

Light new physics at muon factories

Diego Redigolo

Talk based on

[2006.04795](#) [hep-ph] with *L. Calibbi, R. Ziegler, J. Zupan*

[2203.11222](#) [hep-ph] with *S. Knapen and Y. Jho*

work in progress with *S. Knapen, Y. Jho, K. Langhoff, T. Opferkuch*

Rare decays of SM particles as a probe of New Physics

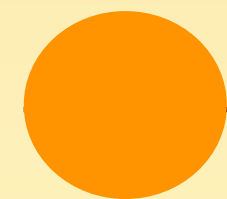
Accidental symmetries play a crucial role in the Standard Model and testing their breaking with increasing precision is one of standard ways we expect New Physics to manifest itself at experiments

Rare decays of SM particles are standard candles to probe NP



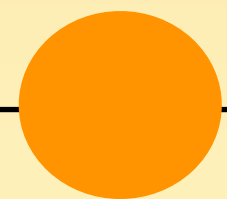
muon rare decays are a great example

1986 $\sim 10^7$ /year(s)



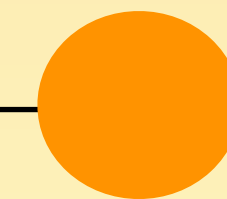
TRIUMF:
Jodidio's

2014 $\sim 10^8$ /year



TRIUMF:
Twist

soon $\sim (10^8 - 10^9)$ /s

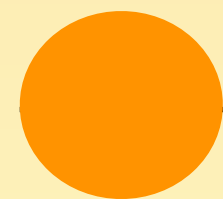


PSI: MEGII+ Mu3e

$t \sim \#\mu's$

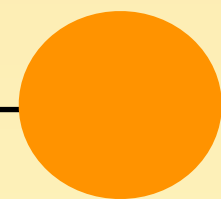
Goal of this talk

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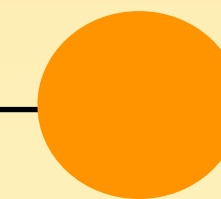
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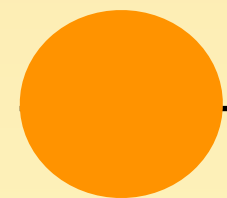
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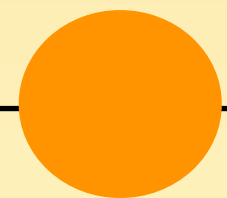
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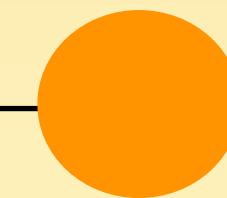
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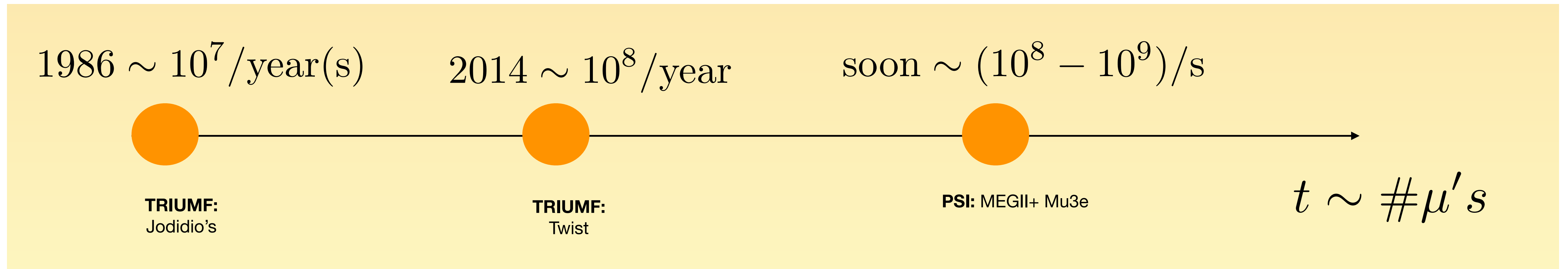
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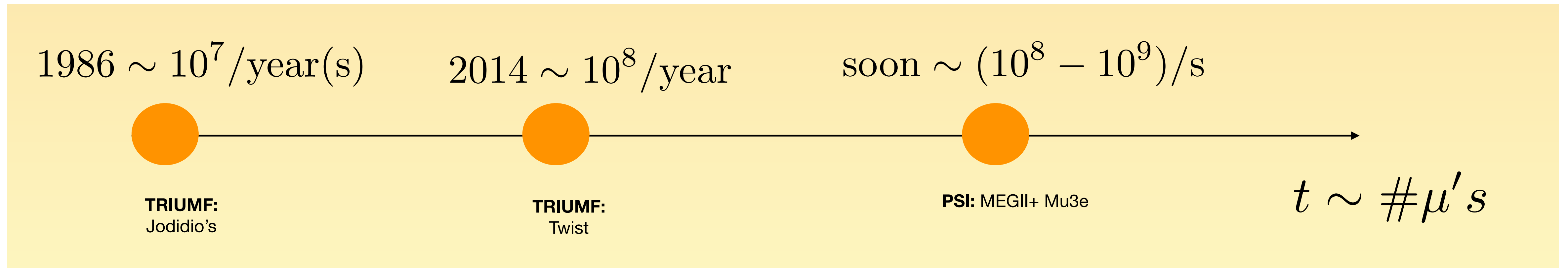


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(especially if data taking requires tweaks at trigger level)



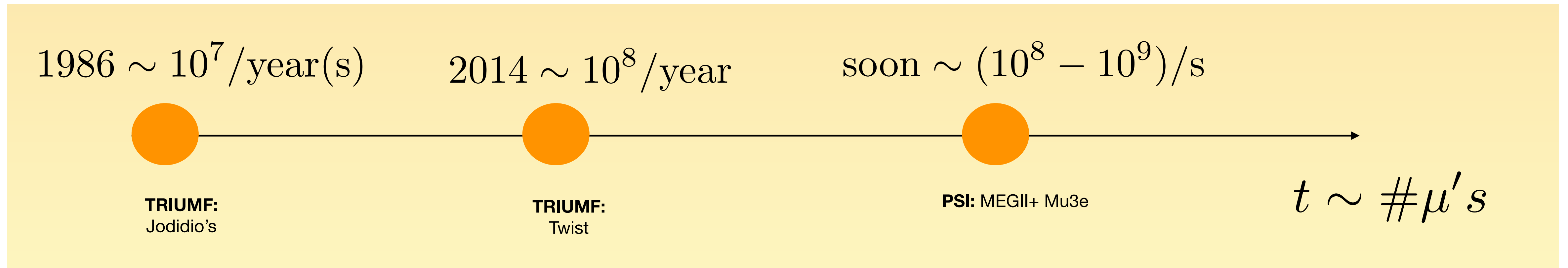
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As a bonus we will show that this program can actually probe reasonable models



Heavy New Physics vs Lepton Flavor

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda} \sum_a C_a^{(5)} Q_a^{(5)} + \frac{1}{\Lambda^2} \sum_a C_a^{(6)} Q_a^{(6)} + \dots$$

LFV @ dimension 6

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$U(1)_e \times U(1)_\mu$

Forbids $\mu \rightarrow e\gamma$

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$$\frac{C_{ij}^{(5)}}{\Lambda} (L_i H)(L_j H)$$

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$$\text{BR}(\mu \rightarrow e\gamma) \sim \frac{\alpha}{8\pi} \times \left(\frac{m_\nu}{M_W}\right)^4$$

obscenely small!

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Sensitive to $\gtrsim 10^4$ TeV NP for $C_{e\mu}^{(6)} \sim 1$

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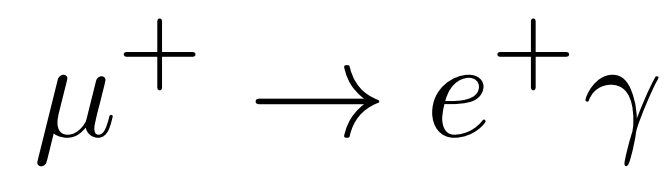
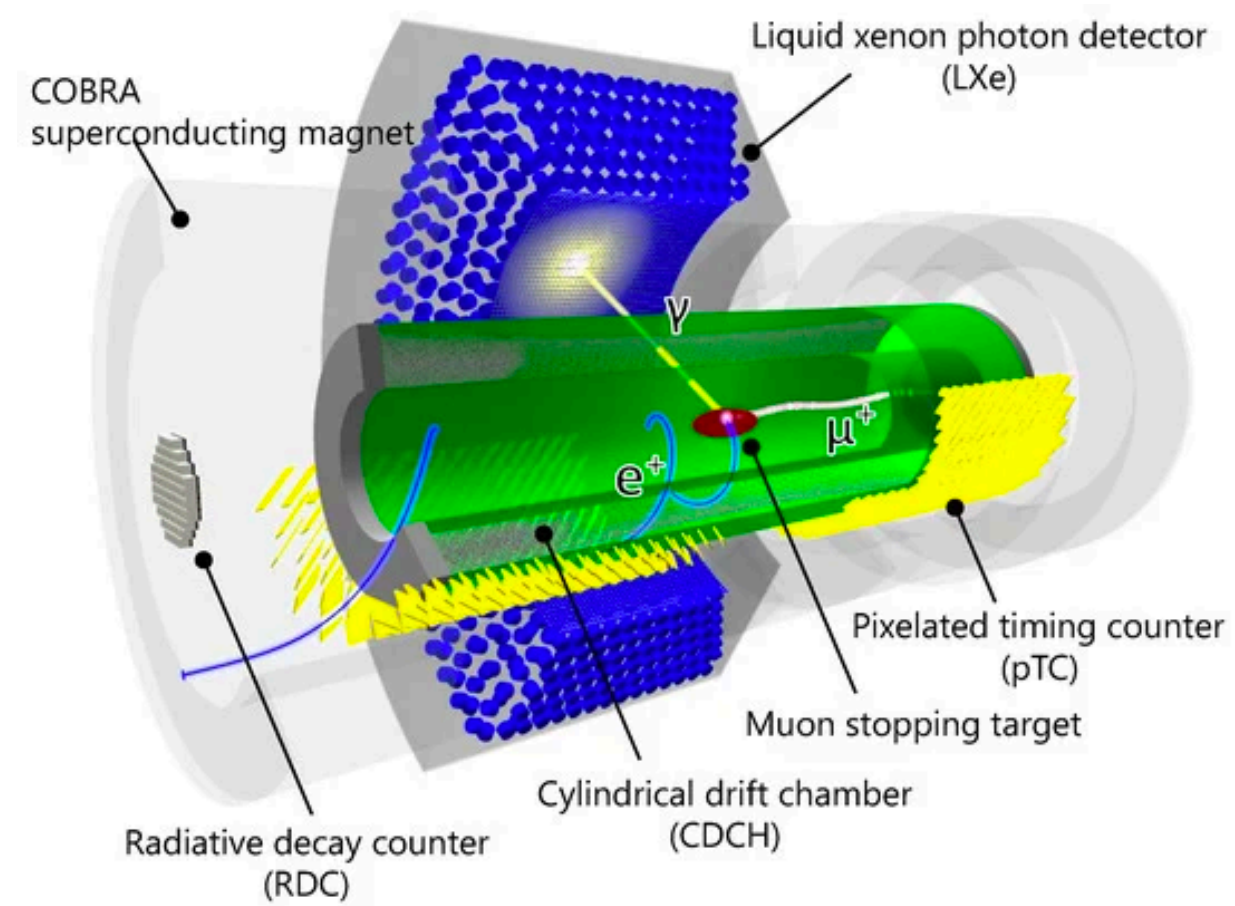
Sensitive to $\gtrsim 10^4$ TeV NP for $C_{e\mu}^{(6)} \sim 1$

The UV models probed by these observables are related to NP at the TeV scale: (SUSY, Compositeness...)

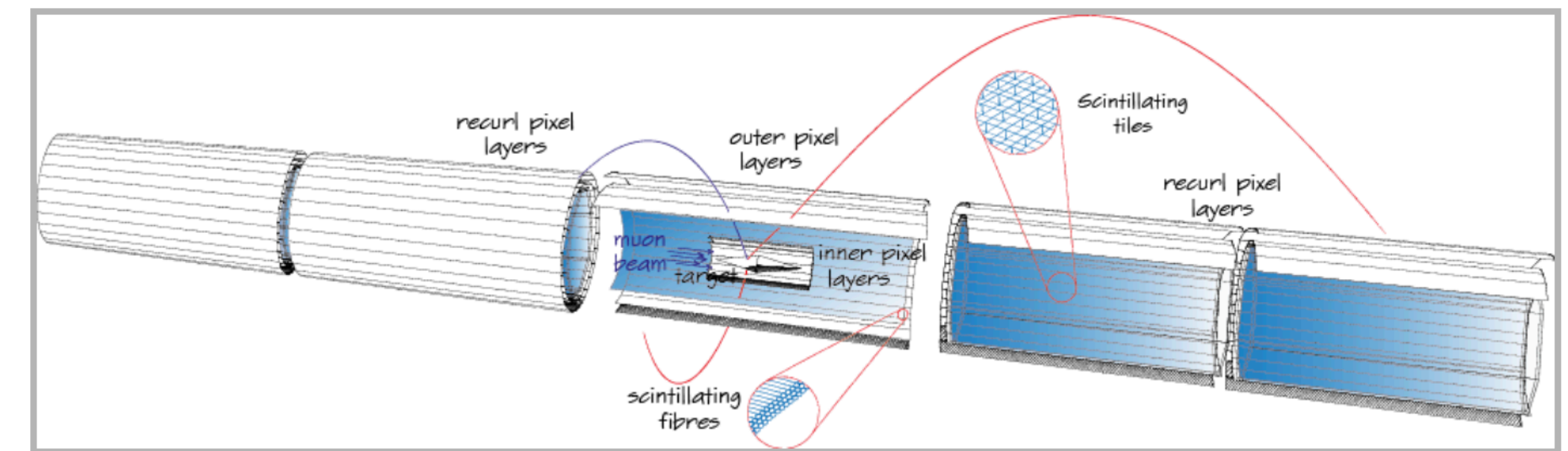
L. Calibbi's talk this morning

Heavy new physics experimental paradigm

MEG II



Mu3e



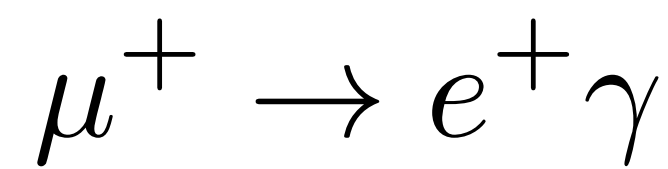
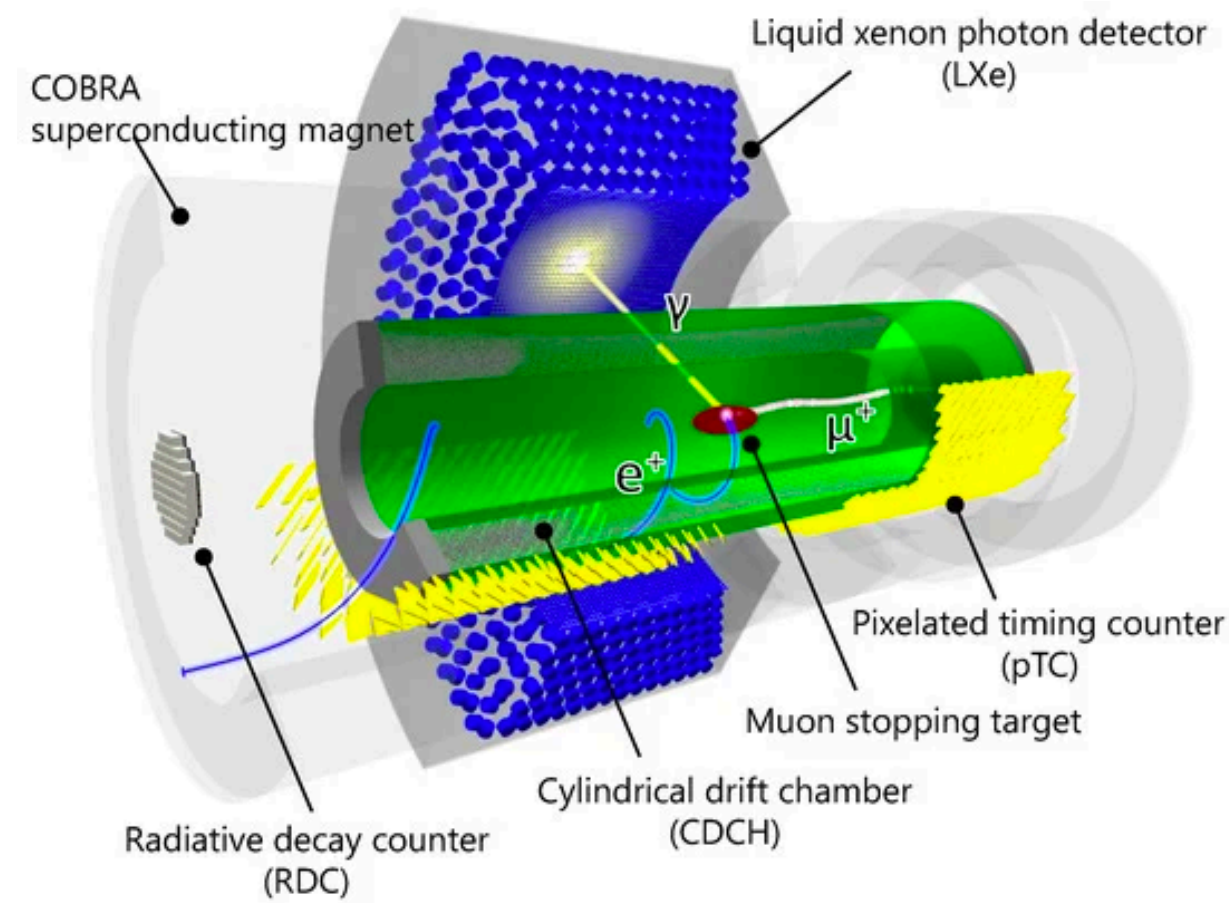
Schematic view of the planned Mu3e experiment



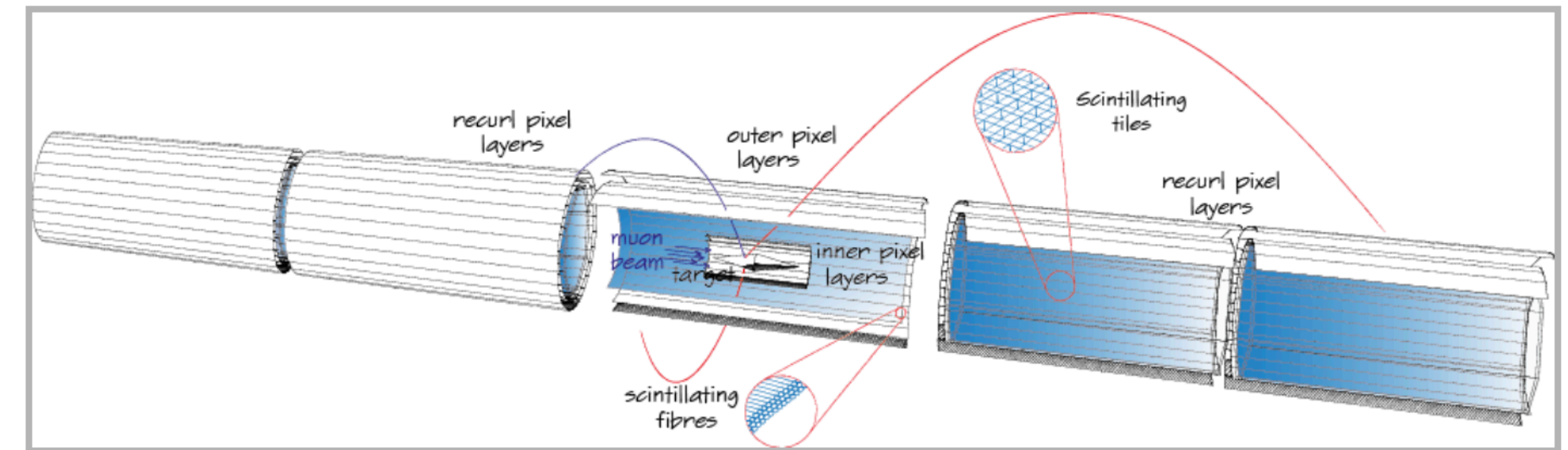
Signal characterised by **no missing energy**

Heavy new physics experimental paradigm

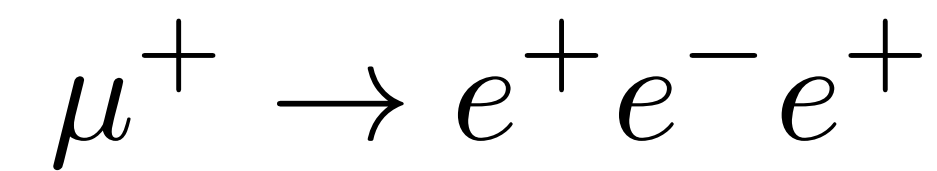
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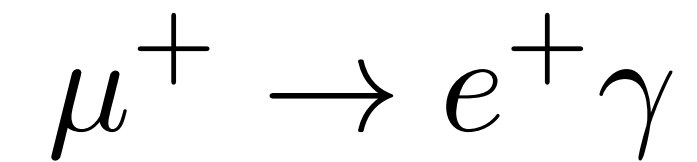
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Signal characterised by **no missing energy**

The challenge is to **increase the luminosity** while suppressing coincidences
tightening the kinematic + timing requirements

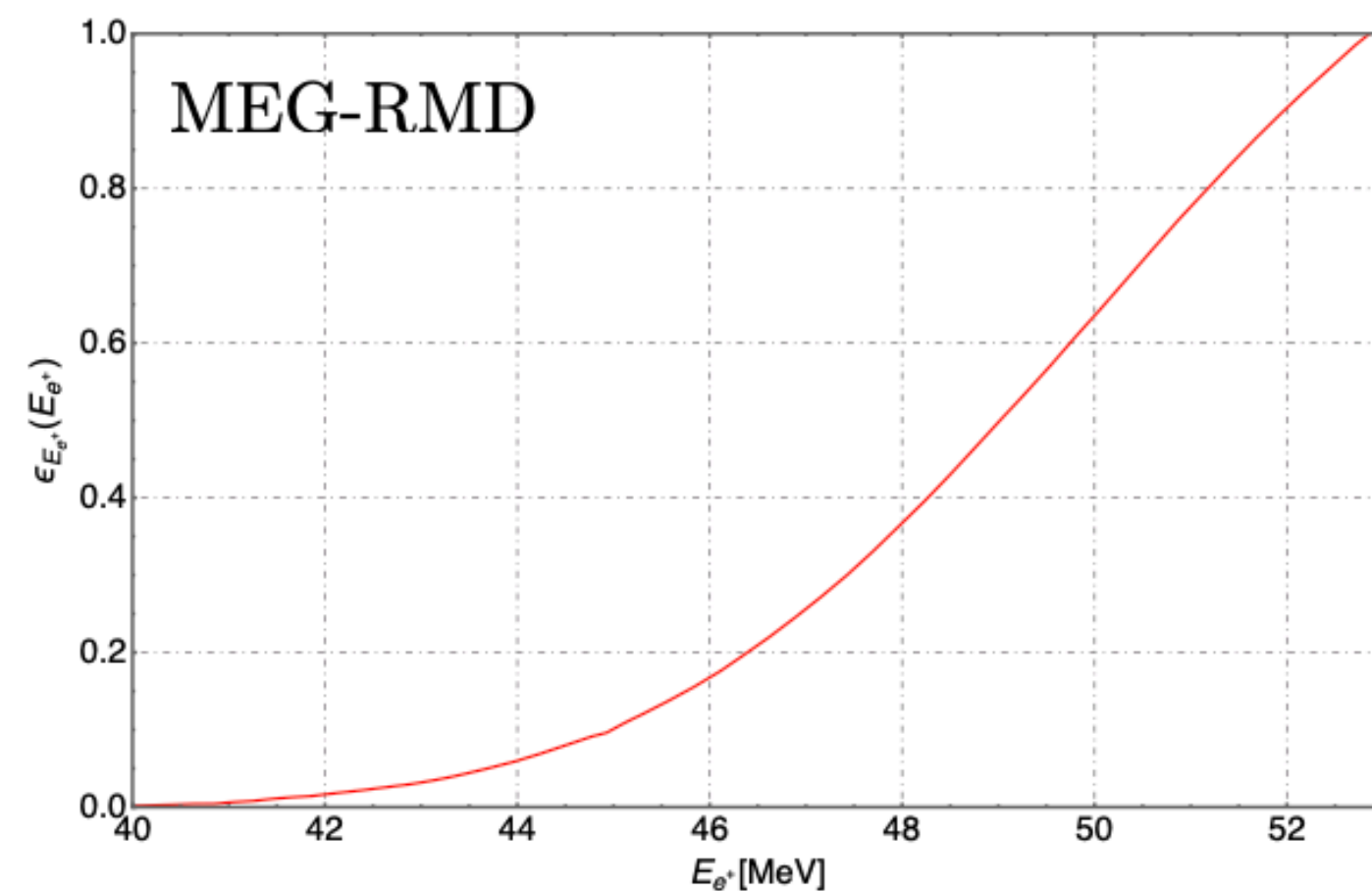
Example: trigger selection at MEG



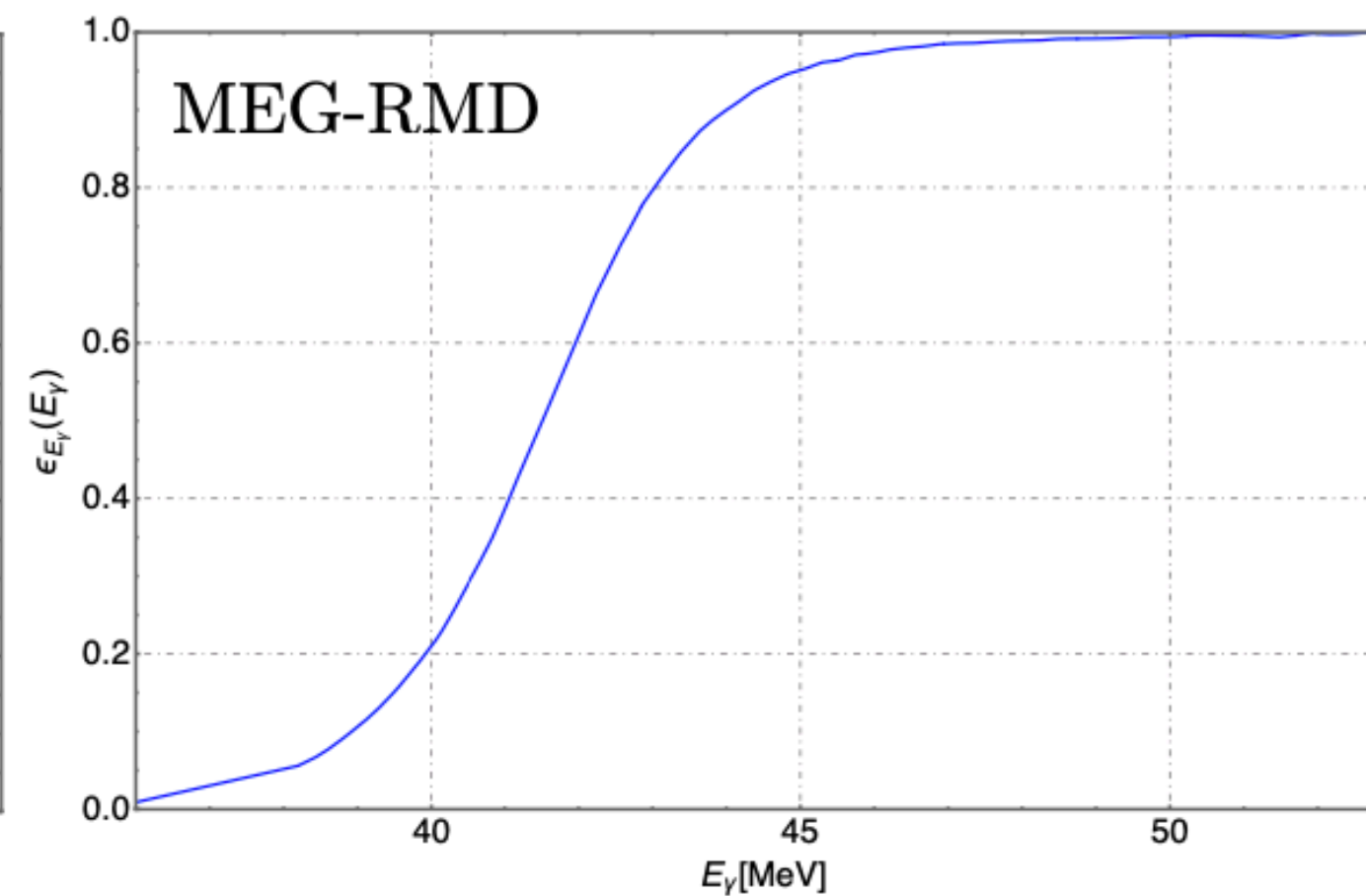
The trigger maximize the efficiency to back to back positron-photon of $E = m_\mu/2$

See Galli et al. *JINST* 9 (2014)

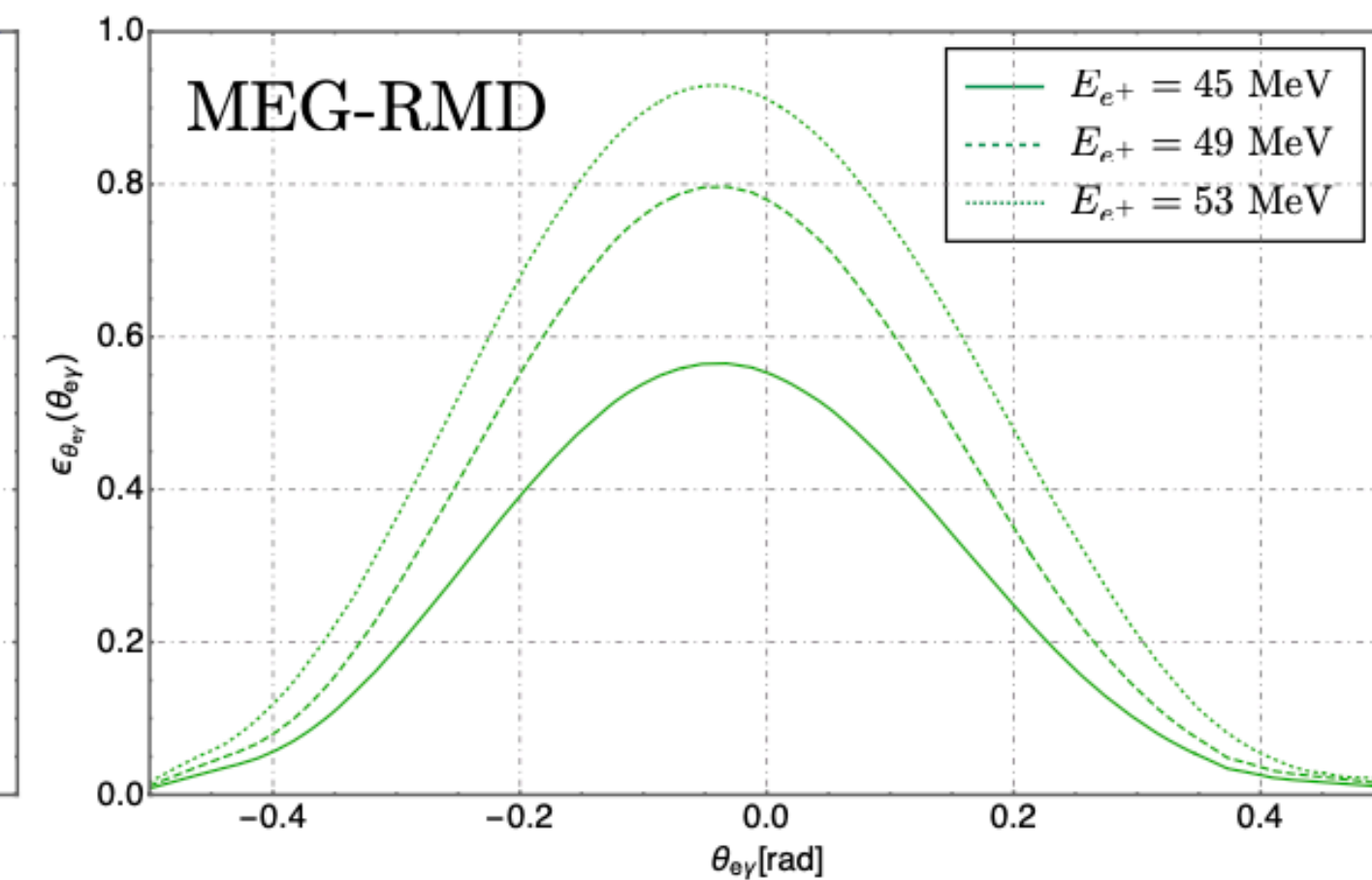
Positron energy >45 MeV @ hardware level



Photon >45 MeV @ trigger level



back to back topology @ trigger level



Taken from *MEG-RMD measurement 1312.3217*

Light new physics vs Lepton Flavor

Accidental symmetries of the Standard Model might be broken by light new particles feebly coupled to the SM

These light particles naturally emerges in models where Lepton Flavor is broken spontaneously at high scale
(familon, axion, axion-like particles, majorons)

See. *L. Calibbi, D.R., R. Ziegler, J. Zupan 2006.04795*

Light pseudo-Goldstone bosons (or ALP) $m_a \ll m_\mu$

LFV @ dimension 5

$$\mathcal{L}_{\text{eff}}^{\text{LFV}} \supset \frac{\partial_\mu a}{2f_a} \bar{\mu} \gamma^\mu (C_{\mu e}^V + C_{\mu e}^A \gamma_5) e + \frac{\partial_\mu a}{f_a} \bar{e} \gamma^\mu \gamma_5 e + \frac{m_a^2}{2} a^2 + \frac{1}{f_a^2} \dots$$

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Heavy scale
not accessible
with dipoles

$$c\tau \sim \frac{8\pi f_a^2}{m_a^3} \gg L_{\text{detector}}$$

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axial and vectorial LFV of the ALP

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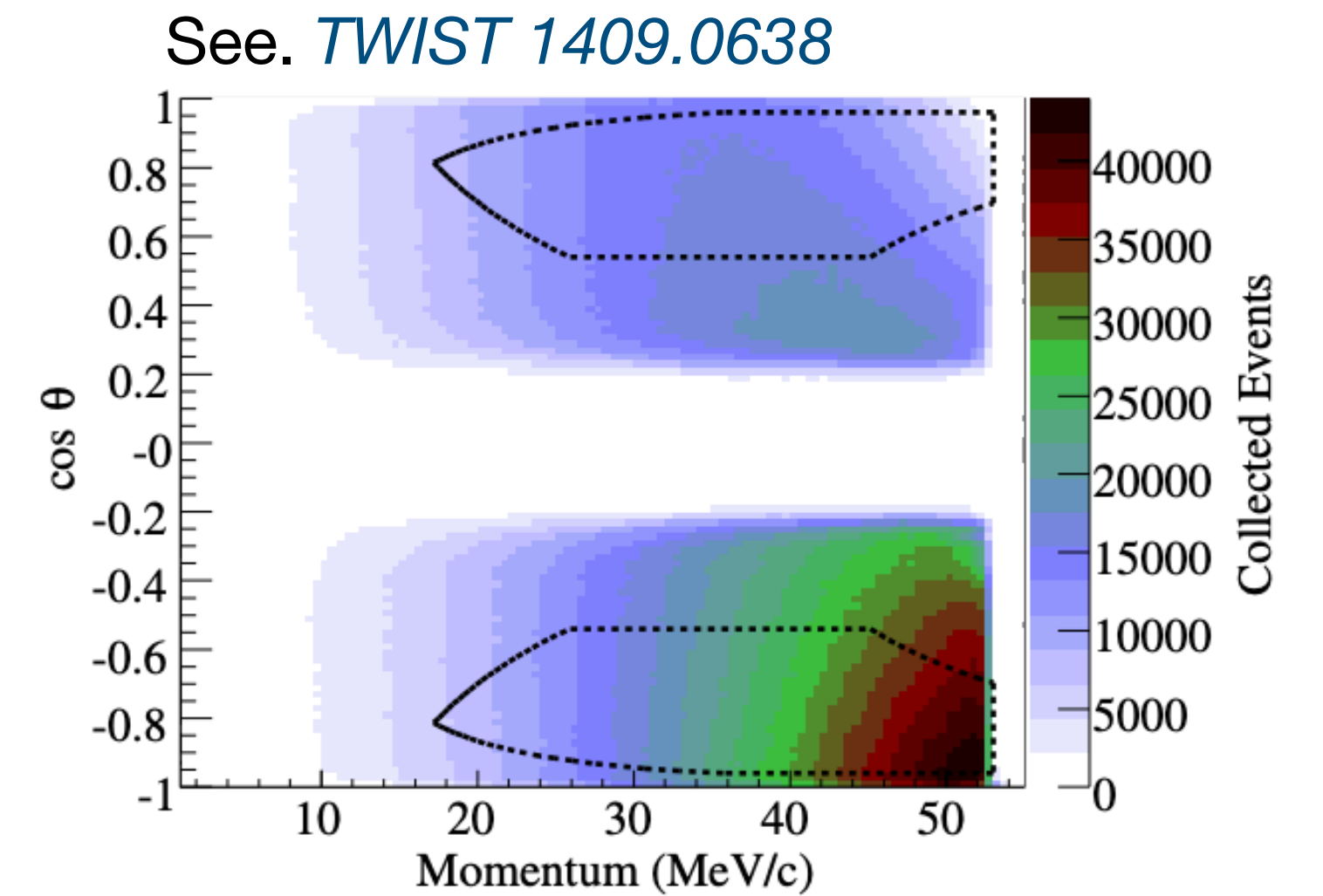
Interplay between flavor experiments and astrophysics

Light new physics experimental paradigm

Hunt for rare muon decays with missing energy

$$\mu \rightarrow ea$$

Huge irreducible background from Michel $\mu \rightarrow e\nu\bar{\nu}$



Light new physics experimental paradigm

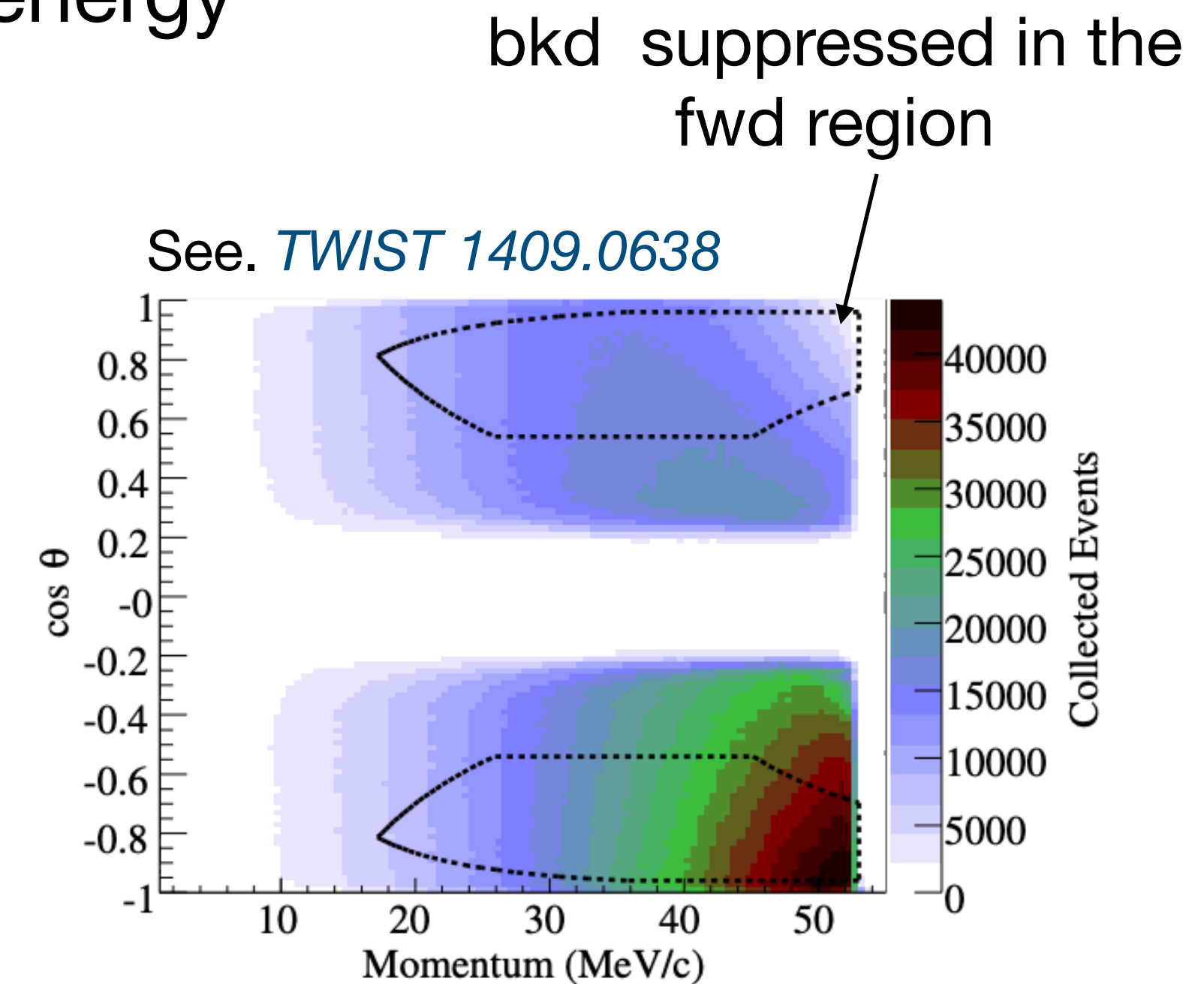
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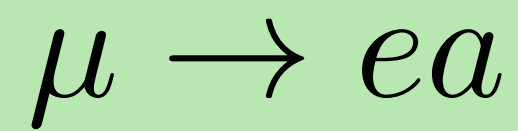
Muon polarization can help discriminating the signal

See. *L. Calibbi, D.R., R. Ziegler, J. Zupan 2006.04795*



Light new physics experimental paradigm

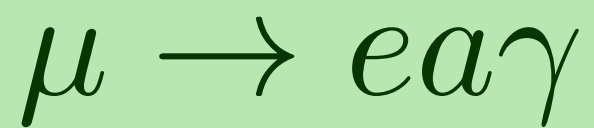
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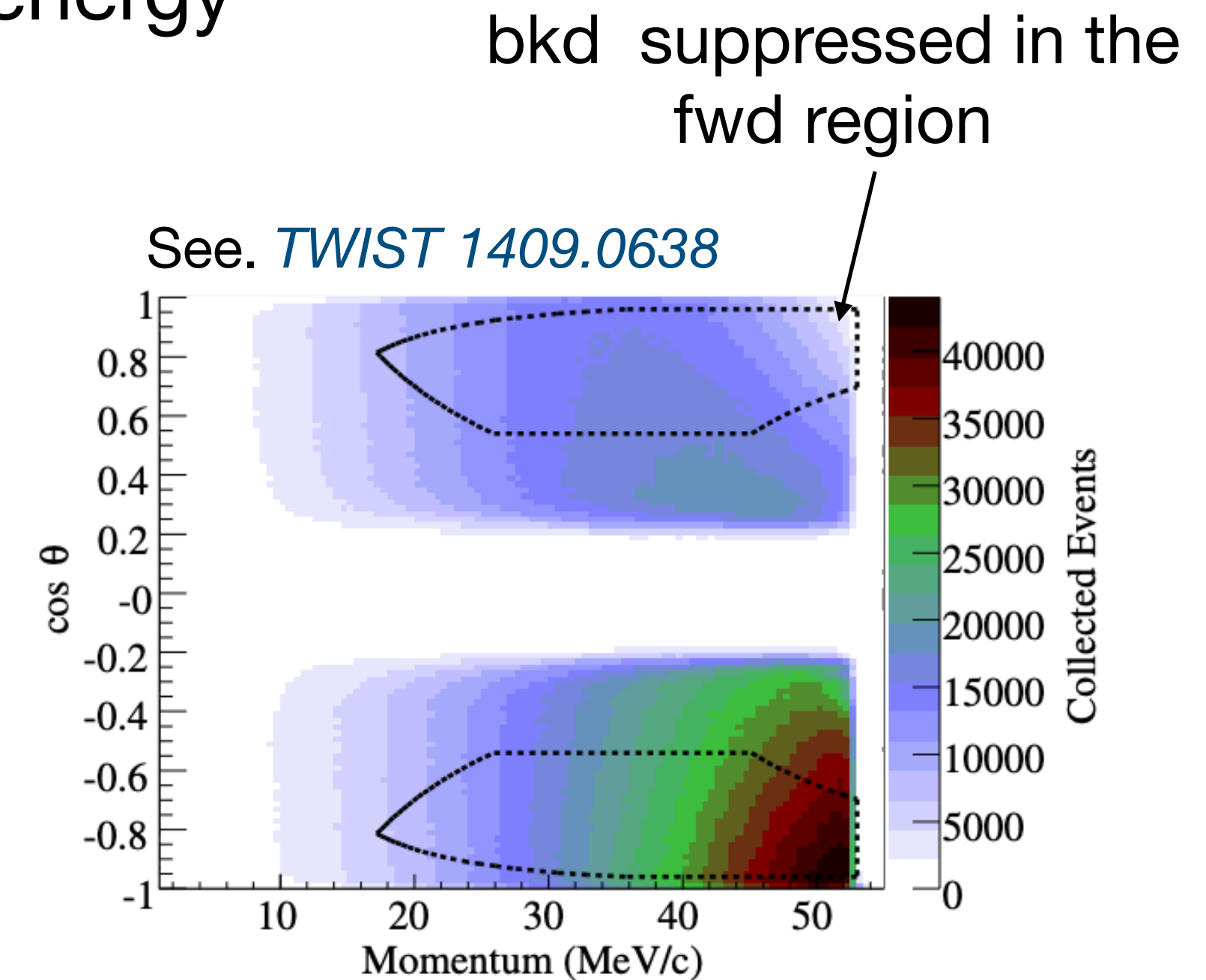
See. [L. Calibbi, D.R., R. Ziegler, J. Zupan 2006.04795](#)



The extra photon helps constructing a missing mass distribution which is not used for calibration

The price to pay is a reduced signal by $\sim \frac{\alpha}{2\pi} \log \frac{2E_\gamma}{m_\mu}$

See. [S. Knapen and Y. Jho 2203.11222](#)



Light new physics experimental paradigm

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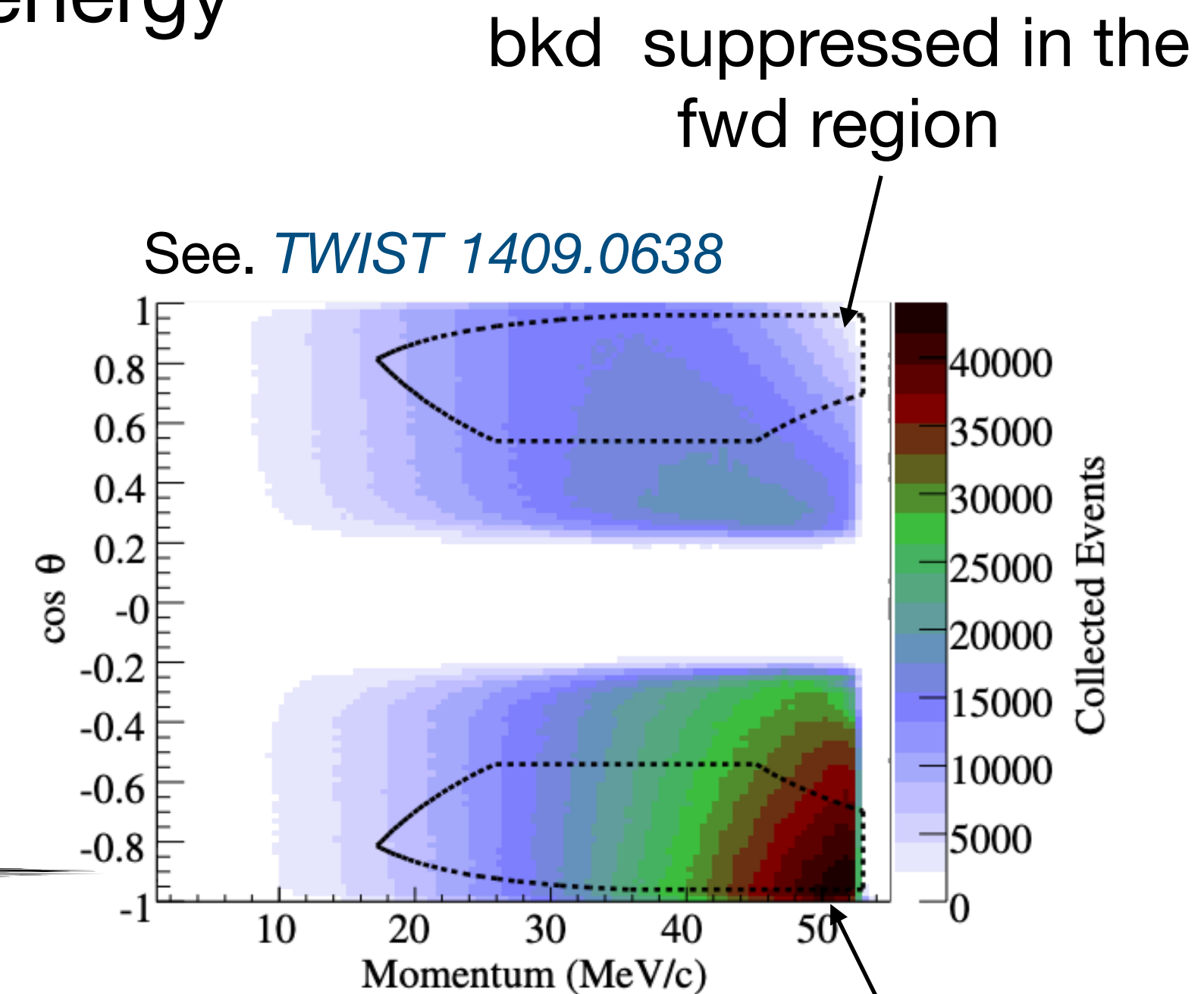
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For left-handed couplings the limitation are large systematics



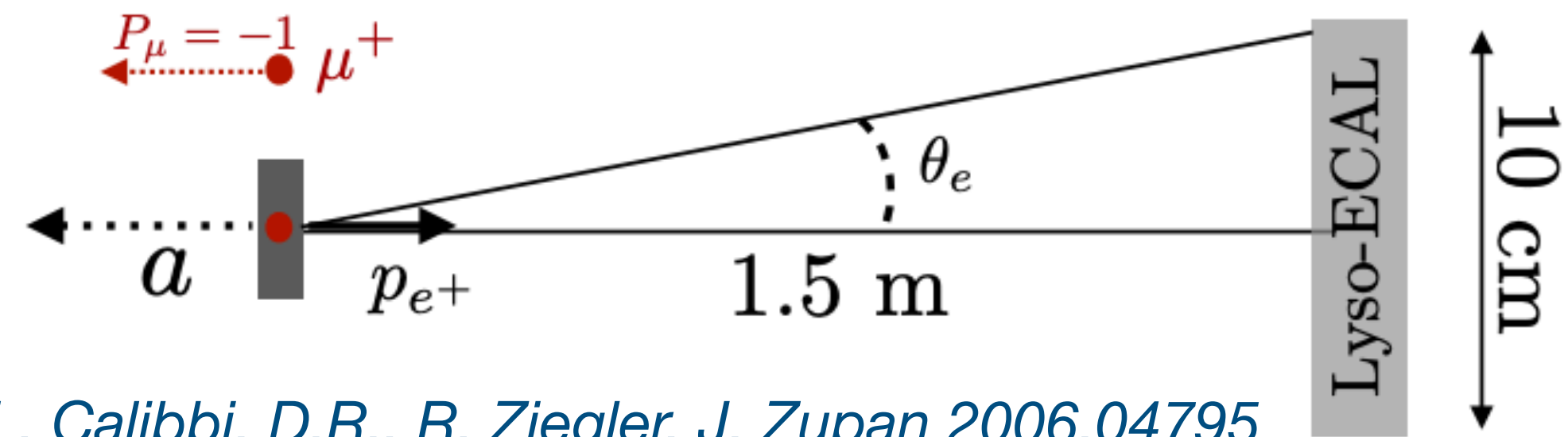
too large bkd

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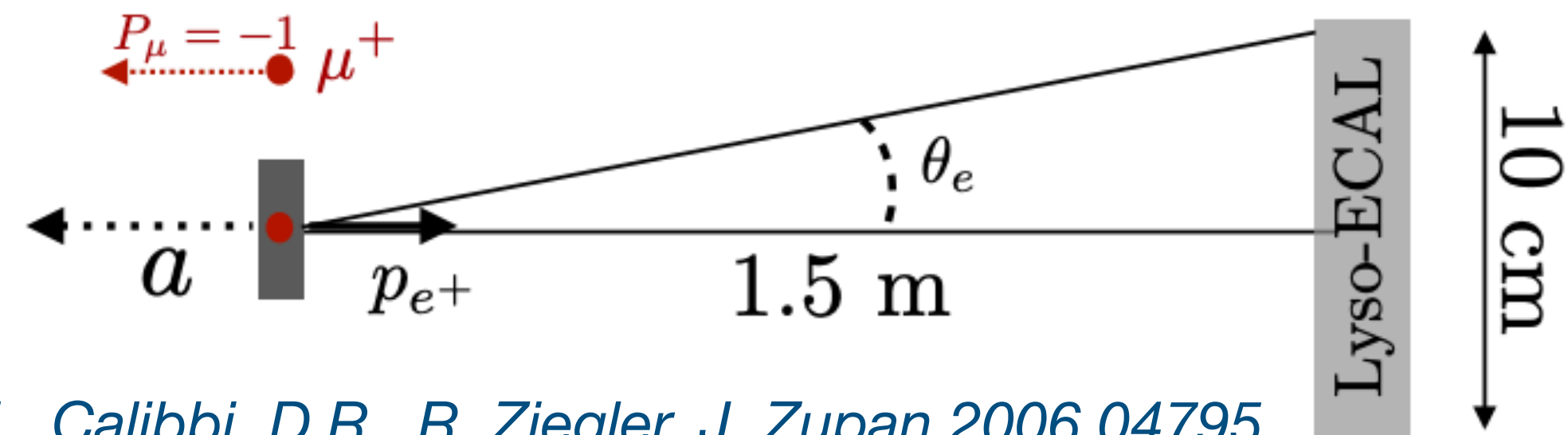
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Looking forward for right-handed ALPs



L. Calibbi, D.R., R. Ziegler, J. Zupan 2006.04795

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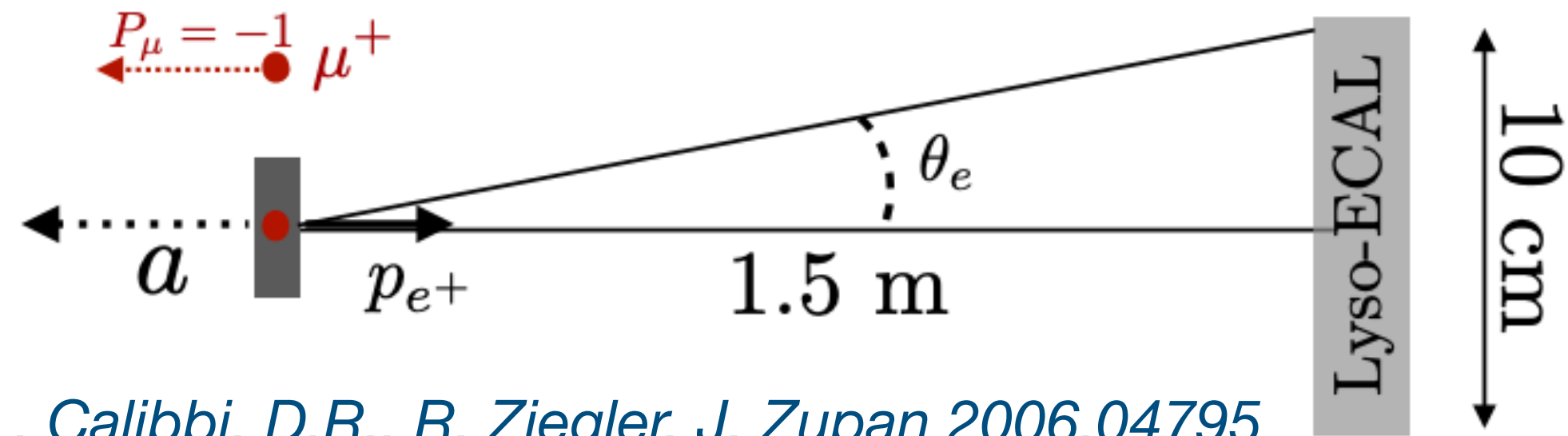


L. Calibbi, D.R., R. Ziegler, J. Zupan 2006.04795

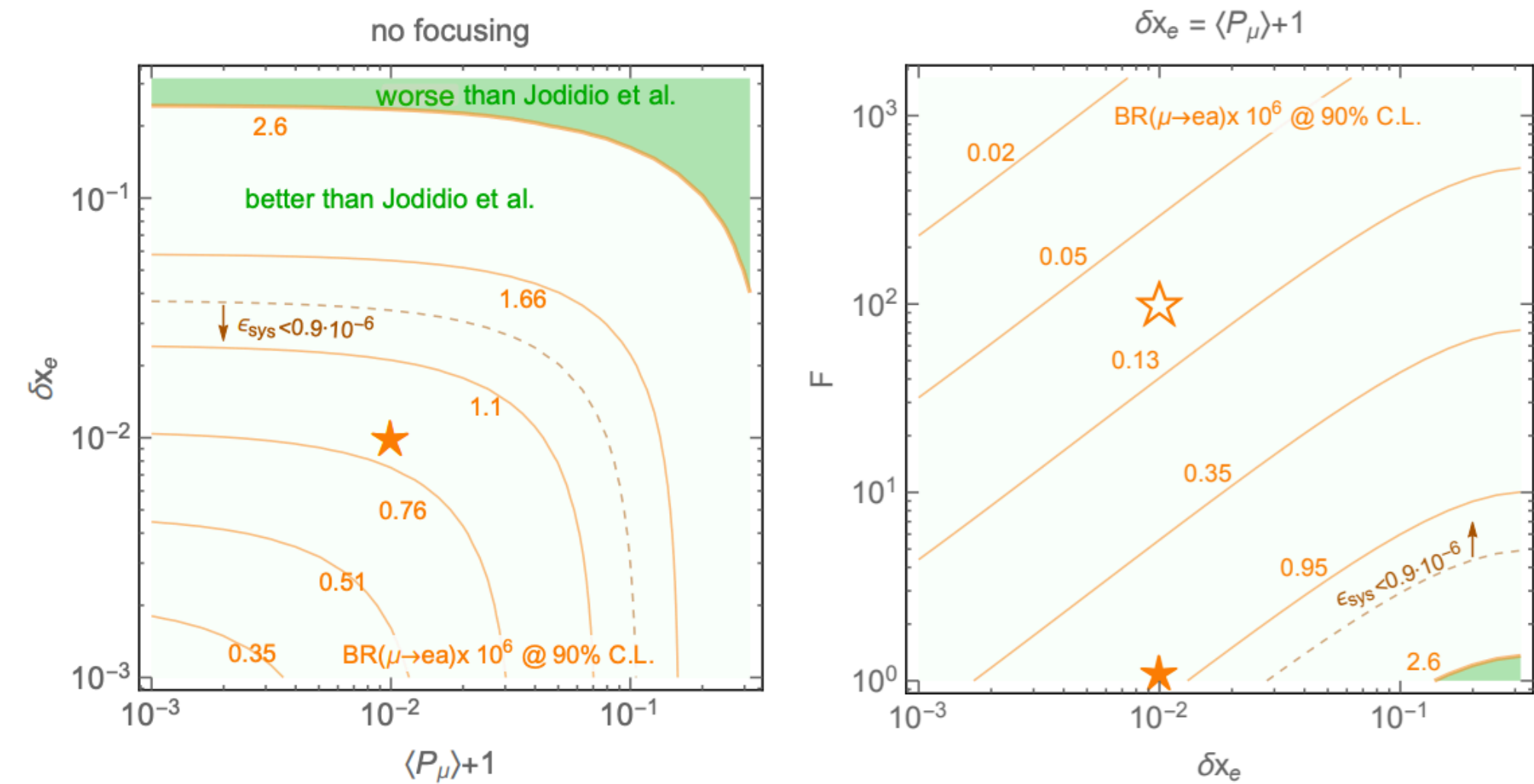
Background suppression in the fwd direction requires:

- 1) good momentum resolution $\delta x_e \sim \%$
- 2) purely polarized muon beam $\delta P_\mu \sim 10^{-2}$

Looking forward for right-handed ALPs



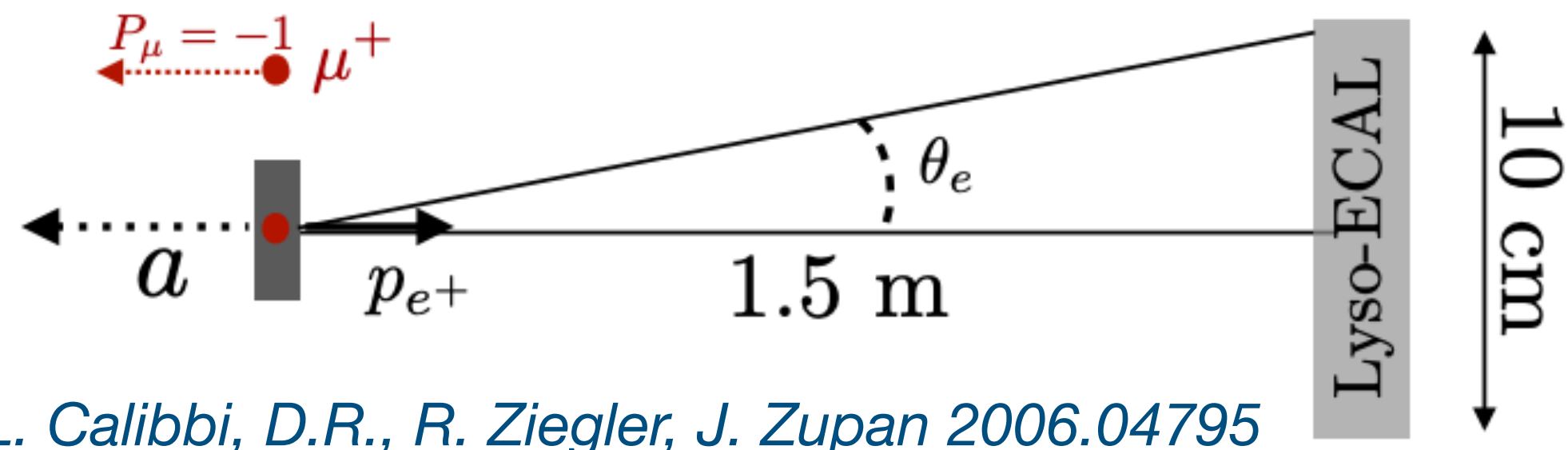
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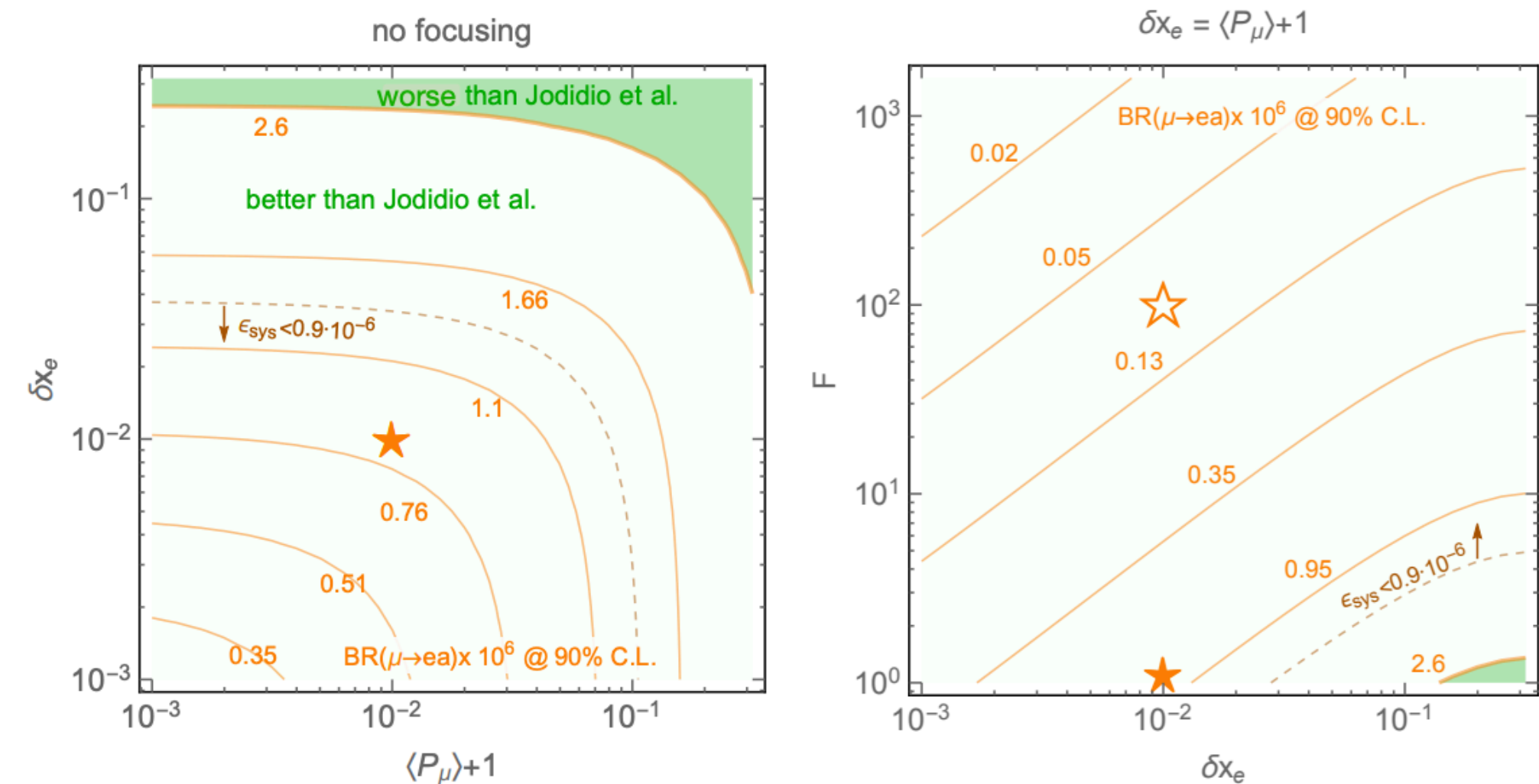
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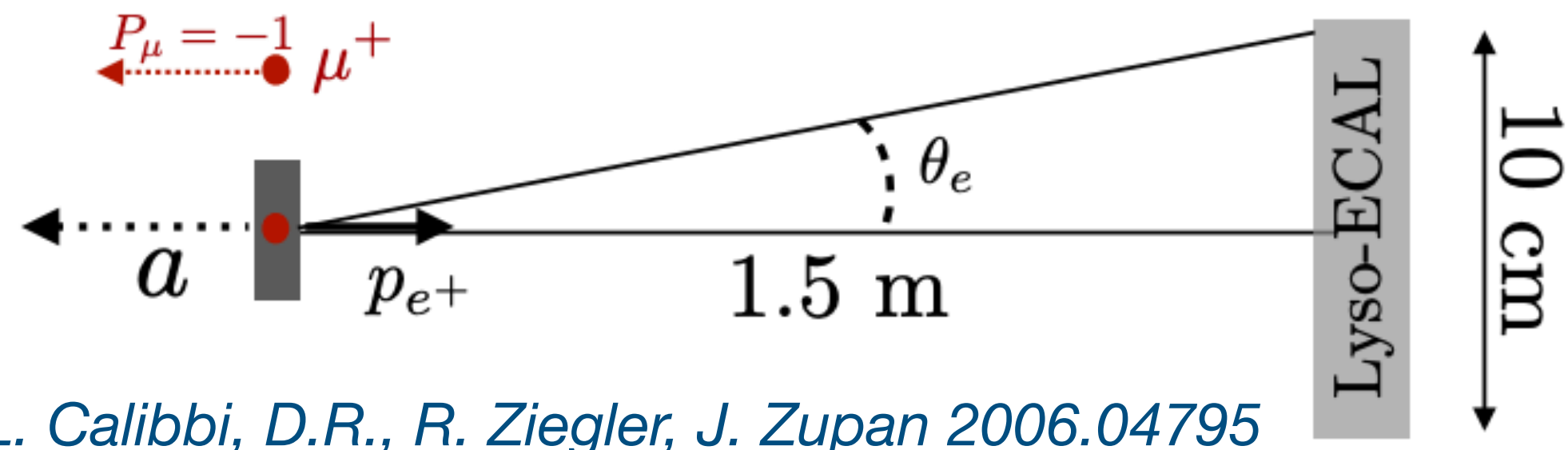


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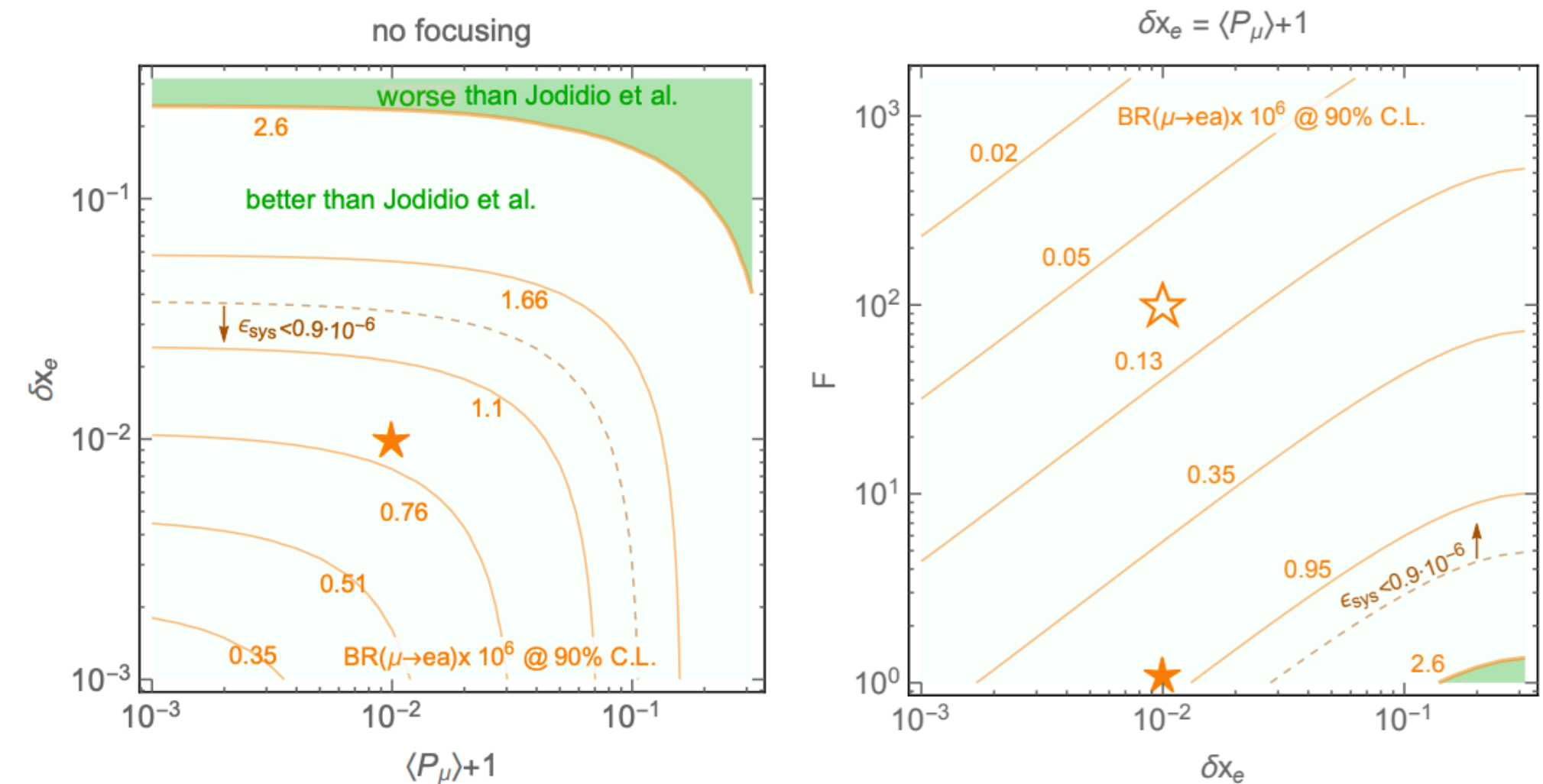


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- A good signal reach requires further:
- 3) magnetic field focusing $F \sim 10^2$
 - 4) large luminosity $N_\mu \sim 10^{14} \mu^+$
 - 5) very low systematics

Looking forward for right-handed ALPs



L. Calibbi, D.R., R. Ziegler, J. Zupan 2006.04795



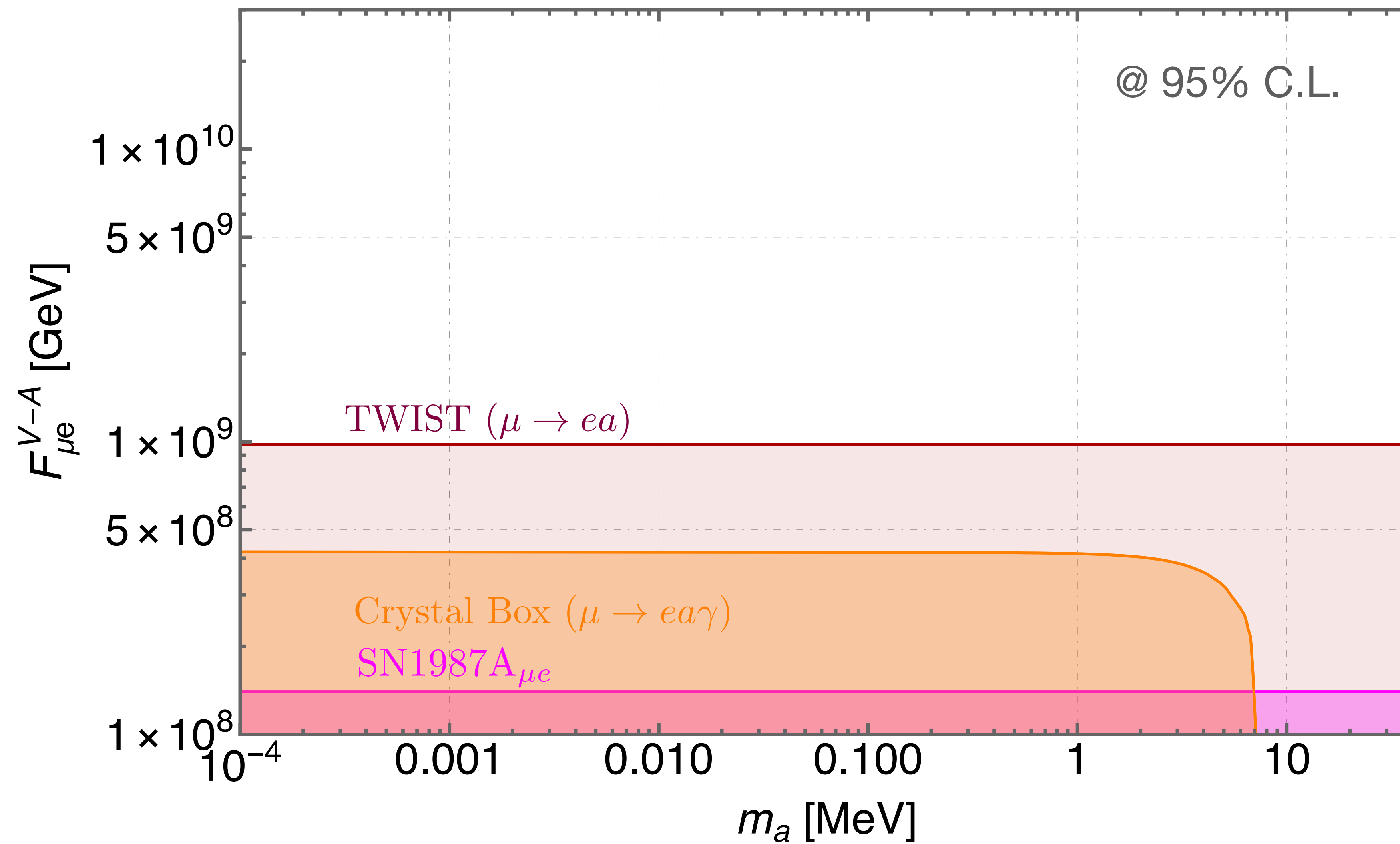
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These conditions have been realized before in the Jodidio's exp.

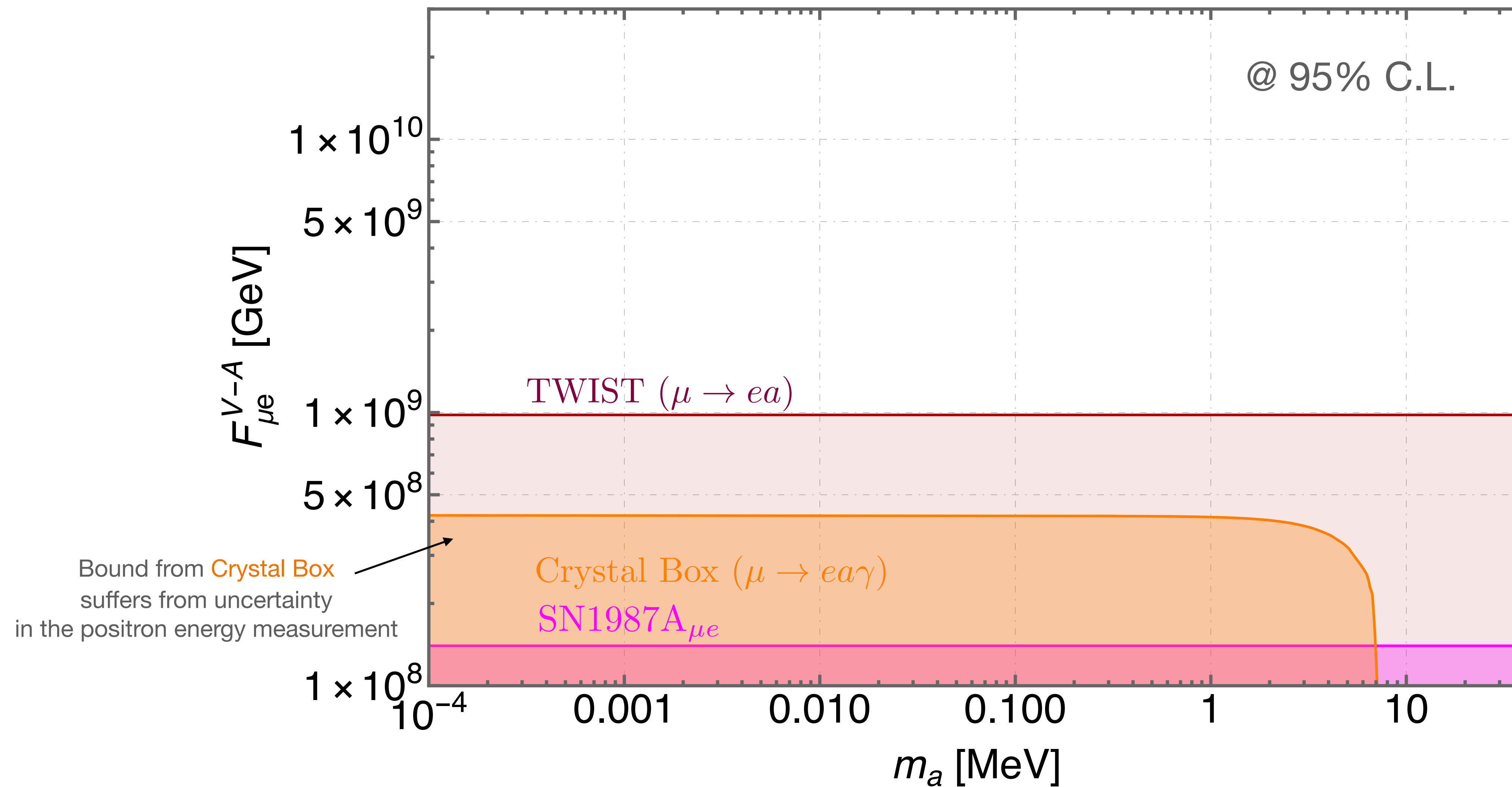
Jodidio et al. (1986)

Can it be done again?

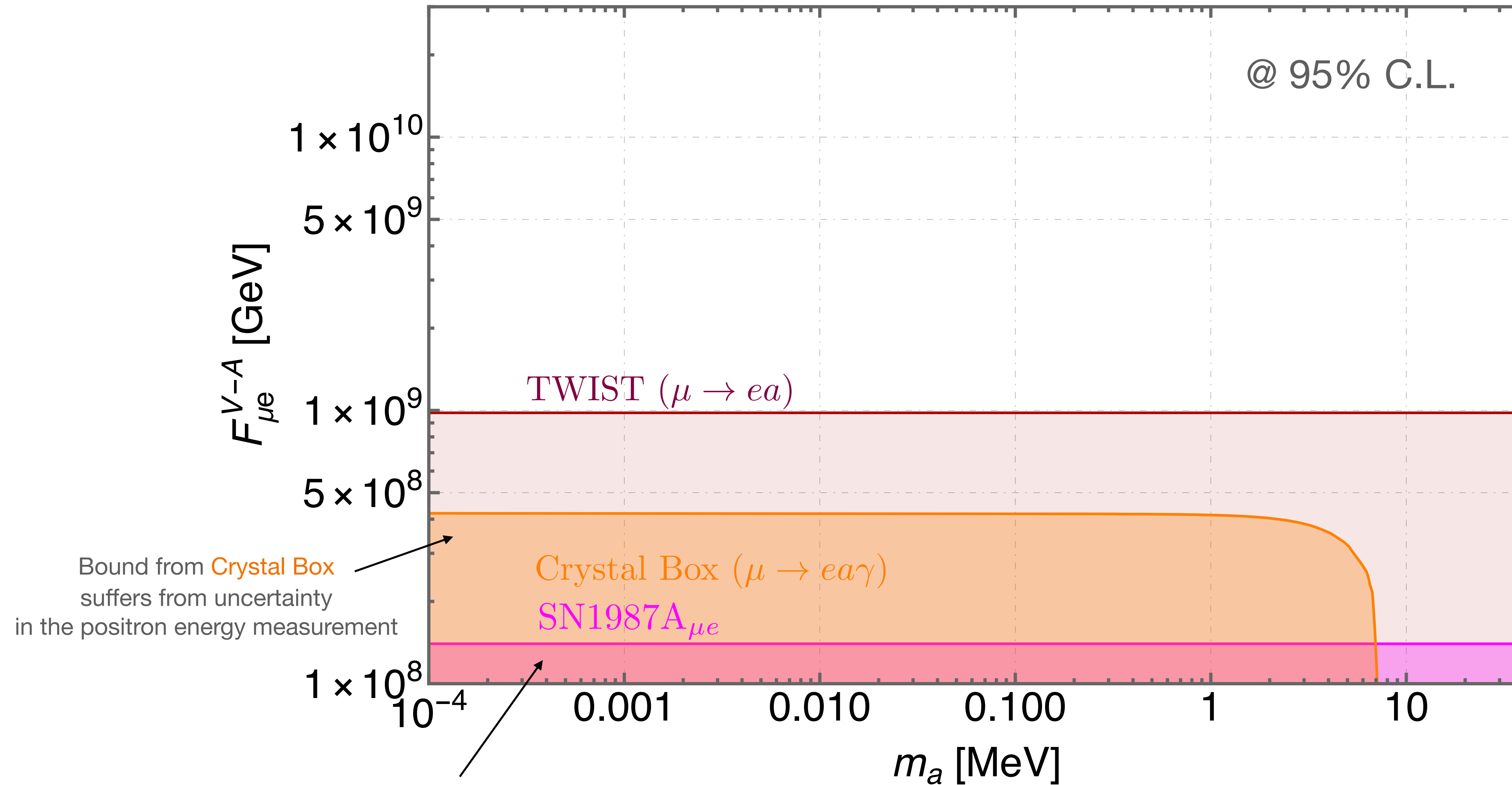
STATUS for left-handed LFV ALP



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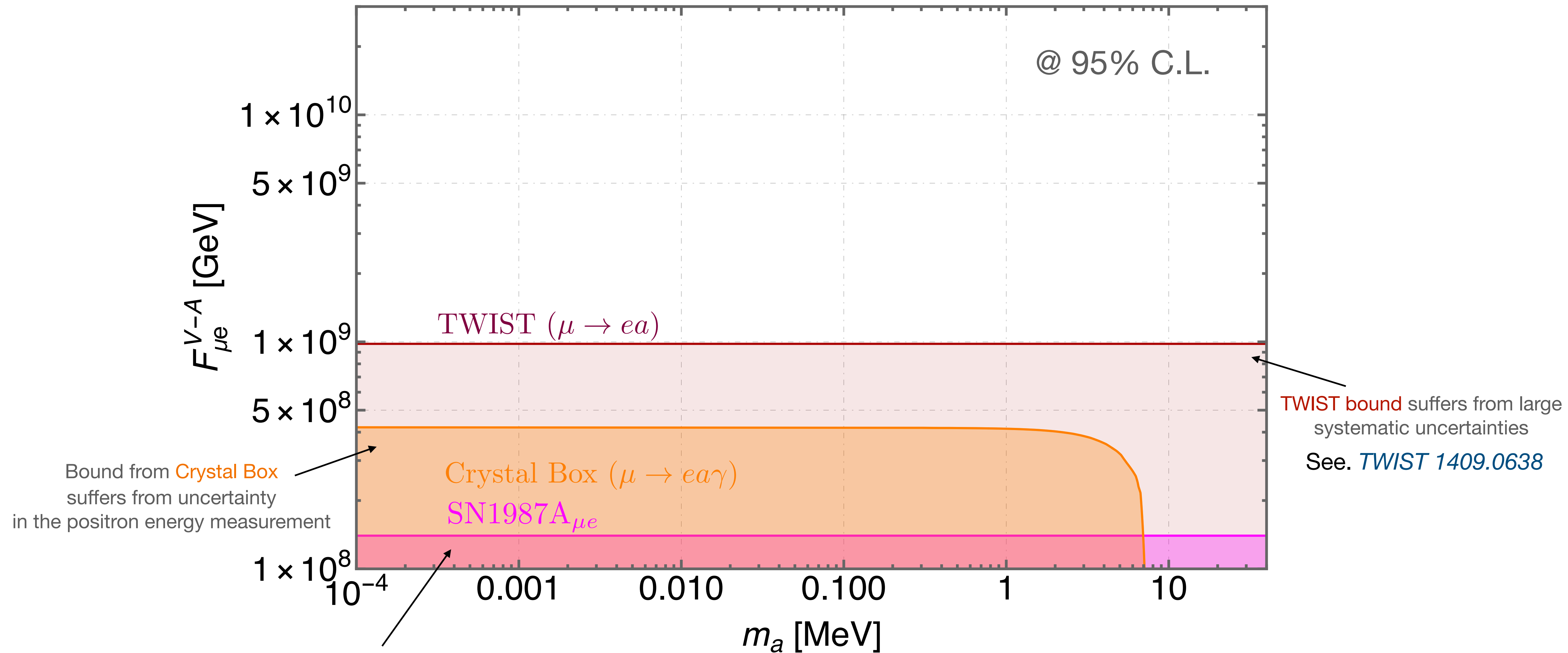


STATUS for left-handed LFV ALP



It is remarkable that present experiments are already stronger than cooling bounds!

STATUS for left-handed LFV ALP



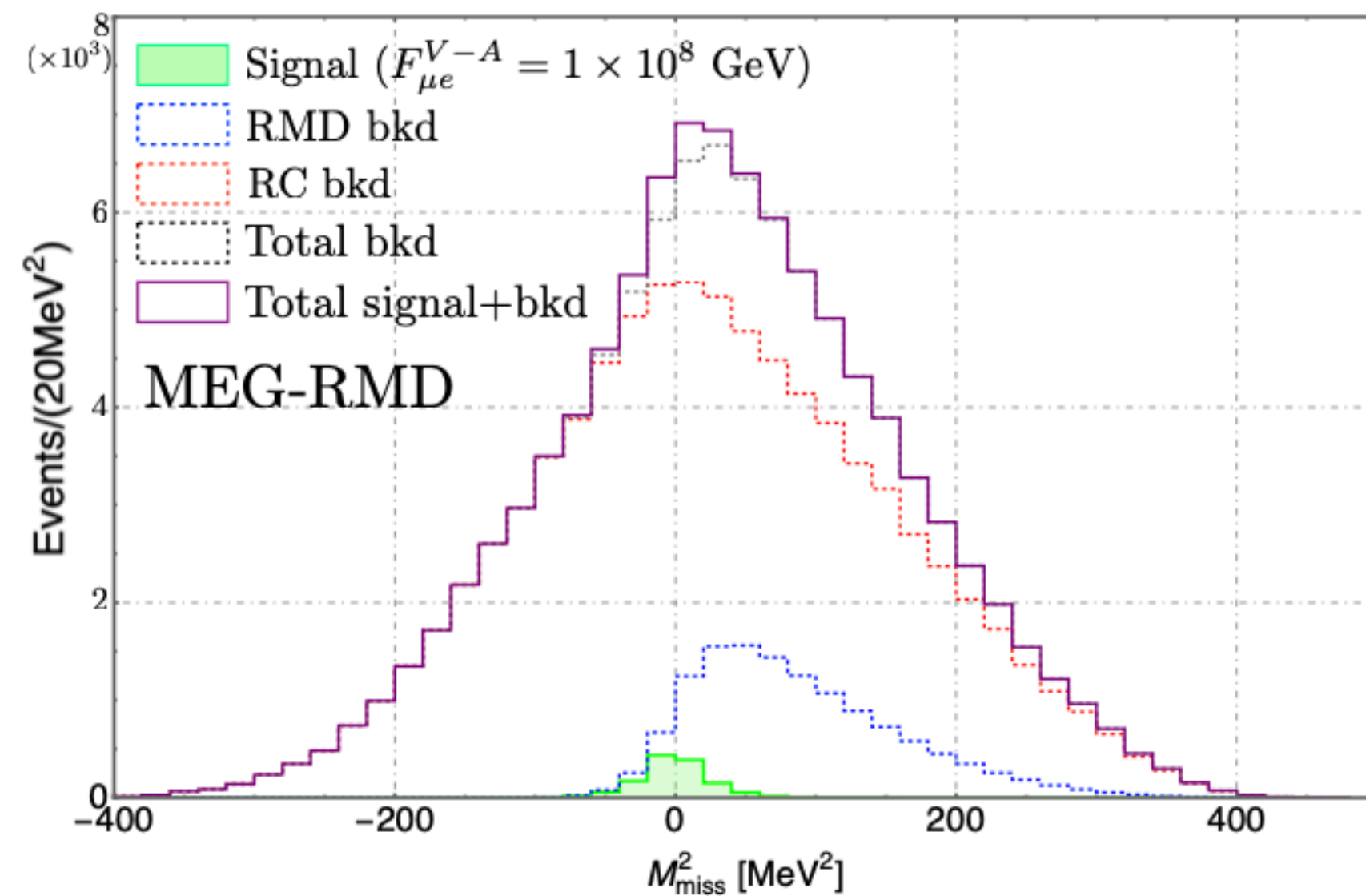
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What **MEG** can teach us about LFV axions?

MEG-RMD search [1312.3217](#) $N_\mu = 1.8 \times 10^{14}$ collected in 2009-2010

The goal was to observe the RMD $\mu \rightarrow e\nu\bar{\nu}\gamma$

over the bkd of random coincidences (RC): pileup of $\mu \rightarrow e\nu\bar{\nu}\gamma + \mu \rightarrow e\nu\bar{\nu}$

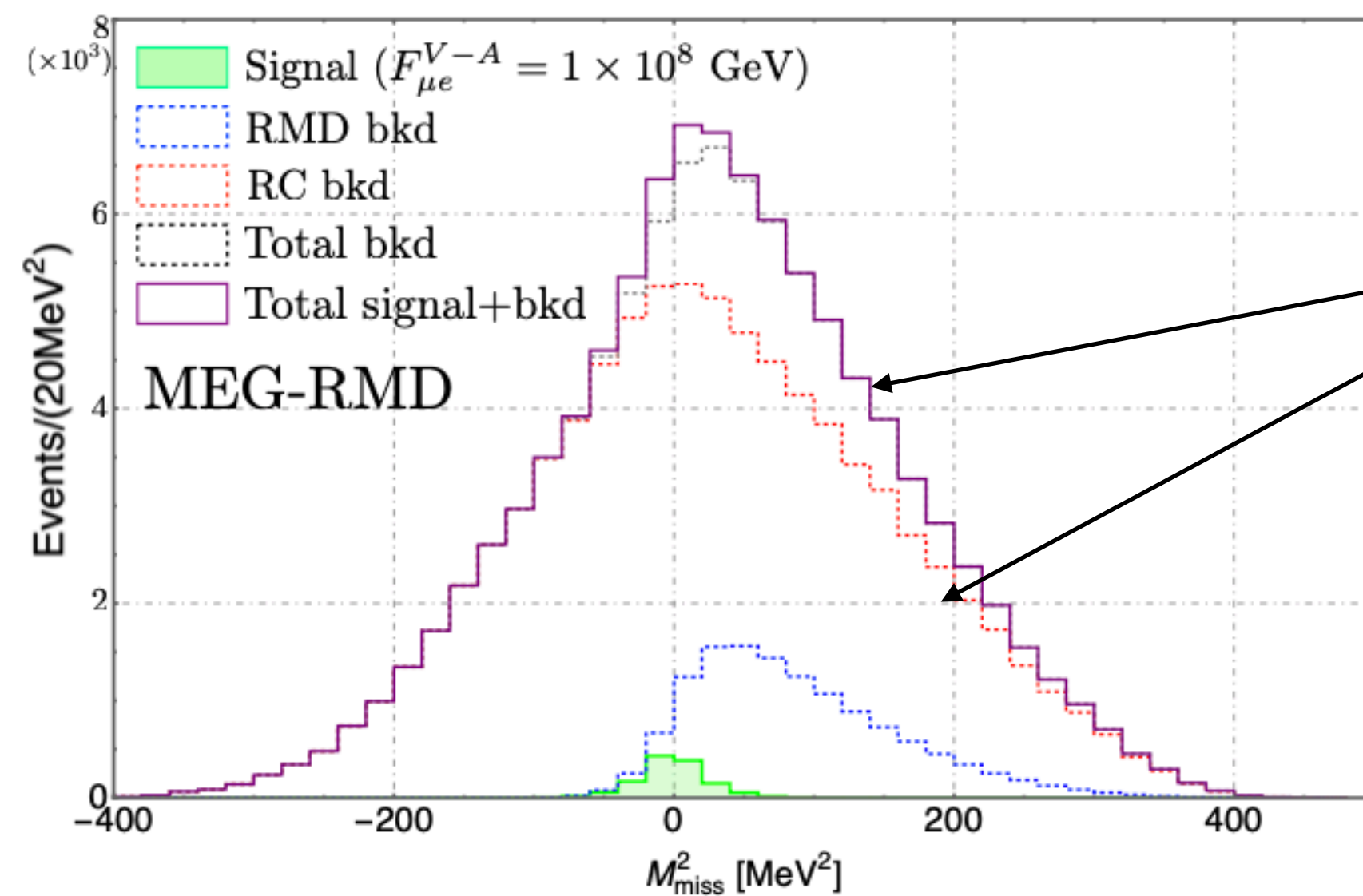


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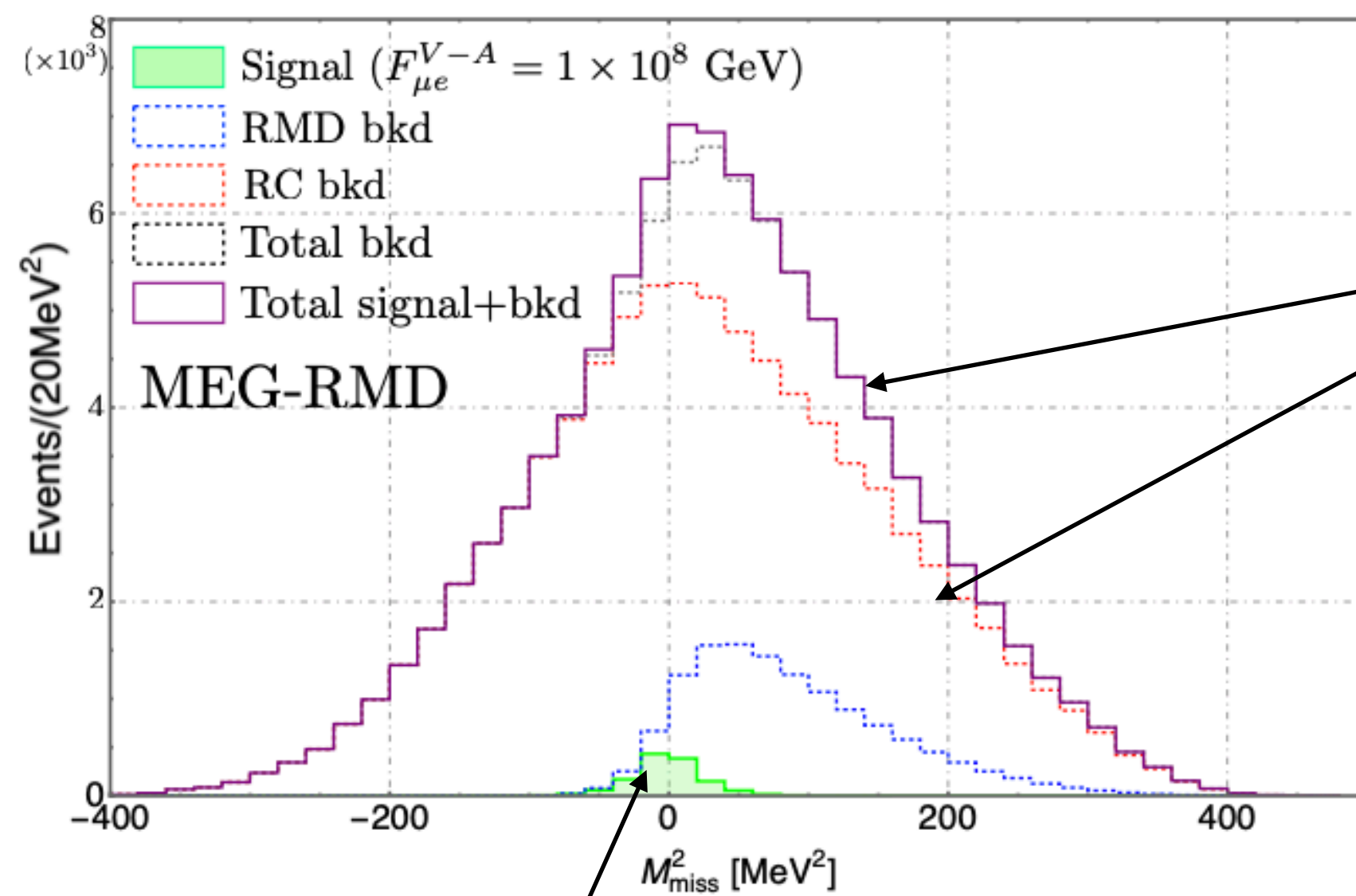
missing mass distribution of the bkd
obtained with an home brew MC normalized
to the offline number of RMD and RC events
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Signal with a rate $\mu \rightarrow ea\gamma$ with mu polarisation See. [MEG 1510.04743](#) with mu polarisation

We account for detector smearing

$$\langle P_\mu \rangle = -0.856 \pm 0.021$$

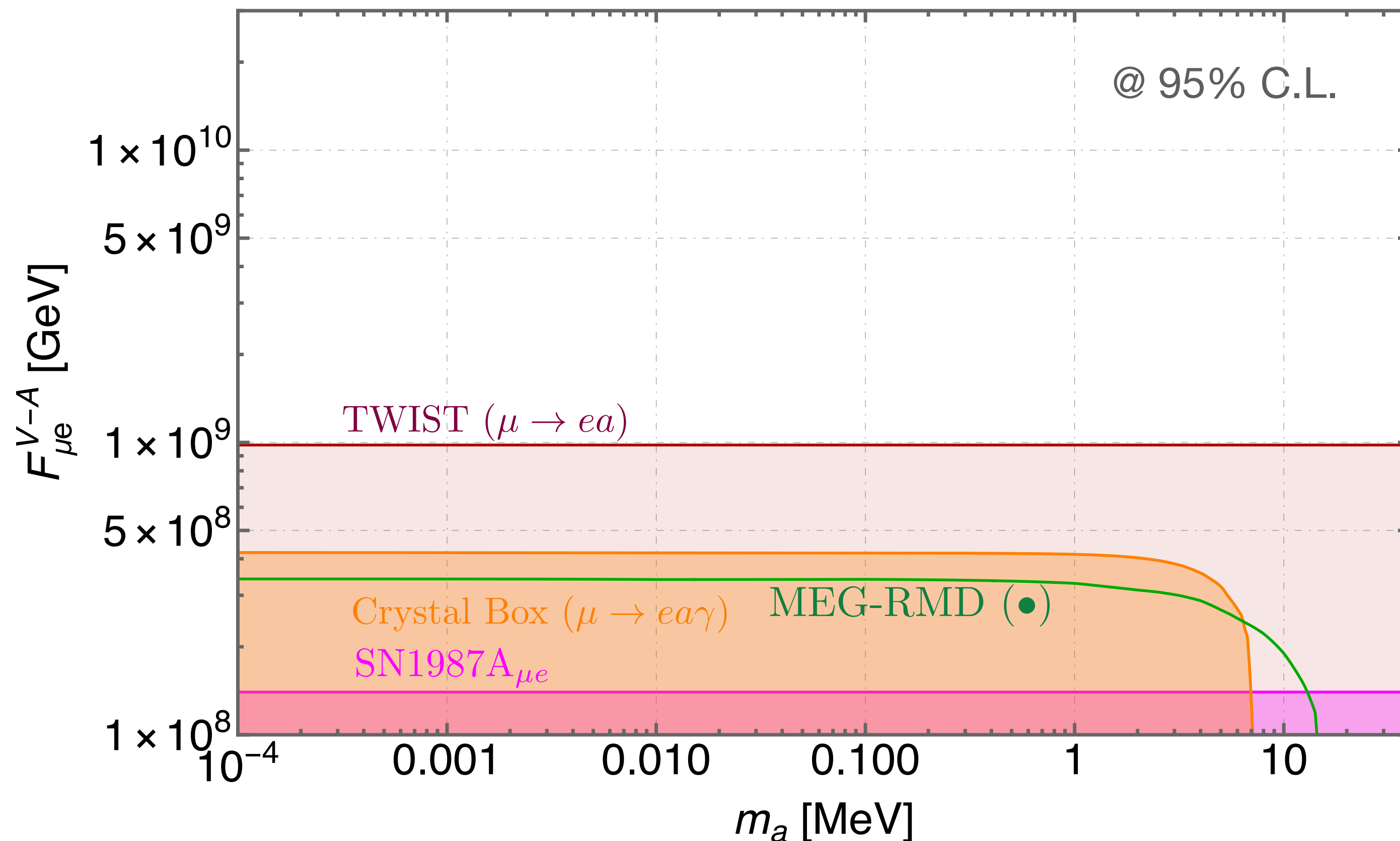
See. [MEG 2005.00339](#)

What **MEG** can teach us about LFV axions?

with $N_\mu = 1.8 \times 10^{14}$ collected in 2009-2010

The topology is already close to back to back at trigger level

Very low signal efficiency for $\mu \rightarrow ea\gamma$ limits the MEG reach

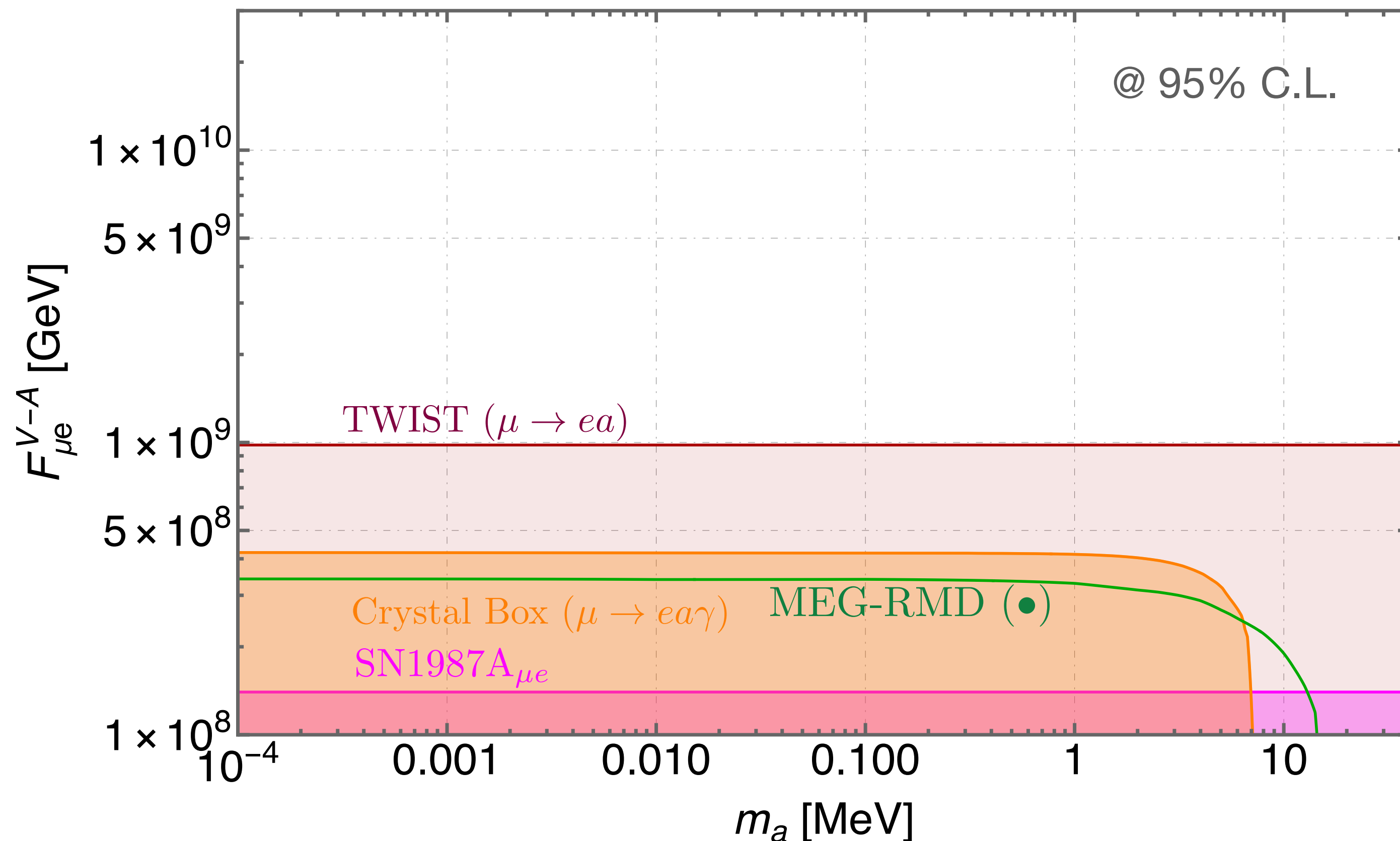


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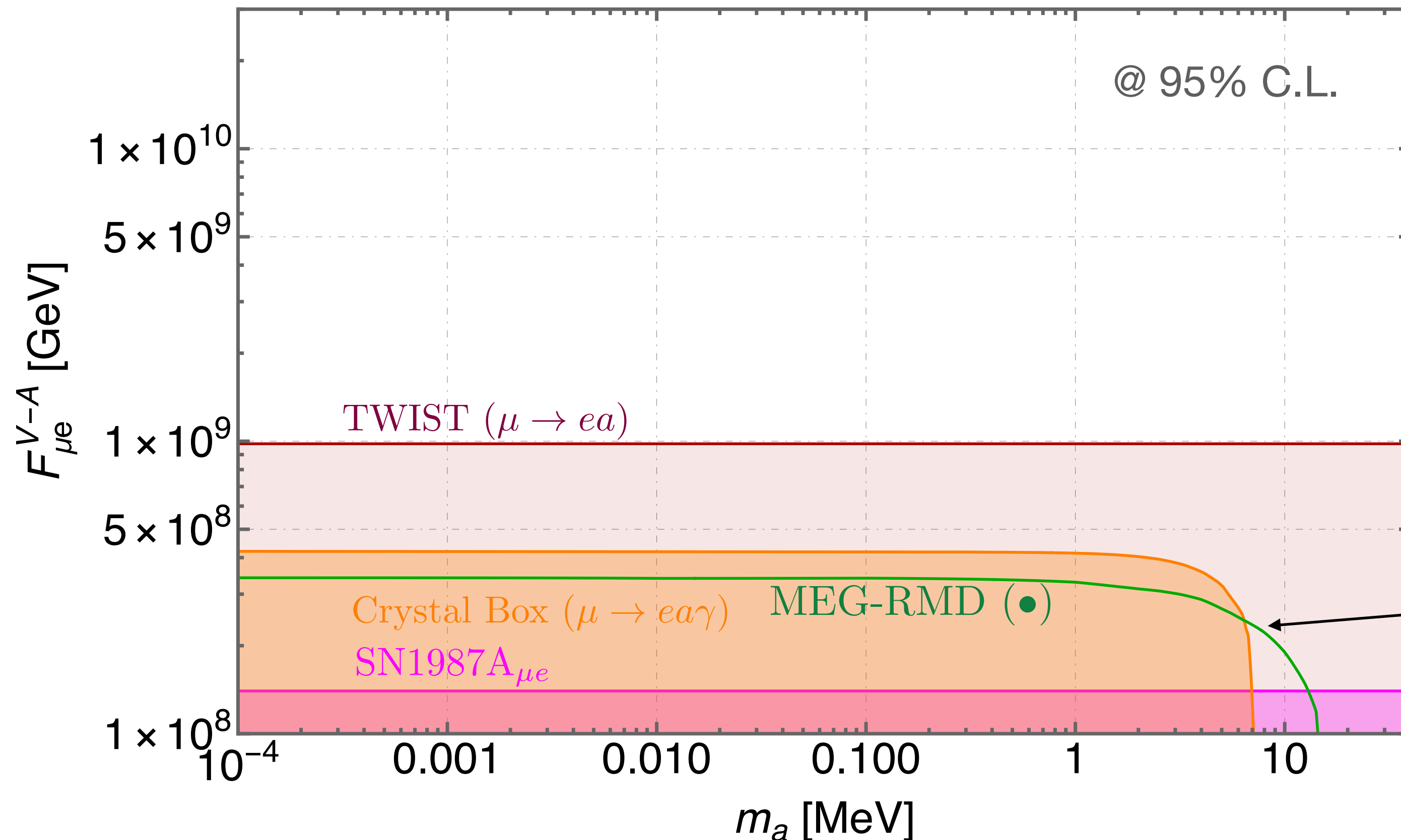
| | BR_i^{base} | $\epsilon_i^{\text{trig.}}$ | | |
|------------------|-----------------------|-----------------------------|-----------------------|-------------------------------|
| | | ϵ_{E_e} | ϵ_{E_γ} | $\epsilon_{\theta_{e\gamma}}$ |
| B_{RMD} | 1.44×10^{-5} | 0.15 | 5.3×10^{-4} | 0.49 |
| B_{RC} | 7.08×10^{-4} | 0.34 | 0.01 | 0.03 |

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3×10^{-4}

MEG RMD data expected sensitivity is already comparable to **Crystal Box**!

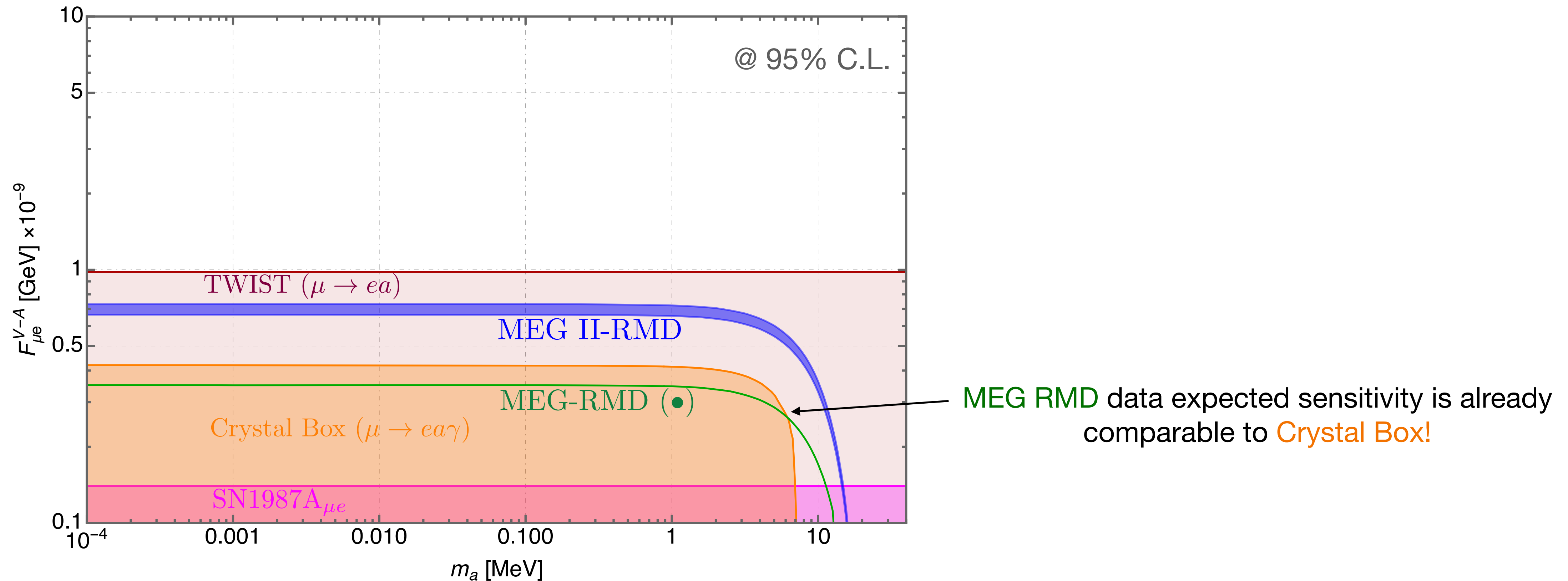
What about **MEG II**?

Keeping the same analysis as the MEG-RMD one but accounting for:

1) larger luminosity $N_\mu = 1.8 \times 10^{15} \mu^+$

2) Improved energy and angular resolution See [MEG II 1801.04688](#)

3) Reduced RC background by 50% after installation of radiative decay counter (RDC)



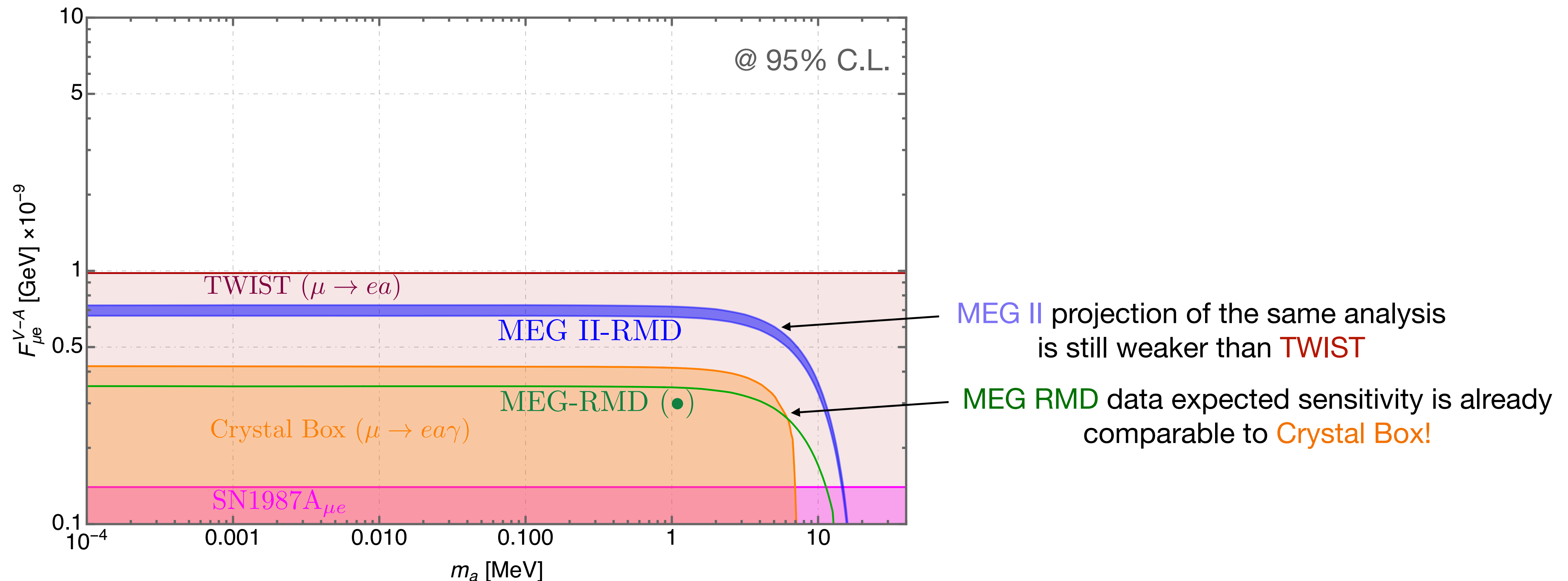
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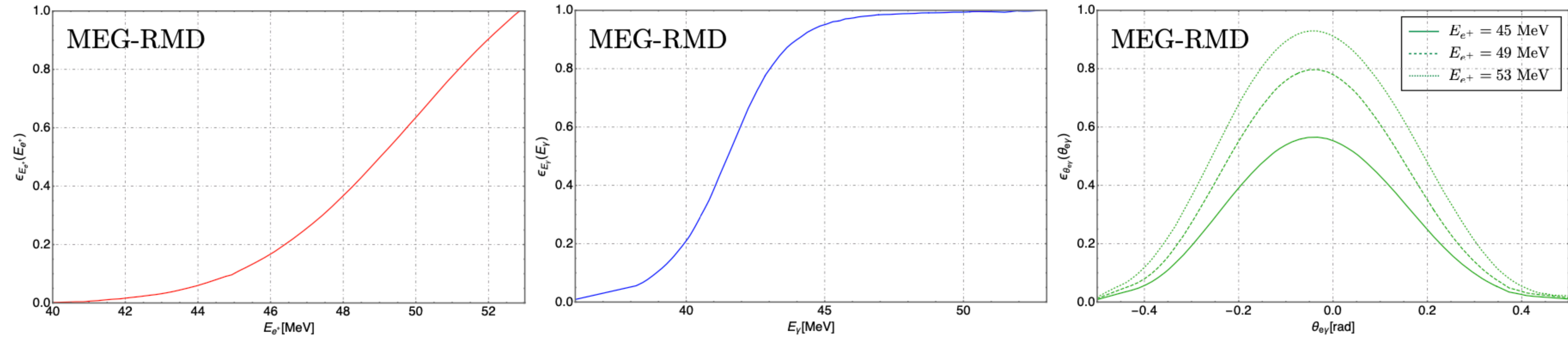
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Towards a new data taking strategy

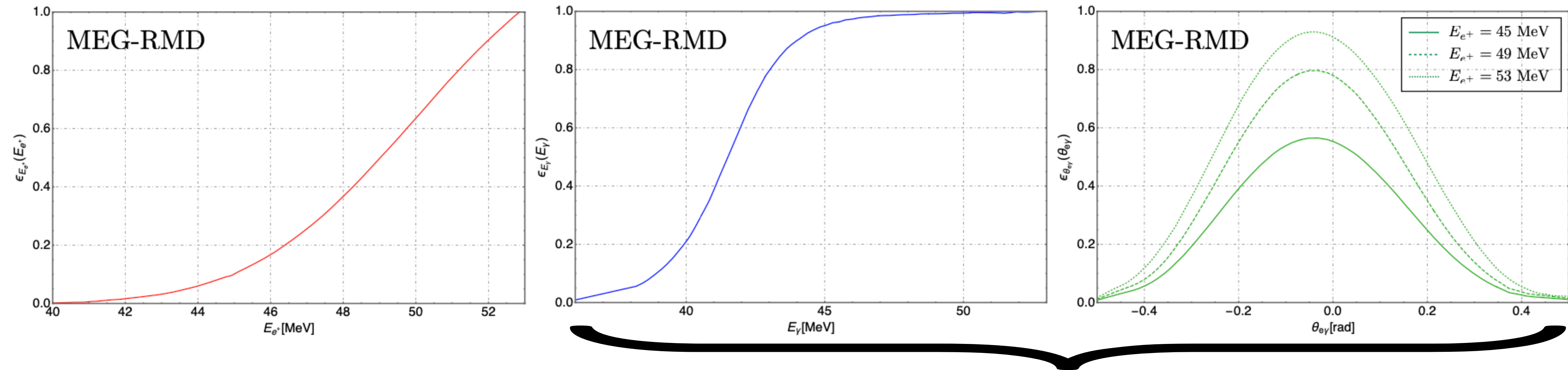
Logic: the trigger requirements are killing the ALP signal



*

Towards a new data taking strategy

Logic: the trigger requirements are killing the ALP signal

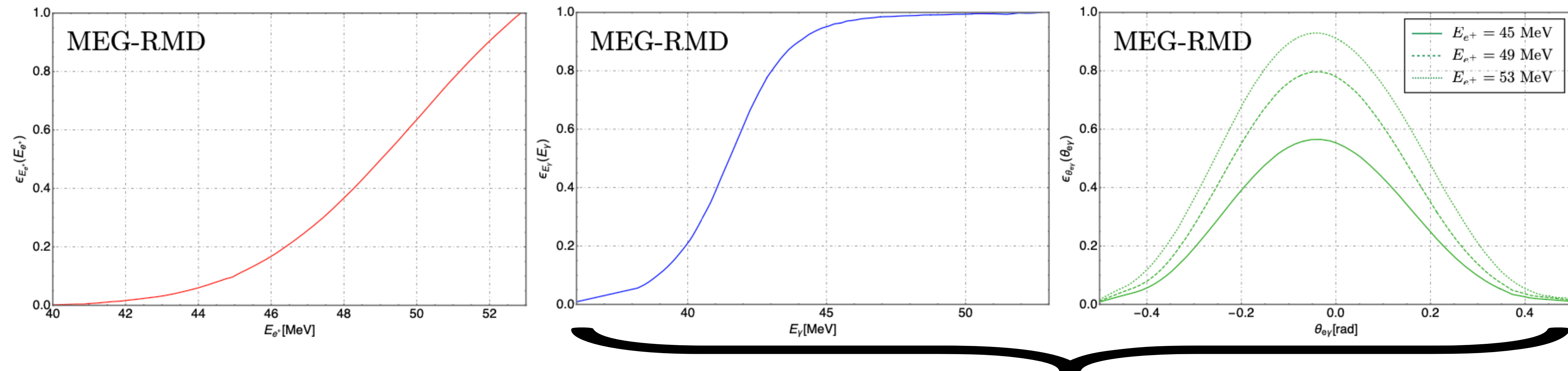


- 1) Eliminating the matching of the TC hit which assumes back to back topology
- 2) Lowering the photon trigger threshold reducing the beam intensity

*

Towards a new data taking strategy

Logic: the trigger requirements are killing the ALP signal



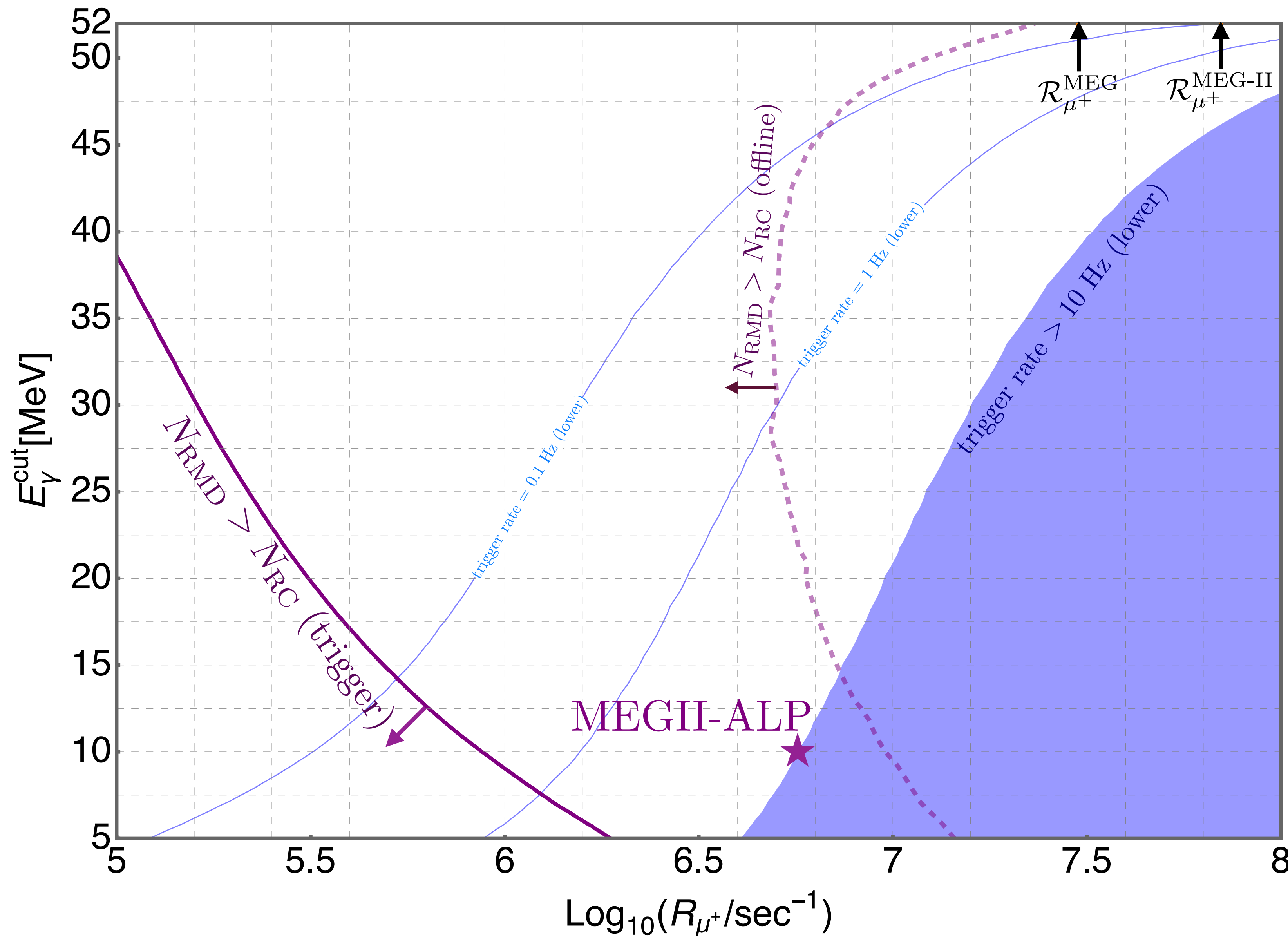
- 1) Eliminating the matching of the TC hit which assumes back to back topology
- 2) Lowering the photon trigger threshold reducing the beam intensity

The RC dominates the trigger rate but it can be suppressed by reducing the intensity R_μ^*

$$\text{RC} \sim R_\mu^2 \quad \text{RMD} \sim R_\mu$$

*many thanks to Luca for teaching us all this!

Towards a new data taking strategy



Max trigger rate 10 Hz

fixes the intensity vs photon cut

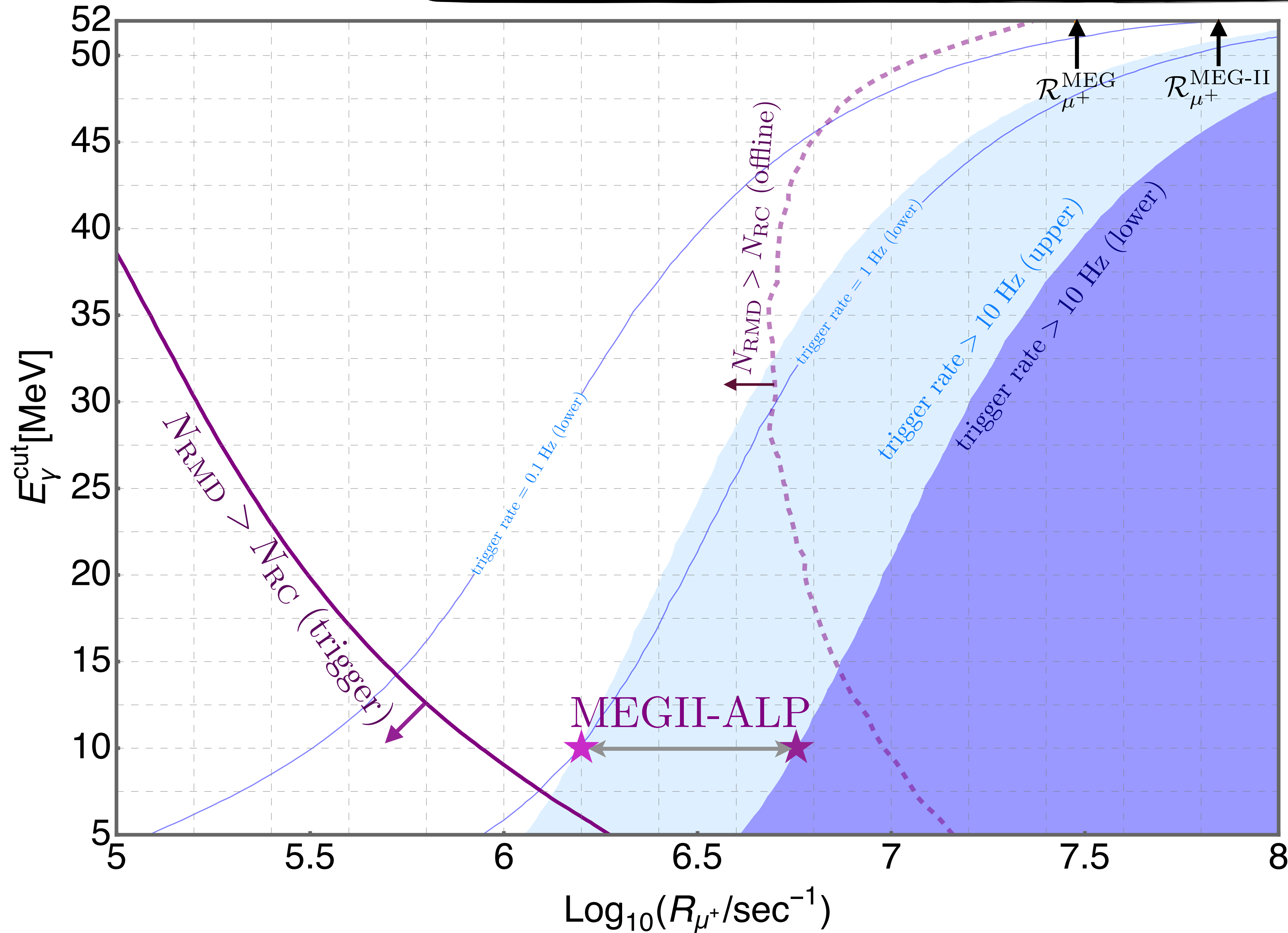
RMD becomes the dominant bed

below a certain intensity

(harder to suppress RMD online)

★ Benchmark fixed to the highest intensity for photon energy of 10 MeV given our estimate of the trigger rate

Towards a new data taking strategy II



Max trigger rate 10 Hz

fixes the intensity vs photon cut

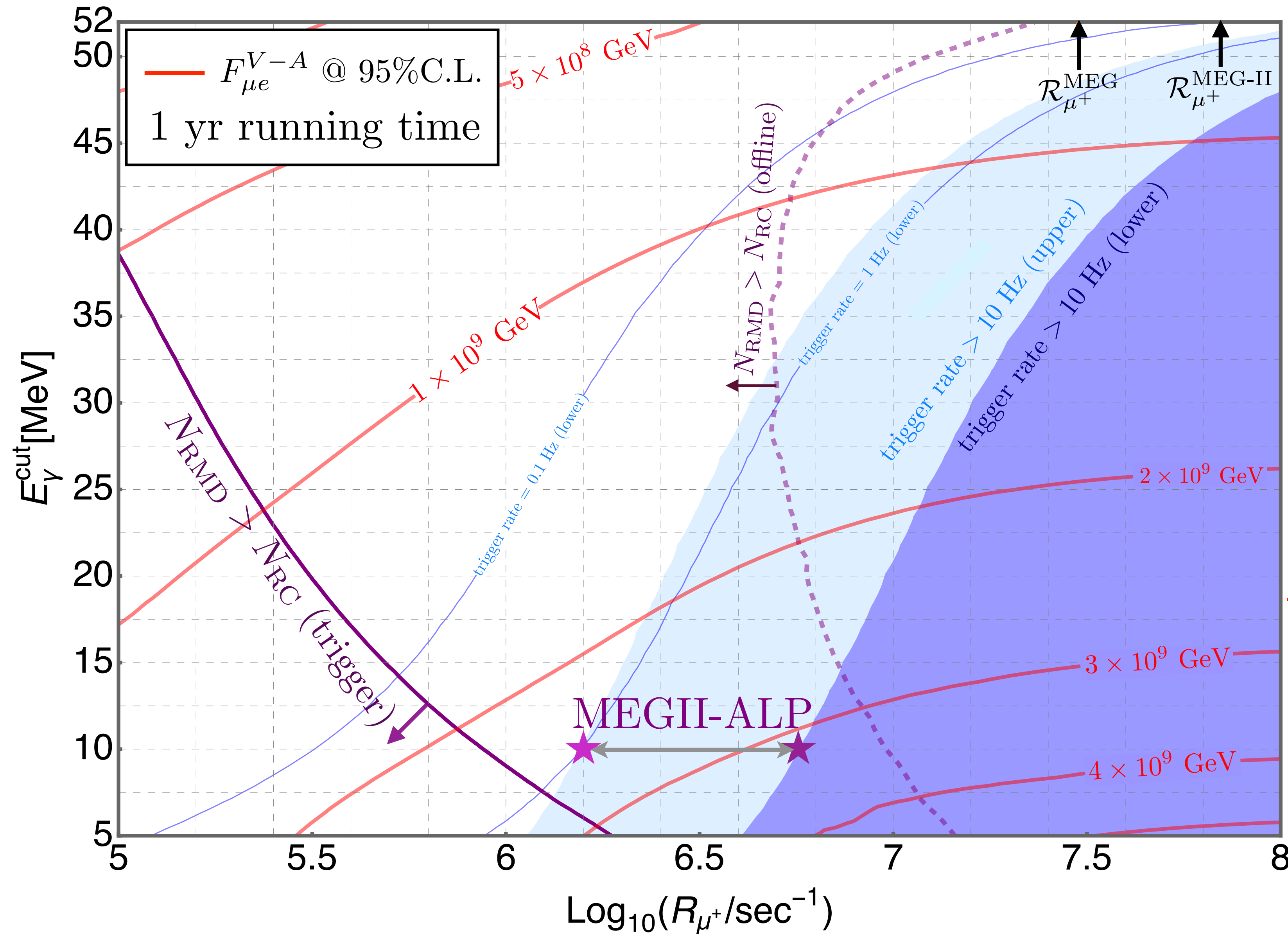
RMD becomes the dominant bkd

below a certain intensity

(harder to suppress RMD online)

★ Uncertainty in trigger rate results in two different benchmark for the same photon energy

Towards a new data taking strategy III



Max trigger rate 10 Hz

fixes the intensity vs photon cut

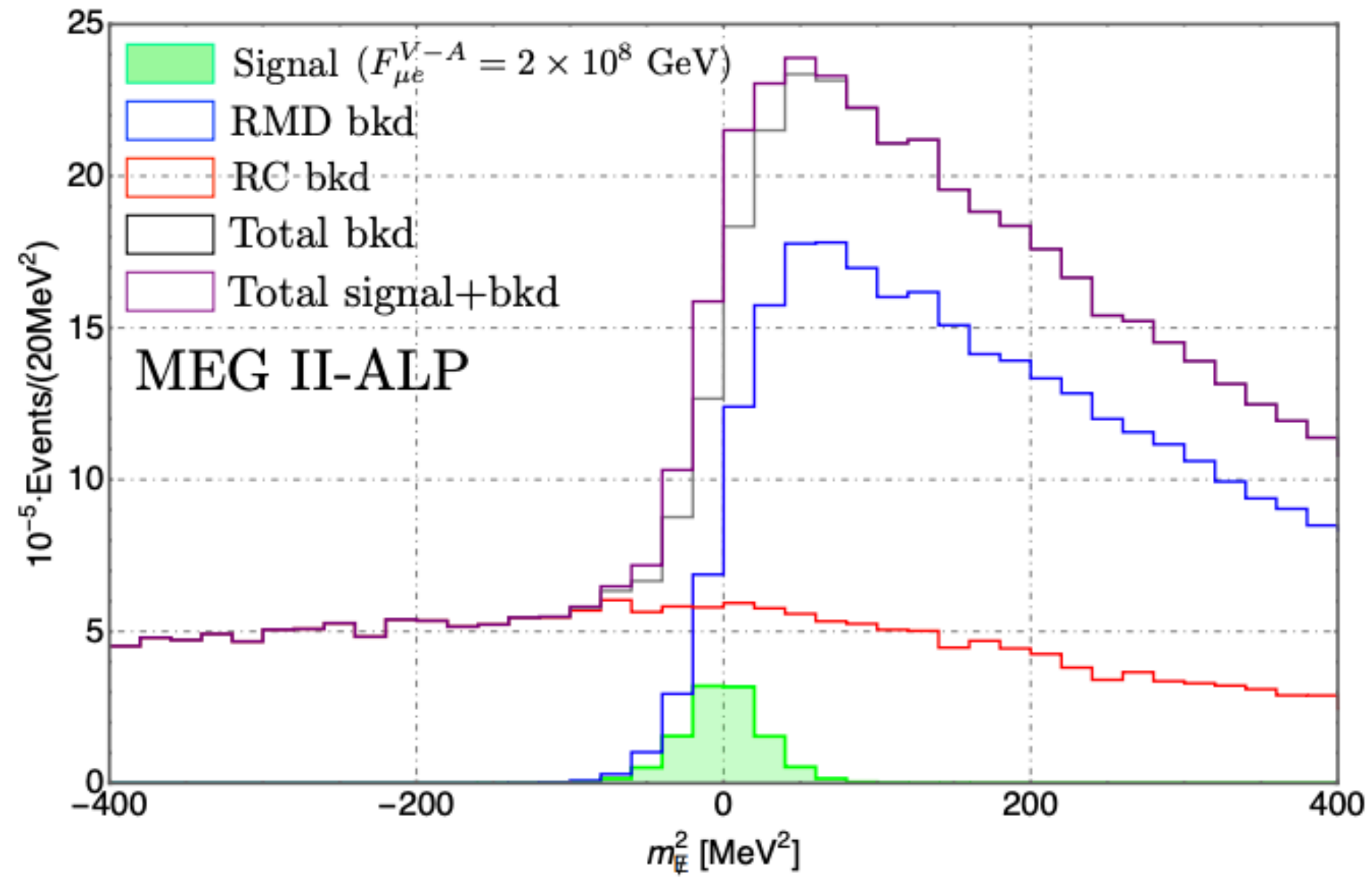
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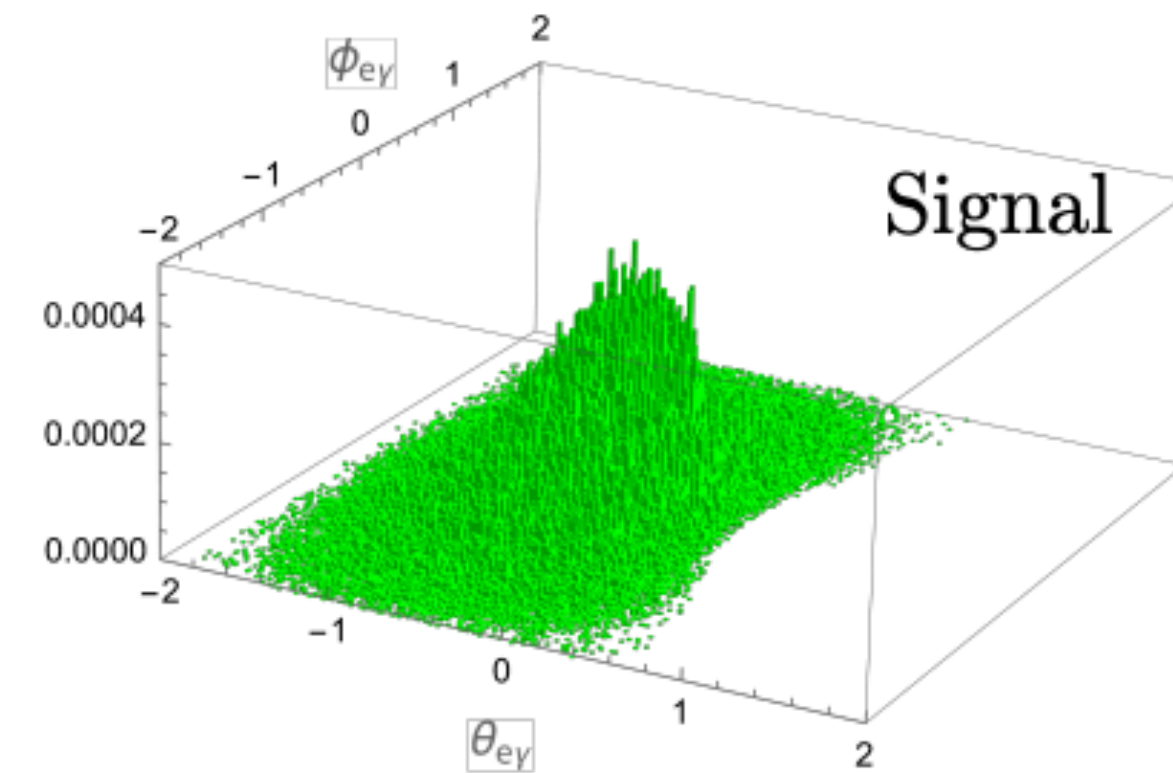
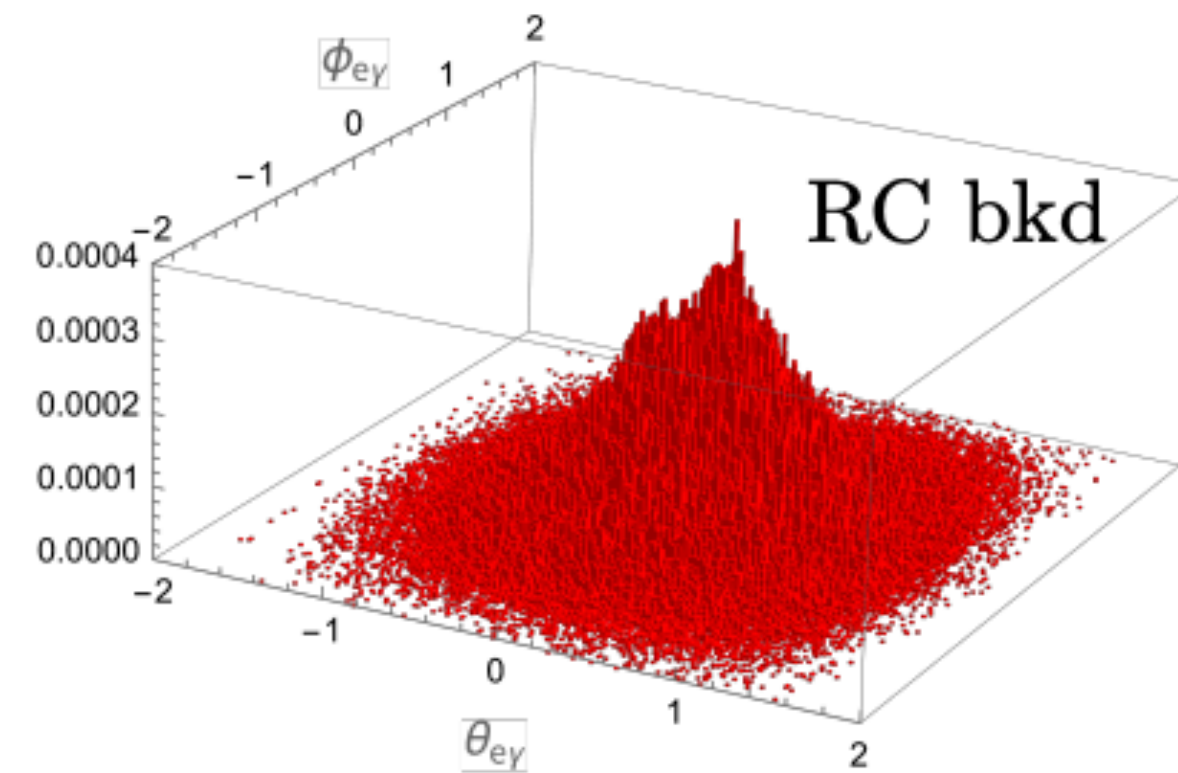
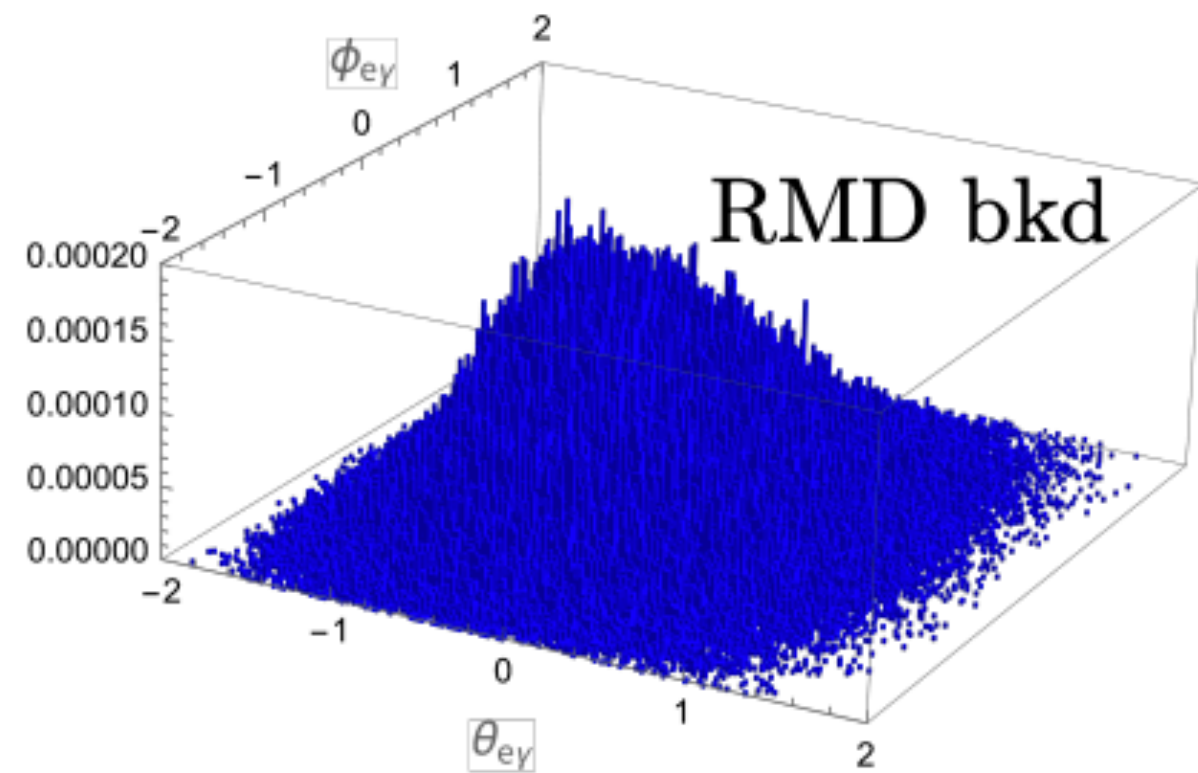
The reach is extracted at each point!

Final reach



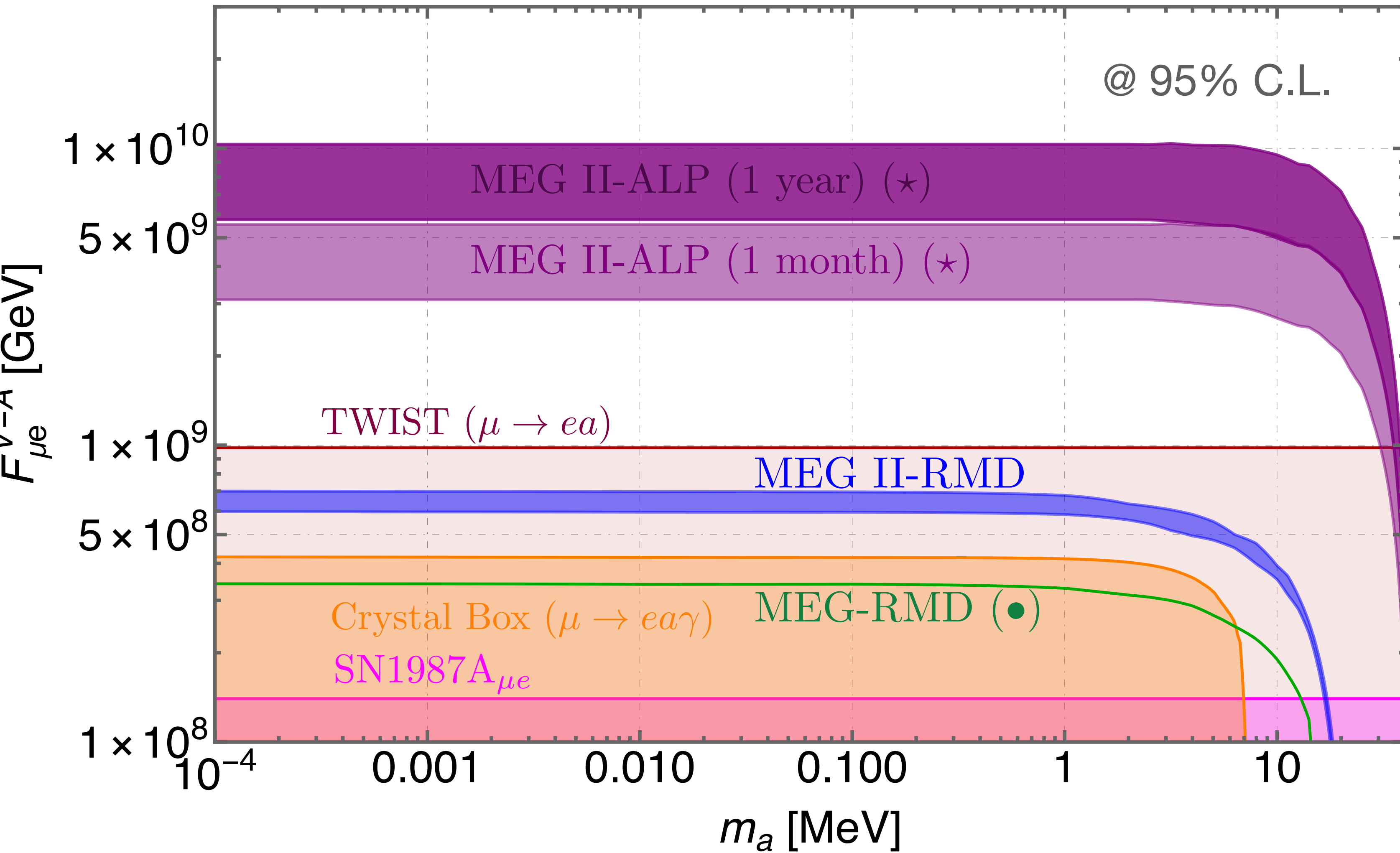
Bump hunt in missing mass*

*for a massless object we are close to a cliff of the bkd (systematics has to be taken into account)

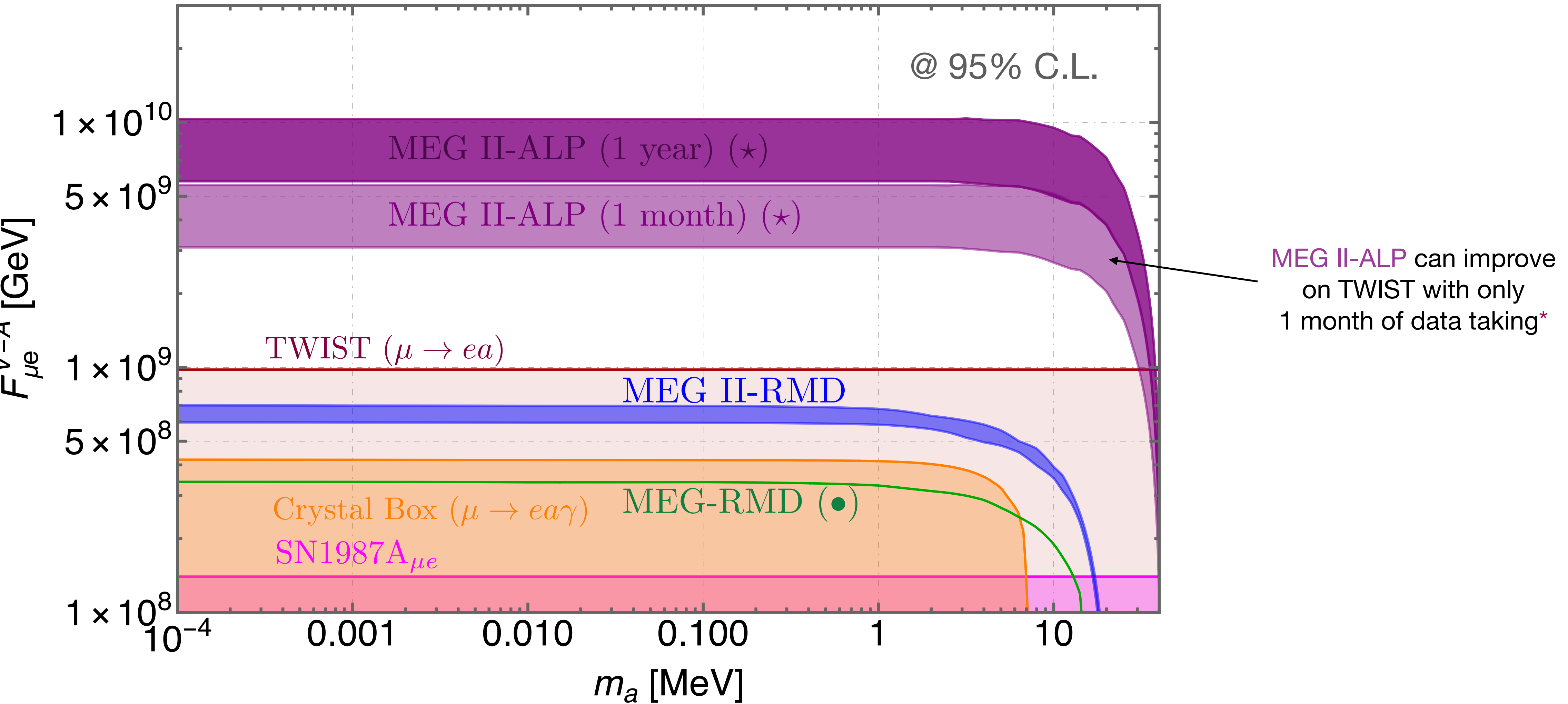


Log-likelihood on angular variables

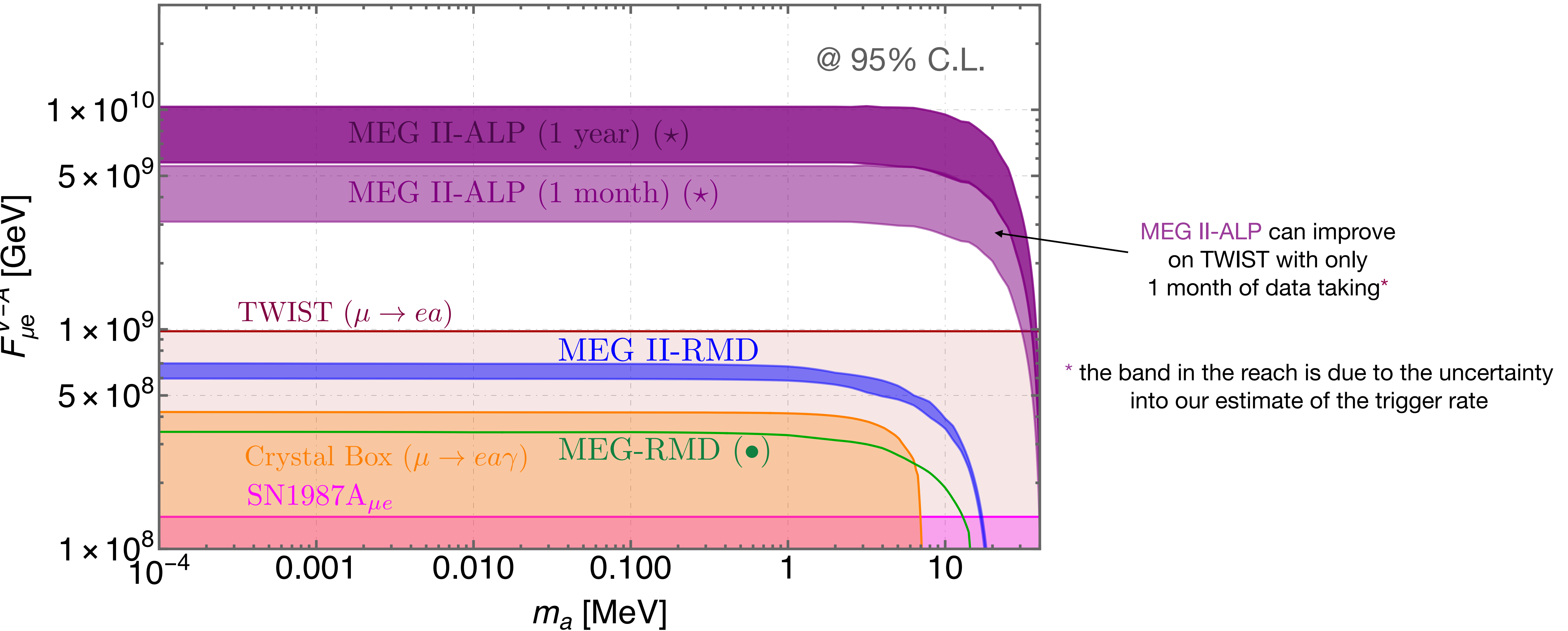
What can we test?



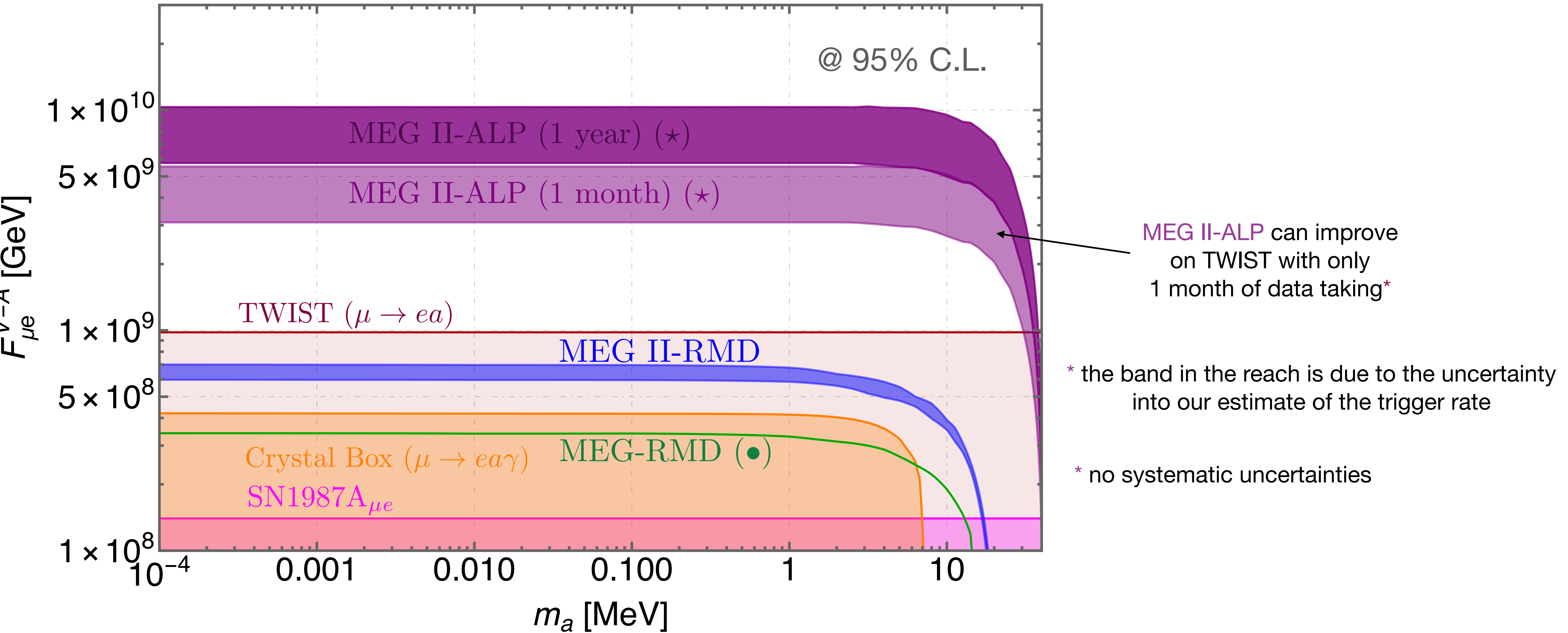
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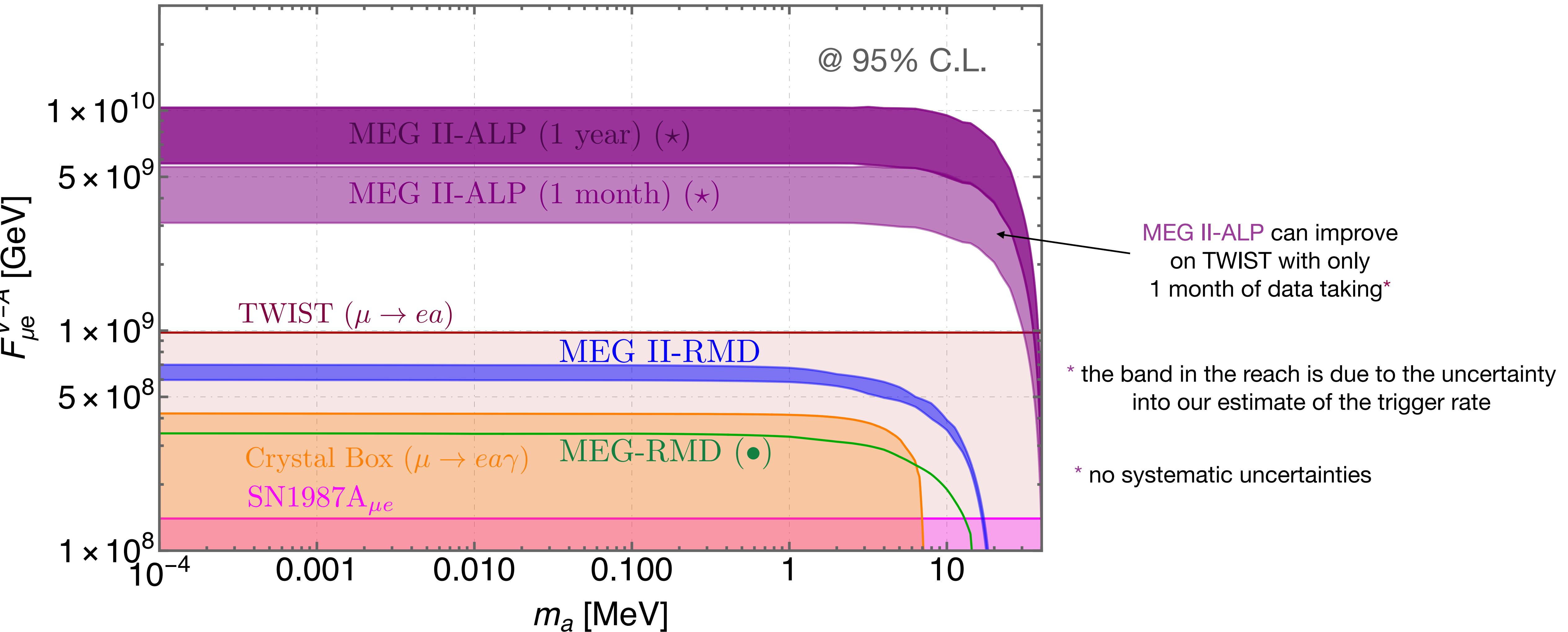
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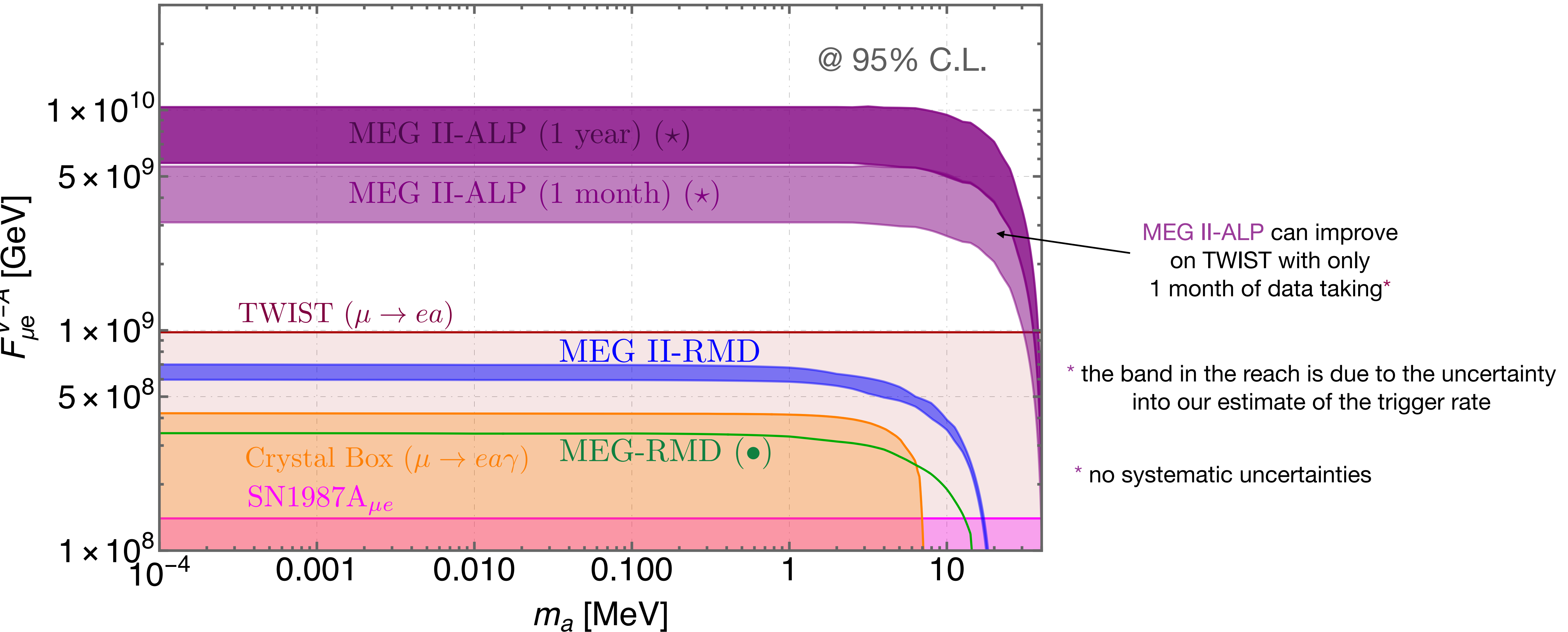


What can we test?



Preliminary results of the MEG-II collaboration show that our estimate on the trigger performances has been conservative!

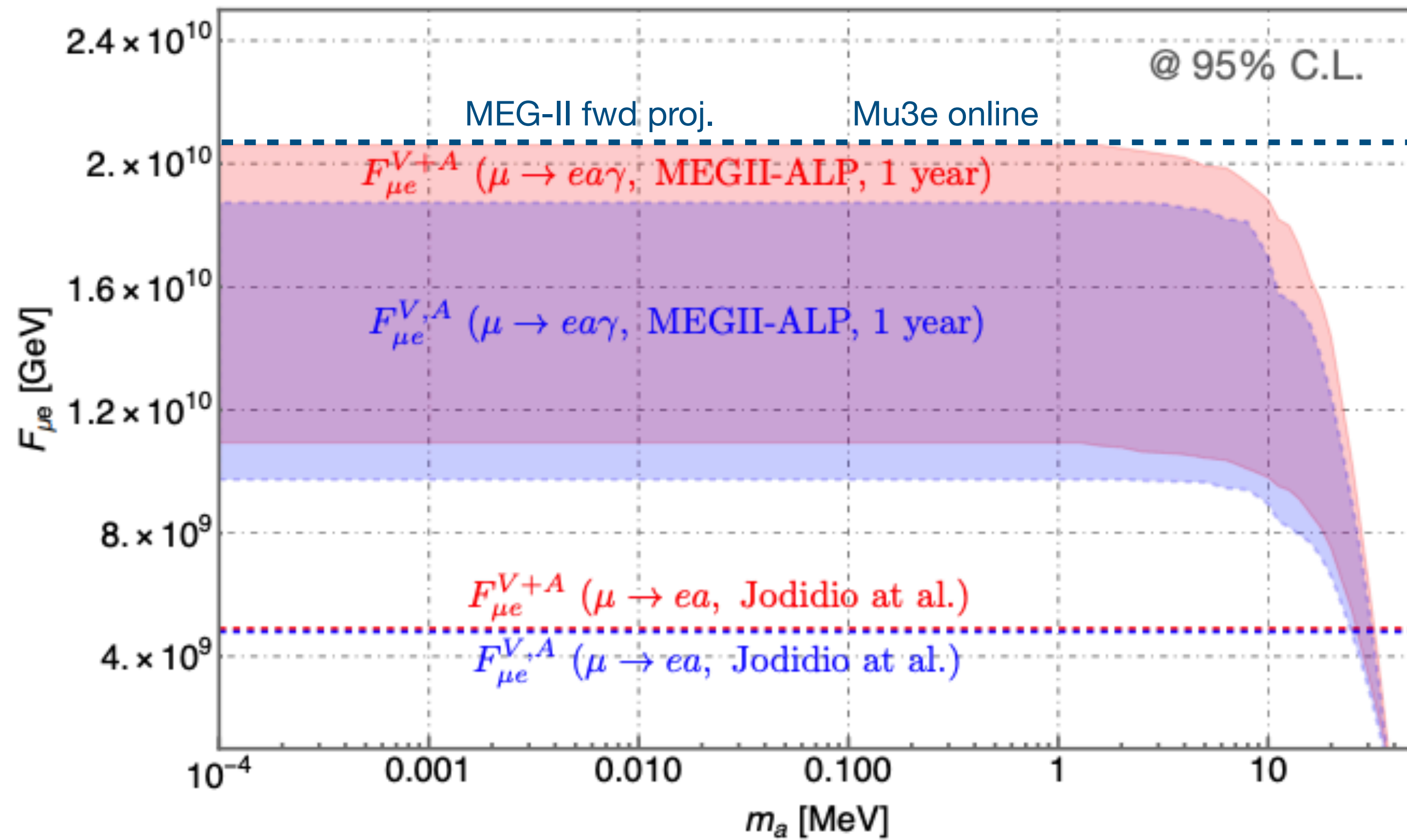
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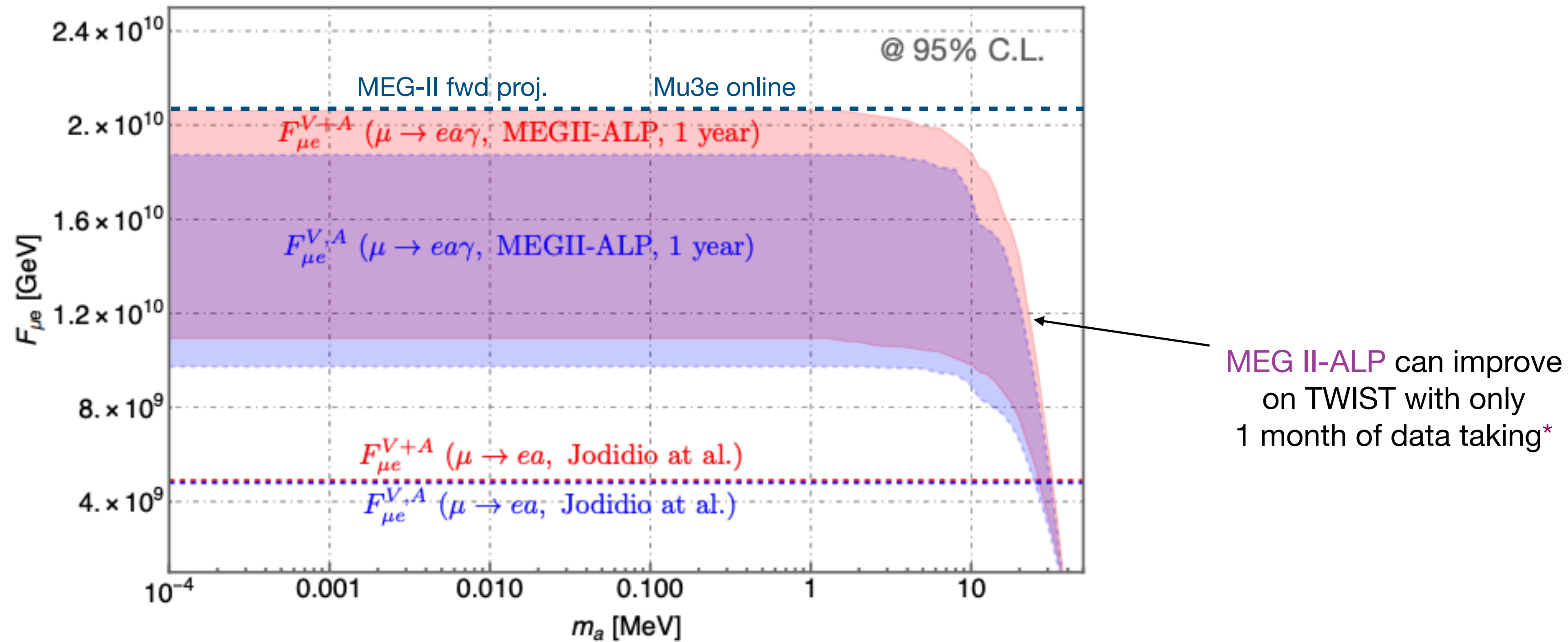
(Luca Galli for more infos)

Different ALP chiral structures



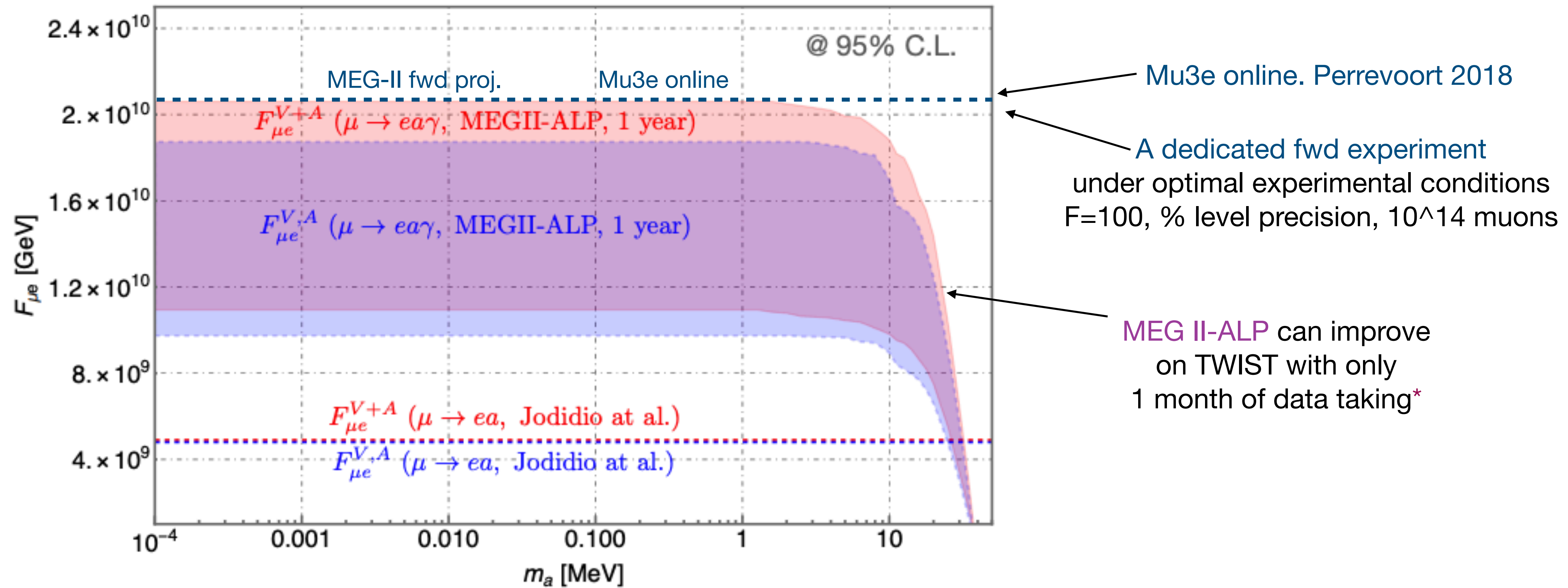
MEG-II ALP is the best way to explore ~ a decade of unconstrained parameter space ~ NOW

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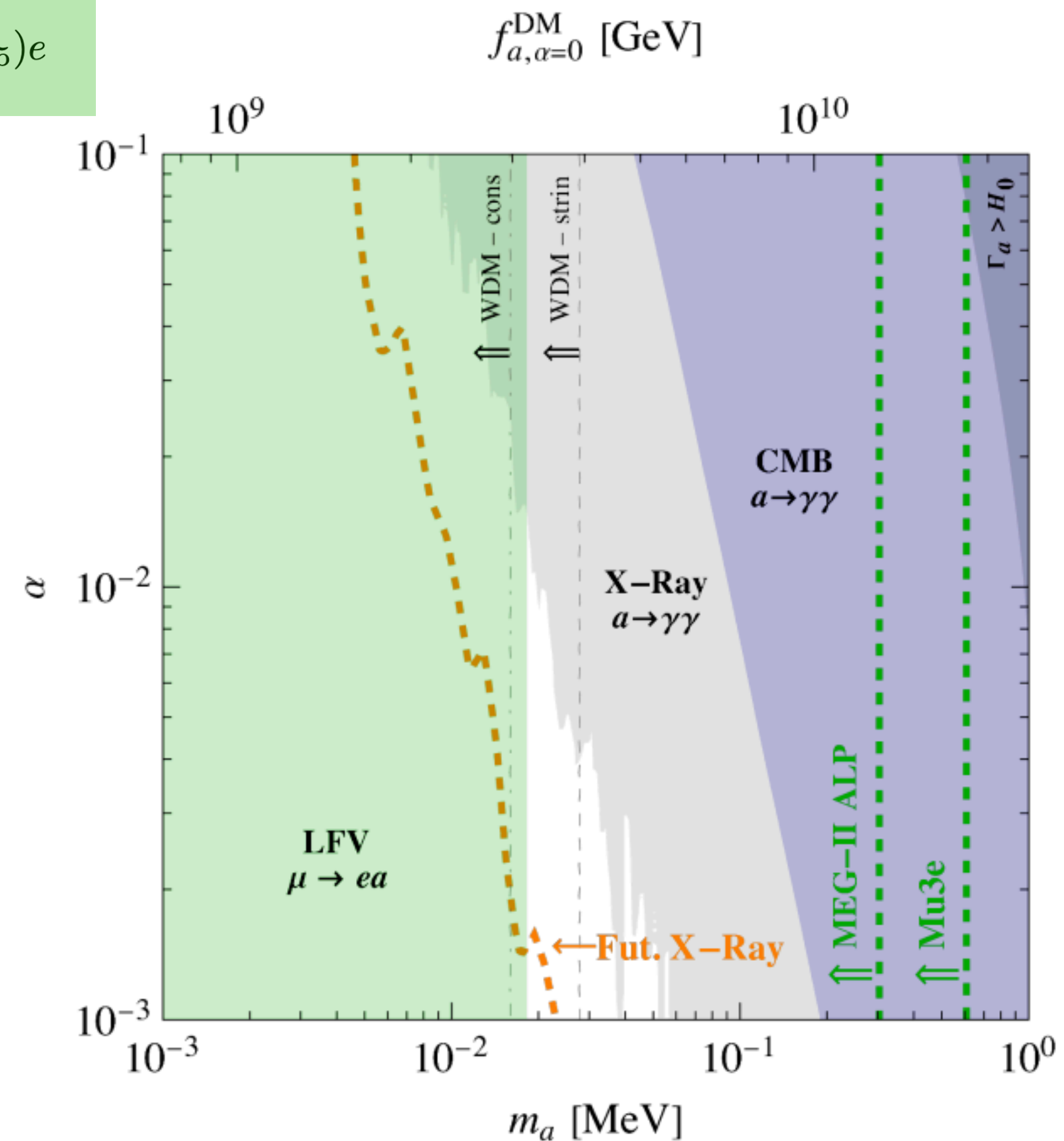
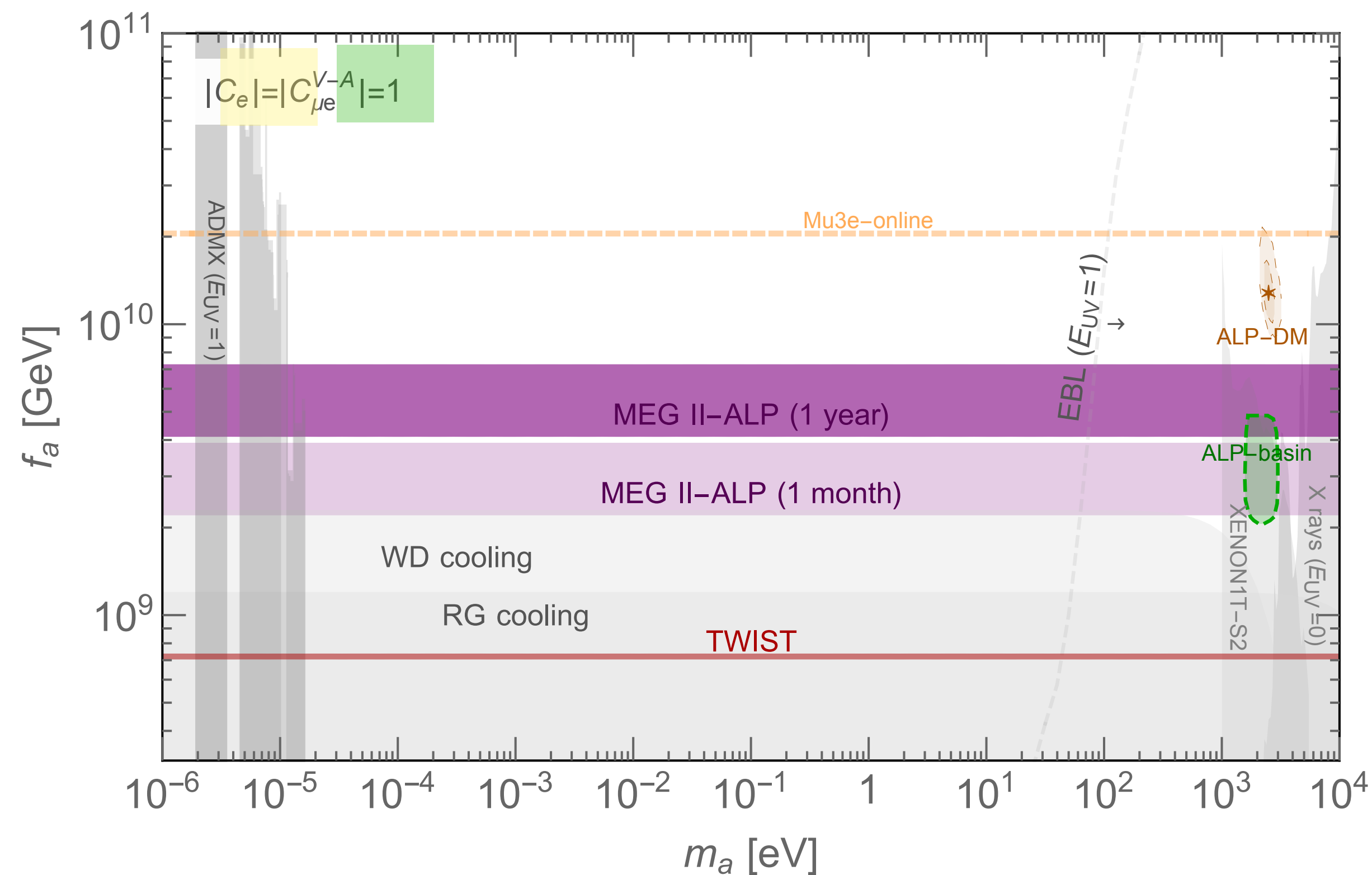
Back to theory land

axions coupled to leptons anarchically: *flavor diagonal* = *flavor off-diagonal*

$$\frac{\partial_\mu a}{f_a} \bar{e} \gamma^\mu \gamma_5 e$$

$$\frac{\partial_\mu a}{2f_a} \bar{\mu} \gamma^\mu (C_{\mu e}^V + C_{\mu e}^A \gamma_5) e$$

Panci, Redigolo, Schwetz Ziegler 2209.03371



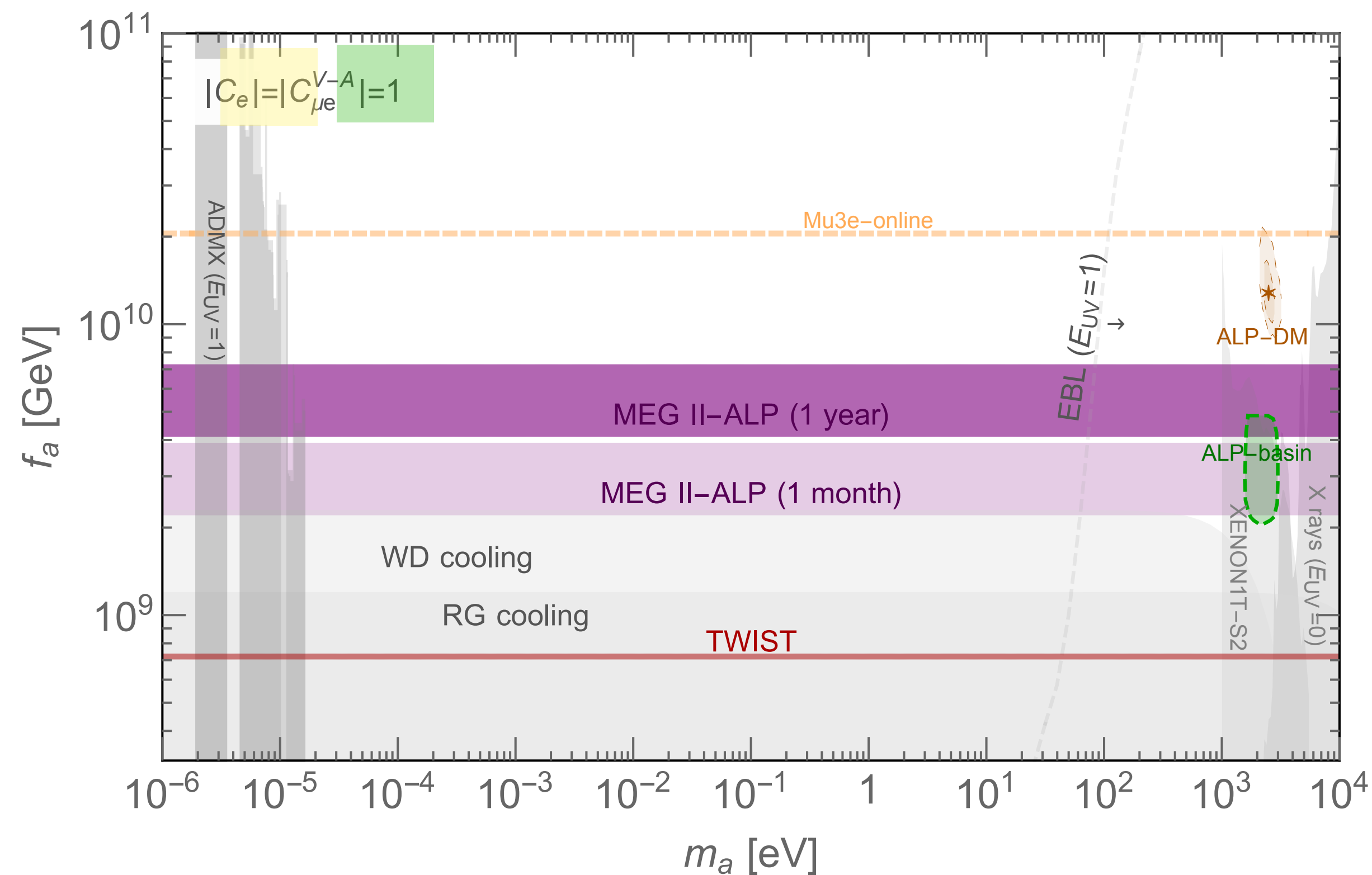
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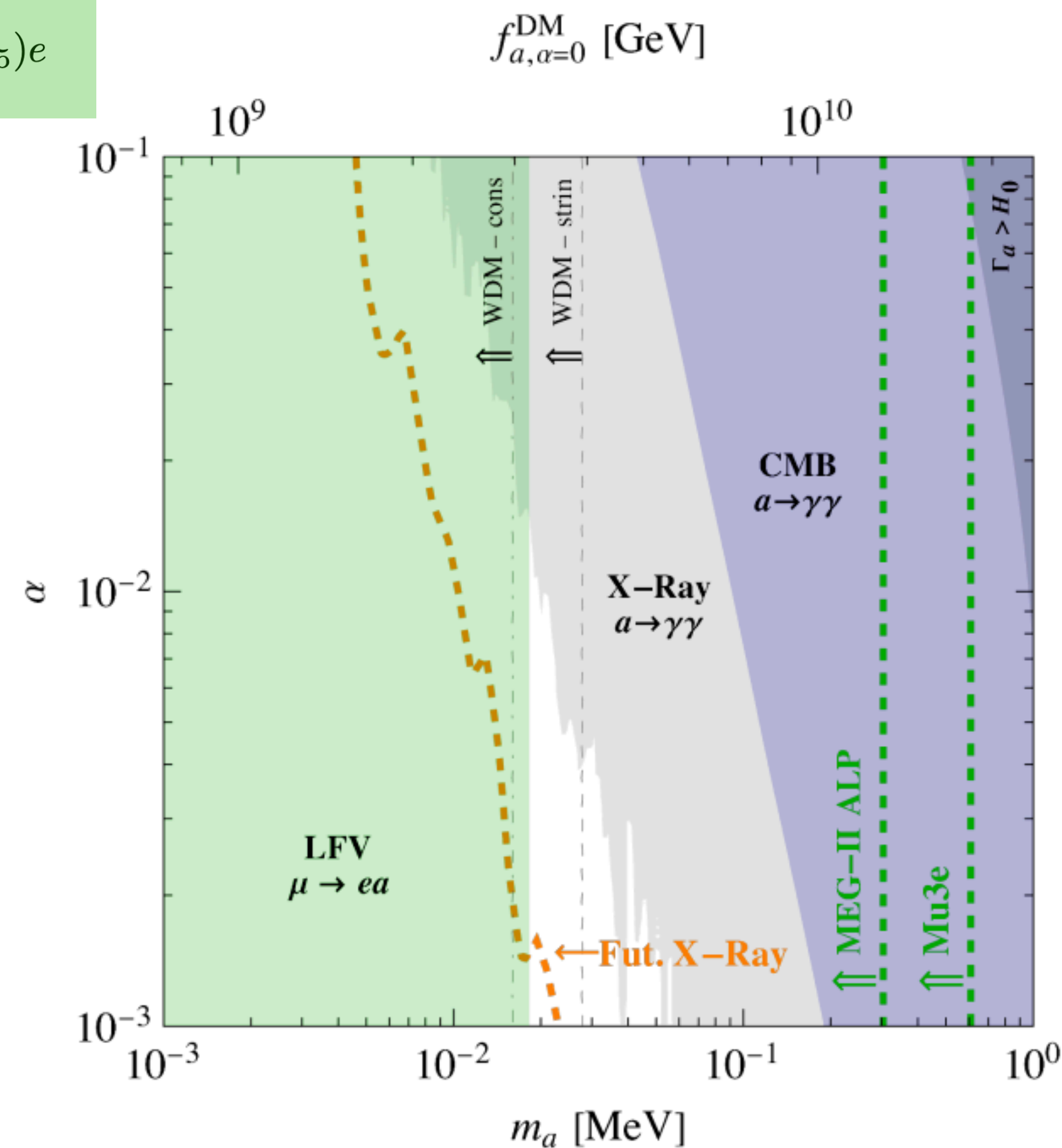
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Panci, Redigolo, Schwetz Ziegler 2209.03371



MEG-II can surpass bounds from star cooling!



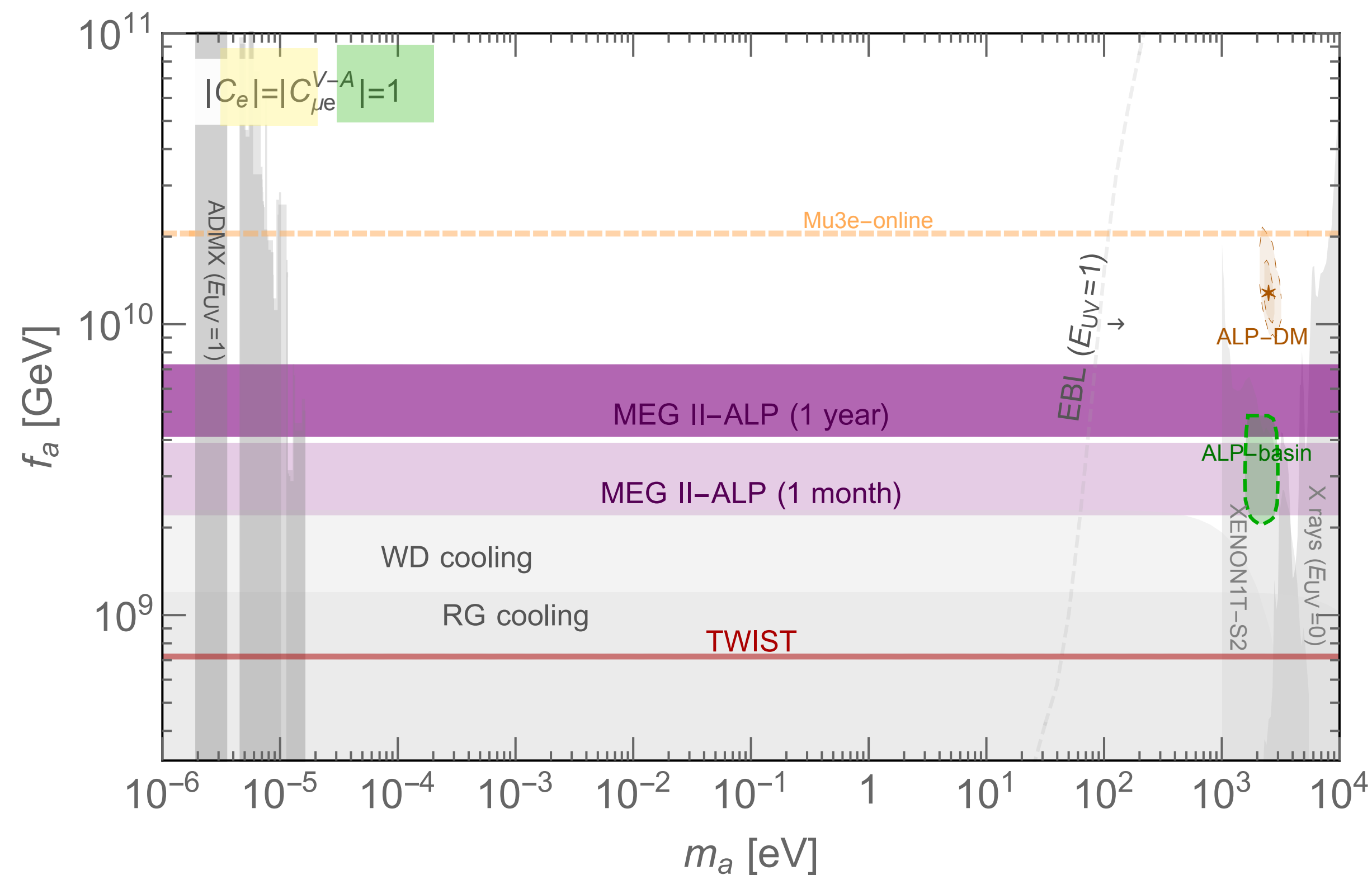
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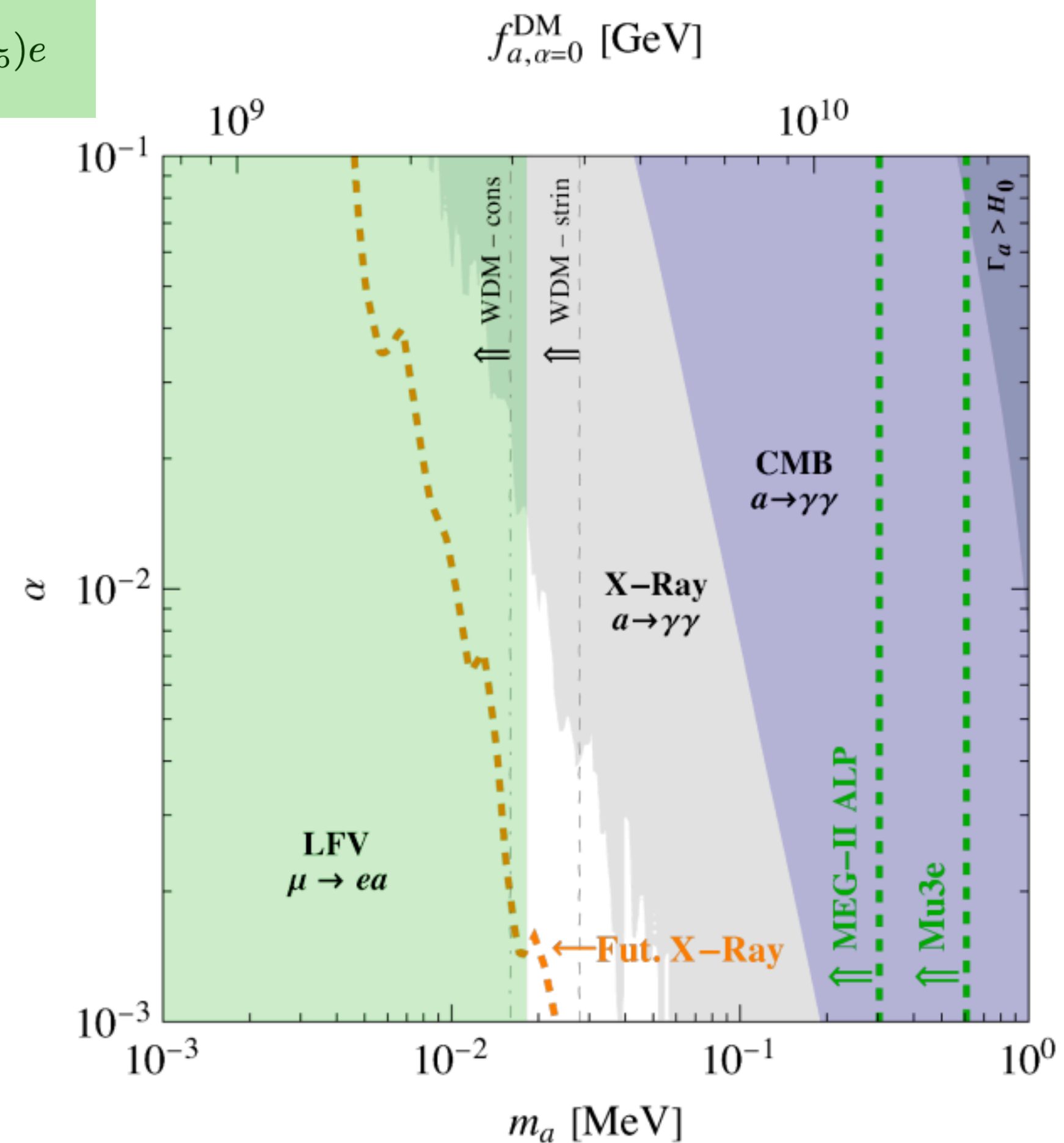
$$\frac{\partial_\mu a}{f_a} \bar{e} \gamma^\mu \gamma_5 e$$

$$\frac{\partial_\mu a}{2f_a} \bar{\mu} \gamma^\mu (C_{\mu e}^V + C_{\mu e}^A \gamma_5) e$$

Panci, Redigolo, Schwetz Ziegler 2209.03371



MEG-II can surpass bounds from star cooling!



MEG-II can completely test Freeze-in model based on LFV decays

Flavor preserving light new physics

The presence of light states can give final states close enough to FV muon decays without breaking any SM accidental symmetry!

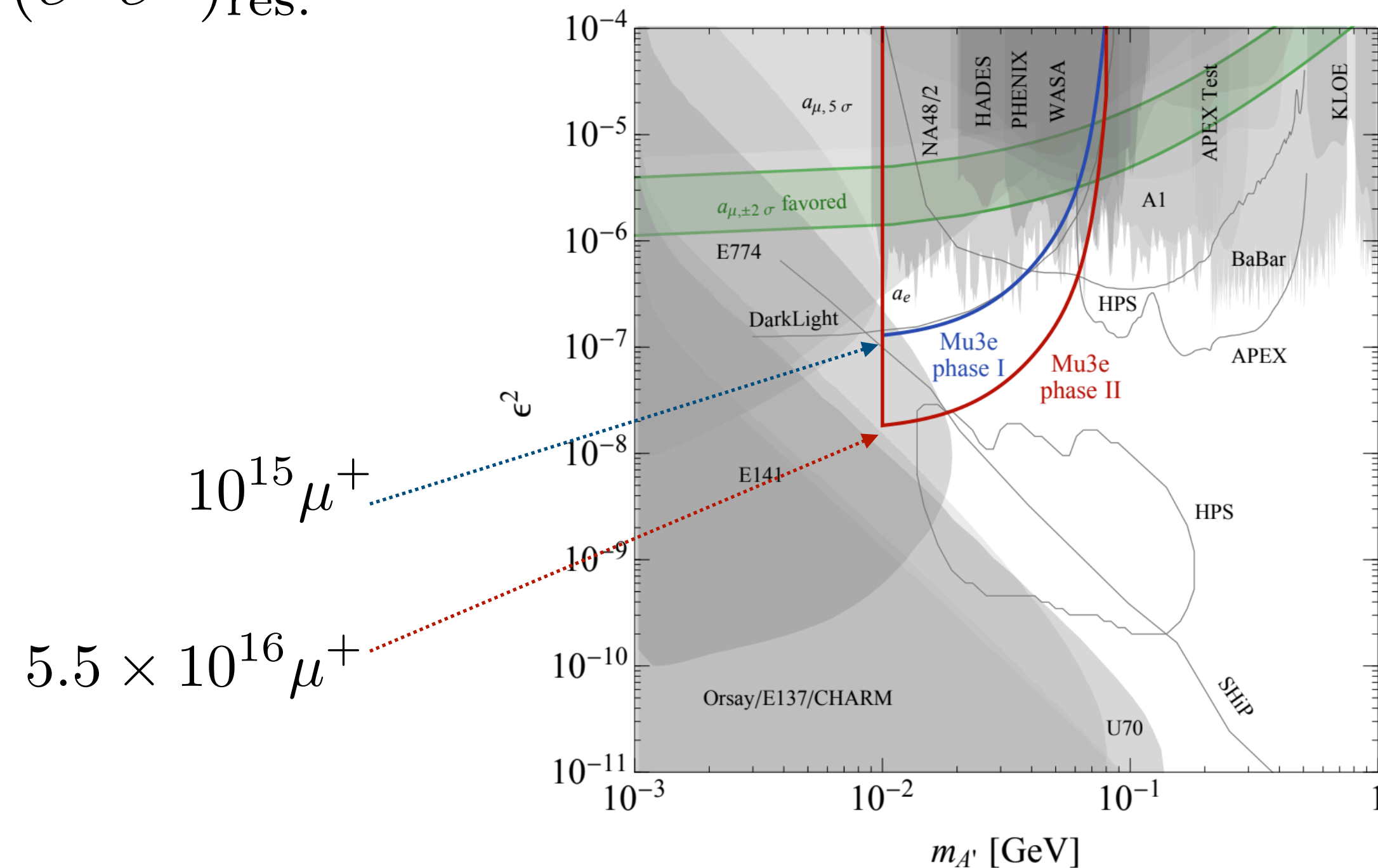
Example: $\mathcal{L}_{\gamma'} \supset F'^{\mu\nu} F'_{\mu\nu} + m_A A'^{\mu} A'_{\mu} + \epsilon A'^{\mu} J_{\mu}^{\text{SM}}$

$$\mu^+ \rightarrow e^+ \nu \bar{\nu} \gamma' \rightarrow e^+ \nu \bar{\nu} (e^+ e^-)_{\text{res.}}$$

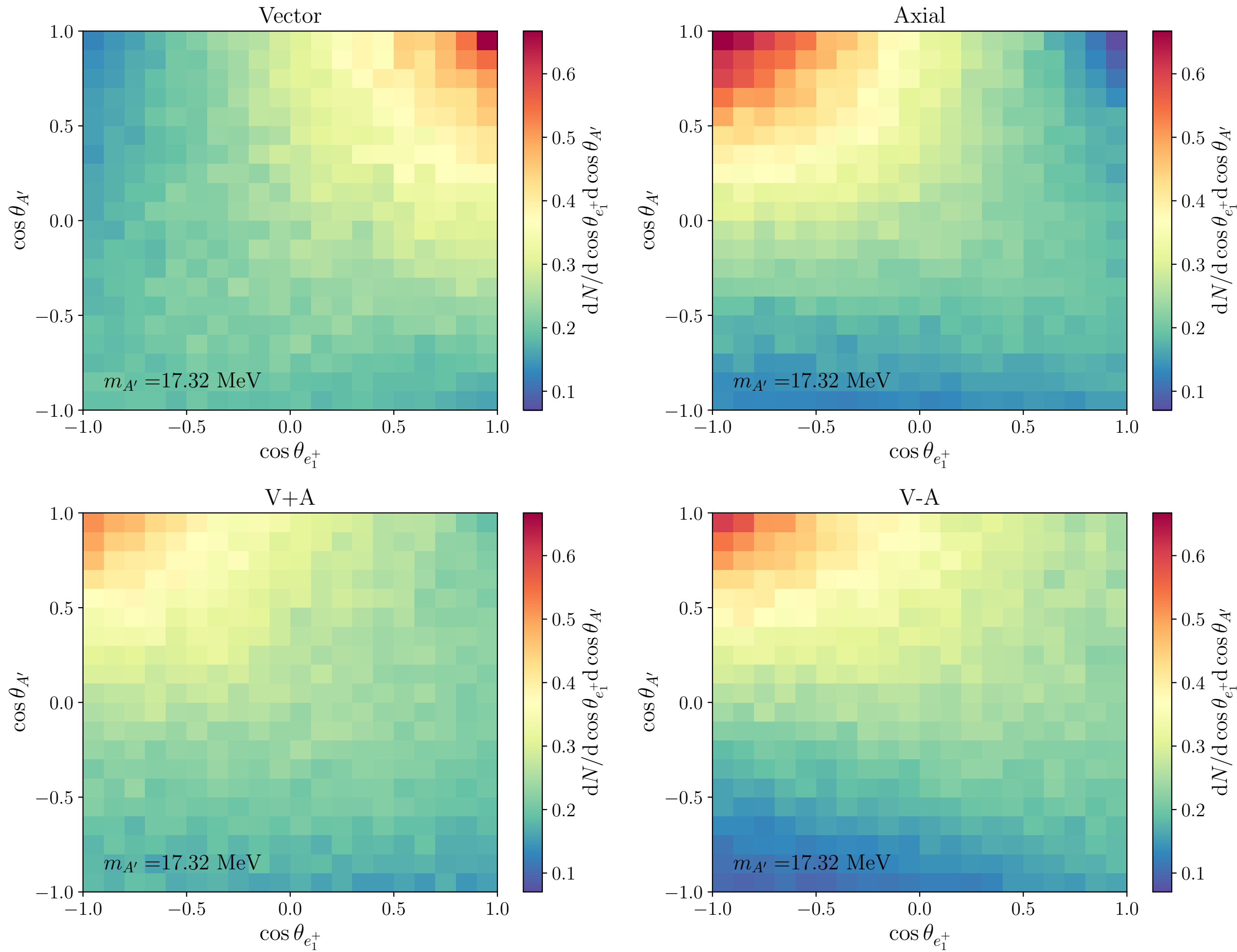
Bkd: $\mu \rightarrow e^+ \nu \bar{\nu} e^+ e^-$

bump-hunt in invariant mass

See. [Echenard, Essig and Zhong 1409.0638](#)

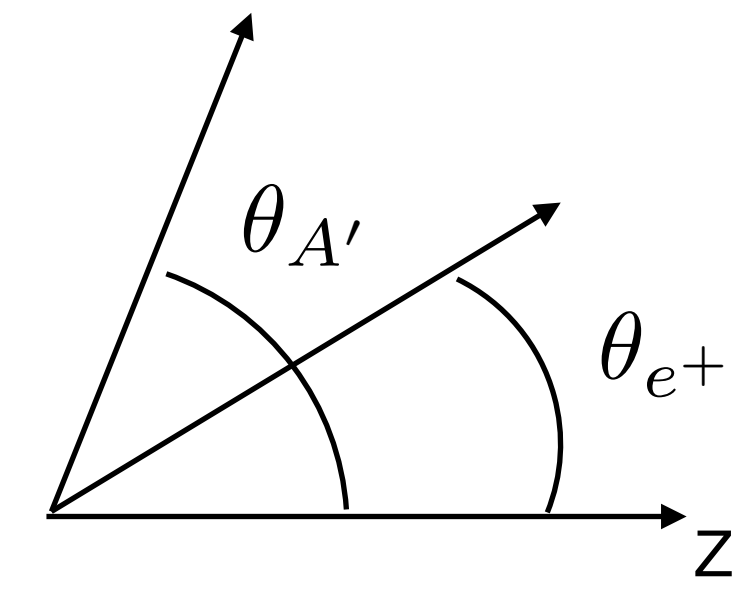


Distinguish vector chirality through angular distributions



$$m_A A'^{\mu} A'_{\mu} + \epsilon A'^{\mu} J_{\mu}^{\text{SM}}$$

- Vector: $J_{\mu}^{\text{SM}} = \bar{f} \gamma_{\mu} f$
- Axial: $J_{\mu}^{\text{SM}} = \bar{f} \gamma_{\mu} \gamma_5 f$
- Right: $J_{\mu}^{\text{SM}} = \bar{f} \gamma_{\mu} (1 + \gamma_5) f$
- Left: $J_{\mu}^{\text{SM}} = \bar{f} \gamma_{\mu} (1 - \gamma_5) f$

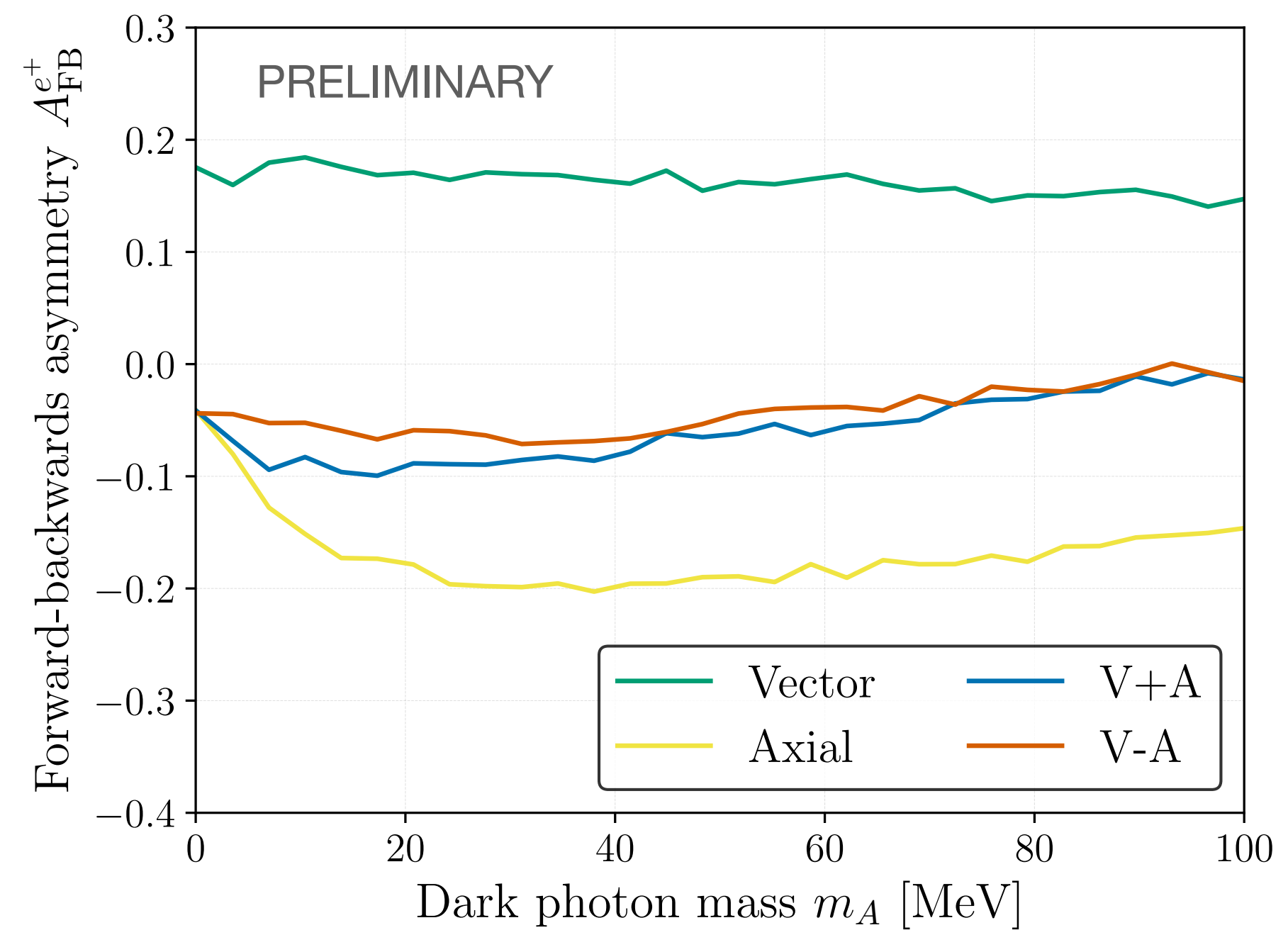
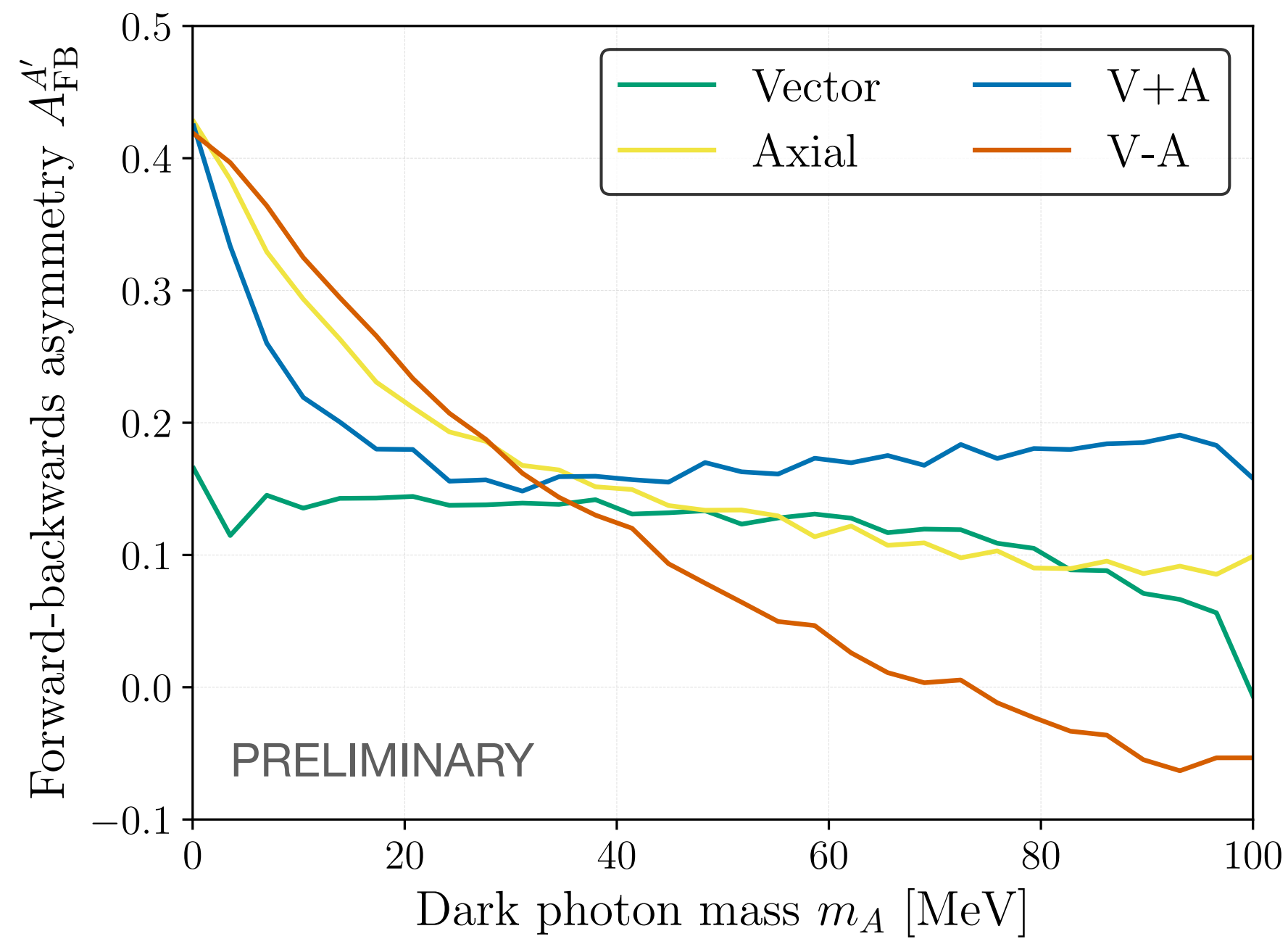


PRELIMINARY

Angular cuts

to suppress background and to distinguish signals

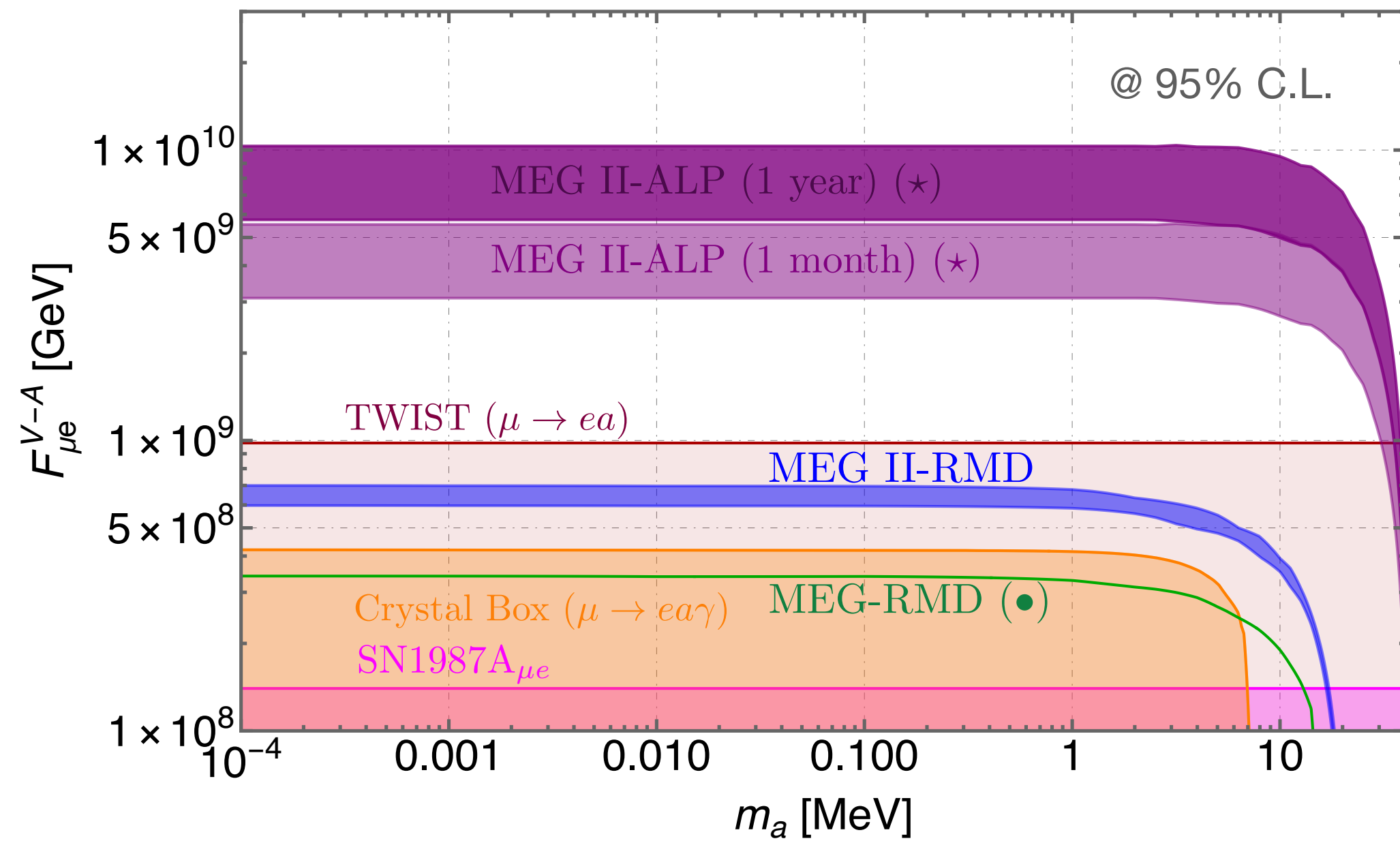
$$A \equiv \frac{F - B}{F + B} \quad F = \int d \cos \theta \frac{d\Gamma}{d \cos \theta} \quad B = \int_{-1}^0 d \cos \theta \frac{d\Gamma}{d \cos \theta}$$



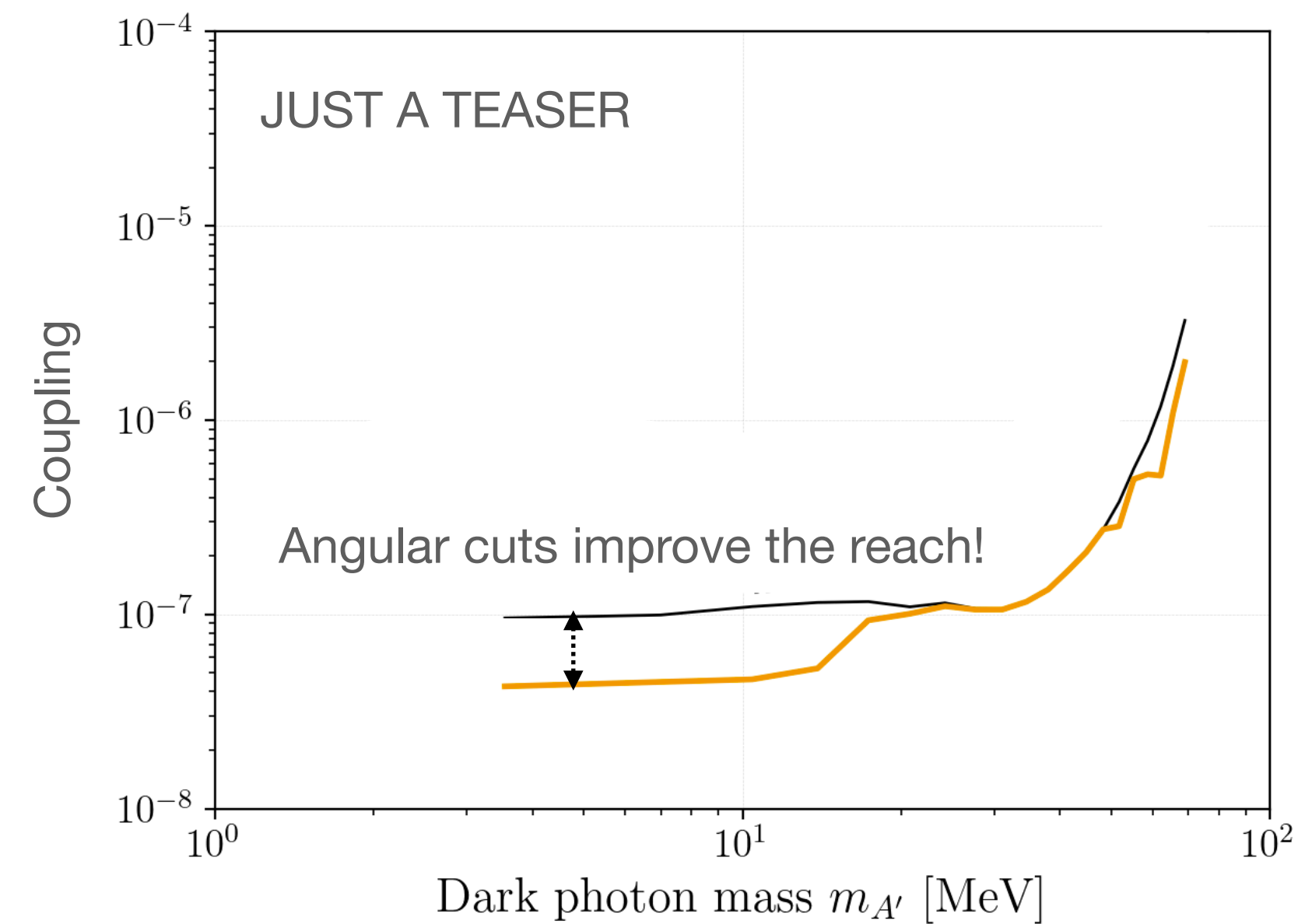
Conclusions

Light new physics opens new experimental opportunities for LFV experiments!

Testing the flavor properties of the axion at MEG II



Testing the chirality of the dark photon at Mu3e

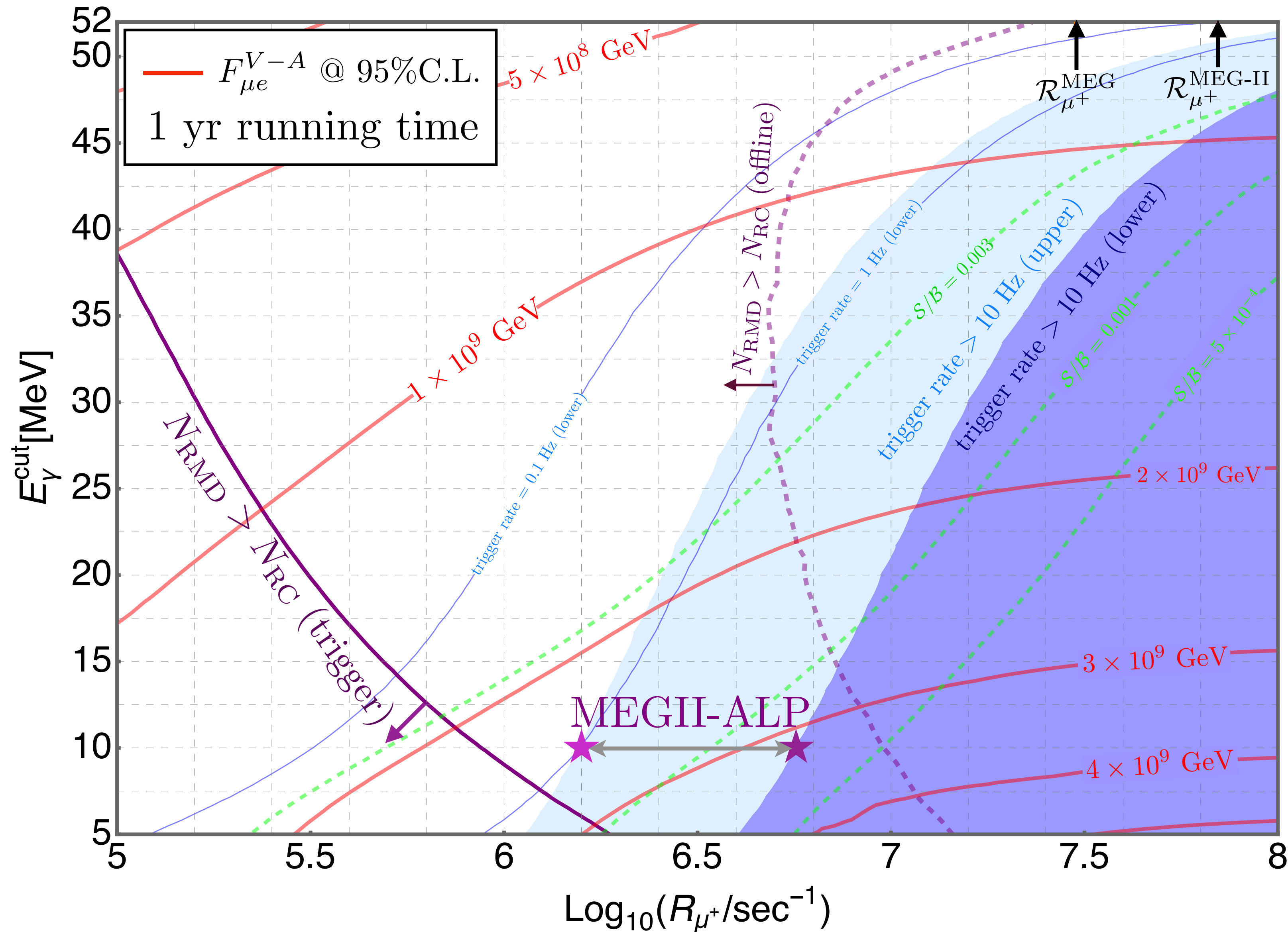


To be continued...

ВАСИКУР



Final reach II



Systematics have to be controlled at per mill level (S/B is a good proxy)

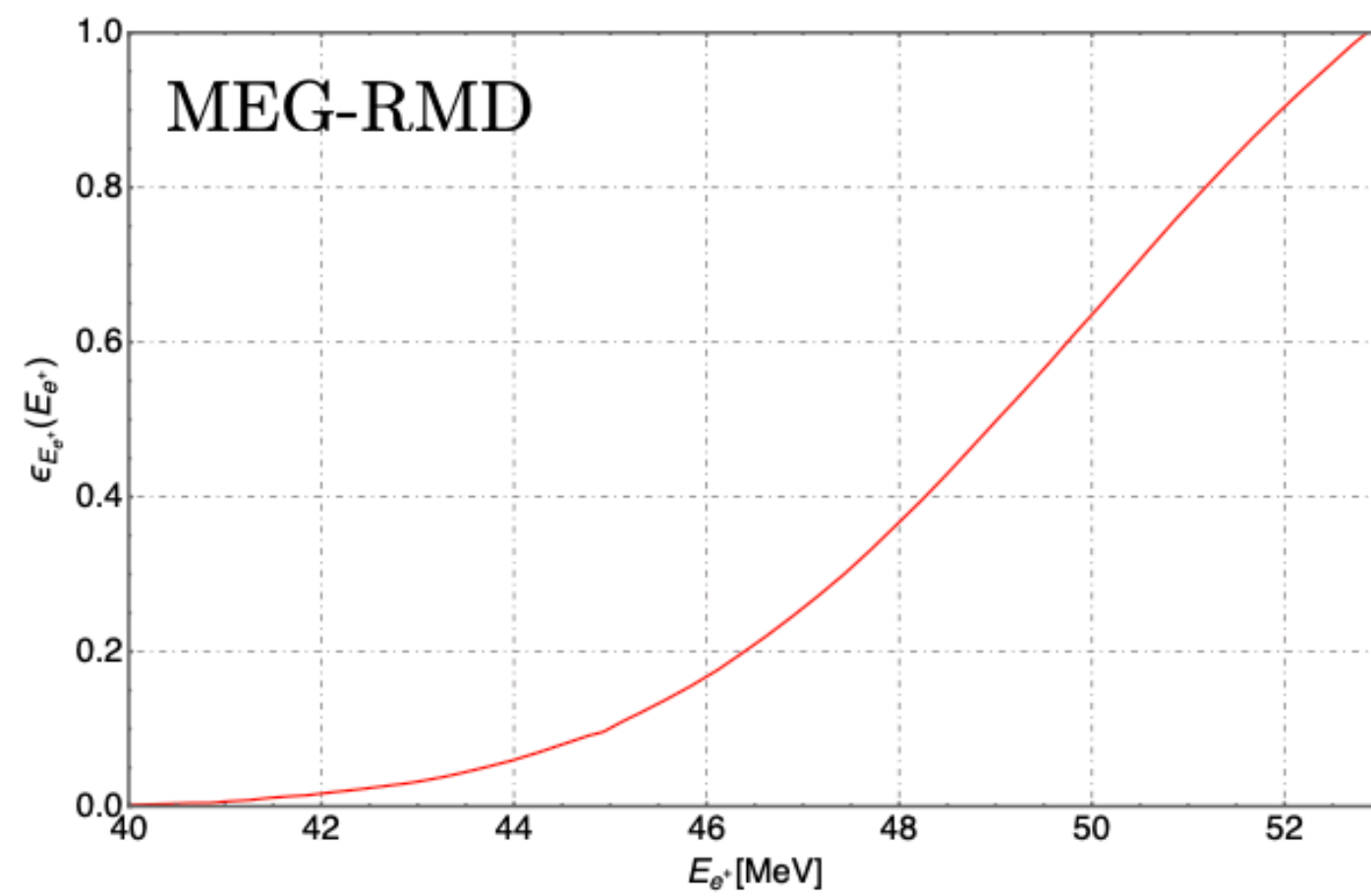
How the trigger rate was estimated I

$$C_{\text{RMD}} \equiv \frac{N_{\text{RMD}}|_{\text{obs.}}}{N_{\mu^+, \text{tot}}^{\text{MEG}} \cdot \text{BR}_{\text{RMD}}^{\text{base}} \cdot \epsilon_{\text{RMD}}^{\text{trig.}} \cdot \epsilon_{\text{RMD}}^{\text{off.}} / \epsilon_{\text{RMD}}^{\text{trig.}}}$$

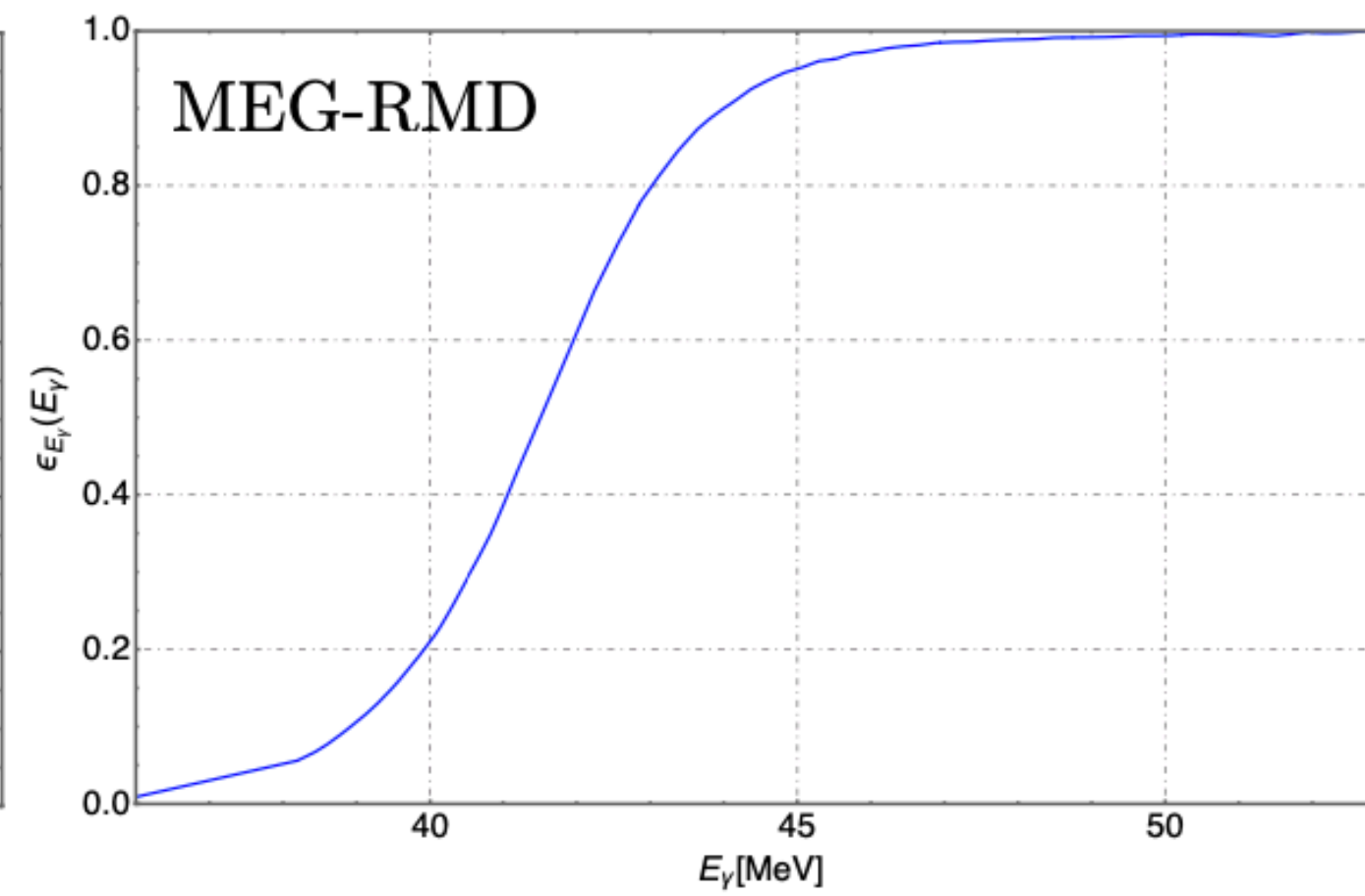
$C_{\text{RMD}} \simeq 0.35$

Known from MEG F \rightarrow $N_{\mu^+, \text{tot}}^{\text{MEG}}$
 Known from MEG RMD paper \rightarrow $N_{\text{RMD}}|_{\text{obs.}}$
 $E_e > 40 \text{ MeV}, E_\gamma > 5 \text{ MeV.}$ \rightarrow $\text{BR}_{\text{RMD}}^{\text{base}}$
 trigger efficiencies \rightarrow $\epsilon_{\text{RMD}}^{\text{trig.}}$
 trigger efficiencies+offline selection \rightarrow $\epsilon_{\text{RMD}}^{\text{off.}} / \epsilon_{\text{RMD}}^{\text{trig.}}$

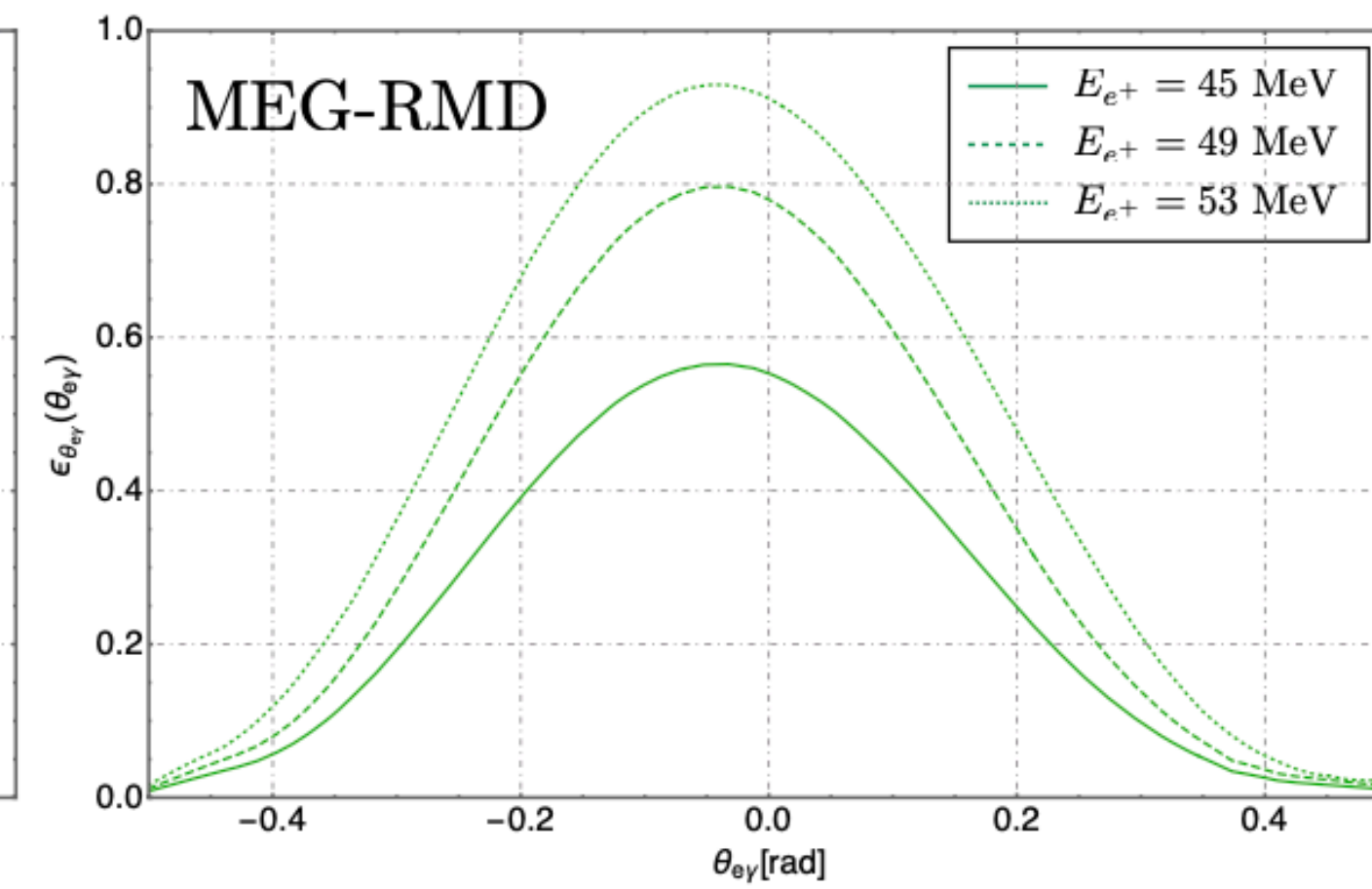
Positron energy >45 MeV @ hardware level



Photon >45 MeV @ trigger level



back to back topology @ trigger level



$$\epsilon_{\text{trigger}}^{\text{MEG}} \equiv \epsilon_{E_e}(E_e) \times \epsilon_{E_\gamma}(E_\gamma) \times \epsilon_{\theta_{e\gamma}}(E_e, \theta_{e\gamma})$$

How the trigger rate was estimated II

$$N_{\mu^+, \text{tot}}^{\text{MEG}} \cdot \text{BR}_{\text{RC}}^{\text{base}} \cdot \epsilon_{\text{RC}}^{\text{trig.}} \cdot \epsilon_{\text{RC}}^{\text{off.}} / \epsilon_{\text{RC}}^{\text{trig.}} = N_{\text{RC}}|_{\text{obs.}}$$

$$\text{BR}_{\text{RC}}^{\text{base}} = c_{\text{RC}} \cdot \text{BR}_{\text{RMD}}^{\text{base}'} \cdot \text{BR}_{\text{Mich.}}^{\text{base}} \cdot R_{\mu^+} \cdot \Delta t_{e\gamma}^{\text{trig.}}$$

ALL IN ALL

$$R_{\text{RMD}}^{\text{trig.}} \in \frac{N_{\text{RMD}}|_{\text{obs.}}}{\epsilon_{\text{RMD}}^{\text{off.}} \cdot t_{\text{run}}^{\text{MEG}}} \left(1, \frac{1}{c_{\text{RMD}}} \right) = (1.7 - 4.8) \cdot 10^{-2} \text{ Hz}$$

$$R_{\text{RC}}^{\text{trig.}} \in \frac{N_{\text{RC}}|_{\text{obs.}}}{\epsilon_{\text{RC}}^{\text{off.}} \cdot t_{\text{run}}^{\text{MEG}}} \left(1, \frac{1}{c_{\text{RC}}} \right) = (0.7 - 10) \text{ Hz} .$$

The c's factors give the size of our uncertainties in the trigger estimate