# Experimental Tests of Multimetric Gravity Gravitational Waves and the Cosmos

#### Manuel Hohmann

II. Institut für theoretische Physik

Universität Hamburg der forschung | der lehre | der bildung

DESY Theory Workshop 29 September 2011

- ACDM model: only 5% visible matter. [Komatsu et al. '09]
  - Dark matter needed to explain "missing mass" in galaxies.
  - Dark energy needed to explain accelerating expansion.
  - Constituents of dark universe are unknown!

- ACDM model: only 5% visible matter. [Komatsu et al. '09]
  - Dark matter needed to explain "missing mass" in galaxies.
  - Dark energy needed to explain accelerating expansion.
  - Constituents of dark universe are unknown!
- Idea here: Additional "dark, negative mass" standard model copy.
- Only interaction between both copies: repulsive gravity.
- Universe contains equal amounts of both types of mass.

- ACDM model: only 5% visible matter. [Komatsu et al. '09]
  - Dark matter needed to explain "missing mass" in galaxies.
  - Dark energy needed to explain accelerating expansion.
  - Constituents of dark universe are unknown!
- Idea here: Additional "dark, negative mass" standard model copy.
- Only interaction between both copies: repulsive gravity.
- Universe contains equal amounts of both types of mass.
- Dark galaxies "push" visible matter & light towards visible galaxies.
   ⇒ Potential explanation of dark matter!
- Mutual repulsion between galaxies drives accelerating expansion.
   ⇒ Potential explanation of dark energy!

- ACDM model: only 5% visible matter. [Komatsu et al. '09]
  - Dark matter needed to explain "missing mass" in galaxies.
  - Dark energy needed to explain accelerating expansion.
  - Constituents of dark universe are unknown!
- Idea here: Additional "dark, negative mass" standard model copy.
- Only interaction between both copies: repulsive gravity.
- Universe contains equal amounts of both types of mass.
- Dark galaxies "push" visible matter & light towards visible galaxies.
   ⇒ Potential explanation of dark matter!
- Mutual repulsion between galaxies drives accelerating expansion.
   ⇒ Potential explanation of dark energy!
- Simple idea: two metrics  $g_{ab}^{\pm}$ , two standard model copies  $\Psi^{\pm}$ .
- No-go theorem forbids bimetric repulsive gravity! [MH, M. Wohlfarth '09]
- Solution:  $N \ge 3$  metrics  $g_{ab}^{l}$  and standard model copies  $\Psi^{l}$ .

# Action and equations of motion

- Matter action: sum of standard model actions.
- Gravitational action:

$$S_G[g^1, ..., g^N] = rac{1}{2} \int d^4 x \sqrt{g_0} \left[ \sum_{l,J=1}^N (x + y \delta^{lJ}) g^{lij} R^J_{ij} + F(S^{lJ}) 
ight]$$

- Symmetric volume form  $g_0 = (g^1 g^2 \dots g^N)^{1/N}$ .
- $F(S^{IJ})$  quadratic in connection difference tensors  $S^{IJ} = \Gamma^{I} \Gamma^{J}$ .

# Action and equations of motion

- Matter action: sum of standard model actions.
- Gravitational action:

$$S_G[g^1, ..., g^N] = rac{1}{2} \int d^4 x \sqrt{g_0} \left[ \sum_{l,J=1}^N (x + y \delta^{lJ}) g^{lij} R^J_{ij} + F(S^{lJ}) 
ight]$$

• Symmetric volume form  $g_0 = (g^1 g^2 \dots g^N)^{1/N}$ .

*F*(*S<sup>IJ</sup>*) quadratic in connection difference tensors *S<sup>IJ</sup>* = Γ<sup>I</sup> − Γ<sup>J</sup>.
 ⇒ Equations of motion:

$$T'_{ab} = \sqrt{\frac{g_0}{g'}} \left[ -\frac{1}{2N} g'_{ab} \sum_{J,K=1}^N (x + y\delta^{JK}) g^{Jij} R^K_{ij} + \sum_{J=1}^N (x + y\delta^{JJ}) R^J_{ab} \right]$$
  
+ terms linear in  $\nabla^I S^{JK}$  + terms quadratic in  $S^{IJ}$ .

# Action and equations of motion

- Matter action: sum of standard model actions.
- Gravitational action:

$$S_G[g^1, \dots, g^N] = rac{1}{2} \int d^4 x \sqrt{g_0} \left[ \sum_{l,J=1}^N (x + y \delta^{lJ}) g^{lij} R^J_{ij} + F(S^{lJ}) 
ight]$$

• Symmetric volume form  $g_0 = (g^1 g^2 \dots g^N)^{1/N}$ .

*F*(*S<sup>IJ</sup>*) quadratic in connection difference tensors *S<sup>IJ</sup>* = Γ<sup>I</sup> − Γ<sup>J</sup>.
 ⇒ Equations of motion:

$$T'_{ab} = \sqrt{\frac{g_0}{g'}} \left[ -\frac{1}{2N} g'_{ab} \sum_{J,K=1}^N (x + y\delta^{JK}) g^{Jij} R^K_{ij} + \sum_{J=1}^N (x + y\delta^{JJ}) R^J_{ab} \right]$$
  
+ terms linear in  $\nabla^I S^{JK}$  + terms quadratic in  $S^{IJ}$ .

⇒ Repulsive Newtonian limit for  $N \ge 3$ . [MH, M. Wohlfarth '10] ⇒ Post-Newtonian limit consistent with solar system. [MH, M. Wohlfarth '10]

# Simple cosmological model

- Standard cosmology:
  - Homogeneous and isotropic universe (FLRW metric).
  - Matter content: perfect fluid.
  - Early universe: radiation; late universe: dust.
  - Copernican principle: common evolution for all matter sectors.

# Simple cosmological model

- Standard cosmology:
  - Homogeneous and isotropic universe (FLRW metric).
  - Matter content: perfect fluid.
  - Early universe: radiation; late universe: dust.
  - Copernican principle: common evolution for all matter sectors.
- Friedmann equation:

$$\rho = \frac{3}{2-N} \left( \frac{\dot{a}^2}{a^2} + \frac{k}{a^2} \right).$$

- ⇒ Positive matter density  $\rho$  > 0 requires k = -1 and  $\dot{a}^2 < 1$ .
- $\Rightarrow$  No solutions for k = 0 or k = 1.

# Simple cosmological model

- Standard cosmology:
  - Homogeneous and isotropic universe (FLRW metric).
  - Matter content: perfect fluid.
  - Early universe: radiation; late universe: dust.
  - Copernican principle: common evolution for all matter sectors.
- Friedmann equation:

$$\rho = \frac{3}{2 - N} \left( \frac{\dot{a}^2}{a^2} + \frac{k}{a^2} \right).$$

- ⇒ Positive matter density  $\rho$  > 0 requires k = -1 and  $\dot{a}^2 < 1$ .
- $\Rightarrow$  No solutions for k = 0 or k = 1.
  - Acceleration equation:

$$\frac{\ddot{a}}{a}=\frac{N-2}{6}\left(\rho+3p\right).$$

 $\Rightarrow$  Acceleration must be positive.

# **Explicit solution**

Equation of state: p = ωρ; dust: ω = 0, radiation: ω = 1/3.
General solution using conformal time η defined by dt = a dη:



# Cosmological parameters

- Friedmann equation:  $(2 N)\Omega_M + \Omega_K = 1$ .
- Matter density:

$$\Omega_{M}=rac{
ho_{0}}{3H_{0}^{2}}\sim {
m sinh^{-2}}\left(rac{3\omega+1}{2}\eta_{0}
ight).$$

- Curvature parameter:  $\Omega_{K} = -\frac{k}{a_{0}^{2}H_{0}^{2}} = \frac{1}{\dot{a}^{2}(t_{0})} \rightarrow 1.$
- Fitting of supernova data: [Amanullah et al. '10]



- Structure formation in ACDM not fully understood:
  - Missing dwarf problem. [Moore et al. '99]
  - Core-cusp-problem. [Dubinski, Carlberg '91; Navarro et al. '96]

- Structure formation in ACDM not fully understood:
  - Missing dwarf problem. [Moore et al. '99]
  - Core-cusp-problem. [Dubinski, Carlberg '91; Navarro et al. '96]
- Structure formation in multimetric gravity:
  - Perturbation of cosmological background solution.
  - Model dust matter by point particles.
  - Interaction between point particles given by Newtonian limit.

- Structure formation in ACDM not fully understood:
  - Missing dwarf problem. [Moore et al. '99]
  - Core-cusp-problem. [Dubinski, Carlberg '91; Navarro et al. '96]
- Structure formation in multimetric gravity:
  - Perturbation of cosmological background solution.
  - Model dust matter by point particles.
  - Interaction between point particles given by Newtonian limit.
- Implementation:
  - Large particle number requires high computing power.
  - ⇒ Use GPU computing!

- Structure formation in ACDM not fully understood:
  - Missing dwarf problem. [Moore et al. '99]
  - Core-cusp-problem. [Dubinski, Carlberg '91; Navarro et al. '96]
- Structure formation in multimetric gravity:
  - Perturbation of cosmological background solution.
  - Model dust matter by point particles.
  - Interaction between point particles given by Newtonian limit.
- Implementation:
  - Large particle number requires high computing power.
  - ⇒ Use GPU computing!
- Results:
  - Galactic clusters and filament-like structures.
  - Seemingly empty voids contain "invisible" matter.
  - $\Rightarrow$  Repulsive gravity effects from galactic voids.
  - ⇒ Negative gravitational lenses in galactic voids?



Manuel Hohmann (Uni Hamburg)

Multimetric gravity



Manuel Hohmann (Uni Hamburg)

Multimetric gravity





Manuel Hohmann (Uni Hamburg)

Multimetric gravity



Manuel Hohmann (Uni Hamburg)

Multimetric gravity





Manuel Hohmann (Uni Hamburg)

Multimetric gravity





Manuel Hohmann (Uni Hamburg)

Multimetric gravity





Manuel Hohmann (Uni Hamburg)

Multimetric gravity







Manuel Hohmann (Uni Hamburg)

Multimetric gravity













Manuel Hohmann (Uni Hamburg)

Multimetric gravity













































# Gravitational waves

- Gravitational waves are null. [MH '11]
- Polarizations classified by reps. of E(2). [Eardley, Lee, Lightman et al. '73]
- E(2) class of multimetric gravity depends on 3 parameters: [MH 11]



- Idea: Repulsive gravity might explain dark matter & dark energy.
- ⇒ Linearly PPN consistent multimetric repulsive gravity for  $N \ge 3$ .
- $\Rightarrow$  Cosmology features late-time acceleration and big bounce.
- ⇒ Structure formation features clusters and voids.
- $\Rightarrow$  Voids contain repulsive mass concentrations.
- $\Rightarrow$  Gravitational waves are null.
- $\Rightarrow$  E(2) class can be one of N<sub>2</sub>, N<sub>3</sub>, III<sub>5</sub>, II<sub>6</sub>.

# Outlook

- Solar system physics:
  - Develop full non-linear multimetric PPN formalism.
  - Examine further exact solutions (single point mass...).
- Cosmology:
  - Calculate further cosmological matching parameters.
  - Perform cosmological perturbation analysis to understand CMB.
- Structure formation:
  - Advanced simulation of structure formation including thermodynamics using GADGET-2 (Millennium Simulation).
  - Search for repulsive gravity sources in galactic voids through gravitational lensing.
- Gravitational waves:
  - Emission of gravitational waves from various sources.
  - Shapiro delay of gravitational waves in weak gravitational fields.