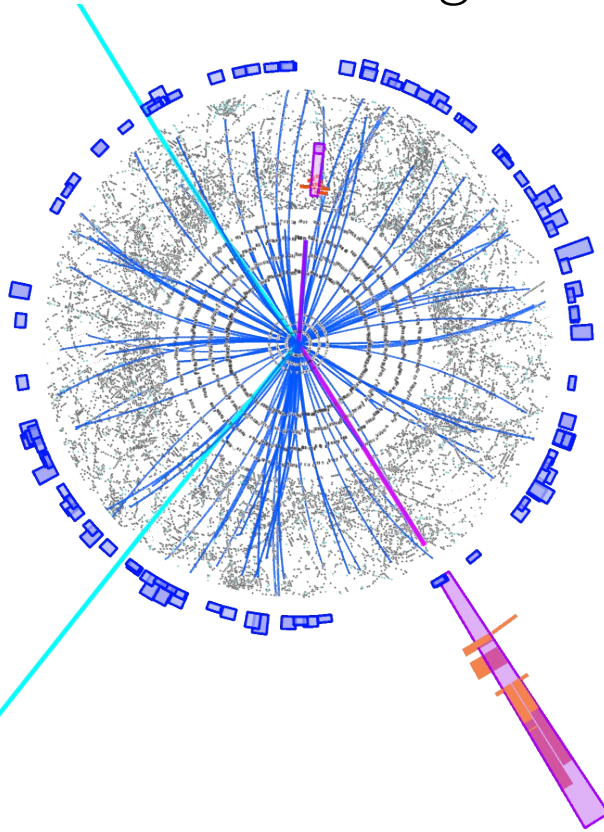


Detector Physics I

Particle Detection in High Energy Physics



Simon Spannagel, DESY
Introduction to the Terascale
06 March 2023

Event-Display:

ATLAS
EXPERIMENT
<http://atlas.ch>

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Prologue: Units

- Elementary charge: $e = 1.602\,176\,6208 \times 10^{-19} \text{ C}$
- Energy unit eV: $1 \text{ eV} = 1.602\,176\,6208 \times 10^{-19} \text{ J (CV)}$
Energy of an electron, accelerated through 1V
- Electron mass: $m_e = 9.1 \times 10^{-31} \text{ kg}$
 $= 511 \text{ keV} / c^2$
- Factor c^2 often dropped, units defined with $c = 1$ (and $\hbar = 1$):
MeV, GeV, TeV...
for energy, momentum, (rest-) mass
- Scientific notation: Mantissa plus exponent to Base 10: $a \times 10^b$

What do we measure?

Determination of particle properties

Which particles can we measure?

- Particles have to be long-lived enough to reach the detectors...

- Many elementary particles have very short life time (Higgs, W, Z...)

- Measuring their decay products

$$d = c \tau \gamma, \quad \gamma = \frac{1}{\sqrt{(1-\beta^2)}}, \quad \beta = \frac{v}{c}$$

- Particles have to interact with detector material!

- Not every particle participates in every force

- Possibilities of detection depends strongly on forces

- Elementary particles: $e^\pm, \mu^\pm, \nu^e, \bar{\nu}^e, \nu^\mu, \bar{\nu}^\mu, \nu^\tau, \bar{\nu}^\tau, \gamma$

- Baryons: $p^\pm, n, \Sigma^\pm, \Xi_0^\pm, \Xi^\pm, \Omega^\pm$

- Mesons: $\pi^\pm, K^\pm, K_0 (K_0^S, K_0^L)$

$$\tau_n \approx 15 \text{ min}$$

$$\tau_\mu \approx 2 \cdot 10^{-6} \text{ s}$$

Observables

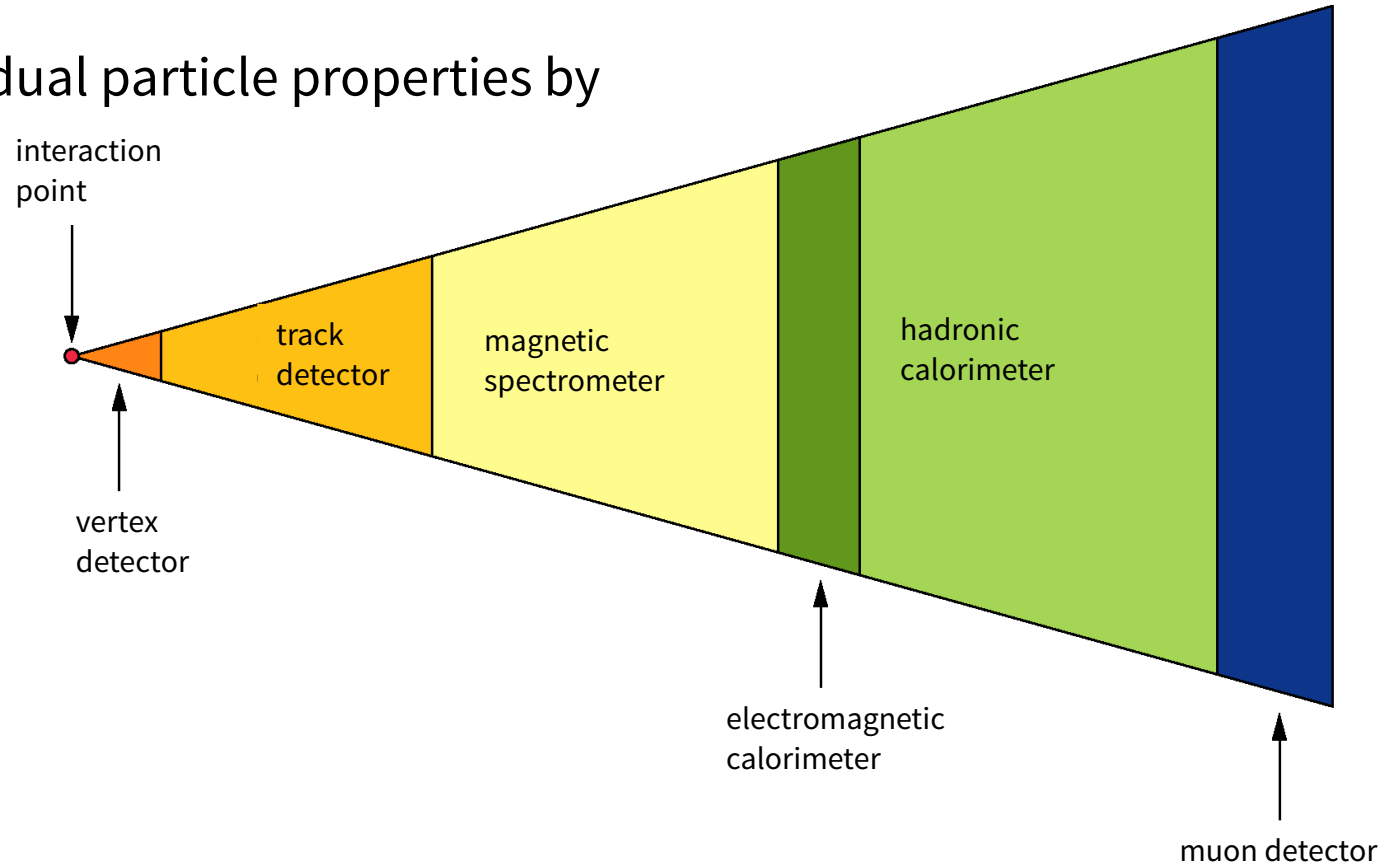
- Momentum **p** Bending radius of track in magnetic field
- Velocity **v** Time of flight, RICH, etc.
- Charge **Q** Bending radius in magnetic field
- Lifetime **τ** Measurement of decay length
- Energy **E** Absorption in calorimeters
- Rest mass **m** Indirect measurement e.g. from momentum and energy or velocity

$$E^2 = m^2 c^4 + p^2 c^2 \quad p = \gamma m v = \frac{m v}{\sqrt{(1 - v^2/c^2)}}$$



Typical Design of an Experiment

- Combination of complementary detection methods
- Measurement of individual particle properties by separate detectors
- Order is important!
Some measurements are “destructive“
- Design of many experiments is very similar



Detector Terminology

- **Dead time:**

Time period immediately after the detection of a particle during which the detector is not yet ready again to detect another particle.

- *Non-paralyzable detector: newly occurring event does nothing*
- *Paralyzable detector: newly occurring event extends dead time*

- **Resolution:**

Achievable uncertainty on the observable

- **Efficiency:**

Number of recorded/detected events divided by number of events that occurred

Interaction of Radiation with Matter

Energy loss and interaction processes

Interaction with Matter

- High energy particles interact via different processes with matter, depending on
 - Particle type
 - Energy
 - Material
- Energy loss of the particles via interaction
 - Energy transfer to material or other (free) particles
 - In Detectors: ***Energy loss = signal!***

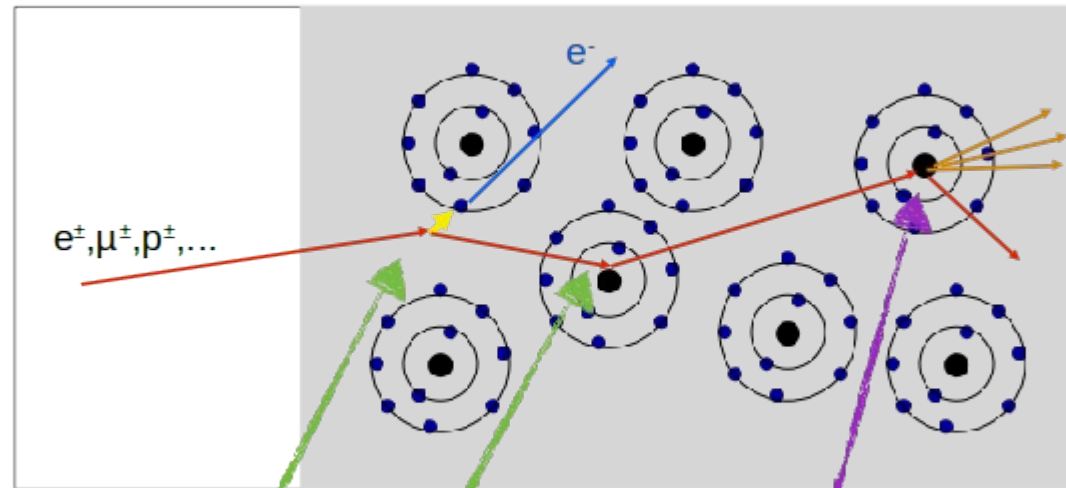
Charged Particles

Ionization

Elastic scattering

Recoil from atom
lattice \rightarrow photons

Inelastic scattering



Electromagnetic interaction

Electromagnetic,
weak or
strong Interaction

Mean Energy Loss

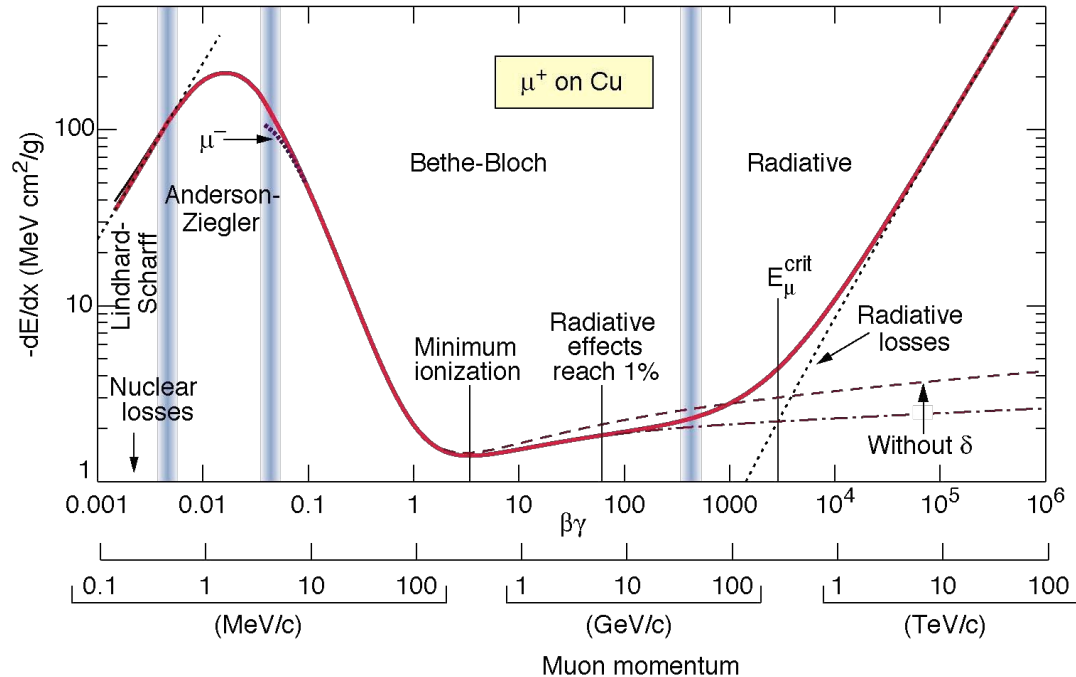
- Charged particles interact with electrons in matter

- For heavy charged particles
Bethe-Formula

$$-\left\langle \frac{dE}{dx} \right\rangle \approx K q^2 \frac{1}{\beta^2} \frac{Z}{A} \left[\ln \left(\frac{2 m_e c^2}{I} \beta^2 \gamma^2 \right) - \beta^2 \right]$$

- Energy loss depends on

- Properties of projectile: charge, energy
- Target properties: atomic number, ionization energy, density

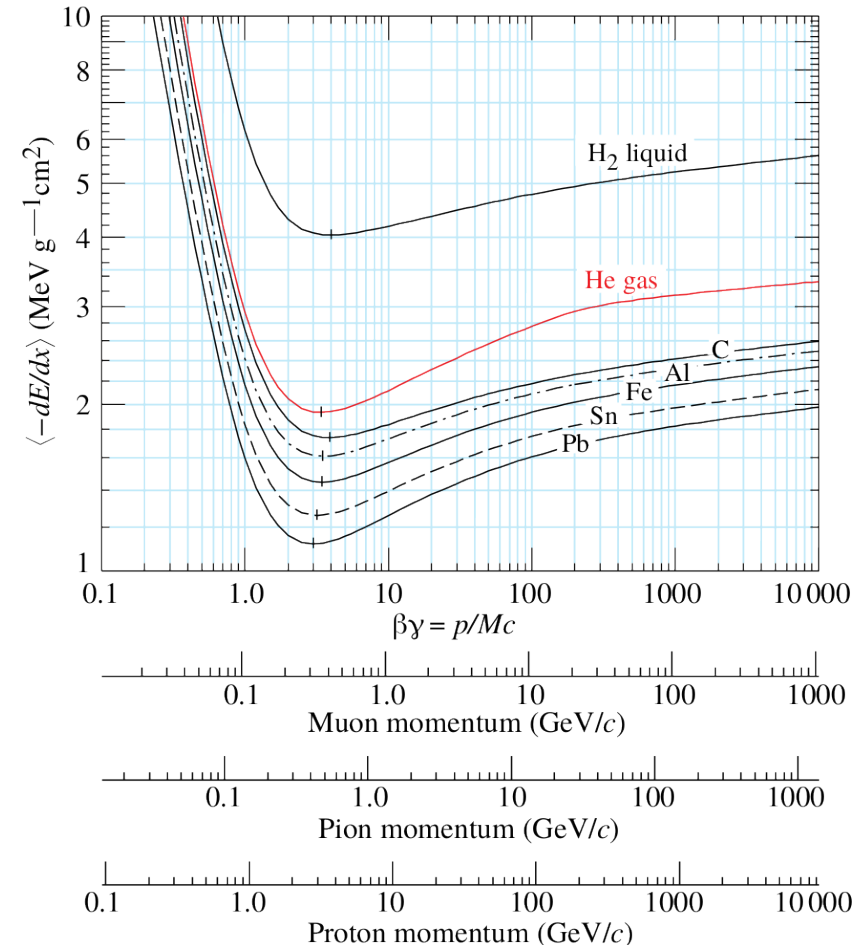


Mean Energy Loss

- Different components are dominating:
 - At low energies: $\sim 1/\beta^2$
 - At high energies: $\sim \ln \gamma$

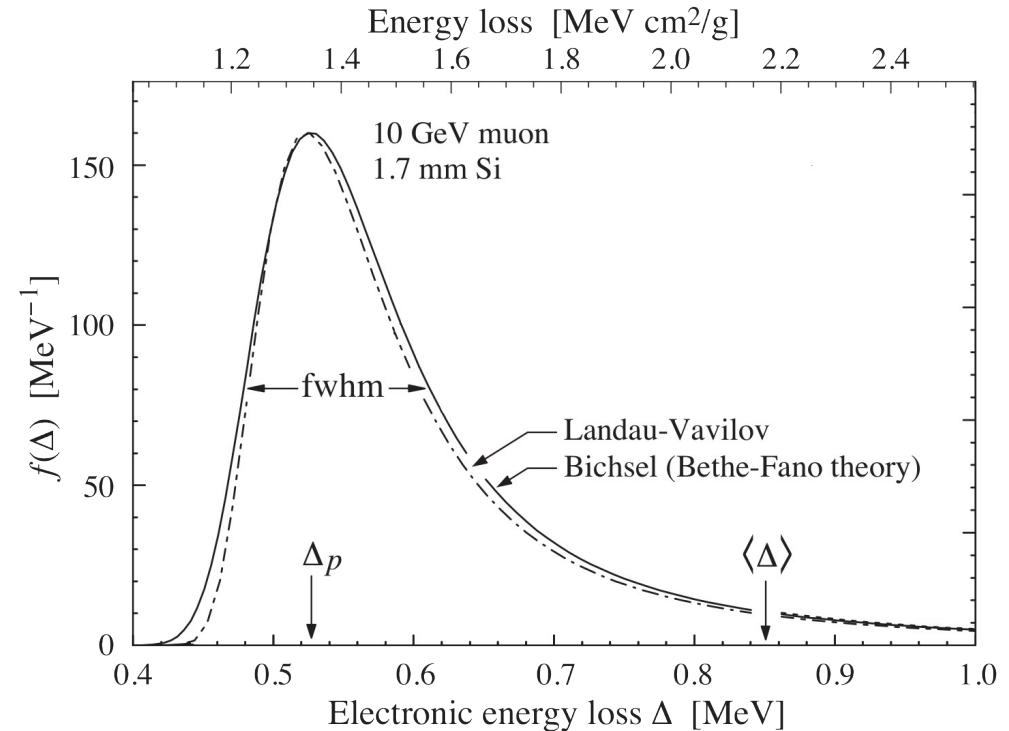
$$-\left\langle \frac{dE}{dx} \right\rangle \approx K q^2 \frac{1}{\beta^2} \frac{Z}{A} \left[\ln \left(\frac{2 m_e c^2}{I} \beta^2 \gamma^2 \right) - \beta^2 \right]$$

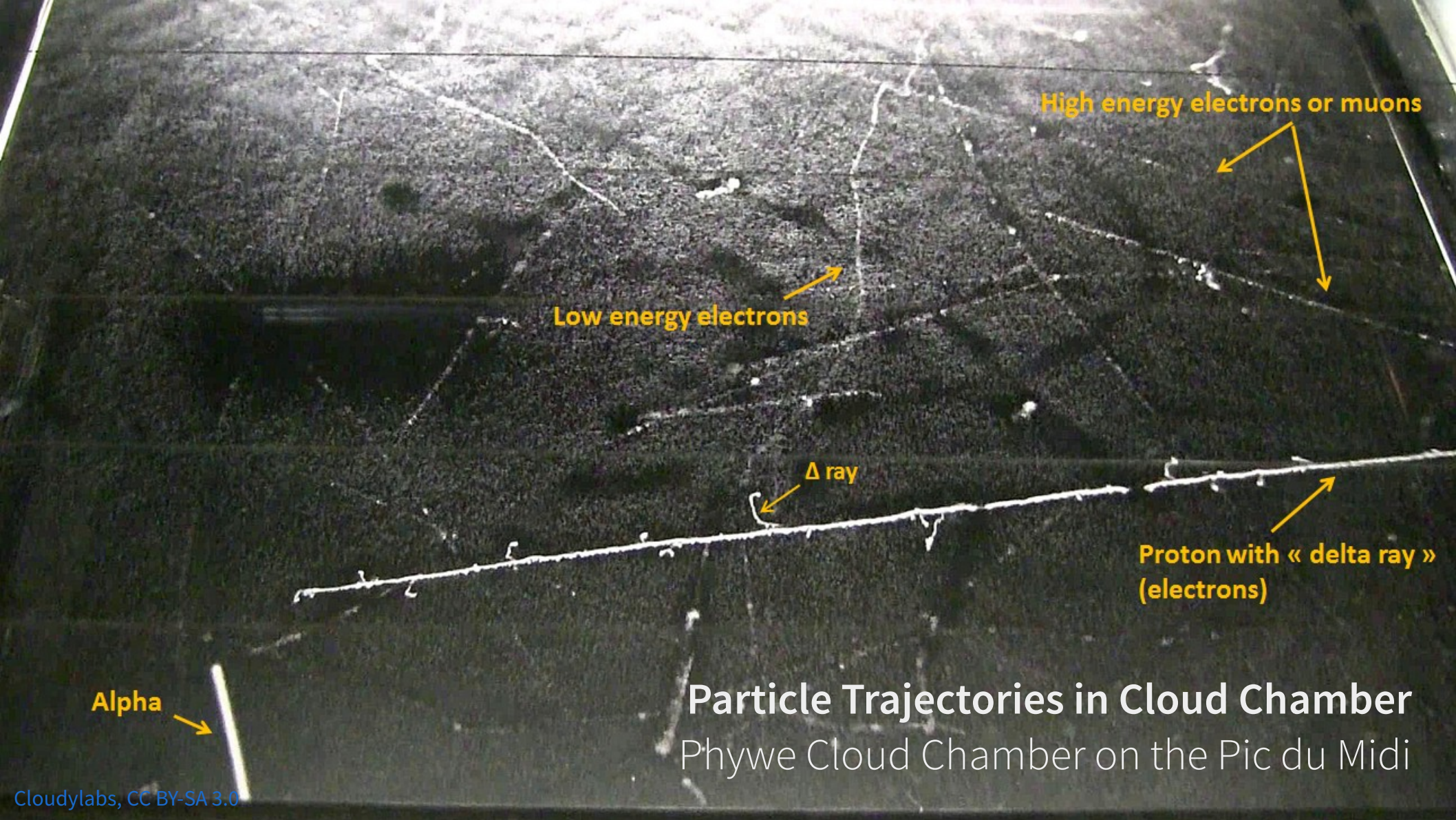
- Minimal energy loss for $\sim \beta\gamma = 3$
MIP: Minimum Ionizing Particle
- $[dx] = \text{g} / \text{cm}^2 = \text{cm} \times \text{g} / \text{cm}^3$



Fluctuations in Energy Loss

- Actual energy loss fluctuates around mean value
- **Landau-Vavilov distribution** with long tails to high energies
 - Most probable value \ll mean value
- Orbital electrons can receive very large energy transfers
 - Creation of delta electrons
 - Delta electrons have enough energy for further ionization





High energy electrons or muons

Low energy electrons

Δ ray

Proton with « delta ray »
(electrons)

Alpha

Particle Trajectories in Cloud Chamber

Phywe Cloud Chamber on the Pic du Midi

Exception: Electron & Positron

- Special case: low mass

$$m_e = 0.511 \text{ MeV} / c^2$$

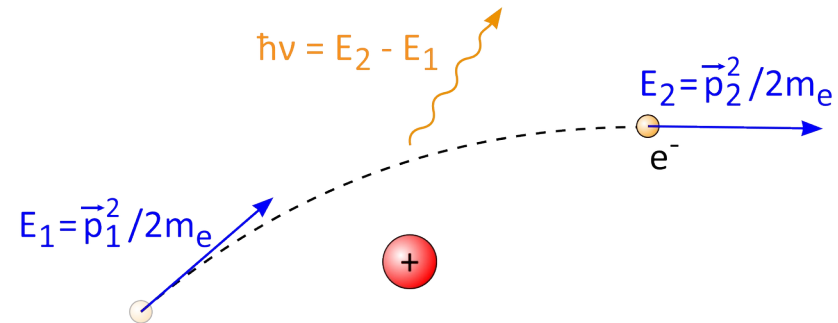
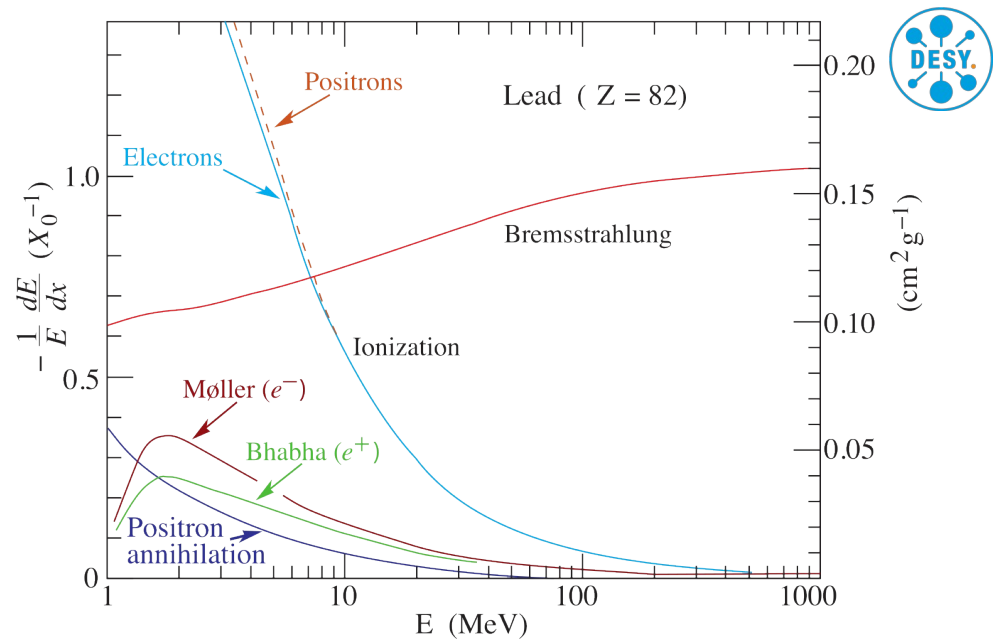
$$m_\mu = 106 \text{ MeV} / c^2 \approx 200 m_e$$

- At high energies: Bremsstrahlung

- occurs when the momentum of a charged particle changes, e.g.
 - ...in the Coulomb field of a nucleus
 - ...in the magnetic field: synchrotron radiation

- Here: particles slowed down in matter

- Relevant for electrons: $-\frac{dE}{dx} \sim E \cdot \frac{1}{m^2}$



Cherenkov Radiation

- Is emitted, when particle velocity > speed of light

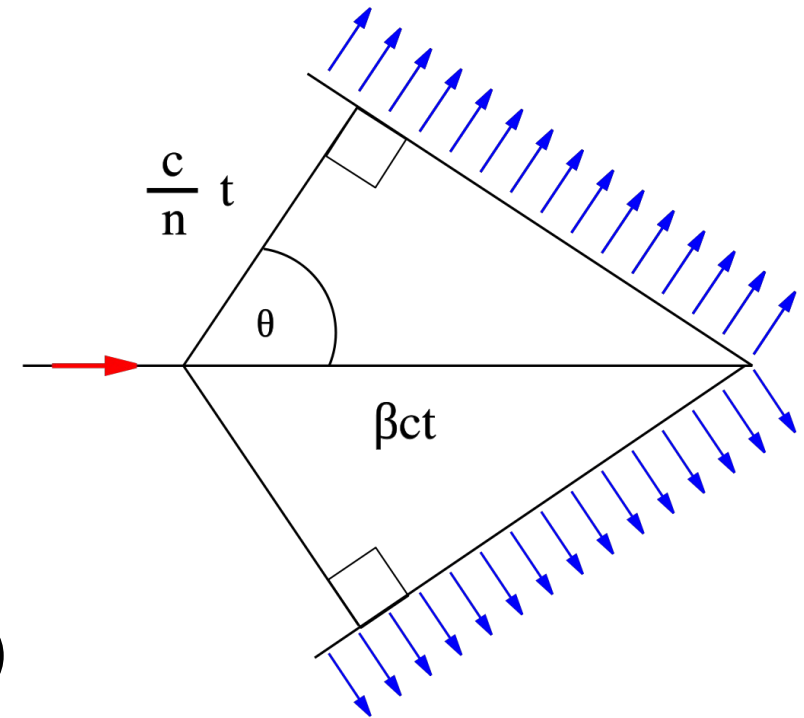
$$c_{\text{Medium}} = \frac{c_{\text{Vacuum}}}{n} < v_{\text{Particle}} < c_{\text{Vacuum}}$$

n : refractive index of medium

- Electromagnetic shockwave with conical shape is emitted under angle θ :

$$\cos(\theta) = \frac{1}{\beta n}, \quad \beta = \frac{v}{c}$$

- Very low energy loss (ca. 1% of total energy loss)



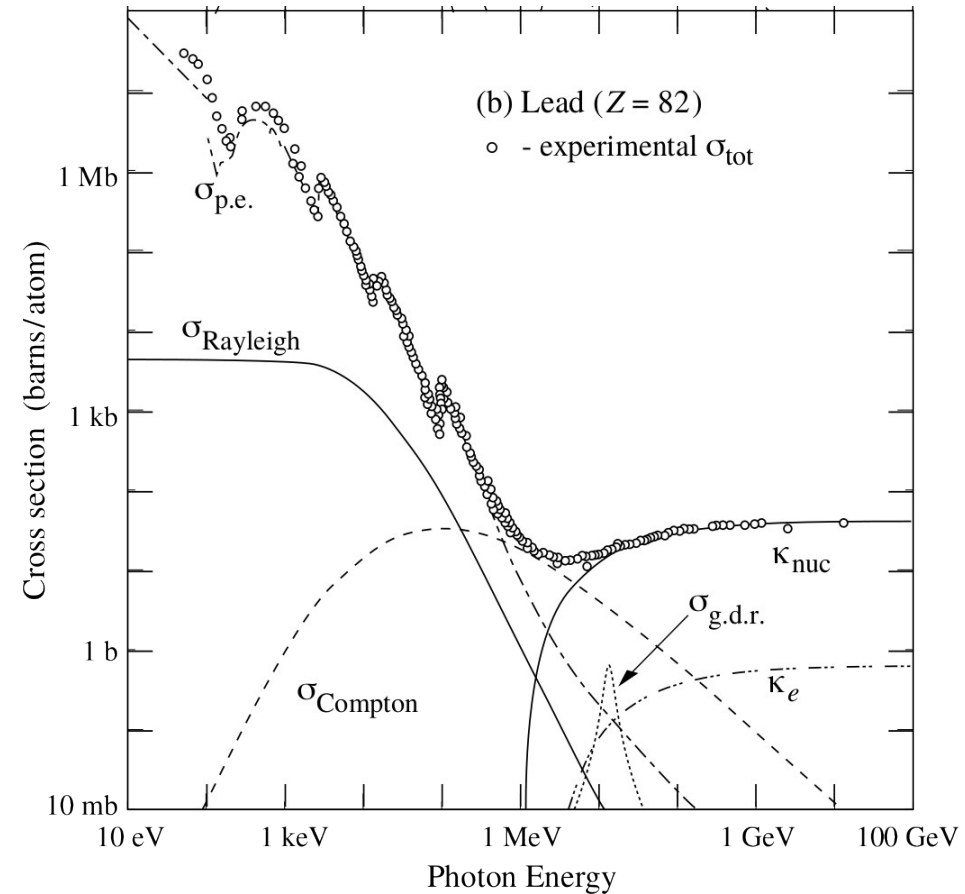


Cherenkov Radiation of a Nuclear Reactor

Advanced Test Reactor, INL

Photons

- Electromagnetic interaction
- Different processes dominate, depending on the photon energy:
 - Photo(electric) effect
 - Rayleigh scattering
 - Compton effect
 - Pair production



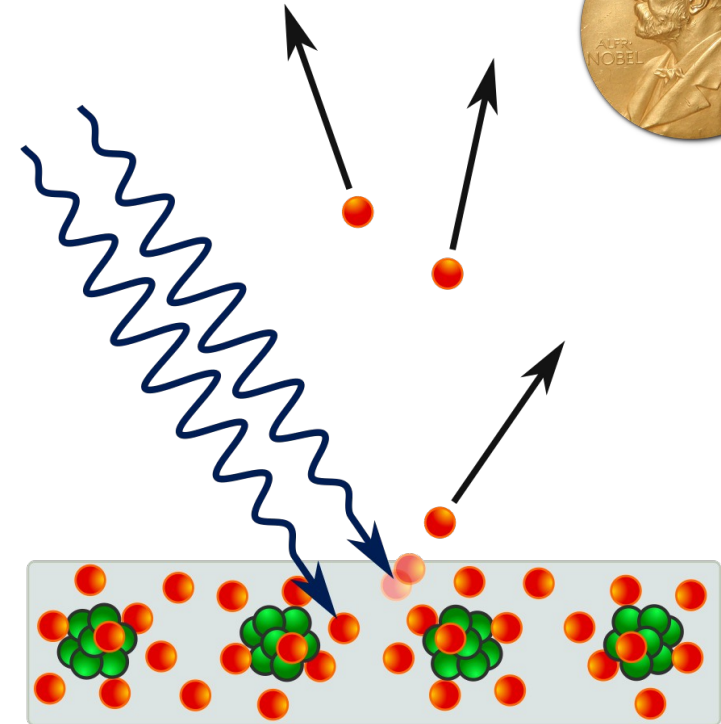
Photoelectric Effect

- Theoretical description of photo effect:
- Photon is absorbed by electron in atomic shell
- Transferred energy releases/frees electron

$$\gamma + Atom = e^{-} + Ion$$

- Process only possible in the field of the nucleus:
 - Momentum conservation
 - Nucleus absorbs recoil
- Cross section of the photoelectric effect shows shell structure of the atom

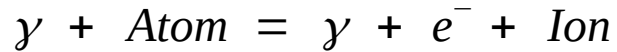
Nobel price 1921 for Albert Einstein



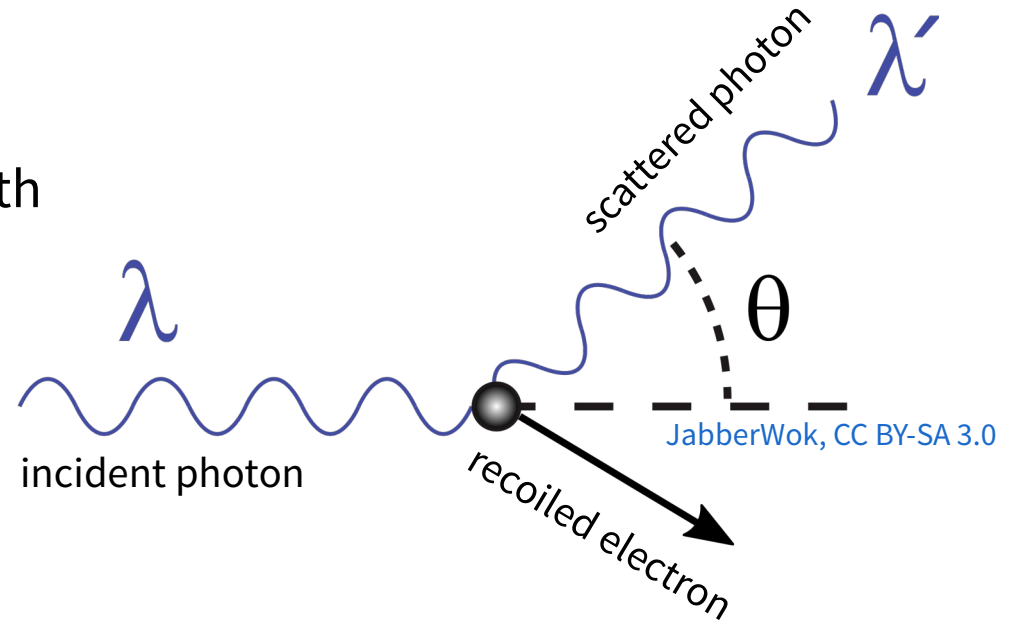
Ponor, CC BY-SA 4.0

Compton Effect

- Describes scattering of photon on “quasi-free” electron

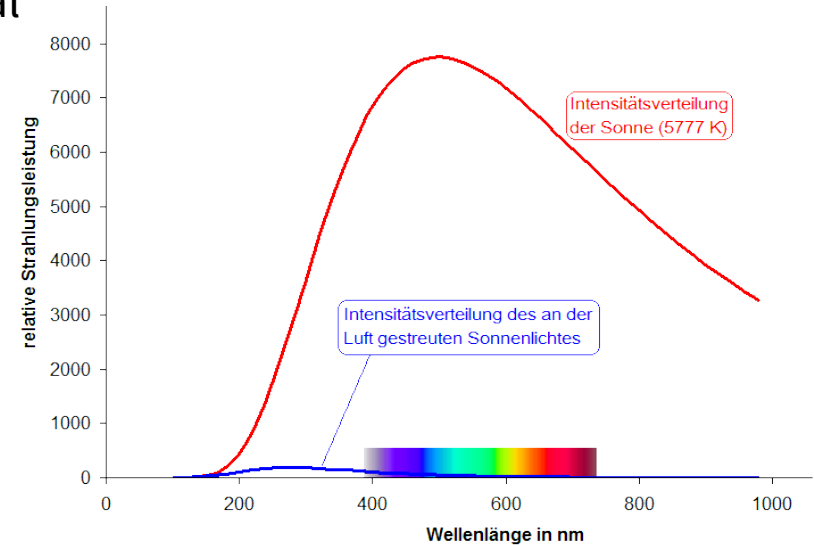


- Photon is deflected from its original path
- Wavelength of photon changes through energy transfer to electron



Photons: Thomson/Rayleigh Scattering

- Elastic scattering: almost no energy transfer to material
- Thomson scattering: Photon scattering on free electron
 - Low-energy limit of Compton scattering
- Rayleigh scattering: Photon scattering off an entire atom
 - Scattering cross section $\sigma_{\text{Rayleigh}} \sim f^4$
- Reason for blue / red coloration of the sky depending on the zenith angle
 - Noon: short path through atmosphere, hardly any blue light scattered
 - Morning/evening: long path through atmosphere, much blue light scattered



$$\frac{\sigma_{\text{blue}}}{\sigma_{\text{red}}} = \frac{1/\lambda_{\text{blue}}^4}{1/\lambda_{\text{red}}^4} = \left(\frac{650 \text{ nm}}{450 \text{ nm}} \right)^4 \approx 4.4$$

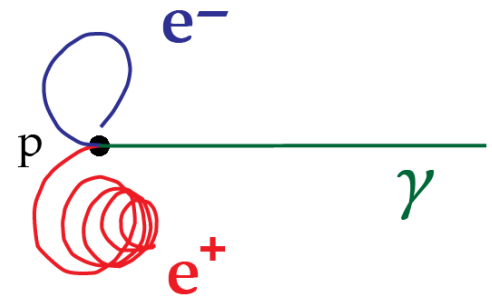
Photons: Pair Production

- Pair production is the generation of an electron-positron pair by the photon
- Pair production occurs in the field of a partner, which absorbs the recoil (atomic nucleus, but also shell electron)

$$\gamma + p = e^+ + e^- + p$$

- Photon must provide at least rest mass of e⁺e⁻ pair plus recoil energy:

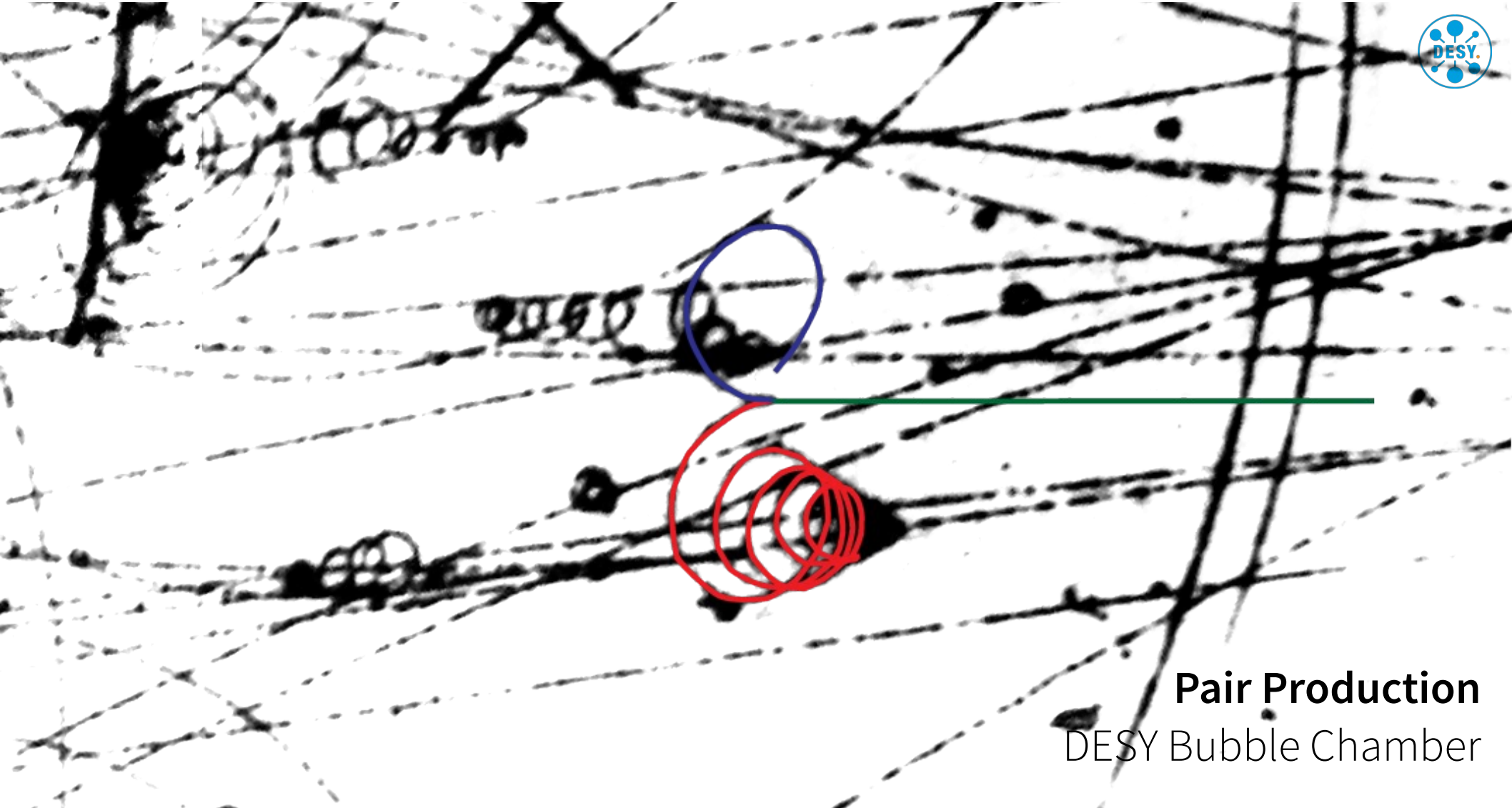
$$E_\gamma \geq 2m_e c^2 \left(1 + \frac{m_e}{M}\right)$$



$$\gamma p \rightarrow e^- e^+ p$$

Ivan Baev, CC BY-SA 3.0

- Recoil energy can often be neglected, e.g. Ge detector: $\frac{m_e}{M} \approx 7.6 \cdot 10^{-6}$



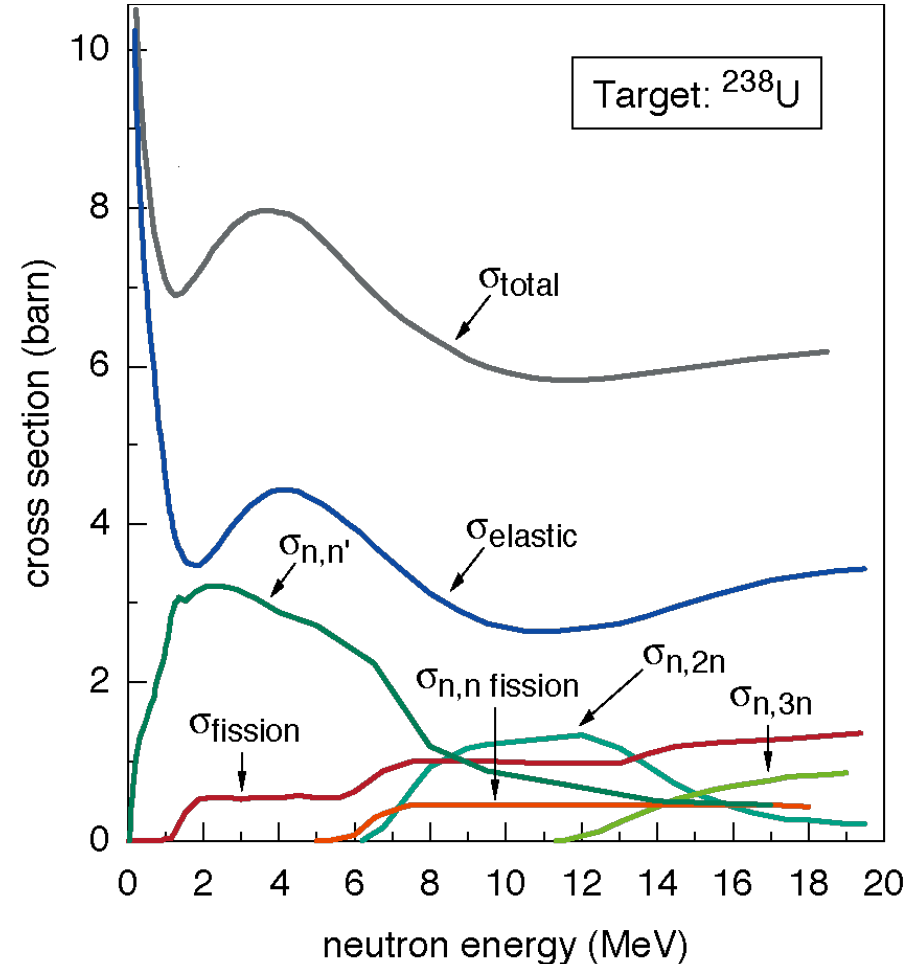
Pair Production

DESY Bubble Chamber

Hadronic Interaction

- Interactions of hadrons with nuclei
- Based on the **strong interaction, QCD**
 - Low range
 - Low probability of hadronic reactions
 - Neutrons can only interact strongly: very penetrating
- Many possible processes (energy dependent)

Elastic, inelastic scattering; neutron capture; reactions with radiation of charged particles; nuclear fission.



Interaction of Neutrinos

- Neutrinos are exclusively subject to the weak interaction
- Possible Interactions

$$\bar{\nu}_l + p \rightarrow n + l^+$$

$$\nu_l + n \rightarrow p + l^-$$

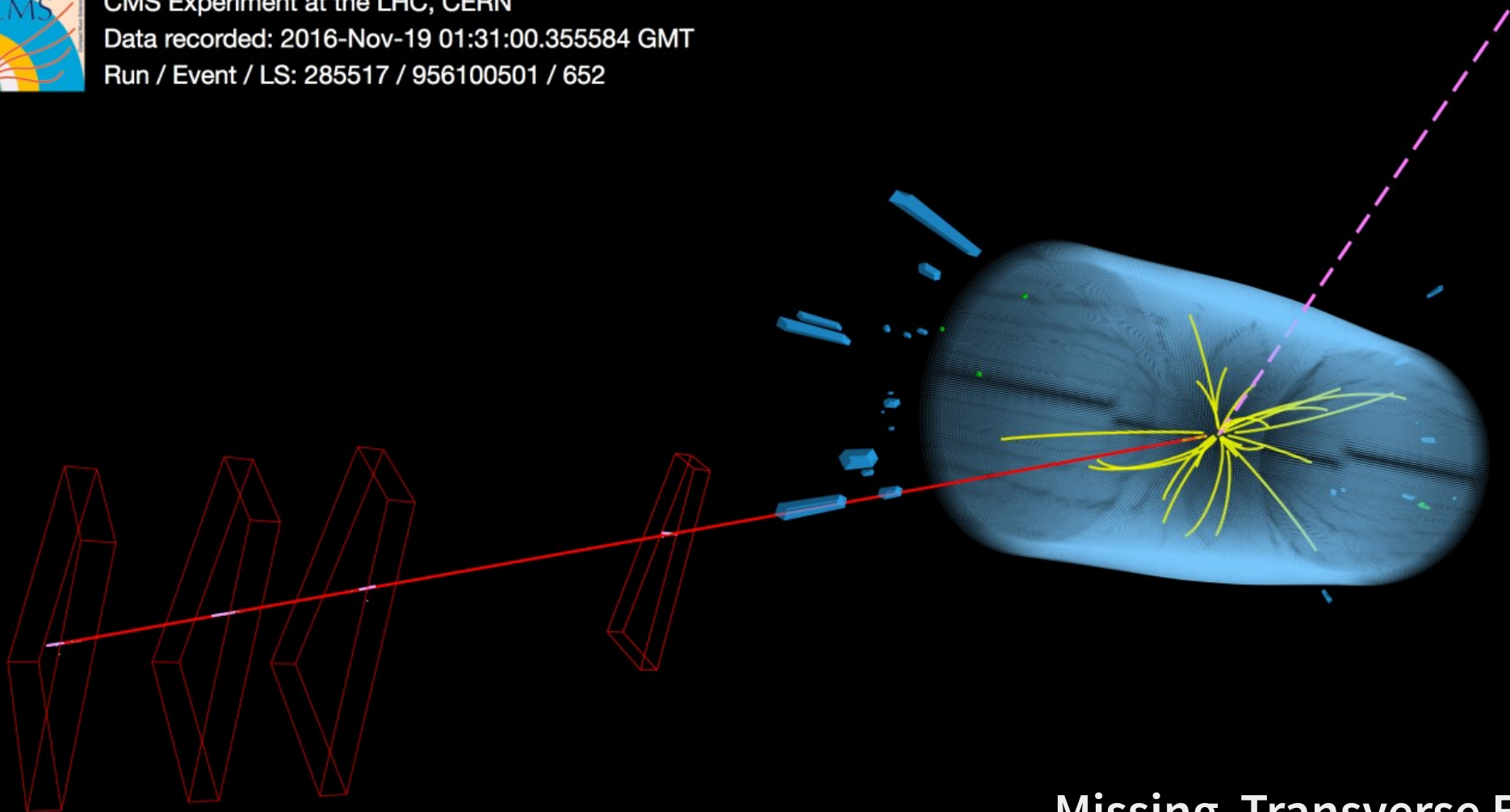
- Neutrino interactions have very (very!) low cross sections
- Detection of neutrinos requires
 - Very large detector and high neutrino fluxes (direct detection) or
 - Hermetic detector for measurement of missing energy (indirect evidence)



CMS Experiment at the LHC, CERN

Data recorded: 2016-Nov-19 01:31:00.355584 GMT

Run / Event / LS: 285517 / 956100501 / 652



Missing Transverse Energy
Detection of Neutrinos

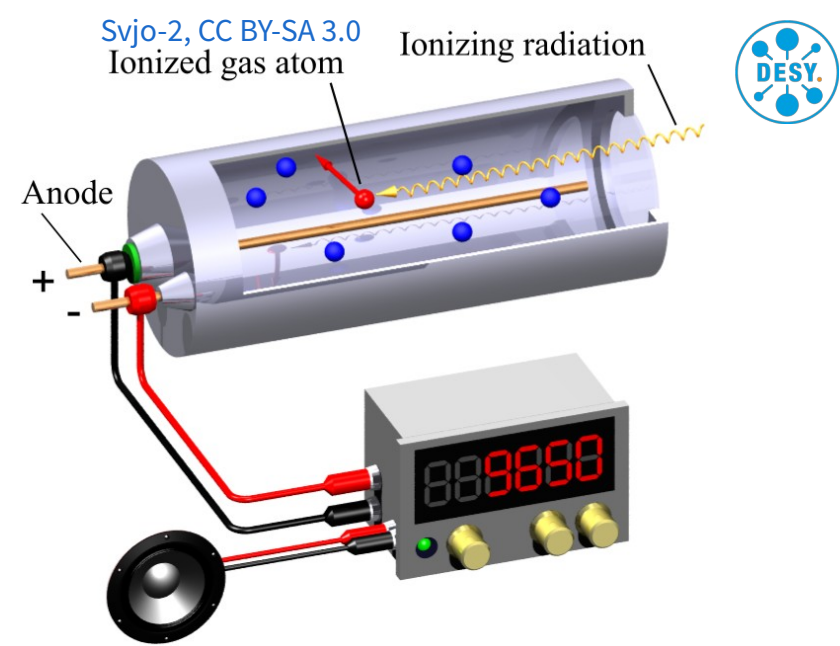
Particle Detectors

Historic Overview

Geiger-Müller Counter

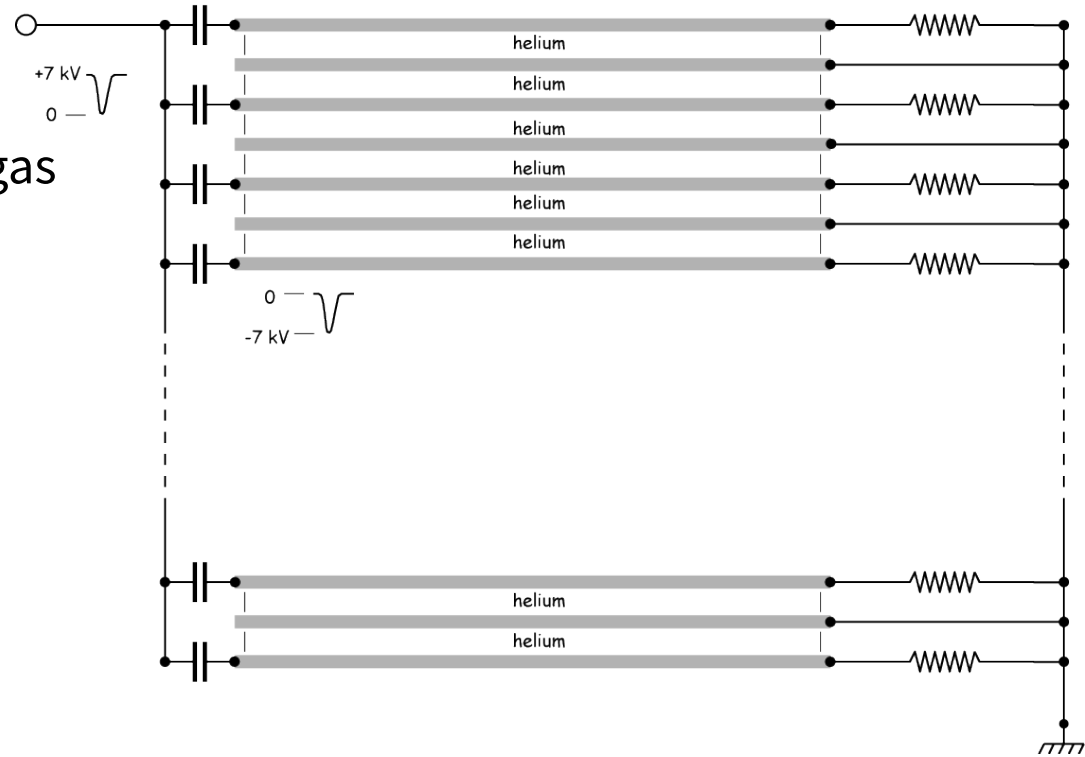
- „Click“ = Particle passing
- „Many Click“ = Many particles
- „A Great Many Click“ = Run...!

- Detector filled with noble gas
 - Charged particles ionize noble gas atoms
 - High voltage applied between electrodes amplifies
- Signal: Current pulse to loudspeaker: Click!



Spark Chamber

- „Many Geiger counters“
- Transparent chamber filled with inert gas and many parallel plates
 - Voltage (\sim kV) between plates
 - Particles ionize noble gas atoms
 - Small sparks along the particle track
- Analysis by photos or microphones
- Relatively large dead time: quenching time of the avalanche
- Used during the 1930s-1960s



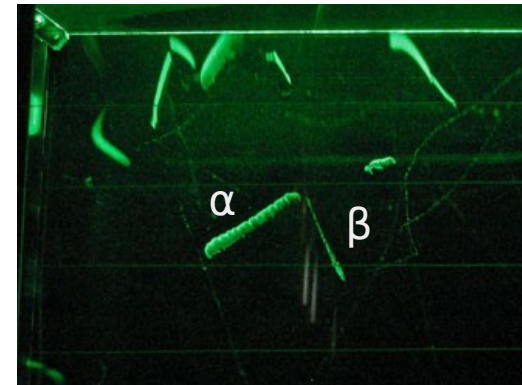
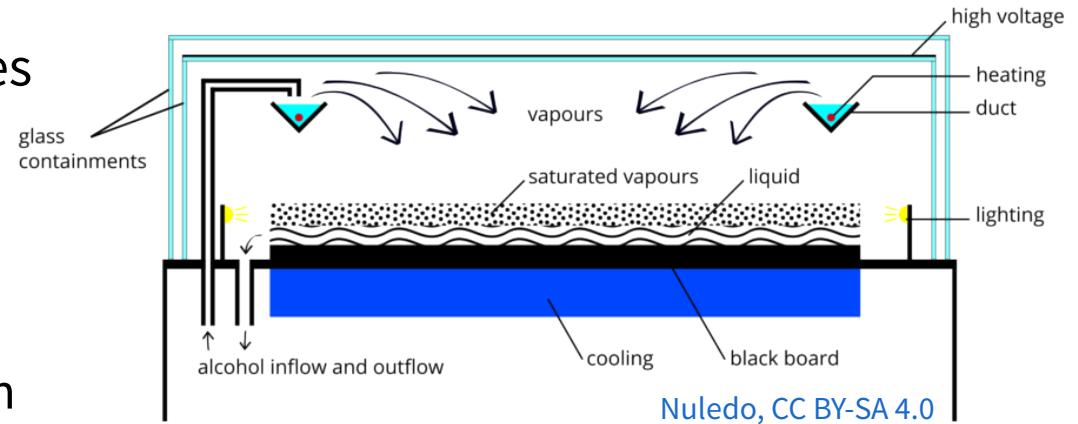


Spark Chamber

Melvin Schwartz at BNL AGS

Cloud Chamber

- Optical detection of charged particles
 - Transparent chamber with supersaturated air-alcohol mixture
 - High-energy charged particle generates ions by impact ionization
 - Ions act as condensation nuclei, droplet formation in the gas mixture
- Type of traces can (sometimes) be identified & assigned to particles
- Nobel Prize 1927 for Charles Thomson Rees Wilson





Cloud Chamber Image
Rendering of particle interaction



Detail of droplets

Cloud chamber tracks after Radon injection

Photo emulsion / Nuclear emulsion

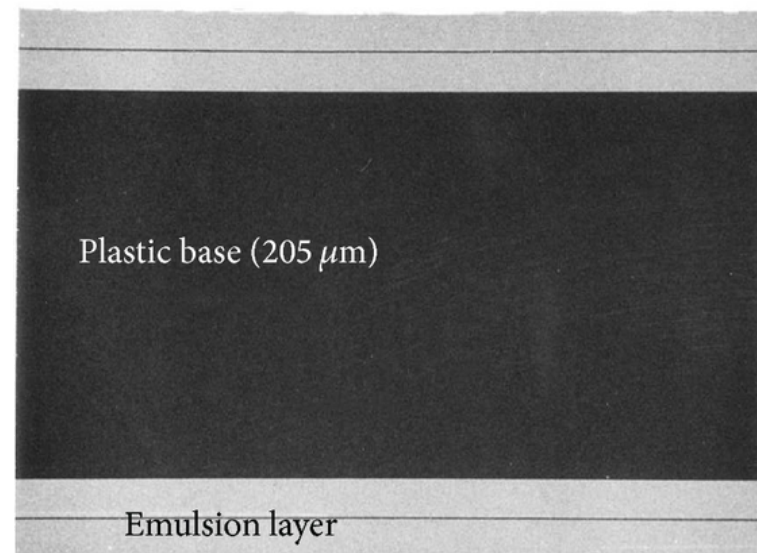
- Photographic plate with thick sensitive layer and very uniform grain size.
 - Ionizing radiation leaves traces
 - Development of the plate
 - Trajectories of particles (blackened by silver) visible with microscope

- Nobel Prize 1950 for Cecil Powell



- OPERA experiment in Gran Sasso

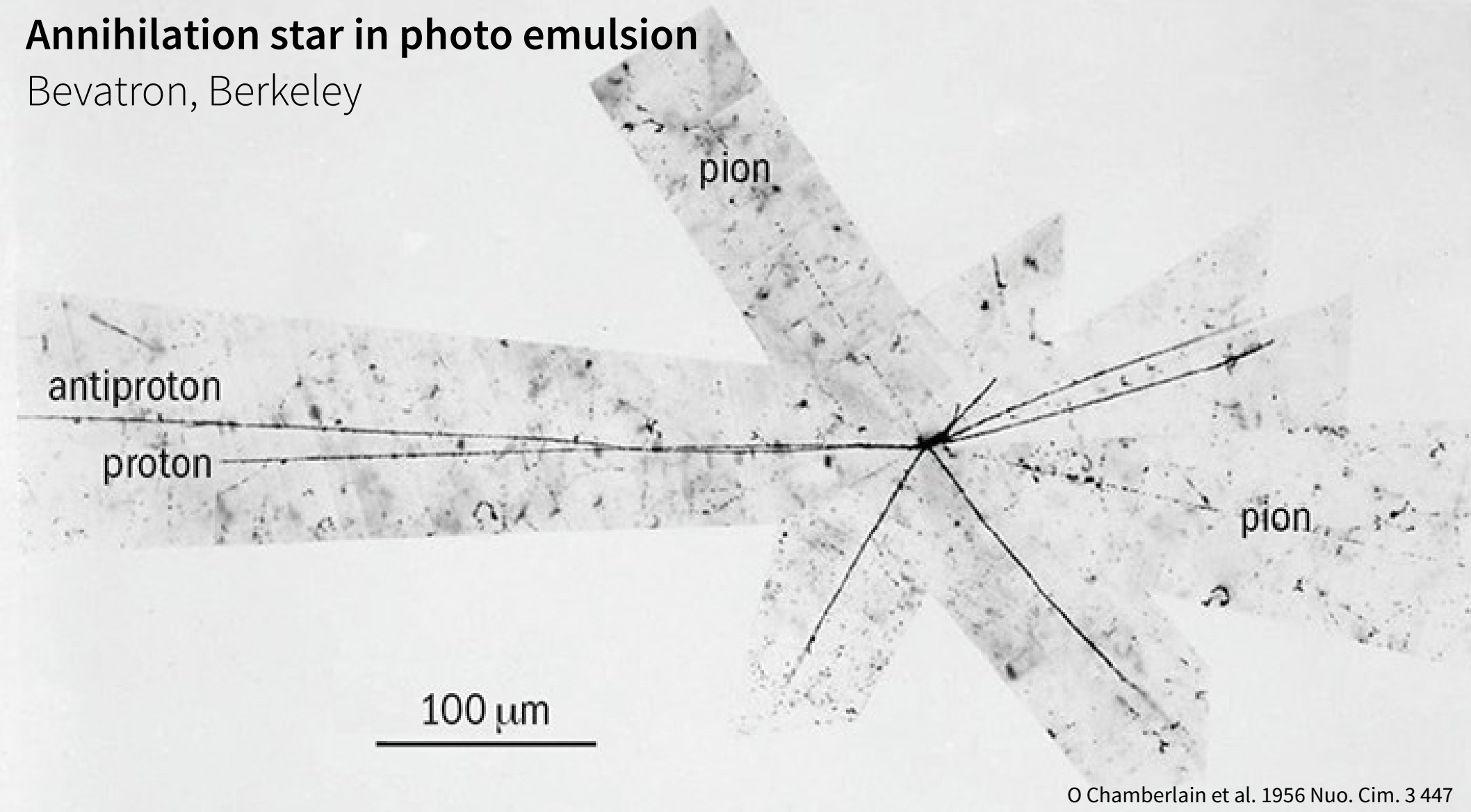
- Photoemulsion plates for particle reconstruction
- AgBr emulsion, cooperation with Fuji Film: 9 million films, each about 10 x 12 cm



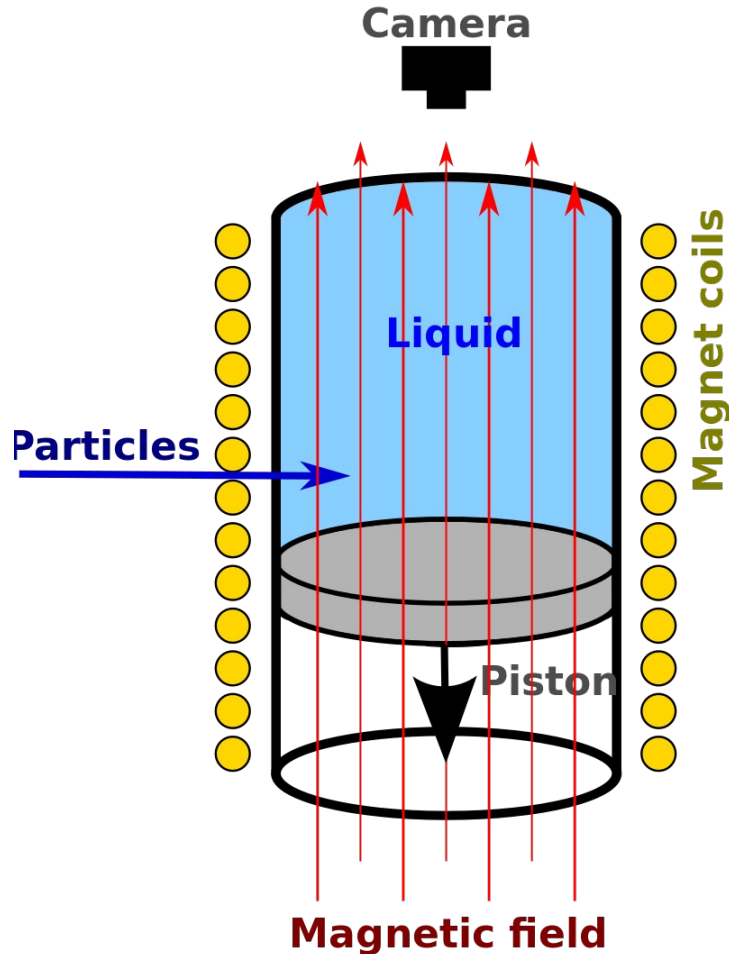
OPERA

Annihilation star in photo emulsion

Bevatron, Berkeley



Bubble Chamber



- Optically transparent chamber filled e.g. with liquid hydrogen
 - Temperature of the liquid close to the boiling point
 - Reduction of chamber pressure with piston
 - Temperature of liquid is now above boiling point
 - Charged particles generate ions along track
 - Ions serve as nuclei for gas bubbles
- Analysis of photos of the traces
- Nobel Prize 1960 for Donald A. Glaser



Prominent Bubble Chambers



- Most important particle detector type in the 1970s
- Particle source: proton synchrotron
- **Gargamelle** - 1970-1978
 - 4.8 m x 1.88 m, 12 000 l
 - First detection of the Z boson
- **BEBC** - 1971-1984
 - Big European Bubble Chamber
 - 3.5 T magnetic field of superconducting coils
 - Discovery of the D meson

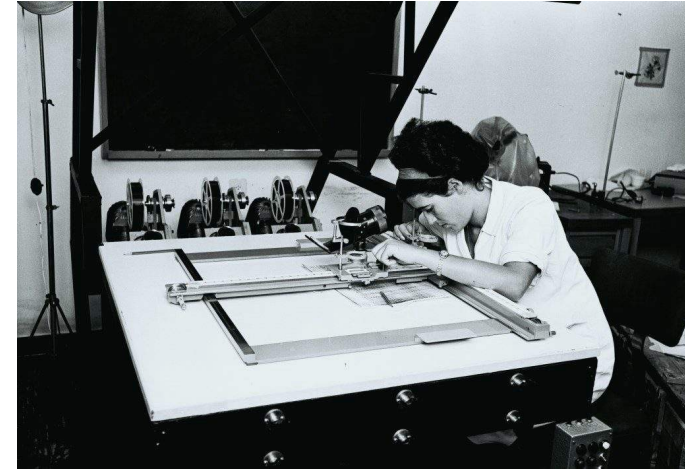
Bubble Chamber Image

Big European Bubble Chamber @ CERN



Problems of Early Particle Detectors

- Many based on photographic images
 - Data evaluation complex and only possible by hand
 - Limitation of the amount of data
- Long detector dead times
 - No further measurement possible until gas avalanche has been quenched/bubbles have disappeared
 - Only low particle rates possible
- Only little information in measurement
 - Information about particle location (& momentum), but no energy/time measurement



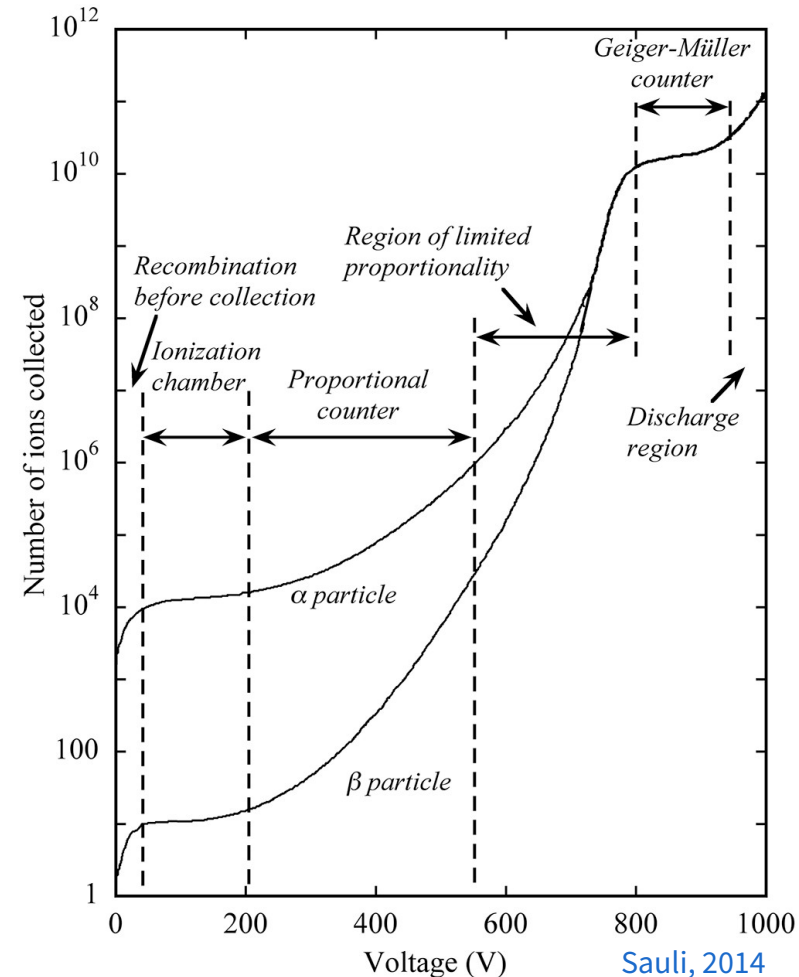
Weizmann Institute Archives

Gaseous Detectors

Primary & secondary ionization of gas atoms

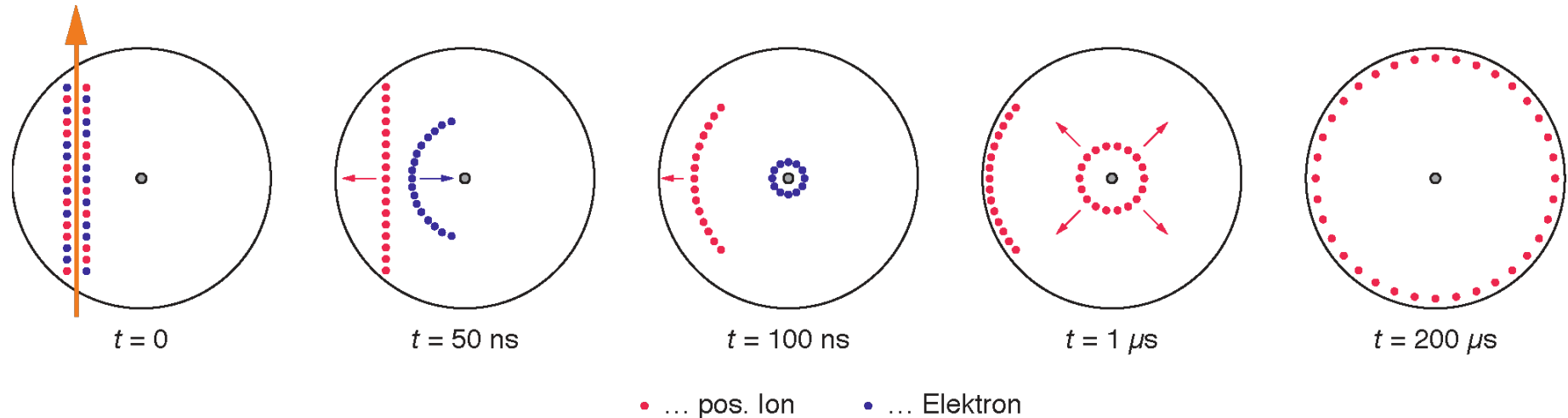
Gaseous Detectors: Operating Principle

- **Primary signal:** charged particle generates electron-ion pairs by ionization
 - Noble gases: relatively low ionization energy
 - Average energy to generate a pair ~ 30 eV
 - Number proportional to deposited energy
- **Amplification:** different working ranges depending on applied voltage
 - Medium voltages: proportional amplification
 - High voltages: Avalanche formation due to secondary ionizations

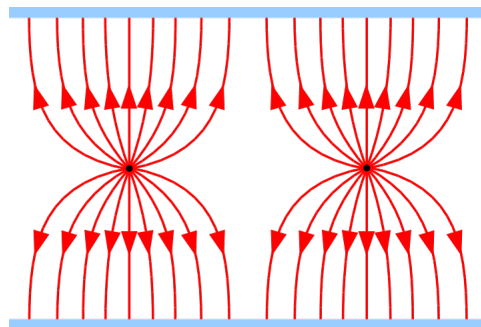
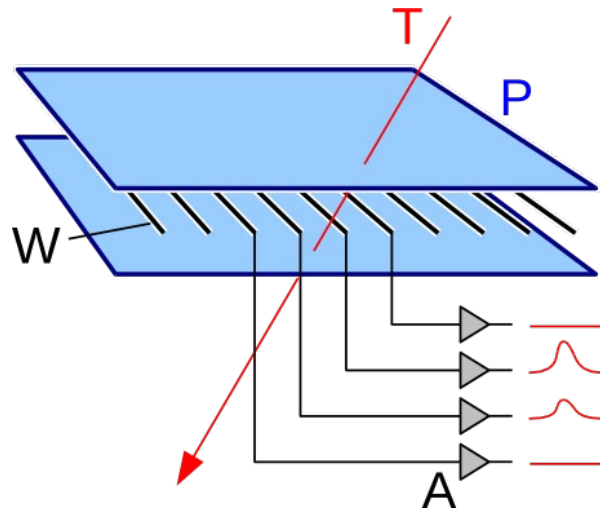


Proportional Counter

- Very similar to the Geiger counter: anode in the form of thin wire
- High field near wire leads to electron multiplication / signal amplification
- Choice of voltage: proportional range
 - Output signal proportional to original number of ionizations



Multi Wire Proportional Chamber (MWPC)

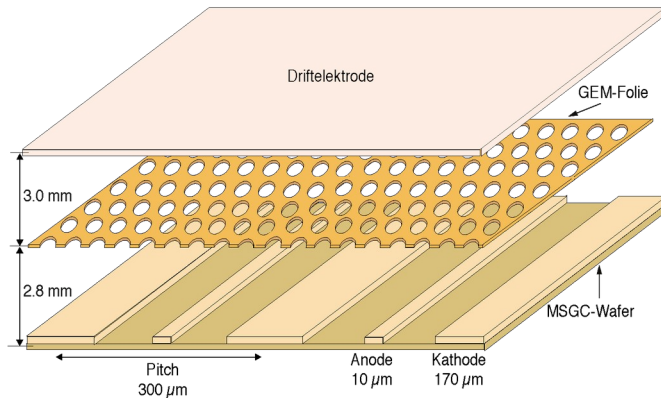


- Essentially many proportional counters next to each other, without separating walls
- Wires spaced a few millimeters apart
 - Good spatial resolution of a traversing particle
 - Large areas possible
 - Electronic selection
- High rates possible: 1000 particles/s
for comparison, bubble chamber: 1-2 particles/s
- Nobel Prize 1992 for Georges Charpak



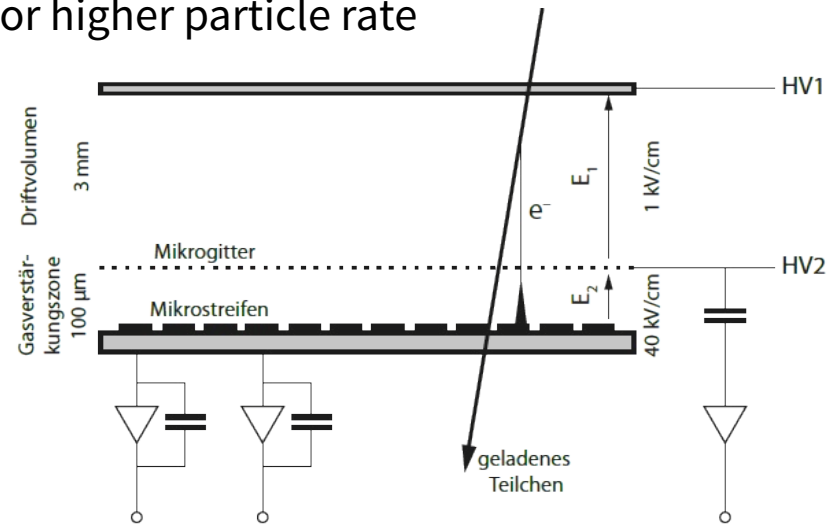
Micropattern Gas Detectors

- Replacement of fragile wires by micro structures
- Potentially better spatial resolution and applicable for higher particle rate



Gas Electron Multiplier (GEM)

- Perforated, metallized Kapton foil, High voltage between electrodes
- Strong dipole field in perforation holes: Gas amplification

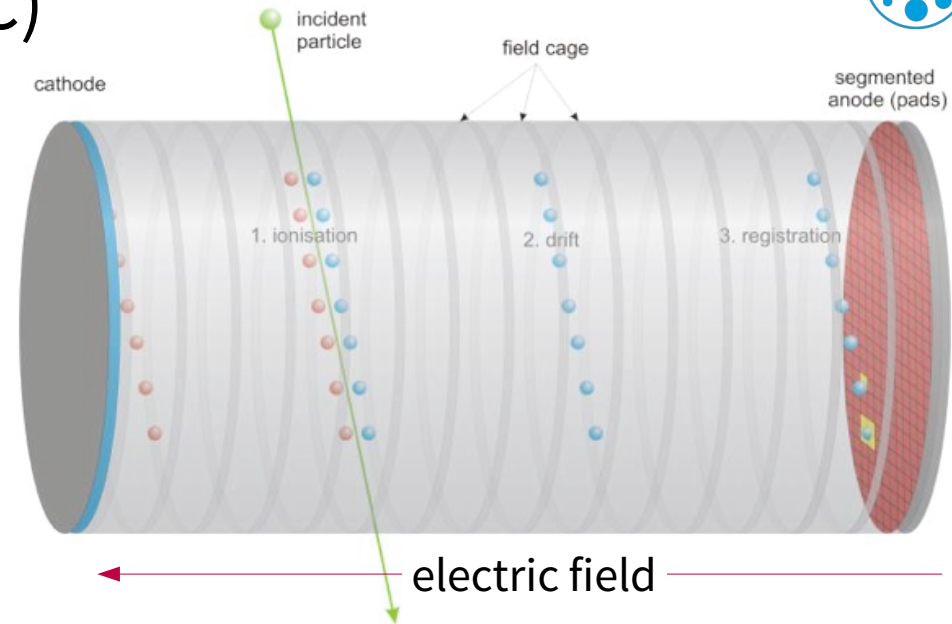


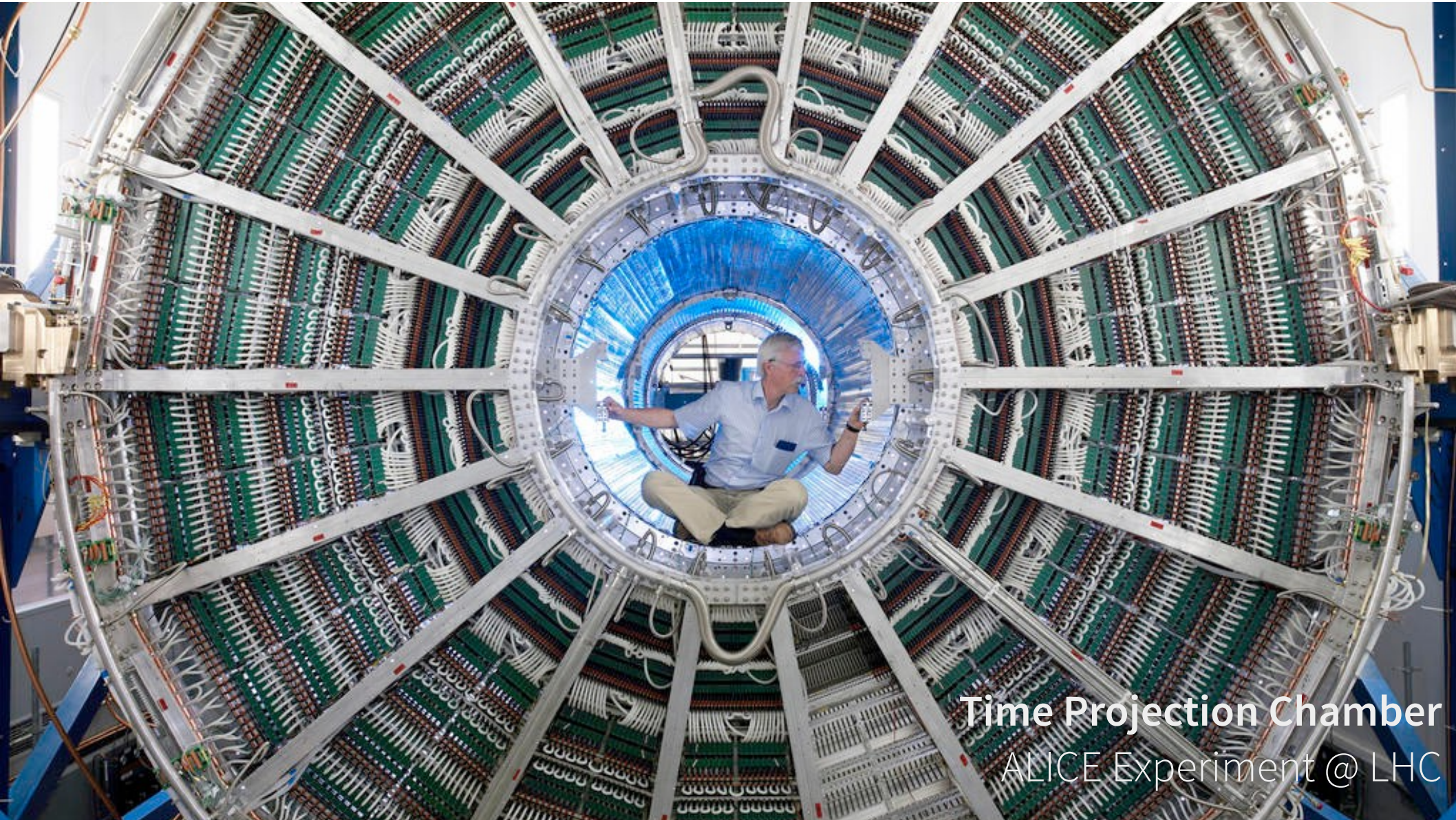
Micro-Mesh Gas Detectors (Micromegas)

- Metallic micro-grid
- Electron avalanche evolution near the lattice: Gas amplification

Time Projection Chambers (TPC)

- Large gas detector system
- Ionization along the particle track
 - Electrons and ions drifting in the E-field
 - Segmented anode: 2D information
 - Measurement of drift time: 3D information
- Readout at anode side e.g. via multi-wire proportional chamber, GEMs, ...



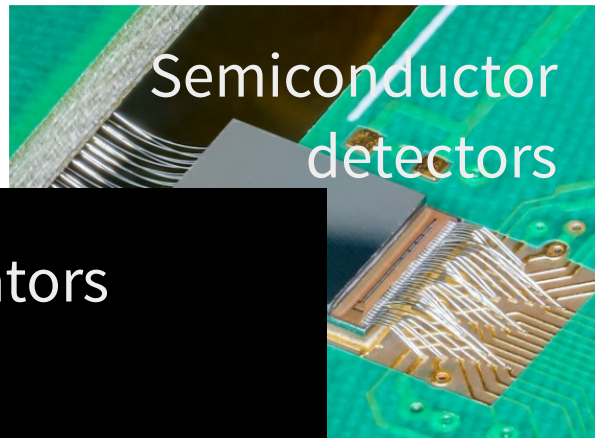


Time Projection Chamber
ALICE Experiment @ LHC

Summary

- Particle detection
 - Measurement is performed by interaction of the particle to be measured with detector material
 - There is no single detector concept that can detect all particle types / properties
 - One needs several detector technologies and concepts
- Interaction of radiation and matter
 - Ionization and excitation of detector atoms, Bremsstrahlung, Cherenkov radiation
 - Photoelectric effect, Compton effect, pair production,
 - Hadronic interaction, missing energy, ...
- Historical detectors
 - Cloud chamber, photo emulsion, bubble chamber, spark chamber
- Gas detectors
 - Operating principle
 - Proportional counters, multi-wire proportional chamber
 - Micropattern gas detectors
 - Time Projection Chamber

Outlook



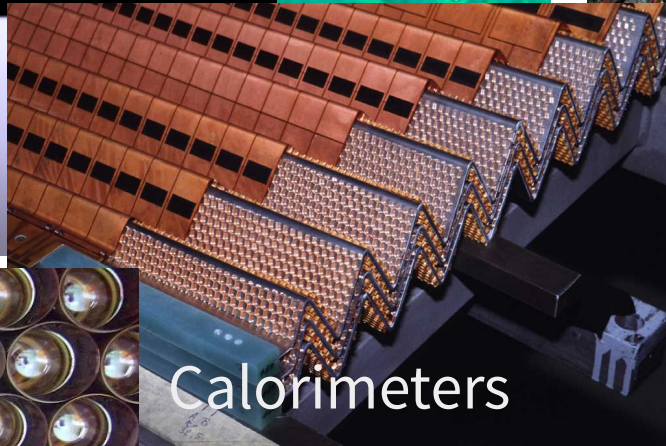
Semiconductor detectors



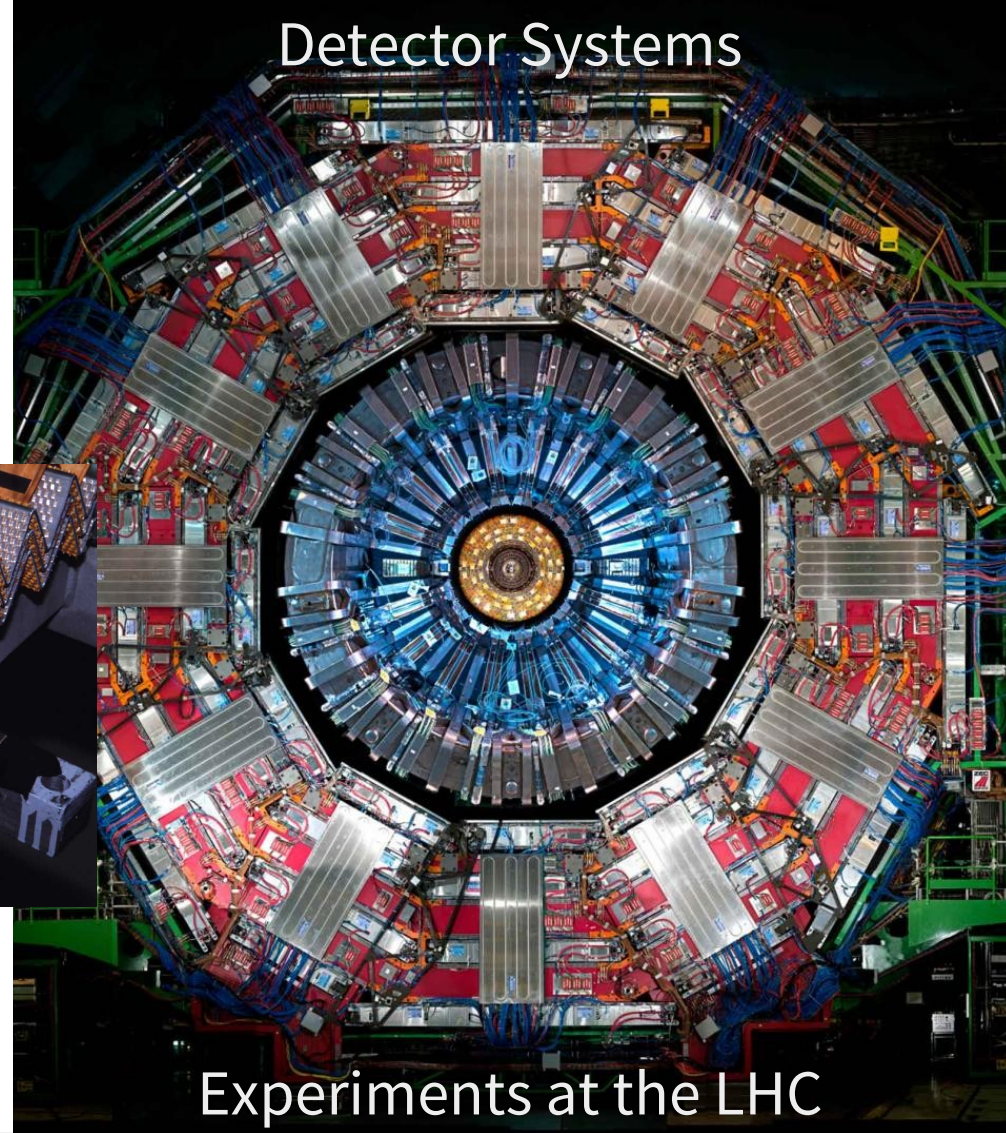
Scintillators



Cherenkov Detectors



Calorimeters



Detector Systems

Experiments at the LHC

