Domain Walls in extended Higgs Sector

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In collaboration with Gudrid Moortgat-Pick

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Introduction to Domain Walls

- Domain walls are a type of topological defects that arise after a spontaneous symmetry breaking of a theory with a discrete symmetry.
- After spontenous symmetry breaking, different regions of the universe can get different vacua which are degenerate with each other.

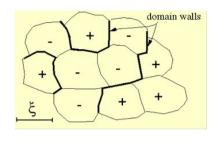
A theory with the symmetry group G broken to a subgroup H

$$f \to H$$

For h element of unbroken group H and a vacuum point Φ_0 : $h\Phi_0 = \Phi_0$

For g element of symmetry group G : $g \Phi_0 = \Phi'$

Acting on Φ_0 with G generates other degenerate vacuum points



H keeps vacuum

points invariant

- The vacuum manifold M of the **standard model** is a **3-Sphere**
- M is not **disconnected** and does **not contain holes** : No **domain walls or cosmic strings or monopoles** in the standard model.
- Beyond standard model physics can have different types of topological defects
- Presence of domain walls in a theory is problematic because they dominate the energy of the universe at some point in time as well as should cause anisotropies in the CMB
 Leads to strong bounds on the energy of domain walls allowed < 0.93 MeV (B. Zel'Dovich, I. Y. Kobzarev, L. B. Okun', Soviet Journal of Experimental and Theoretical Physics 40 (1975) 1)
- Domain walls from **approximate discrete symmetries** get **annihilated** and therefore these models are allowed if the annihilation occurs **before** dominating the energy density of the universe.

Two-Higgs-Doublet Model (2HDM) potential

$$V(\Phi_{1}, \Phi_{2}) = m_{11}^{2} |\Phi_{1}|^{2} + m_{22}^{2} |\Phi_{2}|^{2} - m_{12}(\Phi_{1}^{\dagger}\Phi_{2} + \Phi_{2}^{\dagger}\Phi_{1}) + \frac{\lambda_{1}}{2} |\Phi_{1}|^{4} + \frac{\lambda_{2}}{2} |\Phi_{2}|^{4} + \lambda_{3} |\Phi_{1}|^{2} |\Phi_{2}|^{2} + \lambda_{4}(\Phi_{1}^{\dagger}\Phi_{2})(\Phi_{2}^{\dagger}\Phi_{1}) + \lambda_{5} \left[(\Phi_{2}^{\dagger}\Phi_{1})^{2} + (\Phi_{2}^{\dagger}\Phi_{1})^{2} \right]$$

In the following, focus on Z_2 symmetry (softly broken by m_{12}^2 term)

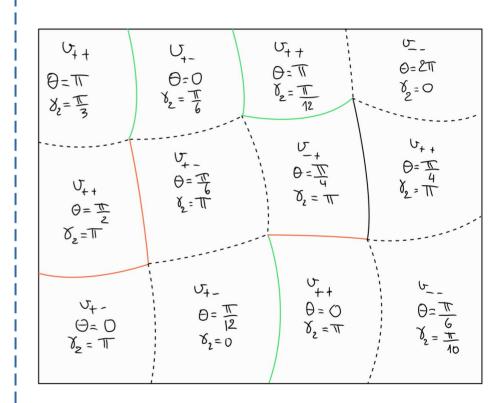
$$\mathbf{Z}_{\mathbf{2}} \quad \Phi_1 \longrightarrow \Phi_1 \quad \Phi_2 \longrightarrow -\Phi_2$$

Full symmetry of the model:

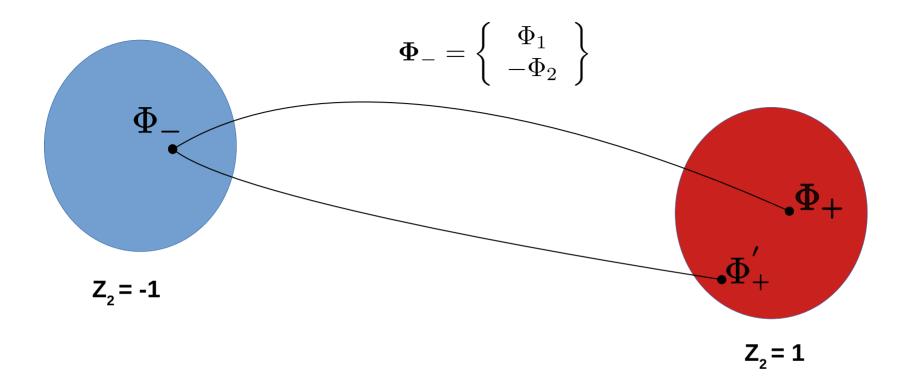
$$SU_L(2) \otimes U_Y(1) \otimes Z_2 \longrightarrow U_{em}(1)$$

 $M = SU(2) \otimes Z_2$ Two disconnected 3-Spheres

In the early universe after SSB, regions with a typical correlation length acquire a random VEV.



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• To obtain domain wall solutions, both minima at -∞ and +∞ should lie on different sectors of the vacuum manifold.

•
$$\Phi_+ = \left\{ egin{array}{c} \Phi_1 \ \Phi_2 \end{array}
ight\}$$
 is a solution, then $\ \Phi_+^{'} = {f U} \Phi_+$ is also a valid solution.

Possible Parametrizations of the Higgs Vacua:

$$\Phi_1 = \begin{pmatrix} \phi_1 + i\phi_2 \\ \phi_3 + i\phi_4 \end{pmatrix}, \quad \Phi_2 = \begin{pmatrix} \phi_5 + i\phi_6 \\ \phi_7 + i\phi_8 \end{pmatrix}$$

Linear Parametrization

Can be rotated using a matrix $U(x) \in SU_L(2) \otimes U_Y(1)$

$$U = e^{i\theta(x)} exp(\frac{i\gamma_i(x)\sigma_i}{2})$$

Vacuum Parametrization

$$\Phi_{i}(x) = U(x)\hat{\Phi}_{i}(x)$$

$$\hat{\Phi}_{1}(x) = \begin{pmatrix} 0\\v_{1}(x) \end{pmatrix}, \hat{\Phi}_{2}(x) = \begin{pmatrix} v_{+}(x)\\v_{2}(x)e^{i\xi(x)} \end{pmatrix}$$

$$U = e^{i\theta} \begin{pmatrix} \cos(\gamma_{1})\exp(i\gamma_{2}) & \sin(\gamma_{1})\exp(i\gamma_{3})\\-\sin(\gamma_{1})\exp(-i\gamma_{3}) & \cos(\gamma_{1})\exp(-i\gamma_{2}) \end{pmatrix}$$

Possible Vacua in the 2HDM:

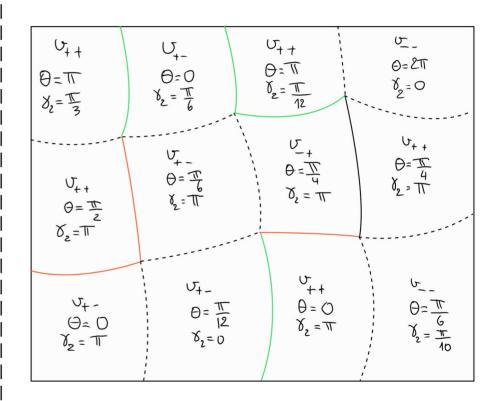
- Neutral Vacua $v_+ = 0, \xi = 0$
- CP breaking Vacua $\xi \neq 0, v_+ = 0$
- Charge breaking Vacua $v_+ \neq 0$

Here we only consider **neutral vacua at the boundaries** and take the general Vacuum Parametrization at each point in x (**possibility of getting CP and/or Charge violation inside the domain wall**)

$$\Phi_1(\pm\infty) = \frac{1}{\sqrt{2}} \begin{pmatrix} 0\\v_1^0 \end{pmatrix}, \qquad \Phi_2(\pm\infty) = \frac{1}{\sqrt{2}} \begin{pmatrix} 0\\\pm v_2^0 \end{pmatrix}$$

$$U = e^{i\theta} exp(\frac{i\gamma_i\sigma_i}{2})$$

- After SSB, different patches of the universe with length scale ξ (correlation length) can acquire VEVs in the two different sectors of the vacuum manifold.
- Because we break the electroweak symmetry at the same time as the discrete symmetry, we can get Abelian and Non-Abelian domain wall solutions. In contrast to the standard domain walls models where only the discrete symmetry gets broken.
 (arXiv:hep-th/0105128 for the case of SU(5)xZ₂
 Domain Walls)
- We end up with **several classes of domain walls** which can **break CP, Charge** or both **inside** the domain wall.



How to find static domain wall solutions ?

Minimize the energy functional of the vacuum configuration $\Phi(x)$

$$E = \int dx \ \mathcal{E}(x), \qquad \mathcal{E}(x) = \frac{d\Phi_i}{dx} \frac{d\Phi_i^{\dagger}}{dx} + V(\Phi_1, \Phi_2)$$

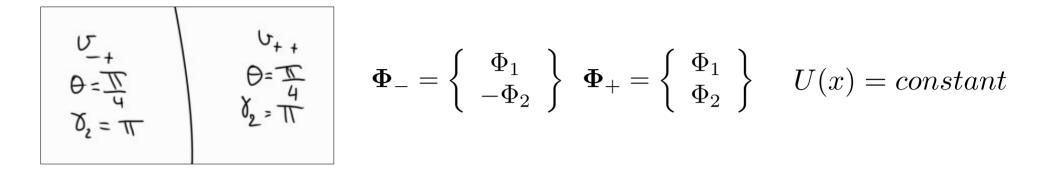
Using $\Phi_i(x) = U(x)\hat{\Phi}(x)$ and $\frac{dU}{dx} = i(\frac{d\theta}{dx} + \frac{1}{2}\frac{d\gamma_i}{dx}\sigma_i)U(x),$ we get:

$$\mathcal{E}(x) = \sum_{k} i \frac{d\hat{\Phi}_{k}^{\dagger}}{dx} U^{\dagger}(x) \left(\frac{d\theta}{dx} + \frac{1}{2}\frac{\gamma_{i}}{dx}\sigma_{i}\right) U(x)\Phi_{k}(x) + \text{h.c} + \left\|\frac{d\hat{\Phi}_{k}}{dx}\right\|^{2} + \left\|\hat{\Phi}_{k}\right\|^{2} \left[\left(\frac{d\theta}{dx}\right)^{2} + \frac{1}{4}\left(\frac{d\gamma_{i}}{dx}\right)^{2}\right] + \hat{\Phi}_{k}\frac{d\theta}{dx}\frac{d\gamma_{i}}{dx}\sigma_{i}\hat{\Phi}_{k}^{\dagger} + V(\Phi_{1},\Phi_{2})$$

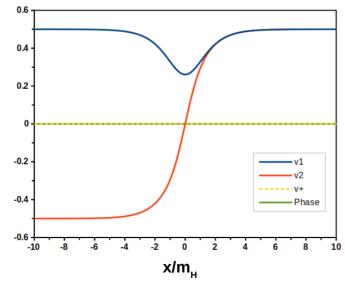
Minimization of E yields the **equation of motion for static fields** (analogous to minimizing the action)

$$\frac{d}{dx} \left(\frac{\partial \mathcal{E}}{\partial \left(d\phi_n / dx \right)} \right) - \frac{\partial \mathcal{E}}{\partial \phi_n} = 0$$

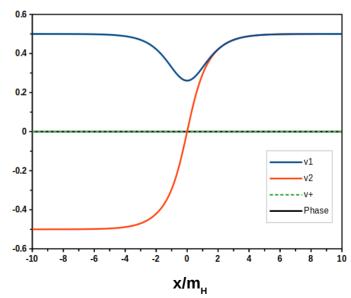
Simplest case : Standard Domain Walls Solution



Domain Wall Solution in the Linear Parametrization



Domain Wall Solution in the non-Linear Parametrization

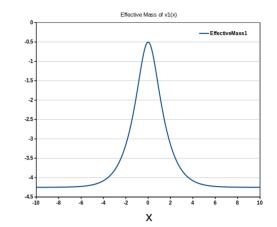


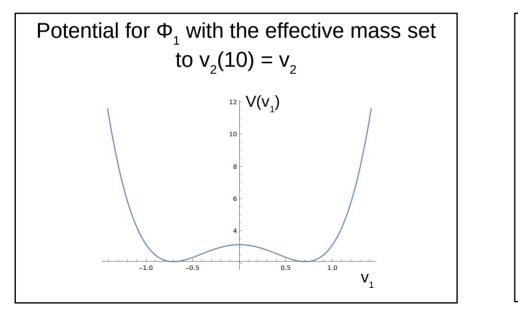
Explaining the behavior of $v_1(x)$

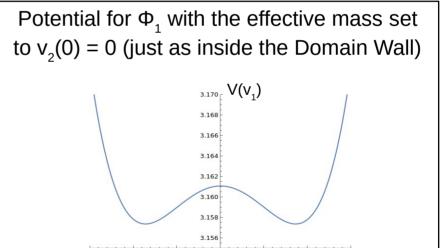
Calculate the **effective mass** term for $v_1(x)$:

$$M_{eff1} = \frac{1}{2}m_{11}^2 + \frac{1}{4}\left[(\lambda_3 + \lambda_4)v_2^2(x) + \lambda_3 v_+^2(x) + \lambda_5 v_2^2(x)\cos(2\xi(x))\right]$$

Becomes bigger (less negative) inside the domain wall !







0.0

0.1

0.2

0.3

V₁

-0.2

-0.1

Leads to $v_1(0)$ getting a smaller value inside the Domain Wall

Example for simplified case where only $\theta(x)$ is non constant :

$$U(x) = e^{i\theta(x)}$$

$$\mathcal{E}(x) = \frac{1}{2}\left(\frac{dv_1}{dx}\right)^2 + \frac{1}{2}\left(\frac{dv_2}{dx}\right)^2 + \frac{1}{2}\left(\frac{dv_+}{dx}\right)^2 + v_2^2\left(\frac{d\xi}{dx}\right)^2 + \frac{1}{2}\left(\frac{d\theta}{dx}\right)^2\left(v_1^2 + v_2^2 + v_+^2\right) + v_2^2\frac{d\theta}{dx}\frac{d\xi}{dx} + V(v_1, v_2, v_+, \xi)$$

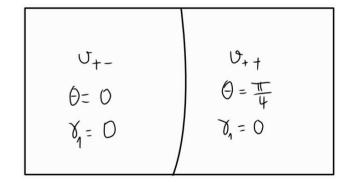
Leads to the equations of motion:

 $\frac{d^2 v_1}{dx^2} - \left(\frac{d\theta}{dx}\right)^2 v_1 - \frac{dV}{dv_1} = 0$ $\frac{d^2 v_2}{dx^2} - v_2 \left(\frac{d\xi}{dx} + \frac{d\theta}{dx}\right)^2 - \frac{dV}{dv_2} = 0$ $\frac{d^2 v_+}{dx^2} - v_+ \left(\frac{d\theta}{dx}\right)^2 - \frac{dV}{dv_+} = 0$ $w^2 \frac{d^2 \xi}{dx^2} + w^2 \frac{d^2 \theta}{dx^2} + 2w \frac{dv_2 d\xi}{dx^2} + 2w$

Solve numerically using Gradient flow method Richard A. Battye, Gary D. Brawn, Apostolos Pilaftsis (1106.3482) JHEP Richard A. Battye, Apostolos Pilaftsis, Dominic G. Viatic (2006.13273) JHEP

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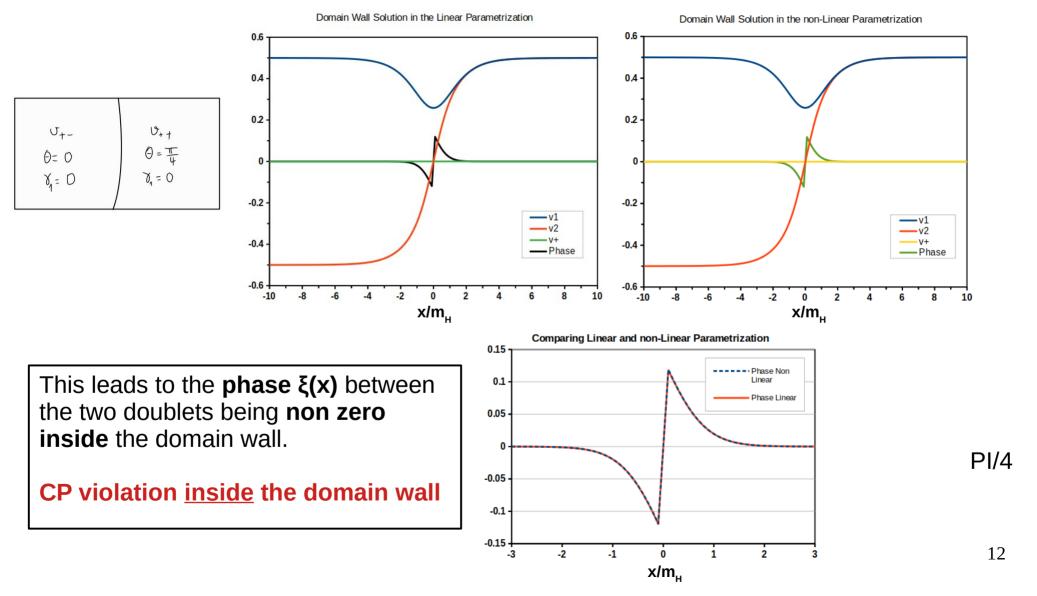
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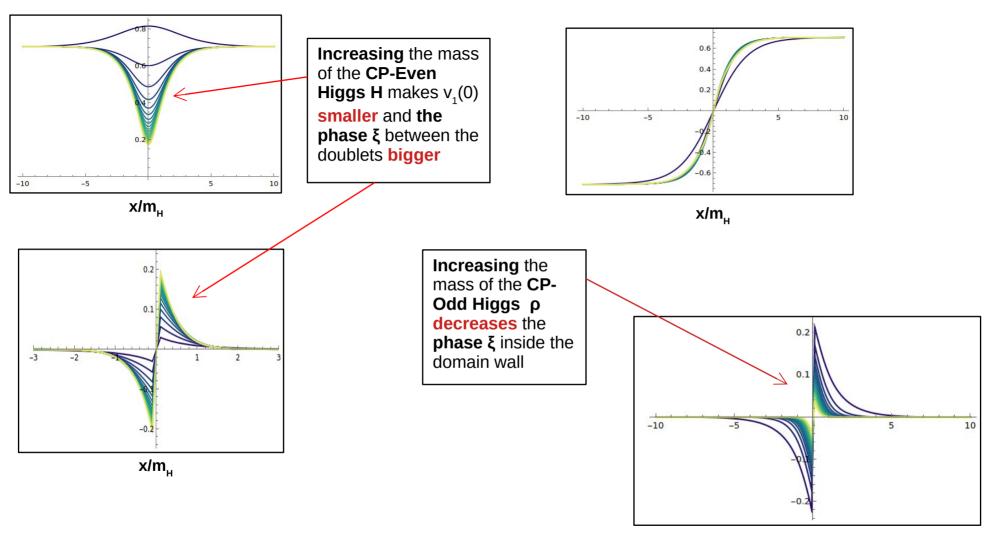


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$$v_2^2 \frac{d^2\xi}{dx^2} + v_2^2 \frac{d^2\theta}{dx^2} + 2v_2 \frac{dv_2}{dx} \frac{d\xi}{dx} + 2v_2 \frac{dv_2}{dx} \frac{d\theta}{dx} - \frac{dV}{d\xi} = 0$$

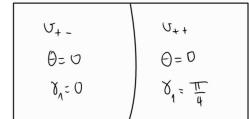
$$(v_1^2 + v_2^2 + v_+^2)\frac{d^2\theta}{dx^2} + 2\frac{d\theta}{dx}(v_1\frac{dv_1}{dx} + v_2\frac{dv_2}{dx} + v_+\frac{dv_+}{dx}) + v_2^2\frac{d^2\xi}{dx^2} + 2v_2\frac{dv_2}{dx}\frac{d\xi}{dx} = 0$$

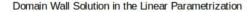


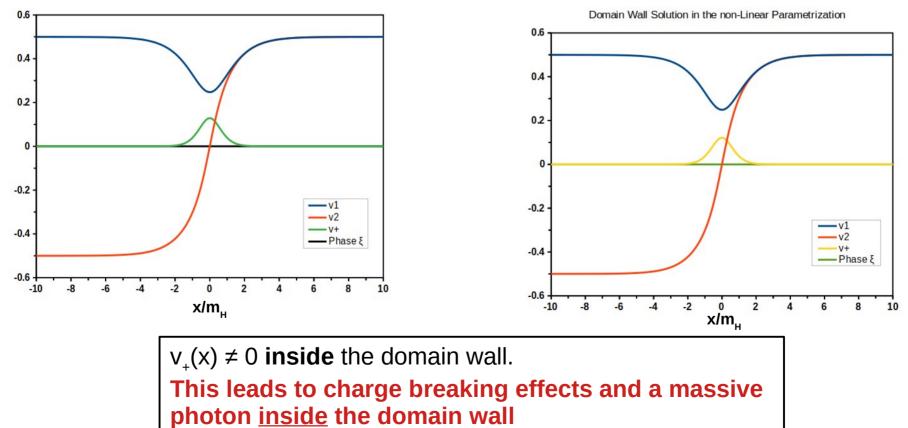


Case where y₂ is non constant :

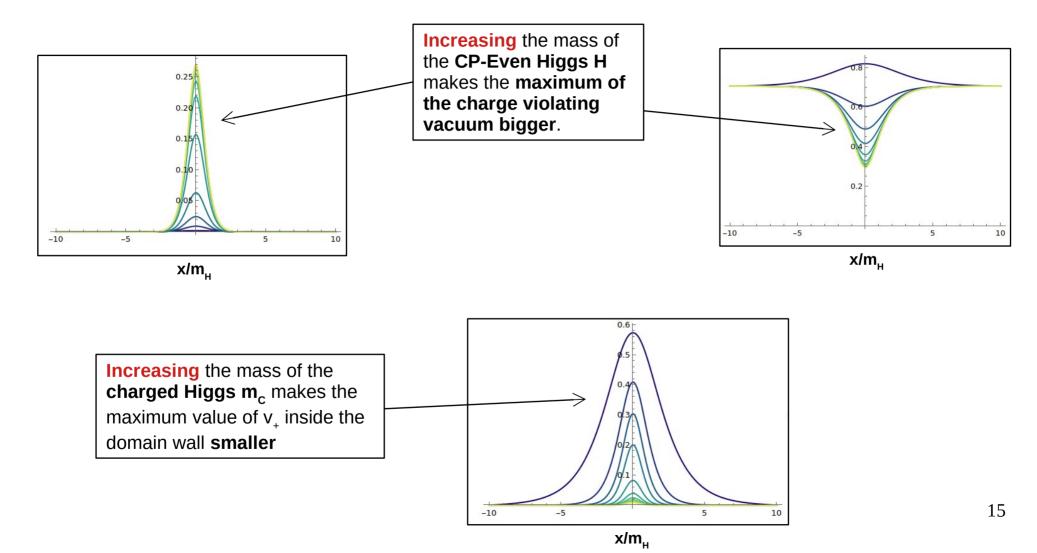
$$U(x) = exp(\frac{i\gamma_1(x)\sigma_1}{2}) \quad \text{As an example take } \gamma_1(x) = \left\{ \begin{array}{c} 0 \text{ at } -\infty \\ \frac{\pi}{4} \text{ at } +\infty \end{array} \right\} \quad \left| \begin{array}{c} \theta_z \\ \theta_z \\ \eta_z \end{array} \right|_{\eta_z}$$



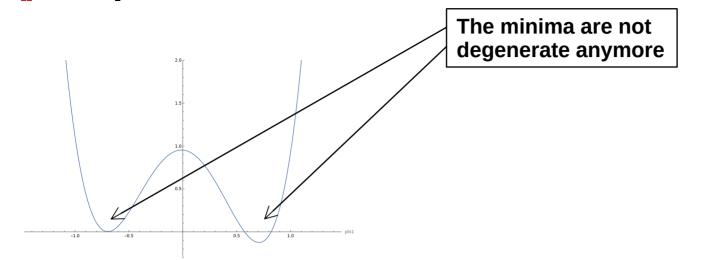


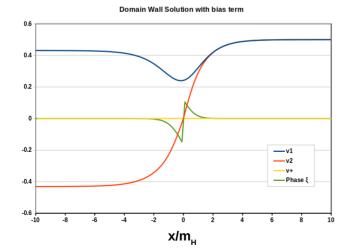


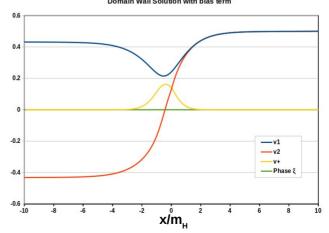
Variation of mC and mH for the charge breaking DW



For $m_{12} \neq 0$, the Z_2 symmetry is then approximate and we get asymmetric domain walls:







Domain Wall Solution with bias term

Baryogenesis with Topological Defects

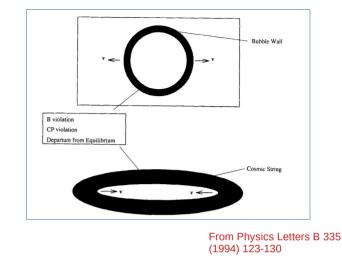
Idea discussed in the 90s :

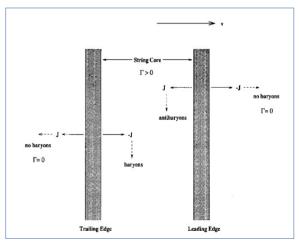
- Local and nonlocal defect-mediated electroweak baryogenesis hep-ph/9409281
- Baryogenesis from Domain Walls in the Next-to-Minimal Supersymmetric Standard Model hep-ph/9505241
- Electroweak Baryogenesis with Cosmic Strings ?
 Hep-ph/9901310

Main idea:

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- The topological defect acts as the bubble wall.
- Sphalerons are less supressed inside the topological defect
- CP violation in the defect walls.





Main Problems discussed in past papers:

Volume suppression factor due to defect not spanning the whole universe.

$$\Delta n_B = \frac{1}{V} \frac{\Gamma_B}{T} V_{\rm BG} \Delta \theta$$

• Symmetry restoration region <u>not large enough</u> to contain Sphalerons.

$$R_{restoration} \sim \frac{1}{\sqrt{\lambda}v} \qquad R_{Sphalerons} \sim \frac{1}{g^2T}$$

For N2HDM/2HDM Domain walls thickness 5-10 times smaller than Sphalerons

For cosmic strings

String-mediated electroweak baryogenesis: A critical analysis

J. M. Cline,^{1,*} J. R. Espinosa,^{2,†} G. D. Moore,^{1,‡} and A. Riotto^{2,§} ¹Department of Physics, McGill University, 3600 University Street, Montréal, Québec, Canada H3A 278 ²CERN TH-Division, CH-1211 Geneva 23, Switzerland (Received 6 October 1998; published 22 February 1999)

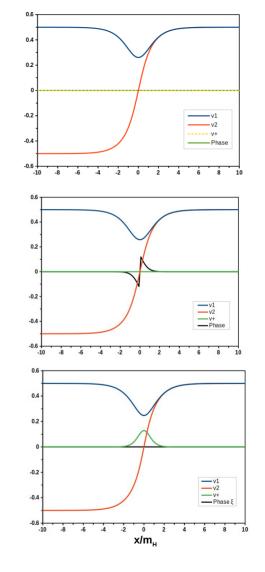
Very Suppressed!

$$\left[\frac{N_B}{N_\gamma}\right]_{strings} \lesssim 10^{-10} \left[\frac{N_B}{N_\gamma}\right]_{observed}.$$
 (4)

That is, the mechanism just studied is uncapable of generating a sufficiently large matter-antimatter asymmetry.

What about Domain Walls ?

- Due to breaking SU(2)xU(1) symmetry along the Z₂ symmetry, the domain walls in the 2HDM can have CP and Charge violation inside the wall.
- Bias term for the Z_2 symmetry leads to asymmetry in the breaking of CP in the different parts of the wall.
- Possible future directions include :
- 1)Electroweak Baryogenesis using the domain walls.
- 2)Using charge violation inside the domain wall to constrain the model.
- 3)Probing gravitational wave spectrum from annihilating domain walls.



Backup

Increasing m_{H} leads to effective mass becoming **more negative** outside the Domain wall and **less negative inside the domain wall**, which leads to $v_1(0)$ becoming smaller with higher m_{H} . However, for the first parameter point (m_{H} = 80 GeV), the effective mass is more negative inside the domain wall.

Terms influencing $\xi(x)$:

Always positive

The second term needs to be negative to make sure that the behaviour of $\xi(x)$ minimizes the energy.

We find from the equation of motion for $\theta(x)$:

$d heta$ _	v_2^2	$d\xi$
$\frac{dx}{dx}$ – –	$\overline{v_1^2 + v_2^2 + v_+^2}$	dx

Second term is more negative for smaller $v_1(x)$ and bigger gradient for $\xi(x)$, gives a more negative contribution than the third term for bigger $\xi(x)$

