

TB2020 Bremsstrahlung Study

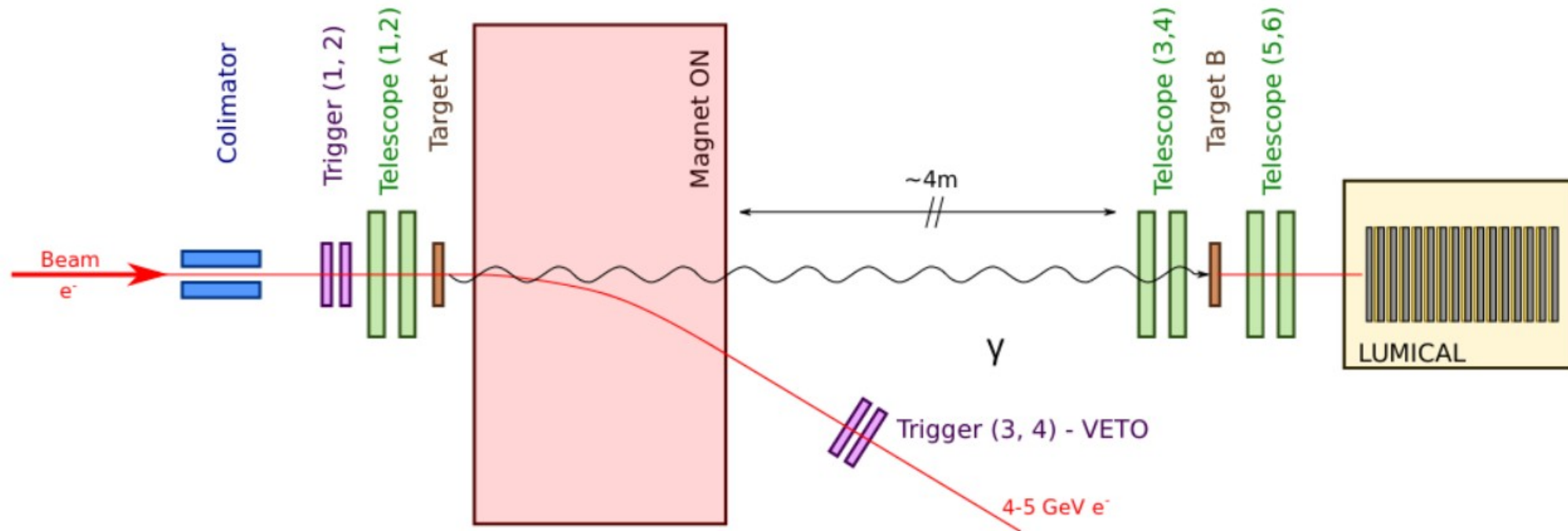
John Hallford

University College London

14/07/2022

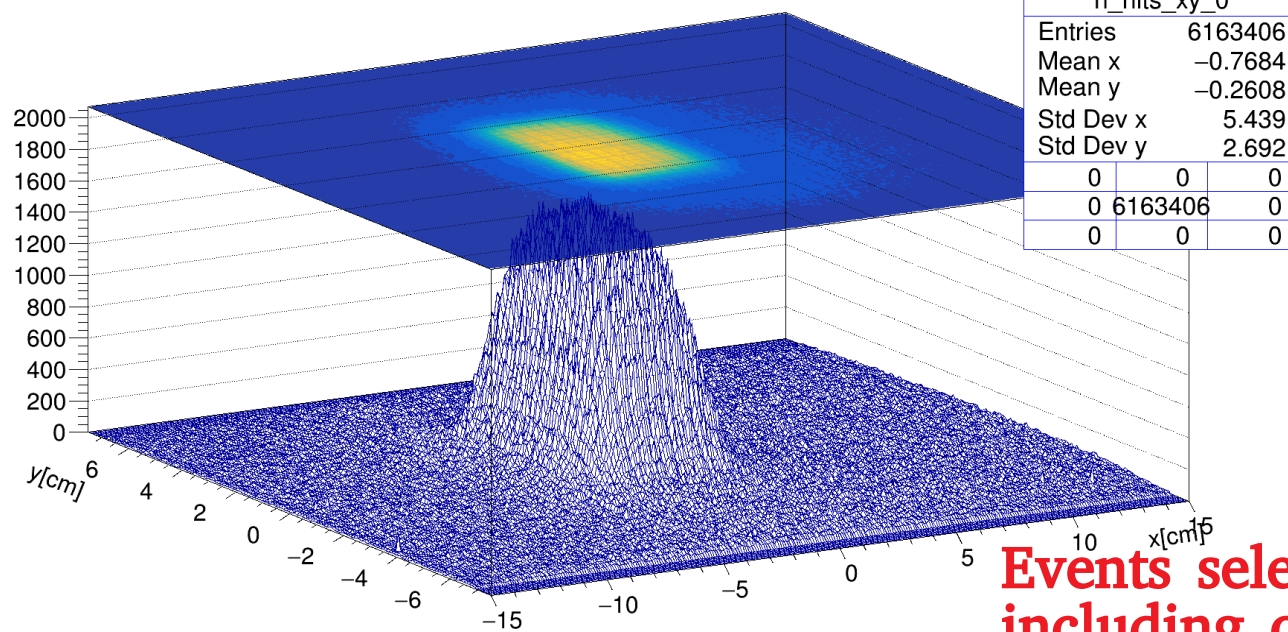
The logo for the LUXE experiment, featuring the word "LUXE" in a bold, blue, sans-serif font. The letter "X" is stylized with a white starburst or spark-like graphic in the center.

TB2020 Alpide Telescope Alignment

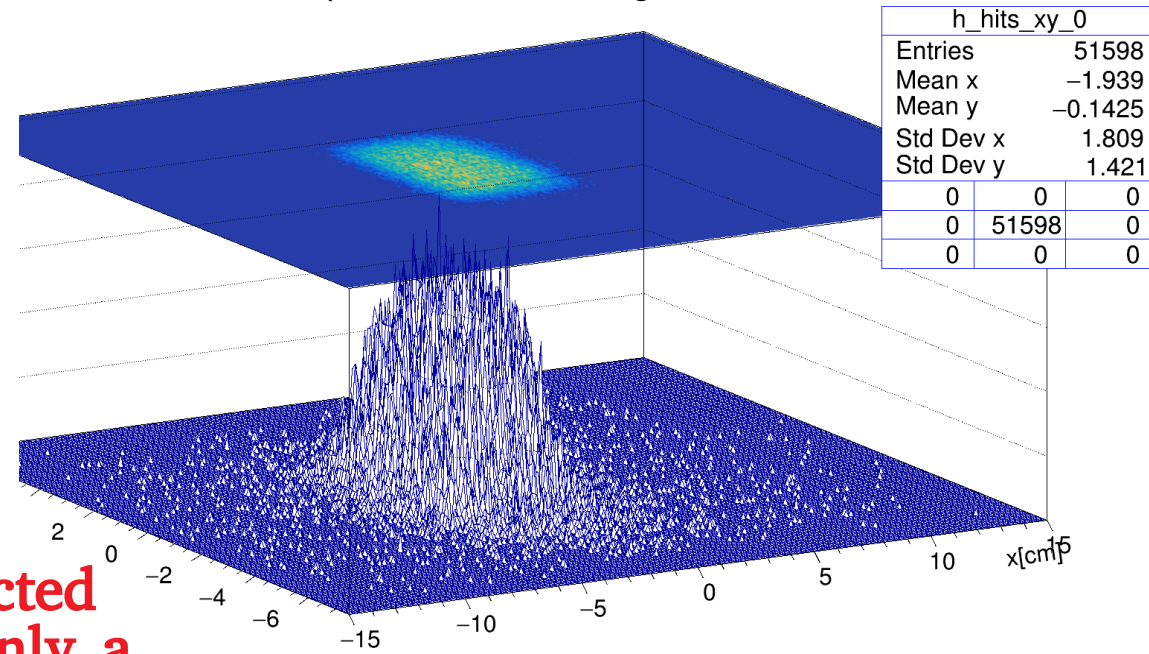


Idea is to use this data to reconstruct the dispersion of bremsstrahlung photons, in particular the polar angle.

Telescope Plane 0 - Pre-target1 - z=0

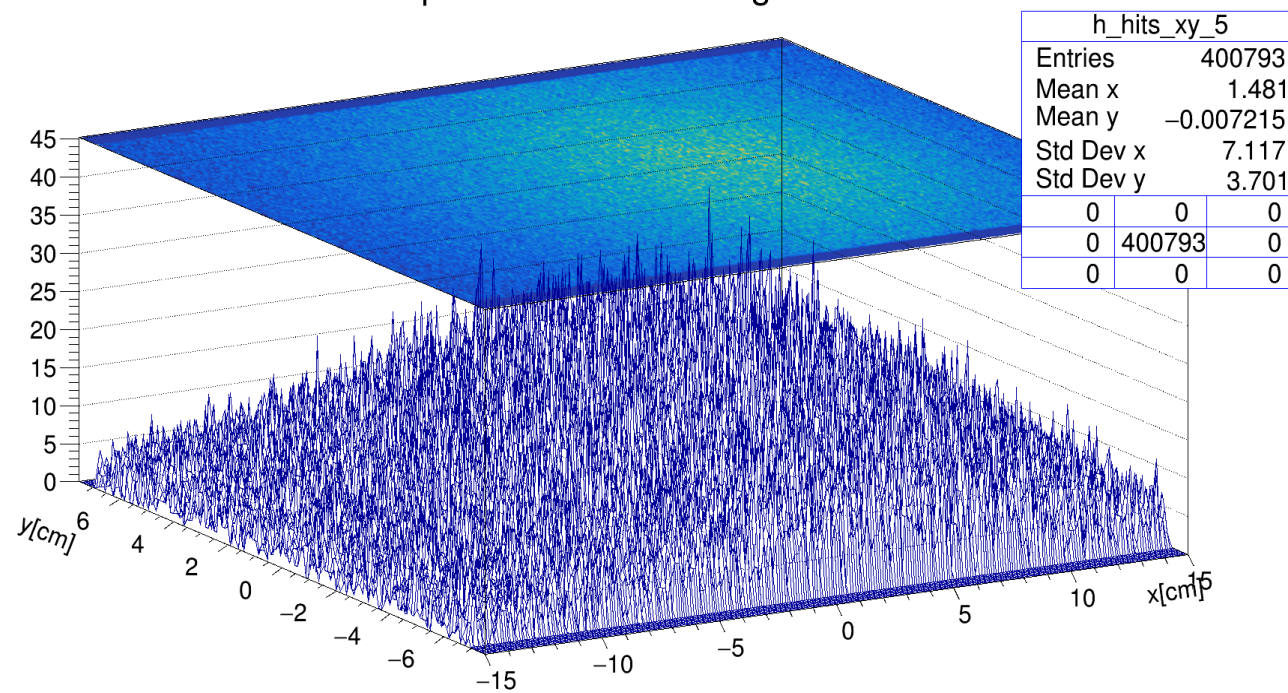


Telescope Plane 0 - Pre-target1 - z=0

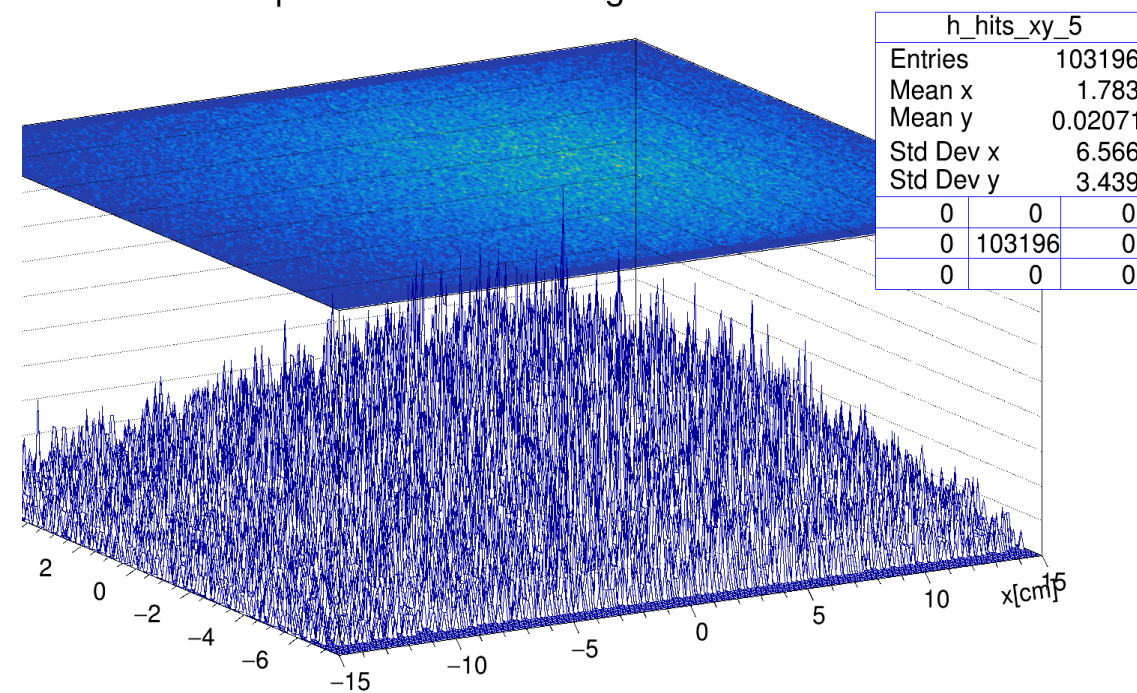


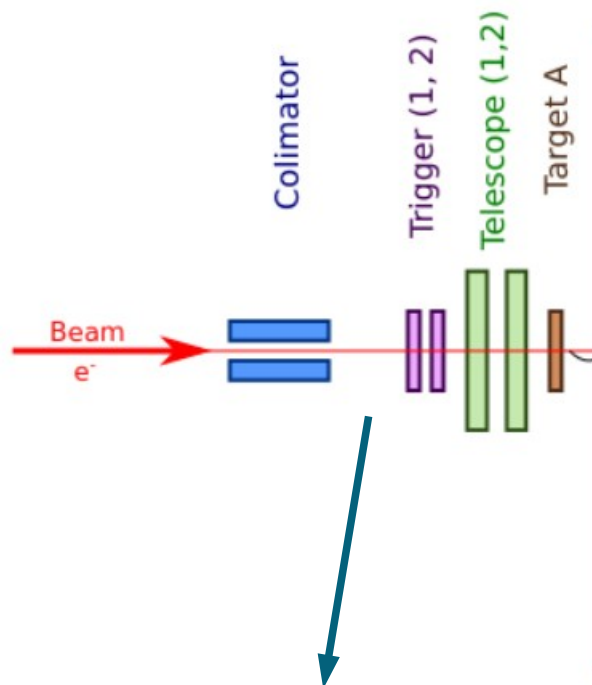
Events selected
including only a
1:1:0:0:2:2 hit pattern

Telescope Plane 5 - Post-target2 - z=7988



Telescope Plane 5 - Post-target2 - z=7988

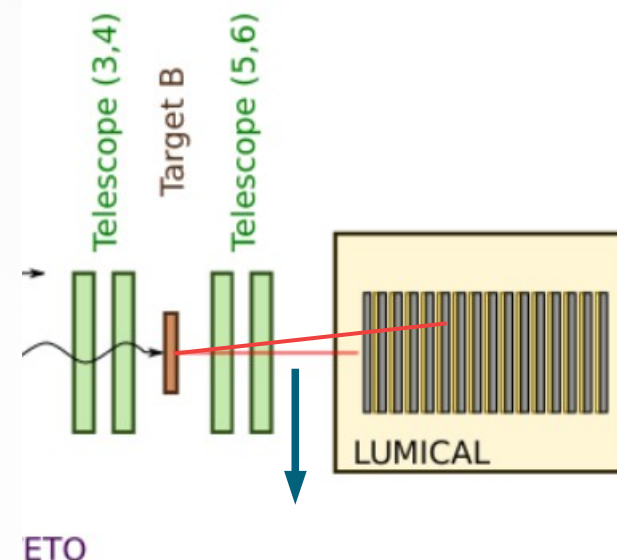




Trace trajectory from hitpoints in planes 1,2 to reconstruct incidence of target 1

Alignment is important!
and this is found by
minimisation of straight-
line fits chi-squared sum
for data with targets
removed, magnet off

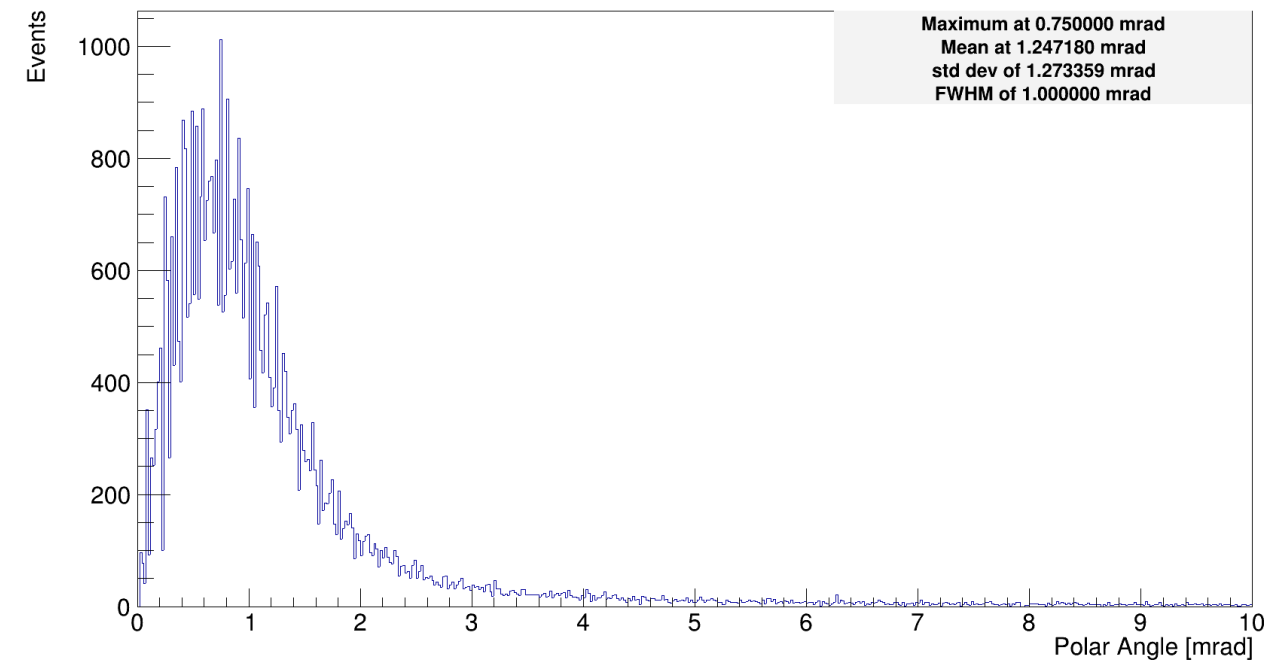
Reconstruct
Bremsstrahlung gamma
momentum vector,
compare to incident e-
vector



Reconstruct gamma
incidence on target 2
using two hits of each
of e-/e+ pair

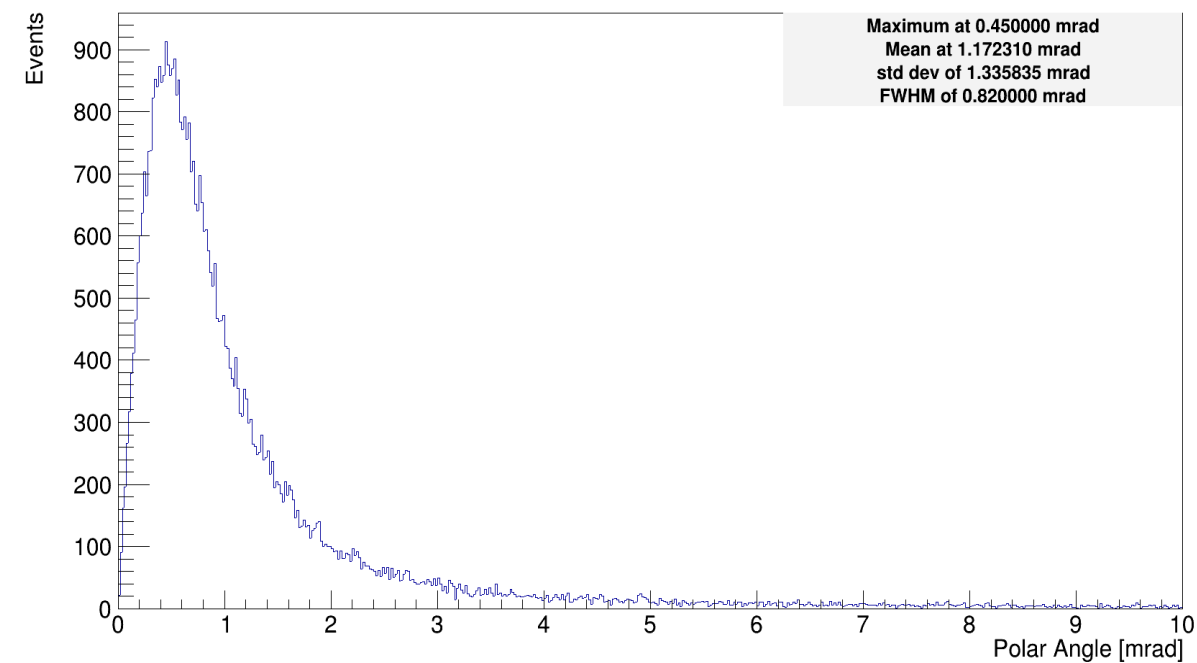
Track + correlate hits to correct
tracks by extrapolating to target
2, choosing minimal distance
between tracks

Initial Electron Polar Angle

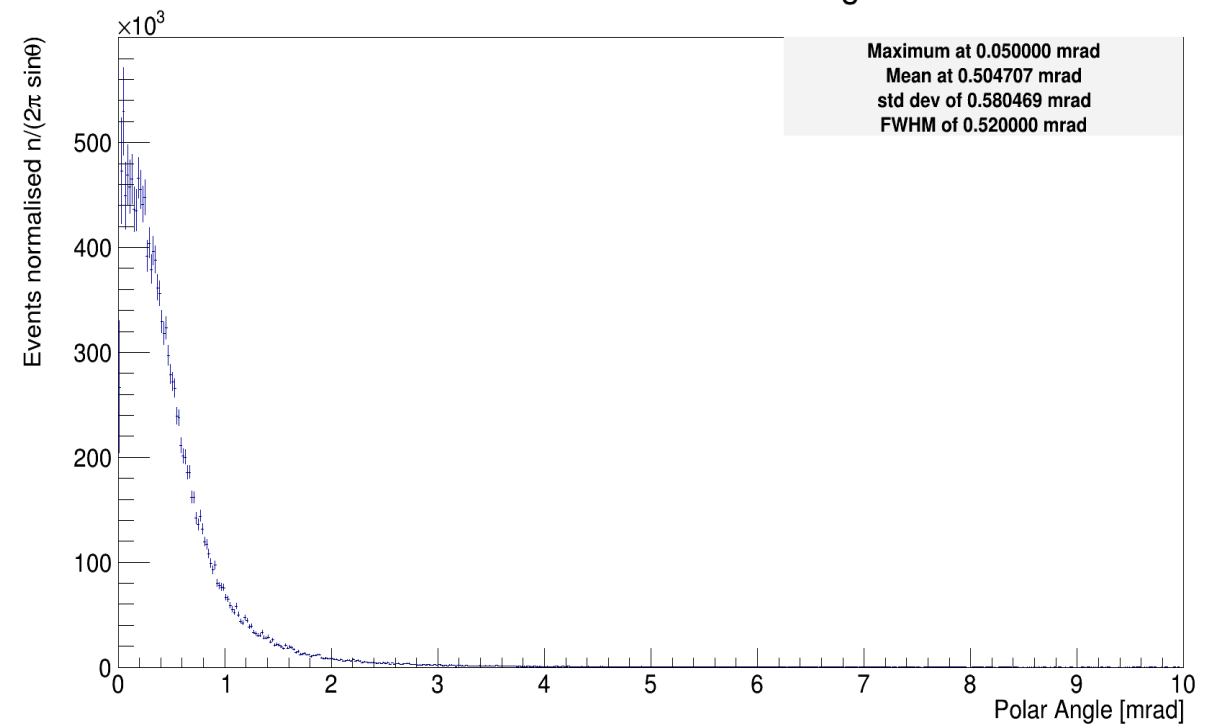


So we find a Brem emission profile of theta with peak 0.45 mrad, FWHM 0.82 mrad

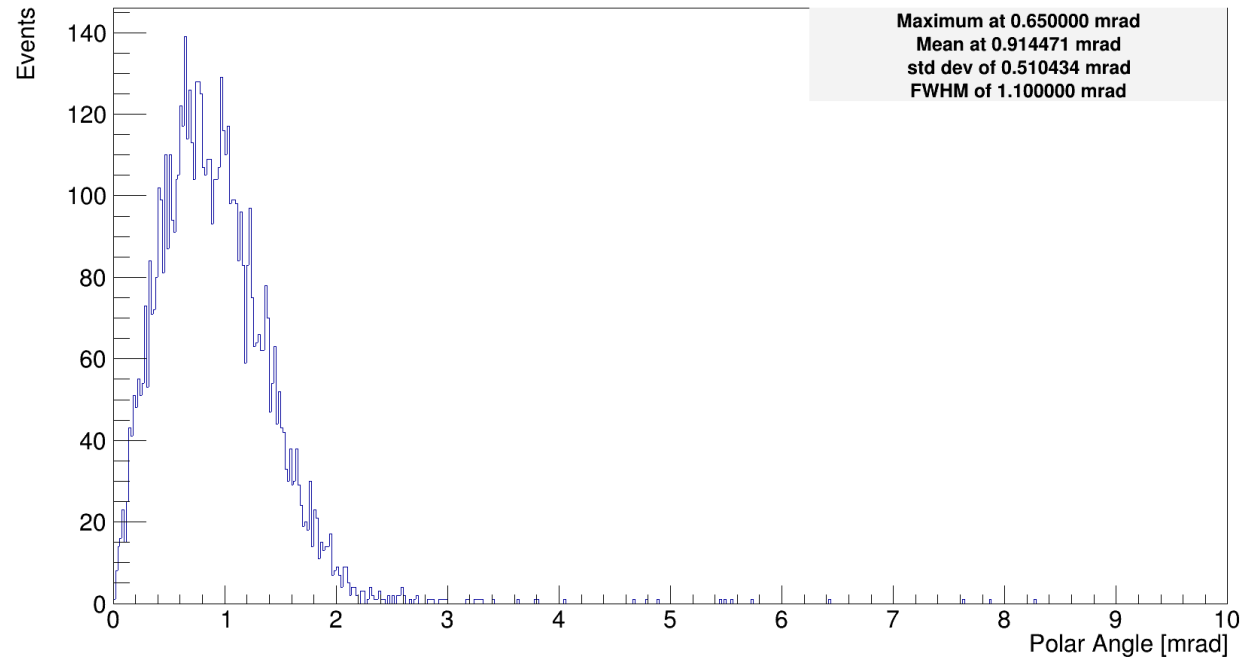
Intermediate Brem Gamma Polar Angle



Intermediate Brem Gamma Polar Angle Fluence



Monte Carlo Initial Electron Polar Angle

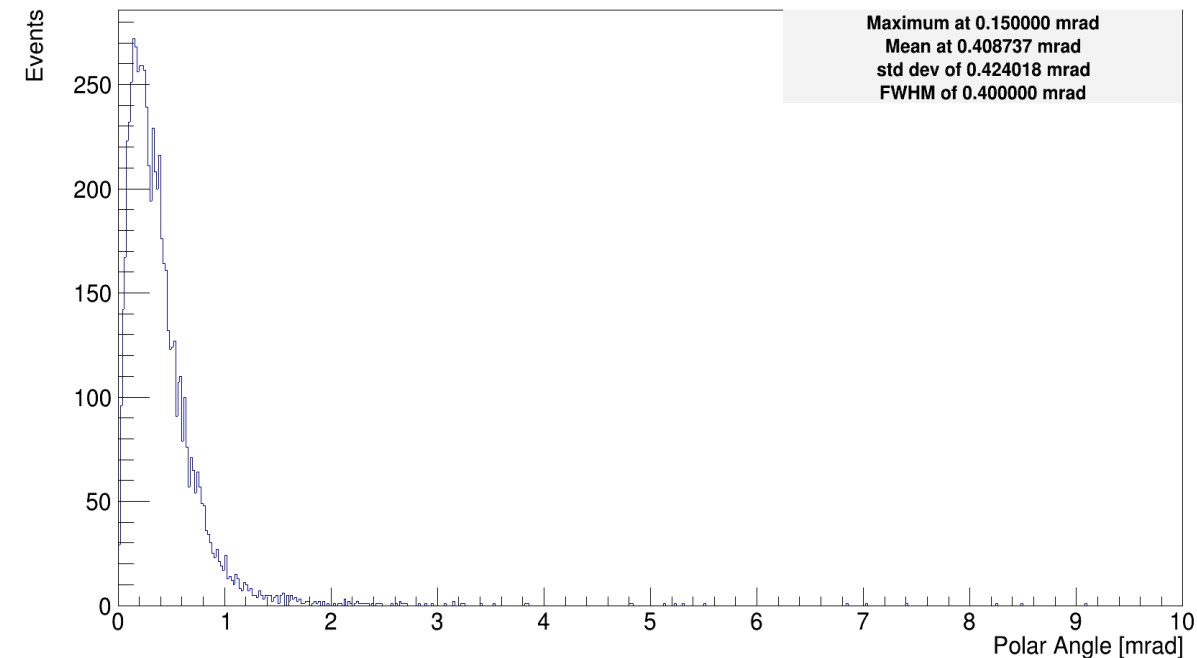


But we also have a more complete Monte Carlo! Including measurement of the plane hits (e-,e+ with $E > 2\text{MeV}$), a complete reconstruction using the same analysis technique is performed

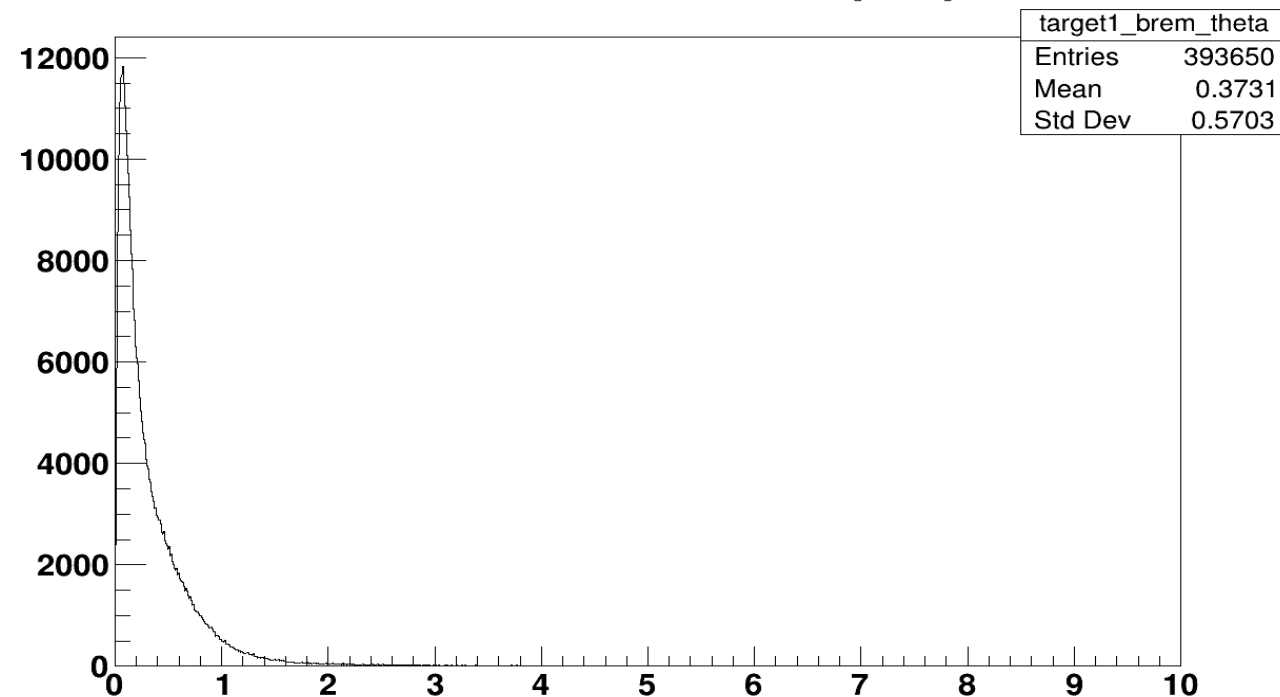
Specifically telescope resolution ($\sim 2.88\mu\text{m}$) and multiple scattering in the environment are modelled

Reconstructed theta distribution is half the size of the real data

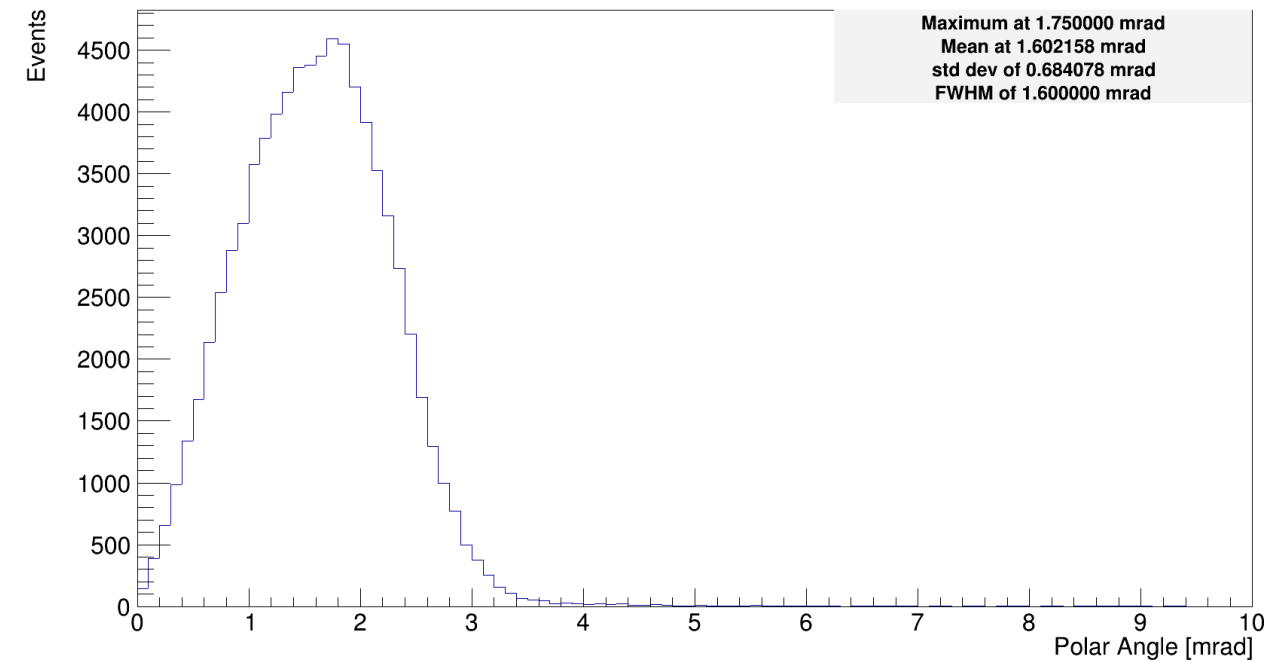
Monte Carlo Intermediate Brem Gamma Polar Angle



Monte Carlo Inst. Brem Photon Theta; θ [mrad]; Events



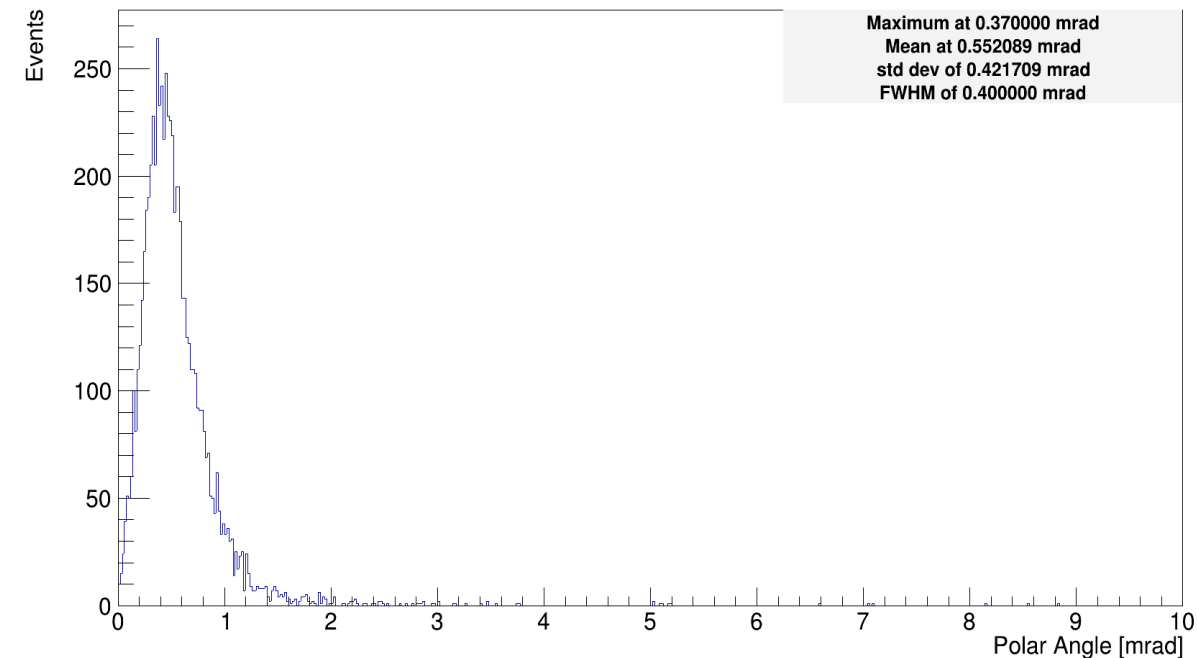
Monte Carlo Initial Electron Polar Angle



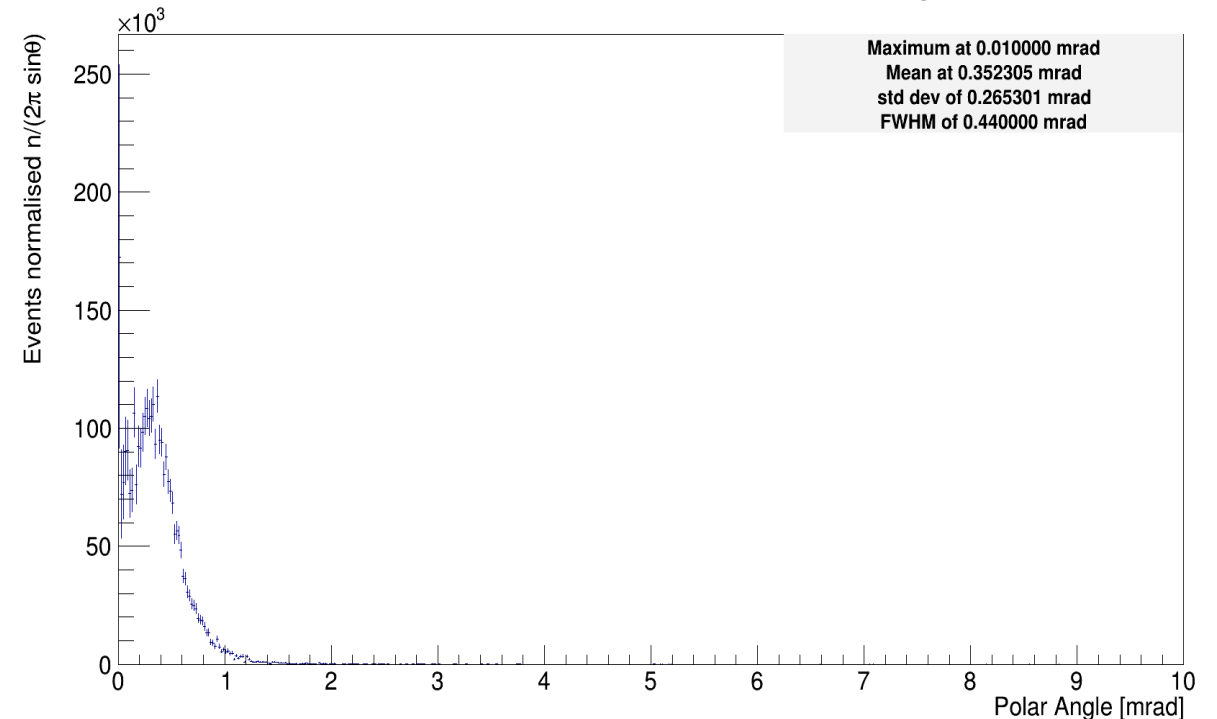
With a Monte Carlo at our disposal we can try an intentional misalignment...

Misaligning each plane by some random value in x,y between $\pm 0.03\text{mm}$...

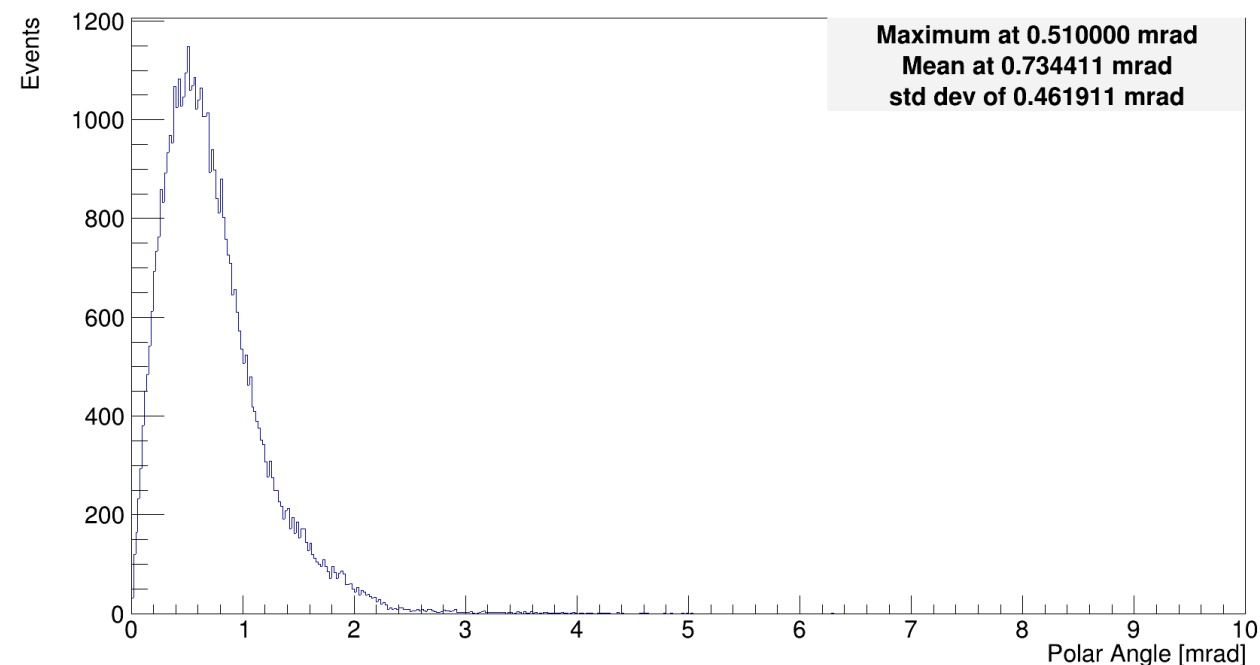
Monte Carlo Intermediate Brem Gamma Polar Angle



Monte Carlo Intermediate Brem Gamma Polar Angle Fluence

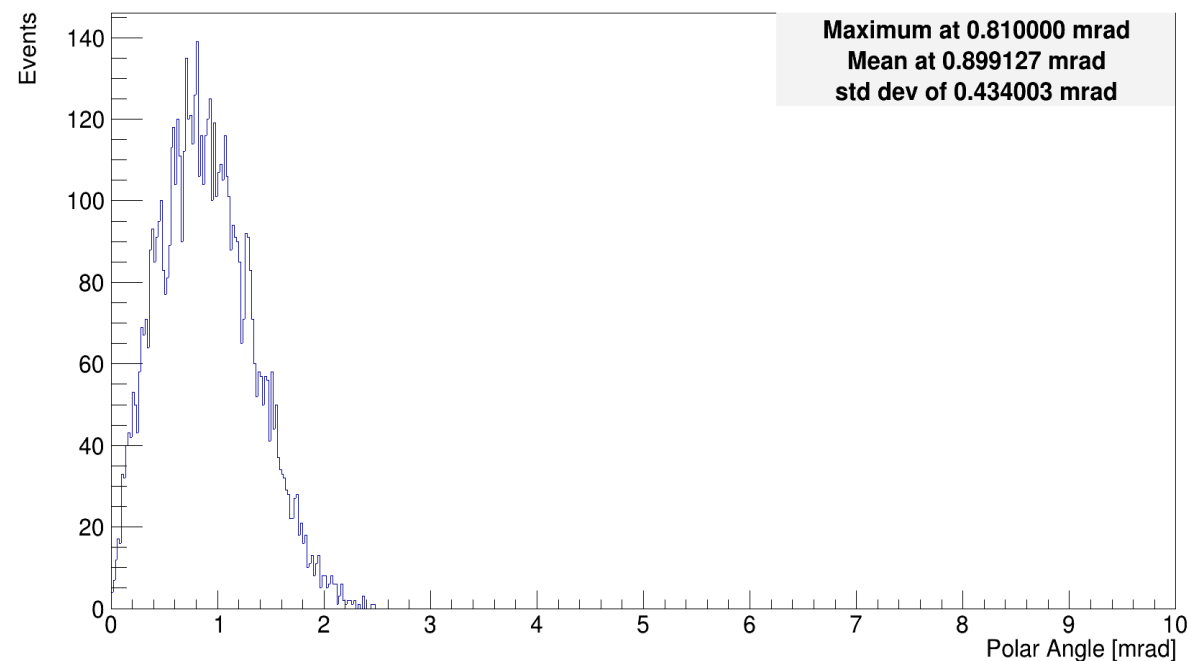


Intermediate Brem Gamma Polar Angle wrt z axis

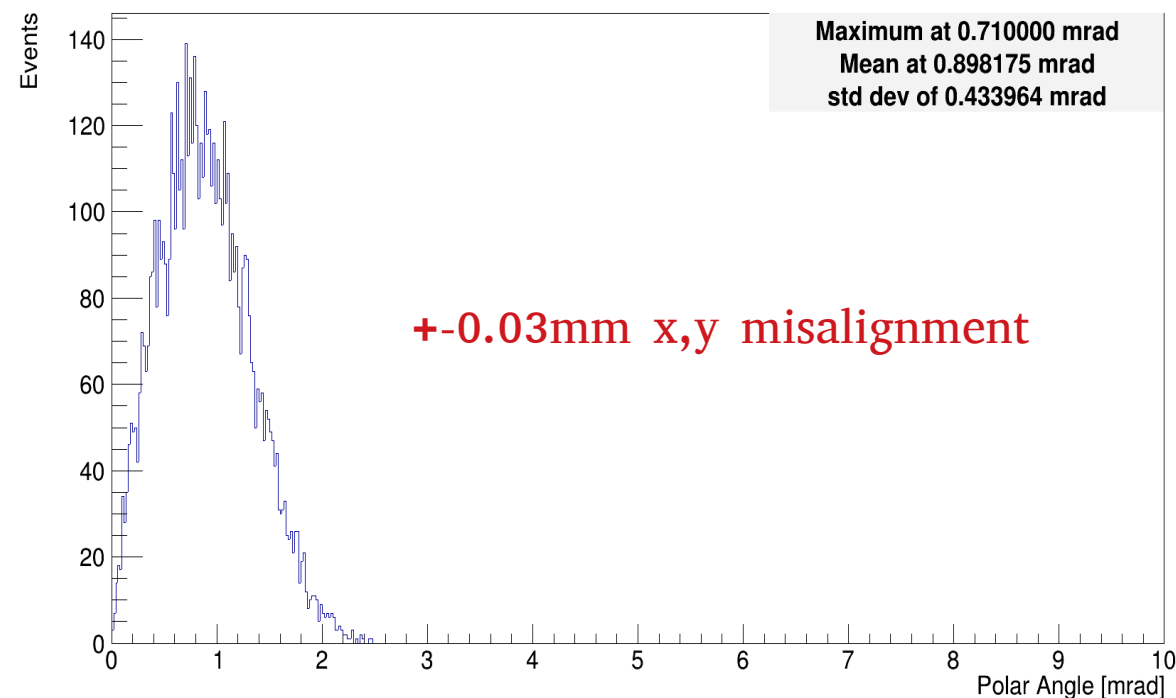


We can use the polar angle of the Brem. Photons w.r.t. a Z axis to get a measurement dependent less on misalignment, but more on initial electron beam distribution

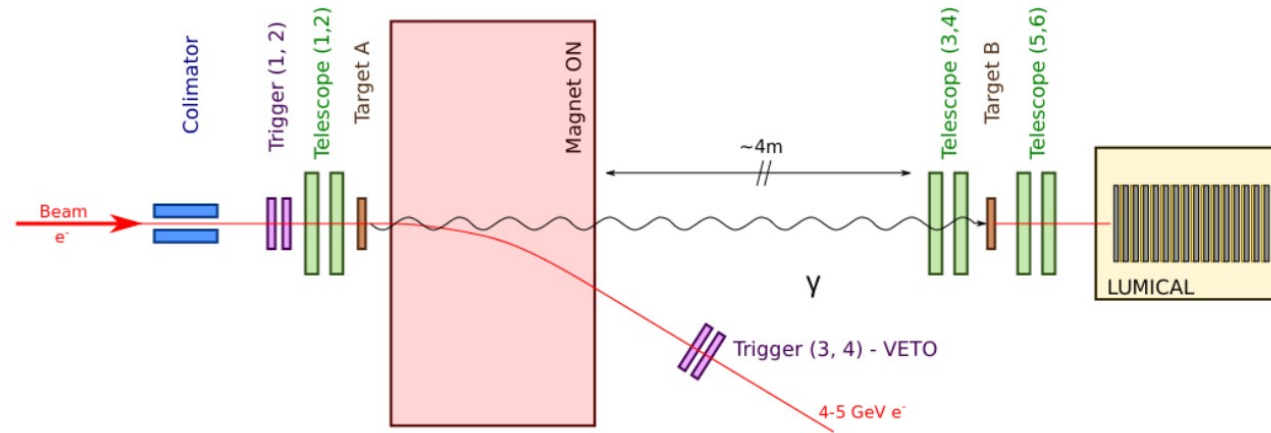
Monte Carlo Intermediate Brem Gamma Polar Angle wrt z axis



Monte Carlo Intermediate Brem Gamma Polar Angle wrt z axis

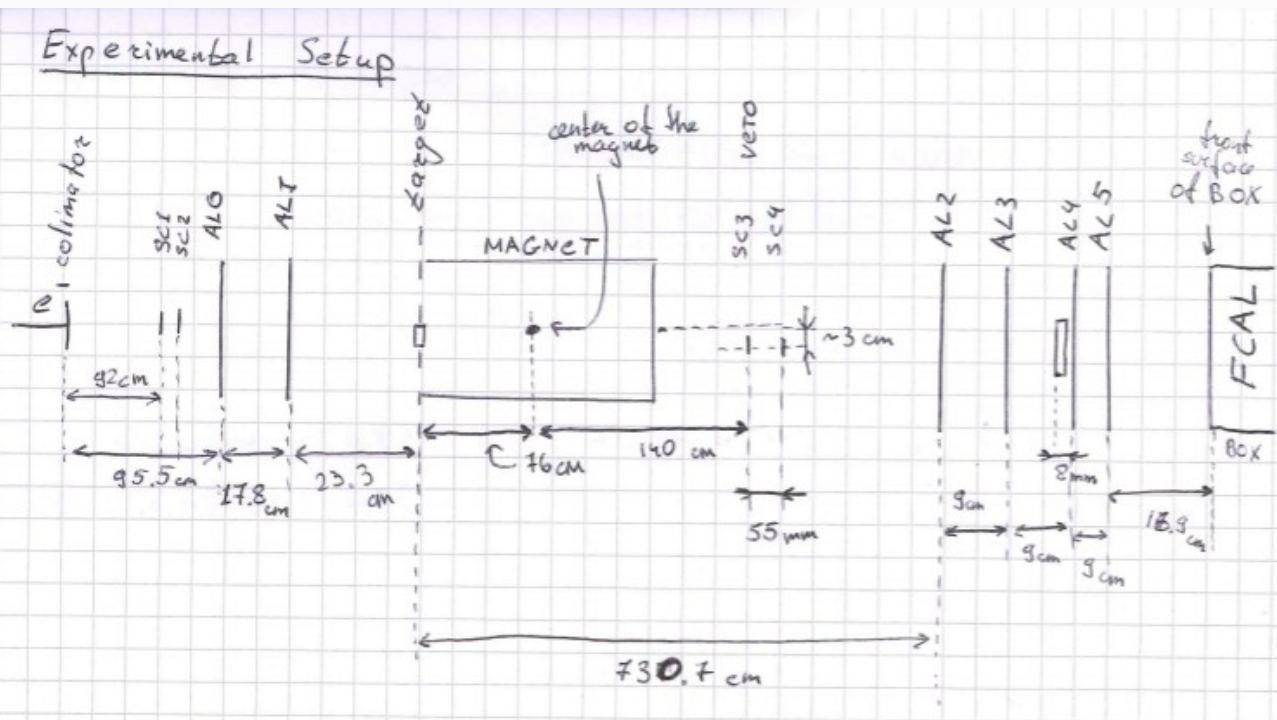


TB2020 Alpile Telescope Alignment



- Continue realignment, especially with 1st two planes
- Try 4-5 GeV e⁻ veto
- Calorimetry analysis

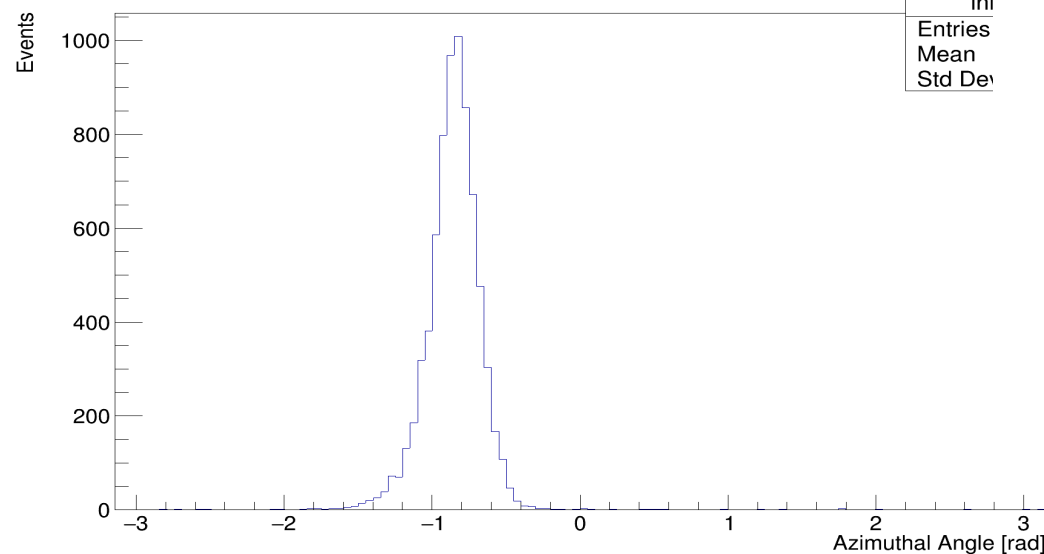
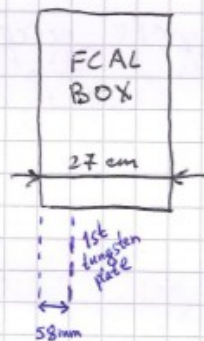
backup



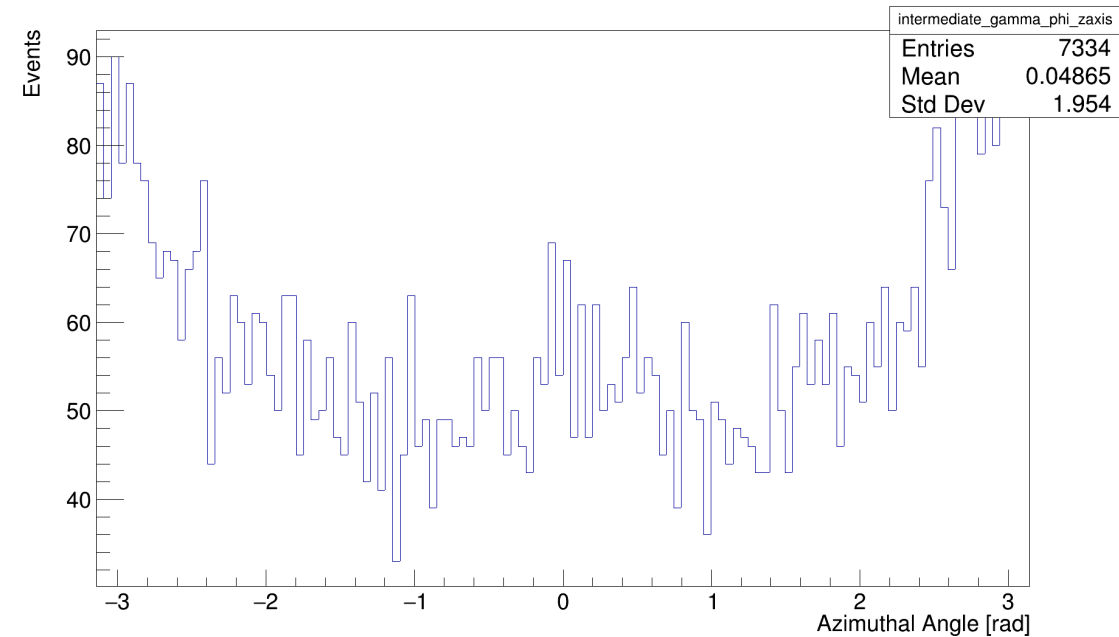
Misaligning each plane by some random value in x,y between $\pm 0.25\text{mm}$...

Initial electron vector seems particularly affected; likely because of the far smaller displacement between planes 1 & 2

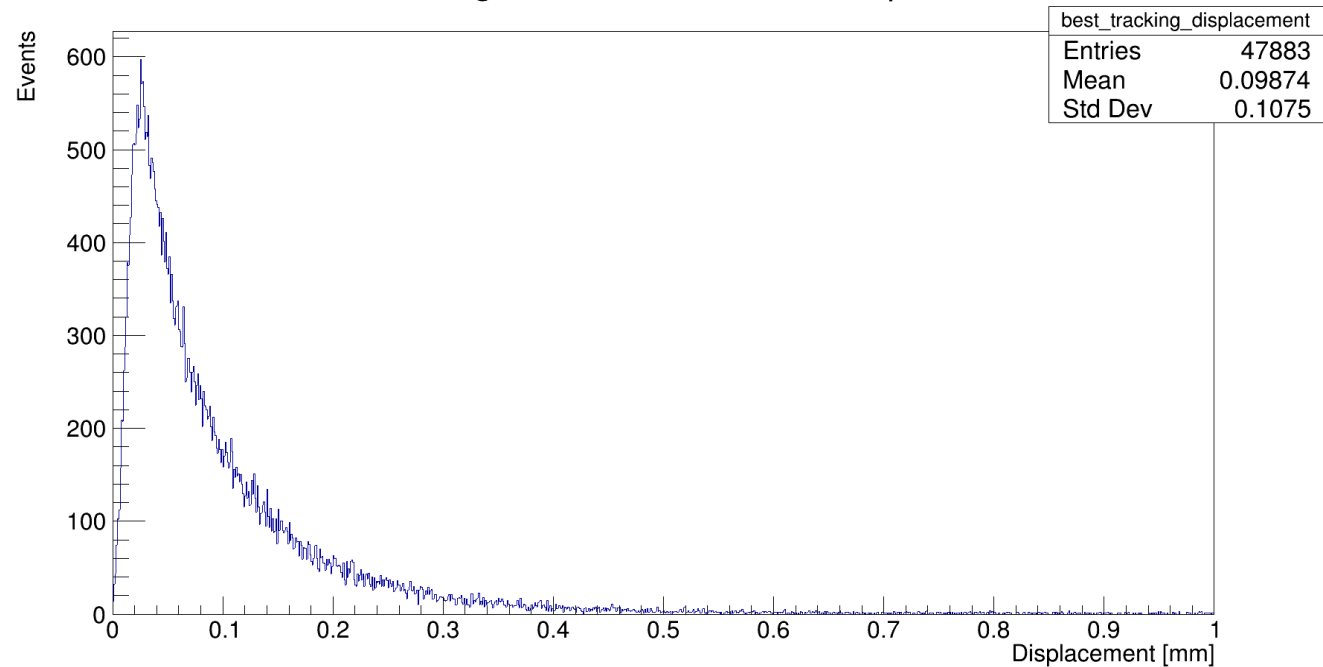
Monte Carlo Initial Electron Azimuthal Angle



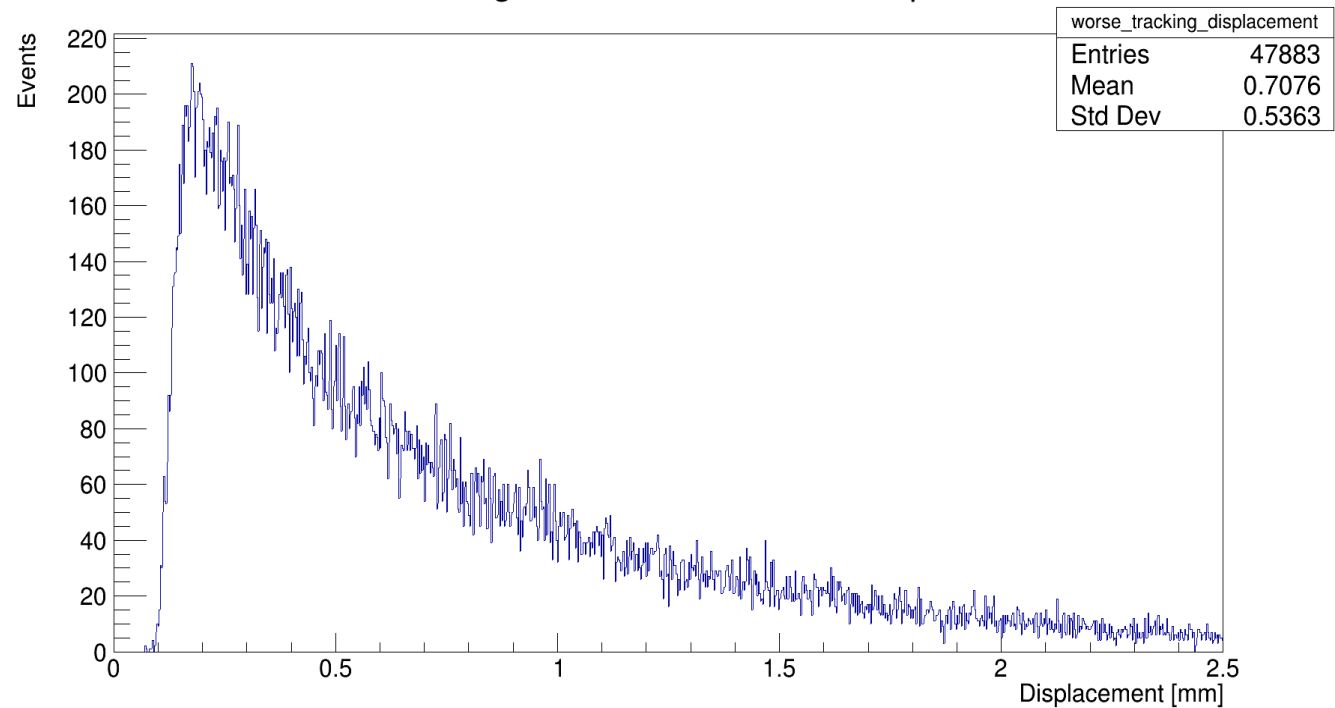
Monte Carlo Intermediate Brem Gamma Azimuthal Angle wrt z axis



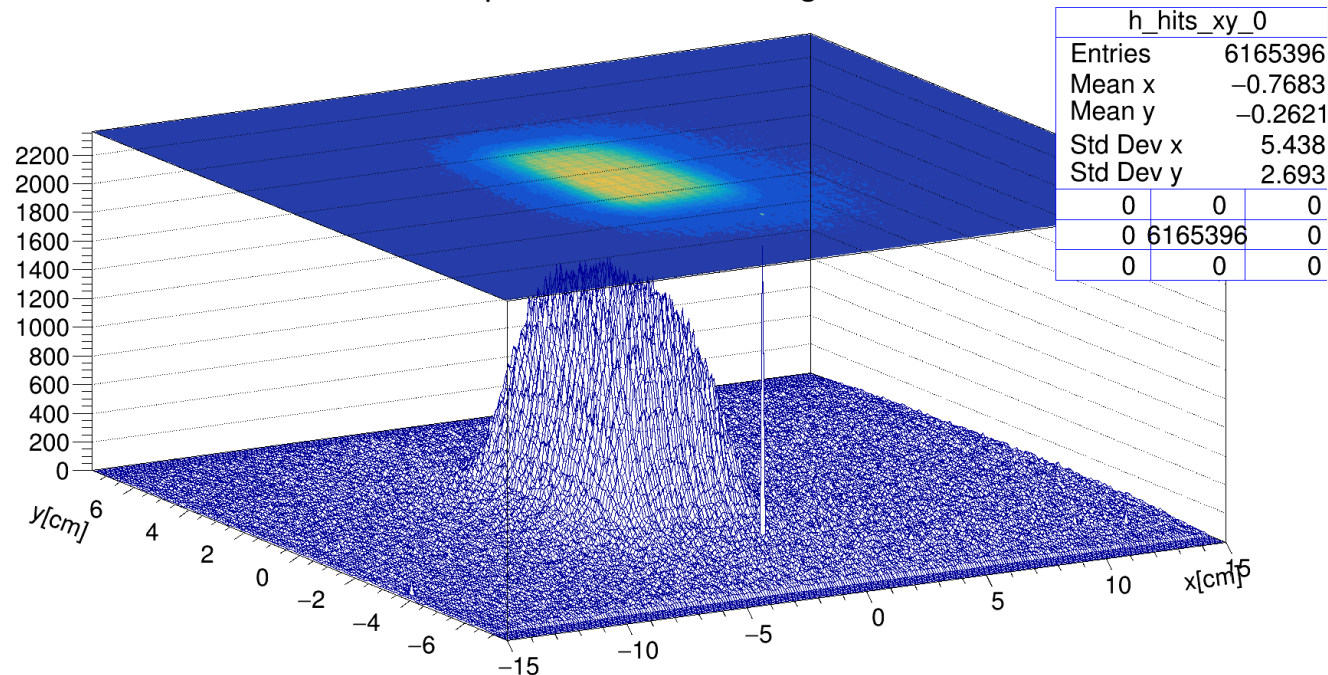
Best Tracking Solution e- e+ vertex displacement



Worse Tracking Solution e- e+ vertex displacement

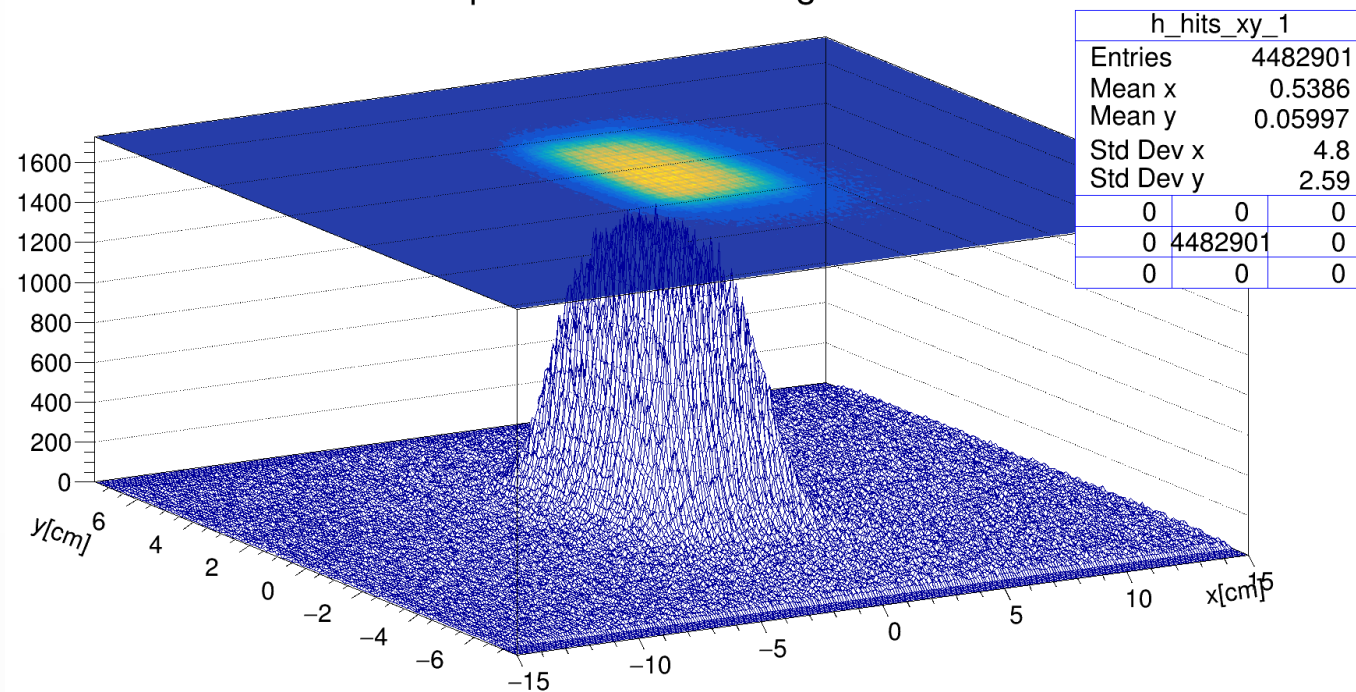


Telescope Plane 0 - Pre-target1 - z=0



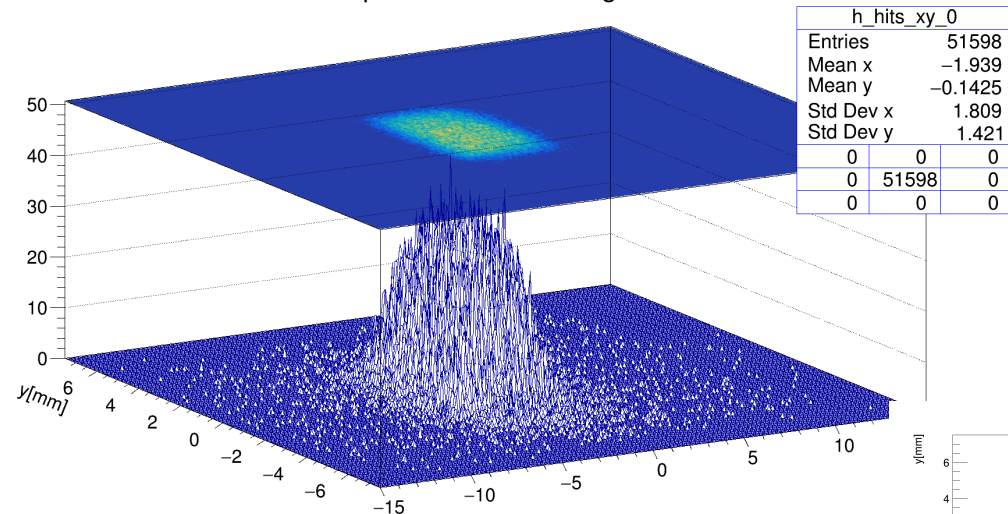
Have cleaned data of hot pixels

Telescope Plane 1 - Pre-target1 - z=178

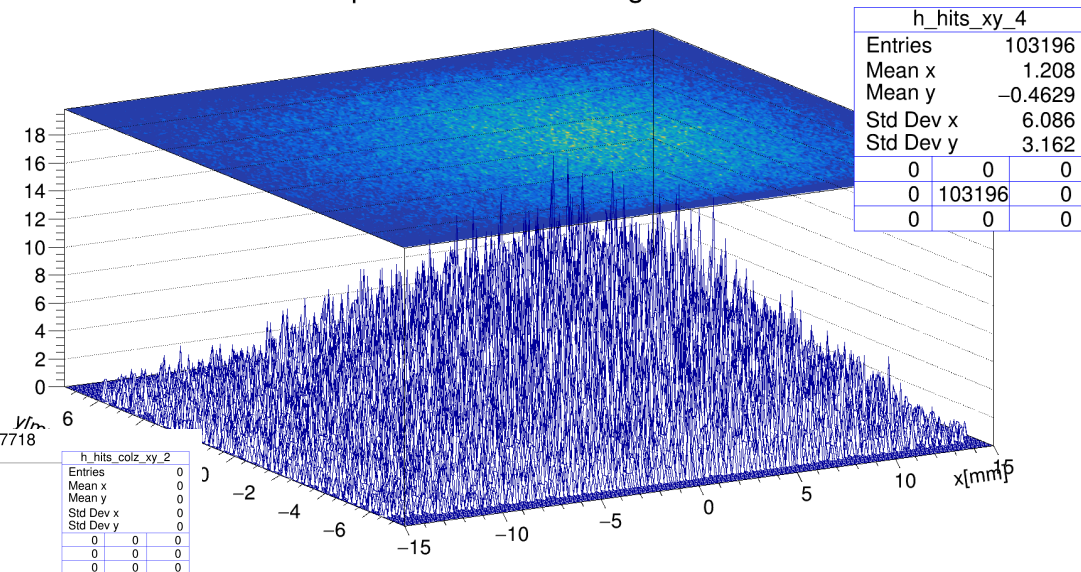


Resulting analysis could exclude real data within these 'hot' pixels but this result is likely small

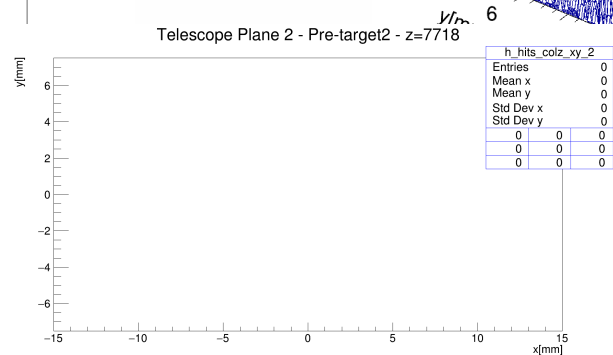
Telescope Plane 0 - Pre-target1 - z=0



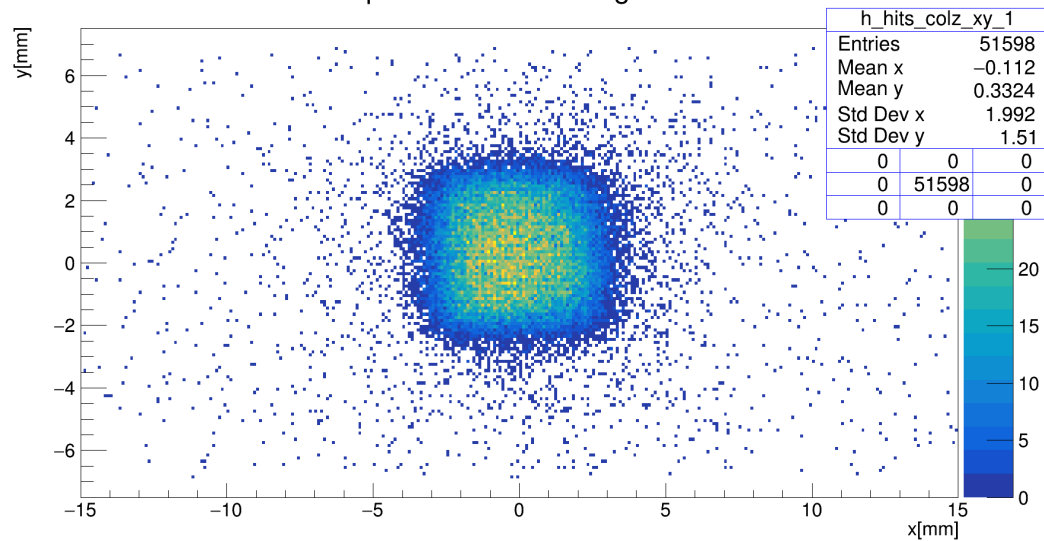
Telescope Plane 4 - Post-target2 - z=7898



Telescope Plane 2 - Pre-target2 - z=7718

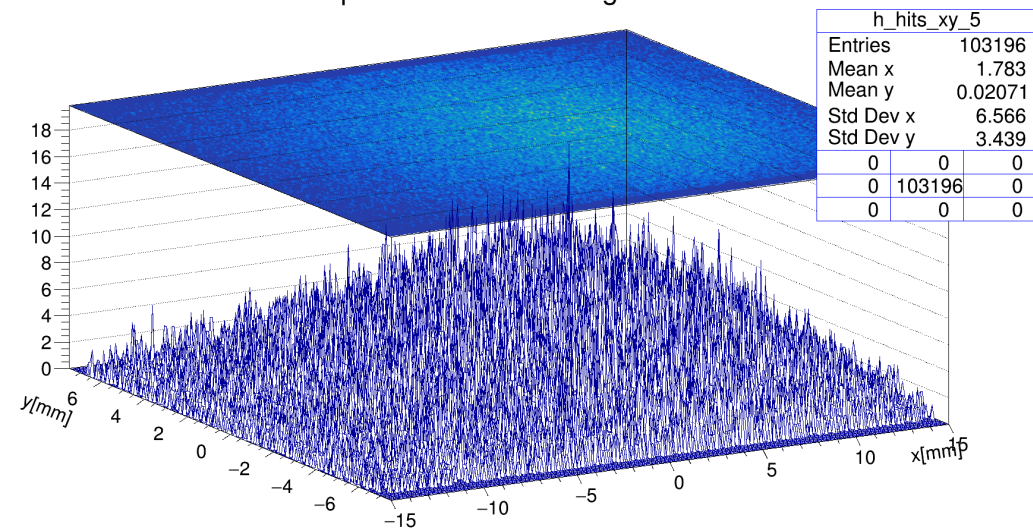


Telescope Plane 1 - Pre-target1 - z=178

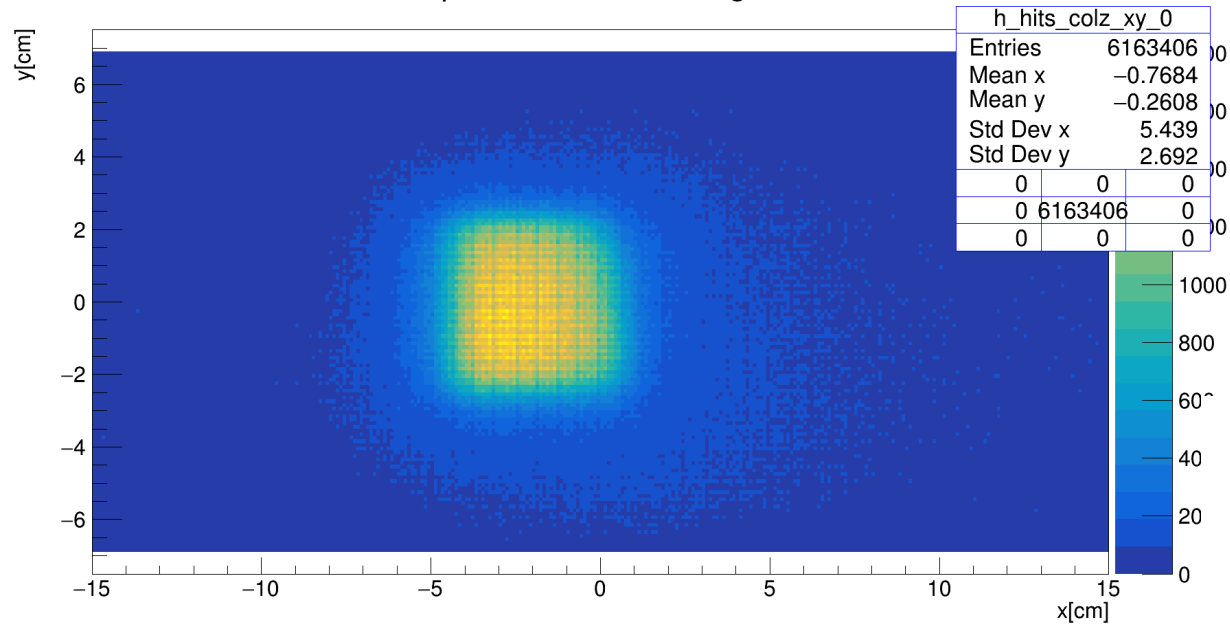


Look for
110022 hit
pattern

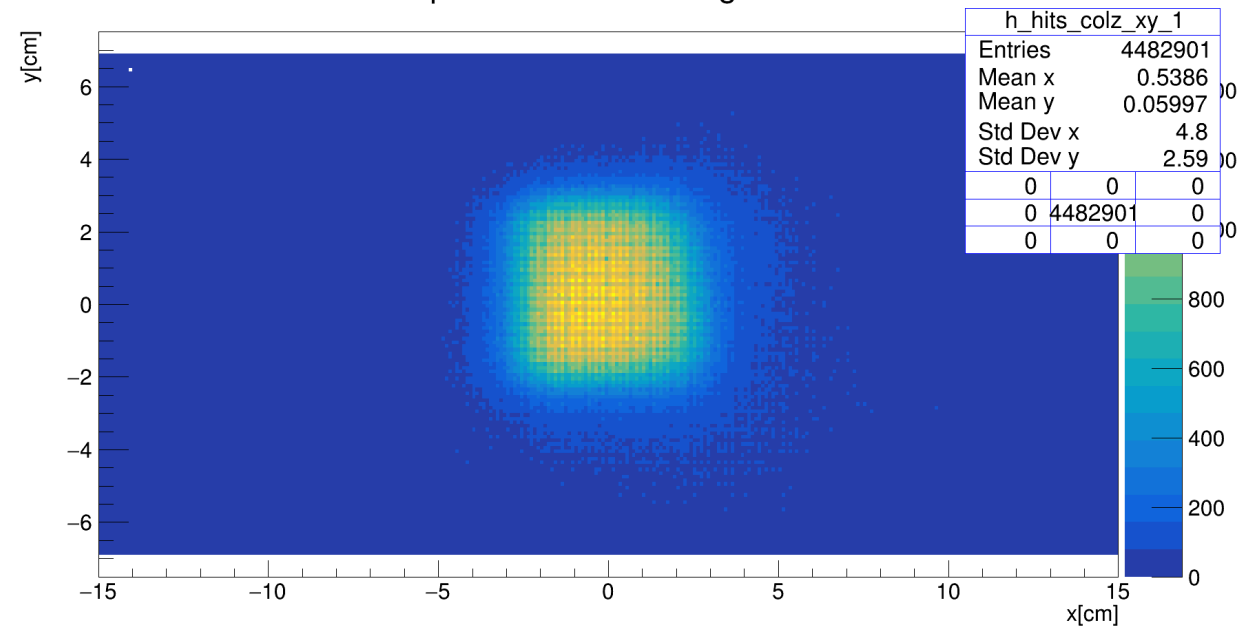
Telescope Plane 5 - Post-target2 - z=7988



Telescope Plane 0 - Pre-target1 - z=0

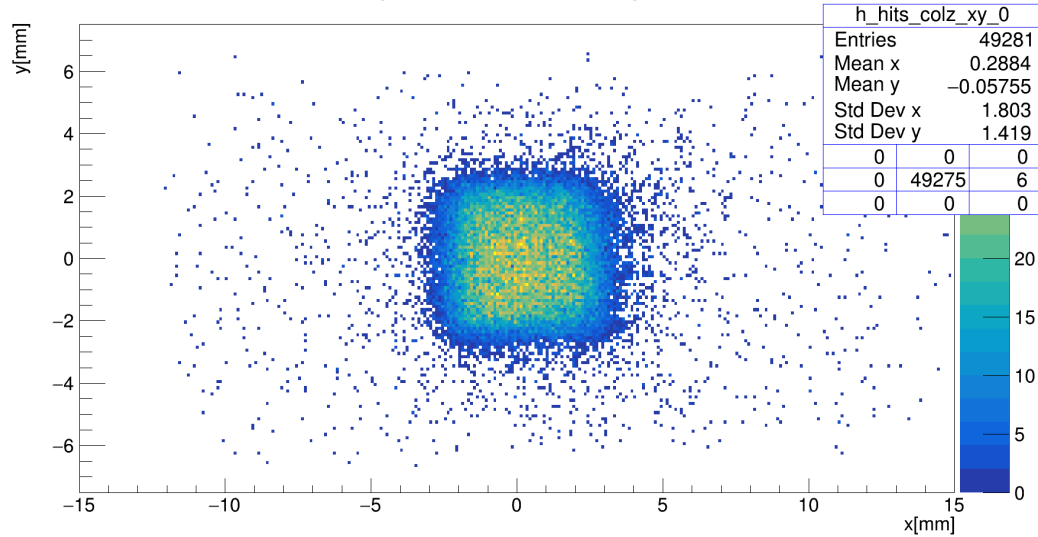


Telescope Plane 1 - Pre-target1 - z=178

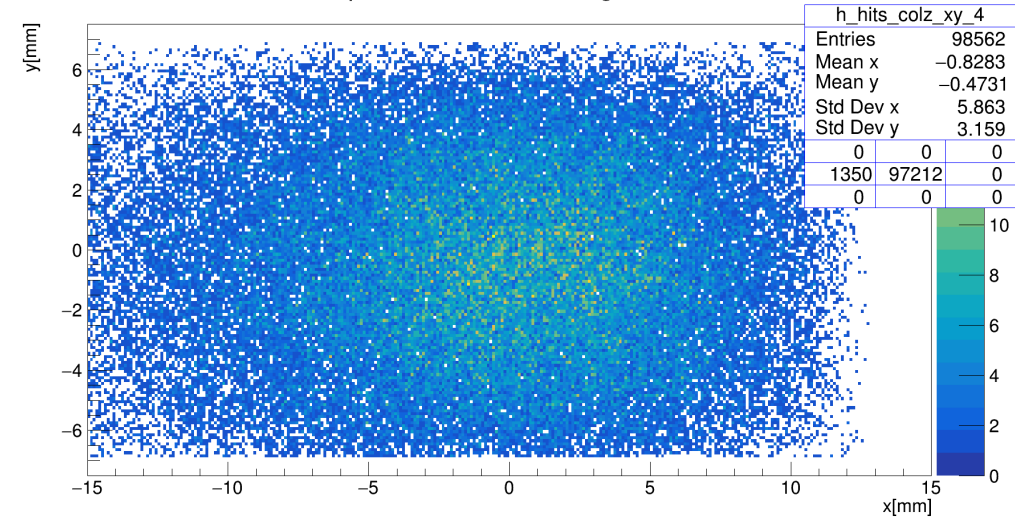


This needs above all,
good alignment!

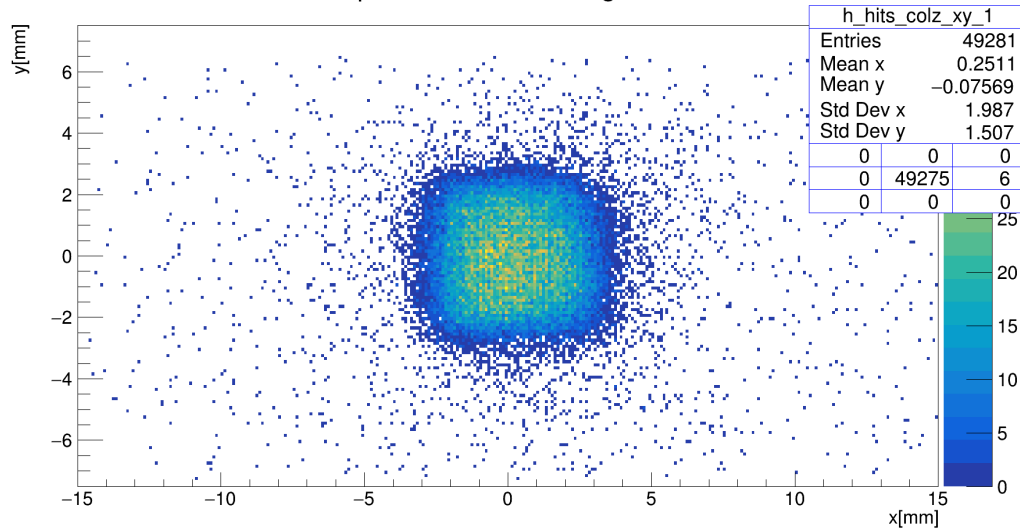
Telescope Plane 0 - Pre-target1 - z=0



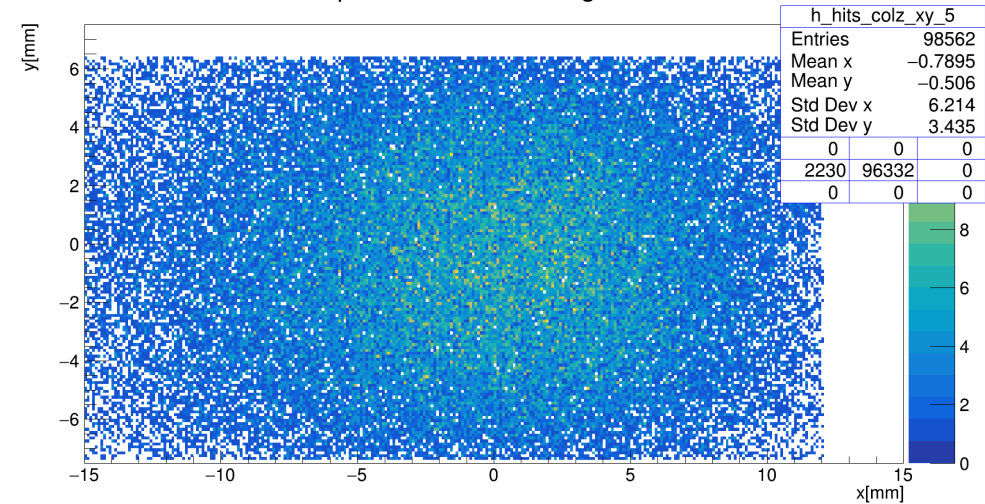
Telescope Plane 4 - Post-target2 - z=7898



Telescope Plane 1 - Pre-target1 - z=178

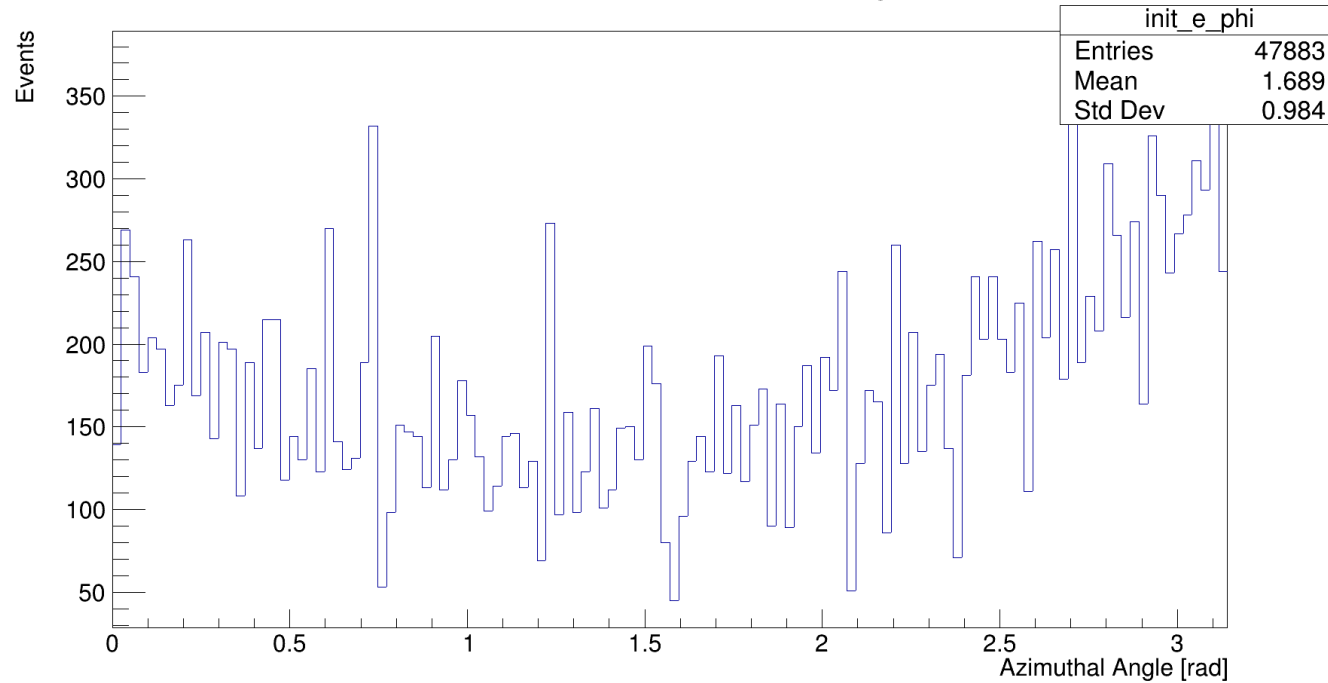


Telescope Plane 5 - Post-target2 - z=7988



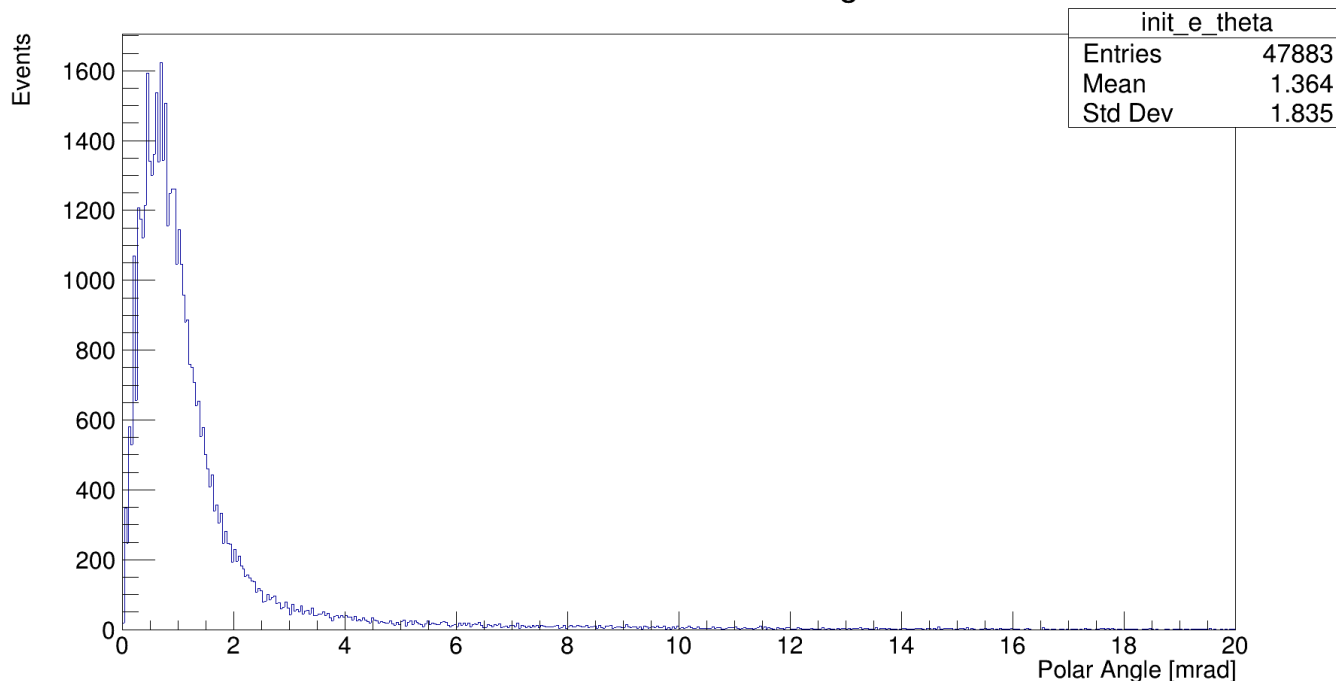
There are runs with no targets, no B-field, with telescope planes in place; use mean position of distribution for alignment

Initial Electron Azimuthal Angle



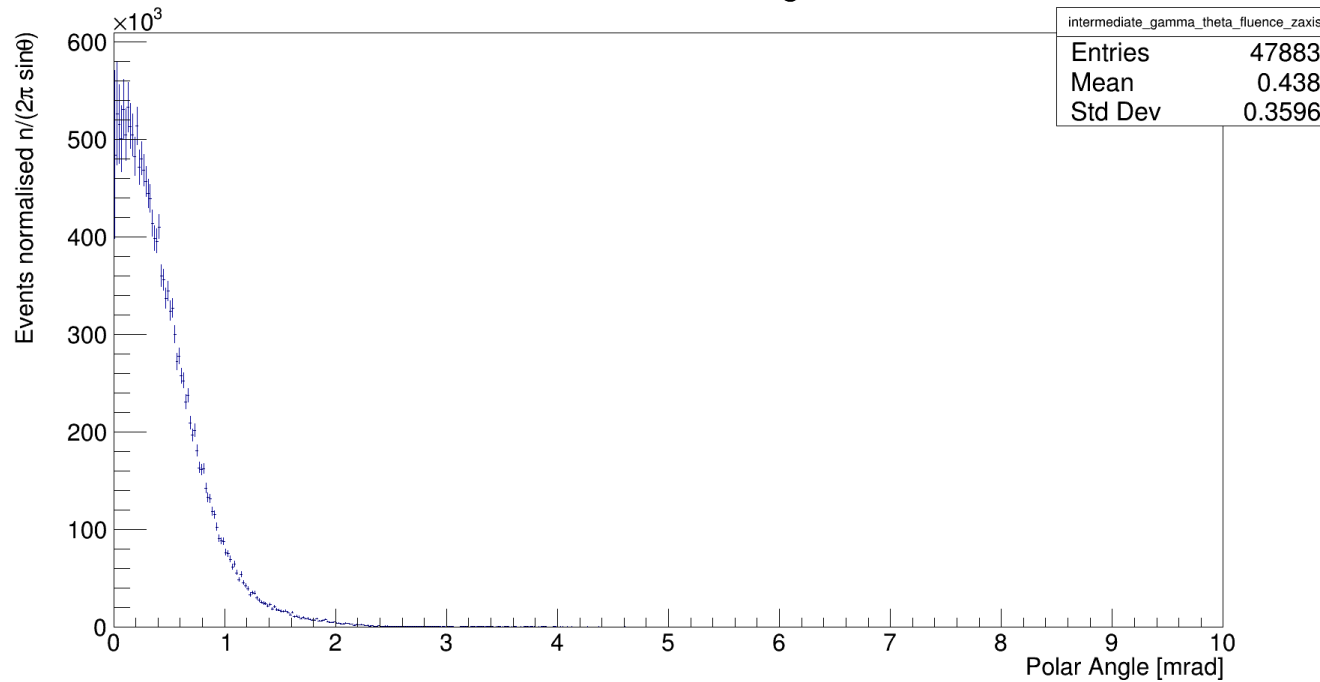
← Unclear with statistics, but Azimuthal distribution does not look flat. Can be explained by pixelated telescope structure, and selection of tracks where all are incident within the rectangular detector plane

Initial Electron Polar Angle



← Polar angle distribution of initial electron w.r.t. center of aligned detector planes. The DESY-II test beam itself is actually created using Brem-radiation then pair-production.

Intermediate Brem Gamma Polar Angle Fluence wrt z axis



← Reconstructed photon track polar angle vs. (somewhat arbitrary) z axis... results in steeper trend?

Alignment must be bettered... previously just used mean position.

Have thought about a more involved event-by-event alignment strategy

Intermediate Brem Gamma Polar Angle Fluence

