



The Theory Landscape of Dark Matter



Yonit Hochberg

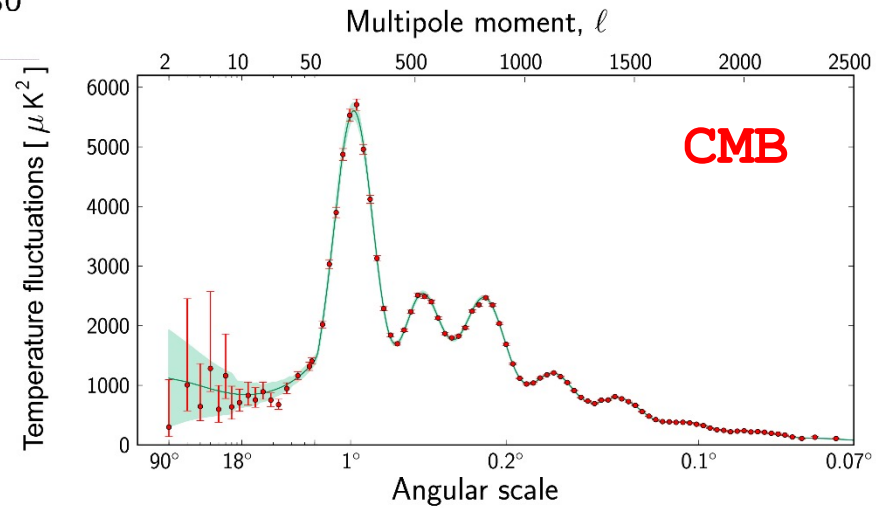
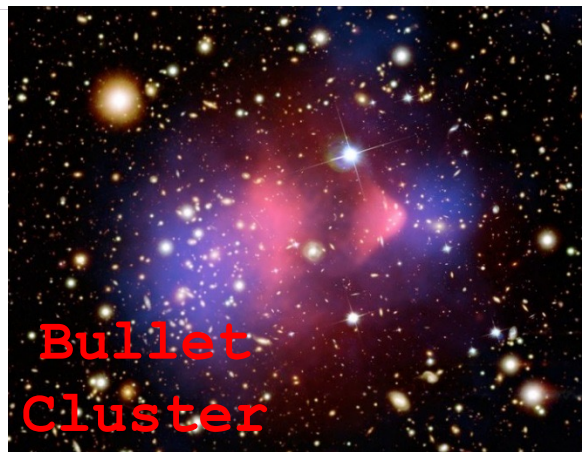
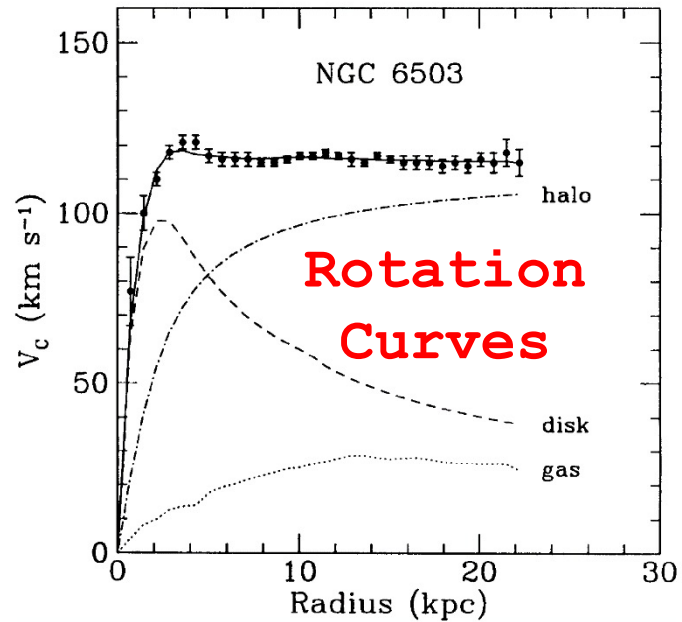


האוניברסיטה העברית בירושלים
THE HEBREW UNIVERSITY OF JERUSALEM

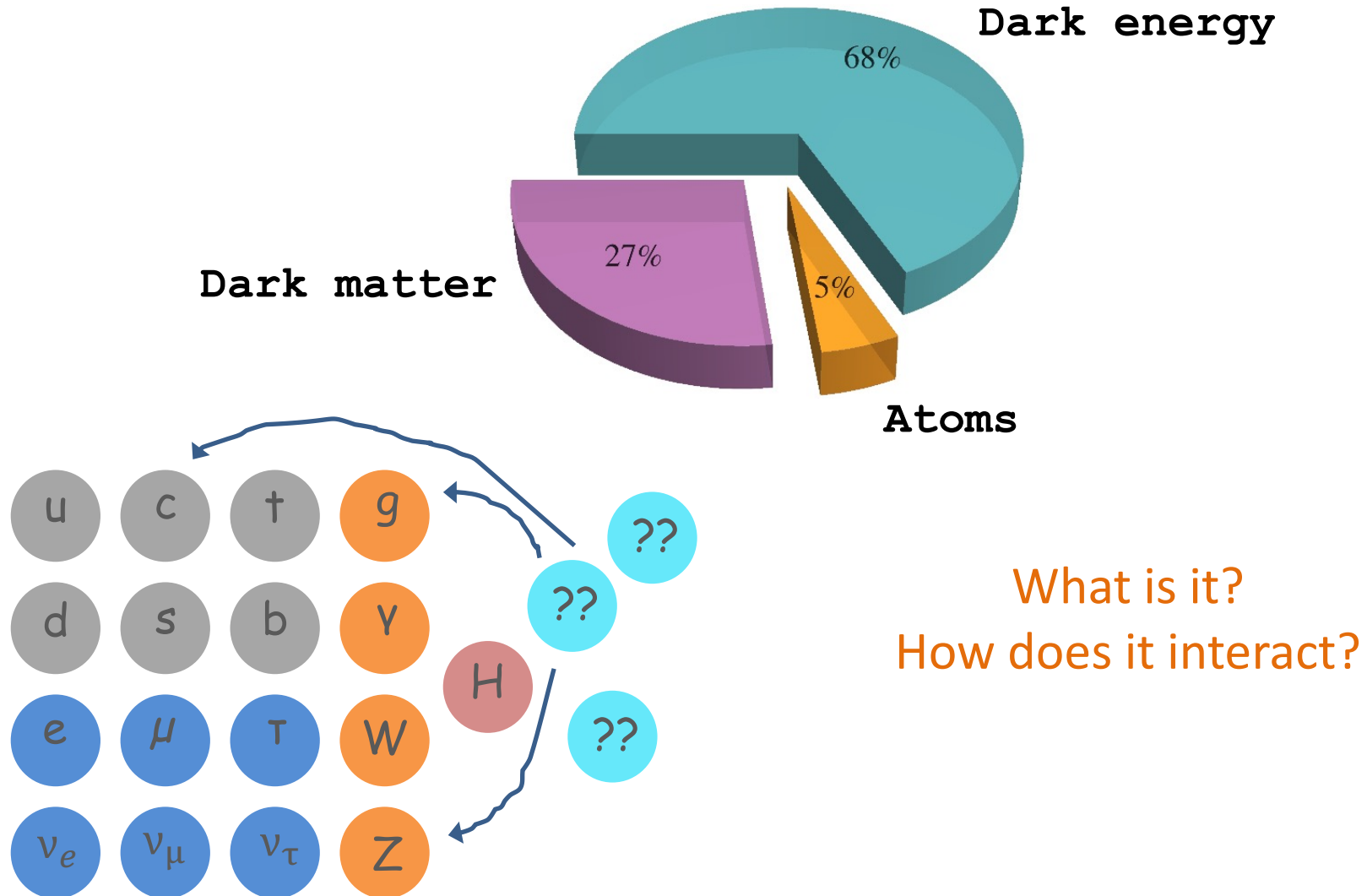


מכון רקח
The Racah Institute
לפיסיקה
of Physics

Evidence for Dark Matter



The Universe is Dark



What do we know?

- Dark matter has 5 times the mass density of baryons
- Massive ($m=???$)
- Can't interact too strongly with QED and QCD
- Doesn't interact too strongly with itself

Past 40 years

WIMP, glorious WIMP^{*}

{ *Also axions, of course
also axions :-) }

The WIMP Miracle

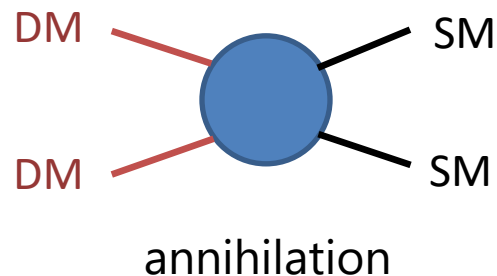
Correct thermal relic abundance:

$$m_{\text{DM}} \sim \alpha \times 30 \text{ TeV}$$

For weak coupling, weak scale emerges.

Weakly Interacting Massive Particle (WIMP)

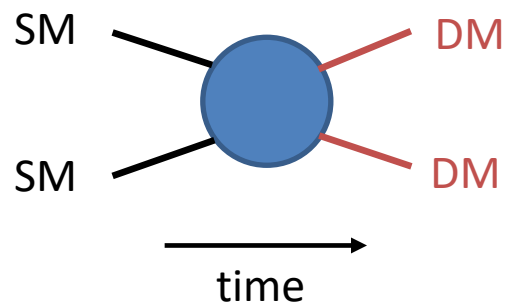
The dominant paradigm for 40+ years.



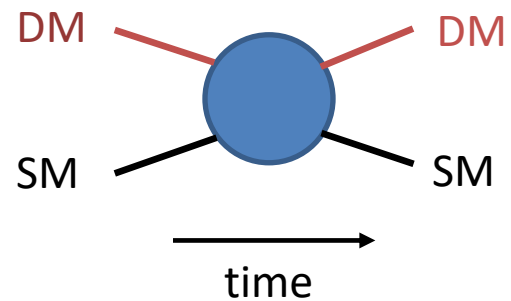
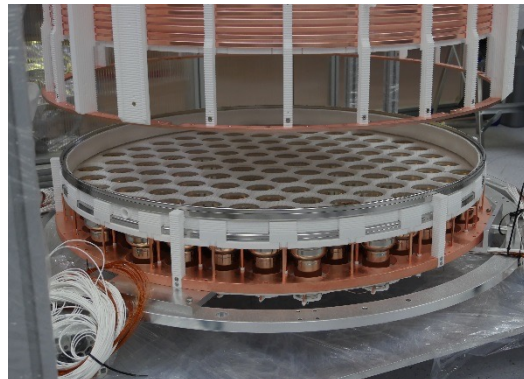
$$\langle \sigma_{\text{ann}} v \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

Searching for WIMPs

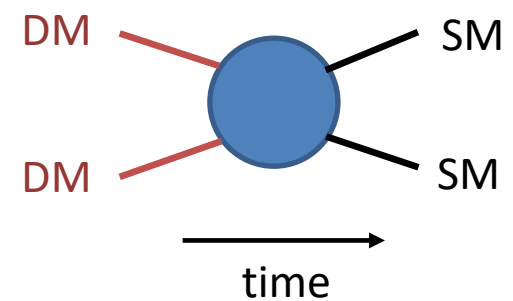
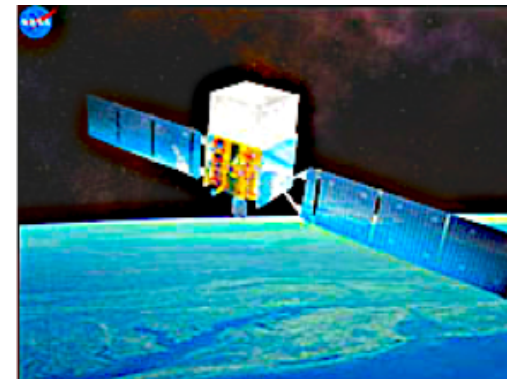
Direct production



Direct detection



Indirect detection

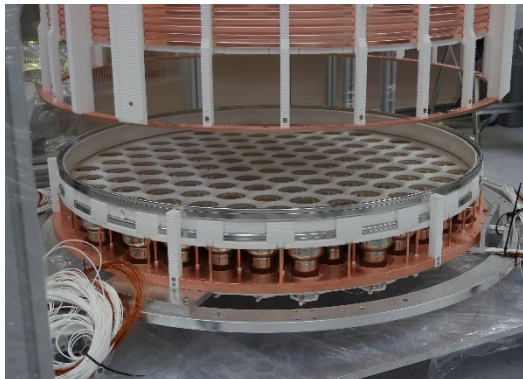


Searching for WIMPs

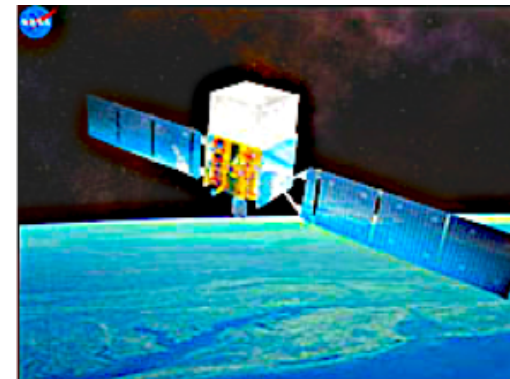
Direct production



Direct detection



Indirect detection



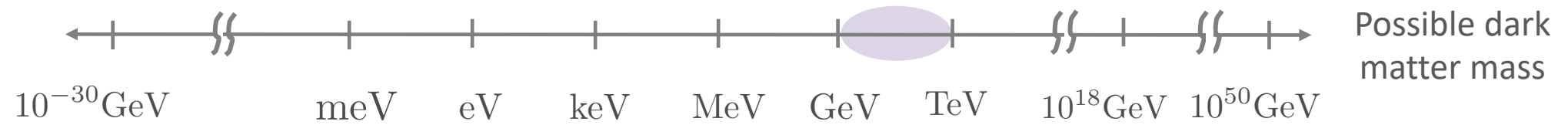
Experiments getting increasingly sensitive

Haven't yet detected dark matter

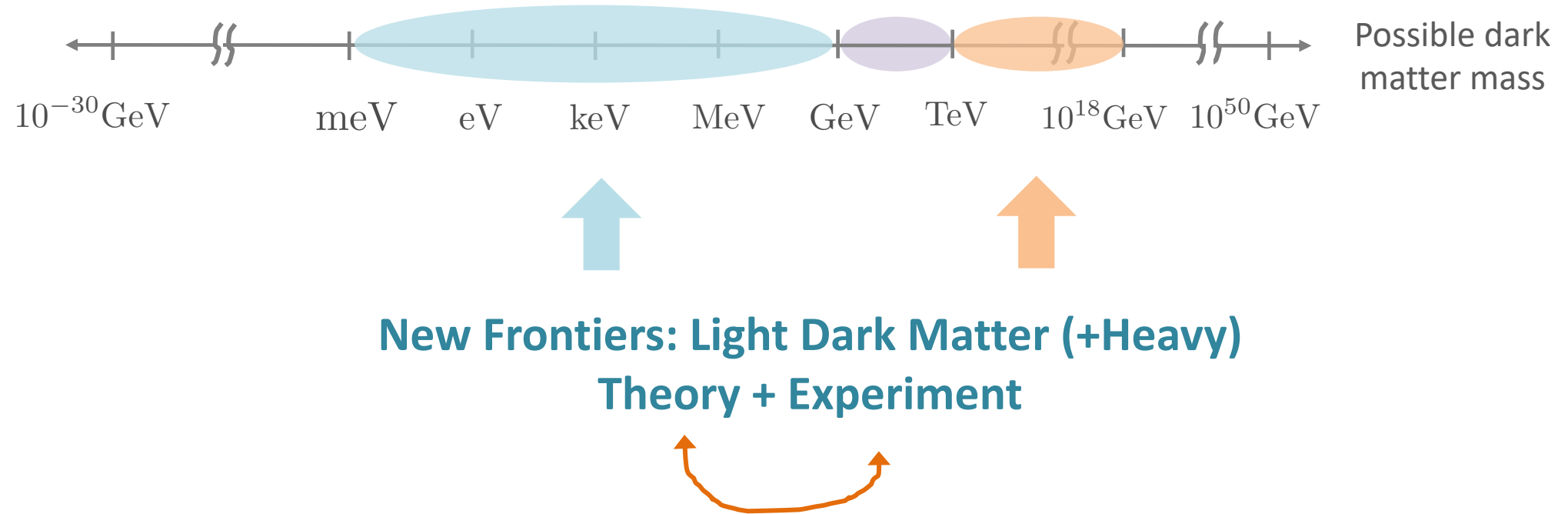


Great opportunity for new ideas.

Beyond the WIMP



Beyond the WIMP



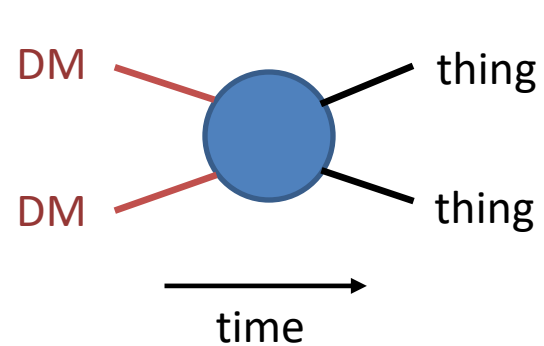
New Theory Ideas

-
- Weakly coupled WIMPs Pospelov, Ritz, Voloshin 2007; Feng, Kumar 2008
- Asymmetric dark matter Kaplan, Luty, Zurek, 2009
- Freeze-in dark matter Hall, Jedamzik, March-Russell, West, 2009
- SIMPs YH, Kuflik, Volansky, Wacker, 2014 | YH, Kuflik, Murayama, Volansky, Wacker, 2015
- ELDERS Kuflik, Perelstein, Rey-Le Lorier, Tsai, 2016 & 2017
- Forbidden dark matter Griest, Seckall, 1991 | D'Agnolo, Ruderman, 2015
- Co-decaying dark matter Dror, Kuflik, Ng, 2016
- Co-scattering dark matter D'Agnolo, Pappadopulo, Ruderman, 2017
- Zombie dark matter Kramer, Kuflik, Levi, Outmezguine, Ruderman, 2021
- Chain dark matter Kim, Kuflik, 2019 | Frumkin, Kuflik, Lavie, Silverwater, PRL 2021
- INDY dark matter Frumkin, YH, Kuflik, Murayama, PRL 2022
-

... Are abundant

By no means a comprehensive list

Ex. #1: Weakly coupled $2 \rightarrow 2$



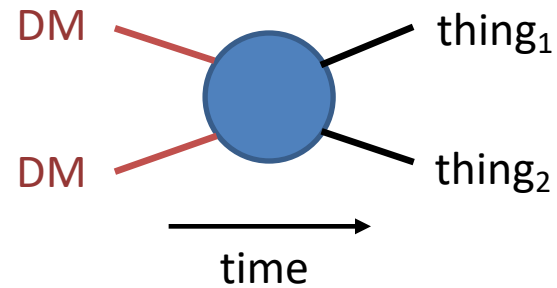
$$\langle \sigma_{\text{ann}} v \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

$$\alpha \ll 1$$

$$m_{\text{DM}} \sim \alpha \times 30 \text{ TeV}$$

Pospelov, Ritz, Voloshin 2007 | Feng, Kumar 2008

Ex. #2: Forbidden Channels



$$m_{\text{DM}} \sim \alpha \times (30 \text{ TeV}) \times e^{-x_F \Delta}$$

freezeout
temp'

mass
difference

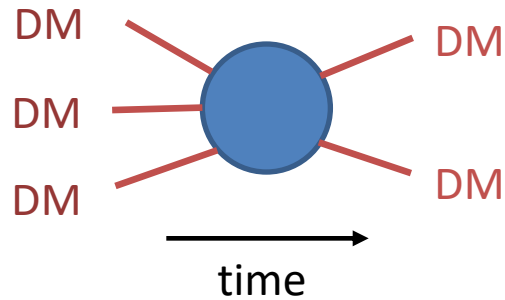
$$2m_{\text{DM}} < m_{\text{thing}_1} + m_{\text{thing}_2}$$

Forbidden @ T=0;
Boltzmann suppressed @ finite T

Griest, Seckel, 1991 | D'Agnolo, Ruderman, 2015

Ex. #3: SIMPs

What if dark matter mostly interacted with itself?



$$\langle \sigma v^2 \rangle_{3 \rightarrow 2} \equiv \frac{\alpha^3}{m_{\text{DM}}^5}$$

$$m_{\text{DM}} \sim \alpha \times 100 \text{ MeV}$$

3 $\rightarrow$ 2 self-annihilations

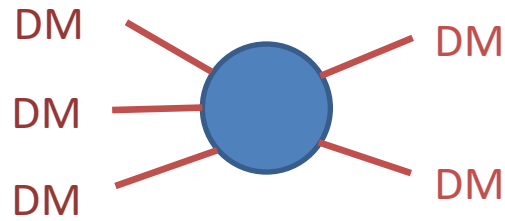
For strong coupling, the strong scale emerges.

SIMP = Strongly (self) Interacting Massive Particle

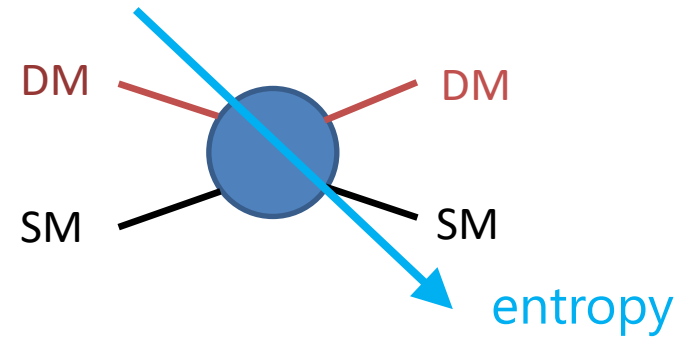
Carlson, Hall, Machacek, 1992 | YH, Kuflik, Volansky, Wacker, 2014

Ex. #3: SIMPs

Pumps heat into the system: need to shed the heat



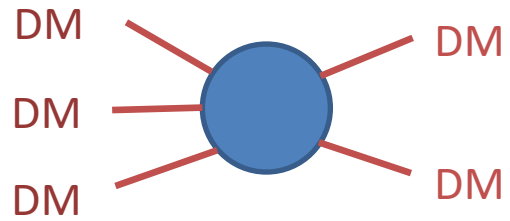
3 $\rightarrow$ 2 self-annihilations



thermalize with
light SM species
(active during freeze-out)

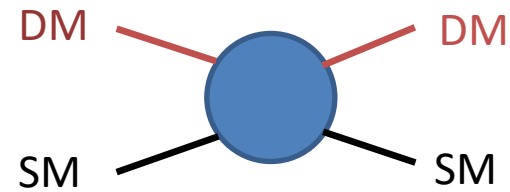
YH, Kuflik, Volansky, Wacker, 2014

Ex. #3: SIMPs



decouples 1st

Determines
DM relic density

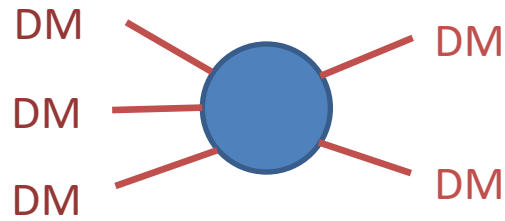


decouples 2nd

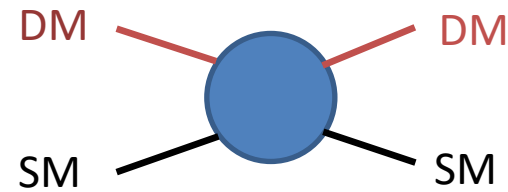
What if the order was reversed?

Ex. #4: ELDERs

ELastically DEcoupling Relic (ELDER)



decouples 2nd



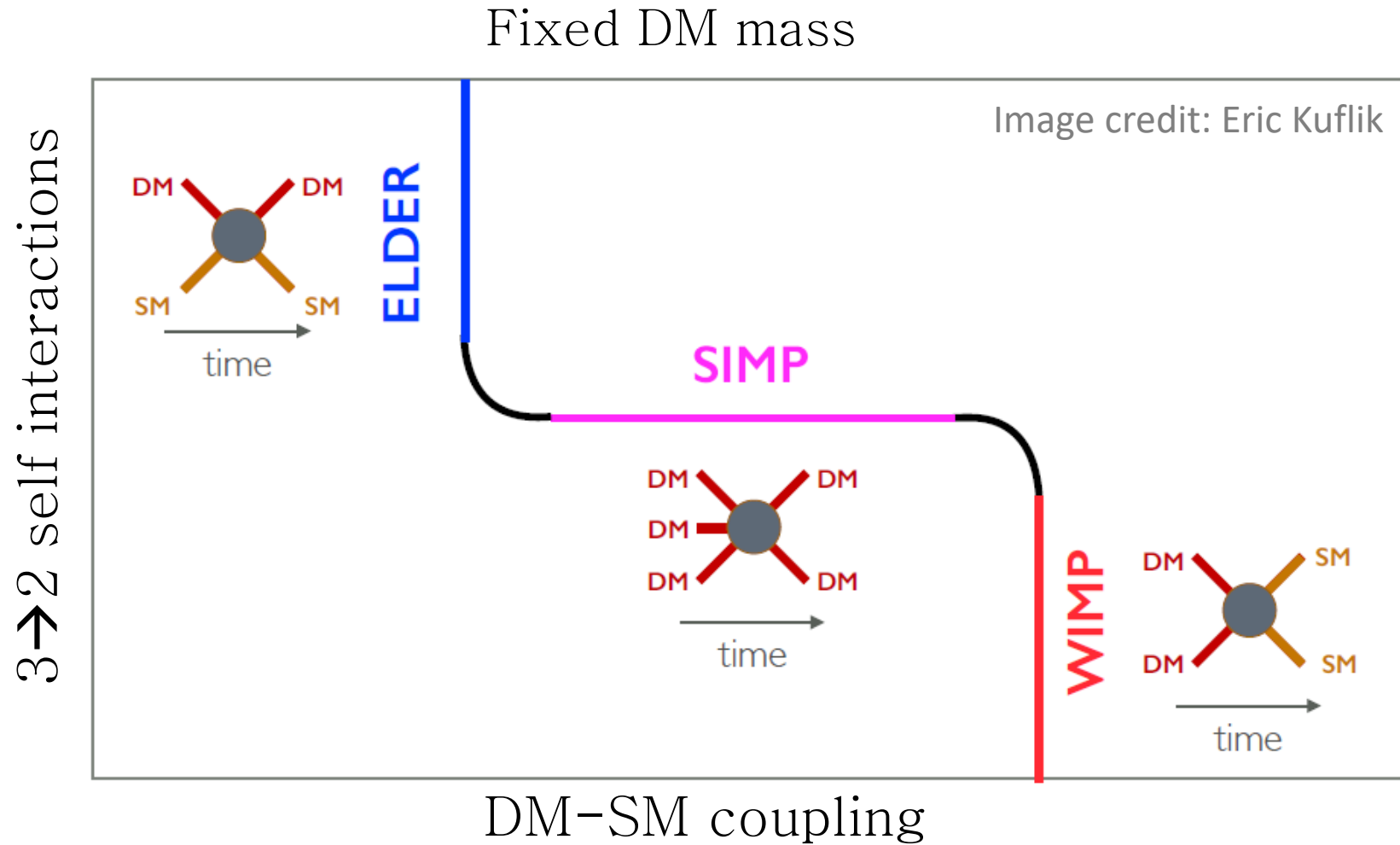
decouples 1st

Determines
DM relic density

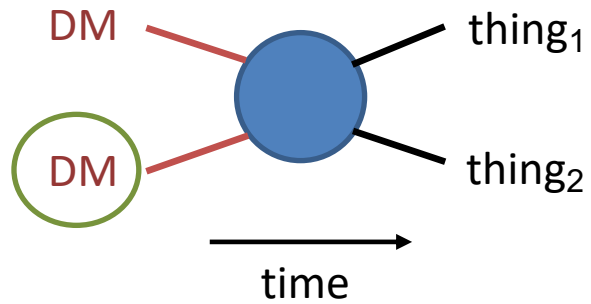
$$\Omega_{\text{DM}} \propto e^{-\langle \sigma v \rangle_{\text{el}} \#}$$

Kuflik, Perelstein, Rey-Le Lorier, Tsai, 2016 & 2017

WIMP/SIMP/ELDER

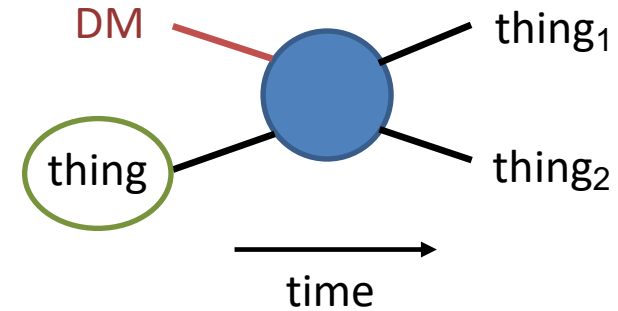


Heavy?



$$\text{Rate} \propto (n_{\text{DM}} \times \text{cross section})$$

vs.



$$\text{Rate} \propto (n_{\text{thing}} \times \text{cross section})$$

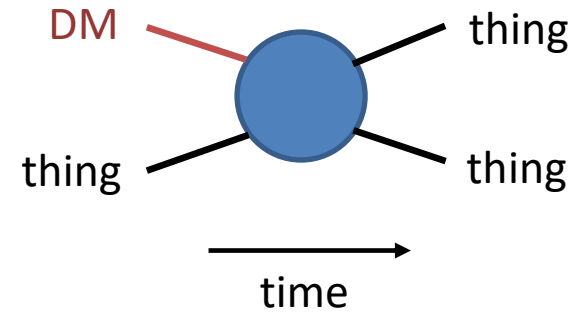
If thing = light, less Boltzmann suppressed!

Can reach heavier dark matter masses

Example #5 + #6: Zombies + Chains

Zombie DM

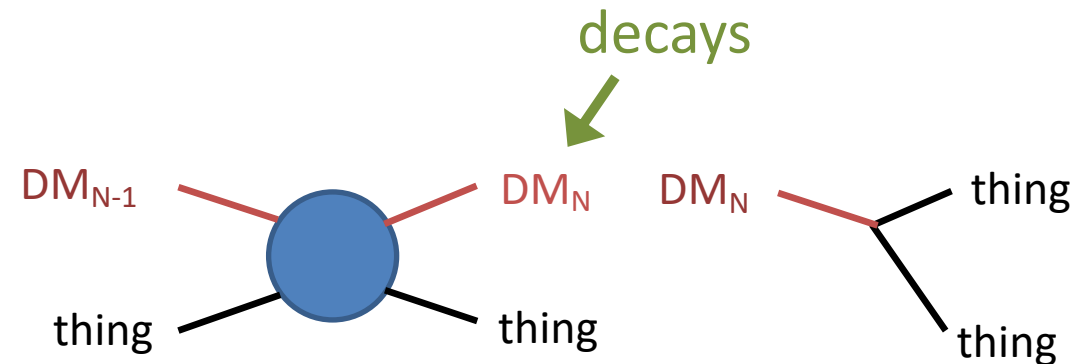
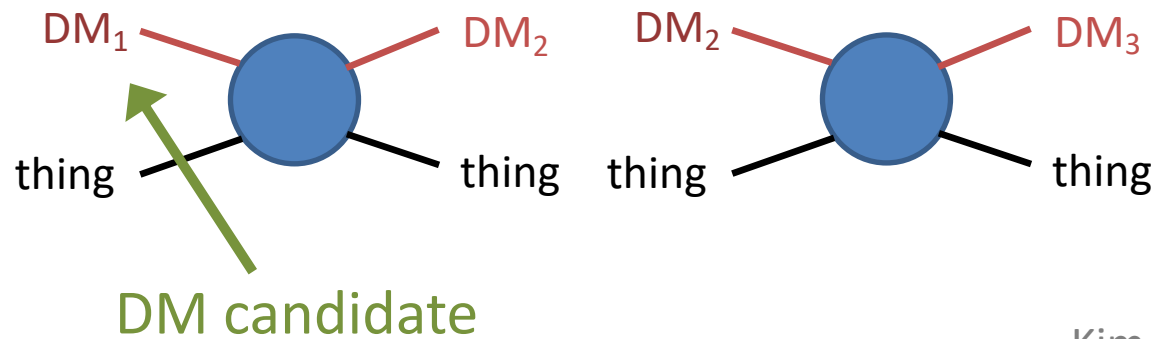
$$m_{\text{DM}} \lesssim 10^{10} \text{ GeV}$$



Kramer, Kuflik,
Levi, Outmezguine,
Ruderman,
PRL 2021

Chain DM

$$m_{\text{DM}} \lesssim 10^{18} \text{ GeV}$$



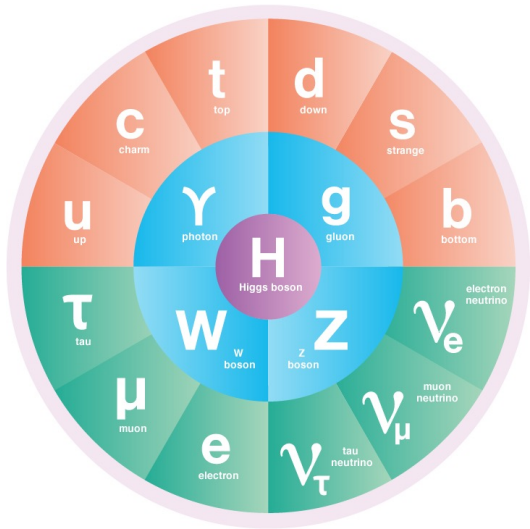
Kim, Kuflik, PRL 2019 | Frumkin, Kuflik, Lavie, Silverwater, PRL 2021



Generic.

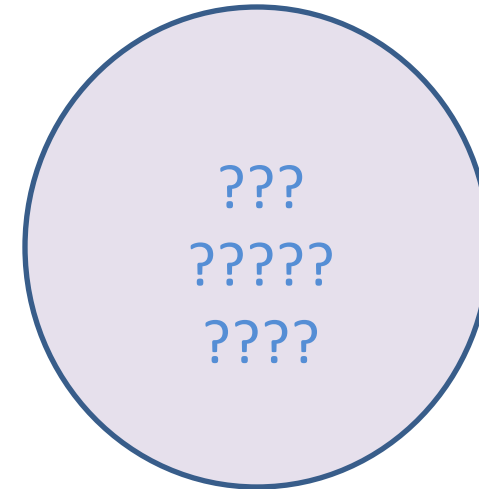
Dark Sectors

Visible sector



Zoo of particles
w/structure
 $SU(3)_C \times SU(2)_L \times U(1)_Y$

Dark sector



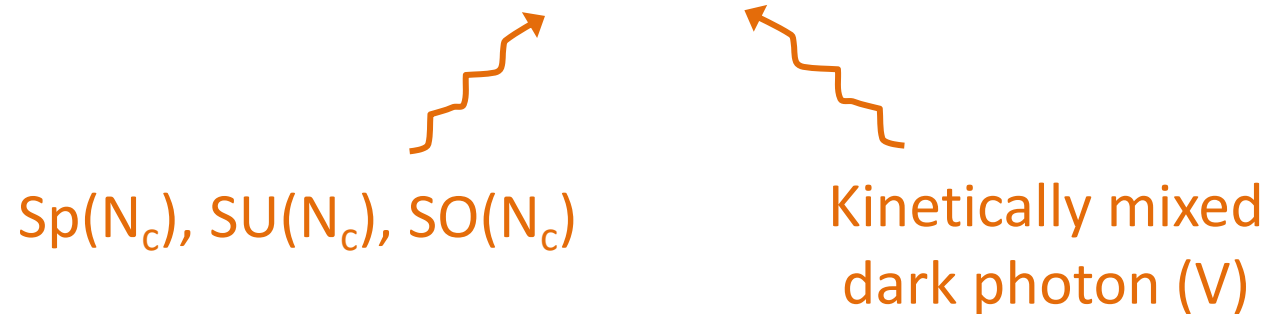
Why not in the
dark sector too?
New gauge symmetries?

Dark Sectors

Think Standard Model!

Dark matter from strongly coupled gauge theories

E.g. $SU(3)_{\text{dark}} \times U(1)_{\text{dark}}$

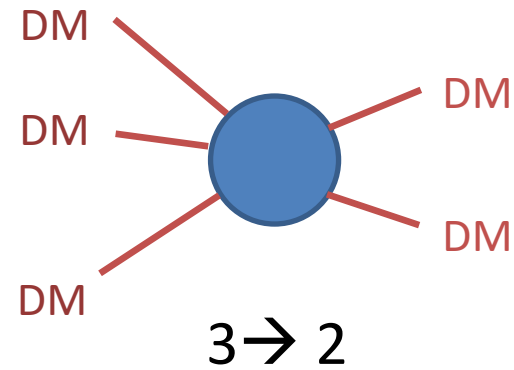


QCD-like theories, pions = dark matter

Many processes, many dark matter mechanisms.

E.g. SIMPs

Think QCD!



QCD has 5-point interactions! $K^+ K^- \rightarrow \pi^+ \pi^0 \pi^-$

WZW term

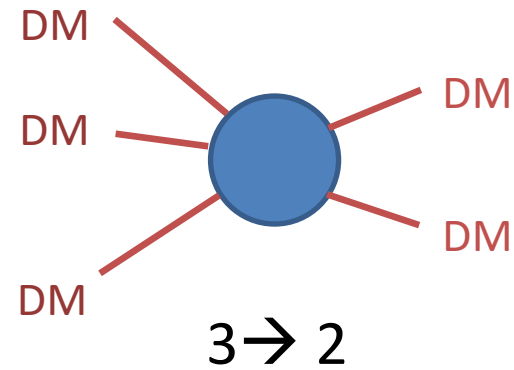
If calculate the rate, find that is just right for SIMPs with
mass $\sim$ few hundred MeV

YH, Kuflik, Murayama, Volansky, Wacker, 2015

E.g. SIMPs

Think QCD!

$\text{Sp}(N_c), \text{SU}(N_c), \text{SO}(N_c)$



$$\mathcal{L}_{\text{WZW}} = \frac{2N_c}{15\pi^2 f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \text{Tr}[\pi \partial_\mu \pi \partial_\nu \pi \partial_\rho \pi \partial_\sigma \pi]$$



pion decay constant

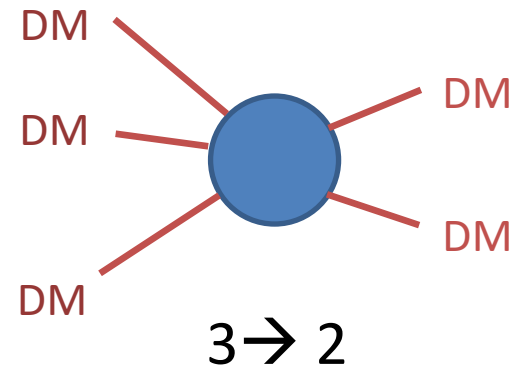
WZW term

Wess and Zumino, 1971 | Witten x 2, 1983

YH, Kuflik, Murayama, Volansky, Wacker, 2015

E.g. SIMPs

Think QCD!
 $Sp(N_c), SU(N_c), SO(N_c)$

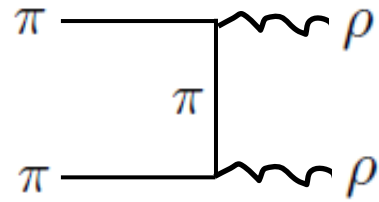


Stable dark matter = dark pions
mass $\sim$ few hundred MeV
Non-exotic!

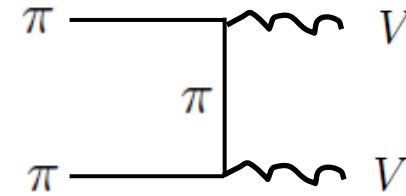
{ 3 → 2 dark glueballs: Carelson, Hall, Machacek, 1992 | Soni, Zhang, 2016 | Forestell, Morrissey, Sigurdson, 2017 }

YH, Kuflik, Murayama, Volansky, Wacker, 2015

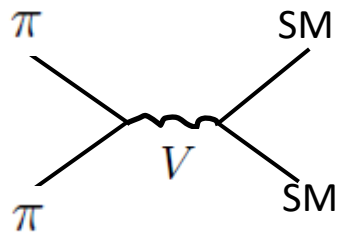
E.g. $2 \rightarrow 2$



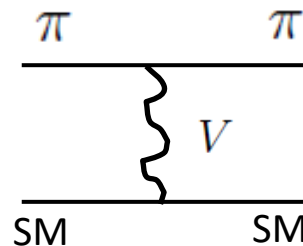
forbidden annihilations



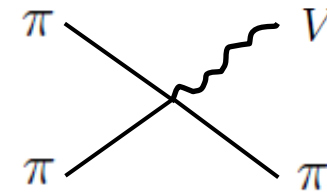
$2 \rightarrow 2$ annihilations



$2 \rightarrow 2$ annihilations



elastic scattering



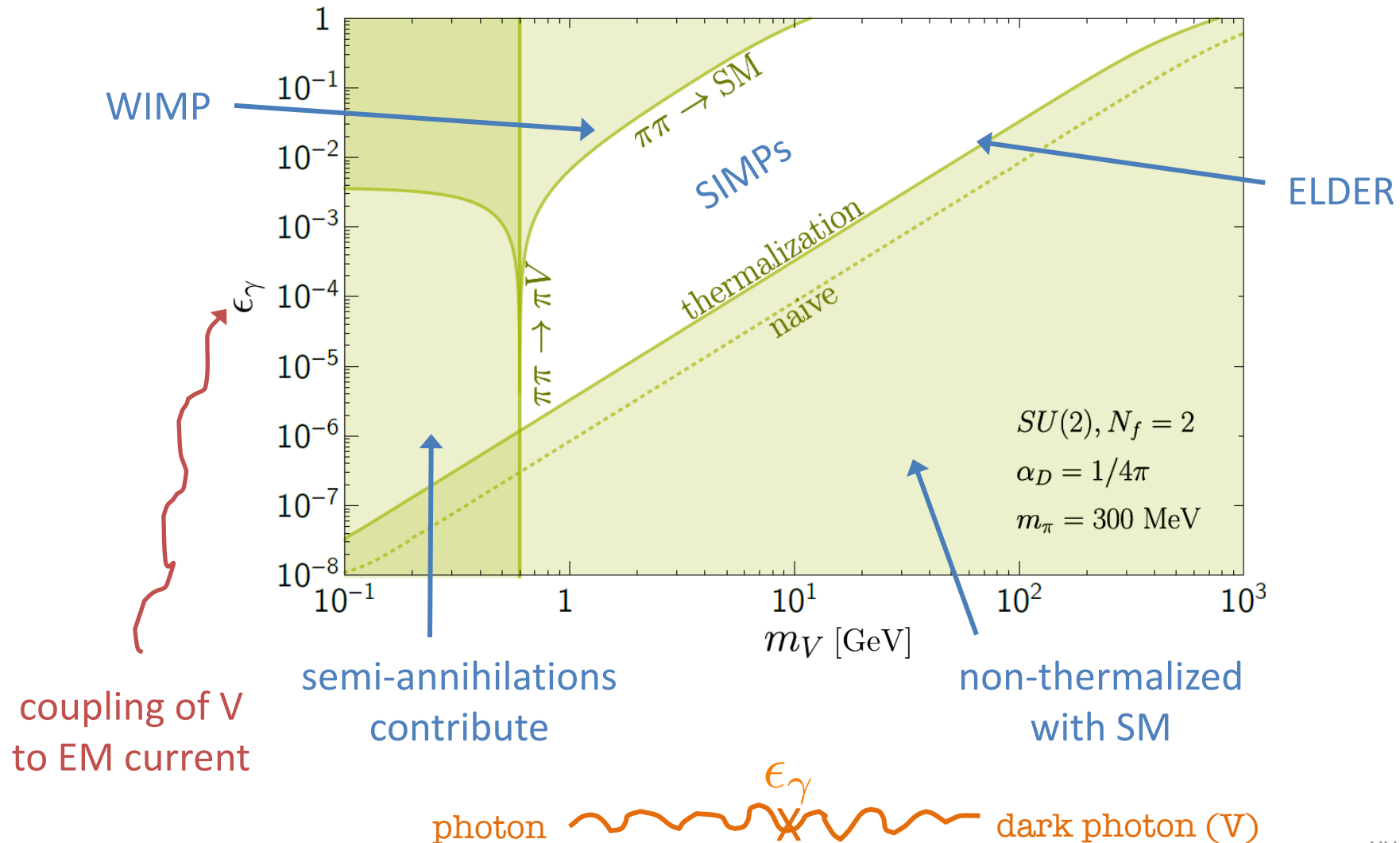
semi-annihilations

Lee, Seo, 2015 | YH, Kuflik, Murayama, 2016 | Harigaya, Nomura, 2016
Berlin, Blinov, Gori, Schuster, Toro, 2018

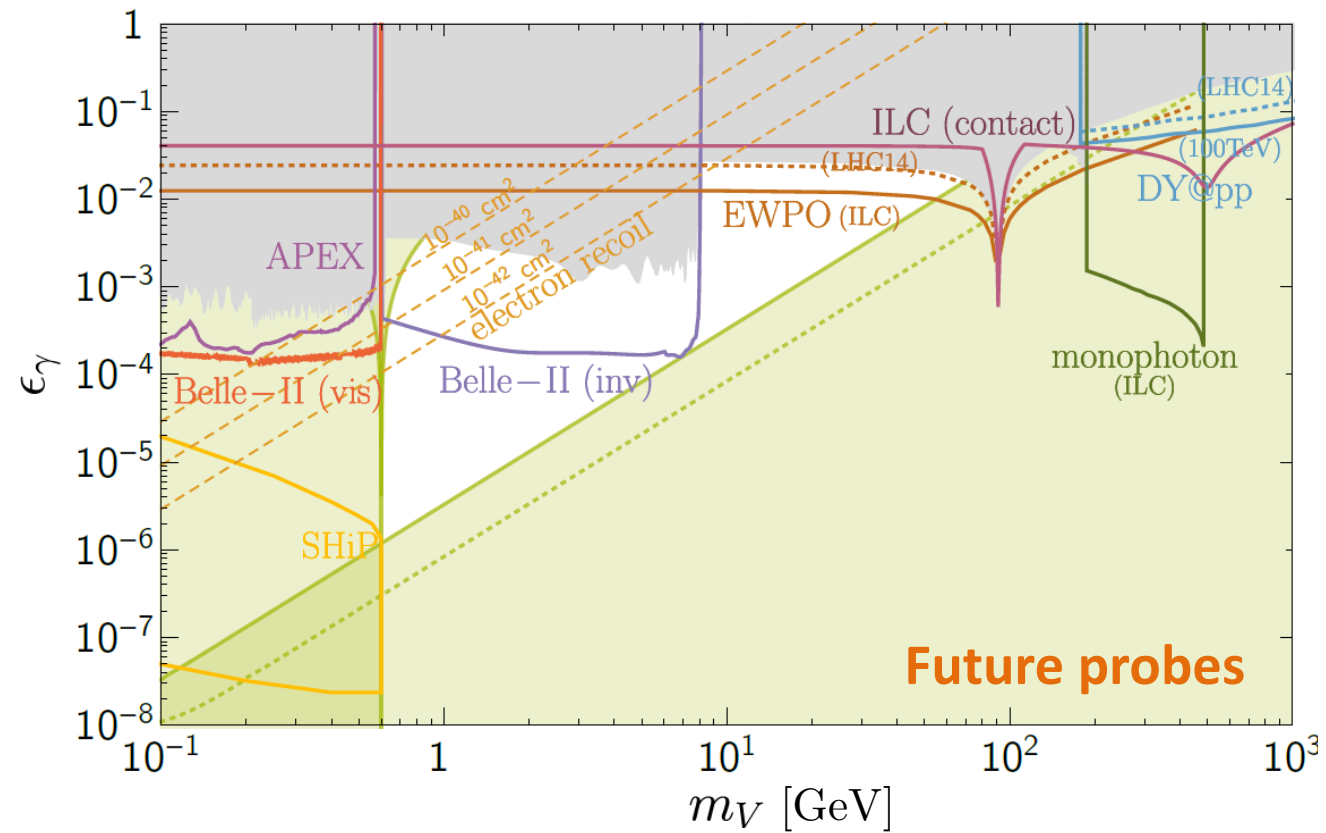


Predictive.

Dark Electromagnetism



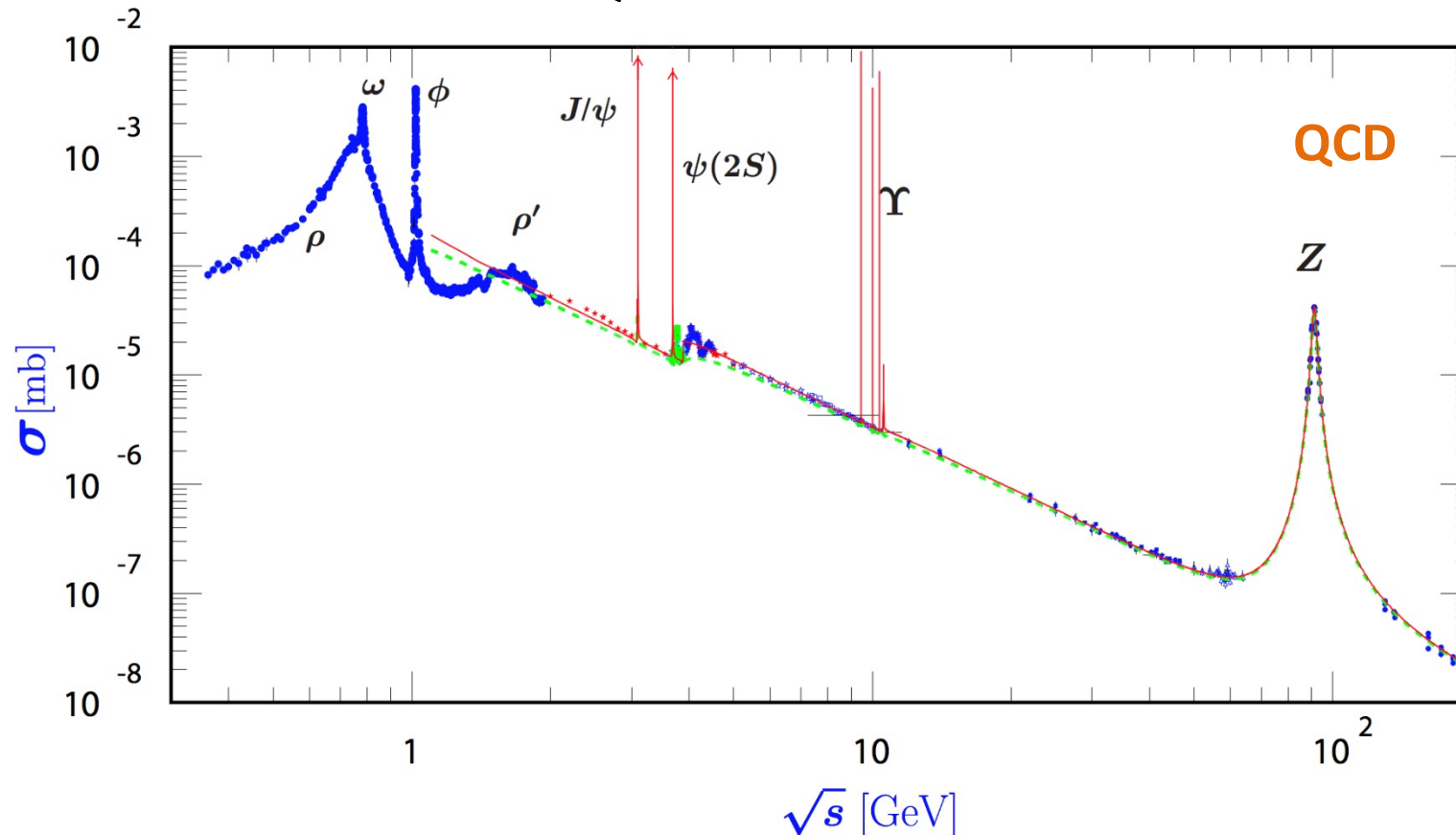
Dark Electromagnetism



Plot: YH, Kuflik, Murayama, 2016

Dark Spectroscopy

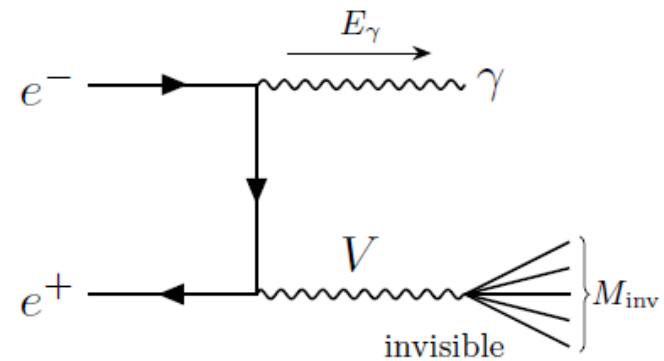
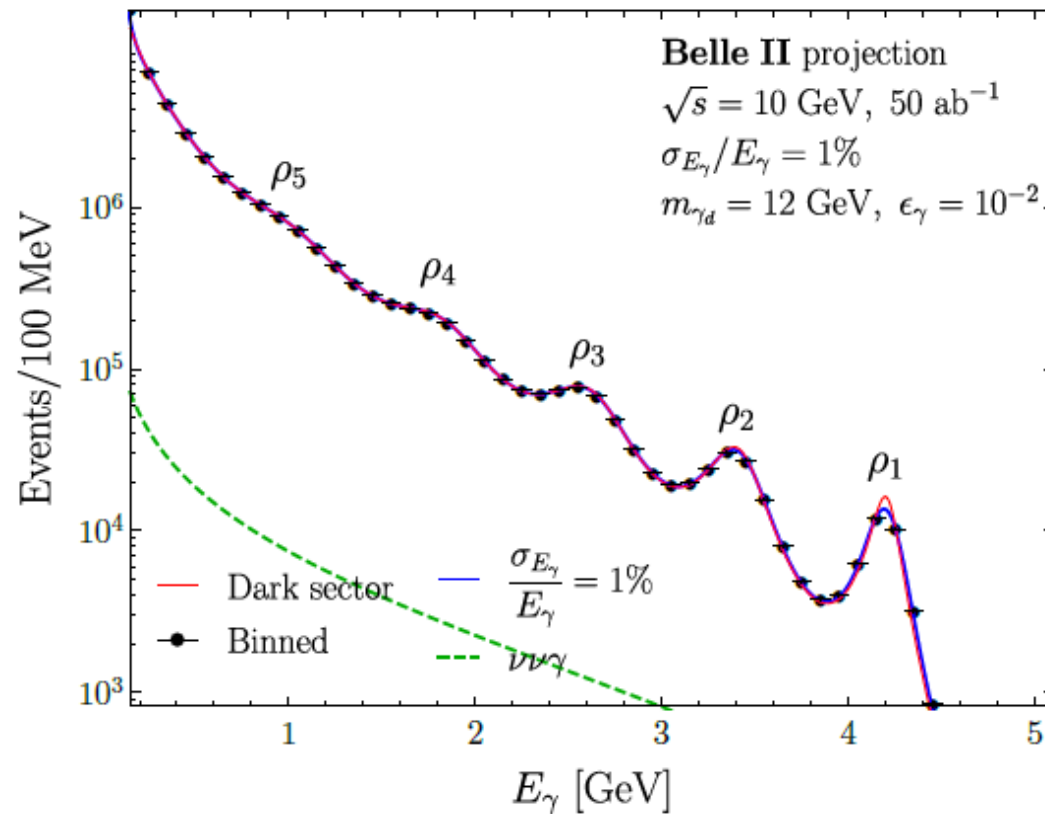
$e^+e^- \rightarrow \text{resonances}$: center of mass energy
traces the QCD resonance structure



Dark Spectroscopy

$$e^+e^- \rightarrow \gamma + \text{inv} :$$

Mono-photon traces the resonance structure of the dark sector



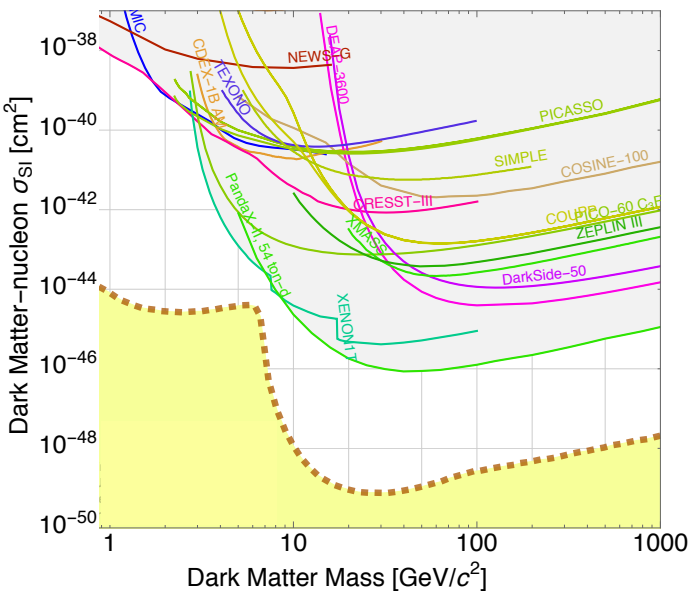
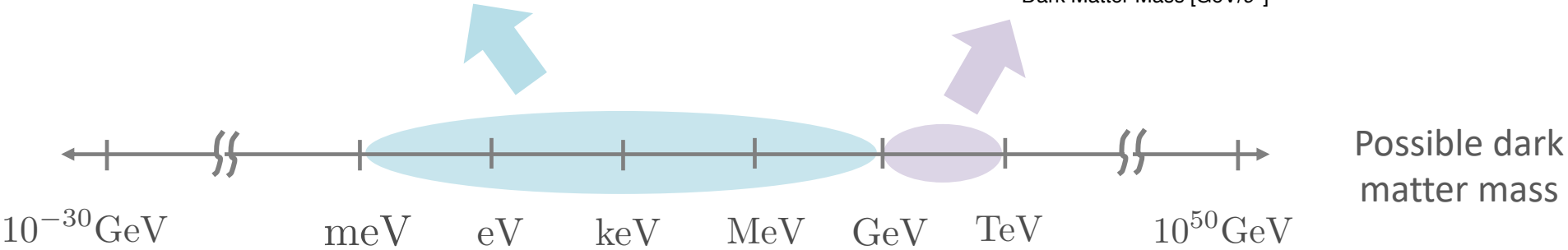
$$E_\gamma = \frac{\sqrt{s}}{2} \left(1 - \frac{M_{\text{inv}}^2}{s} \right)$$

Direct Detection

Explosion of new ideas
Superconductors, semiconductors,
2D targets, scintillators, polar materials,
superfluid helium...

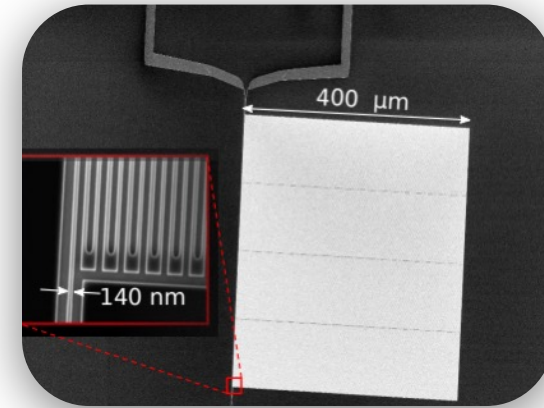
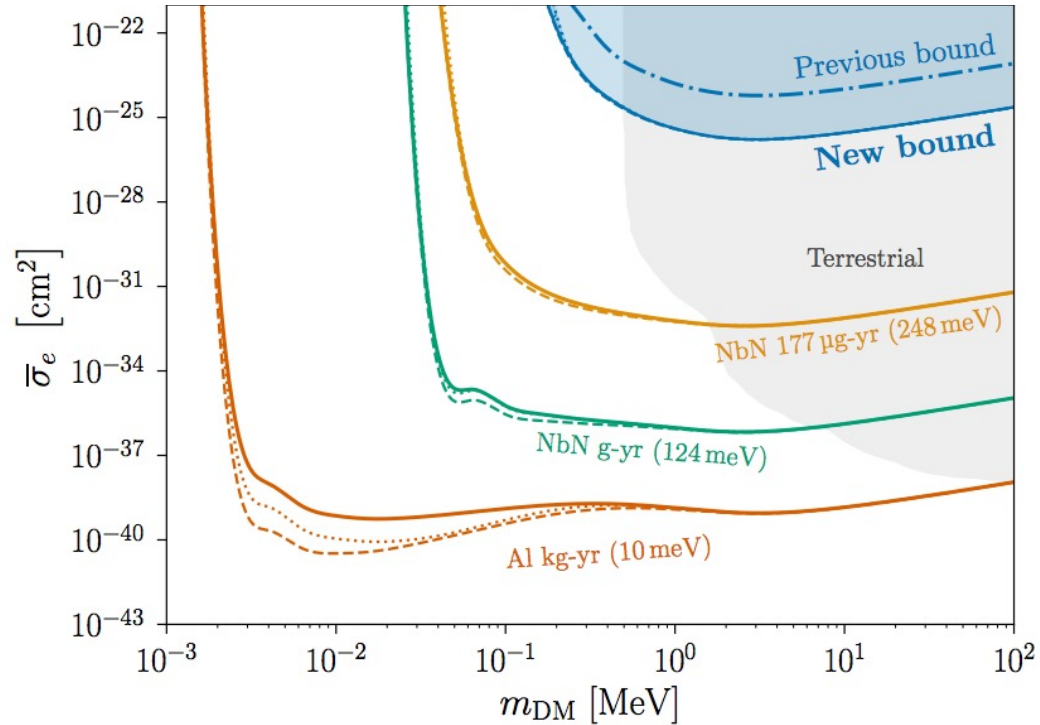


Becoming experimental reality



E.g. Superconductors

Superconducting Nanowire Single Photon Detector (SNSPD) (QIS)



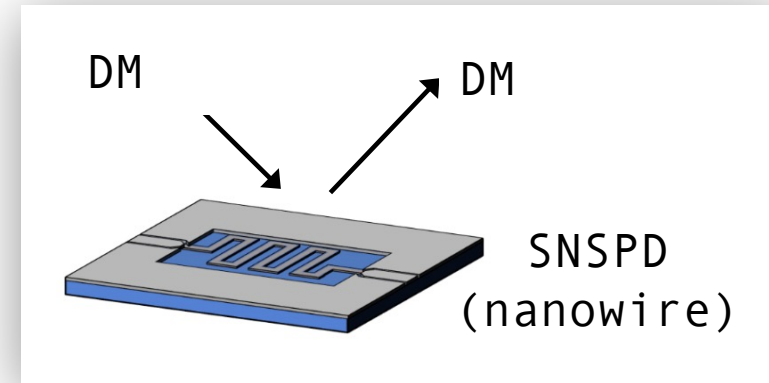
**WSi, 4.3ng, 0.73 eV threshold,
180 hours of data**

World leading limits on sub-MeV DM

YH, Charaev, Nam, Verma, Colangelo, Berggren, PRL 2019 + w/ Lehmann, PRD Editor's Choice 2022

E.g. Superconductors

Superconducting Nanowire Single Photon Detector (SNSPD) (QIS)



Newly forming interdisciplinary collaboration
(particle theory | condensed matter | DM experiment | quantum sensing)

Quantum sensor cryogenic search for Dark matter in Light mass range



Massachusetts
Institute of
Technology



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THE HEBREW UNIVERSITY OF JERUSALEM



Universität
Zürich^{UZH}

YH @ ICFA, DESY 2023

E.g. Superconductors

"How can
insert-new-technology-you-hear-of
be used for the dark matter hunt?"



Institute of
Technology



האוניברסיטה העברית בירושלים
THE HEBREW UNIVERSITY OF JERUSALEM



University of
Zürich^{UZH}

YH @ ICFA, DESY 2023

Outlook

- Lots of activity for dark matter
- Theory $\leftrightarrow$ experiment
- By no means exhausted...
- It's ok for an idea to seem crazy at first
- The best ideas might still be ahead

Prospects

INDY

WIMP

SIMP

Co-scattering

Co-decay

Asymmetric

Chain

Zombie

Different processes
Different properties
Different signatures

Axions

Freeze-in

Forbidden

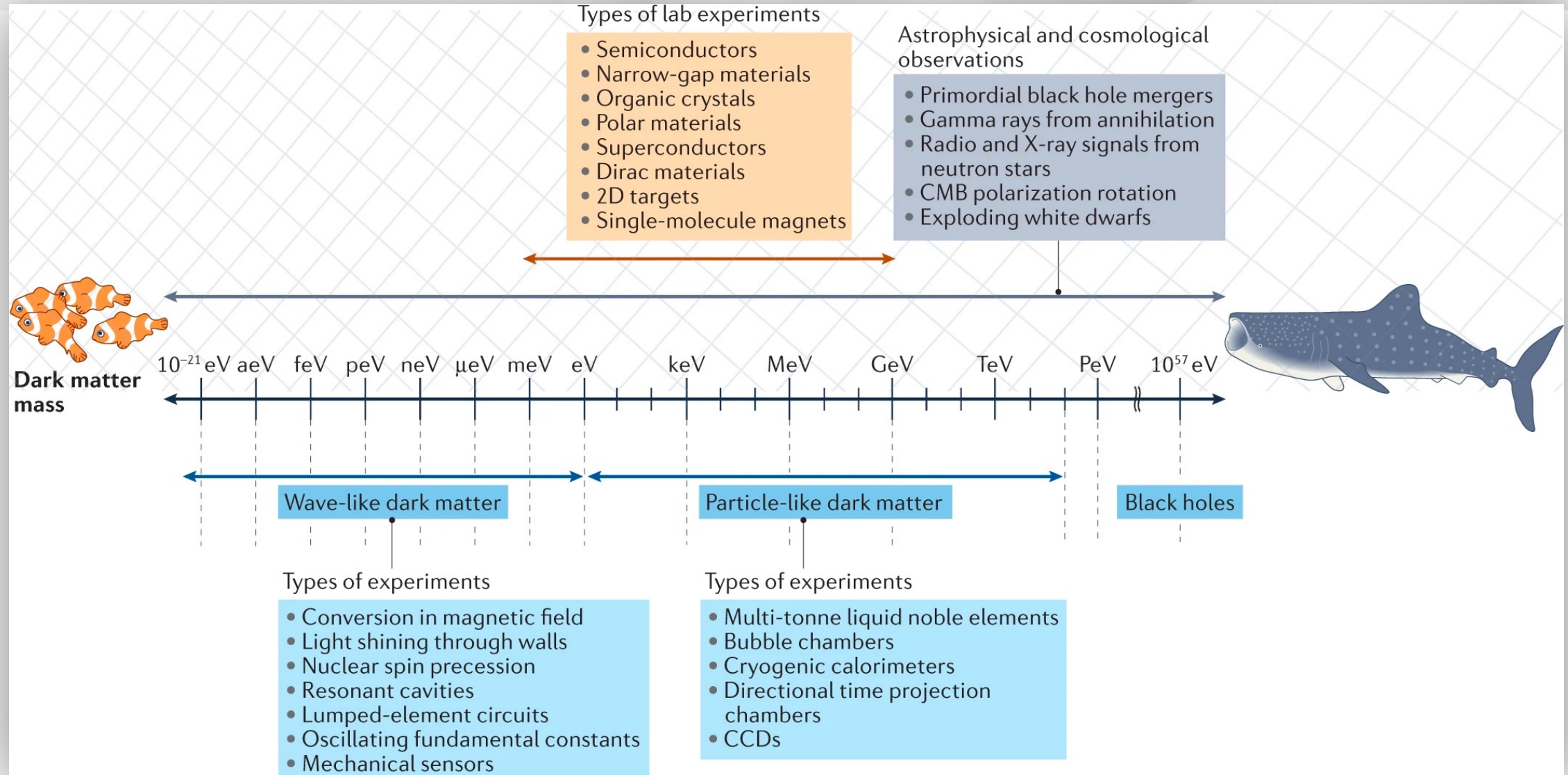
ELDER

Co-annihilation

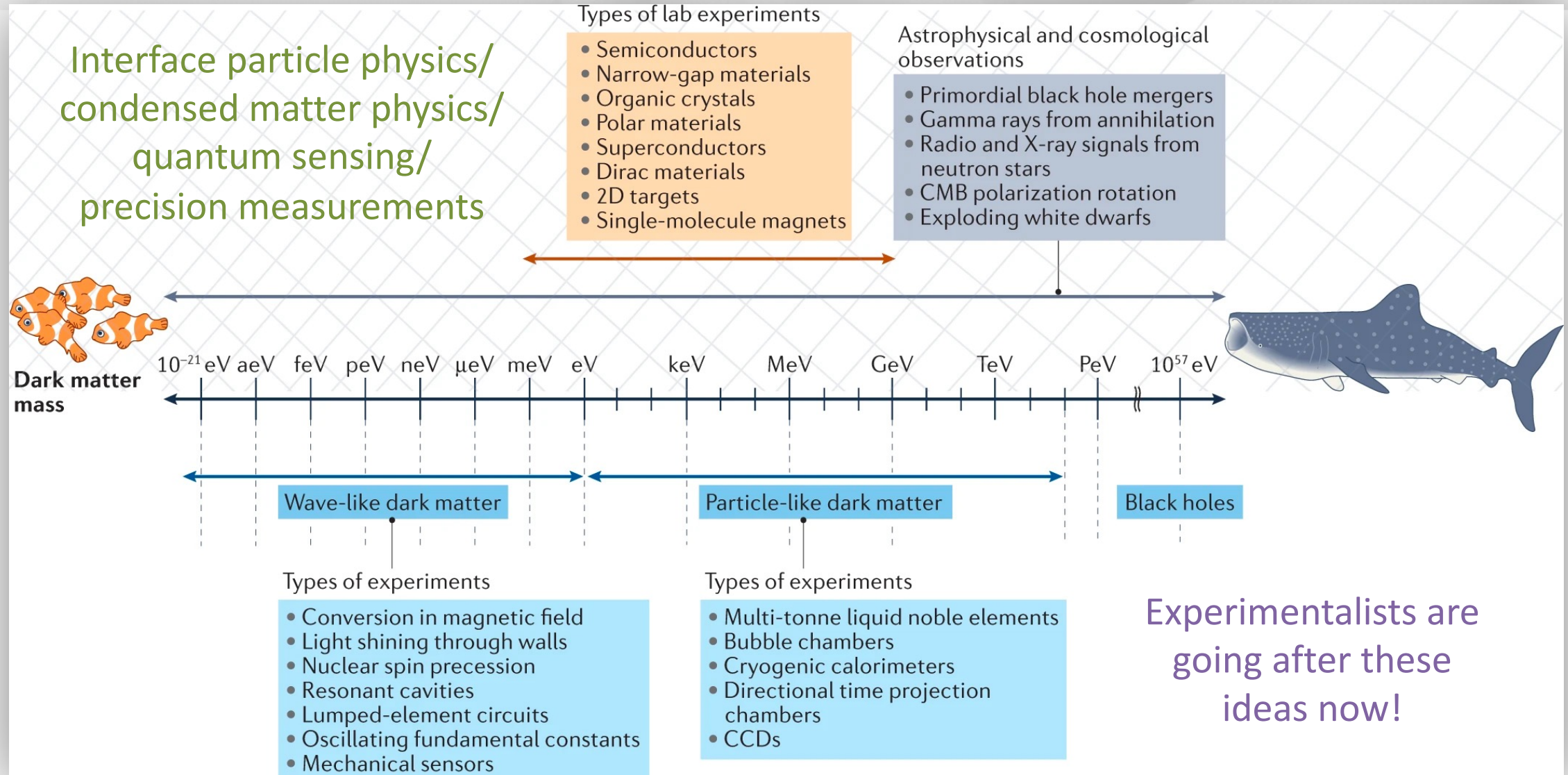
Squeeze-out

.... And many more...

Prospects



Prospects





Think outside the box.
(Fund outside the box.)



If you have any (crazy) new ideas,
please be in touch :-)

Thanks!

