

Holographic Correlators for all Λ s

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AdS-CFT

Quantum Gravity
in AdS_{d+1}

=

(non-gravitational)
CFT in \mathbb{M}^d

Observables ?!



Correlation functions

Constrained non-perturbatively by
the **Conformal Bootstrap**:

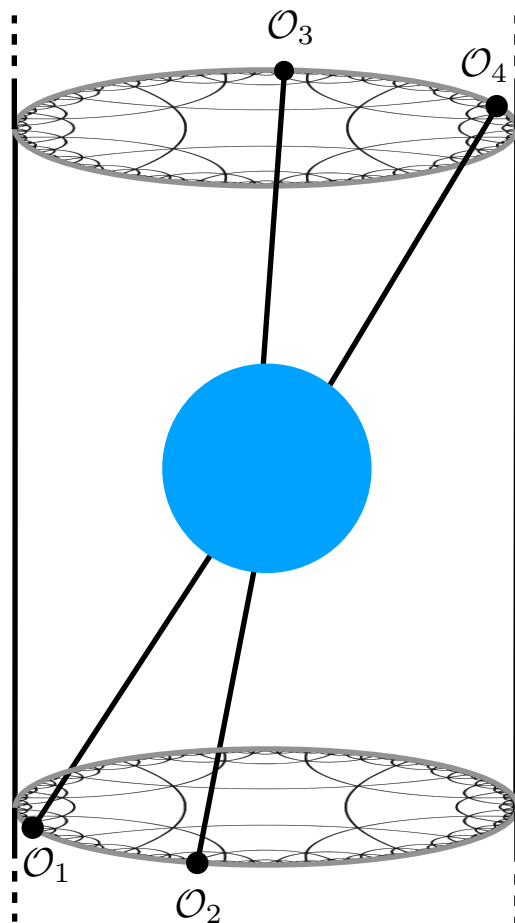
- Conformal symmetry
- Unitarity
- Associative OPE

$$(\mathcal{O}_1 \mathcal{O}_2) \mathcal{O}_3 = \mathcal{O}_1 (\mathcal{O}_2 \mathcal{O}_3)$$

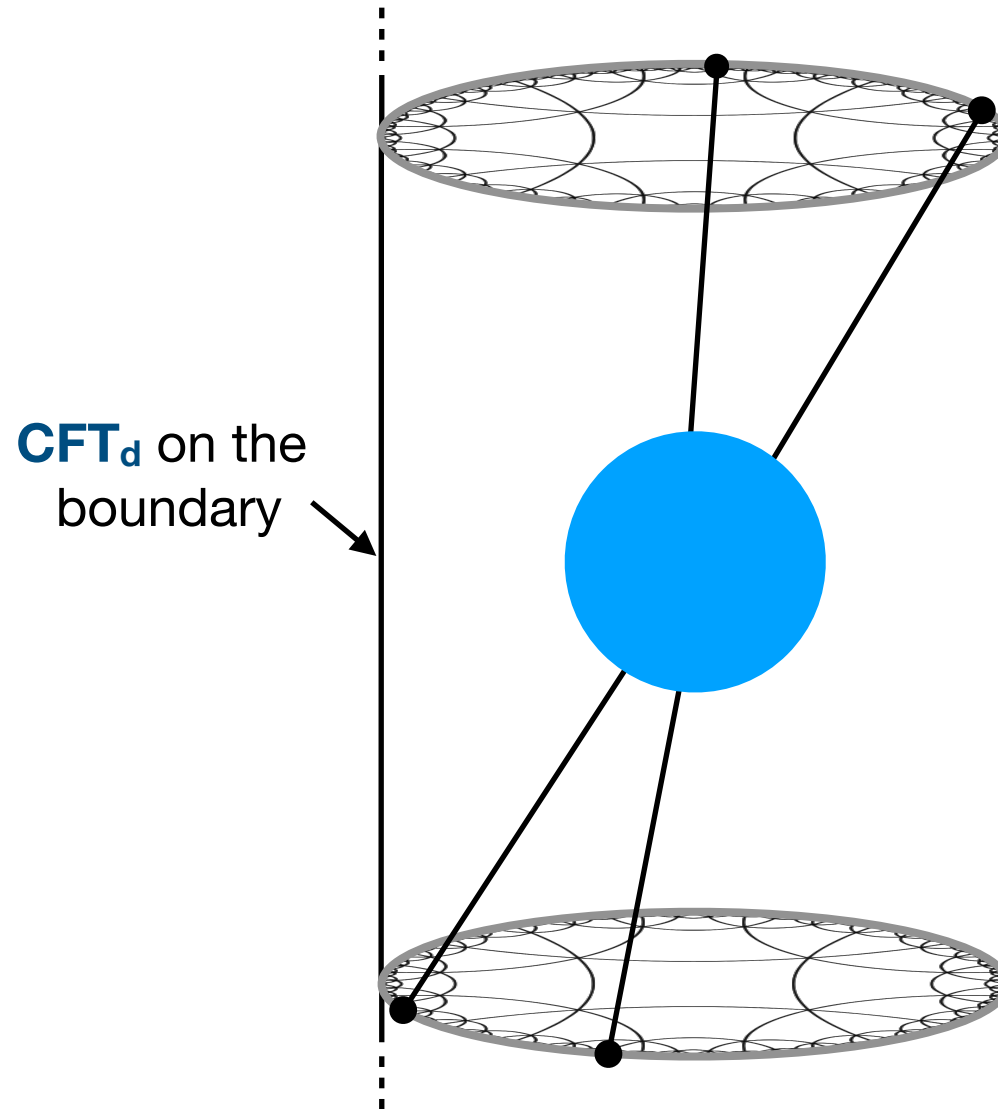
[Belavin, Polyakov, Zamolodchikov 1984;
Rattazzi, Rychkov, Tonni, Vichi 2008]

CFT_d on the
boundary

time



AdS-CFT

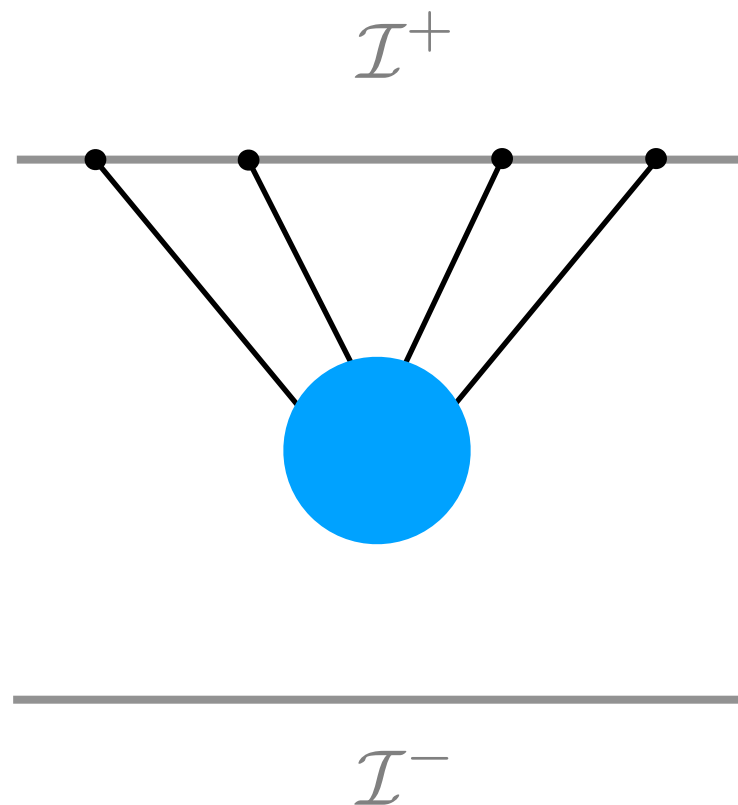


Can we extend this understanding to our own universe?

Holography for all Λ s?

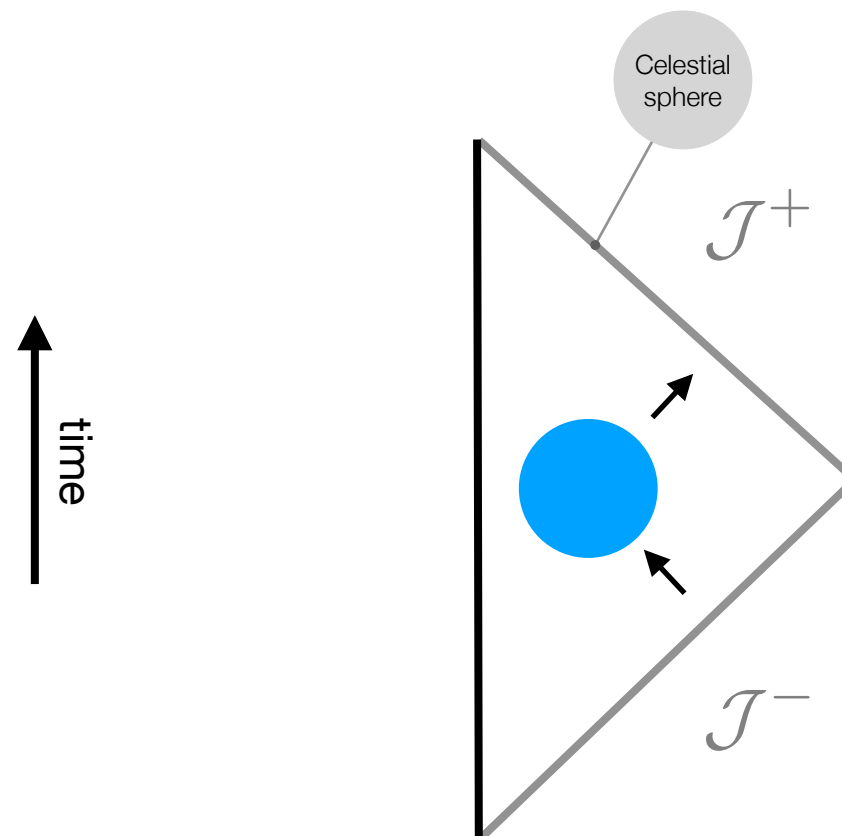
The maximally symmetric cousins of AdS

$\Lambda > 0$ de Sitter



- Cosmological scales
- Primordial inflation

$\Lambda = 0$ Minkowski

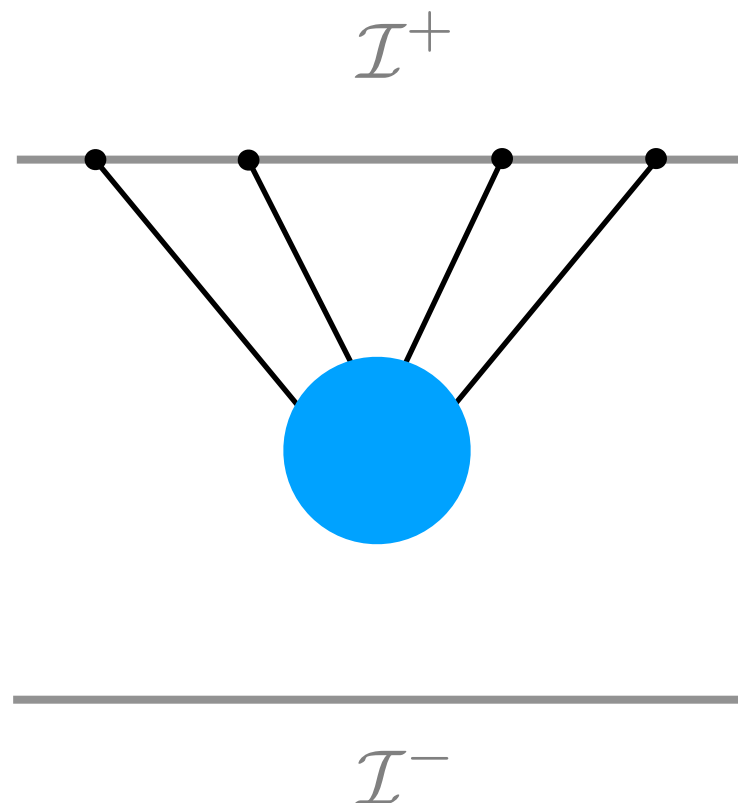


- intermediate scales

Holography for all Λ s?

The maximally symmetric cousins of AdS

$\Lambda > 0$ de Sitter



Cosmological Bootstrap

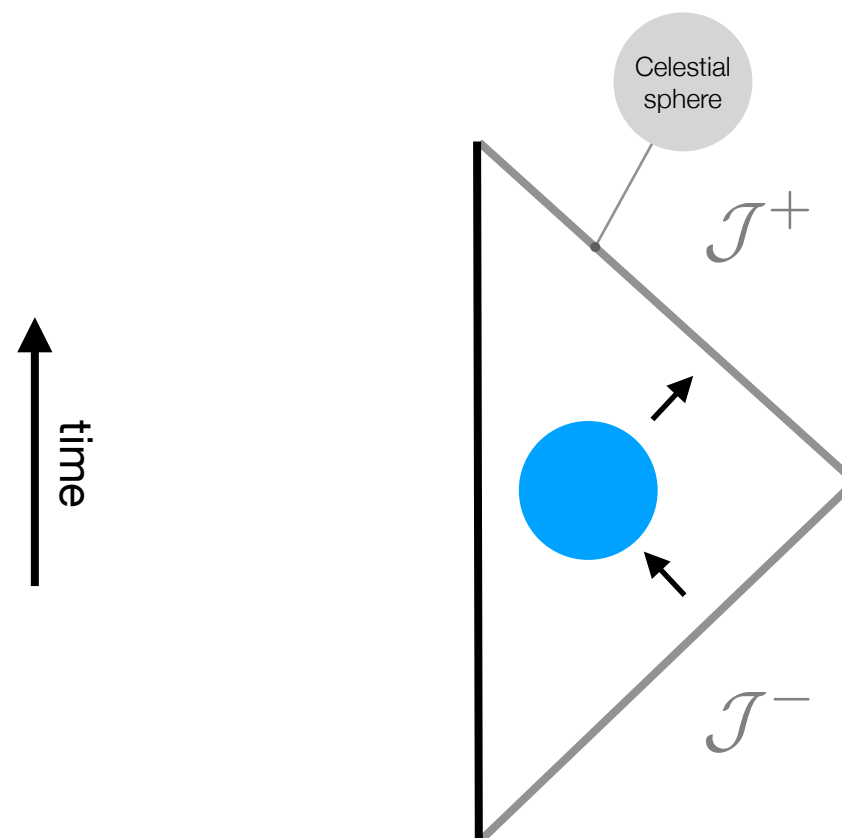
[Arkani-Hamed and Maldacena '15]

[Arkani-Hamed and Benincasa '17]

[Arkani-Hamed, Baumann, Lee and Pimentel '18]

[Sleight and Taronna '19] [Pajer et al '20] [...]

$\Lambda = 0$ Minkowski



Celestial holography

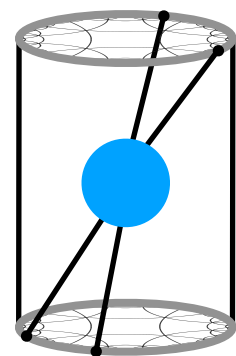
[de Boer and Solodukhin '03]

[Strominger '17] [Pasterski, Shao, Strominger '17]

[Pasterski, Shao '17] [...]

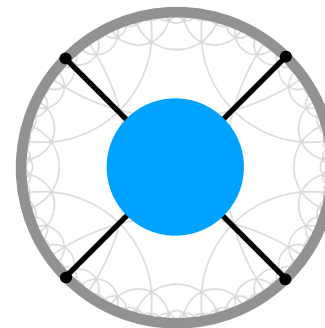
Holography for all Λ s?

Boundary correlators in AdS, dS and on the celestial sphere can be reformulated as boundary correlators in Euclidean AdS:

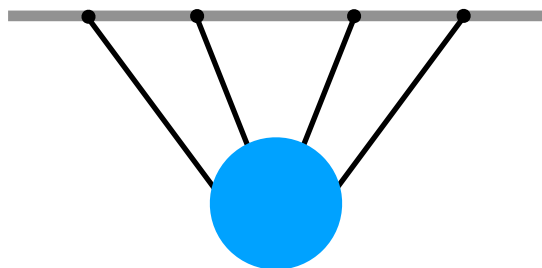


$$\Lambda < 0$$

Wick rotation

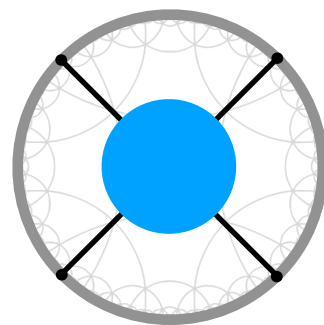


Perturbatively:

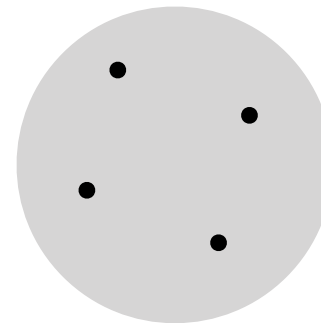


$$\Lambda > 0$$

$$= \sum$$

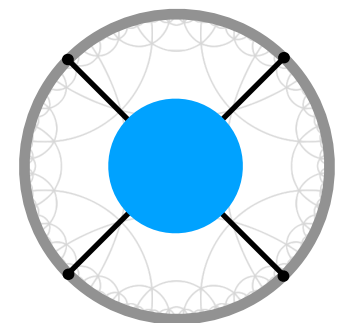


[C.S. & Taronna '19, '20, '21]



$$\Lambda = 0$$

$$= \int$$



[Iacobacci, C.S. & Taronna '22]

[C.S. & Taronna '23]

dS and Celestial correlators therefore have a similar analytic structure to their EAdS counterparts!
On a practical level, can use such identities to import techniques and understanding from AdS.

Outline

I. $\Lambda > 0$

II. $\Lambda = 0$

III. Some Applications.

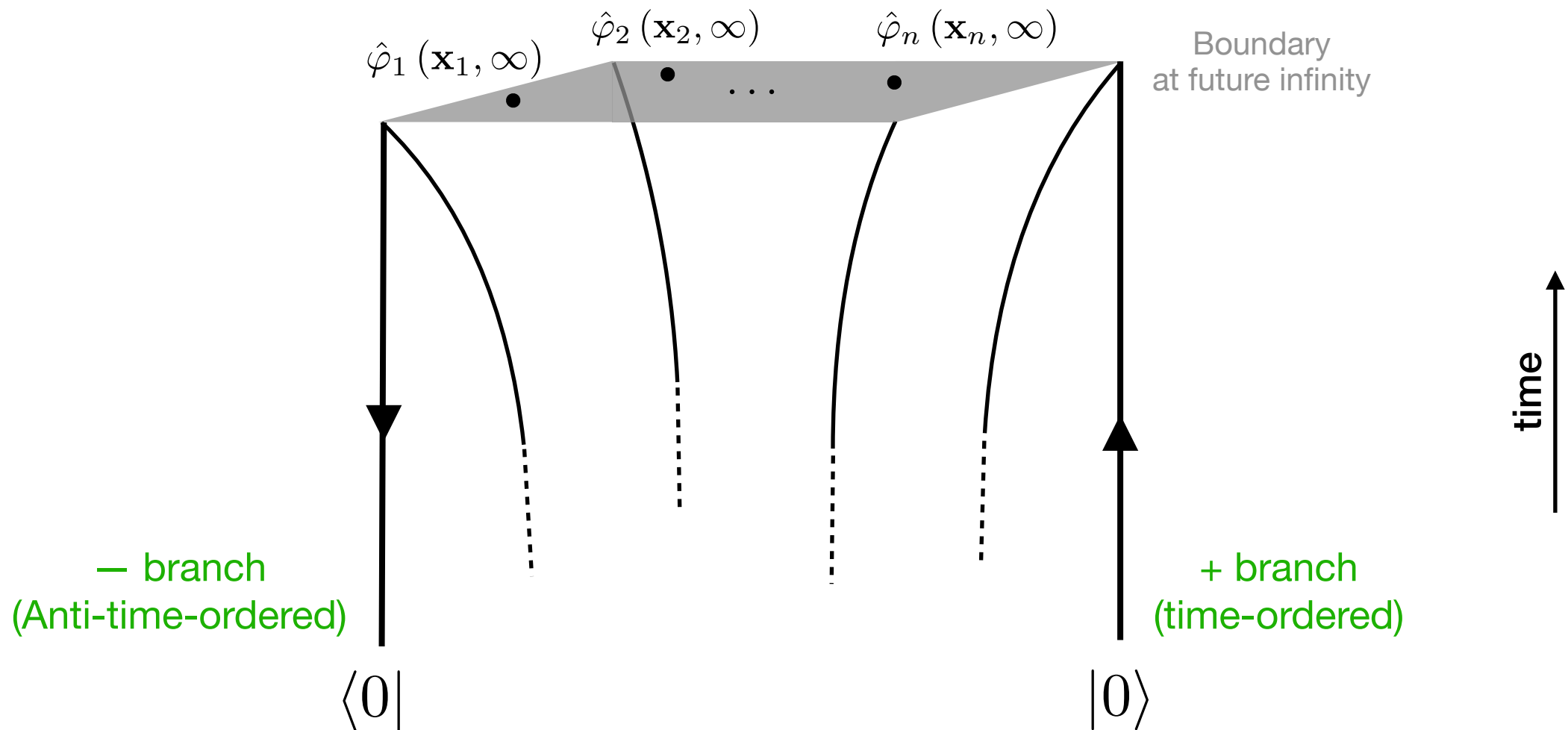
$$\Lambda > 0$$

dS Boundary Correlators

in-in formalism

[Maldacena '02, Weinberg '05]

$$\lim_{\tau \rightarrow \infty} \langle 0 | \hat{\varphi}_1(\mathbf{x}_1, \tau) \dots \hat{\varphi}_n(\mathbf{x}_n, \tau) | 0 \rangle$$



Take $| 0 \rangle$ to be the de Sitter vacuum which reduces to the Minkowski vacuum at early times.

(Bunch Davies vacuum)

dS Boundary Correlators

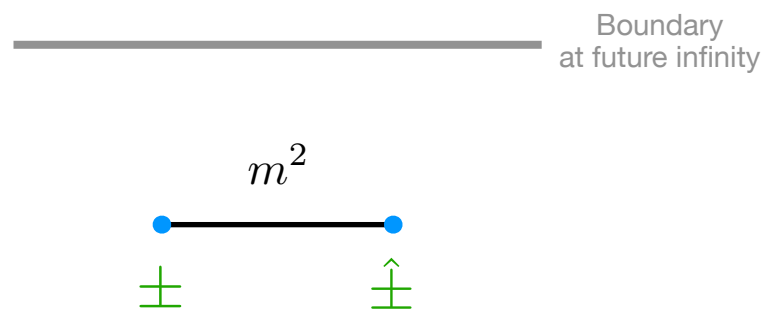
in-in formalism

[Maldacena '02, Weinberg '05]

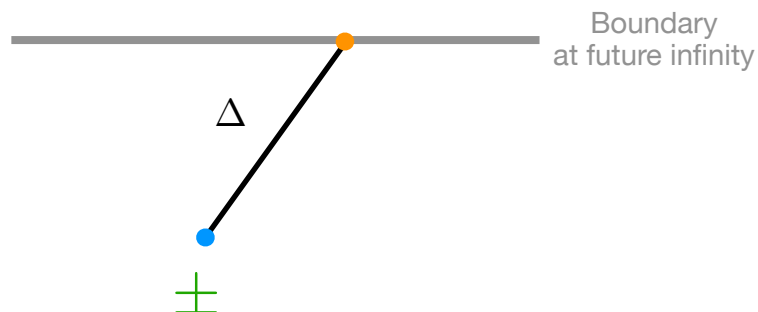
$$\lim_{\tau \rightarrow \infty} \langle 0 | \hat{\varphi}_1(\mathbf{x}_1, \tau) \dots \hat{\varphi}_n(\mathbf{x}_n, \tau) | 0 \rangle$$

Feynman rules:

\pm bulk-to- $\hat{\pm}$ bulk propagator:



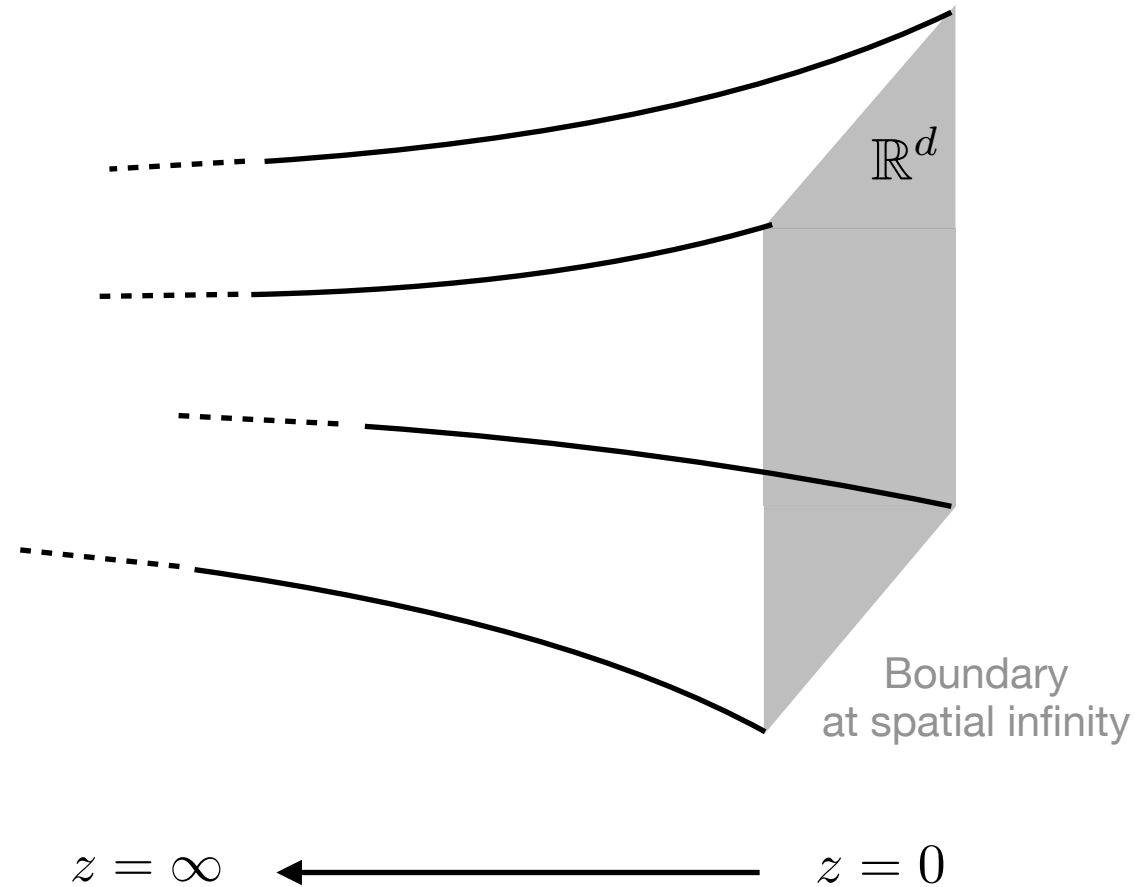
\pm bulk-to-boundary propagator:



Sum contributions from each **branch** (\pm) of the time (in-in) contour!

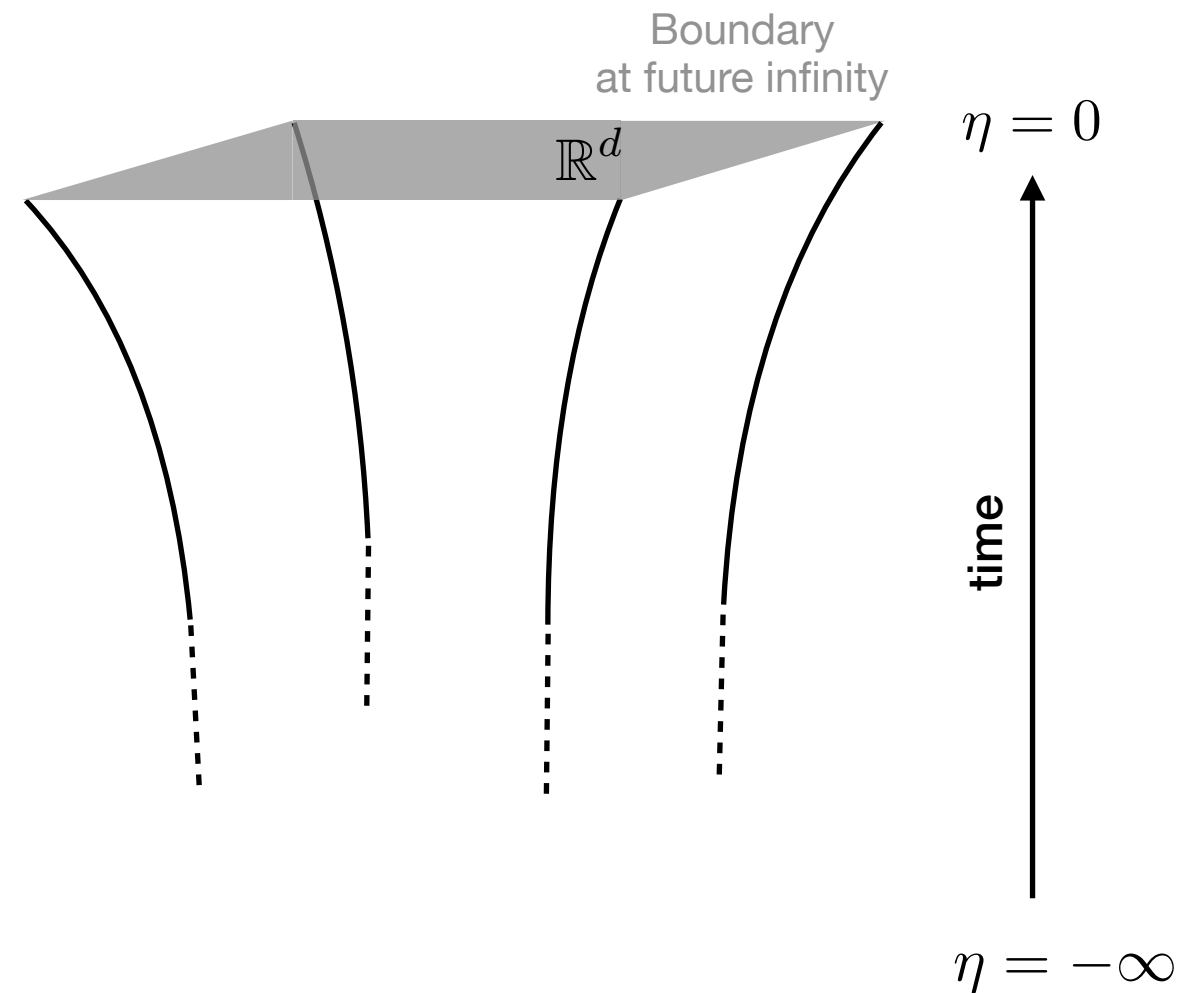
From dS to Euclidean AdS

Euclidean AdS



$$ds^2 = R_{\text{AdS}}^2 \frac{dz^2 + d\mathbf{x}^2}{z^2}$$

dS



$$ds^2 = R_{\text{dS}}^2 \frac{-d\eta^2 + d\mathbf{x}^2}{\eta^2}$$

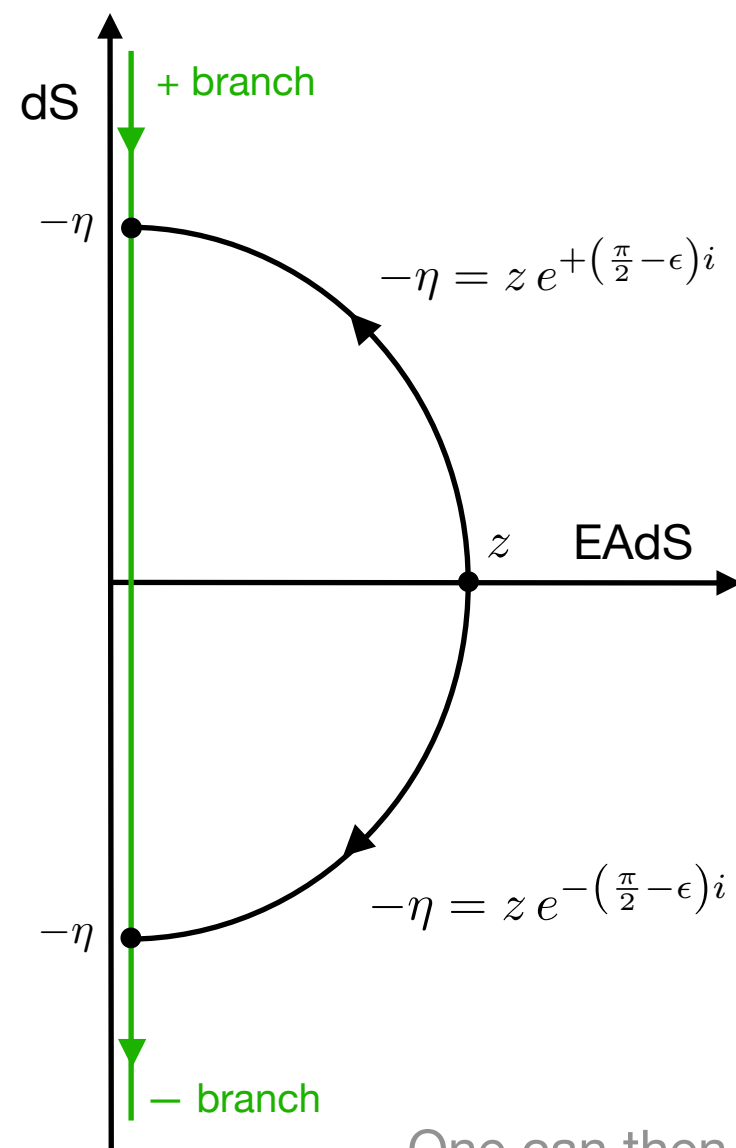
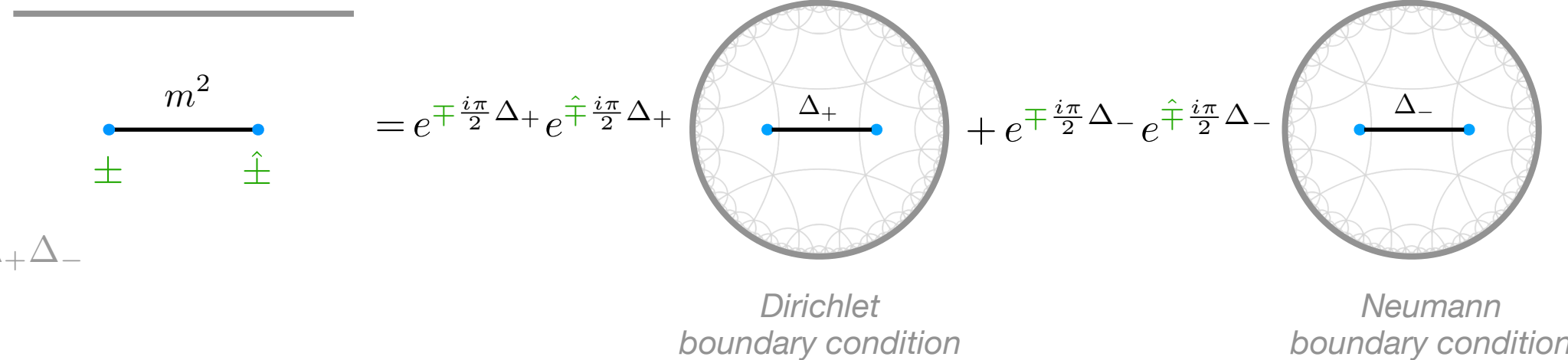
EAdS and dS are identified under:

$$R_{\text{AdS}} = \pm i R_{\text{dS}} \quad z = \pm i (-\eta)$$

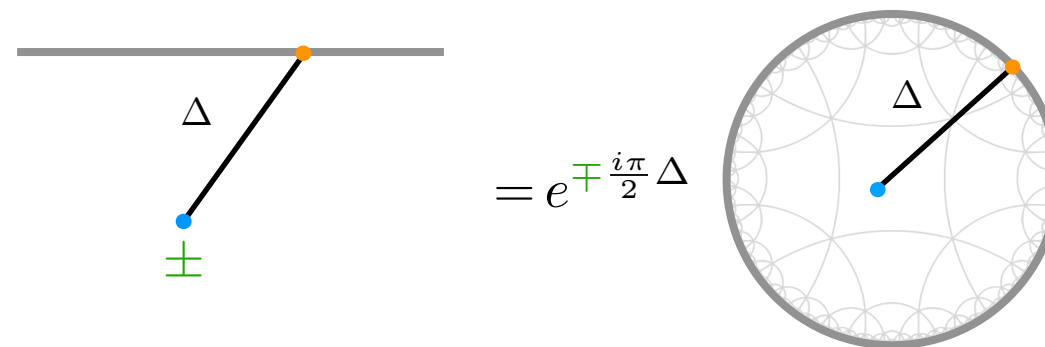
From dS to Euclidean AdS

\pm bulk-to- $\hat{\pm}$ bulk propagator:

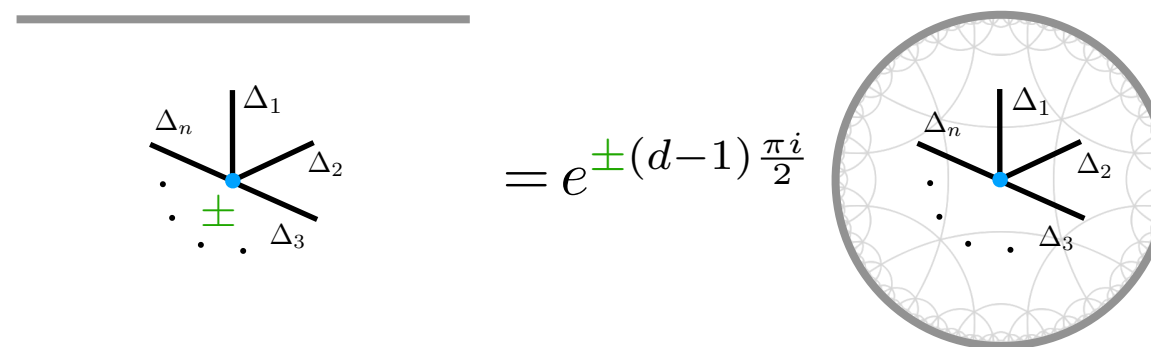
[C.S. and M. Taronna '19, '20, '21]



\pm bulk-to-boundary propagator:



\pm bulk integrals:



One can then write an EAdS Lagrangian for dS correlators [Gorbenko, Komatsu, di Pietro '21]

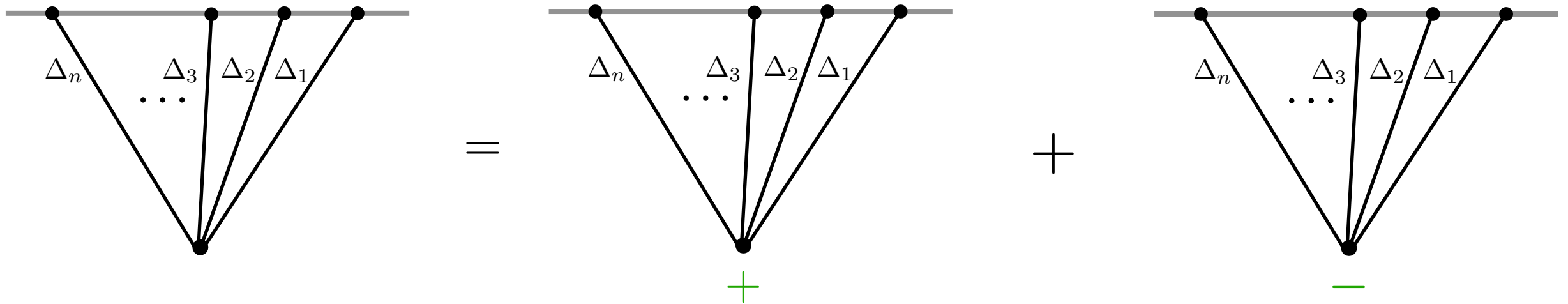
From dS to Euclidean AdS

Examples.

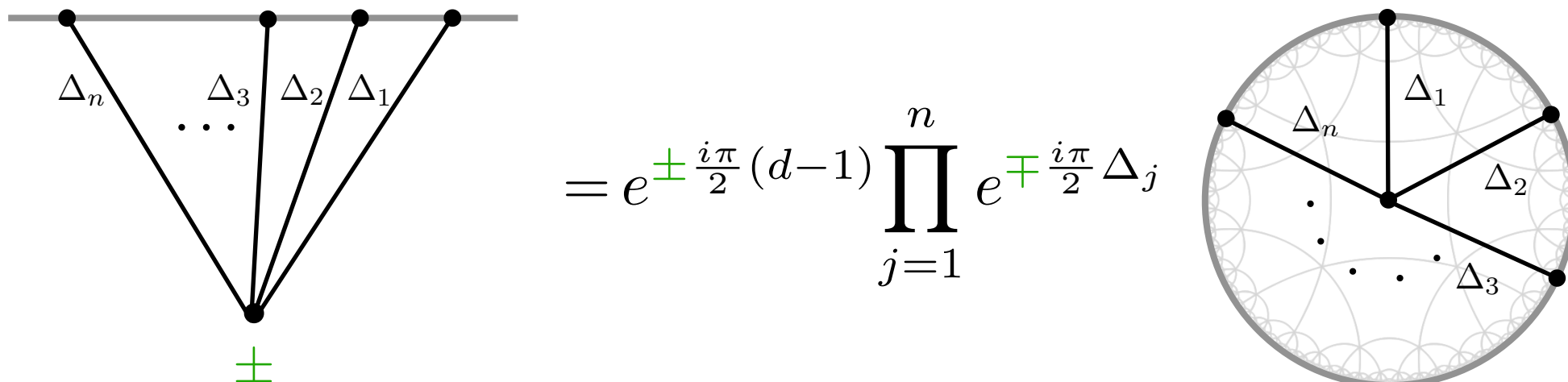
[C.S. and M. Taronna '19]

Non-derivative vertex of scalars fields $\mathcal{V}(X) = g\phi_1(X) \dots \phi_n(X)$

Contact diagram:



Where



Same contact diagram in EAdS

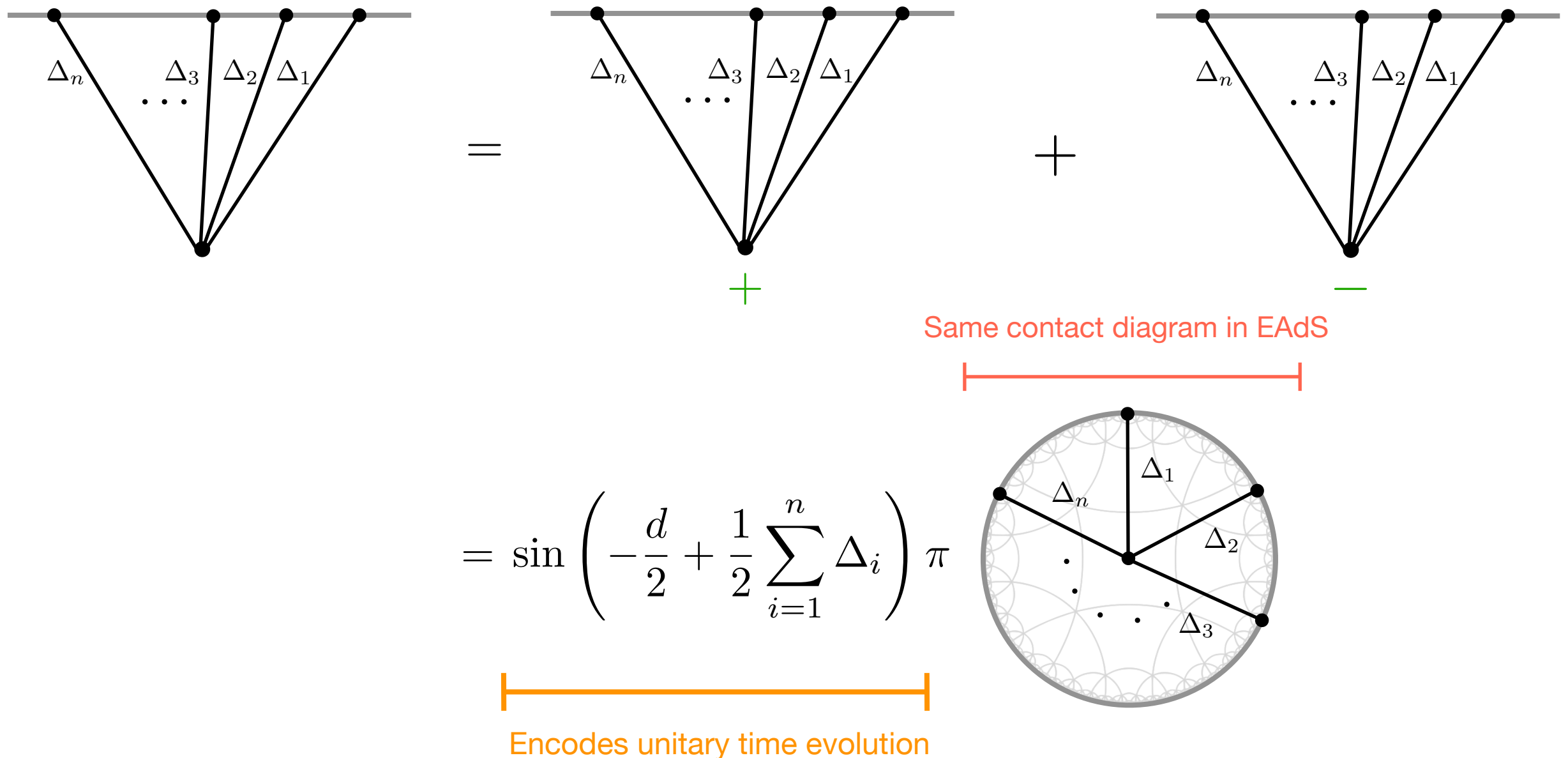
From dS to Euclidean AdS

Examples.

[C.S. and M. Taronna '19]

Non-derivative vertex of scalars fields $\mathcal{V}(X) = g\phi_1(X) \dots \phi_n(X)$

Contact diagram:

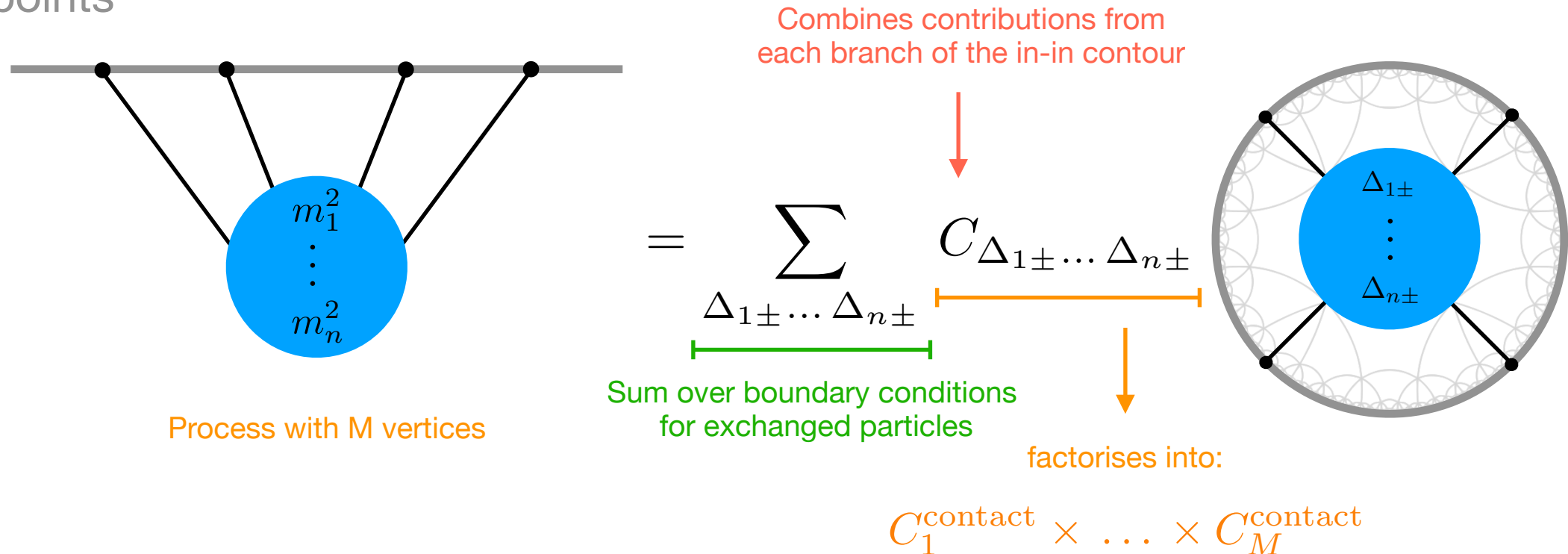


c.f. "Cosmological Optical Theorem" [Goodhew et al, 2020]

From dS to EAdS, and back

dS boundary correlators are perturbatively recast as Witten diagrams in EAdS:

e.g. four-points



Notes:

- Contributions from both Δ_{\pm} modes, which is not always possible in AdS
- $\Delta_{i\pm} \in$ Unitary Irreducible Representation of **dS** isometry

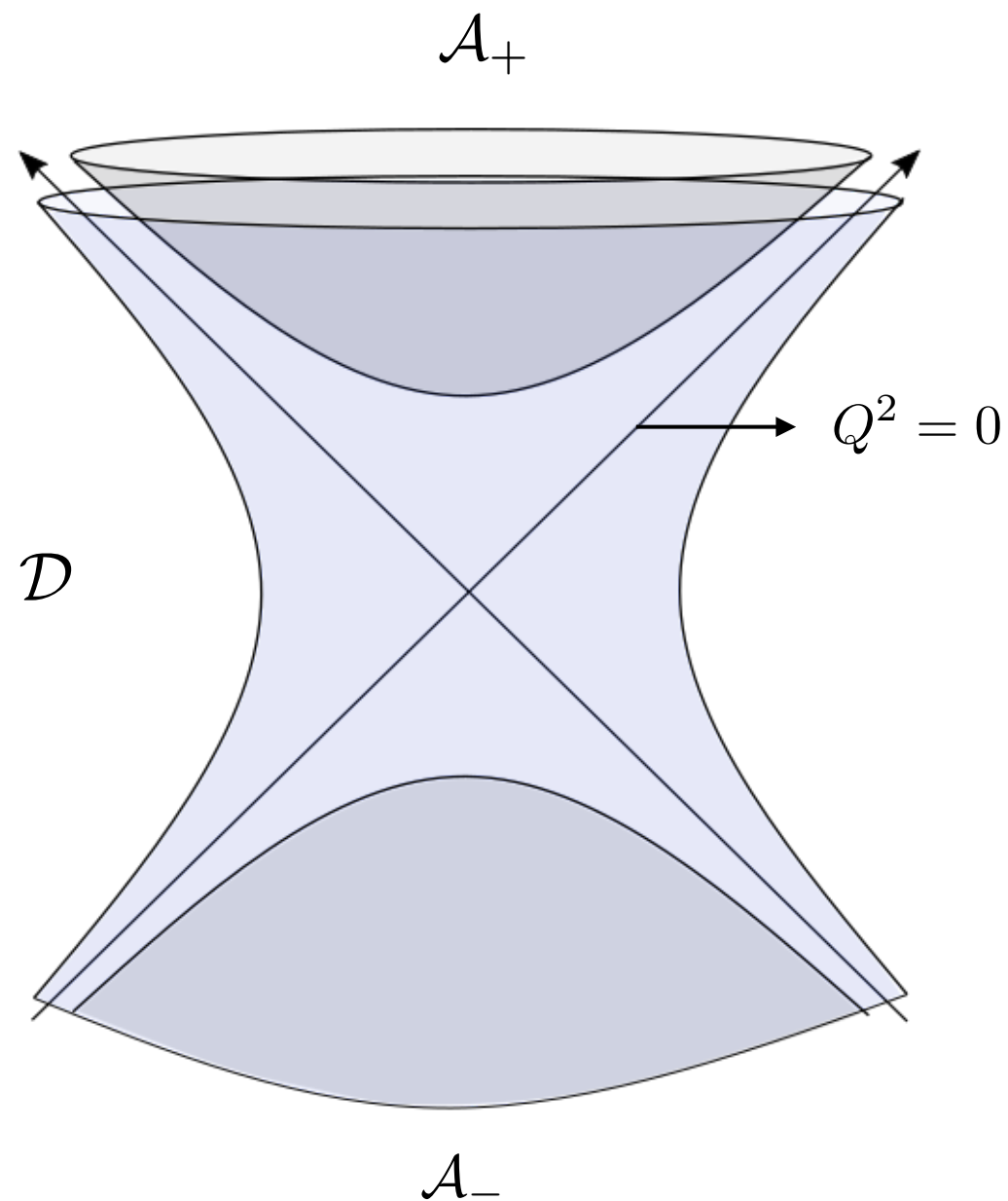
Can use to import techniques and results from AdS to dS!

$$\Lambda = 0$$

Hyperbolic slicing of Minkowski space

[de Boer and Solodukhin '03]

(d+2)-dimensional Minkowski space \mathbb{M}^{d+2} , coordinates X^A , $A = 0, \dots, d+1$



$$\mathcal{A}_{\pm} : X^2 = -t^2 \quad (\text{EAdS}_{d+1}, \text{radius } t)$$

$$\mathcal{D} : X^2 = R^2 \quad (\text{dS}_{d+1}, \text{radius } R)$$

Conformal boundary:

$$Q^2 = 0, \quad Q \equiv \lambda Q, \quad \lambda \in \mathbb{R}^+$$

Introduce projective coordinates:

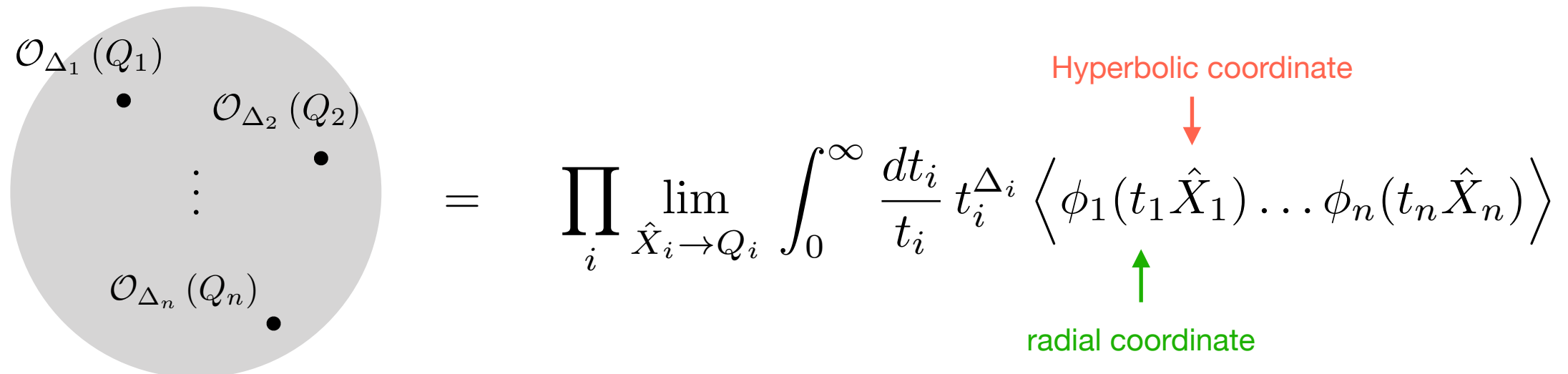
$$\xi_i = Q^i / Q^0, \quad i = 1, \dots, d+1$$

$$\xi_1^2 + \dots + \xi_{d+1}^2 = 1 \quad \left[\begin{array}{l} \text{d-dimensional} \\ \text{unit sphere} \end{array} \right]$$

Minkowski boundary correlators

[C.S. and M. Taronna '23]

Radial **Mellin transform** of Minkowski correlators implements a radial reduction onto the hyperbolic slicing:



$$\mathcal{O}_{\Delta_1}(Q_1) \dots \mathcal{O}_{\Delta_n}(Q_n) = \prod_i \lim_{\hat{X}_i \rightarrow Q_i} \int_0^\infty \frac{dt_i}{t_i} t_i^{\Delta_i} \left\langle \phi_1(t_1 \hat{X}_1) \dots \phi_n(t_n \hat{X}_n) \right\rangle$$

Hyperbolic coordinate \downarrow
radial coordinate \uparrow

Celestial correlators then arise in the boundary limit $\hat{X}_i \rightarrow Q_i$!

Mellin transform

$$\int_0^\infty \frac{dt}{t} t^\Delta (\dots)$$

Inverse Mellin transform

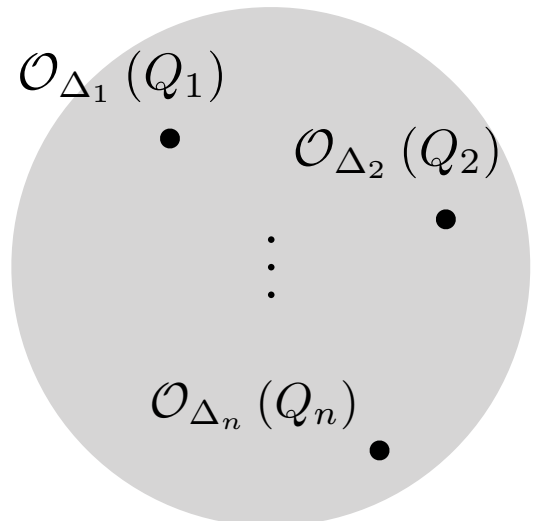
$$\int_{\frac{d}{2}-i\infty}^{\frac{d}{2}+i\infty} \frac{d\Delta}{2\pi i} t^{-\Delta} (\dots)$$

Unitary Principal Series
representations of $\text{SO}(d+1,1)$

Minkowski boundary correlators

[C.S. and M. Taronna '23]

Radial **Mellin transform** of Minkowski correlators implements a radial reduction onto the hyperbolic slicing:



$$= \prod_i \lim_{\hat{X}_i \rightarrow Q_i} \int_0^\infty \frac{dt_i}{t_i} t_i^{\Delta_i} \left\langle \phi_1(t_1 \hat{X}_1) \dots \phi_n(t_n \hat{X}_n) \right\rangle$$

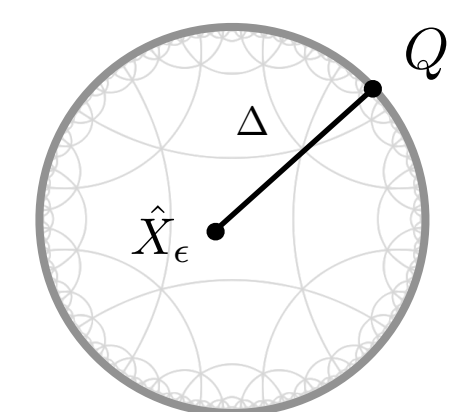
Hyperbolic coordinate radial coordinate

Celestial correlators then arise in the boundary limit $\hat{X}_i \rightarrow Q_i$!

“Celestial” bulk-to-boundary propagator:

$$G_{\Delta}^{\text{flat}}(X, Q) = \lim_{\hat{Y} \rightarrow Q} \int_0^\infty \frac{dt}{t} t^{\Delta} G_F(X, t\hat{Y}) = \overbrace{\mathcal{K}_{i(\frac{d}{2}-\Delta)}^{(m)}(\sqrt{X^2 + i\epsilon})}^{\text{Kernel of the radial reduction (Bessel-K function)}} \times$$

bulk-to-boundary propagator in EAdS



From the Celestial Sphere to EAdS

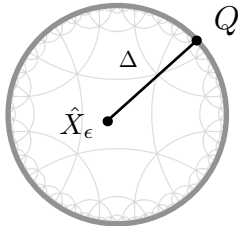
[C.S. and M. Taronna '23]

Examples.

Non-derivative vertex of scalars fields $\mathcal{V}(X) = g\phi_1(X) \dots \phi_n(X)$

Contact diagram:

$$\langle \mathcal{O}_{\Delta_1}(Q_1) \dots \mathcal{O}_{\Delta_n}(Q_n) \rangle = -ig \int d^{d+2}X G_{\Delta_1}^{\text{flat}}(X, Q_1) \dots G_{\Delta_n}^{\text{flat}}(X, Q_n).$$

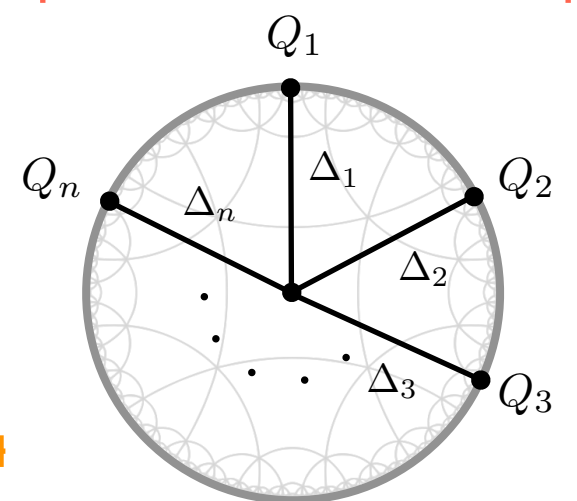
$$G_{\Delta}^{\text{flat}}(X, Q) = \kappa_{i(\frac{d}{2}-\Delta)}^{(m)}(\sqrt{X^2 + i\epsilon}) \times$$


$$= \underbrace{R_{\Delta_1 \dots \Delta_n}(m_1, \dots, m_n)}_{\text{Contribution from radial integral. Encodes all mass dependence. (Generalised hypergeometric)}} \times \underbrace{\sin\left(-\frac{d}{2} + \frac{1}{2} \sum_{i=1}^n \Delta_i\right) \pi}_{\text{Same factor as for dS contact diagrams. Comes from combining contributions from regions inside and outside lightcone}}$$

Contribution from radial integral.
Encodes all mass dependence.
(Generalised hypergeometric)

Same factor as for dS contact diagrams.
Comes from combining contributions
from regions inside and outside lightcone

(Analytically continued)
EAdS contact diagram

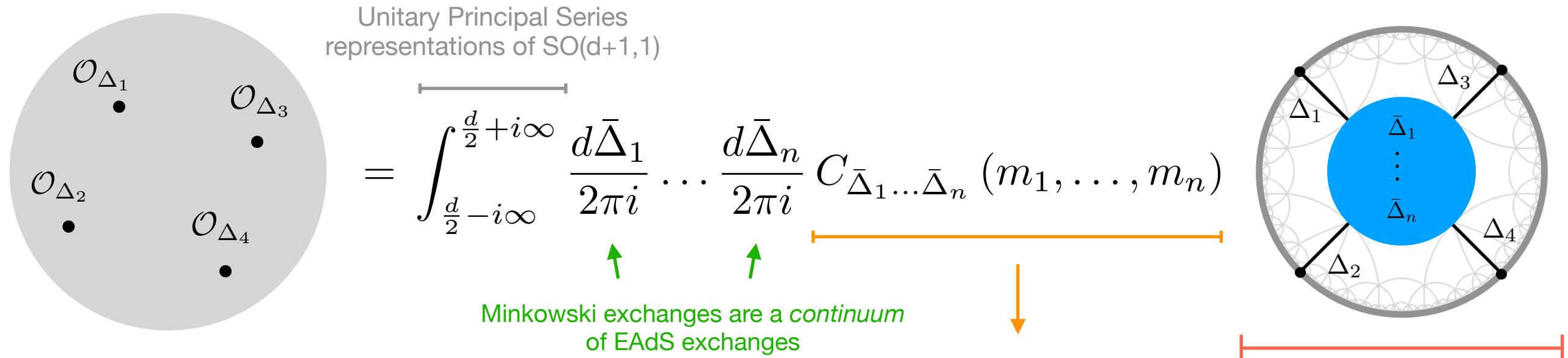


Like in dS, Celestial contact diagrams are proportional to their EAdS counterparts

From the Celestial Sphere to EAdS

[C.S. and M. Taronna '23]

In general, for exchanges of particles of mass m_i , $i = 1, \dots, n$



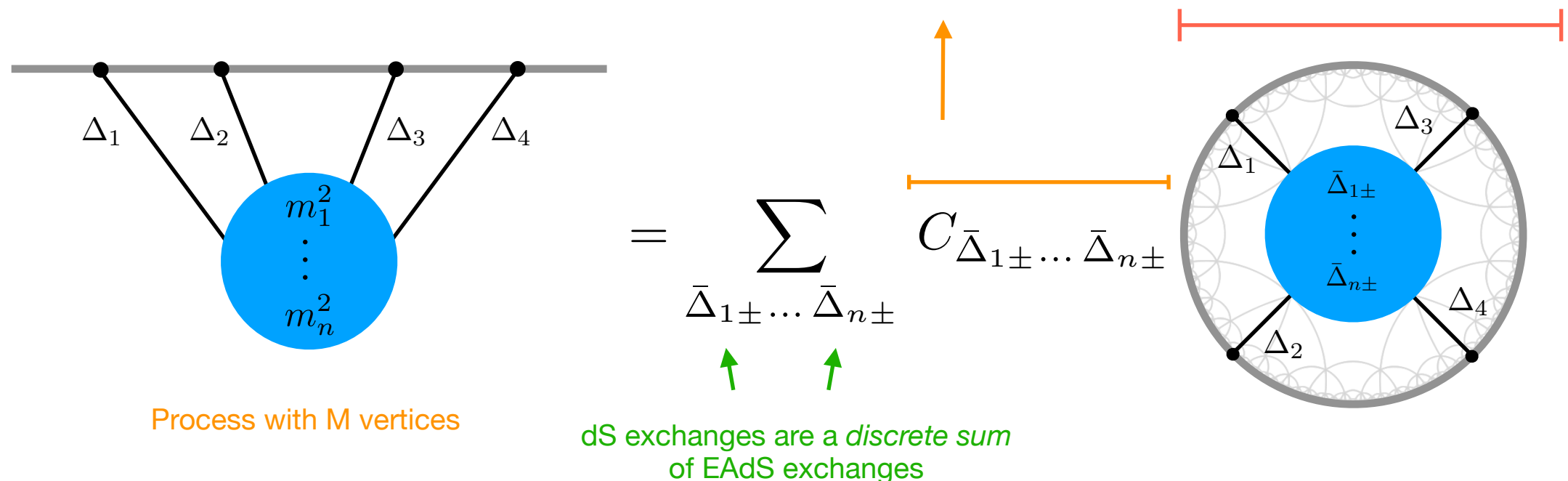
Process with M vertices

factorises into:

$$C_1^{\text{contact}} \times \dots \times C_M^{\text{contact}}$$

Makes manifest conformal symmetry

Compare with de Sitter:

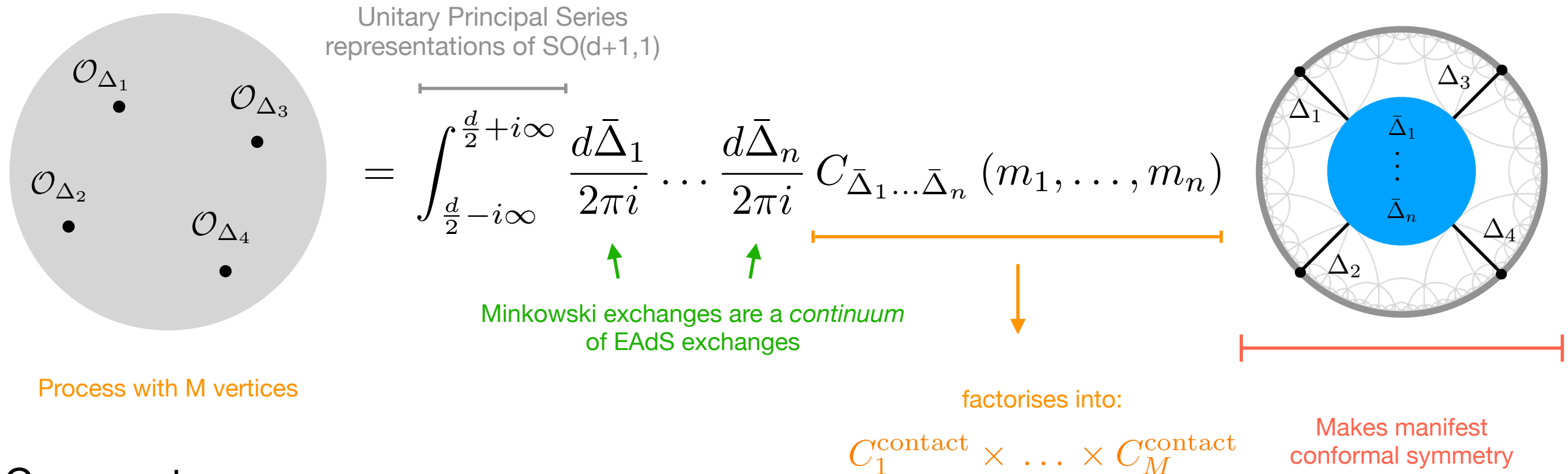


Process with M vertices

From the Celestial Sphere to EAdS

[C.S. and M. Taronna '23]

In general, for exchanges of particles of mass m_i , $i = 1, \dots, n$



Comments:

- Relation to definition [Pasterski, Shao, Strominger '17] of celestial correlators as scattering amplitudes in a conformal basis?

[Pasterski, Shao, Strominger '17] = LSZ ([Sleight, Taronna '23]) ?

- Celestial correlators defined as an extrapolation of bulk Minkowski correlators give a definition of celestial correlators for theories without an S-matrix.

What lessons can we draw from Minkowski CFT?

Some applications.

Perturbative OPE data

Perturbative OPE data on the boundary of dS and Minkowski space from EAdS

E.g. Composite operators on the boundary

[C.S. and M. Taronna '20]

$$[\mathcal{O}\mathcal{O}]_{n,\ell} \sim \mathcal{O} (\partial^2)^n \partial_{i_1} \dots \partial_{i_\ell} \mathcal{O} + \dots \quad \text{scaling dimension: } \Delta_{n,\ell} = \underbrace{2\Delta + 2n + \ell}_{\text{Free theory}} + \underbrace{\gamma_{n,\ell}}_{\text{anomalous dimension}}$$

- $\gamma_{n,\ell}$ induced by bulk ϕ^4 contact diagram in dS:

$$\text{Boundary diagram} = \sin\left(-\frac{d}{2} + 2\Delta\right) \pi \times \text{Bulk contact diagram} \rightarrow \gamma_{n,\ell}^{\phi^4} = \sin\left(-\frac{d}{2} + 2\Delta\right) \pi \times (\text{EAdS}) \gamma_{n,\ell}^{\phi^4}$$

- $\gamma_{n,\ell}$ induced by an exchange diagram in dS:

$$\text{Boundary diagram} = \sin\left(\frac{-d + 2\Delta + \Delta_+}{2}\right) \pi \sin\left(\frac{-d + 2\Delta + \Delta_+}{2}\right) \pi \times \text{Bulk exchange diagram} + (\Delta_+ \rightarrow \Delta_-)$$

$$\rightarrow \gamma_{n,\ell}^{\phi^3 \text{ exch}} = \sin\left(\frac{-d + 2\Delta + \Delta_+}{2}\right) \pi \sin\left(\frac{-d + 2\Delta + \Delta_+}{2}\right) \pi \times (\text{EAdS}) \gamma_{n,\ell}^{\phi^3 \text{ exch } \Delta_+} + (\Delta_+ \rightarrow \Delta_-)$$

Perturbative OPE data

Perturbative OPE data on the boundary of dS and Minkowski space from EAdS

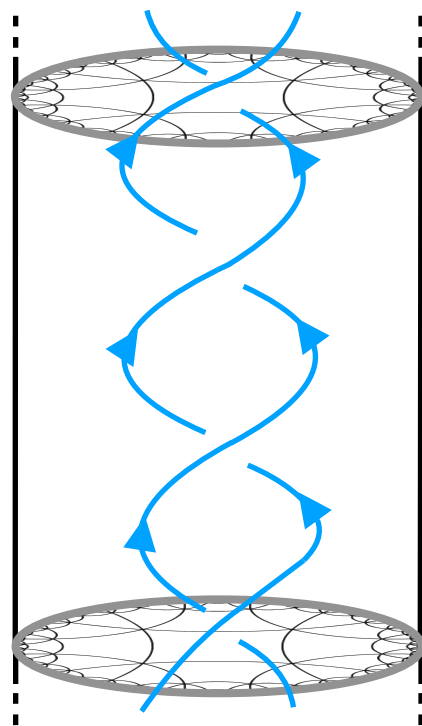
E.g. Composite operators on the boundary

$$[\mathcal{O}\mathcal{O}]_{n,\ell} \sim \mathcal{O} (\partial^2)^n \partial_{i_1} \dots \partial_{i_\ell} \mathcal{O} + \dots$$

scaling dimension: $\Delta_{n,\ell} = 2\Delta + 2n + \ell + \gamma_{n,\ell}$

Free theory anomalous dimension

AdS

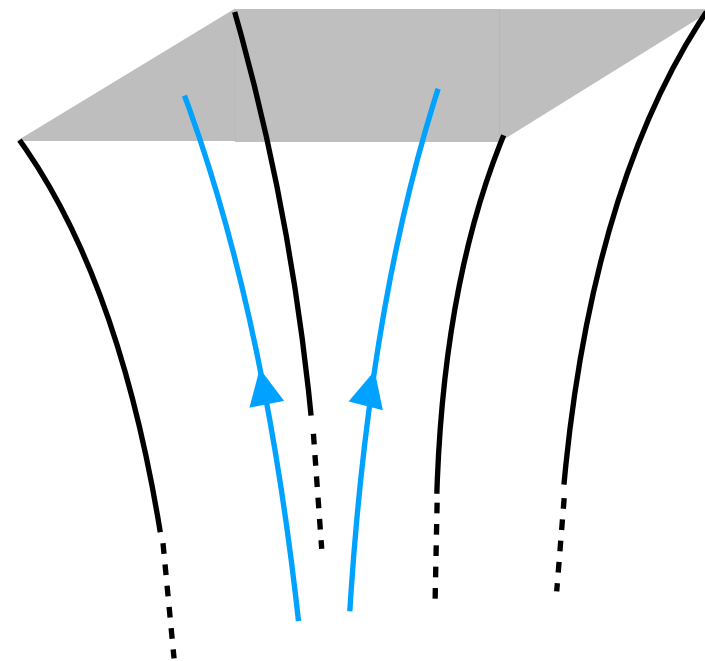


$\Delta_{n,\ell}$ is unitary

→ stable particle (bound state)

vs.

dS



$\Delta_{n,\ell}$ is (generally) non-unitary

→ resonance

Conformal Partial Wave Expansion

[Sleight, Taronna '20] [Hogervorst, Penedones, Vaziri '21] [di Pietro, Komatsu, Gorbenko '21]

Perturbative dS and celestial correlators have a similar analytic structure to those in AdS.

→ Like in AdS they admit a conformal partial wave expansion

$$\langle \mathcal{O}(\mathbf{x}_1) \mathcal{O}(\mathbf{x}_2) \mathcal{O}(\mathbf{x}_3) \mathcal{O}(\mathbf{x}_4) \rangle = \sum_{J=0}^{\infty} \int_{\frac{d}{2}-i\infty}^{\frac{d}{2}+i\infty} \frac{d\Delta}{2\pi i} \overset{\text{Spectral density, meromorphic in } \Delta}{\rho_J(\Delta)} \underbrace{\mathcal{F}_{\Delta,J}(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3, \mathbf{x}_4)}_{\text{Conformal Partial Wave}}$$

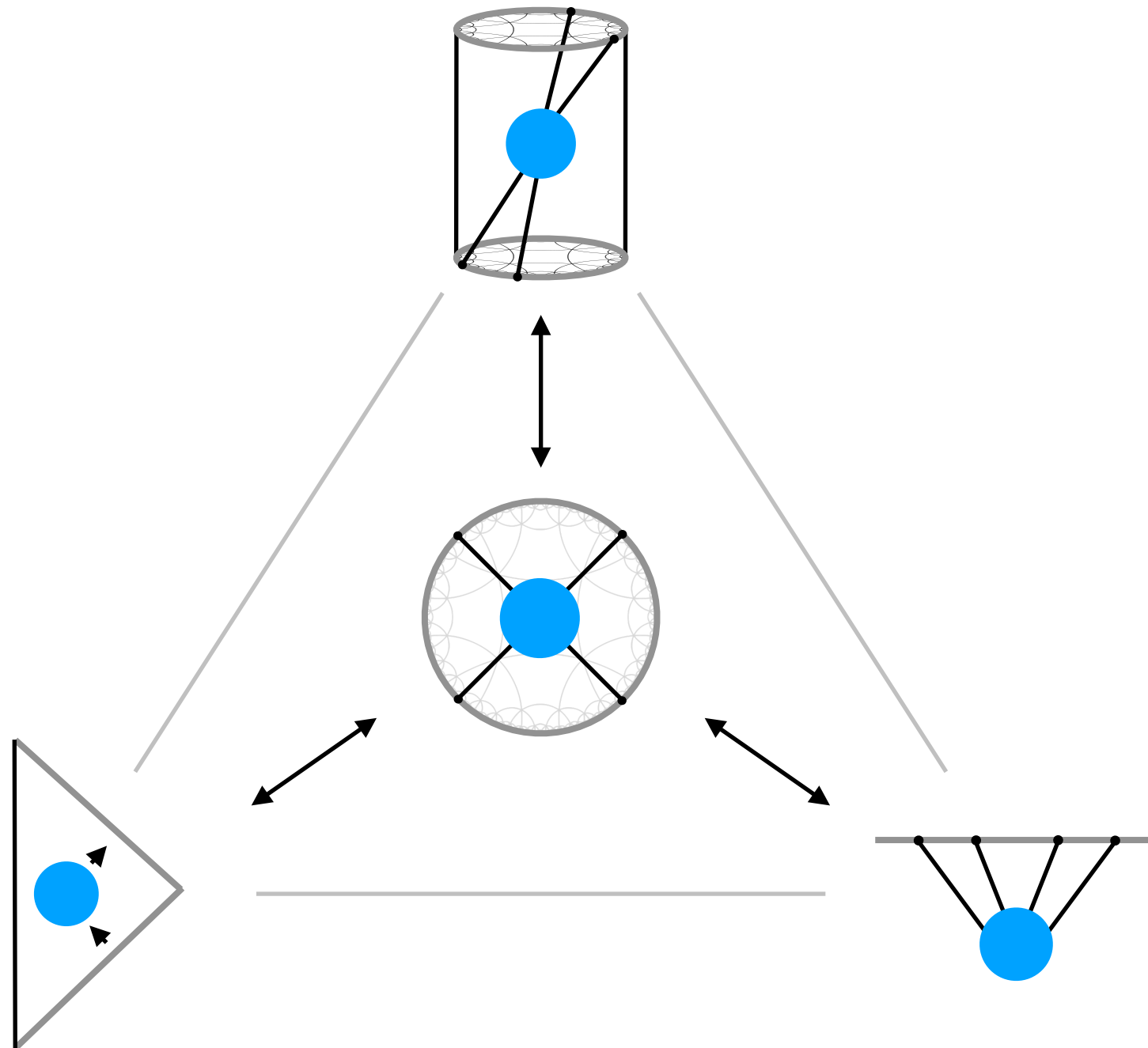
This has been argued to hold non-perturbatively as well [Hogervorst, Penedones, Vaziri '21, di Pietro, Komatsu, Gorbenko '21]

Unitarity: $\rho_J(\Delta) \geq 0$ + crossing → Bootstrap for Euclidean CFTs?

Cf. Lorentzian CFT:

$$\langle \mathcal{O}(\mathbf{x}_1) \mathcal{O}(\mathbf{x}_2) \mathcal{O}(\mathbf{x}_3) \mathcal{O}(\mathbf{x}_4) \rangle = \sum_{\Delta, J} C_{\Delta, J}^2 \underbrace{G_{\Delta, J}(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3, \mathbf{x}_4)}_{\text{Conformal Block}}$$

Unitarity: $C_{\Delta, J}^2 \geq 0$ + crossing → Conformal Bootstrap



Thank you.