## **GBP Main Meeting** Ptarmigan LMA Simulations

K Fleck - 21/09/2022

#### **Ptarmigan LMA simulations Simulation parameters**

- Laser
  - $\lambda = 0.8 \,\mu\text{m}$
  - $\tau_{FWHM} = 30 \, \mathrm{fs}$
  - $E_I = 1.2 \, \text{J}$
  - Linear polarisation

- Electrons
  - $1.5 \times 10^9 e^{-1}$
  - 16.5 GeV
  - $\Theta_{rms} = 8.672 \,\mu rad$
  - $r_h = 5.0 \,\mu m$
- Radiation reaction = on, pair production = off

#### **Ptarmigan LMA simulations** Simulation processing

Nominal ξ	Number of entries processed (1e8)	Number of entries in 10 BX (1e8)	Number of BXs processed
0.5	9.22	9.22	10.0
1.0	28.91	28.91	10.0
2.0	15.35	62.23	2.47
3.0	14.79	75.99	1.95
5.0	5.87	76.01	0.77
7.0	14.05	68.13	2.06
10.0	15.35	58.31	2.63

#### **Ptarmigan LMA simulations** Laser intensity at creation



 $N_{\gamma}$  / BX

#### **Ptarmigan LMA simulations** Laser intensity at creation



#### Functional form of fit taken from <u>Blackburn et. al. 2020</u>

## **Ptarmigan LMA simulations** Laser intensity at creation with energy



#### **Ptarmigan LMA simulations Compton harmonics for IPWs**



#### General case $\nu k \cdot p$ $\omega' =$ $(p + \nu k) \cdot n'$

$$\nu = \frac{k' \cdot p}{k \cdot p'} = \frac{k^{'-} + p^{'-} - k^{'-}}{k^{-}}$$

Infinite plane wave case

$$\nu \to \nu_n = n - \frac{\xi^2}{4\eta} \frac{s}{1 - s}$$
$$\omega'(n) = \frac{n\omega e^{2\zeta}}{1 + 2n\frac{\omega}{m}e^{\zeta} + \xi}$$

 $\zeta = \operatorname{arccosh} \gamma$ 





### **Ptarmigan LMA simulations** Laser intensity at creation with energy with harmonics



#### Same plots as slide 6 but with first, second and third IPW harmonics

overlayed

### Ptarmigan LMA simulations Energy spectra



#### **Ptarmigan LMA simulations Spatial distribution at creation**





#### **Ptarmigan LMA simulations** Spatial distribution at profiler



Projection taken at z = 11.5 m downstream of IP



### **Ptarmigan LMA simulations Energy weighted radiation profile**



ξ = 10.0

#### **Ptarmigan LMA simulations Inference of laser intensity**

• Inference of laser intensity follows <u>Blackburn et. al. 2020</u>

$$\xi = g(\rho) \,\xi_{inf} = \xi_{inf} \sqrt{\frac{1+8\rho^2}{1+4\rho^2}} \quad \rho = \xi_{inf}^2 = 4\sqrt{2} \langle \gamma_i \rangle \langle \gamma_f \rangle (\sigma_{\parallel}^2 - \sigma_{\perp}^2)$$

- Variance of angular profiles constructed using different methods: •
  - Variance of data
  - Variance of Gaussian fit
  - Variance of approximation to Cauchy fit (see slides 15-17)
  - Approximation of variance assuming a Gaussian FWHM
- Fittings are performed on a central region of width 0.06 mrad

$$=\frac{r_b}{w_0}$$

#### Ptarmigan LMA simulations Inference of laser intensity



reasonable based on slide 4



Reconstruction needs more work - alright for low xi but completely off for larger xi

# Additional slides

#### Extra **Approximation of Cauchy distribution**

Standard Cauchy distribution is given as

$$X \sim \text{Cauchy}(0,1) \Rightarrow f_X(x) = \frac{1}{\pi}$$

• A first order approximation to this is a triangle function of the form

$$g(x) = \begin{cases} \frac{1}{a} \left( 1 - \frac{|x|}{a} \right) & \text{for } - a \\ 0 & \text{elsewb} \end{cases}$$

- *a* is determined by the requirements of the approximation; two are considered here:
  - Peak values of distributions coincide  $\Rightarrow a = \pi$
  - Function must pass through FWHM points  $\Rightarrow a = 2$





#### **Extra** Triangular distribution

• The second moment of a distribution is defined as

$$m_2 = \mathbb{E}_X[X^2] = \int_{\mathbb{R}} dx \, x^2 f_X(x)$$

• Using  $f_X(x) = g(x)$  from previous slide

$$m_2 = \frac{a^2}{6} \text{ if } a > 0$$

• Hence, the RMS for each value of *a* is:

$$\operatorname{rms} = \frac{\pi}{\sqrt{6}} \text{ for } a = \pi$$
$$\operatorname{rms} = \sqrt{\frac{2}{3}} \text{ for } a = 2$$

#### Extra **General triangular distribution**

• For a Cauchy distribution with general width parameter  $\gamma$ , the PDF is

$$X \sim \text{Cauchy}(0,\gamma) \Rightarrow f_X(x) = \frac{1}{\pi\gamma} \frac{1}{1 + \left(\frac{x}{\gamma}\right)^2}$$

• Triangular approximation can be obtained by the substitution  $x \to \frac{x}{\gamma}$  or equivalently  $a \to \gamma a$ :

$$g(x) = \begin{cases} \frac{1}{a\gamma} \left( 1 - \frac{|x|}{a\gamma} \right) & \text{for } -a\gamma \le x \le a\gamma \\ 0 & \text{elsewhere} \end{cases}$$

• Hence, the second moments are given by  $m_2 = \frac{a^2 \gamma^2}{6}$  so

• rms = 
$$\frac{\pi\gamma}{\sqrt{6}} \approx 1.283 \,\gamma$$
 for  $a = \pi$ 

• rms = 
$$\gamma \sqrt{\frac{2}{3}} \approx 0.816 \gamma$$
 for  $a = 2$