SOMMERFELD ENHANCEMENT, SPIN EFFECTS and RADIATIVE CORRECTIONS in dark matter interactions

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SOMMERFELD EFFECT for coupled channels

Consider a process of the type

 $\chi_a \chi_b \to \chi_i \chi_j \to \chi_i' \chi_j' \to \ldots \to SM$ final states

where the intermediate pairs $\chi_i \chi_j$ can be the same or different as the initial pair $\chi_a \chi_b$ (fig.a)





recurrence relation for the annihilation amplitudes

In the non-relativistic limit

 $(A_{ij} \text{ full amplitude, } A_{ij}^0 \text{ tree level value}):$

$$A_{ij}(p) = A_{ij}^{0}(p) - \sum_{i'j'\phi} N_{ij,i'j'} \frac{g_{ii'\phi}g_{j'j\phi}}{(2\pi)^3} \int \frac{d^3k}{(\vec{p} - \vec{k})^2 + m_{\phi}^2} \frac{A_{i'j'}(k)}{\frac{\vec{k}^2}{2m_r^{i'j'}} - \mathcal{E} + 2\delta m_{i'j'}}$$

 \vec{p} CM three-momentum, $g_{ii'\phi}$ and $g_{j'j\phi}$ coupling constants, $N_{ij,i'j'}$ normalization and combinatorial factors.

Our convention : $\mathcal{E} = \vec{p}^2/2m_r^{ab}$ kinetic energy of the reference incoming pair a, b m_r^{ij} reduced mass, $2\delta m_{ij} = m_i + m_j - (m_a + m_b)$

To rewrite the expression above in the form of an inhomogeneous Schrödinger equations define :

$$A_{ij}(\vec{p}) = \left(\frac{\vec{p}^2}{2m_r^{ij}} - \mathcal{E} + 2\delta m_{ij}\right) \tilde{\psi}_{ij}(\vec{p})$$
(1)

$$U_{ij}^{0}(\vec{r}) = \int d^{3}\vec{p} \, e^{i\vec{p}\cdot\vec{r}} A_{ij}^{0}(\vec{p}, P_{0})$$
(2)

$$\psi_{ij}(\vec{r}) = \int d\vec{p} \, e^{i\vec{p}\cdot\vec{r}} \tilde{\psi}_{ij}(\vec{p}) \tag{3}$$

which allows us to rewrite the recurrence as a differential equation:

$$-\frac{\partial^2}{2m_r^{ij}}\psi_{ij}(\vec{r}) = U_{ij}^0(\vec{r}) + (\mathcal{E} - 2\delta m_{ij})\psi_{ij}(\vec{r}) + \sum_{i'j'\phi} V_{ij,i'j'}^{\phi}\psi_{i'j'}(\vec{r}),$$

 ϕ : the exchanged particle (scalar, vector or axial vector boson).

The potential has the form

$$V_{ij,i'j'}^{\phi}(r) = rac{c_{ij,i'j'}(\phi)}{4\pi} rac{e^{-m_{\phi}r}}{r},$$

 $c_{ij,i'j'}(\phi)$ depends on the couplings and states involved.

Since the Sommerfeld enhancement factorizes out as a distortion of the incoming wave-function, we first solve the associate homogeneous equation

with an initial state ab with momentum \vec{p} we get a set of

wave functions $\psi_{ij}^{ab,\vec{p}}(\vec{r})$

Finally one reconstructs the amplitude

 $A_{ab}(\vec{p}) = \lim_{p' \to p} (p'^2 - p^2) \int \frac{d^3k}{(2\pi)^6} \sum_{ij} \frac{\left[\int d^3r' e^{-i\vec{p}'\vec{r}} \psi_{ij}^{ab,\vec{k}}(\vec{r}')\right] \left[\int d^3r e^{i\vec{p}\vec{r}} \bar{\psi}_{ij}^{ab,\vec{k}}(\vec{r}) U_{ij}^0(\vec{r})\right]}{k^2 - p^2 - i\epsilon}$ $\to \sum_{ij} \int d^3r \bar{\psi}_{ij}^{ab,\vec{p}}(\vec{r}) \int \frac{d^3q}{(2\pi)^3} e^{i\vec{q}\vec{r}} A_{ij}^0(\vec{r})$

The result gets simple in partial waves.

Decomposing in partial waves, consider the reduced radial wave-function $\varphi(x)$:

$$R_{ij,p,l}(r) = Np \frac{\varphi_{ij,l}(x)}{x}, \qquad x = pr$$

getting equation, for the dominant S-wave, (a, b reference incoming pair)

$$\frac{d^2\varphi_{ij}(x)}{dx^2} + \frac{m_r^{ij}}{m_r^{ab}} \left[\left(1 - \frac{2\delta m_{ij}}{\mathcal{E}} \right) \varphi_{ij}(x) + \frac{1}{\mathcal{E}} \sum_{i'j'\phi} V_{ij,i'j'}^{\phi}(x) \varphi_{i'j'}(x) \right] = 0.$$

Boundary conditions: $\varphi(0) = 0$

At $x \to \infty$ the solution describes one incoming $\chi_a \chi_b$ state and all the possible $\chi_i \chi_j$ states that can be produced in the ladder.

Two cases for the states $\chi_i \chi_j$ at ∞ :

- 1. $2\delta m_{ij} < {\cal E}$ enough energy to produce states $\chi_i \chi_j$ on-shell
- 2. $2\delta m_{ij} > \mathcal{E}$ not enough energy; states $\chi_i \chi_j$ are off-shell

The radial wave functions behave at infinity as:

$$\begin{split} R_{ab}(r) &\to \frac{C_{ab}}{2i} \frac{e^{ik_{ab}r}}{r} - \frac{1}{2i} \frac{e^{-ik_{ab}r}}{r} \quad \text{incoming pair} \\ R_{ij}(r) &\to \left\{ \begin{array}{c} \frac{C_{ij}}{2i} \frac{e^{ik_{ij}r}}{r} & \text{if on-shell} \\ \frac{D_{ij}}{2i} \frac{e^{-ik_{ij}|r}}{r} & \text{if off-shell} \end{array} \right. \text{ every other } \chi_i \chi_j \end{split}$$

(Normalization consistent with $R_{ab} = \sin(k_{ab}r)/r$ for the non-interacting case) For the reduced wave functions at $x \to \infty$ (write φ_{ij}^{ab} for the initial state ab)

with
$$q_{ij} = m_r^{ij}/m_r^{ab} \cdot (1 - 2\delta m_{ij}/\mathcal{E})$$

 $i\varphi_{ab} - \partial_x \varphi_{ab} = -e^{-ix}$
 $\begin{cases} i\sqrt{q_{ij}}\varphi_{ij}^{ab} - \partial_x \varphi_{ij}^{ab} = 0 & \text{if on - shell} \\ \sqrt{-q_{ij}}\varphi_{ij}^{ab} + \partial_x \varphi_{ij}^{ab} = 0 & \text{if off - shell} \end{cases}$

these b.c.'s determine the wave functions $\varphi^{ab}_{ij}(x)$

One then reconstructs the amplitude and find the cross section.

The (co-)annihilation cross section of the pair $\chi_a \chi_b$ is determined, up to the kinematical factor, by

$$\sigma_{(ab)} \propto \sum_{ij} S^{ab}_{ij} \cdot |A^0_{ij}|^2$$

where the enhancement factors with our normalization are (for the S-wave)

$$S_{ij}^{ab} = |\partial_x \varphi_{ij}^{ab}|_{x=0}^2.$$

The Sommerfeld effect can be very large when there is a resonance.

An example in the case of just one channel for a Yukawa potential $rac{lpha}{r}e^{-\mu r}$



cross-section enhancement as a function of the co-annihilating mass in Tev, for $\mu=90$ Gev, $\alpha=1/30$ and an incoming velocity $v=10^{-3}$

The resonant Sommerfeld enhancement also occurs for higher partial waves. It is computed by taking higher derivatives of the reduced wave-function at the origin.



An example for the P-wave l = 1

enhancement as a function of m(Gev) for $v=10^{-3}$ $v=10^{-4}$ $v=10^{-5}$. Here $\alpha=1/30,~\mu=90Gev.$

In the following applications we consider the dominant l = 0 S-wave.

The enhancement depends on the spin of the initial state, taken to be two non-relativistic fermions.

Project the cross section in the singlet and in the triplet spin-state, and multiply each projection by a different enhancement factor.

Describe a fermion-antifermion (Dirac or Majorana) pair by

$$|\Phi_{\gamma}^{ij}
angle = N_{ij}\int dec{z}\, ar{\psi}_i(ec{z})\mathcal{O}_{\gamma}\psi_j(-ec{z})|0
angle \Phi_{\gamma}^{ij}(z).$$

For a (Dirac) fermion-fermion pair take $\psi \rightarrow \psi^c$.

The spin singlet S = 0 and the spin triplet S = 1 are encoded by (extending M.Drees, J.M.Kim, K.I.Nagao PRD2010 arXiv:0911.3795) :

S = 0: $\mathcal{O}_{\gamma} \equiv \gamma_5$, S = 1: $\mathcal{O}_{\gamma} \equiv \vec{\gamma} \cdot \vec{S}$,

Check, taking the non-relativistic large components :

$$\bar{\psi}_{a} = \begin{pmatrix} a_{\uparrow}^{\dagger} & a_{\downarrow}^{\dagger} & 0 & 0 \end{pmatrix} \quad \psi_{b} = C \bar{\psi}_{b}^{T}$$
$$\bar{\psi}_{a} \gamma_{5} \psi_{b} = -a_{\uparrow}^{\dagger} b_{\downarrow}^{\dagger} + a_{\downarrow}^{\dagger} b_{\uparrow}^{\dagger} \quad \bar{\psi}_{a} \gamma_{3} \psi_{b} = -a_{\uparrow}^{\dagger} b_{\downarrow}^{\dagger} - a_{\downarrow}^{\dagger} b_{\uparrow}^{\dagger}$$

The normalization is $N_{ij} = 1/\sqrt{2}$ $i \neq j$ $N_{ij} = 1/2$ i = j

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To compute the enhancement, evaluate how the coupling in front of the potential depends on the spin state.

Take the interaction (Γ denoting a generic gamma matrices structure)

$$V_{int} = g_{\Gamma}^2 \int d\vec{x} d\vec{y} \, \bar{\psi}_k(x) \Gamma \psi_i(x) V_{ki,jl}^{\phi}(\vec{x} - \vec{y}) \bar{\psi}_j(y) \Gamma \psi_l(y),$$

where $V_{ki,jl}^{\phi}(r) = \frac{g_{ki}^{\Gamma}g_{jl}^{\Gamma}e^{-m_{\phi}r}}{4\pi}$ is the non-relativistic potential due to the propagator of the boson exchanged between the two vertices. In the non-relativistic limit only $\Gamma = 1, \gamma_0, \gamma_j \gamma_5$ can contribute.

Compute
$$V_{int} \cdot \int d\vec{z} N_{ij} \bar{\psi}_i(\vec{z}) \mathcal{O}_\gamma \psi_j(-\vec{z}) |0\rangle \Phi_\gamma^{ij}(z)$$

by doing contractions:

$$\langle \psi_i(x)\bar{\psi}_i(z)\rangle \rightarrow \delta(\vec{x}-\vec{z})rac{1+\gamma_0}{2} \quad \langle \bar{\psi}_l(y)\psi_l(w)\rangle \rightarrow -\delta(\vec{y}-\vec{w})rac{1-\gamma_0}{2},$$

getting

$$V_{int} \cdot \Phi_{\gamma}^{ij}(z) \to \Phi_{\gamma}^{kl}(x) = c(klij,\gamma)V_{kl,ij}^{\phi}(|\vec{x}|) \ \Phi_{\gamma}^{ij}(x)$$

The evaluation of $c(klij, \gamma)$ fixes the strength of the potential in the Schrödinger equations.

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A benchmark model

only the fermions of the MSSM and their $SU(2) \times U(1)$ interactions :

- a triplet of Majorana fermions in the adjoint representation call them Wino's with a gauge invariant mass ${\cal M}$
- -two left-handed doublets in the fundamental representation call them Higgsino's with a gauge invariant μ -mass term

One can reorganize those fermions and their Lagrangians into:

- a Wino triplet
$$\lambda^+, \lambda^- = (\lambda^+)^c, \lambda^0 = (\lambda^0)^c$$
 $(Y = 0)$
 $L_{Wino} = \bar{\lambda}^+ (i\gamma^\mu \partial_\mu - g\gamma^\mu W^3_\mu - M)\lambda^+ + \frac{1}{2}\bar{\lambda}^0 (i\gamma^\mu \partial_\mu - M)\lambda^0$
 $+ g\bar{\lambda}^+ \gamma^\mu W^+_\mu \lambda^0 + g\bar{\lambda}^0 \gamma^\mu W^-_\mu \lambda^+$

– a Higgsino Dirac-fermions-doublet $\psi = (\psi^0, \psi^-)$ (Y = -1)

$$L_h = ar{\psi}(i\gamma^\mu\partial_\mu - g\gamma^\mu W^a_\mu au^a + rac{g'}{2}B_\mu - M_h)\psi$$

where $M_h = \mu$ of the μ -term.

RELIC DENSITY COMPUTATION

Take a set of N particles χ_1 , χ_2 , ... χ_N , with mass $m_i \ m_1 \le m_2 \le ... \le m_M$) such that:

i) inelastic scattering on SM thermal bath turn each state into another one ii) χ_1 is stable.

A system of N coupled Boltzmann equations for each number density n_i

as after freeze out all heavier states decay into the lightest \Rightarrow a single equation for $n = \sum_{i} n_i$, (e.g. J. Edsjo P.Gondolo PRD1997 he-ph/9704361) :

$$\frac{dn}{dt} + 3Hn = -\langle \sigma_{\text{eff}} v \rangle \left[n^2 - (n^{eq})^2 \right]$$

P.Gondolo J.Edsjo P.Ullio M.Schelke E.A.Baltz JCAP2004 astro-ph/0406204

The effective thermally averaged annihilation cross section

$$\langle \sigma_{\rm eff} v \rangle = \sum_{i,j} \langle \sigma_{ij} v_{ij} \rangle \frac{n_i^{eq}}{n^{eq}} \frac{n_j^{eq}}{n^{eq}}$$

is a weighted sum over the thermally averaged annihilation cross sections $\chi_i + \chi_j \rightarrow X$ (in the dilute limit, two-body states dominate):

$$\langle \sigma_{ij} v_{ij} \rangle = \frac{1}{n_i^{eq} n_j^{eq}} \sum_X \int \frac{d^3 p_i}{2 E_i} \frac{d^3 p_j}{2 E_j} \frac{d^3 p_X}{2 E_X} \delta^4(p_i + p_j - p_X) f_i^{eq}(p_i) f_j^{eq}(p_j) |A_{ij \to X}|^2,$$

 $n_i^{eq}=\int d^3p f_i^{eq}(p_i)$ is the thermal equilibrium number density for the species i, $n^{eq}=\sum_i n_i^{eq}~.$

Here we will include the Sommerfeld effect: $|A_{ab}|^2 = \sum_{ij} S_{ij}^{(ab)} |A_{ij}^0|^2$,

with $S_{(ab)}^{ij}$ computed as said before.



Relic density Ωh^2 in the $\mu = M_{Higgsino} - M_2 = M_{Wino}$ plane, *left*): perturbative case, *right*) with Sommerfeld effect included. The brighter the colour the higher Ωh^2 . The solid line corresponds to the value (with 1 σ) for relic density (7-year WMAP) $\Omega h^2 = 0.1123 \pm 0.0035$.

call Wino-region $M_2 \ll \mu$ and Higgsino-region $\mu \ll M_2$.

The Sommerfeld enhancement gives a sharp shift in the WMAP-Wino-region to heavier masses; a pure Wino is found at $\Omega h^2 = 0.11$ for a mass of about 2.8 TeV.

Much milder effects of the Sommerfeld enhancement are seen in the Higgsino-region.

The results found for pure Winos and pure Higgsino are similar to J.Hisano S.Matsumoto M.M. Nojiri O.Saito PRD2006 hep-ph/0212022 and M.Cirelli A.Strumia M.Tamburini NPB 2007 arXiv:0706.4071 in the same limiting cases;

we have more annihilation channels and a careful dependence on the spin state of the annihilating particle pair (e.g. a different coefficient -3 in the axial vector exchange in agreement with Drees et al:2009) and a better control of the numerical solution of the Boltzmann Eq.

Another interesting feature is a spike in the region connecting the pure Higgsino to the pure Wino limit, towards $M_2 \sim \mu$ but still with a predominant Wino component.



In this region, we show in detail the ratio: relic density including Sommerfeld enhancement divided by relic density without it.



A thin "resonance" slice appears in the plane, starting for pure Winos with mass $m_{\chi^0} \approx 2.5$ TeV and extending to heavier masses into the region with a sizable Higgsino fraction, where the relic density agrees with observations.

The value of the mass we find for a pure Wino $\frac{1}{m_W} \approx \frac{1}{\alpha m_{\chi^0}}$, indicating a possible loosely bound state: Bohr radius ~ interaction range.

Radiative corrections to the Sommerfeld enhancement from corrections to the potential appearing in the Schrödinger equation.

2-body potential ⇔ sum of the 2-particle-irreducible diagrams. The crossed box is 2-particle-irreducible:



it gives a correction $\sim O(\alpha^2) \frac{e^{-2m_{\phi}r}}{r}$ compare with the leading term $O(\alpha) \frac{e^{-m_{\phi}r}}{r}$. It is suppressed by α and it goes faster to zero for large r: it would give a minor correction to the Sommerfeld enhancement.

Other corrections come from the the radiative corrections to the vertex and to the propagators.

They contain large Log's and could re-sum to the running of the coupling constant.

Question: should we put the coupling at the $m \sim Tev$ scale ?

Here two scales: $m \sim Tev$, $m_w \sim 0.1 Tev$ (and momentum transfer $\leq m_W$)

a situation differing from the standard renormalization-group analysis with one large overall scale; here more similar to the case of the forward scattering.

Radiative corrections to the vertex and propagators



Take SU(N) (N=2), call the vertex (suppressing group indices) $-ig\bar{\Psi}\gamma\cdot W\Psi$

m mass of $\Psi,~m_w$ mass of $W,~\mu$ renormalization scale \bar{q}^2 square momentum transfer $\bar{q}^2 \leq m_w^2.$

$$\begin{split} &Va: -ig\bar{\Psi}\gamma \cdot W\Psi \ \frac{g^2}{(4\pi)^2}(-\frac{N}{2}+C(r))\times (\frac{2}{\epsilon}-\\ &2\int_0^1 dx\int_0^{1-x} dy \ Log[\frac{(x+y)^2m^2+(1-x-y)m_w^2+xy\bar{q}^2}{\mu^2}]\\ &+2\int_0^1 dx\int_0^{1-x} dy \ \frac{(-3+4(x+y)+(1-x-y)^2)m^2-(1-x)(1-y)\bar{q}^2}{(x+y)^2m^2+(1-x-y)m_w^2+xy\bar{q}^2}\)\\ &C(r)=N \ \text{adjoint and} \ C=\frac{N^2-1}{2N} \ \text{fundamental}. \end{split}$$

The last term is a $UV\mbox{-finite contribution, however }IR$ divergent for $m_w^2/m^2 \rightarrow 0.$

Large Log's corrections, dropping $2/\epsilon$:

$$Va: -ig\bar{\Psi}\gamma \cdot W\Psi \; \frac{g^2}{(4\pi)^2} (-\frac{N}{2} + C(r)) \times (-Log[\frac{m^2}{\mu^2}] - 2Log[\frac{m^2}{m_w^2}])$$

$$\begin{array}{rcl} Vb: & -ig\bar{\Psi}\gamma \cdot W\Psi \; \frac{g^2}{(4\pi)^2} \frac{N}{2} \times \left(\frac{6}{\epsilon} - 2\int_0^1 dx \; Log[\frac{m_w^2 + x(1-x)\bar{q}^2}{\mu^2}] - \\ & & 2\int_0^1 dx \int_0^{1-x} dy \; Log[\frac{(1-x-y)m^2 + (x+y)m_w^2 + xy\bar{q}^2}{\mu^2}] \\ & & -\int_0^1 dx \int_0^{1-x} \; \frac{4(1-x-y)(2-x-y)m^2 - 2(x+y)\bar{q}^2}{(1-x-y)^2m^2 + (x+y)m_w^2 + xy\bar{q}^2}) \end{array}$$

The last term is a UV-finite contribution, however IR divergent for $m_w^2/m^2 \to 0.$

Large Log's corrections, dropping 6/ ϵ :

$$Vb: -ig\bar{\Psi}\gamma \cdot W\Psi \ \frac{g^2}{(4\pi)^2} \frac{N}{2} \times (-2Log[\frac{m_w^2}{\mu^2}] - Log[\frac{m^2}{\mu^2}] - 2Log[\frac{m^2}{m_w^2}])$$

There are radiative corrections to the vertex induced by the Fermion Ψ wave-function renormalization: $F wf: -ig\bar{\Psi}\gamma \cdot W\Psi \frac{g^2}{(4\pi)^2}C(r) \times (-\frac{2}{\epsilon} + 2\int_0^1 dxx \ Log[\frac{(1-x)^2m^2+xm_w^2}{\mu^2}] + 4m^2\int_0^1 dx \ \frac{(2-x)x(1-x))}{(1-x)^2m^2+xm_w^2})$

We find the large Log's corrections, dropping $2/\epsilon$:

$$F \ wf: \ -ig\bar{\Psi}\gamma \cdot W\Psi \ \frac{g^2}{(4\pi)^2}C(r) \times (Log[\frac{m^2}{\mu^2}] + 2Log[\frac{m^2}{m_w^2}])$$

By summing the Va, Vb, F wf contributions, the large Log's containing m^2 in the argument disappear:

$$-ig\bar{\Psi}\gamma \cdot W\Psi \ \frac{g^2}{(4\pi)^2}\frac{N}{2} \times (-2Log[\frac{m_w^2}{\mu^2}])$$

Actually, by performing integrations by parts in the integrals over the x, y parameters, it is seen that for $\vec{q} = 0$ the sum of the three contributions is exactly equal to

$$-ig\bar{\Psi}\gamma \cdot W\Psi \ \frac{g^2}{(4\pi)^2}\frac{N}{2} \times \left(\frac{4}{\epsilon} - 2\int_0^1 dx \ Log[\frac{xm_w^2}{\mu^2}]\right)$$

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This is consistent with the Slavnov-Taylor identities: $Z_{\bar{\Psi}W\Psi} = 1 + Va + Vb \qquad Z_{\Psi} = 1 - Fp$ $\frac{Z_{\bar{\Psi}W\Psi}}{Z_{\Psi}}$

is the same for any $\Psi \Rightarrow$ it cannot depend on the Ψ mass m.

To complete the survey of the radiative corrections, we include the correction induced by (one-half of the) Gauge boson W wave function renormalization:

$$G wf: -ig\overline{\Psi}\gamma \cdot W\Psi \frac{g^2}{(4\pi)^2} \times (N\frac{5}{3\epsilon} - \frac{5}{6} Log[\frac{m_w^2}{\mu^2}] + "matt")$$

where "*matt*" denotes the contribution of the matter, i.e. leptons, quarks and possibly higgs, of the form

"matt" = $-(\#)(\frac{2}{\epsilon} - Log[\frac{m_w^2}{\mu^2}])$ (here we show large Log's only).

This correction does not depend on the fermion mass \boldsymbol{m}

By summing everything we get the total correction to the coupling constant g (dropping c_0/ϵ absorbed by a renormalization at a reference scale):

$$g(1 + \frac{g^2}{(4\pi)^2} \times (\{N\frac{11}{3} - (\#)\}\{-\frac{1}{2} \ Log[\frac{m_w^2}{\mu^2}]\})) \equiv g(m_w^2)$$

At the lowest order we recover indeed, with $g = g(\mu^2)$:

$$g(m_w^2)^2 = \frac{g(\mu^2)^2}{1 + \frac{g(\mu^2)^2}{(4\pi)^2} \{N\frac{11}{3} - (\#)\} \ Log[\frac{m_w^2}{\mu^2}]}$$

that is the standard running of the SU(N) coupling constant at the scale m_w .

In conclusion:

by taking the coupling constant $g(m_w^2)$ at the m_w^2 scale, the radiative corrections to the vertex $\bar{\Psi}\gamma \cdot W\Psi$ (appearing in the evaluation of the Sommerfeld effect) are negligible for $\bar{q}^2 \leq m_w^2 \ll m^2$ (actually, no correction at all for $\bar{q}^2 \to 0$).

The same result holds for the U(1) interaction due to the U(1) Ward identity.