



Contribution ID: 455

Type: Poster

Dark matter and axion dark radiation properties from the dynamics of self-gravitating collisionless dark matter

We present a new theory to predict dark matter (DM) particle mass, size, lifetime, and properties of possible dark radiation from DM particle decay. In the dynamics of self-gravitating collisionless dark matter, the existence of inverse mass and energy cascade from small to large scales facilitates the hierarchical structure formation. A scale-independent constant rate of energy cascade $\varepsilon_u \approx -4.6 \times 10^{-7} m^2/s^3$ can be identified. The energy cascade leads to a two-thirds law for pairwise velocity and a four-thirds law for halo density. Both scaling laws can be directly confirmed by N-body simulations and galaxy rotation curves. For collisionless dark matter with only gravity involved, scaling laws can be extended down to the smallest scale, where quantum effects become important. Combining ε_u , Planck constant \hbar , and gravitational constant G on that scale, we predict DM particles have a mass $m_X = (\varepsilon_u \hbar^5 G^{-4})^{1/9} = 0.9 \times 10^{12} \text{ GeV}$, a size $l_X = (\varepsilon_u^{-1} \hbar G)^{1/3} = 3 \times 10^{-13} \text{ m}$, and a lifetime $\tau_X = c^2/\varepsilon_u = 10^{16} \text{ yrs}$, where c is the speed of light. The energy scale $E_X = (\varepsilon_u^5 \hbar^7 G^{-2})^{1/9} = 10^{-9} \text{ eV}$ strongly suggests a dark “radiation” field to provide a viable energy dissipation mechanism for energy conservation. If existing, the dark “radiation” should be produced by DM decay at early universe $t_X = (\varepsilon_u^{-5} \hbar^2 G^2)^{1/9} = 10^{-6} \text{ s}$ (quark epoch) with a mass of 10^{-9} eV such that axion can be a very promising candidate. If axion is the dark “radiation” responsible for the energy dissipation, it should have a mass around 10^{-9} eV with a GUT scale decay constant 10^{16} GeV and an effective axion-photon coupling constant $10^{-18} \text{ GeV}^{-1}$. The dark radiation energy density is around $\Omega_a h^2 \approx 2.6 \times 10^{-7}$, which is about 1% of the photon energy in CMB. Parameterized by the increase in the effective number of neutrino, $\Delta N_{eff} = 0.02$ can be obtained. This work suggests a heavy dark matter scenario and a dark radiation field from dark matter decay. More details can be found at arXiv:2202.07240.

Collaboration / Activity

See arxiv

Primary author: XU, Zhijie (Jay) (Pacific Northwest National Laboratory)

Presenter: XU, Zhijie (Jay) (Pacific Northwest National Laboratory)

Session Classification: Poster session

Track Classification: Dark Matter