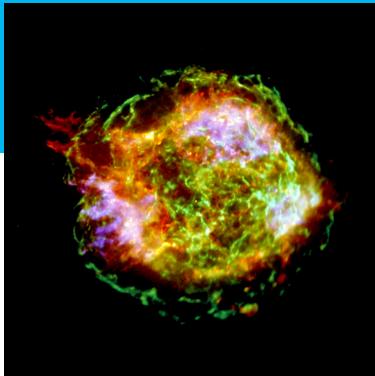


Gamma-Ray Astroparticle Physics: CTA and VERITAS

Stefan Schlenstedt, DESY

71^{rst} PRC, Open session

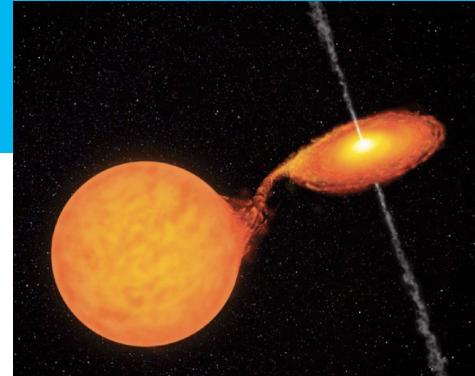
Hamburg, April 28, 2011



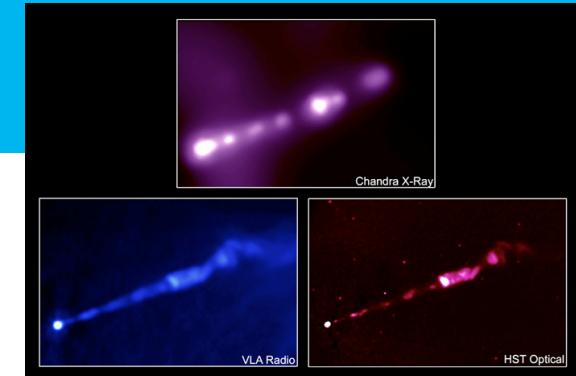
Super Novae R



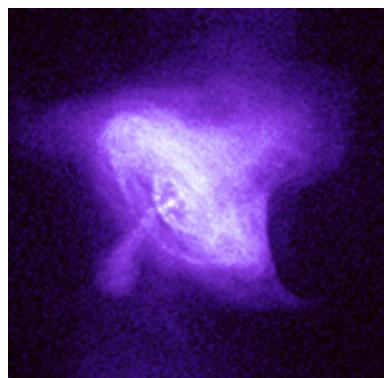
Gamma Ray B



Micro quasars



Active galactic nuclei

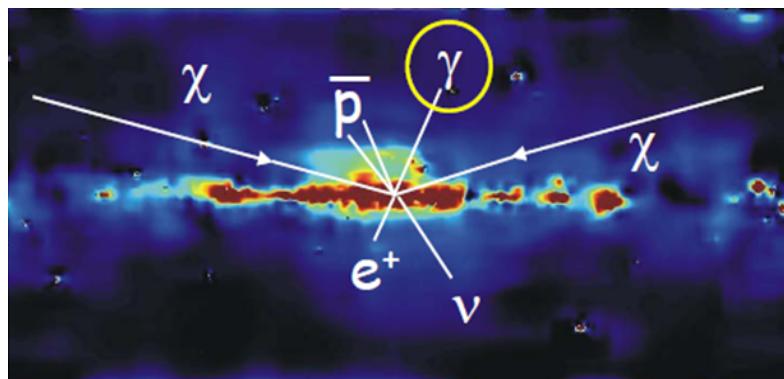


Pulsar (wind nebulae)

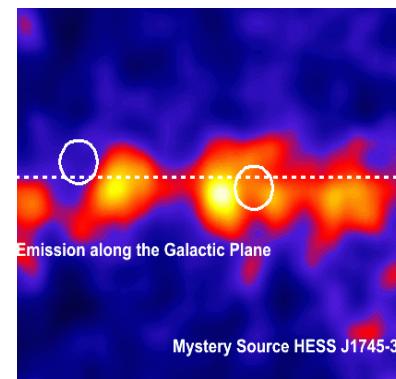
- > Sources of cosmic rays
- > Astrophysics of sources
- > Acceleration and propagation



Starburst galaxies



Dark matter



Unknown sources

The Cherenkov Telescope Array



- Increase sensitivity
- Extend energy range
- Improve angular resolution
- Observatory with flexible operation
- Arrays in North and South for full sky coverage

CTA: An advanced facility
for ground-based γ -ray
astronomy and astro-
particle physics

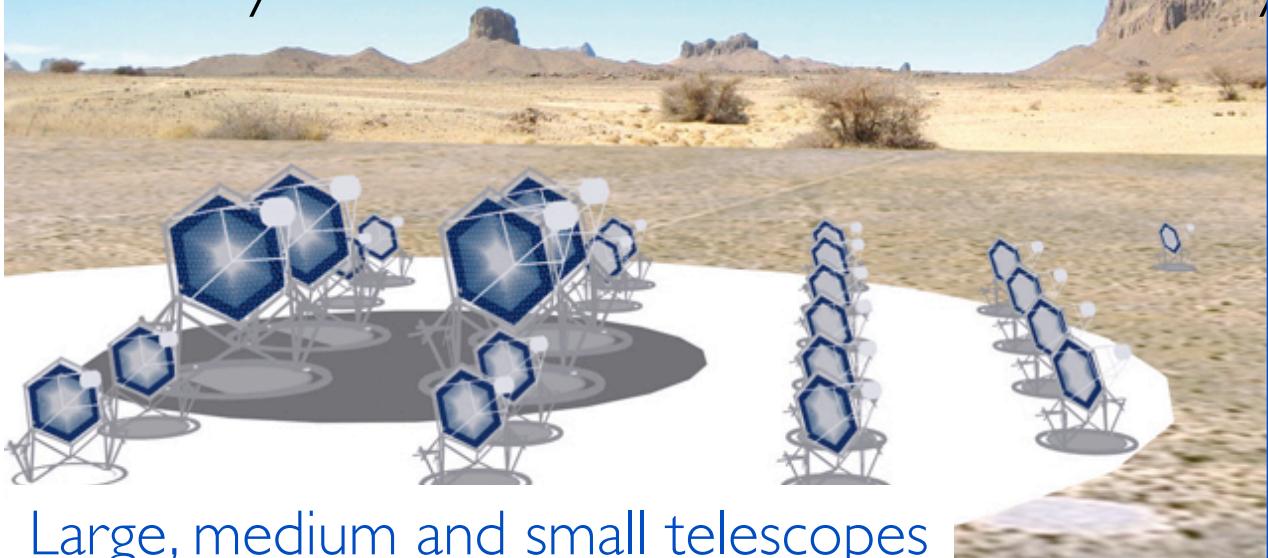


The Cherenkov Telescope Array



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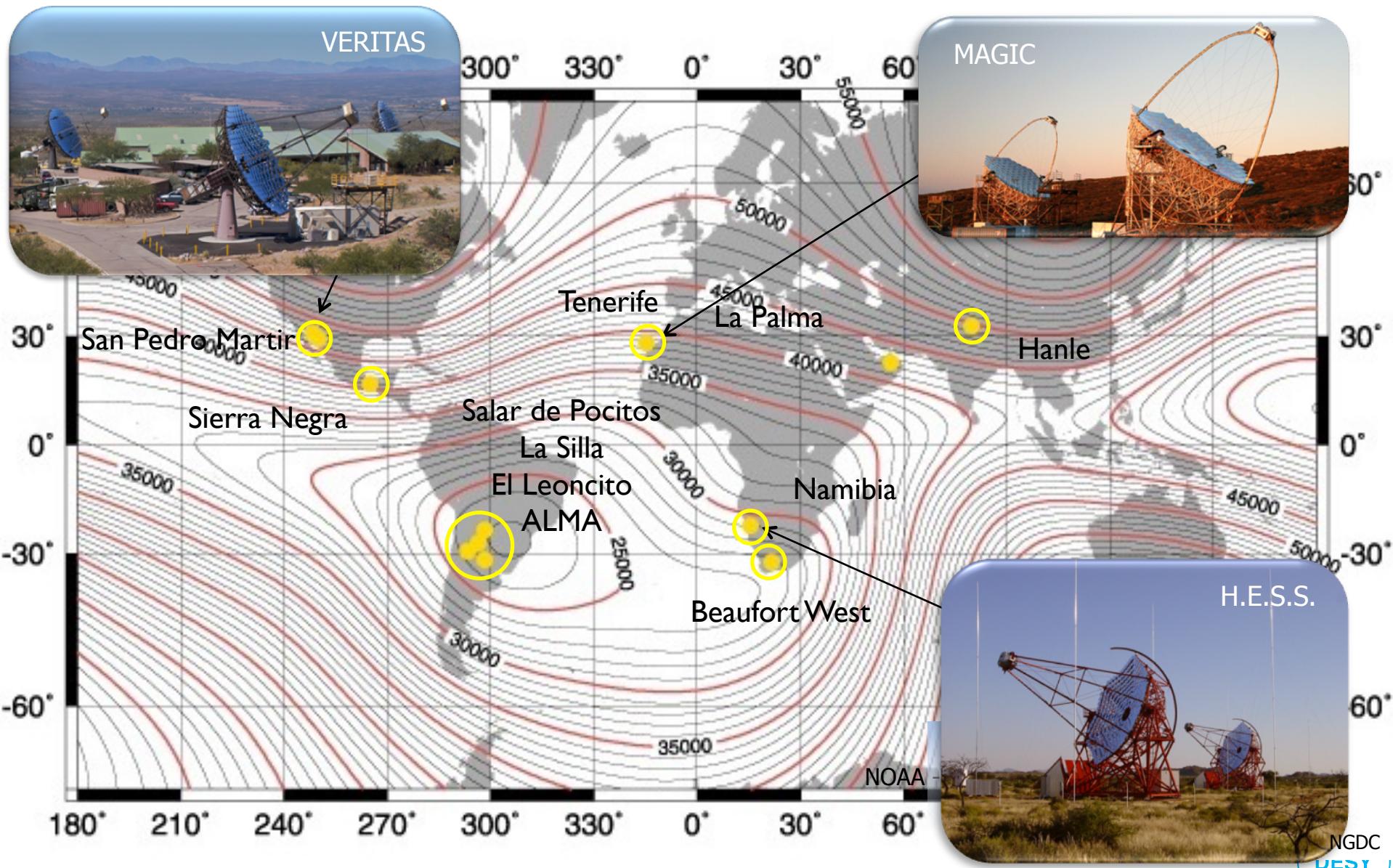


Large, medium and small telescopes

– 2013:
preparatory
phase, prototypes,
site decision

> 2013:
construction and
partial science
operation

CTA Site?



DESY Activities

> Hardware

- medium-size telescope
- electronics

> Array Control

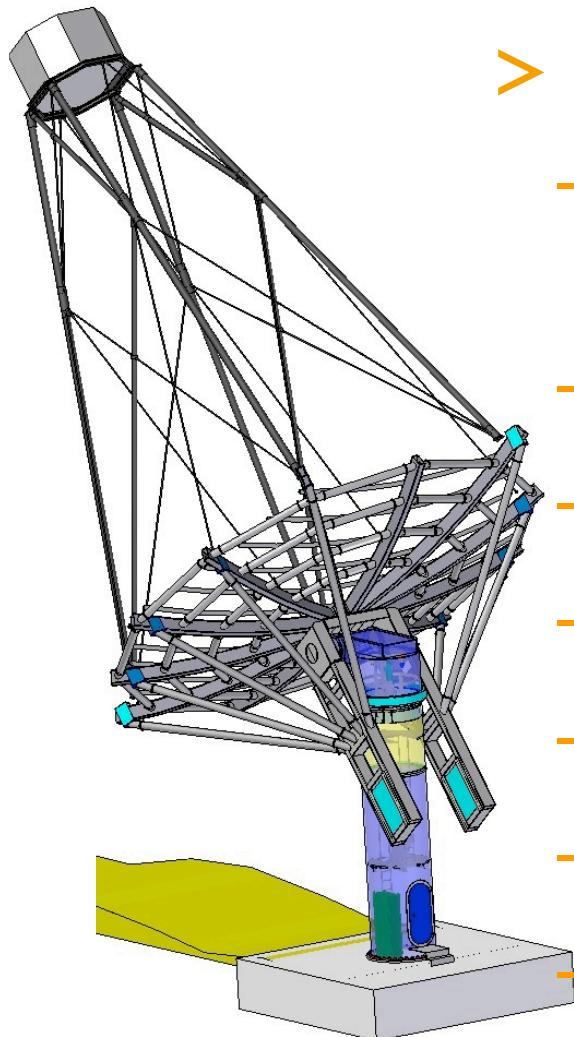
> Sensitivity and Reconstruction methods

> Analysis

> Computing

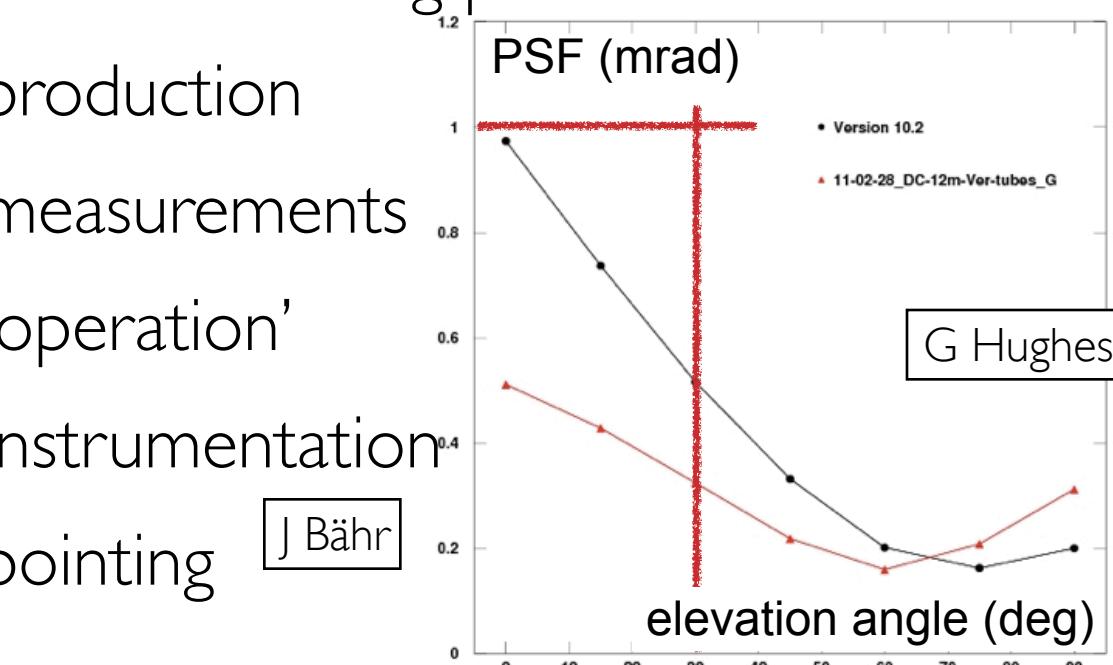


Medium Size Telescope



> Prototype 2011 in Berlin

- design: dish, camera structure, head, tower, foundation
- statics - building permit
- production
- measurements
- 'operation'
- instrumentation
- pointing



J Bähr

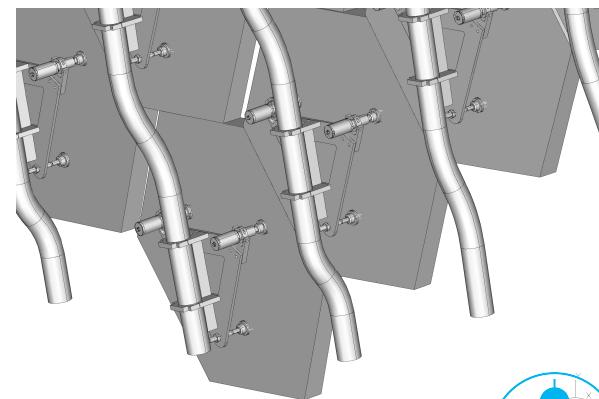
G Hughes

Medium Size Telescope

- > Pre-prototype = quarter-dish in Zeuthen
 - measure with photogrammetry and tachymetry
 - mirror mounting

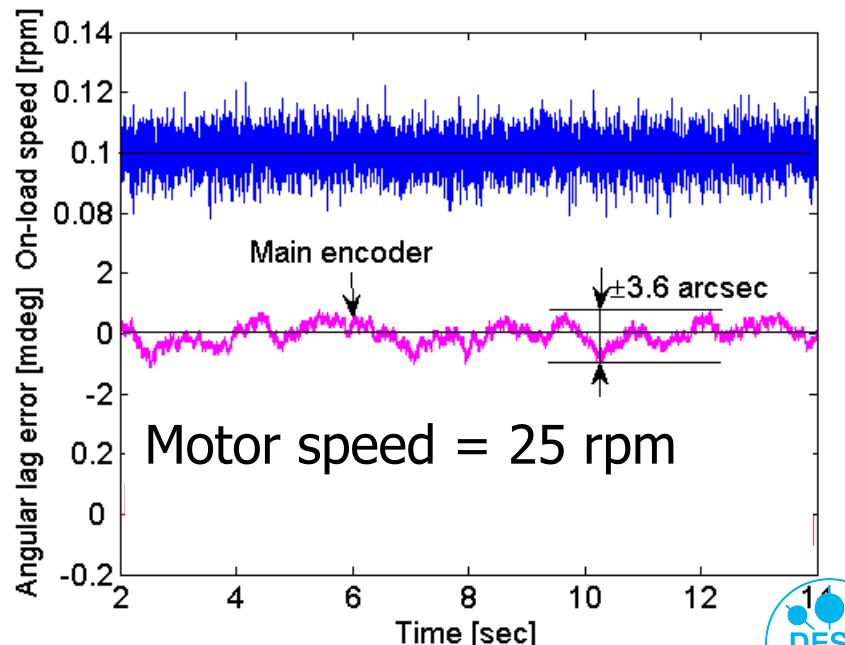
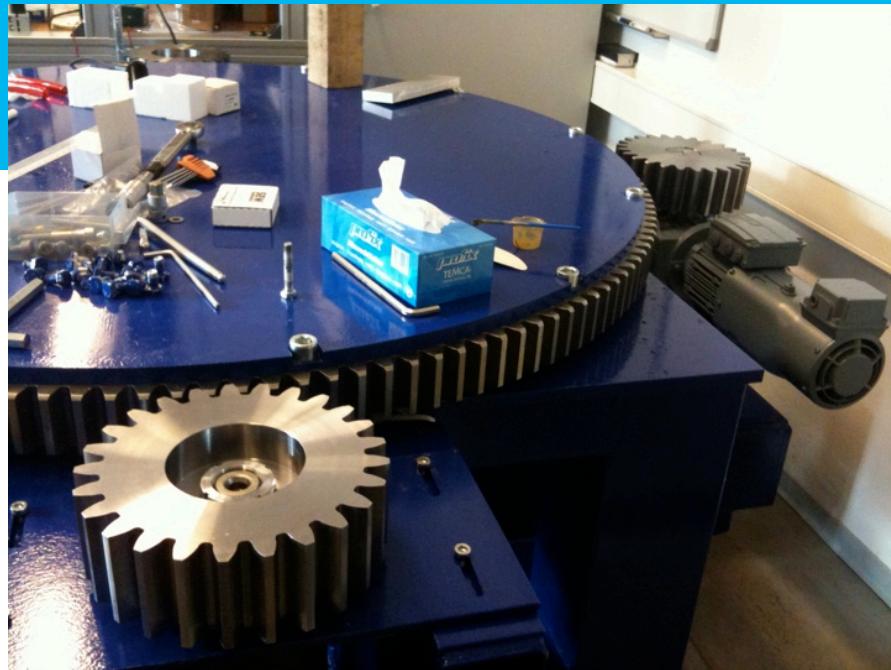


J Bähr



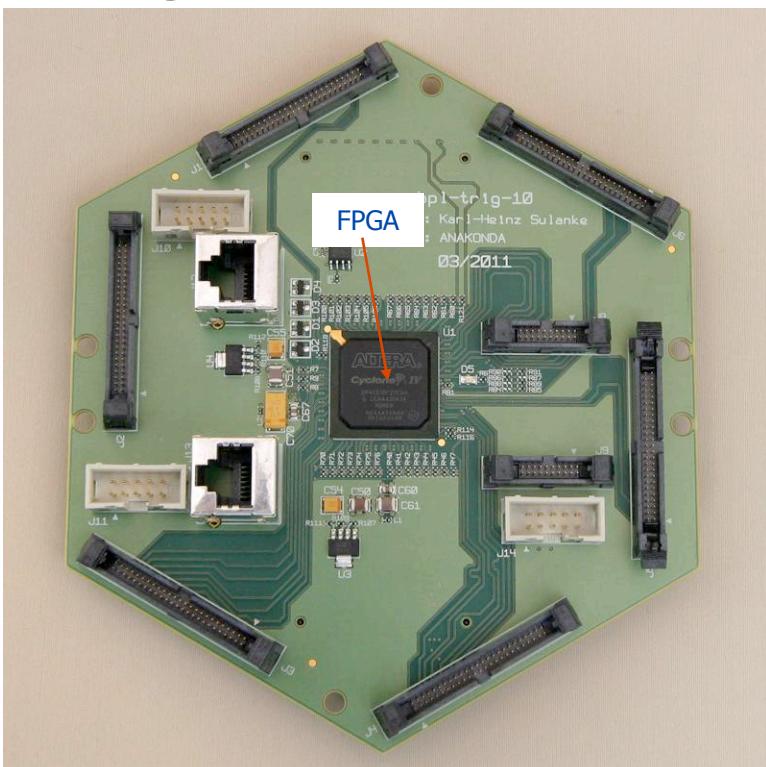
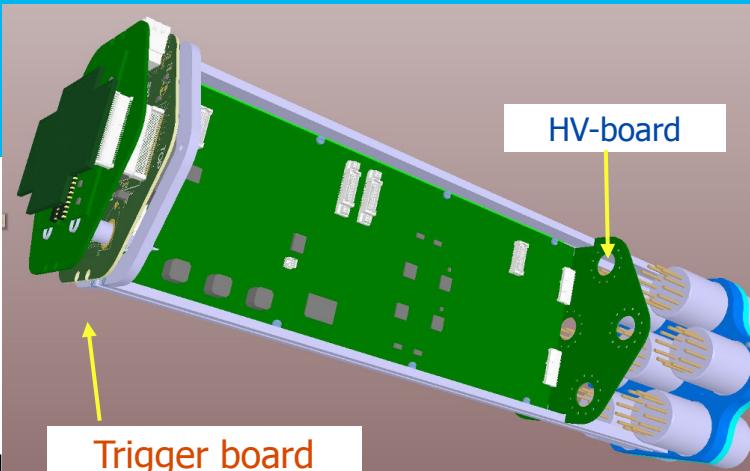
Medium Size Telescope

- > Drive system:
 - 2 azimuth and 4 elevation motors
 - speed range 1 ... 3600 rpm
 - high accuracy
- > Safety concept
 - active and passive
 - risk analysis
- > Energy management system

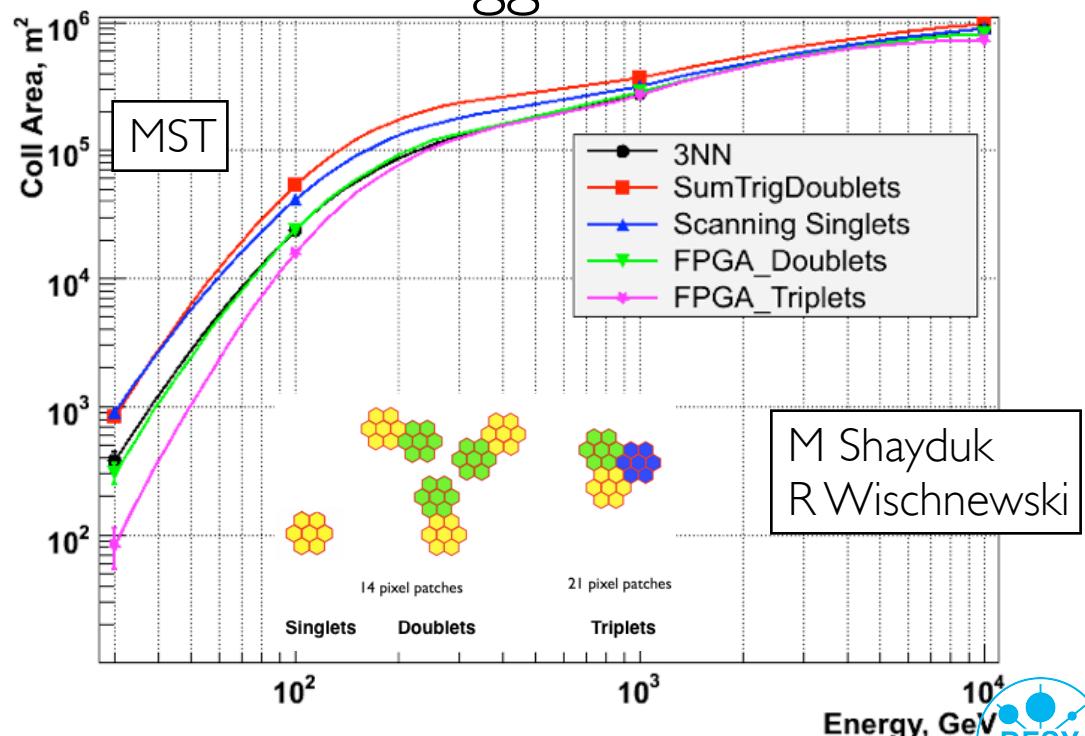


Electronics Development

- Camera photo-tube HV board
 - Digital trigger electronics (1 Hz signal with 120 MHz background)



> Trigger simulation



DESY Activities

➤ Hardware

➤ Array Control

➤ Sensitivity and
Reconstruction
methods

➤ Analysis

➤ Computing



Array Control

- > Evaluate software frameworks (ACS: ESO supported software framework for the ALMA telescopes)
- > Implement a prototype for Array Control Centre
 - drive system
 - mirror control
 - weather station
 - alignment system
 - camera simulator

B Behera
P Kostka
T Schmidt

CTA operation

DESY topics:

- Science management
- Operation planning
- Telescope operation
- Data acquisition
- On-site analysis
- Data reduction

G Maier
S Schlenstedt
P Wegner



DESY Activities

➤ Hardware

➤ Array Control

➤ Sensitivity studies

➤ New reconstruction
methods

➤ Analysis

➤ Computing



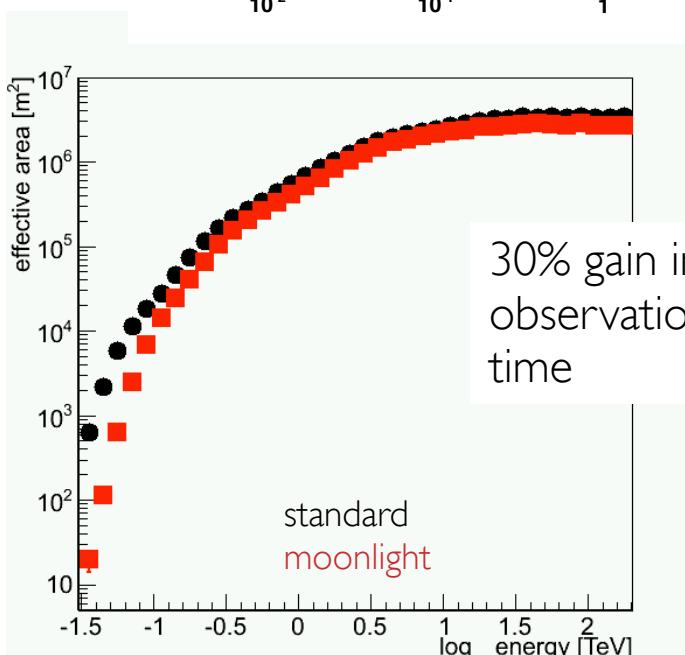
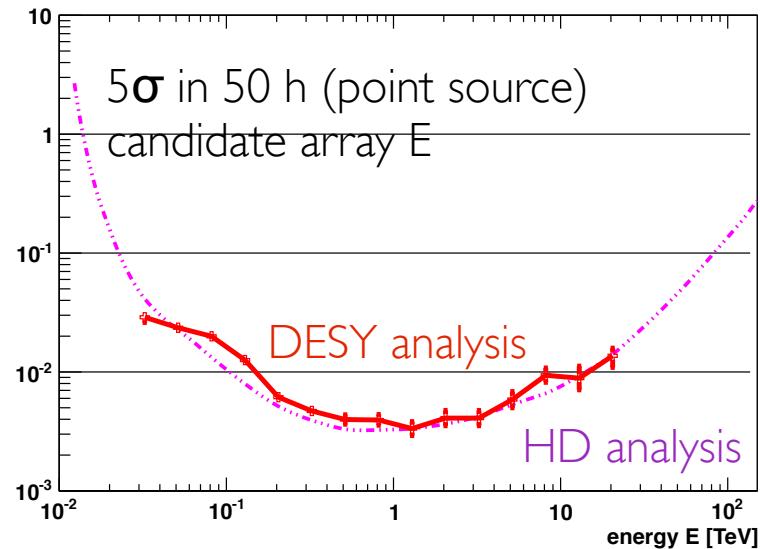
CTA Analysis Chain

- > Complete analysis chain for CTA
 - well tested and used for science publications in VERITAS
 - core developer team at DESY
 - interfaces to DESY trigger simulations and CTA-US simulation packages
 - very flexible in terms of detector geometry (used for VERITAS, AGIS, CTA, ...)
 - available to everybody in CTA
(in contrast to most of the tools in CTA)
- > Processing of and optimization with CTA MC simulations



CTA Sensitivity

- Processing and optimization of CTA MC simulations
- Study of different arrays; different observing (e.g., moonlight) and site conditions (e.g. height)



30% gain in
observation
time

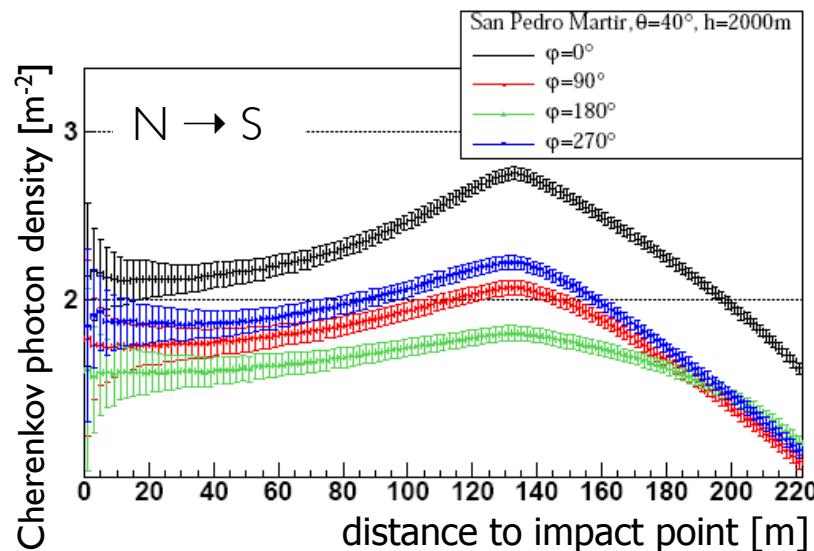
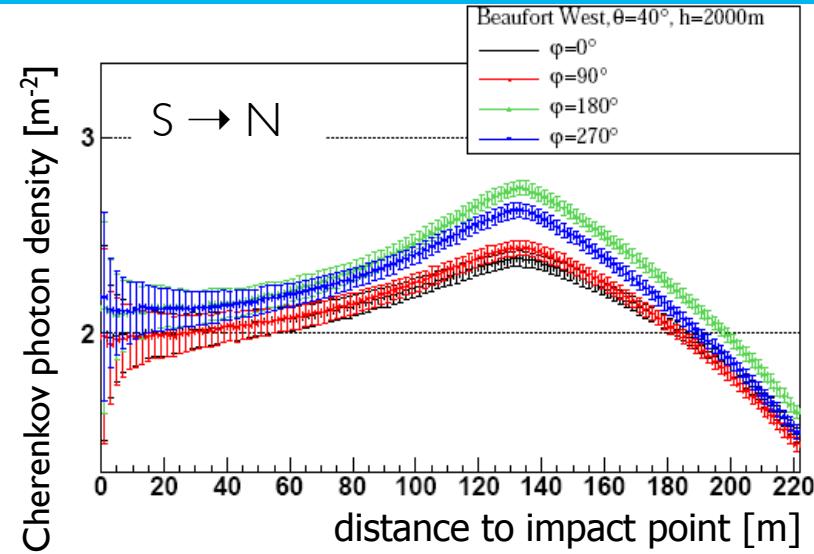
standard
moonlight

Geomagnetic Field (20,70) μ T

Beaufort West
H.E.S.S.
El Leoncito

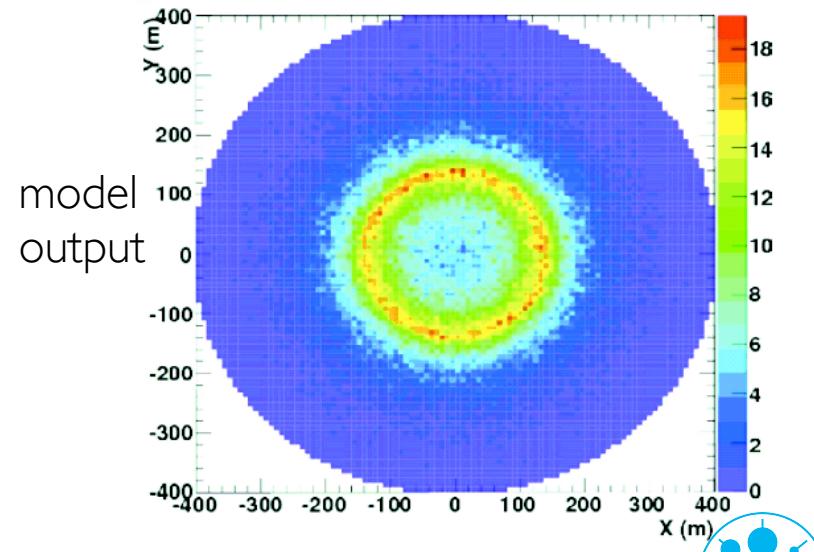
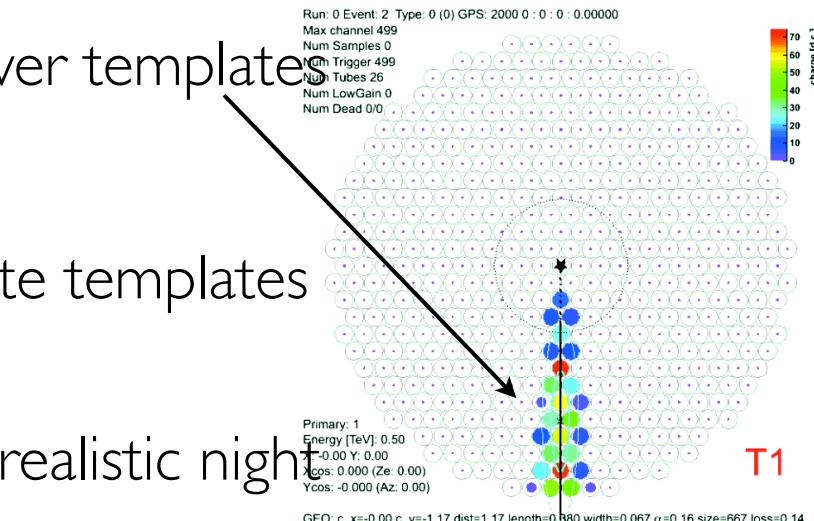
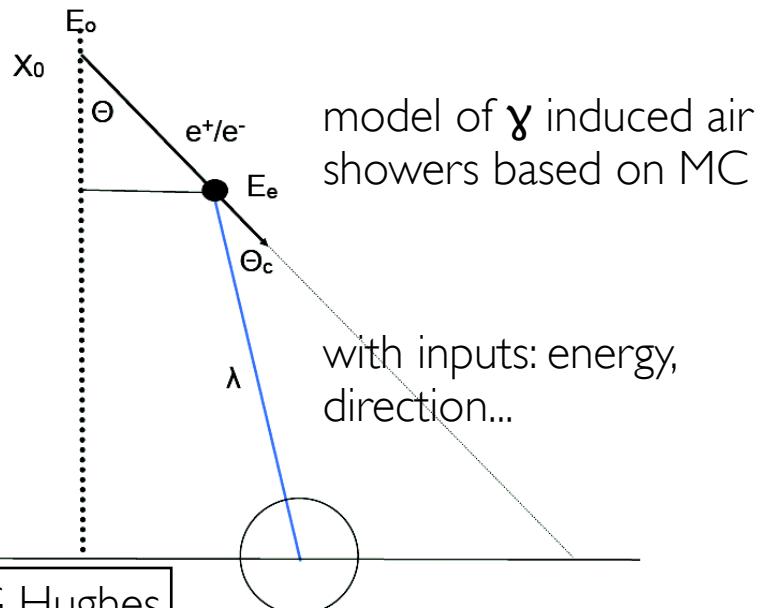
La Palma
San Pedro Martir

showers 50° above horizon



New Reconstruction Techniques

- Reconstruct γ -ray showers using shower templates
- Shown to improve sensitivity by 2x
- Use a semi-analytical model to generate templates
 - large parameter space of showers
- Compare templates to real data (and realistic night sky backgrounds and pedestals)



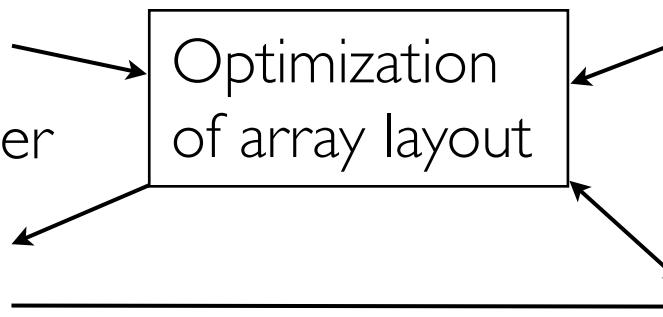
Site Optimization – Dual Mirror Telescopes

Baseline: Davies-Cotton

(H.E.S.S., MAGIC, VERITAS)

US-CTA: Schwarzschild-Couder

Energy & angular resolution

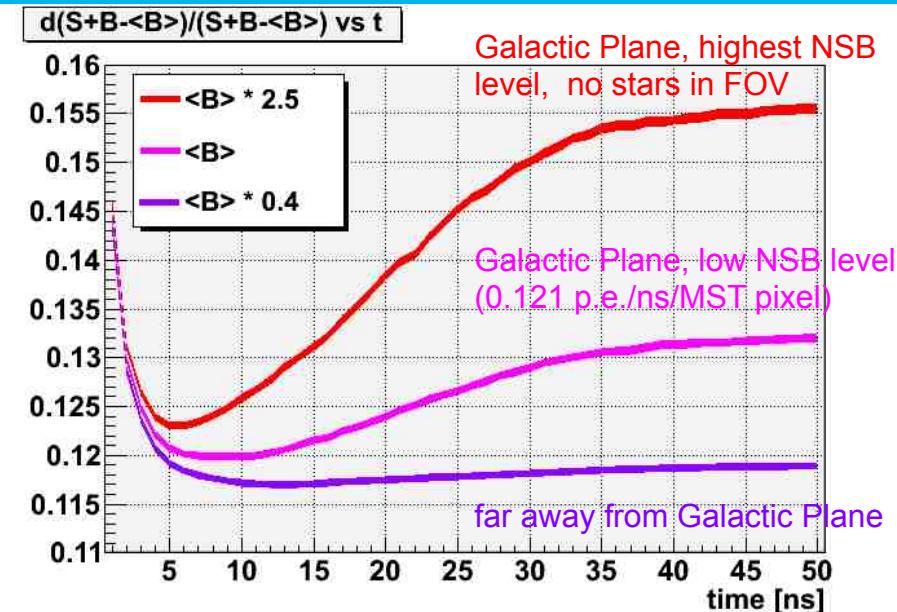
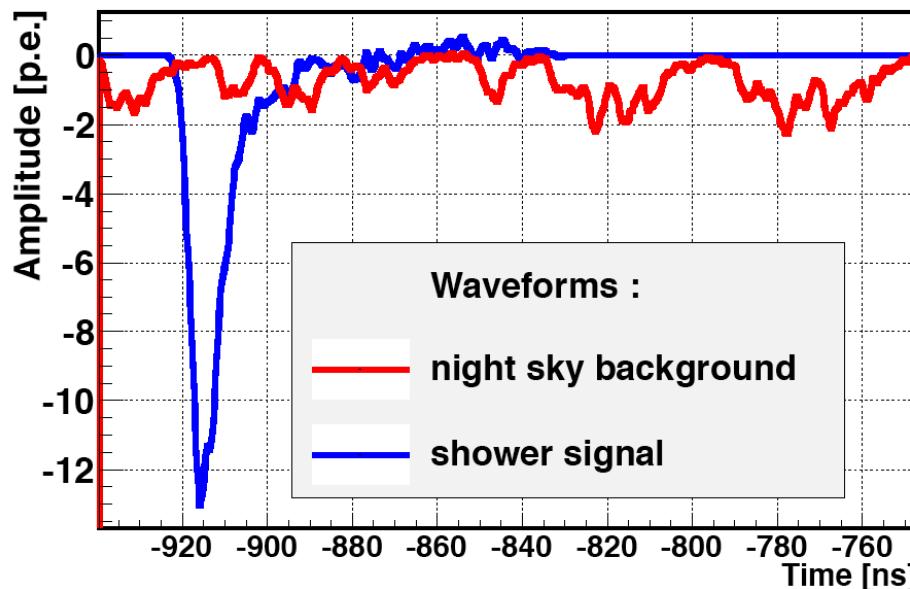


MC tools
(CORSIKA/
Simtelarray/
CARE)
Cost

- > Gain using Schwarzschild-Couder telescopes?
- > How to integrate them?
- > How to arrange them – for a given budget?

G Decerprit

Optimization of Front-End Electronics



- Analyze waveforms (shower+background) for various physical and instrumental parameters using real PMT signals
- Optimal readout time ($\sim 5\text{-}20$ ns) is smaller at lower S/B ratios
- Optimal analog bandwidth is related to trigger parameters (\rightarrow input)
- Optimize CTA performances (energy threshold, sensitivity, resolution) using readout+trigger+analysis



DESY Activities

> Hardware

> Array Control

> Sensitivity studies and
Reconstruction
methods

> Analysis of VERITAS
data
> Preparation of CTA
physics

> Computing







- > Non-affiliated members in VERITAS
- > Real-data and observing experience for postdoc and PhDs
- > Implementation of next-day analysis → fast analysis feedback for variable sources
- > MC development, validation and production
- > Lead authors in papers on observations of binaries

THE ASTROPHYSICAL JOURNAL, 732:1 (10pp), 2011 ???
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[doi:10.1088/0004-637X/732/1/1](https://doi.org/10.1088/0004-637X/732/1/1)

VERITAS Reports Increased Activity of HESS J0632+057 in Very High Energy Gamma Rays

ATel #3153; [Rene A. Ong, for the VERITAS Collaboration](#)
 on 8 Feb 2011; 22:54 UT

Distributed as an Instant Email Notice Transients
 Credential Certification: Rene Ong (rene@astro.ucla.edu)

Subjects: >GeV, TeV, VHE, Binary, Transient

Referred to by ATel #: [3161](#), [3180](#), [3209](#)

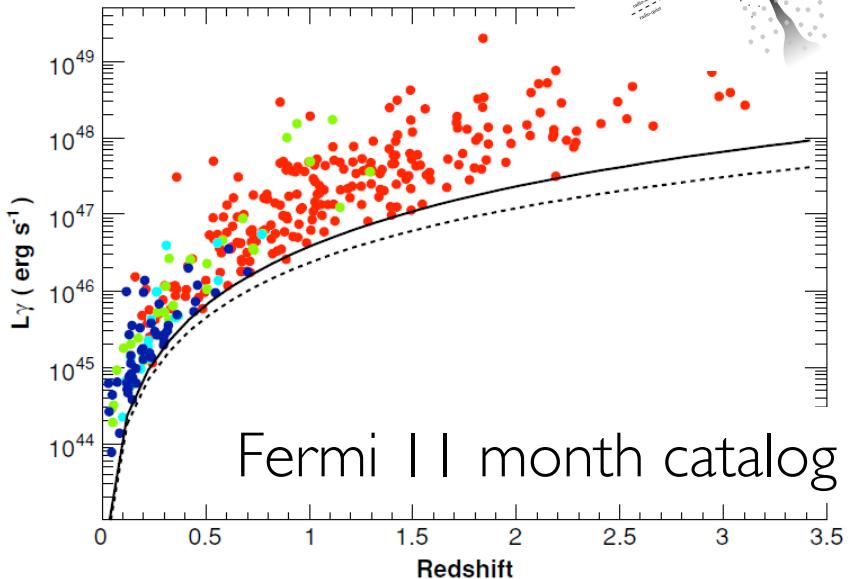
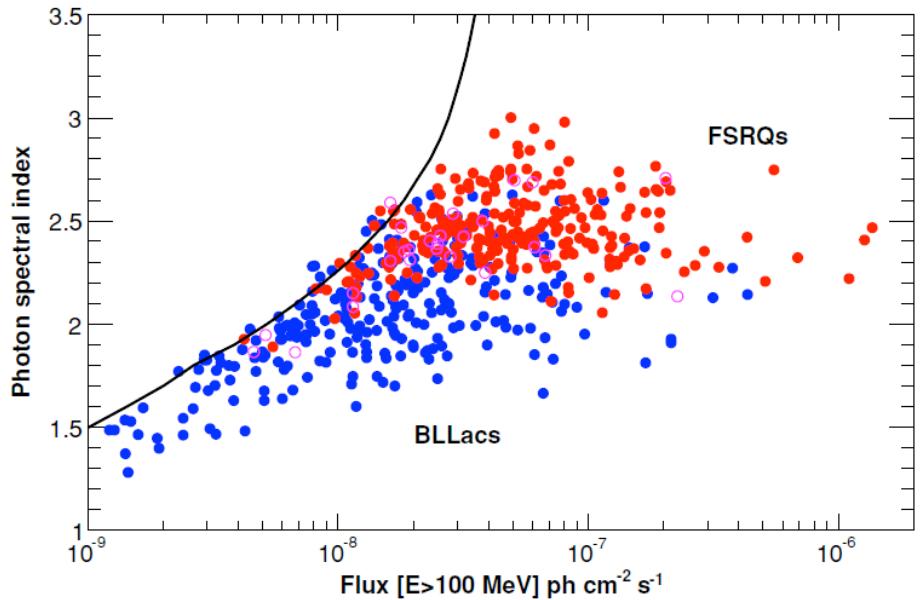
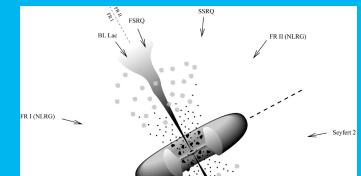
astro-ph/1103.3250

GAMMA-RAY OBSERVATIONS OF THE Be/PULSAR BINARY 1A 0535+262 DURING A GIANT X-RAY OUTBURST

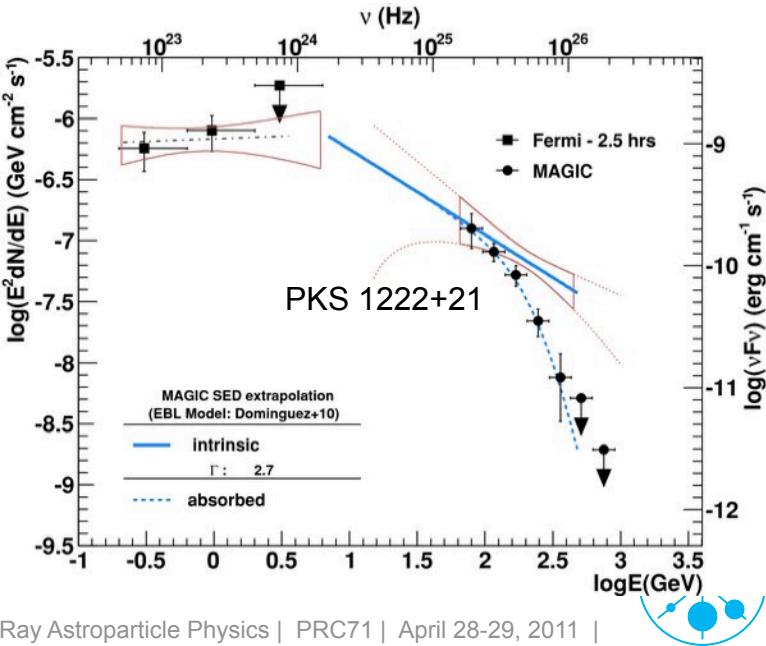
V. A. ACCIARI¹, E. ALIU², M. ARAYA³, T. ARLEN⁴, T. AUNE⁵, M. BEILICKE⁶, W. BENBOW¹, S. M. BRADBURY⁷, J. H. BUCKLEY⁶, V. BUGAEV⁶, K. BYRUM⁸, A. CANNON⁹, A. CESARINI¹⁰, L. CIUPIK¹¹, E. COLLINS-HUGHES⁹, W. CUT^{3,26}, R. DICKHERBER⁶, C. DUKE¹², A. FALCONE¹³, J. P. FINLEY³, L. FORTSON¹¹, A. FURNIS³, N. GALANTE¹, D. GALL³, S. GODAMBE¹⁴, S. GRIFFIN¹⁵, R. GUENETTE¹⁵, G. GYUK¹¹, D. HANNA¹⁵, J. HOLDER¹⁶, G. HUGHES¹⁷, C. M. HUI¹⁴, T. B. HUMENSKY¹⁸, A. IMRAN¹⁹, P. KAARET²⁰, M. KERTZMAN²¹, H. KRAWCZYNSKI⁶, F. KRENNRICH¹⁹, A. S. MADHAVAN¹⁹, G. MAIER^{17,26}, P. MAJUMDAR⁴, S. McARTHUR⁶, P. MORIARTY²², R. A. ONG¹, A. N. OTTE³, D. PANDEL²⁰, N. PARK¹⁸, J. S. PERKINS¹, M. POHL²³, H. PROKOPH¹⁷, J. QUINN⁹, K. RAGAN¹⁵, L. C. REYES¹⁸, P. T. REYNOLDS²⁴, E. ROCACHE¹, H. J. ROSE⁷, D. B. SAXON¹⁶, G. H. SEMBROSKI¹⁶, G. DEMET SENTURK²⁵, A. W. SMITH⁸, G. TEŠIĆ¹⁵, M. THEILING¹, S. THIBADEAU⁶, A. VARLOTTA^{3,26}, S. VINCENT¹⁴, M. VIVIER¹⁶, S. P. WAKELY¹⁸, J. E. WARD⁹, T. C. WEEKES¹, A. WEINSTEIN⁴, T. WEISGARBER¹⁸, S. WENG³, D. A. WILLIAMS⁸, M. WOOD⁶, AND B. ZITZER³



Flat Spectrum Radio Quasars



- Brighter and steeper spectrum blazars
- Detected up to larger distances
- Three quasars detected so far by HESS and MAGIC
- Interesting for VERITAS, HESS II & CTA



DESY Activities

> Hardware

> Array Control

> Sensitivity and
Reconstruction
methods

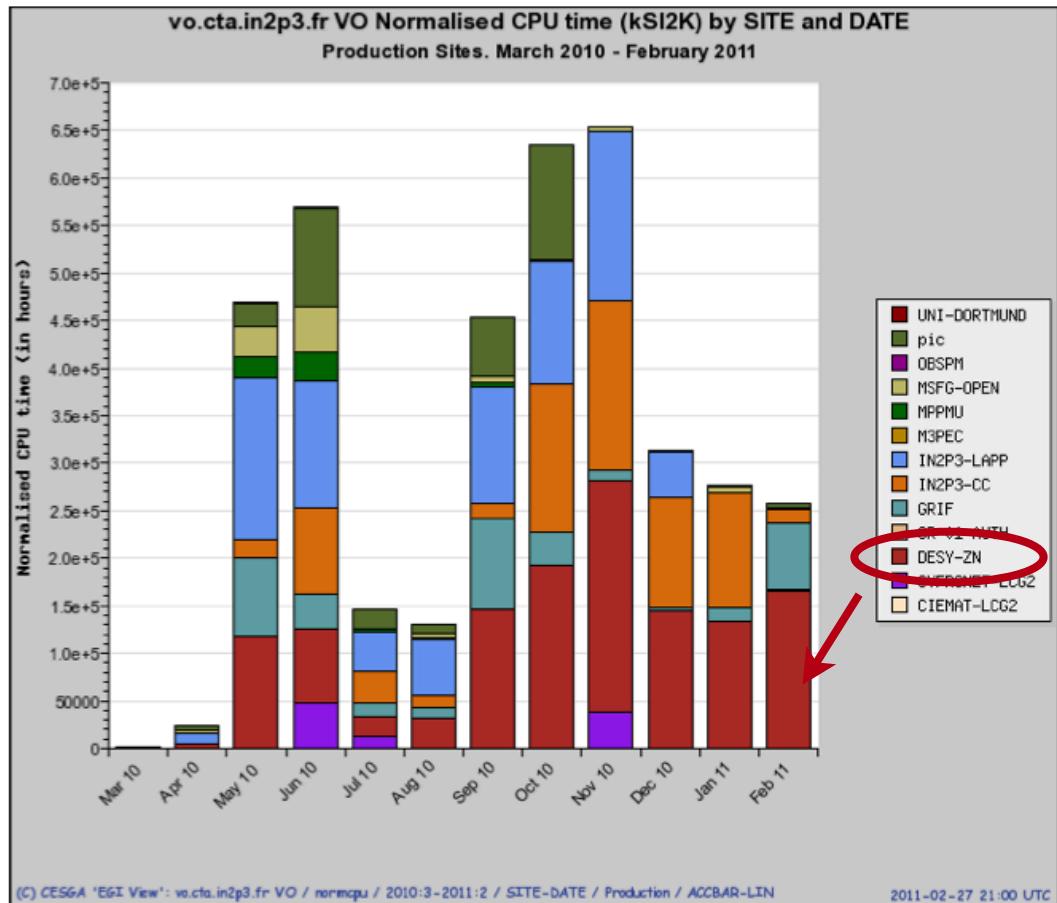
> Analysis

> Computing: massive, grid based computing element
and centre for MC production



Computing

- > Grid MC production of massive signal and background data-sets
- > Prepare CTA data processing and storage element
 - connected to Array Control Centre
 - analysis
 - storage currently: 100TB in dCache/Grid, 150 TB lustre



Summary

- CTA in Preparatory Phase working towards final design
- CTA will be the first VHE observatory open to the community
- CTA will be unique and versatile tool for astronomy and astroparticle physics
- DESY plays a key role in the preparation for CTA in key areas of design and planning
- Rich VERITAS analysis
- Prepare CTA physics using VERITAS experience

Young
Investigator
Group

