# NNLO QCD corrections to SIDIS: <br> unpolarized \& polarized 

Narayan Rana
NISER

Loops and Legs in Quantum Field Theory
18th April 2024
in collaboration with S. Goyal, R. N. Lee, S. Moch, V. Pathak, V. Ravindran

## Electron Ion Collider

a machine to look inside the nucleus

EIC will take precision snapshots of the internal structures of the protons and neutrons, allowing us a better understanding of the strongest force in nature

A precise theoretical description of our current understanding (the Standard Model) is also necessary to find any agreement/disagreement with precise experimental data

The inclusive/semi-inclusive deep inelastic scattering (DIS/SIDIS) plays a crucial role.
Parton model connects the partonic cross-section to the hadronic one through PDFs \& FFs. We compute the partonic cross-section using the framework of perturbative QCD order by order in $\alpha_{s}$.

Higher order corrections are essential to

- achieve sufficiently/comparably precise theoretical estimates
- reduce the uncertainties arising from the factorization scales


## Semi-Inclusive DIS

see the talk by J. Blümlein for DIS


## SIDIS

$$
l+H \rightarrow l+H^{\prime}+X
$$

Phase-space:
Final state hadron is tagged!
Extra constrain on the phase-space
DIS

$$
l+H \rightarrow l+X
$$

Phase-space:
All final states are fully integrated!

$$
\left.d P S\right|_{\mathrm{SIDIS}}=\left.d P S\right|_{\mathrm{DIS}} \times \delta\left(z^{\prime}-\frac{p_{a} \cdot p_{b}}{p_{a} \cdot q}\right)
$$

The hadronic part is characterized by structure functions ( $F_{i}$ ). The hadronic SFs are expressed with the convolution of partonic finite coefficient functions $\left(\mathcal{F}_{i, a b}\right)$, PDFs and FFs.

## Parton model \& perturbative expansion

$$
F_{i}=x^{i-1} \sum_{a, b} \int_{x}^{1} \frac{d x_{1}}{x_{1}} f_{a}\left(x_{1}, \mu_{F}^{2}\right) \int_{z}^{1} \frac{d z_{1}}{z_{1}} D_{b}\left(z_{1}, \mu_{F}^{2}\right) \times \mathcal{F}_{i, a b}\left(\frac{x}{x_{1}}, \frac{z}{z_{1}}, Q^{2}, \mu_{F}^{2}\right)
$$

the finite coefficient functions which can be computed perturbatively

In QCD, we have a series expansion of the partonic cross sections in strong coupling constant $\alpha_{s}$ :

$$
\begin{aligned}
\mathcal{F}_{a b}(z) & =\mathcal{F}_{a b}^{(0)} \sum_{m=0}^{\infty} \alpha_{s}^{m} \mathcal{F}_{a b}^{(m)}(z) \\
& =\mathcal{F}_{a b}^{(0)}\left[1+\alpha_{s} \mathcal{F}_{a b}^{(1)}(z)+\alpha_{s}^{2} \mathcal{F}_{a b}^{(2)}(z)+\alpha_{s}^{3} \mathcal{F}_{a b}^{(3)}(z)+\cdots\right]
\end{aligned}
$$

We are interested in the second order correction

## Parton model \& perturbative expansion: polarized

We consider the asymmetry between the parallel and anti-parallel spin orientations of the colliding electron and hadron. $\Delta f_{a}=f_{a(\uparrow) / H(\uparrow)}-f_{a(\downarrow) / H(\uparrow)}$ are the spin-dependent PDFs.

$$
\begin{gathered}
g_{1}=\sum_{a, b} \int_{x}^{1} \frac{d x_{1}}{x_{1}} \Delta f_{a}\left(x_{1}, \mu_{F}^{2}\right) \int_{z}^{1} \frac{d z_{1}}{z_{1}} D_{b}\left(z_{1}, \mu_{F}^{2}\right) \times \mathcal{G}_{1, a b}\left(\frac{x}{x_{1}}, \frac{z}{z_{1}}, Q^{2}, \mu_{F}^{2}\right) \\
\Downarrow
\end{gathered}
$$

the finite coefficient functions which can be computed perturbatively

$$
\begin{aligned}
\mathcal{G}_{a b}(z) & =\mathcal{G}_{a b}^{(0)} \sum_{m=0}^{\infty} \alpha_{s}^{m} \mathcal{G}_{a b}^{(m)}(z) \\
& =\mathcal{G}_{a b}^{(0)}\left[1+\alpha_{s} \mathcal{G}_{a b}^{(1)}(z)+\alpha_{s}^{2} \mathcal{G}_{a b}^{(2)}(z)+\alpha_{s}^{3} \mathcal{G}_{a b}^{(3)}(z)+\cdots\right]
\end{aligned}
$$

We are interested in the second order correction

## State-of-the-art

## NLO and beyond

```
Altarelli, Ellis, Martinelli, Pi (1979);
Furmanski and Petronzio (1982);
de Florian, Stratmann and Vogelsang (1998); (polarized)
Daleo and Sassot (2003);
```

NNLO
Goyal, Lee, Moch, Pathak, NR, Ravindran $(2023,2024)$
Bonino, Gehrmann, Löchner, Schönwald, Stagnitto (2024)
$\Rightarrow$ This talk
$\Rightarrow$ Next talk by Bonino

Resummation \& Parton shower
Cacciari and Catani (2001);
Anderle, Ringer and Vogelsang (2013);
Abele, de Florian, and Vogelsang (2021,2022);
Borsa, Jäger (2024);
$\Rightarrow$ Talk by Borsa

## This talk is based on

S. Goyal, S. Moch, V. Pathak, NR, V. Ravindran NNLO QCD corrections to semi-inclusive DIS, (arXiv:2312.17711 [hep-ph]).
S. Goyal, R. Lee, S. Moch, V. Pathak, NR, V. Ravindran

NNLO QCD corrections to polarized semi-inclusive DIS, (arXiv:2404.09959 [hep-ph]).

## Schematic diagrams for NNLO contributions to SIDIS



Each individual contribution is divergent: $\frac{1}{\epsilon}$ in dimensional regularization

## Schematic diagrams for NNLO contributions to SIDIS



Sum of all degenerate processes: KLN theorem \& mass factorization

## Computational procedure

$$
d=4+\epsilon
$$

- Diagrammatic approach -> QGRAF to generate Feynman diagrams
- In-house FORM routines for algebraic manipulation: Lorentz, Dirac and Color algebra
- Projectors: different projectors for unpolarized and polarized cases
- Reverse unitarity : phase-space integrals to loop integrals

$$
\delta\left(k^{2}-m^{2}\right) \sim \frac{1}{2 \pi i}\left(\frac{1}{k^{2}-m^{2}-i 0}-\frac{1}{k^{2}-m^{2}+i 0}\right)
$$

- Decomposition of the dot products to obtain scalar integrals
- Identity relations among scalar integrals:IBPs (using LiteRed)
- Algebraic linear system of equations relating the integrals

```
    \Downarrow
Master integrals (MIs)
```

- Computation of MIS : Parametric integration \& Method of diff. eqn. (generic \& canonical)
- UV renormalization and mass factorization
- Numerical evaluation using suitable PDFs and FFs


## Solving master integrals

The main challenges are the MIs appearing in the double real emissions. We follow three different techniques to compute them.

## Parametric integration

Conventional method by choosing an appropriate frame of reference
In the end, we are left with the following type of angular integration

$$
\int_{0}^{\pi} d \theta \int_{0}^{\pi} d \phi \frac{(\sin \theta)^{d-3}(\sin \phi)^{d-4}}{(a+b \cos \theta)(A+B \cos \theta+C \cos \phi \sin \theta)}
$$

- For $a=b$ and $A^{2}=B^{2}+C^{2}$ (which is the case for the RR MIs in leading-color), this angular integration results into simple hypergeometric functions.
- For the rest of the MIs, either $a \neq b$ or $A^{2} \neq B^{2}+C^{2}$ which may result into functions of higher complexity. So, we use the method of DE to solve these MIs.
- However, we consider the threshold limit of these MIs and compute them using parametric integration which will be used as the boundary conditions for the DE.


## The method of differential equations

A Feynman integral is a function of spacetime dimension $d$ and kinematic invariants $x, z$.

$$
J_{i}=\mathcal{N} \int \frac{d^{d} l_{1}}{(2 \pi)^{d}} \frac{d^{d} l_{2}}{(2 \pi)^{d}} \frac{1}{l_{1}^{2} l_{2}^{2}\left(\left(l_{1}-l_{2}\right)^{2}-m^{2}\right)\left(l_{1}-q\right)^{2}\left(l_{2}-q\right)^{2}} \equiv f(d, x, z)
$$

The idea is to obtain a differential eqn. for the integral w.r.t. $x, z$ and solve it.

## The method of differential equations

A Feynman integral is a function of spacetime dimension $d$ and kinematic invariants $x, z$.

$$
J_{i}=\mathcal{N} \int \frac{d^{d} l_{1}}{(2 \pi)^{d}} \frac{d^{d} l_{2}}{(2 \pi)^{d}} \frac{1}{l_{1}^{2} l_{2}^{2}\left(\left(l_{1}-l_{2}\right)^{2}-m^{2}\right)\left(l_{1}-q\right)^{2}\left(l_{2}-q\right)^{2}} \equiv f(d, x, z)
$$

The idea is to obtain a differential eqn. for the integral w.r.t. $x, z$ and solve it.

$$
\begin{aligned}
\frac{d}{d z} J_{i}= & \text { some combinations of integrals } \\
& \Downarrow \text { IBP identities/reduction } \\
& =\sum_{j=1}^{n} c_{i j} J_{j}
\end{aligned}
$$

$c_{i j}$ 's are rational function of $d, x$ and $z$.

## The method of differential equations

A Feynman integral is a function of spacetime dimension $d$ and kinematic invariants $x, z$.

$$
J_{i}=\mathcal{N} \int \frac{d^{d} l_{1}}{(2 \pi)^{d}} \frac{d^{d} l_{2}}{(2 \pi)^{d}} \frac{1}{l_{1}^{2} l_{2}^{2}\left(\left(l_{1}-l_{2}\right)^{2}-m^{2}\right)\left(l_{1}-q\right)^{2}\left(l_{2}-q\right)^{2}} \equiv f(d, x, z)
$$

The idea is to obtain a differential eqn. for the integral w.r.t. $x, z$ and solve it.

$$
d_{z}\left(\begin{array}{l}
J_{1} \\
J_{2} \\
J_{3} \\
J_{4} \\
\vdots \\
J_{n}
\end{array}\right)=\left[\begin{array}{cccccc}
\bullet & \bullet & \bullet & \bullet & \cdots & \bullet \\
\bullet & \bullet & \bullet & \bullet & \cdots & \bullet \\
\bullet & \bullet & \bullet & \bullet & \cdots & \bullet \\
\bullet & \bullet & \bullet & \cdots & \bullet \\
\vdots & \vdots & \vdots & \vdots & \ddots & \vdots \\
\bullet & \bullet & \bullet & \bullet & \cdots & \bullet
\end{array}\right]\left(\begin{array}{l}
J_{1} \\
J_{2} \\
J_{3} \\
J_{4} \\
\vdots \\
J_{n}
\end{array}\right)
$$

$$
d_{z} \mathbb{J}=\mathbb{A}(d, z) \mathbb{J}
$$

## The method of differential equations

A Feynman integral is a function of spacetime dimension $d$ and kinematic invariants $x, z$.

$$
J_{i}=\mathcal{N} \int \frac{d^{d} l_{1}}{(2 \pi)^{d}} \frac{d^{d} l_{2}}{(2 \pi)^{d}} \frac{1}{l_{1}^{2} l_{2}^{2}\left(\left(l_{1}-l_{2}\right)^{2}-m^{2}\right)\left(l_{1}-q\right)^{2}\left(l_{2}-q\right)^{2}} \equiv f(d, x, z)
$$

The idea is to obtain a differential eqn. for the integral w.r.t. $x, z$ and solve it.

$$
d_{z}\left(\begin{array}{l}
J_{1} \\
J_{2} \\
J_{3} \\
J_{4} \\
\vdots \\
J_{n-1} \\
J_{n}
\end{array}\right)=\left[\begin{array}{ccccccc}
\bullet & \bullet & \bullet & \bullet & \cdots & \bullet & \bullet \\
0 & \bullet & \bullet & \bullet & \cdots & \bullet & \bullet \\
0 & \bullet & \bullet & \bullet & \cdots & \bullet & \bullet \\
0 & 0 & 0 & \bullet & \cdots & \bullet & \bullet \\
\vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\
0 & 0 & 0 & 0 & \cdots & \bullet & \bullet \\
0 & 0 & 0 & 0 & \cdots & 0 & \bullet
\end{array}\right]\left(\begin{array}{l}
J_{1} \\
J_{2} \\
J_{3} \\
J_{4} \\
\vdots \\
J_{n-1} \\
J_{n}
\end{array}\right)
$$

The bullets $(\bullet)$ indicate a non-zero rational function of $d, x$ and $z$.
To solve such a system, it would be best to organize it in such a way that it diagonalizes, or at least it takes a block-triangular form. Then, it can be solved using bottom-up approach.

The results are obtained in terms of iterated integrals (HPLs/GPLs).

## The method of differential equations: canonical form

We have a $21 \times 21$ system of DEs, on which we perform canonical transformation using Libra


After getting the $\epsilon$-form, we performed path-ordered integrations. Note that we have 4 square-roots in addition to 10 kernels in $x^{\prime}$ and $z^{\prime}$. As mentioned earlier, we obtained the boundaries at $x^{\prime} \rightarrow 1$ and $z^{\prime} \rightarrow 1$. We have computed all the MIs in this method.

## The method of differential equations: generic

Considering the MIs as given by LiteRed, we find a few sub-systems which are coupled. And, as expected, the homogeneous solutions contain the square-roots. The goal is to avoid them.

## The method of differential equations: generic

Considering the MIs as given by LiteRed, we find a few sub-systems which are coupled. And, as expected, the homogeneous solutions contain the square-roots. The goal is to avoid them.

Iterated integrals are one-dimensional. They can be computed with great precision in a short amount of time. Besides, they have the following properties:
(a) Shuffle algebra: Allows to obtain a basis for a set of iterated integrals. Reduction to such a basis is extremely effective to reduce the computation time by few times.
(b) Scaling invariance : Allows to convert the limit of these integrals from kinematical variables ( $z$ ) to constants (1). This makes the integration really precise.

## The method of differential equations: generic

Considering the MIs as given by LiteRed, we find a few sub-systems which are coupled. And, as expected, the homogeneous solutions contain the square-roots. The goal is to avoid them.

Iterated integrals are one-dimensional. They can be computed with great precision in a short amount of time. Besides, they have the following properties:
(a) Shuffle algebra: Allows to obtain a basis for a set of iterated integrals. Reduction to such a basis is extremely effective to reduce the computation time by few times.
(b) Scaling invariance : Allows to convert the limit of these integrals from kinematical variables ( $z$ ) to constants (1). This makes the integration really precise.

First few MIs were solved in terms of GPLs
$\Downarrow$

What happens when we have (multiple) square-roots?!

## Rationalization

## Rationalizable

- Find a suitable transformation

Let's consider $\sqrt{\left(1-z^{\prime}\right)^{2}+4 x^{\prime} z^{\prime}}$.
We can use the transformation $x^{\prime} \rightarrow$ $\left(1-u^{2}\right) / 4 / t /(1-t), z^{\prime} \rightarrow 1-1 / t$ for this.

## Partly-rationalizable

- A single transformation can not rationalize all square-roots simultaneously.
$\Downarrow$
Square-roots will be present in the iterated integrals.

1) We can accept 'the fact' and evaluate them with appropriate analytic continuation.
or
2) Instead of using a single transformation rule to rationalize them, we write the system (each MI) as sum of functions of dependent variables and treat them separately. As a result, each sub-system has alphabet with 'good' letters with different argument.

## Solving master integrals: summary

- We followed three different techniques to compute the MIs for double real emissions.
- We have perfect agreement among the results from different methods.
- For real-virtual MIs also, we compute them using parametric method, where we use exact expressions for one-loop integrals followed by parametric phase-space integration. We also use method of DE and find perfect agreement.

To obtain the finite partonic cross-section

## Combine everything, UV renormalization and mass factorization



- We compute relevant Feynman diagrams \& corresponding Feynman integrals analytically.
- After summing all contributions for the Feynman diagrams, we can express the results in terms of MPLs with square-root arguments.
- The resultant diagrams contain two types of divergences: dimensional and kinematical


## Combine everything, UV renormalization and mass factorization

- The contributions also have divergences for $x^{\prime} \rightarrow 1$ and $z^{\prime} \rightarrow 1$. We use

$$
(1-x)^{-1+n \epsilon}=\frac{1}{n \epsilon} \delta(1-x)+\sum_{k=0}^{\infty} \frac{\epsilon^{k}}{k!}\left[\frac{\ln ^{k}(1-x)}{(1-x)}\right]_{+}
$$

where

$$
\int_{0}^{1} d x f[x] g[x]_{+}=\int_{0}^{1} d x(f[x]-f[1]) g[x]
$$

- Of course, for our case, we have double distributions.
- We also find functions that require special care for analytic continuation. We use partial fractioning and theta function to separate different sectors and implement correct analytic continuation. For example,

$$
\left(z^{\prime}-x^{\prime}\right)^{\epsilon}=\left|z^{\prime}-x^{\prime}\right|^{\epsilon}\left(\theta\left(z^{\prime}-x^{\prime}\right)+(-1+i \varepsilon)^{\epsilon} \theta\left(x^{\prime}-z^{\prime}\right)\right)
$$

- The divergences at $x^{\prime}=z^{\prime}$ or $x^{\prime}+z^{\prime}=1$ are spurious and eventually they cancel.


## Combine everything, UV renormalization and mass factorization

## unpolarized

- The sources of dimensional divergences are of UV \& IR type. UV renormalization takes care of the UV divergences. In case of unpolarized SIDIS, the UV renormalization is trivial.
- KLN theorem ensures the cancellation of all soft and some collinear divergences. The residual collinear divergences are absorbed by mass factorization, using appropriate splitting kernels.

$$
\frac{\hat{\sigma}_{I, a b}(\epsilon)}{x^{I-1}}=\Gamma_{c \leftarrow a}\left(\mu_{F}^{2}, \epsilon\right) \otimes \mathcal{F}_{I, c d}\left(\mu_{F}^{2}, \epsilon\right) \tilde{\otimes} \tilde{\Gamma}_{d \leftarrow b}\left(\mu_{F}^{2}, \epsilon\right)
$$

polarized

- The $\gamma_{5}$ in polarized SIDIS needs special care. We use Larin's prescription and we appropriately multiply the finite renormalization constants.
- For polarized case, we use the spin-dependent Altarelli-Parisi kernels for initial states. Since, the partonic cross-section is computed in Larin's scheme, the polarized space-like kernels also need to be considered in the same scheme. For final states, the spin-averaged time-like kernels are considered in the $\overline{\mathrm{MS}}$ scheme.
- In the end, we transform our results obtained in Larin's scheme to $\overline{M S}$ scheme.



## Results

- We have computed full NNLO corrections to unpolarized SIDIS (all processes) and the polarized SIDIS where the initial state is polarized.
- In both cases, we find good perturbative convergence of the NNLO corrections, and reduction of renormalization and factorization scale dependence.



The K-factors for corrections from different perturbative order for quark initiated process to a fragmenting quark. We have shown here only the nonsinglet contributions for NNLO.

Detailed phenomenology is on the way!

## Results

- We have computed full NNLO corrections to unpolarized SIDIS (all processes) and the polarized SIDIS where the initial state is polarized.
- In both cases, we find good perturbative convergence of the NNLO corrections, and reduction of renormalization and factorization scale dependence.


The contributions from different perturbative orders in polarized SIDIS with $\sqrt{s}=140 G e V$. We have integrated over $x$ between 0.40 to 0.62 and $y$ between 0.30 to 0.50 .

Detailed phenomenology is on the way!

## Checks

- The master integrals were computed using different methods with perfect agreement!
- Mass factorization (universal) removes all remaining infrared singularities!
- Successful checks with available results in the threshold limit! SV and NSV results match with the literature for $F_{1}, F_{2}$ and $g_{1}$.
- The non-singlet part of $g_{1}$ is equal to $F_{3}$ which we have explicitly checked.
- The constraint $\left(z^{\prime}\right)$ can be integrated in our analytic result. We found perfect agreement with the fully inclusive result.
- We also have perfect agreement for unpolarized SIDIS with the results by Bonino et al.
talk by Bonino

- EIC will unravel the mysteries of strong force. Theoretical precision studies are extremely necessary to fully exploit the EIC data.
- Our current (well-tested) theoretical understanding (the SM) is constrained by its perturbative nature and hence, higher order perturbative corrections are necessary to achieve precise theoretical predictions.
- In this talk, we have presented the first results on NNLO QCD corrections to unpolarized and polarized SIDIS.
- The NNLO predictions will be extremely important to extract PDFs and FFs very precisely.
- Precision predictions for polarized PDFs will be crucial to solve the proton spin puzzle.
- Aside the phenomenological impact of the result, it also sets a milestone for the computational technique.

Thank you for your attention!

