## SUSY, BSM and Future Colliders at Snowmass 2021

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<sup>1</sup>DESY - FTX/SLB

FH particle physics discussion, DESY, April, 2023



CLUSTER OF EXCELLENCE
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#### **Outline**

- Outline
- Introduction
- The Energy Frontier at Snowmass
- Direct BSM at the Energy Frontier at Snowmass
- Snapshot of the contents of the BSM report
  - SUSY at high energy lepton colliders
    - SUSY with no loop-holes
    - SUSY@lepton colliders: non-Exclusion = Discovery
  - SUSY In The Briefing-book/Snowmass report
  - Z', ALPs, HNL, ...
- Conclusions



#### The frontiers at Snowmass

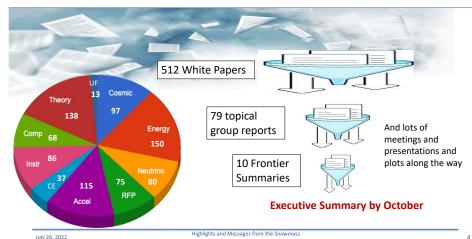
10 Frontiers	80 Topical Groups
Energy	Higgs Boson properties and couplings, Higgs Boson as a portal to new physics, Heavy flavor and top quark physics, EW Precision Phys. & constraining new phys., Precision QCD, Hadronic structure and forward QCD, Heavy Ions, Model specific explorations, More general explorations, Dark Matter at colliders
Neutrino Physics	Neutrino Oscillations, Sterile Neutrinos, Beyond the SM, Neutrinos from Natural Sources, Neutrino Properties, Neutrino Cross Sections, Nuclear Safeguards and Other Applications, Theory of Neutrino Physics, Artificial Neutrino Sources, Neutrino Detectors
Rare Processes	Weak Decays of b and c, Strange and Light Quarks, Fundamental Physics and Small Experiments. Baryon and Lepton Number Violation, Charged Lepton Flavor Violation, Dark Sector at Low Energies, Hadron spectroscopy
Cosmic	Dark Matter: Particle-like, Dark Matter: Wave-like, Dark Matter: Cosmic Probes, Dark Energy & Cosmic Acceleration: The Modern Universe, Dark Energy & Cosmic Acceleration: Cosmic Dawn & Before, Dark Energy & Cosmic Acceleration: Complementarity of Probes and New Facilities
Theory	String theory, quantum gravity, black holes, Effective field theory techniques, CFT and formal QFT, Scattering amplitudes, Lattice gauge theory, Theory techniques for precision physics, Collider phenomenology, SSM model building, Astro-particle physics and cosmology, Quantum information science, Theory of Neutrino Physics
Accelerator	Beam Physics and Accelerator Education, Accelerators for Neutrinos, Accelerators for Electroweak and Higgs Physics, Multi-TeV Colliders, Accelerators for Physics Beyond Colliders & Rare Processes, Advanced Accelerator Concepts, Accelerator Technology R&D: RF, Magnets, Targets/Source
Instrumentation	Quantum Sensors, Photon Detectors, Solid State Detectors & Tracking, Trigger and DAQ, Micro Pattern Gas Detectors, Calorimetry, Electronics/ASICS, Noble Elements, Cross Cutting and System Integration, Radio Detection
Computational	Experimental Algorithm Parallelization, Theoretical Calculations and Simulation, Machine Learning, Storage and processing resource access (Facility and Infrastructure R&D), End user analysis
Underground Facilities	Underground Facilities for Neutrinos, Underground Facilities for Cosmic Frontier, Underground Detectors
Community Engagement	Applications & Industry, Career Pipeline & Development, Diversity & Inclusion, Physics Education, Public Education & Outreach, Public Policy & Government Engagement
Snowmass Early Career	Snowmass Early Career to represent early career members and promote

7/17/22

Snowmass Greeting, July 17, JB



## The Snowmass process: Getting there





- Will touch on the uptake on BSM from 'our" Frontier
  - The Energy Frontier
- ... even though also
  - The Neutrino Frontier
  - The Cosmic Frontier
  - The Rare Processes Frontier
- ... of course also includes BSM aspects.
- I won't talk about the "How?" frontiers (Instrumentation, Accelerator, Computing, ...), sorry.

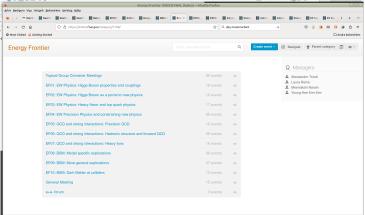
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#### The Energy Frontier at Snowmass

Lots of meetings for  $\sim$  two years *before* the final Seattle work-shop



269 open meetings!

#### Studied projects

## Higgs-boson factories (up to 1 TeV c.o.m. energy)

Collider	Type	$\sqrt{s}$	P[%]	$\mathcal{L}_{\mathrm{int}}$	Start Date	
			$e^-/e^+$	$\mathrm{ab}^{-1}/\mathrm{IP}$	Const.	Physics
HL-LHC	pp	14 TeV		3		2027
ILC & C <sup>3</sup>	ee	250 GeV	$\pm 80 / \pm 30$	2	2028	2038
		350 GeV	$\pm 80 / \pm 30$	0.2		
		500 GeV	$\pm 80 / \pm 30$	4		
		1 TeV	$\pm 80/ \pm 20$	8		
CLIC	ee	380 GeV	±80/0	1	2041	2048
CEPC	ee	$M_Z$		50	2026	2035
		$2M_W$		3		
		240 GeV		10		
		360 GeV		0.5		
FCC-ee	ee	$M_Z$		75	2033	2048
		$2M_W$		5		
		240 GeV		2.5		
		$2 M_{top}$		0.8		
$\mu$ -collider	μμ	125 GeV		0.02		

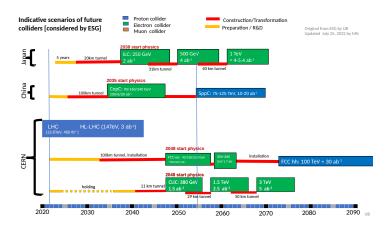
## Multi-TeV colliders (> 1 TeV c.o.m. energy)

Collider	Type	$\sqrt{s}$	P[%]	$\mathcal{L}_{\mathrm{int}}$	Start Date	
			. e-/e+	$\mathrm{ab^{-1}/IP}$	Const.	Physics
HE-LHC	pp	27 TeV		15		
FCC-hh	pp	100 TeV		30	2063	2074
SppC	pp	75-125 TeV		10-20		2055
LHeC	ер	1.3 TeV		1		
FCC-eh		3.5 TeV		2		
CLIC	ee	1.5 TeV	±80/0	2.5	2052	2058
		3.0 TeV	±80/0	5		
μ-collider	$\mu\mu$	3 TeV		1	2038	2045
		10 TeV		10		

Large Experiments Panel @CSS, Seattle, July 26, 2022

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#### The Energy Frontier: Timelines



#### The Energy Frontier: Timelines



- · Find a contact lab(s)
- · Successful R&D and feasibility demonstration for CCC and Muon Collider
- · Evaluate CCC progress in the international context, and consider proposing an ILC/CCC [ie CCC used as an upgrade of ILC1 or a CCC only option in the US.
- · International Cost Sharing
- · Consider proposing hosting ILC in the US.

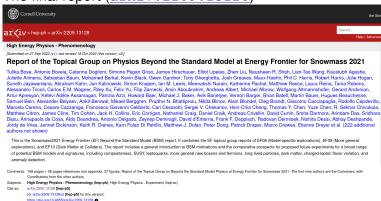
#### Direct BSM at the Energy Frontier at Snowmass

	October 2020
	Oct 29 EF08: BSM: Model specific explorations Biweekly Meeting
July 2021	Oct 15 EF08: BSM: Model specific explorations Biweekly Meeting
Jul 22 EF08: BSM: Model specific explorations Biweekly Meeting	September 2020
Jul 08 EF08: BSM: Model specific explorations Biweekly Meeting	Sop 17 EF08: BSM: Model specific explorations Biweekly Meeting
May 2021	Book Sep 03 EF08: BSM: Model specific explorations Biweekly Meeting
May 14 EF08: BSM: Model specific explorations: Anomalies (g-2, etc) chat	August 2020
December 2020	
Dec 10 EF08: BSM: Model specific explorations Biweekly Meeting	June 2020
November 2020	Jun 25 EF08: BSM: Model specific explorations Biweekly Meeting
Nov 12 EF08: BSM: Model specific explorations Biweekly Meeting	Jun 11 EF08: BSM: Model specific explorations Biweekly Meeting
Nov 11 EF08 : Snowmass pMSSM scans	May 2020
■ Nov 04 EF08 : Snowmass pMSSM scans	May 28 EF08: BSM: Model specific explorations - Kick off meeting with community

15 meetings in EF08 (Direct BSM in specific models - my focus).

#### Direct BSM at the Energy Frontier at Snowmass

The final report (arXiv:2209.13128)



300+ authors, most of whom really did contribute (talks, White papers, discussions): No tourists!

#### Direct BSM: SUSY

#### In this talk: Concentrating on

- SUSY:
  - The most complete theory of BSM.
  - Most studied model with serious simulation: In most cases, full simulation, with all SM backgrounds, all beam-induced backgrounds included.
  - Serves as a boiler-plate for BSM: almost any new topology can be obtained in SUSY...
  - Under some stress(?) by LHC. However, in particular lepton colliders offers
    - Complete coverage of Compressed spectra the most interesting case.
  - Loop-hole free searches.
- + A few slides on non-SUSY BSMs...

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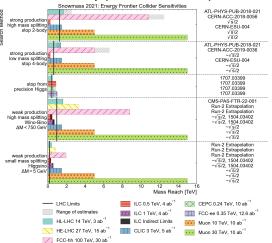
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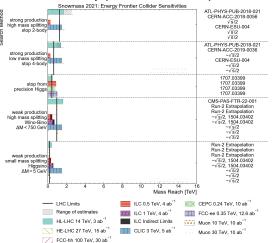
### SUSY in the Energy Frontier report

#### SUSY summary plot in the EF report



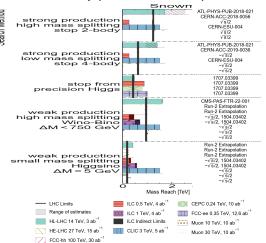
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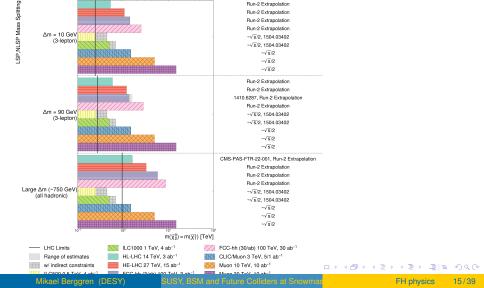
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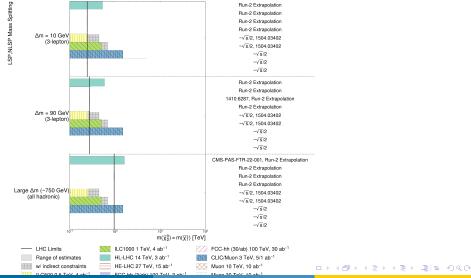


#### Details from the BSM topical group report: Winos

Bun-2 Extrapolation Run-2 Extrapolation Run-2 Extrapolation

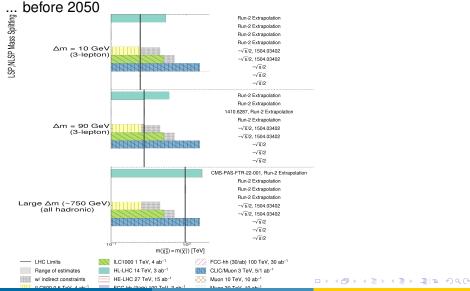


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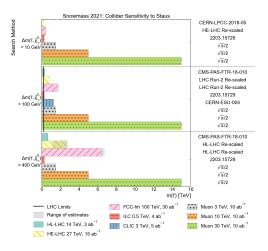


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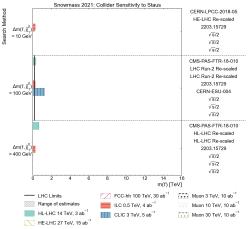


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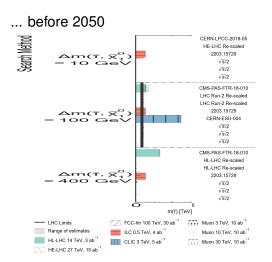


## Details from the BSM topical group report: $\tilde{\tau}$ :s

#### ... before 2050



### Details from the BSM topical group report: $\tilde{\tau}$ :s



## SUSY at high energy lepton colliders - ILC as an example (but relevant for C<sup>3</sup>, HELEN, CLIC, ... )

- $e^+e^-$  collider with  $E_{CMS} = 250 500 (-1000++)$  GeV, and polarised beams
- $e^+e^-$  means EW-production  $\Rightarrow$  Low background.
  - Detectors w/  $\sim 4\pi$  coverage.
  - Rad. hardness not needed: only few % X<sub>0</sub> in front of calorimeters.
  - No trigger
- $e^+e^-$  means colliding point-like objects  $\Rightarrow$  initial state known
- 22 year running  $\rightarrow$  2 ab<sup>-1</sup> @ 250 GeV + 4 ab<sup>-1</sup> @ 500 GeV.
- Construction under political consideration in Japan.



# SUSY: What *do* we know? And why does that give lepton colliders an edge?

Naturalness, hierarchy, DM, g-2 all prefer light electroweak sector.

- Except for 3rd gen. squarks, the coloured sector doesn't enter the game.
- Many models and the global set of constraints from observation points to a compressed spectrum.
- So, most sparticle-decays are via cascades, with small  $\Delta(M)$  at the end.
- For this, current LHC limits are for specific models. LEP2 sets the scene.

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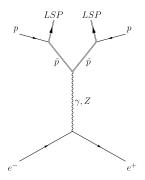
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### SUSY@lepton colliders: Loop-hole free searches

- All is known for given masses, due to SUSY-principle: "sparticles couples as particles".
- This doesn't depend on the SUSY breaking mechanism!
- Obviously: There is one NLSP, and it must have 100 % BR to it's SM-partner and the LSP.

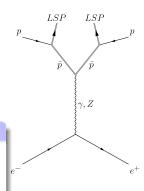


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#### So, at ILC:

- Model independent exclusion/ discovery reach in M<sub>NLSP</sub> – M<sub>LSP</sub> plane.
- Repeat for all NLSP:s.
- Cover entire parameter-space in a few plots
- No fine-print!

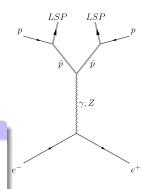


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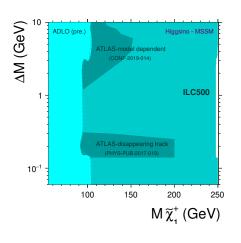
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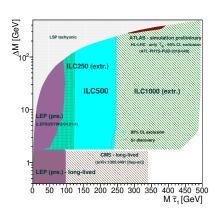
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## ILC projection for Higgsino or $\tilde{\tau}$ NLSP

From arXiv:2002.01239

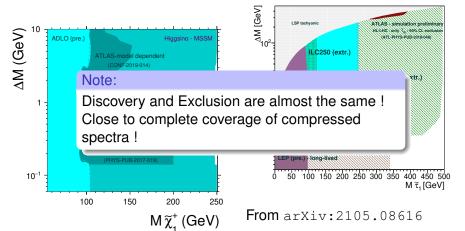




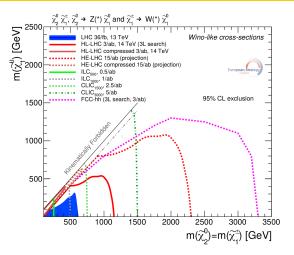
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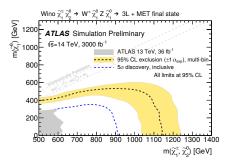
# SUSY In The Briefing-book ( $\approx$ Snowmass) : Bino LSP (ie. large $\Delta_M$ )



(This is referred to, and not updated @ Snowmass)

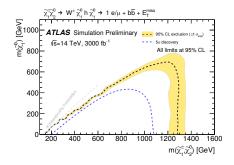
### SUSY In The Briefing-book: Bino LSP - Sources

- From PHYS-PUB-2018-04 (ATLAS HL-LHC projection). Then extrapolated (up and down)
- Note that the BB curve is exclusion, not discovery!
- This is for the best decay mode!
- The other decay mode
- Better at  $M_{LSP}$ =0, weaker at lower  $\Delta_M$ .

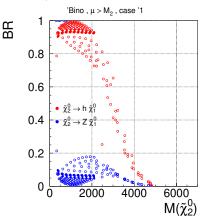


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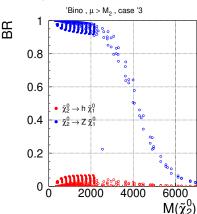
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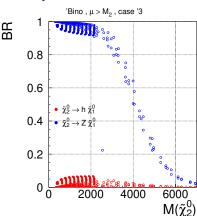
- Vary relative signs of μ, M<sub>1</sub>, and M<sub>2</sub>, for μ > M<sub>2</sub>
- Conclusion: Whether the Z or the H decay-mode of  $\tilde{\chi}_2^0$  dominates is pure speculation and
- The exclusion-region is the intersection of the two plots, not the union!



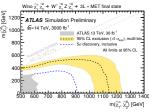
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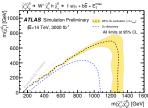


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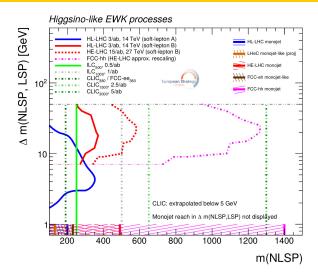


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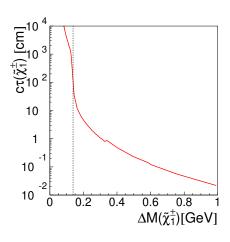


### SUSY In The Briefing-book: Wino/Higgsino LSP

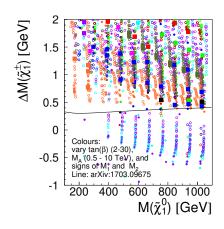


(This, too, is refered to, but also gets an update @ Snowmass)

- $c\tau$  vs.  $\Delta(M)$  for charginos. Note where 1 cm is...
- Higgsino LSP. The line is the absolute limit mentioned in the BB.
- Let other parameters vary, any signs,  $M_1$  and  $M_2$  close to  $\mu$  .... Note that the LSP often would be the  $\tilde{\chi}_1^{\pm}$ !
- Reason: 1703.09675 considers *only SM* effects on the mass-splitting, ie. that  $M_1$  and  $M_2 >> \mu$
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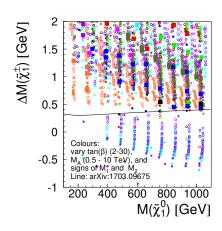


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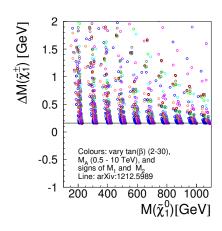


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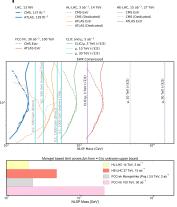


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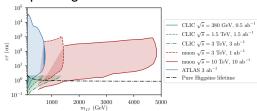


### Wino/Higgsino LSP: Snowmass update

#### Leptons and Mono-X

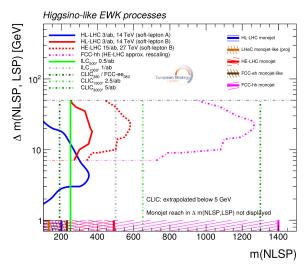


#### Disapearing tracks



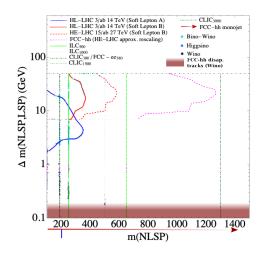
NB. Irrelevant for lepton colliders - The standard search gives stronger limits.

### SUSY In The Briefing-book: Wino/Higgsino LSP



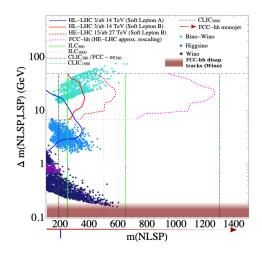
So: Disappearing tracks exclusion is actually off the scale!

### SUSY In The Briefing-book: Re-boot



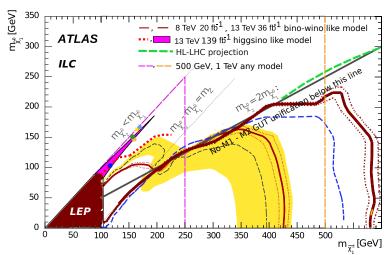


### SUSY In The Briefing-book: Re-boot



With models that are consitent with g-2 and no over-production of DM From arXiv: 2103.13403.

#### SUSY bosinos - All-in-one

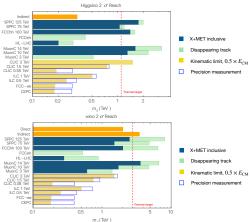


ATLAS Eur Phys J C 78,995 (2018), Phys Rev D 101,052002 (2020), arXix:2106.01676;

ATLAS HL-LHC ATL-PHYS-PUB-2018-048: ILC arxiv: 2002.01239: LEP LEP LEPSUSYWG



#### Bosinos as DM

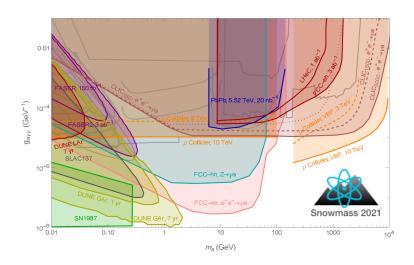


The thermal targets are assuming that the Wino/Higgsino is *all* of DM. And the 'Disapearing tracks' limits only apply in the pure Wino/Higgsino case, remember.

Machine	Туре	√s (TeV)	JL dt (ab-1)	Source	Z' Model	5σ (TeV)	95% CL (TeV)
HL-LHC				R.H.	Z' <sub>SSM</sub> → dijet	4.2	5.2
	pp	14	3	ATLAS	Z' <sub>SSM</sub> → l+ l-	6.4	6.5
				CMS	Z' <sub>SSM</sub> → l+ l-	6.3	6.8
				EPPSU*	Z' <sub>Unix</sub> (g <sub>2</sub> '=0.2)		6
ILC250/	e+ e-	0.25	2	ILC	Z' <sub>SSM</sub> → f+ f-	4.9	7.7
CLIC380/ FCC-ee				EPPSU*	Z' <sub>Unix</sub> (g <sub>2</sub> '=0.2)	-	7
HE-LHC/ FNAL-SF	pp	27	15	EPPSU*	Z' <sub>Unix</sub> (g <sub>2</sub> '=0.2)		11
				ATLAS	Z' <sub>SSM</sub> → e <sup>+</sup> e <sup>-</sup>	12.8	12.8
ILC	e+ e-	0.5	4	ILC	$Z'_{SSM} \rightarrow f^+ f^-$	8.3	13
				EPPSU*	Z' <sub>Unix</sub> (g <sub>Z</sub> '=0.2)		13
CLIC	e+ e-	1.5	2.5	EPPSU*	Z' <sub>Unix</sub> (g <sub>2</sub> '=0.2)	-	19
Muon Collider	μ+ μ-	3	1	IMCC	Z' <sub>Unix</sub> (g <sub>z</sub> '=0.2)	10	20
ILC	e+ e-	1	8	ILC	Z' <sub>SSM</sub> → f+ f-	14	22
				EPPSU*	Z' <sub>Unix</sub> (g <sub>2</sub> '=0.2)		21
CLIC	e+ e-	3	5	EPPSU*	Z' <sub>Unix</sub> (g <sub>Z</sub> '=0.2)		24
FCC-hh	рр	100	30	R.H.	$Z'_{SSM} \rightarrow dijet$	25	32
				EPPSU*	Z' <sub>Unix</sub> (g <sub>z</sub> '=0.2)		35
				EPPSU	Z' <sub>SSM</sub> → l+ l-	43	43
Muon Collider	μ+ μ-	10	10	IMCC	Z' <sub>Unix</sub> (g <sub>Z</sub> '=0.2)	42	70
VLHC	pр	300	100	R.H.	Z' <sub>SSM</sub> → dijet	67	87
Coll. In the Sea	pp	500	100	R.H.	Z' <sub>SSM</sub> → dijet	96	130

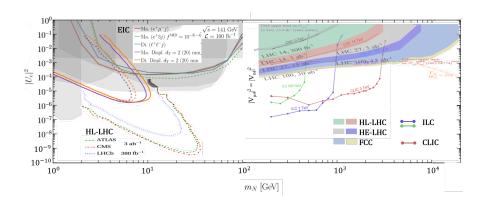
Already the Higgs factories are expected to go beyond the HL-LHC reach....

### Z', ALPs, HNL, ...





### Z', ALPs, HNL, ...



### The Energy Frontier: The vision

#### The Energy Frontier vision in a nutshell

#### It is essential to

- · Complete the HL-LHC program,
- Start now a targeted program for <u>detector R&D for Higgs Factories</u>
- Support a fast start of the construction of a Higgs factory
- Ensure the long-term viability of the field by <u>developing a multi-TeV energy</u> frontier facility such as a muon collider or a hadron collider.

Support to AF, CEF, CompF, IF, and TF is crucial to the realization of the EF vision



### The Energy Frontier: in 5-10-15 years

#### **EF Resources and Timelines**

#### Five year period starting in 2025

- o Prioritize HL-LHC physics program, including auxiliary experiments
- Establish a targeted e+e- Higgs Factory detector R&D for US participation in a global collider
- o Develop an initial design for a first stage TeV-scale Muon Coll. in the US (pre-CDR)
- o Support critical detector R&D towards EF multi-TeV colliders

#### Five year period starting in 2030

- o Continue strong support for HL-LHC program
- Support and advance construction of an e+e- Higgs Factory
- Demonstrate principal risk mitigation and deliver CDR for a first-stage TeV-scale Muon Coll.

#### ➤ After 2035

- Support continuing HL-LHC physics program to the conclusion of archival measurements
- Begin and support the physics program of the Higgs Factories
- o Demonstrate readiness to construct and deliver TDR for a first-stage TeV-scale Muon Coll.
- o Ramp up funding support for detector R&D for EF multi-TeV colliders



18

- Very intense 10 days with no day off.
- Great organisation:
  - Mornings with Frontier/topical group parallels (Meaning that I was almost only following EF-BSM parallels)
  - Afternoons with plenaries each frontier got its, non-shared, plenary.
  - Also specific cross-frontier parallels eg. Energy/Accelerator
- 735 on-site participants (+654 remote). All having a 2 hour lunch on University Street, just off-campus ⇒ lots of opportunities for off-the-record cross-frontier discussions.
- About 35 Europeans, 10 Japanese on-site.
- Lab directors (US of course, but also CERN, KEK, IHEP, Triumf) , APS, ICFA, STFC and IDT chairs present ....



- The Americans didn't "make the Wave" about FCC more noted with interest the activities in Europe.
- Fabiola's sobering presentation on the FCC time-line probably contributed to that.
- Surprises:
  - US wants to get back with a domestic Energy Frontier facility.
  - ILC in US on the table!
  - Great revival of the interest in the muon collider
  - Little mention of Plasma Wakefields, at least outside the AF
  - And: The closest to a mention of the war in Ukraine in any talk was a mention of current "supply-chain difficulties" in the DoE talk - quite

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#### What now?

- The P5 is working. Inform yourself on
  - The (beutifully old-school) P5 Web page
- Note that Beate is the only european on the pannel!
- Join the P5 town-hall meeting next week. It is the one devoted to the Energy Frontier.
  - P5 Town Hall Meeting at BNL
- P5 is asked to deliver its report to HEPAP by end of summer.



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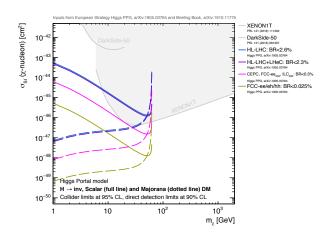


## Thank You!



#### Backup

# **BACKUP SLIDES**



ILD fast detector simulation studies: Selectrons in a co-annihilation model ( $_{\text{EPJC}}$  76,183 (2016)), after:

- 5 fb<sup>-1</sup>  $\approx$  1 week
  - 500 fb<sup>-1</sup>  $\approx$  2 years.

Will never be in "3  $\sigma$  limbo"!

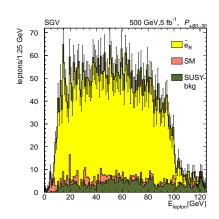
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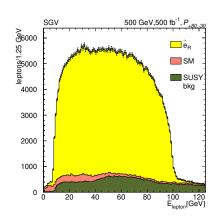
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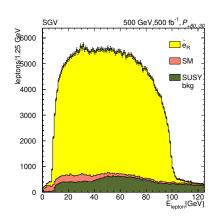
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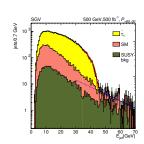
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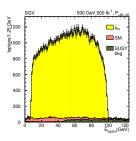
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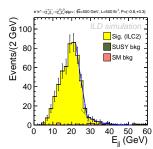
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- ... and typical neutralino signal, higgsino-LSP model, with moderate ΔM (FullSim) (Phys Rev D 101,095026 (2020))
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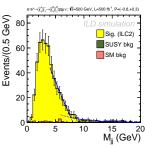
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(EPJC 73,2660 (2013))
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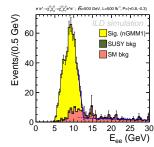
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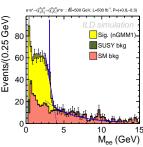




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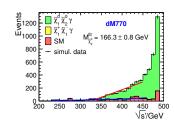
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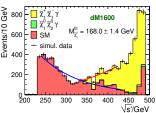




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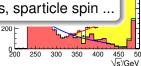




#### ILD detector simulation studies:

- Typical slepton signal ( $\tilde{\tau}$  and  $\tilde{\mu}$ ), in a co-annihilation model (FastSim). (EPJC
- Typical chargin In all cases:
- ... and typical r signal, higgsing with moderate (Phys Rev D 101,0950
- SUSY masses to sub-percent
- Cross-sections to few percent
- Also: Branching fractions, mixing angles, sparticle spin ...
- Typical chargino/neutralino signal, higgsino-LSP model, with very low ΔM (Fast/FullSim).

```
(EPJC 73,2660 (2013))
```



66.3 ± 0.8 GeV

√s'/GeV

simul data

300

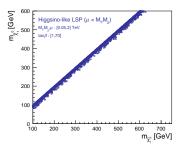
600 400

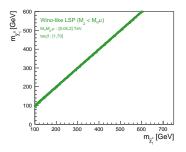
- Higgsino or Wino LSP:
  - If the LSP is Higgsino or a Wino, several other bosinos must be close to the LSP.
  - ⇒ Compressed spectrum.
    - In addition: if the LSP is higgsino: Natural SUSY:

 $\bullet \ m_Z^2 = 2 \frac{m_{H_U}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - 2 |\mu|^2$ 

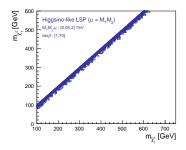
• Low fine-tuning  $\Rightarrow \mu = \mathcal{O}(m_Z)$ 

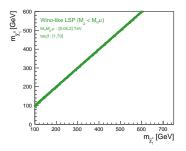
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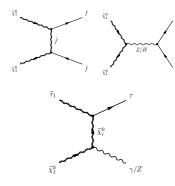


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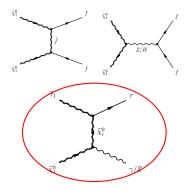




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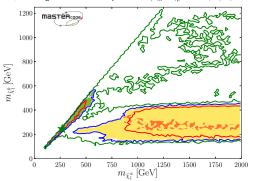


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# Why compressed spectra? Global fits

pMSSM11 fit by Mastercode to LHC13/LEP/g-2/DM(=100% LSP)/precision observables (arXiv:1710.11091):



$$M_{{ ilde \chi}_1^\pm}$$
 -  $M_{{ ilde \chi}_1^0}$  plane

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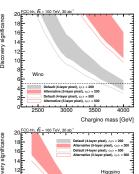
Low  $\Delta(M)$ !

 $M_{\widetilde{\chi}_1^\pm}$  -  $M_{\widetilde{\chi}_1^0}$  plane

## SUSY In The Briefing book: Wino/Higgsino LSP -Sources

(Don't look at the pink curves - they correspond to a detector that is never considered anywhere else i the CDR)

- The "Disappearing tracks" was done by FCChh (in the CDR)
  - FCChh-detector (better than ATLAS in this case: first layer of VD closer.)
  - FCChh-ish PU (but still to small: 500 vs. CDR number 955)
  - For higgsinos: Only just



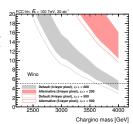
1000

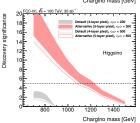


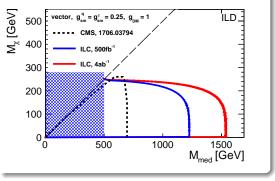
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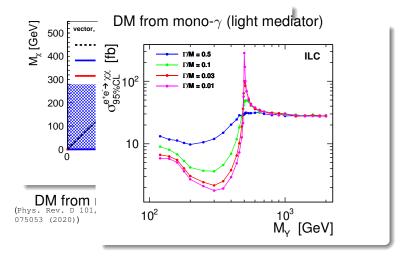
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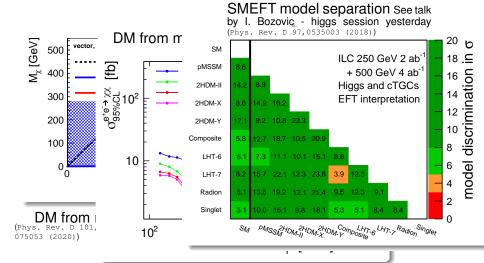


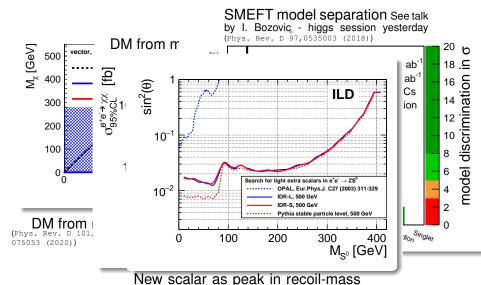




 $\underset{\text{(Phys. Rev. D 101,}}{\text{DM from mono-}\gamma} \text{ (EFT)}$ 



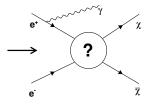




(arXiv:2005.06265)

# Only WIMPs

- What if this is the only accessible NP?
- Search for direct WIMP pair-production at collider: Need to make the invisible visible:
  - Require initial state radiation which will recoil against "nothing" ⇒ Mono-X search.
  - At ILC:  $e^+e^- \rightarrow \chi \chi \gamma$ , ie. X is a  $\gamma$



- ILC simulation studies: arXiv:1206.6639v1, A. Chaus, Thesis, M. Habermehl, Thesis,in preparation.
- Model-independent Effective operator approach to "?"
  - Analyse as an effective four-point interaction. Strength =  $\Lambda$ .
    - Allowable if direct observation the mediator is beyond reach. Mostly true at ILC, but not at LHC!
  - Write down all possible Lorentz-structures of the operators.
  - Exclusion regions in  $M_{\chi}/\Lambda$  plane, for each operator.



## ILC and LHC exclusion

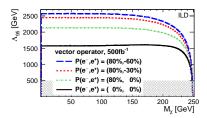
- Examples:
  - Vector operator ("spin independent"), Note how
- useful beam-polarisation is!

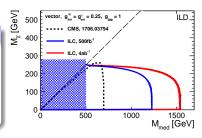
   At LHC, EffOp can't be used

  ⇒ use "simplified models"
- Need to translate  $\Lambda$  to  $M_{med}$ :  $M_{med} = \sqrt{g_{SM}g_{DM}}\Lambda$

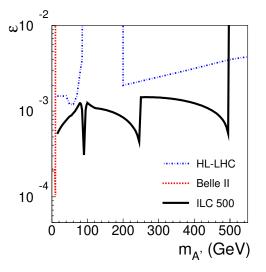
#### ILC/LHC complementarity

- LHC: coupling to hadrons,
   ILC: coupling to leptons.
- LHC has best  $M_{\chi}$  reach, ILC best  $M_{med}$  reach





## Dark photons



(Theory level estimate - FullSim in the works...)

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  - Caveat: also CP-conservation. The experimental implication of CP violation needs study
- sfermions not NLSP (idem, except  $\tilde{\tau}$  but even worse for pp...)
- Then: LSP is Bino, Wino, or Higgsino (more or less pure), same for the NLSP
- $M_1$ ,  $M_2$  and  $\mu$  are the main-players.
- ullet Consider any values, and combinations of signs, up to values that makes the bosinos out-of-reach for any new facility  $\sim$  a few TeV.
- Also vary other parameters ( $\beta$ ,  $M_A$ ,  $M_{sfermion}$ ) with less impact.
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## Aspects of the spectrum

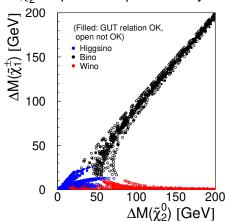
Another angle:  $\Delta(M)$  for  $\tilde{\chi}_1^{\pm}$  vs. that of  $\tilde{\chi}_2^0$ : Important experimentally

• Three regions:

• Bino: Both the same, but can be anything.

 $\begin{tabular}{ll} \bullet & \mbox{Wino: } \Delta_{\widetilde{\chi}_1^\pm} \mbox{ small, while } \Delta_{\widetilde{\chi}_2^0} \\ \mbox{ can be anything.} \\ \end{tabular}$ 

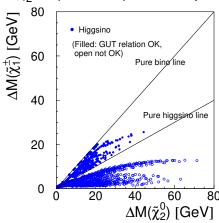
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- But note, seldom on the "Higgsino line", ie. when the chargino is exactly in the middle of mass-gap between the first and second neutralino



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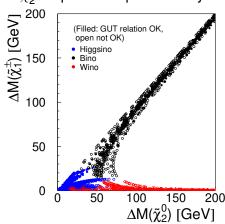
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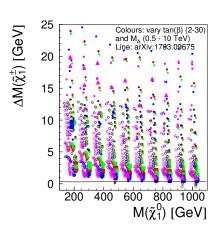
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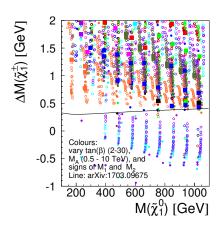


#### Higgsino LSP.

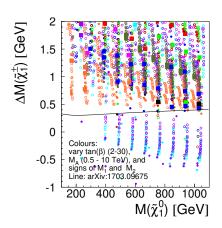
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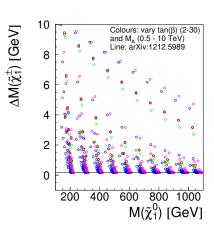
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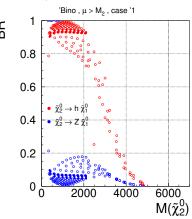
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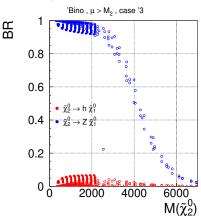
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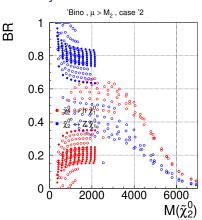
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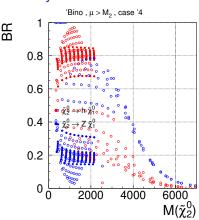
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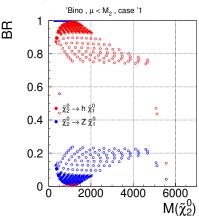
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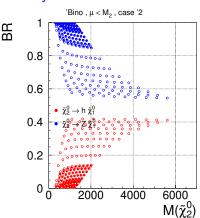
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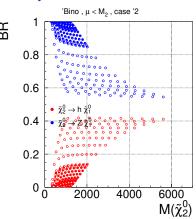
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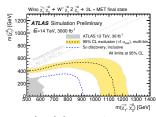
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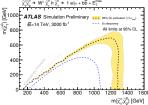


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  - Extendability in energy and polarised beams
  - Detectors factors more precise, hermetic, and with no need for triggering
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ILC input to the european strategy update

The Potential of the ILC for Discovering New Particles and references therein ...

# Thank You!