

Reactor Neutrino Energy Reconstruction with Machine Learning Techniques for the JUNO Experiment Arsenii Gavrikov^{1,2} on behalf of the JUNO collaboration ¹The University of Padova, Italy & ²INFN — Sezione di Padova, Italy

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Introduction

- Jiangmen Underground Neutrino Observatory (JUNO):
 - **multipurpose** experiment;
 - 53 km away from **8 reactor cores** in China
 - \sim 650-meter deep underground
 - data taking expected in \sim 2024
- The main goals of JUNO:
 - neutrino mass ordering (3σ in 6 years)
 - [†]precise measure of oscillation parameters $\sin^2 \theta_{12}, \Delta m_{21}^2, \Delta m_{31}^2$
- The Central Detector:
 - detection channel: $\overline{V}_e + p \rightarrow e^+ + n$
 - **deposited energy** converts to optical light
 - the largest liquid scintillator detector: 20 kt
 - 77.9% photo-coverage by photo-multiplier tubes (PMTs): ~18k 20" (LPMT), ~26k 3" (SPMT)

[†]see poster by V. Cerrone: Prospects for Oscillation Physics in the JUNO Experiment

Problem statement







Feature selection

- **Feature selection** procedure is performed with *a greedy algorithm* using Boosted Decision Trees
- Optimized **set of features** (sorted by *importance*):



Dashed line — average [†]MAPE for BDT trained on all features with its standard deviation

[†]MAPE = $\frac{100\%}{N} \sum_{i=1}^{N} \left| \frac{E_{\text{true},i} - E_{\text{pred},i}}{E_{\text{true},i}} \right|$, where *N* - number of events

Models description

Fully Connected Deep Neural Network (FCDNN):

Input layer	÷	32 hidden layers		
64 units	:	h_1	256 units	h ₃₂
	÷			

• Optimization of the hyperparameters using BayesianOptimization

Data description

To train a model and **to evaluate** model performance we prepared **two datasets** generated by the full detector Monte Carlo method using the official JUNO software:

- **1** Training dataset:
 - 2.25 million positron events
 - uniformly distributed in kinetic energy E_{kin}
 - uniformly spread in the volume of the central detector
 - $E_{\text{kin}} \in [0, 10]$ MeV. $E_{\text{dep}} = E_{\text{kin}} + 1.022$ MeV
- **Testing dataset:**
- subsets with discrete kinetic energies
- 0, 0.1, 0.3, 0.6, 1, 2, ..., 10 [MeV]
- uniform spatial distribution
- each subset contains **50 thousand** events



For energy reconstruction, we use **aggregated information** from the whole array of PMTs **as features** for models. Their *full* set is as follows: $\boldsymbol{5} \quad \boldsymbol{\gamma}_x^{\rm cc} = \frac{x_{\rm cc}}{\sqrt{z_{\rm cc}^2 + y_{\rm cc}^2}}$

- AccumCharge the accumulated charge on fired PMTs
- nPMTs the total number of fired PMTs 2
- Coordinates of the center of charge: 3

$$x_{cc}, y_{cc}, z_{cc}) = \vec{r}_{cc} = \frac{\sum_{i=1}^{N_{LPMTs}} \vec{r}_{LPMT_i} \cdot n_{p.e.,i} + \sum_{i=1}^{N_{SPMTs}} \vec{r}_{SPMT_i} \cdot n_{p.e.,i}}{\sum_{i=1}^{N_{LPMTs}} n_{p.e.,i} + \sum_{i=1}^{N_{SPMTs}} n_{p.e.,i}}$$

and its radial component: $R_{\rm cc} = |\vec{r}_{\rm cc}|$

4 Coordinates of the center of FHT:

$$(x_{\rm cht}, y_{\rm cht}, z_{\rm cht}) = \vec{r}_{\rm cht} = \frac{\sum_{i=1}^{N_{\rm LPMTs}} \frac{\vec{r}_{\rm LPMT_i}}{t_{\rm fht,i}+c} + \sum_{i=1}^{N_{\rm SPMTs}} \frac{\vec{r}_{\rm SPMT_i}}{t_{\rm fht,i}+c}}{\sum_{i=1}^{N_{\rm LPMTs}} \frac{1}{t_{\rm fht,i}+c} + \sum_{i=1}^{N_{\rm SPMTs}} \frac{1}{t_{\rm fht,i}+c}}$$

and its radial component: $R_{\rm cht} = |\vec{r}_{\rm cht}|$

- $\theta_{\rm cc} = \arctan \frac{\sqrt{x_{\rm cc}^2 + y_{\rm cc}^2}}{z_{\rm cc}}$ $\gamma_{y}^{cc} = \frac{\gamma_{cc}}{\sqrt{x_{cc}^{2} + z_{cc}^{2}}}$ $\gamma_{z}^{cc} = \frac{z_{cc}}{\sqrt{x_{cc}^{2} + y_{cc}^{2}}}$ $\rho_{\rm cc} = \sqrt{x_{\rm cc}^2 + y_{\rm cc}^2}$
- with 7 similar features for the components of the center of FHT
- Percentiles of FHT and charge distributions:
 - ${ht_{2\%}, ht_{5\%}, ht_{10\%}, ht_{15\%}, ..., ht_{90\%}, ht_{95\%}}$
 - { $pe_{2\%}, pe_{5\%}, pe_{10\%}, pe_{15\%}, ..., pe_{90\%}, pe_{95\%}$ }
- Ø Differences between percentiles for FHT:
- { $ht_{5\%-2\%}$, $ht_{10\%-5\%}$, ..., $ht_{95\%-90\%}$ }
- Moments for FHT and charge distributions:
 - { ht_{mean} , ht_{std} , ht_{skew} , $ht_{kurtosis}$ }



Boosted Decision Trees (BDT) from XGBoost:



- Training with early stopping
- Validation dataset: 200k events
- The optimized set of features
- The optimized set of features
- Optimized hyperparameters (using Grid Search):
 - The maximum depth of the tree: 11
 - **2** Number of trees in the ensemble: $\simeq 300$
 - **3** Learning rate: 0.08

Results

Metrics:

- Defined by a Gaussian fit of the $E_{\text{predicted}} - E_{\text{dep}}$ distributions
- **2** <u>*Resolution*</u>: σ/E_{dep} , where σ standard deviation of the fit
- **3** <u>*Bias*</u> μ/E_{dep} , where μ mean of the fit



Parameterization:

 $\frac{\sigma}{E_{\rm dep}} = \sqrt{\left(\frac{a}{\sqrt{E_{\rm dep}}}\right)^2 + b^2 + \left(\frac{c}{E_{\rm dep}}\right)^2}$

Model $a \pm \Delta a$ $b \pm \Delta b$ $c \pm \Delta c$ BDT $2.56 \pm 0.11 \mid 0.67 \pm 0.05 \mid 1.18 \pm 0.32$ FCDNN 2.35 ± 0.11 0.72 ± 0.05 1.62 ± 0.20

(FCDNN and BDT) using aggregated features: • energy reconstruction

Machine learning approaches

Summary

- required energy resolution < 3% @ 1 MeV achieved
- great computation speed

References

- [1] F. An et al., Neutrino Physics with JUNO, J. Phys. G 43 (2016) no.3, 030401.
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Acknowledgments

We are immensely grateful to Yury Malyshkin and Fedor Ratnikov for their invaluable contribution to this work. We also are very thankful to JUNO collaborators, who contributed to developing and validating the JUNO simulation software.

This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie Grant Agreement No. 101034319 and from the European Union – NextGenerationEU.









Transferability of the approach: TAO

- [†]Taishan Antineutrino Observatory (TAO): 1
 - JUNO's satellite near detector
 - 30 m from Taishan core
 - diameter of **1.8 meters**
- The main goals of TAO:
 - **reference spectrum** for JUNO
 - search for sterile neutrinos
 - isotopic yields and spectra
- **3** The detector:
 - the liquid scintillator detector of **2.8 tons**
 - ~94% photo-coverage: $4k 5x5 \text{ cm}^2 \text{ SiPMs}$
 - required energy resolution of 2% @ 1 MeV
- Analogues datasets for training and testing
- BDT & FCDNN trained on an *optimized set* of features:
 - AccCharge ht_{5%} 1 7 $R_{\rm cc}$ $\rho_{\rm cc}$ ht_{35%} 3 9 pestd 10 ht_{75%} pe90% 1 pekurtosis pe_{mean} ht_{15%} nSiPMs 12
- $\sim 2\%$ at 1 MeV *achieved*

[†]more details about TAO detector in poster by **C. Lombardo**: JUNO-TAO design, prototype and its impact for JUNO physics





