# Bubbles from Dark confinement with Holography

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How Fast Does The Bubble Grow

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# SU(N) Yang-Mills



### GWs from FOPTs

#### ► LISA PT working group 2019



5D ED gravity in AdS  $\rightarrow$  4D Large  $N_c$  Yang-Mills Kiritsis, Nitti, Gursoy 2007

$$\mathcal{S}_5 = -M_\rho^3 N_c^2 \int d^5 x \sqrt{g} \left( R - rac{4}{3} (\partial \Phi)^2 + V(\Phi) 
ight) + 2M_
ho^3 \int_{\partial \mathcal{M}} d^4 x \sqrt{h} \mathcal{K},$$

- $V(\Phi)$  dilaton potential
- ▶ 5-D coordinate  $r \iff RG$  scale
- Dilaton  $\lambda = e^{\Phi} \iff$  t'Hooft coupling  $\lambda_t = N_c g_{YM}^2$
- AdS-BH/Thermal AdS  $\iff$  Phases of SU( $N_c$ )

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### The potential $V(\lambda)$ and It's Parameters

UV Asymptotics

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$$V(\lambda)=rac{12}{\ell^2}(1+v_0\lambda+v_1\lambda^2+...)$$

**IR** Asymptotics

$$V(\lambda) \sim \lambda^{rac{4}{3}} \left(\log(\lambda)
ight)^{rac{1}{2}}$$

 $V_0$   $V_2$  YM beta function 2-loops  $V_1$ ,  $V_3$  Phenomenological parameters

 $V_1$  Free gas asymptotics,  $V_3$  Latent heat

$$V_{1} = 170, \qquad V_{3} = 14. \qquad \frac{T}{T_{c}}$$

$$V(\lambda) = \frac{12}{\ell^{2}} \left( 1 + V_{0}\lambda + V_{1}\lambda^{\frac{4}{3}} (\log[1 + V_{2}\lambda^{\frac{4}{3}} + V_{3}\lambda^{2}]^{\frac{1}{2}}) \right) \qquad JG |U|$$

## Finite T Solutions





# Confinement (HP) Phase Transition & Effective Action I



$$V_{\text{eff}}(\lambda_h, T) = \mathcal{F}(\lambda_h) - 4\pi M_p^3 N_c^2 b(\lambda_h)^3 \left(1 - \frac{T_h}{T}\right) . \qquad \text{JG} \left| \mathbf{U} \right|^2$$

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## Confinement (HP) Phase Transition & Effective Action II

Kinetic Term Normalization

We vary  $c \rightarrow \frac{1}{3} - 3$ , Moderate dependence on GW spectrum Thermal Tunneling effective action O(3) symmetric bounce

$$S_{eff} = \frac{4\pi}{T} \int d\rho \, \rho^2 \left[ c \frac{N_c^2}{16\pi^2} (\partial_\rho \lambda_h(\rho))^2 + V_{\text{eff}}(\lambda_h(\rho), T) \right]$$

 $c \frac{N_c^2}{16\pi^2}$ 

Nucleation Rate for Thermal Tunneling

$$\Gamma = T^4 \left(\frac{\mathcal{S}_B}{2\pi}\right)^{3/2} e^{-\mathcal{S}_B}$$

Percolation:  $\mathcal{P}(true) \simeq \mathcal{P}(false)$  (End of PT! GW Emission)

	α	$\beta/H(v_w=1)$	$\beta/H(0.1)$	$\beta/H$ (0.01)	IC	h
$T_c = 50 \mathrm{MeV}$	0.343	$9.0 \times 10^4$	$8.6  imes 10^4$	$8.2  imes 10^4$	JO	
$100{ m GeV}$	0.343	$6.8 imes10^4$	$6.4 imes10^4$	$6.1 \times 10^4$	3	9 Q Q

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#### Gravitational Wave Spectra SU(3) "The Money Plot"



GW spectra for SU(3) at different critical temperatures

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Reference	2103.09827	2202.10503, 2205.06274	2104.12817
Theory	SU(N) YM	Strong Gauge theory	SUSY SU(N) YM
Approach	Pheno	Bottom-up	Top-Down
Supercooling	minimal	minor/moderate	moderate/large
V <sub>W</sub>	$v_w \sim 10^{-4}$	$v_w \sim \mathcal{O}(0.01-0.3)$	$v_w \sim \mathcal{O}(0.01-1)$

Employ steady state approach like Bigazzi 2104.12817

$$\Delta P_{fric}^{tot} = 0 = \frac{F_{fric}}{A} + \Delta P_{bubble}$$

Consider the drag force on an external quark stationary moving in the plasma as the plasma friction  $F_{fric} \sim F_{drag}$ .

Momentum flow along string  $\iff$  Drag force on quark moving in plasma.



From Dewolfe 2013

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$$F_{drag} = \frac{dp_1}{dt} = \pi_{\xi} = -\frac{ve^{2A(\lambda_s)}\lambda_s^{\frac{4}{3}}}{2\pi\ell_s^2}$$
$$\Delta P_{fric}^{tot} = 0 = \frac{F_{drag}}{A} + P_f(T_{boost}) - P_f(T_{perc})$$

 $T_{boost}$  is the 2-D worldsheet BH temperature generated by the induced metric.

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# Preliminary Results!

Access to the pressures in deconfined phase from the lattice fit. Finally for the wall velocity we obtain so far

#### $v_w = 0.07 \pm 0.03$

Further insights based on phenomenological arguments regarding this velocity computation are under consideration,  $F_{fric} \sim N_c^2 F_{drag}$ , reflected gluons of bubble wall ...

Take Home message: Quantitative predictions of GWs at strong coupling are at our grasp thanks to clever use of tools, even for not perfect CFTs!

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