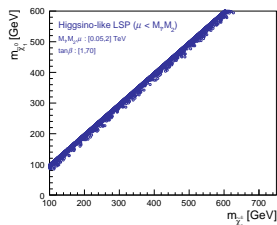
Naturalness, hierarchy, DM, $g-2$ all prefers **light electro-weak** sector.

- Except for 3d gen. squarks, **the coloured sector** - where pp machines excel - **doesn't enter the game**.
- If the LSP is higgsino or wino, EW sector is "compressed". Only for bino-LSP can the difference be large.
- So, most sparticle-decays are **via cascades**, with small $\Delta(M)$ at the end.
- For this, current limits from LHC are only for specific models, and **LEP2** sets the scene.

SUSY: What do we know ?

Naturalness, hierarchy, DM, g-2 all prefers **light electro-weak** sector.

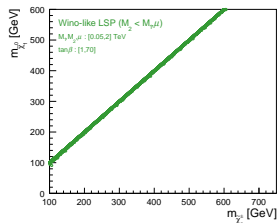
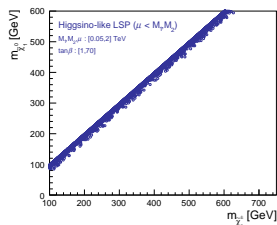
- Except for 3d gen. squarks, **the coloured sector** - where pp machines excel - **doesn't enter the game**.
- If the LSP is higgsino or wino, EW sector is “compressed”. Only for bino-LSP can the difference be large.
- So, most sparticle-decays are **via cascades**, with small $\Delta(M)$ at the end.
- For this, current limits from LHC are only for specific models, and LEP2 sets the scene.



SUSY: What do we know ?

Naturalness, hierarchy, DM, g-2 all prefers **light electro-weak** sector.

- Except for 3d gen. squarks, **the coloured sector** - where pp machines excel - **doesn't enter the game**.
- If the LSP is higgsino or wino, EW sector is “compressed”. Only for bino-LSP can the difference be large.
- So, most sparticle-decays are **via cascades**, with small $\Delta(M)$ at the end.
- For this, current limits from LHC are only for specific models, and **LEP2** sets the scene.



What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
- sfermions not NLSP (**idem**, except $\tilde{\tau}$ but even worse for $pp \dots$)
- Then: LSP is Bino, Wino, or Higgsino (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, up to values that makes the bosinos out-of-reach for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**
- Use SPheno 4.0.5beta to calculate spectra and BR:s, and use Whizard 2.8.0 for cross-sections

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for pp ...)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, up to values that makes the bosinos out-of-reach for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**
- Use SPheno 4.0.5beta to calculate spectra and BR:s, and use Whizard 2.8.0 for cross-sections

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for $pp \dots$)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, **up to values that makes the bosinos out-of-reach** for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**
- Use SPheno 4.0.5beta to calculate spectra and BR:s, and use Whizard 2.8.0 for cross-sections

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for $pp \dots$)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, **up to values that makes the bosinos out-of-reach** for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**
- Use SPheno 4.0.5beta to calculate spectra and BR:s, and use Whizard 2.8.0 for cross-sections

What *would* be seen at colliders in the worst case?

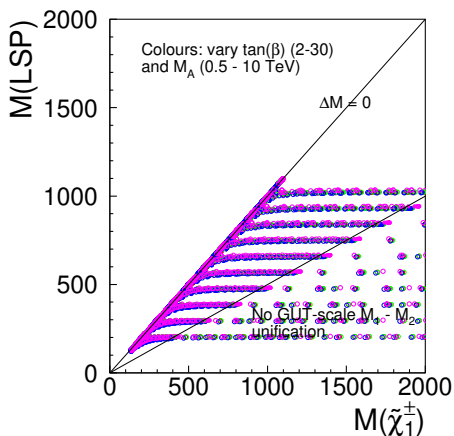
- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for $pp \dots$)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, **up to values that makes the bosinos out-of-reach** for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**
- Use SPheno 4.0.5beta to calculate spectra and BR:s, and use Whizard 2.8.0 for cross-sections

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for pp ...)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ What happens with spectra, cross-sections, BRs when exploiting this “cube”?
- Consider **an** p to values that makes the bosinos **out-of-reach** for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**
- Use `SPheno 4.0.5beta` to calculate spectra and BR:s, and use `Whizard 2.8.0` for cross-sections

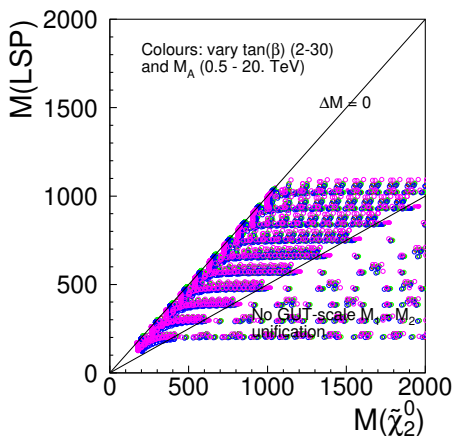
Aspects of the spectrum

- M_{LSP} vs. $M_{\tilde{\chi}_1^\pm}$
- M_{LSP} vs. $M_{\tilde{\chi}_2^0}$
- Colours indicate different settings of the secondary parameters (lesson is that they don't matter much...)
- Open circles indicated cases where GUT-scale unification of M_1 and M_2 is not possible



Aspects of the spectrum

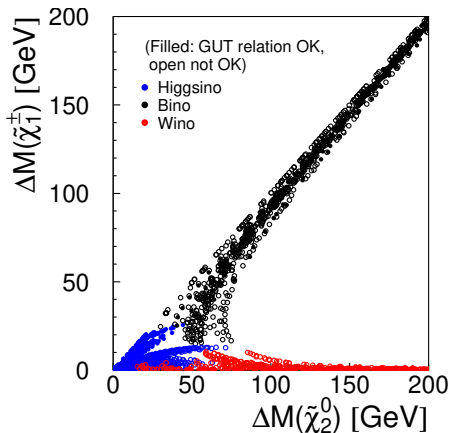
- M_{LSP} vs. $M_{\tilde{\chi}_1^\pm}$
- M_{LSP} vs. $M_{\tilde{\chi}_2^0}$
- Colours indicate different settings of the secondary parameters (lesson is that they don't matter much...)
- Open circles indicated cases where GUT-scale unification of M_1 and M_2 is not possible



Aspects of the spectrum

Another angle: $\Delta(M)$ for $\tilde{\chi}_1^\pm$ vs. that of $\tilde{\chi}_2^0$: Important experimentally

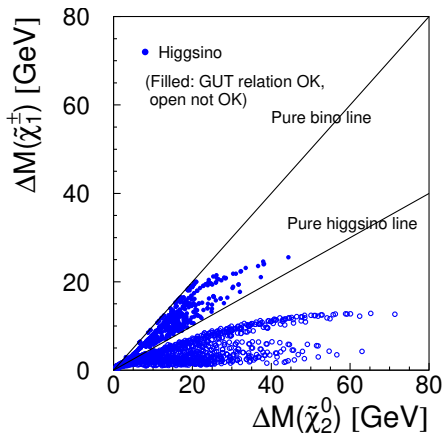
- Three regions:
 - Bino: Both the same, but can be anything.
 - Wino: $\Delta_{\tilde{\chi}_1^\pm}$ small, while $\Delta_{\tilde{\chi}_2^0}$ can be anything.
 - Higgsino: Both often small
- But note, seldom on the “Higgsino line”, ie. when the chargino is *exactly* in the middle of mass-gap between the first and second neutralino.

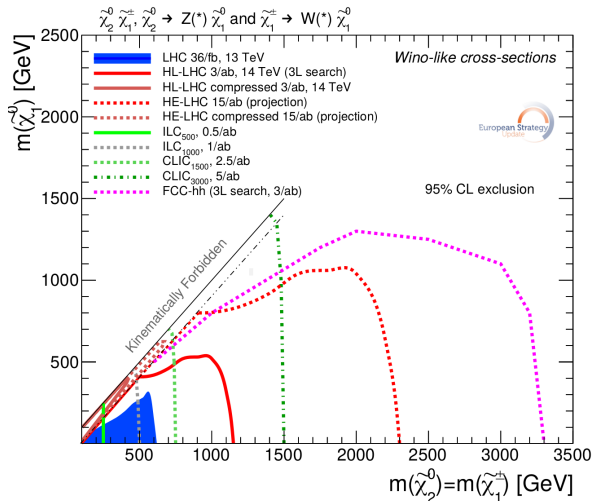


Aspects of the spectrum

Another angle: $\Delta(M)$ for $\tilde{\chi}_1^\pm$ vs. that of $\tilde{\chi}_2^0$: Important experimentally

- Three regions:
 - Bino: Both the same, but can be anything.
 - Wino: $\Delta_{\tilde{\chi}_1^\pm}$ small, while $\Delta_{\tilde{\chi}_2^0}$ can be anything.
 - Higgsino: Both often small
- But note, **seldom on the “Higgsino line”**, ie. when the chargino is *exactly* in the middle of mass-gap between the first and second neutralino.

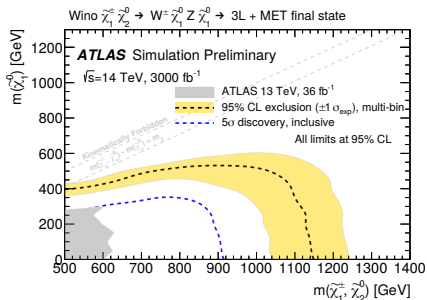


SUSY In The Briefing-book: Bino LSP (ie. large $\Delta(M)$)

NB: e^+e^- curves are **certain discovery**, pp are **possible exclusion** !!!

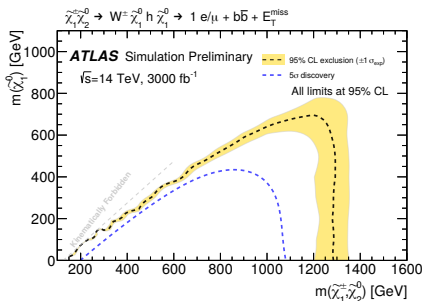
SUSY In The Briefing-book: Bino LSP - Sources

- ATLAS-PHYS-PUB-2018-048, ATLAS HL-LHC projection, extrapolated (up *and down*)
- This is for the best mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .
- Why is the decay-mode an issue? Here's why :
 - Vary signs of μ , M_1 , and M_2
- So: The exclusion-region is the *intersection* of the two plots, not the *union*!



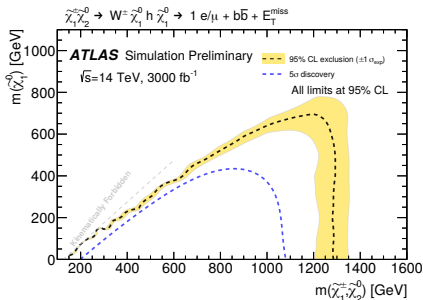
SUSY In The Briefing-book: Bino LSP - Sources

- ATLAS-PHYS-PUB-2018-048, ATLAS HL-LHC projection, extrapolated (up *and* down)
- This is for the best mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .
- Why is the decay-mode an issue? Here's why :
 - Vary signs of μ , M_1 , and M_2
- So: The exclusion-region is the *intersection* of the two plots, not the *union*!



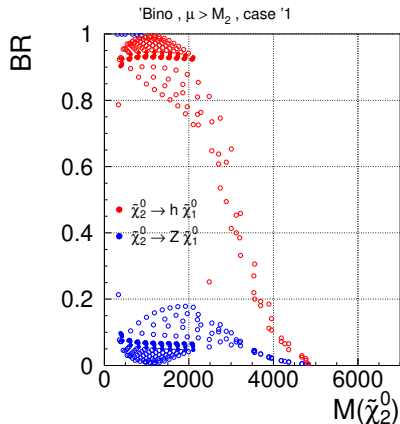
SUSY In The Briefing-book: Bino LSP - Sources

- ATLAS-PHYS-PUB-2018-048, ATLAS HL-LHC projection, extrapolated (up *and* down)
- This is for the best mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .
- Why is the decay-mode an issue? **Here's why** :
 - Vary signs of μ , M_1 , and M_2
- So: The exclusion-region is the *intersection* of the two plots, not the *union*!



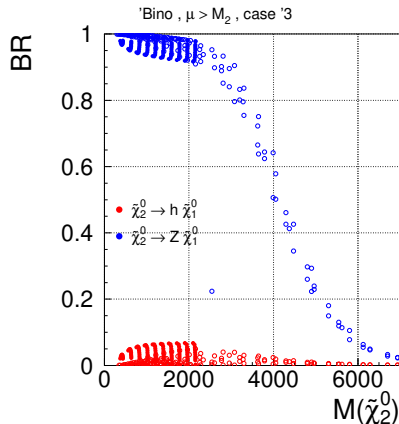
SUSY In The Briefing-book: Bino LSP - Sources

- ATLAS-PHYS-PUB-2018-048, ATLAS HL-LHC projection, extrapolated (up *and down*)
- This is for the best mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .
- Why is the decay-mode an issue? *Here's why* :
 - Vary signs of μ , M_1 , and M_2
- So: The exclusion-region is the *intersection* of the two plots, not the *union*!



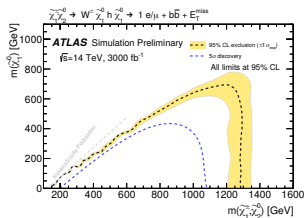
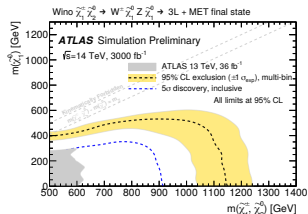
SUSY In The Briefing-book: Bino LSP - Sources

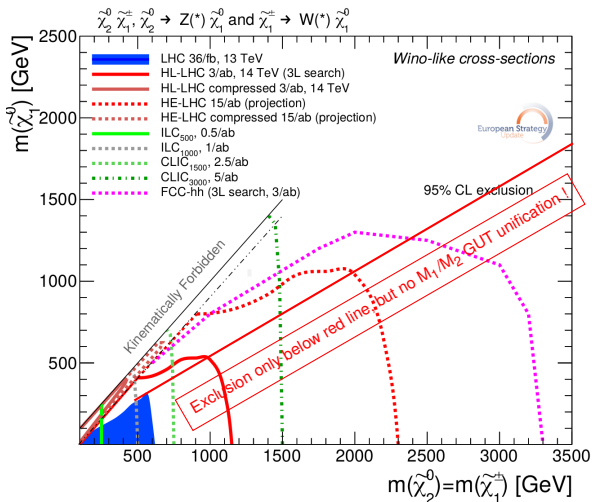
- ATLAS-PHYS-PUB-2018-048, ATLAS HL-LHC projection, extrapolated (up *and* down)
- This is for the best mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .
- Why is the decay-mode an issue? *Here's why* :
 - Vary signs of μ , M_1 , and M_2
- So: The exclusion-region is the *intersection* of the two plots, not the *union*!



SUSY In The Briefing-book: Bino LSP - Sources

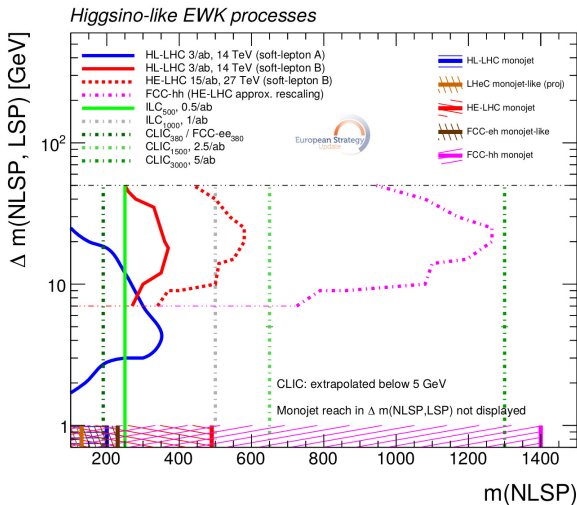
- ATLAS-PHYS-PUB-2018-048, ATLAS HL-LHC projection, extrapolated (up *and* down)
- This is for the best mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .
- Why is the decay-mode an issue? *Here's why* :
 - Vary signs of μ , M_1 , and M_2
- So: The exclusion-region is the *intersection* of the two plots, not the *union*!



SUSY In The Briefing-book: Bino LSP (ie. large Δ_M)

NB: e^+e^- curves are **certain discovery**, pp are **possible exclusion** !!!

SUSY In The Briefing-book: Wino/Higgsino LSP



SUSY In The Briefing-book: Wino/Higgsino LSP - Soft lepton Sources

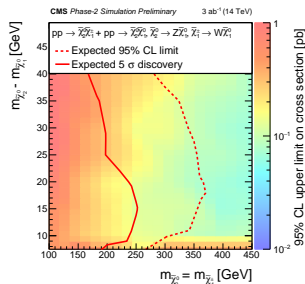
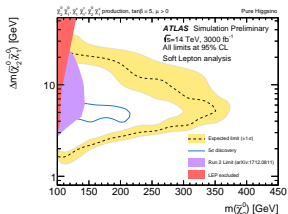
● Soft lepton analysis:

- ATLAS HL-LHC projection
ATL-PHYS-PUB-2018-031.
- CMS HE-LHC projection
(and extrapolated to FCChh)
CMS-PAS-FTR-18-001.

● Crucial experimental issue: lepton ID

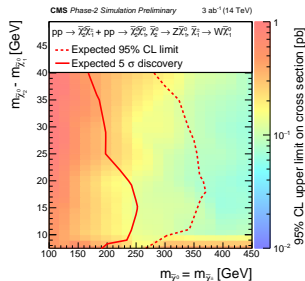
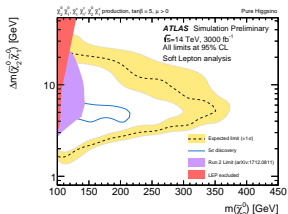
- To separate $e/\mu/\pi$, particles must reach calorimeter.
- ... and FCChh detector has both higher B-field and calorimeter radius (and CMS has that wrt. ATLAS)

● Unlikely that lower $\Delta(M)$ will be excluded in future.



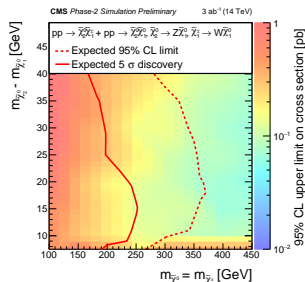
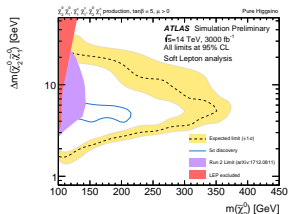
SUSY In The Briefing-book: Wino/Higgsino LSP - Soft lepton Sources

- Soft lepton analysis:
 - ATLAS HL-LHC projection
ATL-PHYS-PUB-2018-031.
 - CMS HE-LHC projection
(and extrapolated to FCChh)
CMS-PAS-FTR-18-001.
- Crucial experimental issue: lepton ID
 - To separate $e/\mu/\pi$, particles must reach calorimeter.
 - ... and FCChh detector has both higher B-field and calorimeter radius (and CMS has that wrt. ATLAS)
- Unlikely that lower $\Delta(M)$ will be excluded in future.



SUSY In The Briefing-book: Wino/Higgsino LSP - Soft lepton Sources

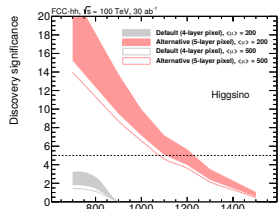
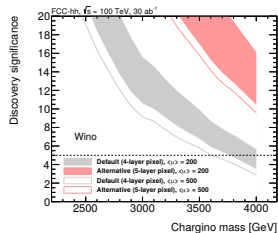
- Soft lepton analysis:
 - ATLAS HL-LHC projection
ATL-PHYS-PUB-2018-031.
 - CMS HE-LHC projection
(and extrapolated to FCChh)
CMS-PAS-FTR-18-001.
- Crucial experimental issue: lepton ID
 - To separate $e/\mu/\pi$, particles must reach calorimeter.
 - ... and FCChh detector has both higher B-field and calorimeter radius (and CMS has that wrt. ATLAS)
- Unlikely that lower $\Delta(M)$ will be excluded in future.



SUSY In The Briefing book: Wino/Higgsino LSP - Very low $\Delta(M)$ sources

(Don't look at the pink curves - they correspond to a detector that is never considered anywhere else i the CDR)

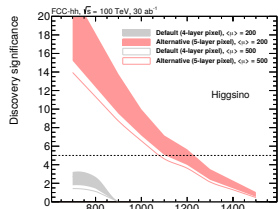
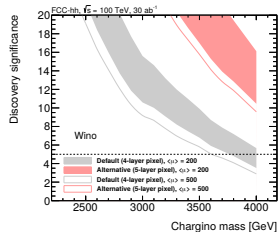
- The “Disappearing tracks” was done by FCChh (in the CDR)
 - FCChh-detector
 - FCChh-ish PU (but still too small: 500 vs. CDR number 955)
 - Assumes **only SM loops** for mass-splitting, i.e. not SUSY mixing: The “other two” mass-parameters very large.
 - For higgsinos: Only *just* reaches 2σ
- A study of the “mono-X” method was done in [arXiv:1805.00015](https://arxiv.org/abs/1805.00015), but it is too rudimentary in the experimental aspects to allow for any conclusions.



SUSY In The Briefing book: Wino/Higgsino LSP - Very low $\Delta(M)$ sources

(Don't look at the pink curves - they correspond to a detector that is never considered anywhere else i the CDR)

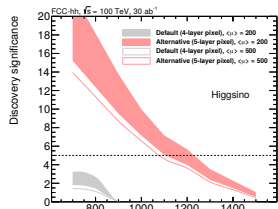
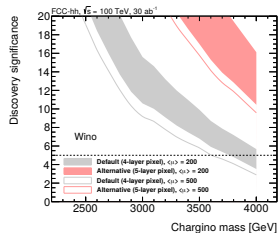
- The “Disappearing tracks” was done by FCChh (in the CDR)
 - FCChh-detector
 - FCChh-ish PU (but still too small: 500 vs. CDR number 955)
 - Assumes **only SM loops** for mass-splitting, i.e. not SUSY mixing: The “other two” mass-parameters very large.
 - For higgsinos: Only *just* reaches 2σ
- A study of the “mono-X” method was done in [arXiv:1805.00015](https://arxiv.org/abs/1805.00015), but it is too rudimentary in the experimental aspects to allow for any conclusions.



SUSY In The Briefing book: Wino/Higgsino LSP - Very low $\Delta(M)$ sources

(Don't look at the pink curves - they correspond to a detector that is never considered anywhere else i the CDR)

- The “Disappearing tracks” was done by FCChh (in the CDR)
 - FCChh-detector
 - FCChh-ish PU (but still too small: 500 vs. CDR number 955)
 - Assumes **only SM loops** for mass-splitting, i.e. not SUSY mixing: The “other two” mass-parameters very large.
 - For higgsinos: Only *just* reaches 2σ
- A study of the “mono-X” method was done in arXiv:1805.00015, but it is too rudimentary in the experimental aspects to allow for any conclusions.



Key element for “Disappearing tracks”: $\Delta(M)$

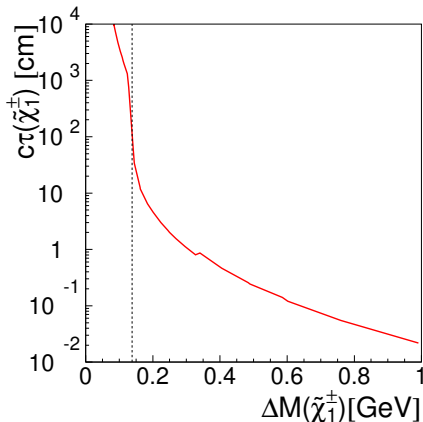
Why is this important?

- Because c_{τ} depends on $\Delta(M)$, and c_{τ} needs to be macroscopic to get “Disappearing tracks”.
- Cf. [arXiv:1712.02118](https://arxiv.org/abs/1712.02118) where ATLAS found that c_{τ} needs to be ~ 6 cm.
- c_{τ} for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.

Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

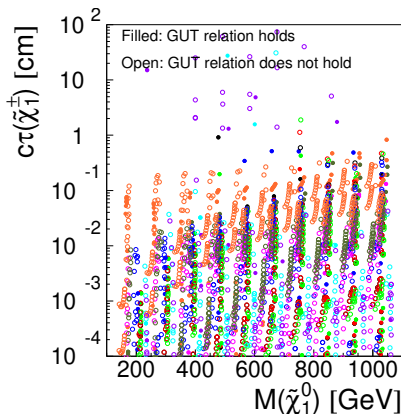
- Because $c\tau$ depends on $\Delta(M)$, and $c\tau$ needs to be macroscopic to get “Disappearing tracks”.
- Cf. arXiv:1712.02118 where ATLAS found that $c\tau$ needs to be ~ 6 cm.
- $c\tau$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.



Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

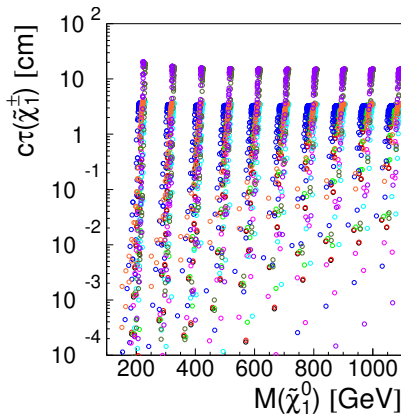
- Because $c\tau$ depends on $\Delta(M)$, and $c\tau$ needs to be macroscopic to get “Disappearing tracks”.
- Cf. arXiv:1712.02118 where ATLAS found that $c\tau$ needs to be ~ 6 cm.
- $c\tau$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.



Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

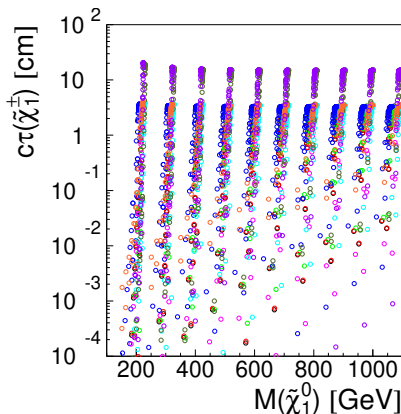
- Because c_{τ} depends on $\Delta(M)$, and c_{τ} needs to be macroscopic to get “Disappearing tracks”.
- Cf. arXiv:1712.02118 where ATLAS found that c_{τ} needs to be ~ 6 cm.
- c_{τ} for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.



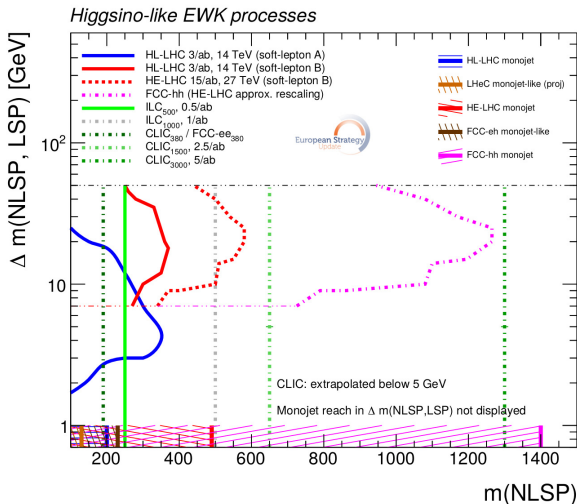
Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

- Because c_{τ} depends on $\Delta(M)$, and c_{τ} needs to be macroscopic to get “Disappearing tracks”.
- Cf. [arXiv:1712.02118](https://arxiv.org/abs/1712.02118) where ATLAS found that c_{τ} needs to be ~ 6 cm.
- c_{τ} for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.

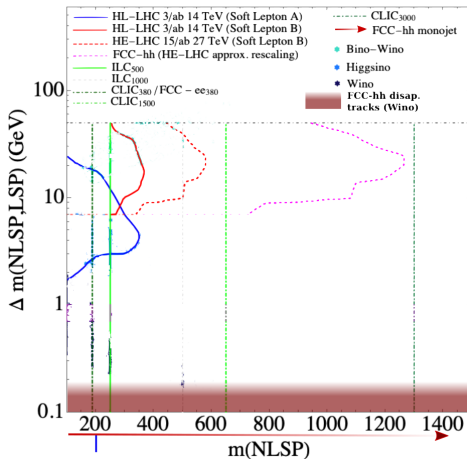


SUSY In The Briefing-book: Wino/Higgsino LSP

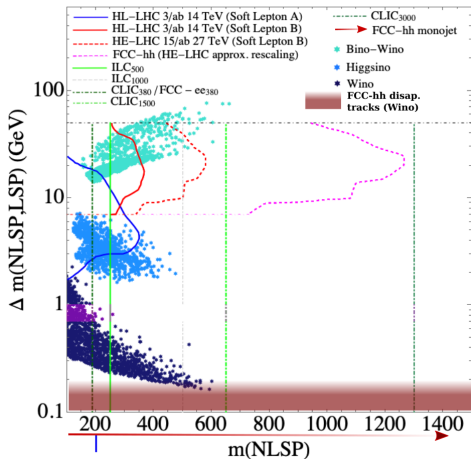


So: Disappearing tracks exclusion is actually off the scale !

SUSY In The Briefing-book: Re-boot

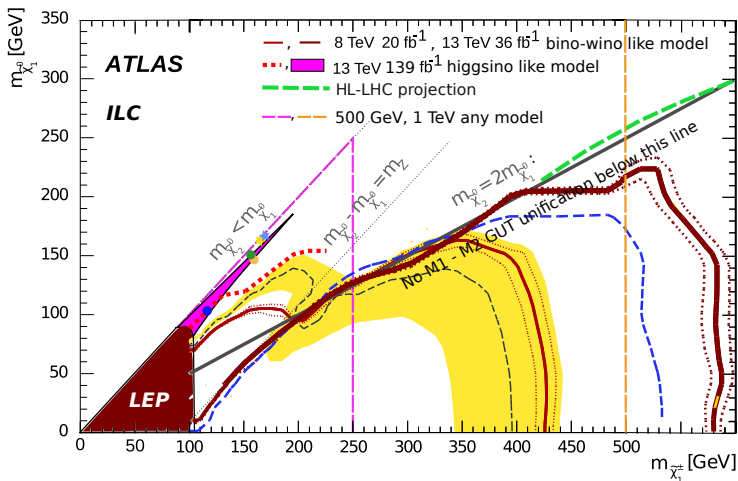


SUSY In The Briefing-book: Re-boot



With models that are consistent with $g-2$ and no over-production of DM
 From [arXiv:2103.13403](https://arxiv.org/abs/2103.13403).

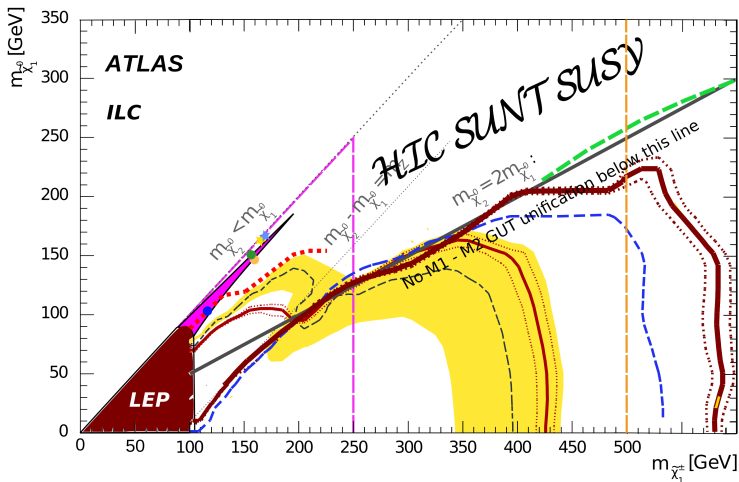
Summary: SUSY - All-in-one



ATLAS Eur Phys J C 78,995 (2018), Phys Rev D 101,052002 (2020), arXiv:2106.01676;

ATLAS HL-LHC ATL-PHYS-PUB-2018-048; ILC arXiv:2002.01239; LEP LEP LEPSUSYWG/02-04.1

Summary: SUSY - All-in-one



ATLAS Eur Phys J C 78,995 (2018), Phys Rev D 101,052002 (2020), arXiv:2106.01676;

ATLAS HL-LHC ATL-PHYS-PUB-2018-048; ILC arXiv:2002.01239; LEP LEP LEP SUSYWG/02-04.1

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - discovery potential to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full discovery **and** exclusion potential up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - discovery potential to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full discovery **and** exclusion potential up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - discovery potential to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full discovery **and** exclusion potential up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full **discovery and exclusion** potential up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full **discovery and exclusion potential** up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full **discovery and** exclusion potential up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: there will always be loopholes.
- Future TeV-scale ee machines - on the other hand - have
 - Full **discovery and exclusion** potential up to the kinematic limit

Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have

Take-home message

- discov
 - but - loop
 - Future Te
 - Full c
- **Without a TeV scale lepton-collider**, we would not be able to exclude SUSY further than today at the end of this century. **LEP2++ would be the final word.**
 - Except if a future pp machine discovers SUSY, which is a **problem we'd like to have!**

ays be

mit

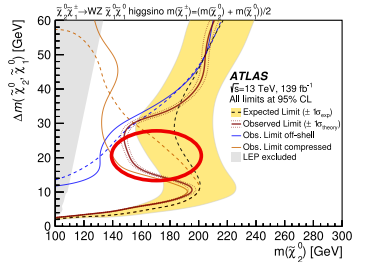
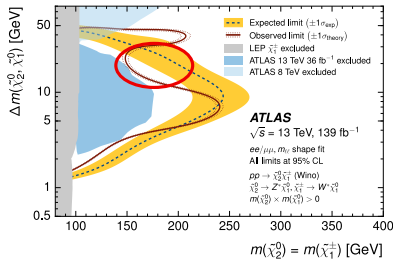
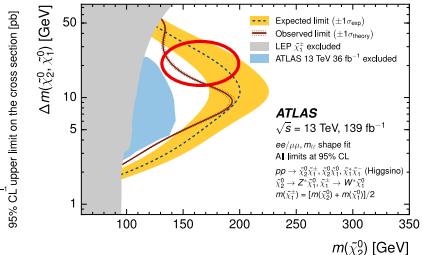
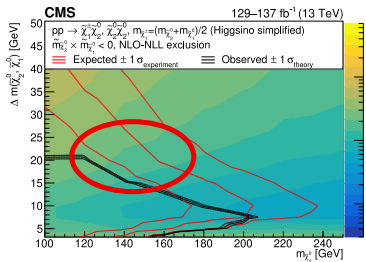
Conclusions

- SUSY is **not** excluded.
- Even Plain vanilla SUSY is **not** excluded.
- HL-LHC might well discover SUSY, because future pp machines have

Take-home message

- discov
 - but - loop
 - Future Te
 - Full c
- **Without a TeV scale lepton-collider**, we would not be able to exclude SUSY further than today at the end of this century. **LEP2++ would be the final word.**
 - **Except** if a future pp machine discovers SUSY, which is a **problem we'd like to have!**
- ays be
mit

LHC Run 3 teaser: Maybe...





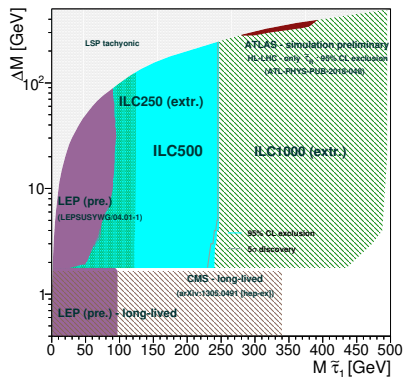
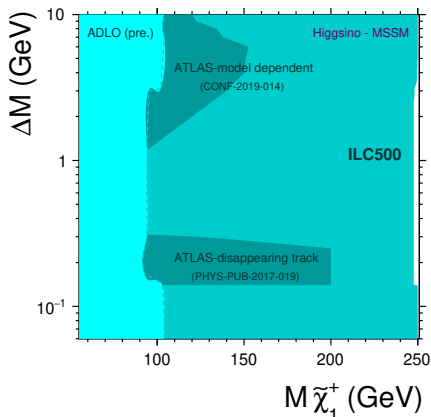
That's all Folks!

Backup

BACKUP SLIDES

Summary: ILC projection on Higgsinos and $\tilde{\tau}$'s

From arXiv:2002.01239

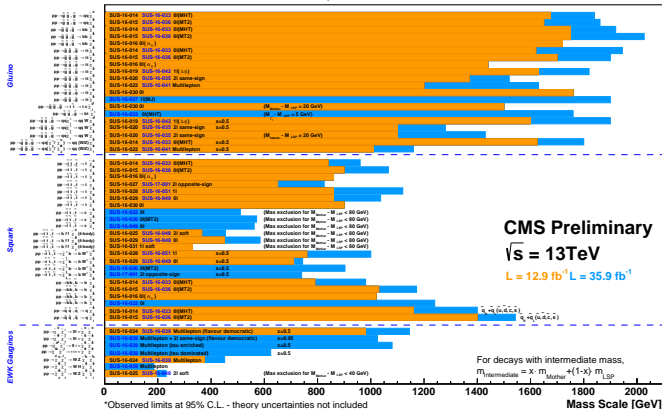


From arXiv:2105.08616

SUSY@LHC: Does this make us depressed ?

Selected CMS SUSY Results* - SMS Interpretation

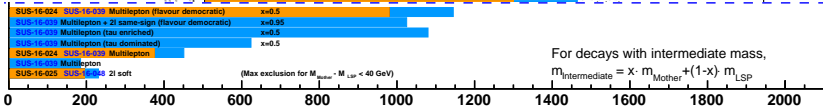
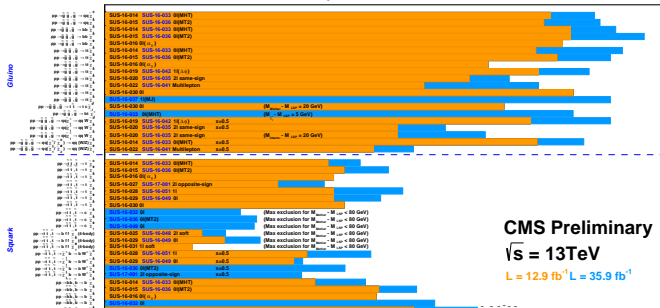
ICHEP '16 - Moriond '17



SUSY@LHC: No! Read the fine-print!

Selected CMS SUSY Results* - SMS Interpretation

ICHEP '16 - Moriond '17

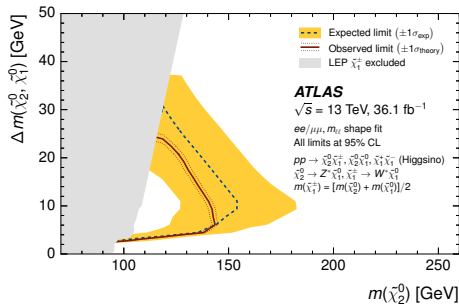


*Observed limits at 95% C.L. - theory uncertainties not included

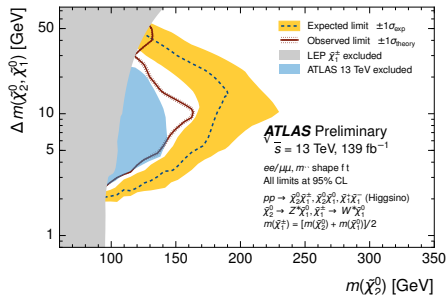
Only a selection of available mass limits. Probe *up to* the quoted mass limit for $m_{\text{LSP}} = 0 \text{ GeV}$ unless stated otherwise

Latest Atlas (13 TeV, 36 and 139 fb⁻¹) on higgsinos

arXiv:1803.02762

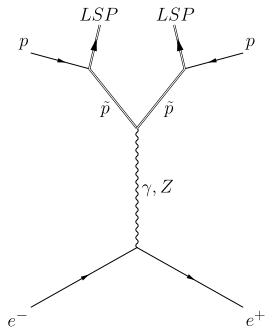


ATLAS-CONF-2019-01



Loop-hole free SUSY searches

- All is **known** for given masses, due to SUSY-principle: “sparticles couples as particles”.
- This doesn't depend on the SUSY breaking mechanism !
- Obviously: There is **one** NLSP.

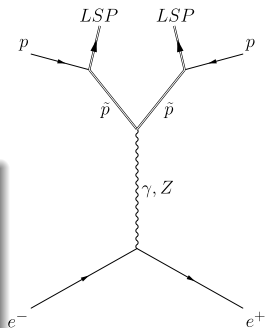


Loop-hole free SUSY searches

- All is **known** for given masses, due to SUSY-principle: “sparticles couples as particles”.
- This doesn't depend on the SUSY breaking mechanism !
- Obviously: There is **one** NLSP.

So, at an LC :

- Model **independent** exclusion/ discovery reach in $M_{NLSP} - M_{LSP}$ plane.
- Repeat for **all** NLSP:s.
- **Cover entire parameter-space in a hand-full of plots**
- NLSP search \leftrightarrow “simplified models” @ LHC!

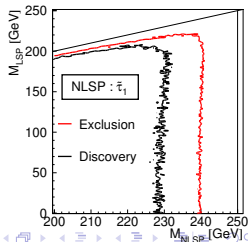
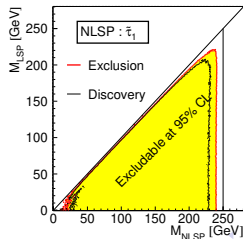
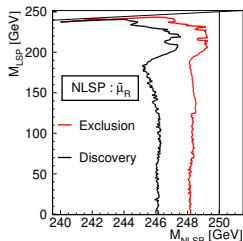
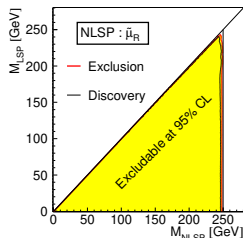


Simplified models

- Simplified methods at hadron and lepton machines are **different beasts**.
- At lepton machines they are quite **model independent**, at LHC **model dependent**.
- A few examples (M.B. arXiv:1308.1461)
 - $\tilde{\mu}_R$ NLSP
 - $\tilde{\tau}_1$ NLSP (minimal σ).

Simplified models

- Simplified methods at hadron and lepton machines are **different beasts**.
- At lepton machines they are quite **model independent**, at LHC **model dependent**.
- A few examples (M.B. arXiv:1308.1461)
 - $\tilde{\mu}_R$ NLSP
 - $\tilde{\tau}_1$ NLSP (minimal σ).



Simplified models

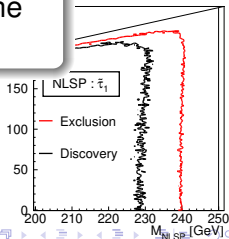
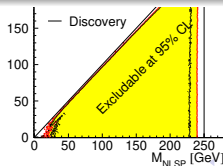
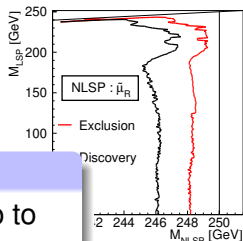
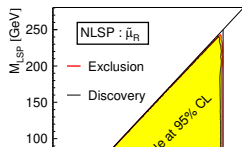
- Simplified methods at hadron and lepton machines are **different beasts**.

- At lepton machines they are called **At ILC**

independent Both **discover** and **exclude** NLSPs up to **model dependent some GeV**:s from the kinematic limit,

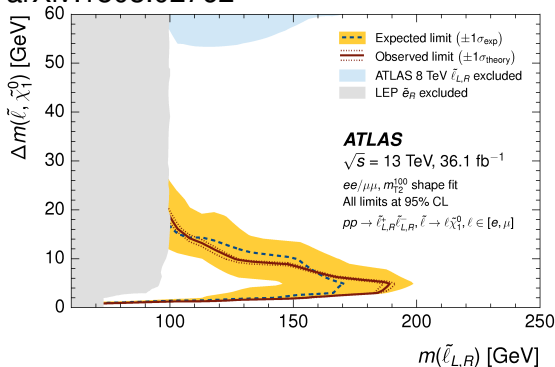
- A few examples whatever the NLSP is, and whatever the **arXiv:1308.1461**) rest of the spectrum is!

- $\tilde{\mu}_R$ NLSP
- $\tilde{\tau}_1$ NLSP (minimal σ).

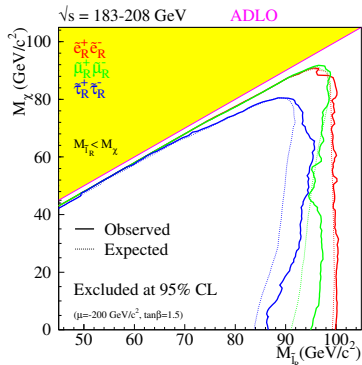


Latest Atlas (13 TeV, 36 fb⁻¹) and LEP on sleptons

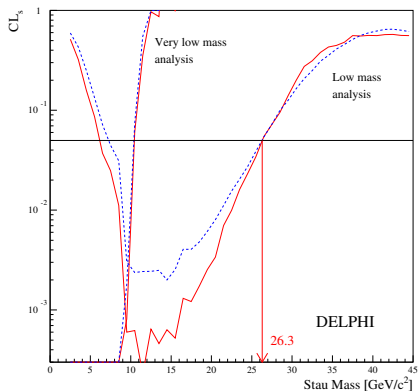
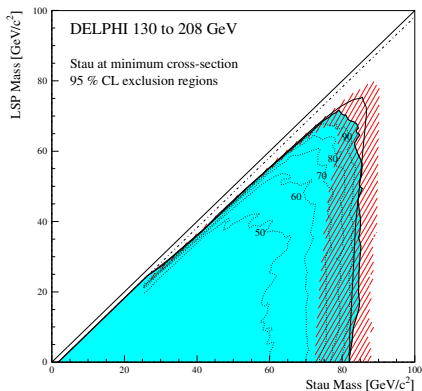
arXiv:1803.02762



This is a *combined* limit, assuming $\tilde{\mu}_L, \tilde{\mu}_R, \tilde{e}_L$ and \tilde{e}_L all have the **same mass** !!!



This is $\tilde{e}_R, \tilde{\mu}_R$ and $\tilde{\tau}_R$ *only*, separately!

In real life: LEP $\tilde{\tau}$ limits

NB: a $\tilde{\tau}$ as light as 26.3 GeV is *not* excluded!

In real life: LEP $\tilde{\tau}$ limits

With 1000 times the luminosity and no trigger, the ILC at 250 will push the limits for all possible NLSPs to close to 125 GeV, and $\Delta(M) \approx 0$. The area covered will \sim double the LEP ones. They are in the most compelling region of parameter-space.

- These will be rock-solid limits.
- Or discoveries!



NB: a $\tilde{\tau}$ as light as 26.3 GeV is *not* excluded!

In real life: LEP $\tilde{\tau}$ limits

With 1000 times the luminosity and no trigger, the ILC at 250 will push the limits for all possible NLSPs to close to 125 GeV, and $\Delta(M) \approx 0$. The area covered will \sim double the LEP ones. They are in the most compelling region of parameter-space.

- These will be rock-solid limits.
- Or discoveries!



NB: a $\tilde{\tau}$ as light as 26.3 GeV is *not* excluded!

Why compressed spectra ? Natural SUSY: Light, degenerate higgsinos

Why would one expect the spectrum to be compressed ?

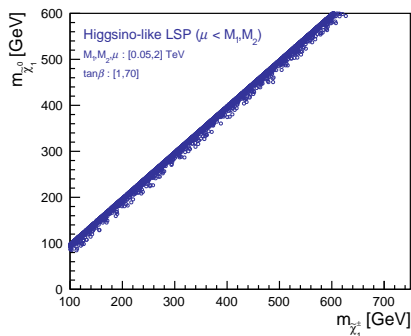
- Natural SUSY:

- $m_Z^2 = 2 \frac{m_{H_u}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - 2 |\mu|^2$
- \Rightarrow Low fine-tuning \Rightarrow
 $\mu = \mathcal{O}(\text{weak scale})$.

- Wino-like LSP: Same conclusion.
- Only for Bino-like LSP, non-compressed occurs
- But also: the data ...

quite generic:

Parameter-scan by T. Tanabe:



Why compressed spectra ? Natural SUSY: Light, degenerate higgsinos

Why would one expect the spectrum to be compressed ?

- Natural SUSY:

- $m_Z^2 = 2 \frac{m_{H_u}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - 2 |\mu|^2$
- \Rightarrow Low fine-tuning \Rightarrow
 $\mu = \mathcal{O}(\text{weak scale})$.

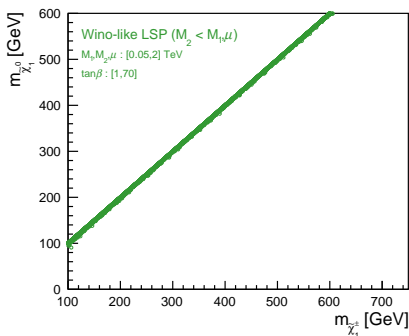
- Wino-like LSP: Same conclusion.

- Only for Bino-like LSP, non-compressed occurs

- But also: the data ...

quite generic:

Parameter-scan by T. Tanabe:



Why compressed spectra ? Natural SUSY: Light, degenerate higgsinos

Why would one expect the spectrum to be compressed ?

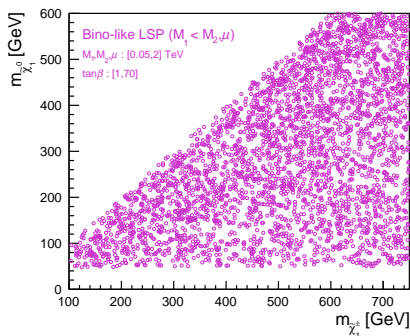
- Natural SUSY:

- $m_Z^2 = 2 \frac{m_{H_u}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - 2 |\mu|^2$
- \Rightarrow Low fine-tuning \Rightarrow
 $\mu = \mathcal{O}(\text{weak scale})$.

- Wino-like LSP: Same conclusion.
- Only for Bino-like LSP, non-compressed occurs
- But also: the data ...

quite generic:

Parameter-scan by T. Tanabe:



Why compressed spectra ? Natural SUSY: Light, degenerate higgsinos

Why would one expect the spectrum to be compressed ?

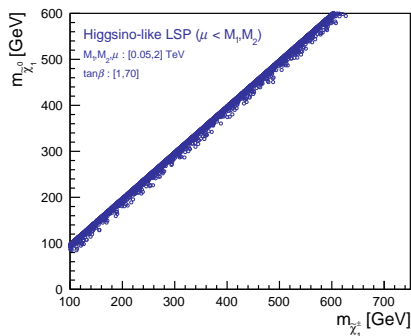
- Natural SUSY:

- $m_Z^2 = 2 \frac{m_{H_u}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - 2 |\mu|^2$
- \Rightarrow Low fine-tuning \Rightarrow
 $\mu = \mathcal{O}(\text{weak scale})$.

- Wino-like LSP: Same conclusion.
- Only for Bino-like LSP, non-compressed occurs
- But also: the data ...

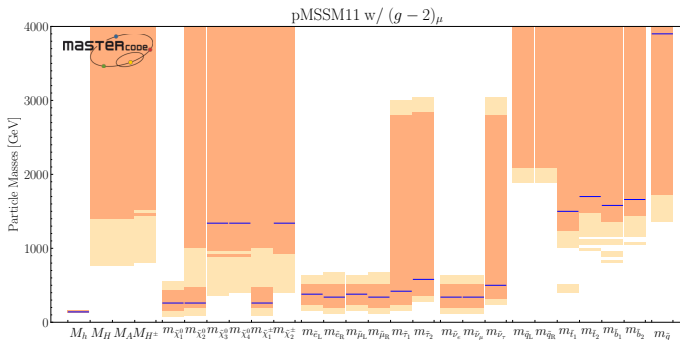
quite generic:

Parameter-scan by T. Tanabe:



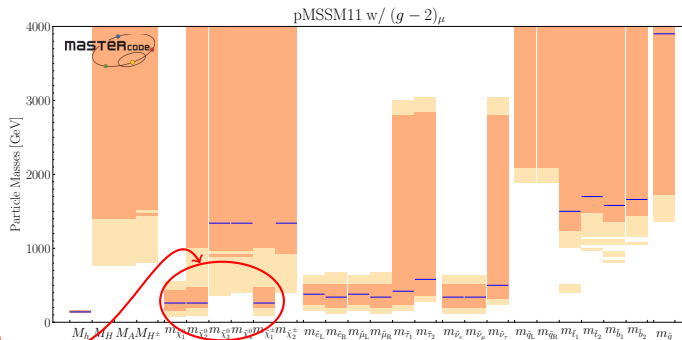
One approach: Global fits with prejudice

pMSSM11 fit by **Mastercode** to
 LHC13/LEP/**g-2/DM(=100% LSP)**/precision observables
 (arXiv:1710.11091):



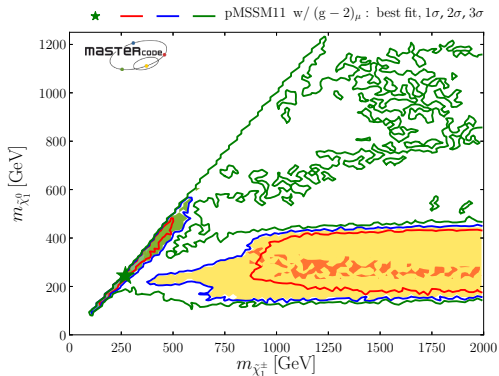
One approach: Global fits with prejudice

pMSSM11 fit by **Mastercode** to
 LHC13/LEP/**g-2/DM(=100% LSP)**/precision observables
 (arXiv:1710.11091):



One approach: Global fits with prejudice

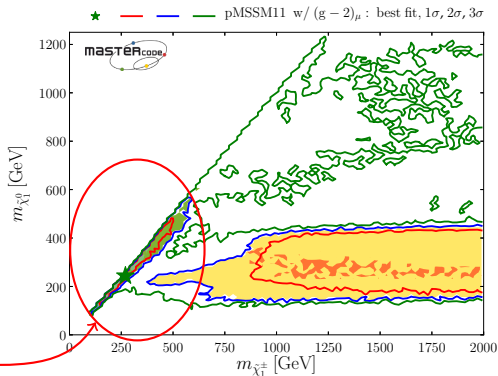
pMSSM11 fit by **Mastercode** to
 LHC13/LEP/g-2/DM(=100% LSP)/precision observables
 (arXiv:1710.11091):



$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0}$ plane

One approach: Global fits with prejudice

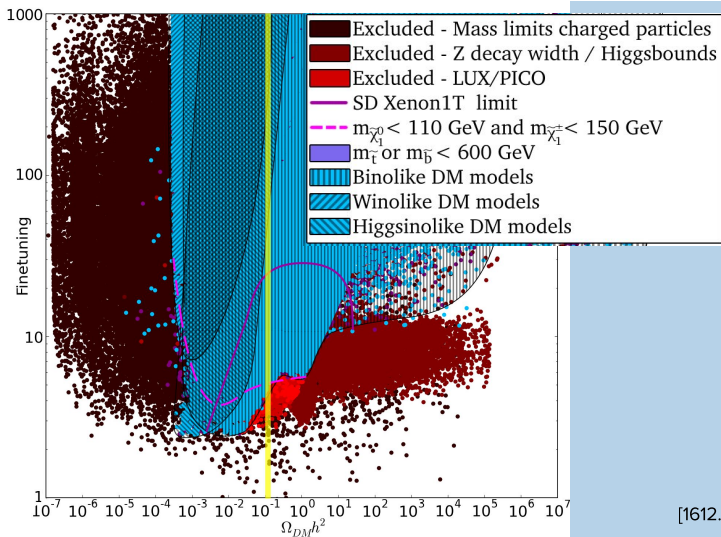
pMSSM11 fit by **Mastercode** to
LHC13/LEP/g-2/DM(=100% LSP)/precision observables
(arXiv:1710.11091):



Low $\Delta(M)$!

$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0}$ plane

One approach: Global fits with prejudice

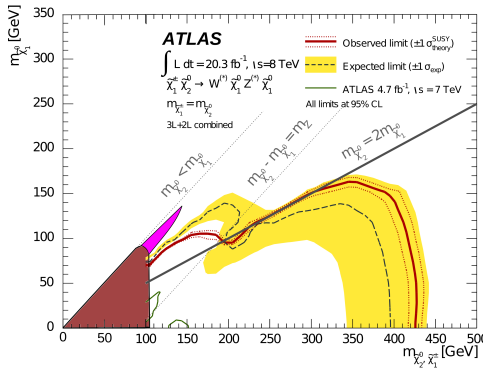


Melissa van Beekveld

[1612.06333]

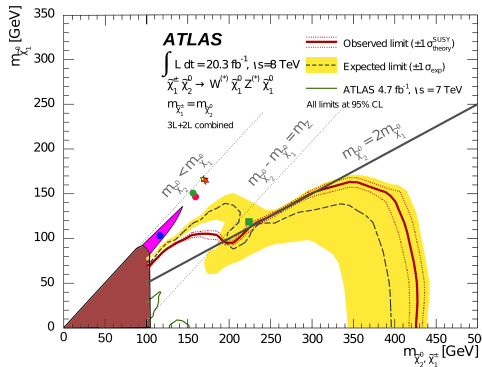
Compare LHC (Atlas) & ILC

- On the 7 TeV plot, with LEP (brown) and the low $\Delta(M)$ search (magenta)...
- At ILC: Various benchmarks studied w/ detailed simulation:
 $M_{\tilde{\chi}_1^0} = 100\text{-}170$ GeV, $\Delta(M) = 0.8$ to 20 GeV.
- Projected discovery reaches for LHC, HL-LHC, ILC-500, and ILC-1000.



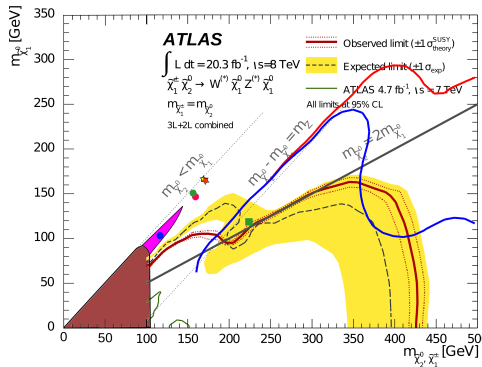
Compare LHC (Atlas) & ILC

- On the 7 TeV plot, with LEP (brown) and the low $\Delta(M)$ search (magenta)...
- At ILC: Various benchmarks studied w/ detailed simulation:
 $M_{\tilde{\chi}_1^0} = 100\text{-}170$ GeV, $\Delta(M) = 0.8$ to 20 GeV.
- Projected discovery reaches for LHC, HL-LHC, ILC-500, and ILC-1000.



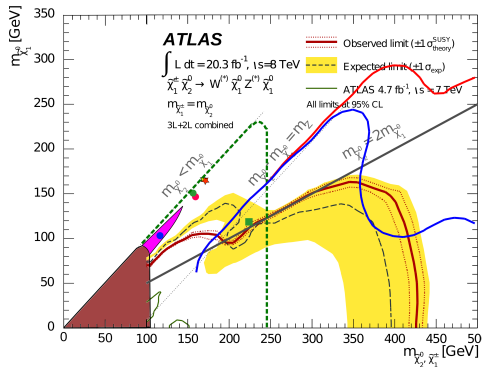
Compare LHC (Atlas) & ILC

- On the 7 TeV plot, with LEP (brown) and the low $\Delta(M)$ search (magenta)...
- At ILC: Various benchmarks studied w/ detailed simulation: $M_{\tilde{\chi}_1^0} = 100\text{-}170$ GeV, $\Delta(M) = 0.8$ to 20 GeV.
- Projected **discovery** reaches for LHC, HL-LHC, ILC-500, and ILC-1000.



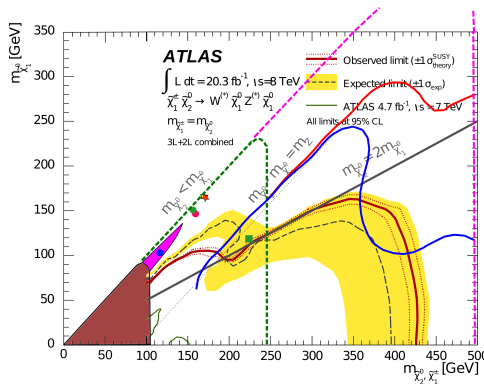
Compare LHC (Atlas) & ILC

- On the 7 TeV plot, with LEP (brown) and the low $\Delta(M)$ search (magenta)...
- At ILC: Various benchmarks studied w/ detailed simulation:
 $M_{\tilde{\chi}_1^0} = 100\text{-}170$ GeV, $\Delta(M) = 0.8$ to 20 GeV.
- Projected **discovery** reaches for LHC, HL-LHC, ILC-500, and ILC-1000.



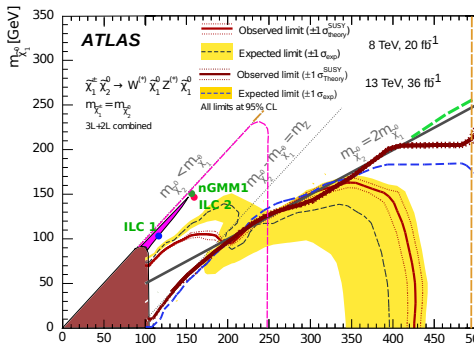
Compare LHC (Atlas) & ILC

- On the 7 TeV plot, with LEP (brown) and the low $\Delta(M)$ search (magenta)...
- At ILC: Various benchmarks studied w/ detailed simulation:
 $M_{\tilde{\chi}_1^0} = 100\text{-}170$ GeV, $\Delta(M) = 0.8$ to 20 GeV.
- Projected **discovery** reaches for LHC, HL-LHC, ILC-500, and ILC-1000.



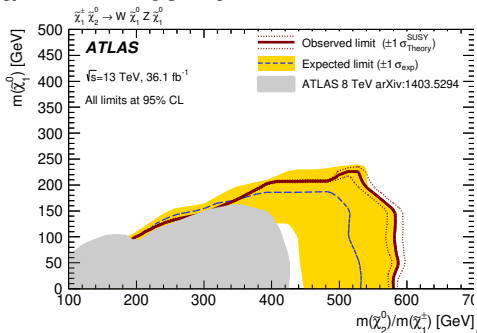
Compare LHC (Atlas) & ILC

- On the 7 TeV plot, with LEP (brown) and the low $\Delta(M)$ search (magenta)...
- At ILC: Various benchmarks studied w/ detailed simulation: $M_{\tilde{\chi}_1^0} = 100\text{-}170$ GeV, $\Delta(M) = 0.8$ to 20 GeV.
- Projected **discovery** reaches for LHC, HL-LHC, ILC-500, and ILC-1000.



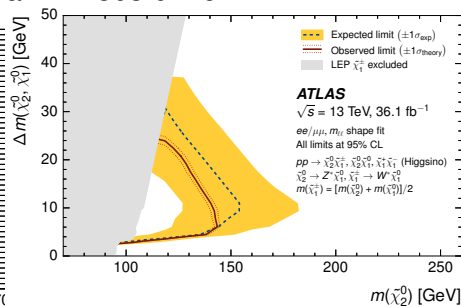
Latest Atlas (13 TeV, 36 fb⁻¹) on EWkinos

arXiv:1712.08119



\sim same analysis as shown in talk.
 Only extends below the $M_{\tilde{\chi}_2^0}$ (or $M_{\tilde{\chi}_1^\pm} > 2M_{\tilde{\chi}_2^0}$) line. *No progress in Higgsino region!*

arXiv:1803.02762

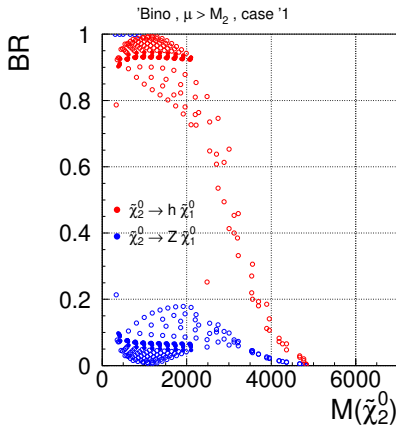


Same channel as in talk. Look at $\Delta(M) \sim 1 \text{ GeV}$ and $M_{\tilde{\chi}_2^0} \sim 160 \text{ GeV}$. The actual limit is the LEP one. *Wrongly represented!*

Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

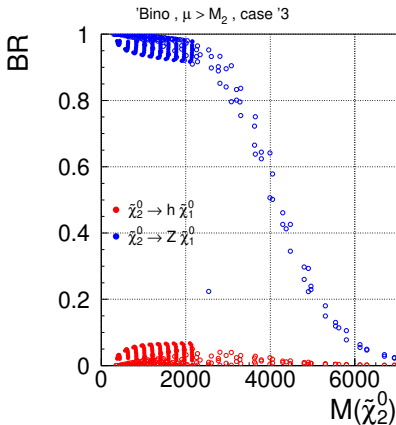
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

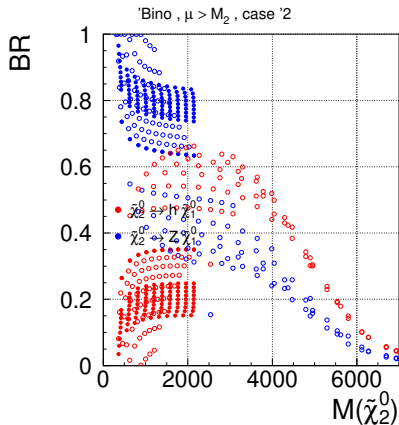
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

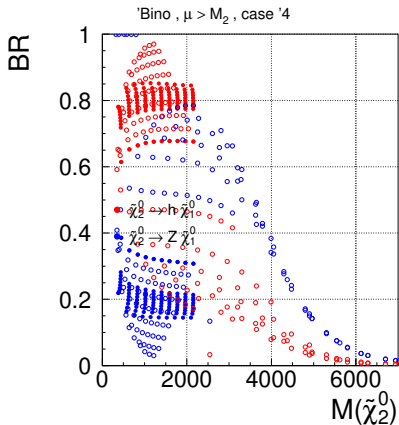
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

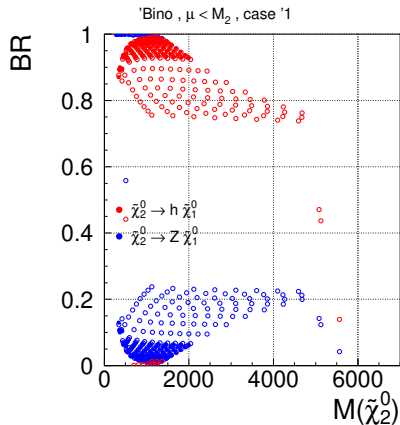
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

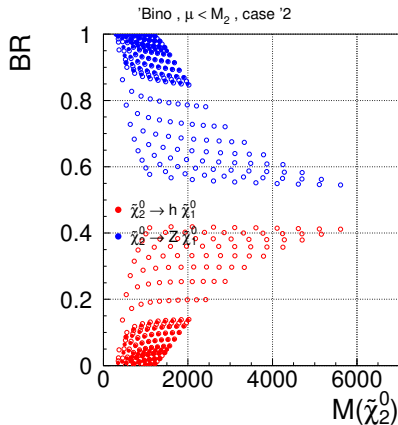
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

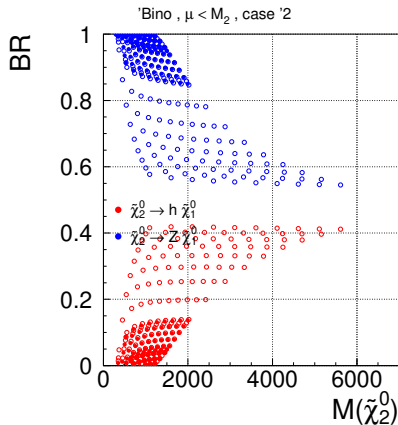
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

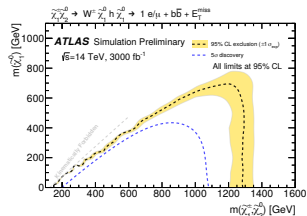
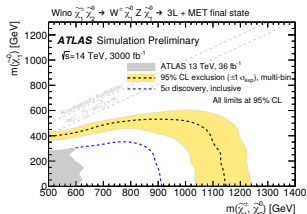
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

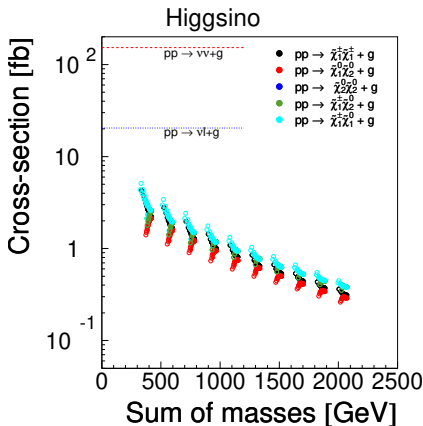
- Vary relative signs of μ , M_1 , and M_2
- For $\mu > M_2$
- or $\mu < M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



SUSY cross-sections at FCChh

Variation of cross-section for $pp \rightarrow$ uncoloured bosinos + gluon
(CTEQ6L1 pdfs)

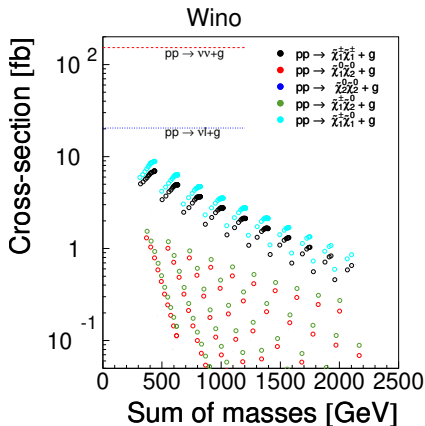
- Higgsino LSP
- Wino LSP
- or Bino LSP
- Note: Can vary by \sim factor 2
- Note: Exponential fall with mass
- \Rightarrow Will extend far beyond current at high $\Delta(M)$, but will stay below the $M_{NLSP} = 2 \times M_{LSP}$ line (see backup...)



SUSY cross-sections at FCChh

Variation of cross-section for $pp \rightarrow$ uncoloured bosinos + gluon
(CTEQ6L1 pdfs)

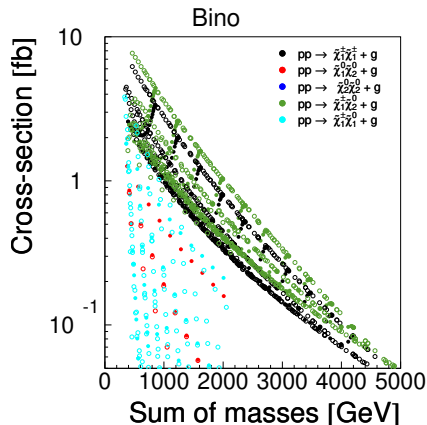
- Higgsino LSP
- Wino LSP
- or Bino LSP
- Note: Can vary by \sim factor 2
- Note: Exponential fall with mass
- \Rightarrow Will extend far beyond current at high $\Delta(M)$, but will stay below the $M_{NLSP} = 2 \times M_{LSP}$ line (see backup...)



SUSY cross-sections at FCChh

Variation of cross-section for $pp \rightarrow$ uncoloured bosinos + gluon
(CTEQ6L1 pdfs)

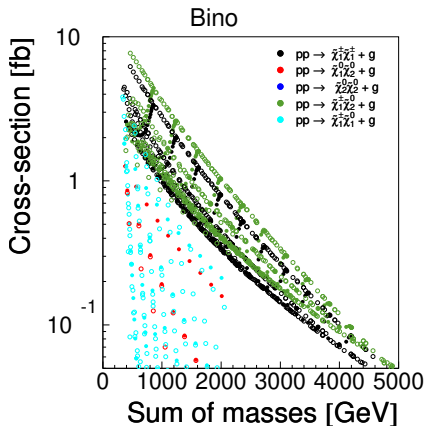
- Higgsino LSP
- Wino LSP
- or Bino LSP
- Note: Can vary by \sim factor 2
- Note: Exponential fall with mass
- \Rightarrow Will extend far beyond current at high $\Delta(M)$, but will stay below the $M_{NLSP} = 2 \times M_{LSP}$ line (see backup...)



SUSY cross-sections at FCChh

Variation of cross-section for $pp \rightarrow$ uncoloured bosinos + gluon
(CTEQ6L1 pdfs)

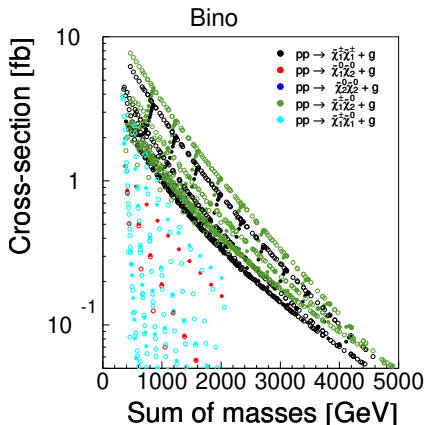
- Higgsino LSP
- Wino LSP
- or Bino LSP
- Note: Can vary by \sim factor 2
- Note: Exponential fall with mass
- \Rightarrow Will extend far beyond current at high $\Delta(M)$, but will stay below the $M_{NLSP} = 2 \times M_{LSP}$ line (see backup...)



SUSY cross-sections at FCChh

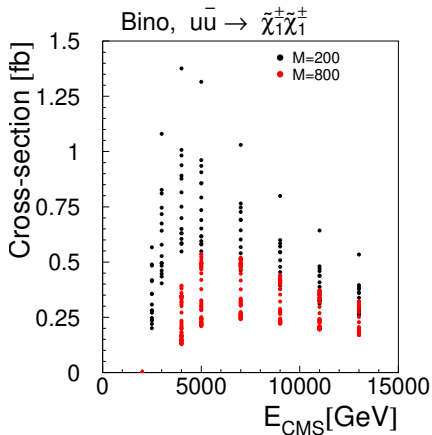
Variation of cross-section for $pp \rightarrow$ uncoloured bosinos + gluon
(CTEQ6L1 pdfs)

- Higgsino LSP
- Wino LSP
- or Bino LSP
- Note: Can vary by \sim factor 2
- Note: Exponential fall with mass
- \Rightarrow Will extend far beyond current **at high $\Delta(M)$** , but will stay **below the $M_{NLSP} = 2 \times M_{LSP}$** line (see backup...)



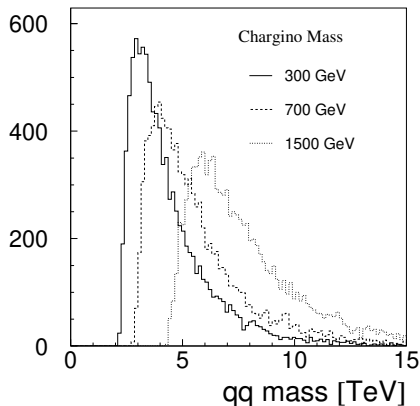
SUSY cross-sections at FCChh: Why exponential fall-off

- Consider *fixed* m_{qq} , at two masses: First rise w/ β , then fall-off w/ $1/s$.
- Fold this with rapidly falling pdf:s (in particular for the sea)
- $\Rightarrow m_{qq}$ (linear) function of bino-mass



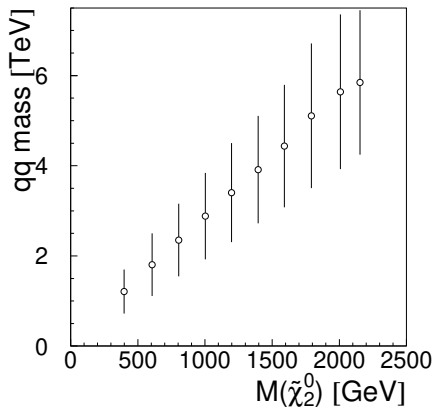
SUSY cross-sections at FCChh: Why exponential fall-off

- Consider *fixed* m_{qq} , at two masses: First rise w/ β , then fall-off w/ $1/s$.
- Fold this with rapidly falling pdf:s (in particular for the sea)
- $\Rightarrow m_{qq}$ (linear) function of bino-mass



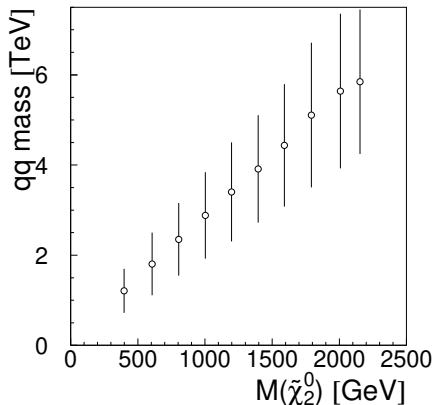
SUSY cross-sections at FCChh: Why exponential fall-off

- Consider *fixed* m_{qq} , at two masses: First rise w/ β , then fall-off w/ $1/s$.
- Fold this with rapidly falling pdf:s (in particular for the sea)
- $\Rightarrow m_{qq}$ (linear) function of bino-mass



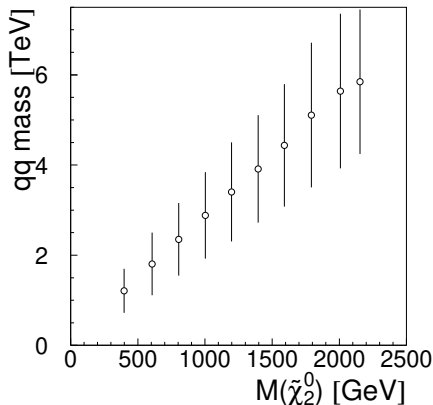
SUSY cross-sections at FCChh: Why exponential fall-off

- m_{qq} (linear) function of bosino-mass
- At these mass-ratios, missing p_T is proportional to m_{qq}
- \Rightarrow missing p_T increases linearly with bosino-mass.
- \Rightarrow can increase missing p_T -cut linearly when looking for higher masses, with the same efficiency
- Then the background decreases as much.
- S/B remains constant along lines in $M_{\tilde{\chi}_1^\pm}$ vs. M_{LSP}



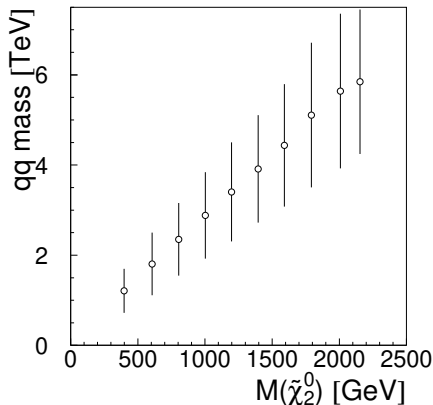
SUSY cross-sections at FCChh: Why exponential fall-off

- m_{qq} (linear) function of bosino-mass
- At these mass-ratios, missing p_T is proportional to m_{qq}
- \Rightarrow missing p_T increases linearly with bosino-mass.
- \Rightarrow can increase missing p_T -cut linearly when looking for higher masses, with the same efficiency
- Then the background decreases as much.
- S/B remains constant along lines in $M_{\tilde{\chi}_1^\pm}$ vs. M_{LSP}



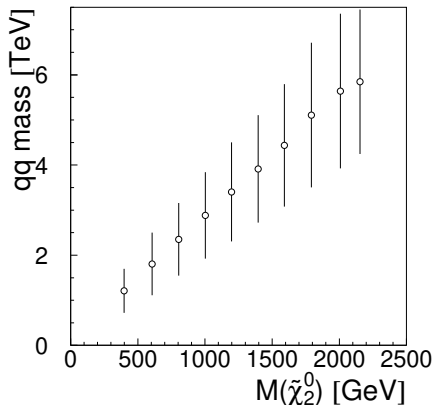
SUSY cross-sections at FCChh: Why exponential fall-off

- m_{qq} (linear) function of bosino-mass
- At these mass-ratios, missing p_T is proportional to m_{qq}
- \Rightarrow missing p_T increases linearly with bosino-mass.
- \Rightarrow can increase missing p_T -cut linearly when looking for higher masses, with the same efficiency
- Then the background decreases as much.
- S/B remains constant along lines in $M_{\tilde{\chi}_1^\pm}$ vs. M_{LSP}



SUSY cross-sections at FCChh: Why exponential fall-off

- m_{qq} (linear) function of bosino-mass
- At these mass-ratios, missing p_T is proportional to m_{qq}
- \Rightarrow missing p_T increases linearly with bosino-mass.
- \Rightarrow can increase missing p_T -cut linearly when looking for higher masses, with the same efficiency
- Then the background decreases as much.
- S/B remains constant along lines in $M_{\tilde{\chi}_1^\pm}$ vs. M_{LSP}



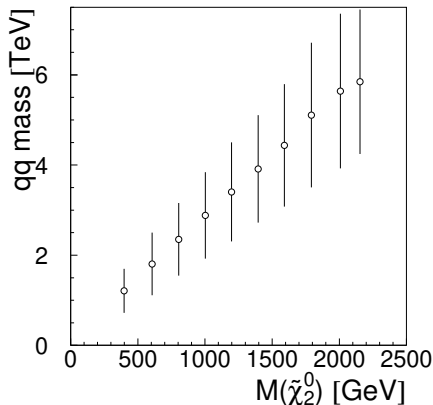
SUSY cross-sections at FCChh: Why exponential fall-off

- m_{qq} (linear) function of bosino-mass
- At these mass-ratios, missing p_T is proportional to m_{qq}
- \Rightarrow missing p_T increases linearly with bosino-mass.

Uptake

Expect that the limit sticks to the **same diagonal** as energy is increased.

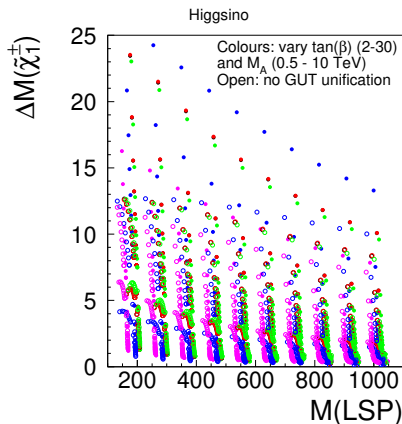
- Then the background decreases as much.
- S/B remains constant along lines in $M_{\tilde{\chi}_1^\pm}$ vs. M_{LSP}



Aspects of the spectrum : $\Delta(M)$

Yet another angle: $\Delta(M)$ for $\tilde{\chi}_1^\pm$ vs. M_{LSP}

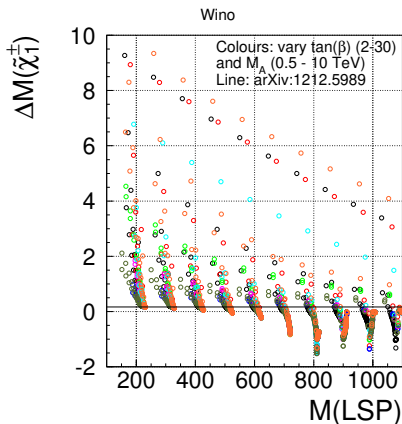
- For Higgsino LSP
- For Wino LSP
- Note large spread possible!



Aspects of the spectrum : $\Delta(M)$

Yet another angle: $\Delta(M)$ for $\tilde{\chi}_1^\pm$ vs. M_{LSP}

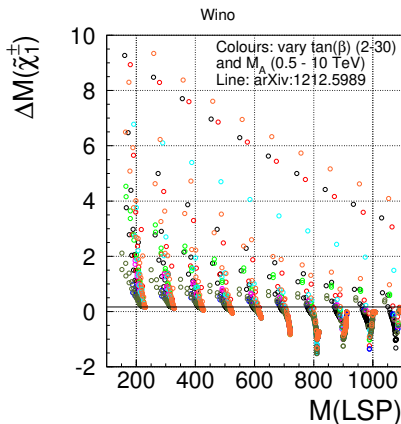
- For Higgsino LSP
- For Wino LSP
- Note large spread possible!



Aspects of the spectrum : $\Delta(M)$

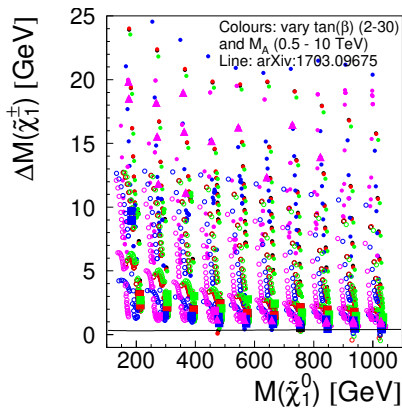
Yet another angle: $\Delta(M)$ for $\tilde{\chi}_1^\pm$ vs. M_{LSP}

- For Higgsino LSP
- For Wino LSP
- Note large spread possible!



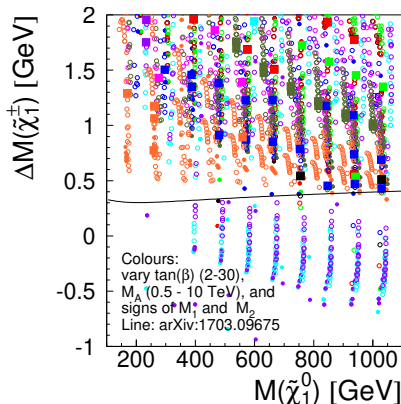
Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP.
- Zoom in. The line is the absolute limit mentioned in the BB.
- Reason:
arXiv:1703.09675
considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.



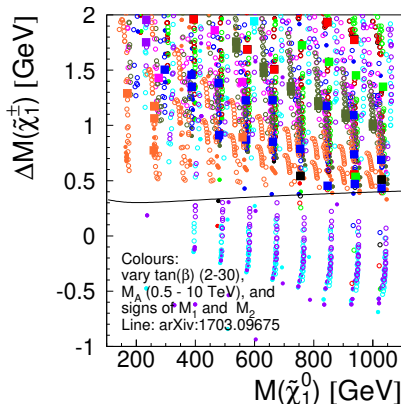
Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP.
- Zoom in. The line is the absolute limit mentioned in the BB.
- Reason:
arXiv:1703.09675
considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.



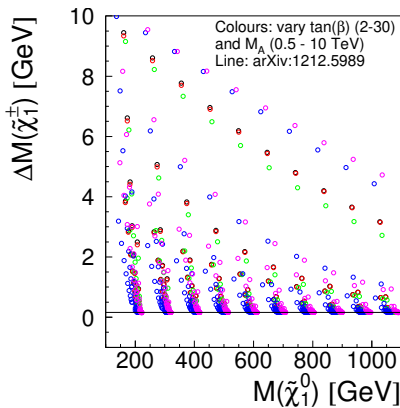
Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP.
- Zoom in. The line is the absolute limit mentioned in the BB.
- Reason:
arXiv:1703.09675
considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.

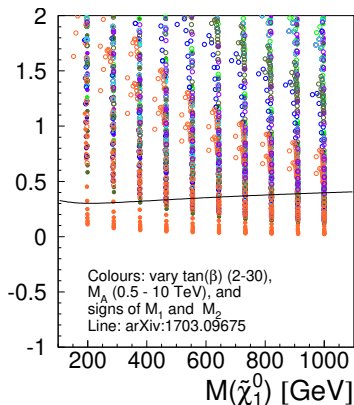
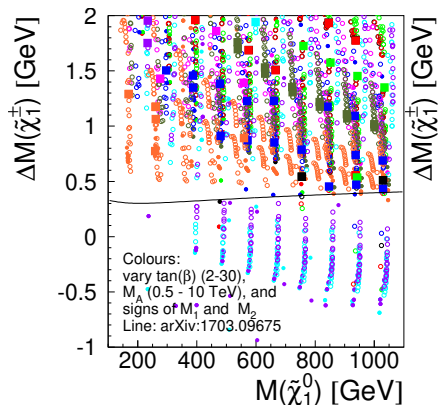


Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP.
- Zoom in. The line is the absolute limit mentioned in the BB.
- Reason:
arXiv:1703.09675
considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.

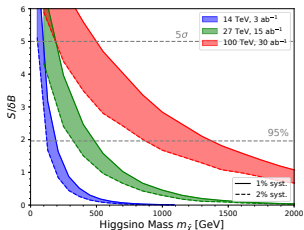
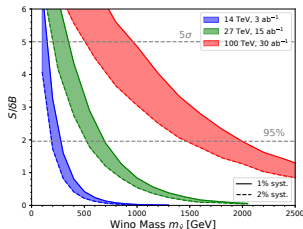


second opinion: feynhiggs



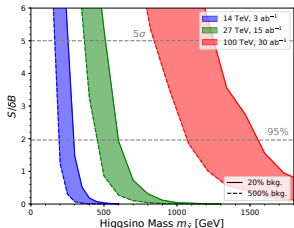
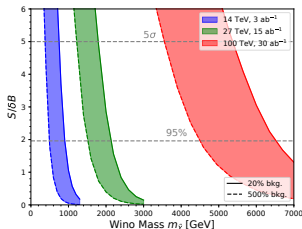
SUSY In The Briefing-book: Wino/Higgsino LSP - Very low $\Delta(M)$ Sources

- Two methods: “Disappearing tracks” and “Mono-X”
 - “Disappearing tracks”
 - “Mono-X”
- [arxiv:1805.00015](https://arxiv.org/abs/1805.00015), Based on DELPHES with ATLAS-card (\Rightarrow LHC PU...)
- Both from the HE/HL-LHC input to ESU (*not* FCChh)
- Systematics-limited. Both ATLAS and CMS state $\sim 10\%$ in existing “Mono-X” searches (PU 1/20 of FCChh)



SUSY In The Briefing-book: Wino/Higgsino LSP - Very low $\Delta(M)$ Sources

- Two methods: “Disappearing tracks” and “Mono-X”
 - “Disappearing tracks”
 - “Mono-X”
- [arxiv:1805.00015](#), Based on DELPHES with ATLAS-card (\Rightarrow LHC PU...)
- Both from the HE/HL-LHC input to ESU (*not* FCChh)
- Systematics-limited. Both ATLAS and CMS state $\sim 10\%$ in existing “Mono-X” searches (PU 1/20 of FCChh)



SUSY In The Briefing-book: Wino/Higgsino LSP - Very low $\Delta(M)$ Sources

- Two methods: “Disappearing tracks” and “Mono-X”
 - “Disappearing tracks”
 - “Mono-X”
- arxiv:1805.00015, Based on DELPHES with ATLAS-card (\Rightarrow LHC PU...)
- Both from the HE/HL-LHC input to ESU (*not* FCChh)
- Systematics-limited. Both ATLAS and CMS state $\sim 10\%$ in existing “Mono-X” searches (PU 1/20 of FCChh)

