

# A Glimpse of Quantum Field Theory in Curved Spacetime

David Lai

DESY Hamburg



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## QUANTUM MECHANICS' RECAP

Before moving to quantum mechanics, let's state well-known results in classical mechanics that will serve as a basis for quantum systems.

Def., let  $\mathcal{M}$  be the phase space of a classical system with finite degrees of freedom. The dynamical evolution of  $y \in \mathcal{M}$  is defined via  $H : \mathcal{M} \rightarrow \mathbb{R}$  through:

$$(1) \quad \frac{dy^\mu}{dt} = \tilde{\omega}_\nu^{\mu\nu} \partial_\nu H \quad \mu, \nu \in \{1, \dots, 2n\}$$

where  $\tilde{\omega}_\nu$  is an antisymmetric  $2d \times 2d$  matrix such that  $\tilde{\omega}_\nu^{\mu\nu} = 1$  when  $\mu = \nu - d$ ,  $\tilde{\omega}_\nu^{\mu\nu} = 0$  when  $|\mu - \nu| \neq d$ .

Remark.  $\tilde{\omega}_\nu$  defines a non-degenerate, closed 2-form on  $\mathcal{M}^{(*)}$  via its inverse. Chosen as local frame on  $\mathcal{M}$ ,  $\{q^\mu, p_\mu\}$ , we have

(\*) symplectic

$$\Omega := dp_\mu \wedge dq_\mu$$

$$\tilde{\omega}^{ab} \tilde{\omega}_{bc} = \delta_a^b$$

Remark. Given  $h^\alpha = \tilde{\omega}^{ab} \nabla_b H$ , then any possible trajectory allowed by  $H$  is an integral curve of  $h^\alpha$ .

Def. (OBSERVABLE) Given a phase space  $\mathcal{M}$  together with a symplectic form  $\Omega$ , the set

$$\Theta := \{ f : \mathcal{M} \rightarrow \mathbb{R} \mid f \text{ is } C^\infty \}$$

enlarges to a Poisson algebra through

$$\{, \} : \Theta \times \Theta \rightarrow \Theta$$

$$\{f, g\} := \Omega^{ab} \nabla_a f \nabla_b g$$

Remark. Consider now  $q_\mu : \mathcal{M} \rightarrow \mathbb{R}$ ,  $p_\nu : \mathcal{M} \rightarrow \mathbb{R}$ . we have:

$$(1) \quad \{q_\mu, p_\nu\} = \delta_{\mu\nu}$$

$$(2) \quad \{q_\mu, q_\nu\} = \{p_\mu, p_\nu\} = 0$$

Remark. (ii) Given  $y_1, y_2 \in \mathcal{M}$ , we have:

$$(3) \quad \Omega(y_1, y_2) := p_{1\mu} q_{2\mu} - p_{2\mu} q_{1\mu}$$

Def. ("QUANTIZATION") By "quantization" of a classical system we mean a correspondence map,

$$\hat{\phantom{f}} : \Theta \rightarrow \hat{\Theta}$$

such that

$$[\hat{f}, \hat{g}] := i \hat{\{f, g\}}$$

where the latter should be understood as an algebra  $(\hat{\Theta}, [\cdot, \cdot])$  of self-adjoint operators on an Hilb.

space  $\mathcal{F}$ .

Remark. For linear dynamical systems, it's possible to associate to each coordinate  $p_\mu, q_\mu$  in  $\text{CM}$  some operators  $\hat{q}_\mu, \hat{p}_\mu$  s.t.

$$[\hat{q}_\mu, \hat{p}_\nu] = i \{q_\mu, p_\nu\} = i \delta_{\mu\nu} \overset{\wedge}{\text{Id}}_{\mathcal{F}}$$

$$[\hat{q}_\mu, \hat{q}_\nu] = [\hat{p}_\mu, \hat{p}_\nu] = 0_{\mathcal{F}}$$

Remark. For observables at most linear in the momentum we can endow the set w/ an algebra structure consisting of bounded self-adjoint operators on the Hilbert space  $\mathcal{F}$ .

Remark. For other kind of observables, factor ordering ambiguities can arise. For theories of oscillators like the one we are interested in this lecture, only  $p, q$  and  $H$  are relevant, and these are not affected by factor ordering ambiguities.

Ex. Schrödinger  $\text{QM}$  is characterized by

$$\mathcal{F} = L^2(\mathbb{R}^3), \quad \hat{q}_\mu := q_\mu \quad \text{and} \quad \hat{p}_\mu := i\partial_\mu$$

Ex. An important example that lead us to field theory is the classical harmonic oscillator ( $H_0$ ).

Given

$$H = \frac{p^2}{2} + \frac{\omega^2}{2} q^2$$

we can aim to a re-writing in terms of the classical analogue of the annihilation operator and its conjugate:

$$a := \sqrt{\frac{\omega}{2}} q + i \sqrt{\frac{1}{2\omega}} p$$

that leads upon quantization to:

$$[\hat{a}, \hat{a}^\dagger] = \text{Id}_{\mathcal{F}} = 1$$

$$\hat{H} = \omega (\hat{a}^\dagger \hat{a} + \frac{1}{2})$$

Claim: Given

$$\hat{O}_H := U_t^\dagger \hat{O} U_t$$

we have:

$$\hat{q}(t) = \sqrt{\frac{1}{2\omega}} (e^{-i\omega t} \hat{a} + e^{i\omega t} \hat{a}^\dagger)$$

We introduce now a very important concept that will return later:

Def. We define as vacuum of the theory  $|0\rangle \in \mathcal{F}$  such that

$$(1) \quad a|0\rangle = 0_F$$

Remark. The  $n$ -th excited state in the theory is given

by:

$$|n\rangle := \frac{1}{\sqrt{n!}} (\hat{a}^\dagger)^n |0\rangle$$

Trivially:

$$\hat{H}|n\rangle = (n + \frac{1}{2})\omega |n\rangle$$

# QFT IN MINKOWSKI RECAP

Def. (KLEIN GORDON FIELD IN  $M_2$ )

Given  $M_2$  as the two-dim. Minkowski spacetime

$$ds^2 = dt^2 - dx^2 = \eta_{\mu\nu} dx^\mu dx^\nu$$

we define as Klein-Gordon field  $\phi: M_2 \rightarrow \mathbb{R}$

a function on  $M_2$  whose dynamics is described by the following action:

$$S = \frac{1}{2} \int d^2x (\partial_\mu \phi \partial^\mu \phi - m^2 \phi^2)$$

Def (QUANTIZED KG FIELD). The quantization procedure of a KG field in  $M_2$  is carried out using the prescription stated before. Considering the "fundm." variables  $\phi(t, x), \pi(t, x)$  (where  $\pi$  is the conjugate of  $\phi$ ) we construct a map

$$\hat{\phantom{\phi}}: \Theta \rightarrow \hat{\Theta}$$

such that

$$\begin{aligned} [\hat{\phi}(t, x), \hat{\pi}(t, y)] &= i \{ \phi(t, x), \pi(t, y) \} \\ &= i S(x-y) \text{Id}_\phi \end{aligned}$$

Remark In terms of modes' expansion: ( $\kappa^\mu := (\omega_\kappa, x)$ )

$$\phi(t, x) = \int \frac{d^3\kappa}{\sqrt{2\omega_\kappa}} (\hat{a}_\kappa e^{-i\kappa^\mu x_\mu} + \hat{a}_\kappa^\dagger e^{i\kappa^\mu x_\mu}) \quad (*)$$

$$\text{where } [\hat{a}_k, \hat{a}_{k'}^\dagger] = \delta(k-k') \text{Id}_F$$

$$[\hat{a}_k, \hat{a}_{k'}] = [\hat{a}_{k'}^\dagger, \hat{a}_{k'}^\dagger] = 0_F$$

Remember: Following the harmonic oscillator @ page 5, we see that a vacuum can be defined as:

$$\forall k, \quad \hat{a}_k |0\rangle = 0_F$$

Comment: Pay attention on the procedure that we did above: we used a time-translational symmetry to decompose  $\phi$  (i.e.  $F$ , our Hilbert space) in positive and negative energy (i.e. the conserved charge corresponding to time translations). This makes the definition of a vacuum (and - as a consequence - the particle interpretation) self-consistent. What if the time-translational symmetry / any other symmetries are not there?

We will come back later to this, but as a first naive comment we can certainly say that the notion of vacuum/particles cannot certainly be a universal concept.

Let's move to more interesting situations.

## QFT IN FRW SPACETIMES

A first example different from QFT in Minkowski is the following

Def. (FRW METRIC)

$$ds^2 := dt^2 - a^2(t) dx^2 \quad (1)$$

Def. (CONFORMAL TIME)

Let  $t_0$  be an arbitrary constant. The conformal time  $\eta$  is defined as:

$$\eta(t) := \int_{t_0}^t dt/a(t)$$

Claim. Let  $d=2$ . Then (1) in the new coordinates is expressed as:

$$ds^2 := a^2(\eta) (d\eta^2 - dx^2)$$

Claim. Consider the action of a KG field in two-dimensions in a FRW background:

$$\begin{aligned} S_{\text{FRW}} &:= \frac{1}{2} \int d\text{Vol}_{\text{FRW}} (\partial^\mu \phi \partial_\mu \phi - m^2 \phi^2) \\ &= \frac{1}{2} \int d\eta dx a^2(\eta) \left( \partial_\eta \phi \partial_\eta \phi - \partial_x \phi \partial_x \phi - m^2 \phi^2 \right) \end{aligned}$$

Using

$$X(\eta, x) := a(\eta) \phi(\eta, x) \Rightarrow$$

(we omit functional dependencies and we denote  $\partial_f := f'$ )

$$\phi' = \partial_\eta \left( \frac{x}{a} \right) = -\frac{a'}{a^2} x + \frac{x'}{a}$$

$$(\phi')^2 = \frac{(a')^2}{a^4} x^2 + \frac{(x')^2}{a^2} - 2 \frac{a' x' x}{a^3}$$

$$a^2 (\phi')^2 = \left( \frac{a'}{a} \right)^2 x^2 + (x')^2 - 2 \frac{a'}{a} x' x$$

BUT...

$$-2 \frac{a'}{a} x' x = \frac{a'}{a} ((x')^2)' = -\left( \frac{a'}{a} (x')^2 \right)' + \left( \frac{a'}{a} \right)' x^2$$

$$\text{i.e. } a^2 (\phi')^2 = (x')^2 + \left( \frac{a''}{a} \right) x^2 + \text{total derivative}$$

We obtain:

$$S_{\text{FRW}} := \frac{1}{2} \int d\eta dx \left( \partial_\eta x \partial_\eta x - \partial_x x \partial_x x - M^2(\eta) x^2 \right)$$

where

$$M(\eta) := m^2 a^2(\eta) - \frac{\partial_\eta^2 a(\eta)}{a(\eta)} \quad \text{"effective mass.."}.$$

Comment. We reduced the dynamics of a KG field in a FRW metric to a problem of a KG field w/ a  $\eta$ -dependent mass in Minkowski. Thus we can

admit a mode expansion as ( $\partial K = \frac{\partial K}{\partial \eta}$ )

$$\hat{\chi}(\eta, x) := \int \frac{d\kappa}{\sqrt{2}} \left( e^{i\kappa \cdot x} V_K^*(\eta) \hat{a}_K + e^{-i\kappa \cdot x} V_K(\eta) \hat{a}_K^\dagger \right)$$

where now  $V_K(\eta)$  are mode functions satisfying the analogue of the ODE describing an harmonic oscillation w/ time-dependent frequency:

Def. (MODE FUNCTION OF A FEW FIELD)

We def. as mode function of a KG field in the few metric in  $(\eta, x)$  coordinates a solution of:

$$V_K'' + \omega_K^2(\eta) V_K = 0, \quad \omega_K^2 := \kappa^2 + M^2(\eta)$$

Comment: Notice that the choice of  $V_K(\eta)$  will lead to different definitions of  $\hat{a}_K$ ,  $\hat{a}_K^\dagger$ , leading to different vacua / particles!

Remark. Let choose two different sets of mode functions,  $\{U(K), V(K)\}$ . It holds that (with the suitable normalisations):

$$(1) \quad V_K^*(\eta) := \alpha_K U_K(\eta) + \beta_K V_K(\eta)$$

$$(2) \quad |\alpha_K|^2 - |\beta_K|^2 = 1$$

Claim: Since

$$\begin{aligned}\hat{x}(\eta, x) &:= \int \frac{d\kappa}{\sqrt{2}} \left( e^{i\kappa \cdot x} v_\kappa^+(\eta) \hat{a}_\kappa \right. \\ &\quad \left. + e^{-i\kappa \cdot x} v_\kappa(\eta) \hat{a}_\kappa^\dagger \right) \\ &= \int \frac{d\kappa}{\sqrt{2}} \left( e^{i\kappa \cdot x} v_\kappa^+(\eta) \hat{b}_\kappa \right. \\ &\quad \left. + e^{-i\kappa \cdot x} v_\kappa(\eta) \hat{b}_\kappa^\dagger \right)\end{aligned}$$

$$\Rightarrow (1) \quad \begin{cases} \hat{b}_\kappa = \alpha_\kappa \hat{a}_\kappa + \beta_\kappa^* \hat{a}_{-\kappa}^\dagger & \text{(notice the } -\kappa, \\ \hat{b}_\kappa^\dagger = \alpha_\kappa^* \hat{a}_\kappa^\dagger + \beta_\kappa \hat{a}_{-\kappa} & \text{we want to} \\ & \text{factor a } e^{ix \cdot x}) \end{cases}$$

Def. (Bogoliubov TRANSFORMATIONS)

Given two solutions of:

$$v_\kappa'' + \omega_\kappa^2(\eta) v_\kappa = 0, \quad \omega_\kappa^2 = \kappa^2 + M^2(\eta)$$

namely  $v_\kappa(\eta)$  corresponding to the operators  $\hat{a}_\kappa, \hat{a}_\kappa^\dagger$  and  $v_\kappa^*(\eta)$  corresponding to  $\hat{b}_\kappa, \hat{b}_\kappa^\dagger$ , these two related by

$$(3) \quad v_\kappa^*(\eta) := \alpha_\kappa^* v_\kappa(\eta) + \beta_\kappa v_\kappa^*(\eta)$$

$$(4) \quad |\alpha_\kappa|^2 - |\beta_\kappa|^2 = 1$$

we define as Bogoliubov transformations the eqs. in (1).

Def. (a-VACUUM / b-VACUUM). It follows from (1) that we can associate a vacuum for each Bogoliubov-transformed set of modes:

i.e.

$$|0\rangle_a \quad |a_k |0\rangle_a = 0 \quad \forall k$$

or

$$|0\rangle_b \quad |b_k |0\rangle_b = 0 \quad \forall k$$

Claim: let  $n_k^{(b)}$  be the density of  $b$  particles ( $k^{\text{th}}$  mode) seen in the  $a$ -vacuum. we have:

$$n_k^{(b)} = |\beta_k|^2$$

Proof:

$$\underbrace{N_b^{(k)}}$$

$$a \langle 0 | b_k^\dagger b_k | 0 \rangle_a =$$

$$a \langle 0 | (\alpha_k^\dagger \alpha_k + \beta_k^\dagger \alpha_k) (\alpha_k \alpha_k + \beta_k^\dagger \alpha_k^\dagger) | 0 \rangle_a$$

$$= a \langle 0 | (\beta_k^\dagger \alpha_k) (\beta_k^\dagger \alpha_k^\dagger) | 0 \rangle_a = |\beta_k|^2 \delta(k-k)$$

$$= |\beta_k|^2 \delta(0)$$

Considering the quantization in a 1d box, we would replace  $\delta(0)$  by  $L$ . Computing the density would have given:

$$n_k^{(b)} = |\beta_k|^2$$

Remark If  $\int |\beta_k|^2 dk$  is divergent, it means that the back reaction of this particle production cannot be neglected!

# FULLING-DAVIES-UNRUH EFFECT

[ FULLING, 1973 ]

[ DAVIES, 1975 ]

[ UNRUH, 1976 ]

Comment. By the equivalence principle:

GRAVITY  $\sim$  NON-INERTIAL FRAME

we should be able to define the a/b vacuum effect of FRW metric also in  $M_2$ , modulo the fact that we adopt a non-inertial frame. The simplest frame that we can adopt is a constant proper acceleration frame. Considering  $M_2$  in the usual patch  $(t, x)$ , we define as usual the proper time as

$$i. \quad | \quad ds^2 := d\tau^2$$

and:

$$u^\mu := \frac{dx^\mu}{d\tau}, \quad a^\mu := \frac{d^2x^\mu}{d\tau^2} \quad | \quad a^\mu a_\mu = -|a|^2 \\ = -a^2$$

that leads to:

$$(1) \quad x(\tau) = x_0 - \frac{1}{a} + \frac{1}{a} \cosh(a\tau)$$

$$t(\tau) = t_0 + \frac{1}{a} \sinh(a\tau)$$

and choosing  $x_0 = \frac{1}{a}$ ,  $t_0 = 0$  we obtain a simpler form of (1).

Obs. We can obtain a frame where the accel.  
observer sit at  $\xi = 0$ , with proper time  $\tau$ :

$$t(\tau, \xi) := \frac{1}{\alpha} e^{\alpha \xi} \sinh(\alpha \tau) \quad (1)$$

$$x(\tau, \xi) := \frac{1}{\alpha} e^{\alpha \xi} \cosh(\alpha \tau)$$

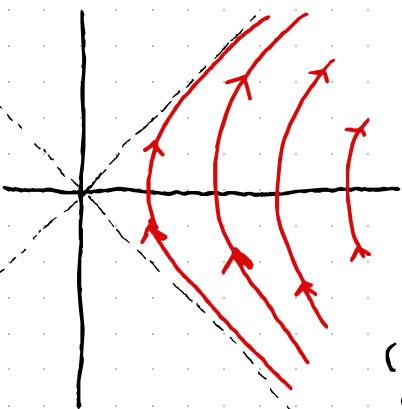
wl metric element:

$$(2) \quad ds^2 := e^{2\alpha \xi} (d\tau^2 - d\xi^2)$$

↑  
conformally  
flat!

Def. (RINDLER SPACETIME) We define as Rindler Spacetime  
the metric obtained by (2) defining a new coord.

$$\hat{\xi} \mid \xi := \frac{1}{\alpha} \ln(1 + \alpha \hat{\xi})$$



(1) means that  
we cover the  
right Rindler wedge.

Def. (RINDLER MODE EXPANSION) We consider a massless  $\text{K}\bar{\phi}$  field.

$$\hat{\phi}(\tau, \xi) := \int dx \frac{1}{\sqrt{2\lambda\omega}} (e^{-ik_1\tau + ik_2\xi} b_{k_1} + e^{ik_1\tau - ik_2\xi} b_{k_1}^\dagger)$$

Def. (RINDLER VACUUM)

$$|0\rangle_R \mid b_k |0\rangle_R = 0 \quad \forall k$$

Question: how  $|0\rangle_R$  is seen by  $|0\rangle_{M_2}$ ?

Def. (LIGHTCONE COORDINATES)

We define the  $M_2$  light-cone coordinates as:

$$x^- := t - x$$

$$x^+ := t + x$$

that reduce the metric to:

$$ds^2 := dx^+ dx^-$$

We define the (conformal) Rindler light-cone coordinates as:

$$u := \tau - \xi$$

$$v := \tau + \xi$$

such that:

$$ds^2 := dx^+ dx^- = e^{a(v-u)} du dv$$

Notice:

$$x^+ = -\frac{1}{a} e^{-av} \quad ; \quad x^- = \frac{1}{a} e^{av} v$$

Obs. The field equations become:

M2 hypothesis:  $\partial_x^+ \partial_x^- \phi(x^+, x^-) = 0$   
 $\Rightarrow \phi(x^+, x^-) = \phi^+(x^+) + \phi^-(x^-)$

Rindler hypothesis:  $\partial_u \partial_v \phi(u, v) = 0$   
 $\Rightarrow \phi(u, v) = \phi^\uparrow(u) + \phi^\downarrow(v)$

Claim. Considering  $\omega := |k\omega|$ , we accomplish to separate the mode expansion in  $u, v$  components /  $x^+, x^-$  comp. components. It focus on  $x^+, u$  parts, respectively. This is consistent with the above observation.

$$\begin{aligned}\phi^+(x^+) &:= \int_0^{+\infty} d\omega \frac{1}{\sqrt{2\omega}} \left( e^{-i\omega x^+} a_\omega + e^{i\omega x^+} a_\omega^+ \right) \\ \phi^\uparrow(u) &:= \int_0^{+\infty} d\omega \frac{1}{\sqrt{2\omega}} \left( e^{-i\omega u} b_\omega + e^{i\omega u} b_\omega^+ \right)\end{aligned}$$

Obs. Since

$$x^+ \rightarrow u \text{ and } x^- \rightarrow v$$

we obtain:

$$\begin{aligned}&\int_0^{+\infty} d\omega \frac{1}{\sqrt{2\omega}} \left( e^{-i\omega x^+} a_\omega + e^{i\omega x^+} a_\omega^+ \right) \\ &= \int_0^{+\infty} d\omega \frac{1}{\sqrt{2\omega}} \left( e^{-i\omega u} b_\omega + e^{i\omega u} b_\omega^+ \right) \quad (\circ)\end{aligned}$$

We have now to take the Fourier transform on both sides.

$$\int_{-\infty}^{+\infty} d\omega e^{i\omega u} \phi^+(x^+) = \int_0^{+\infty} d\omega / \sqrt{2\omega} (F(\omega, \omega) a_\omega + F(-\omega, \omega) a_\omega^+)$$

where

$$\begin{aligned} F(\omega, \omega) &:= \int_{-\infty}^{+\infty} d\omega e^{i\omega u - i\omega x^+} \\ &= \int_{-\infty}^{+\infty} d\omega \exp\left(i\omega u + i\frac{\omega}{\alpha} e^{-\alpha u}\right) \end{aligned}$$

$$\int_{-\infty}^{+\infty} d\omega e^{i\omega u} \phi^+(u) = \frac{1}{\sqrt{2|\alpha|}} \cdot \begin{cases} b_\omega & \text{if } \alpha > 0 \\ b_\omega^+ & \text{if } \alpha < 0 \end{cases}$$

Comparing the above eqs. for  $\omega > 0$ :

$$b_\omega = \int_0^{+\infty} d\omega \alpha(\omega, \omega) a_\omega + \beta(\omega, \omega) a_\omega^+$$

where  $\alpha(\omega, \omega)$ ,  $\beta(\omega, \omega)$  are functions of the frequencies\*. It can be shown that:

$$\begin{aligned} \int d\omega (\alpha(\omega, \omega) \alpha^*(\omega, \omega') - \beta(\omega, \omega) \beta^*(\omega, \omega')) \\ = \delta(\omega - \omega') \end{aligned}$$

\* See the next page

Obs. For completeness:

$$\alpha(\omega, \omega_0) := \sqrt{\frac{\omega_0}{\omega}} F(\omega, \omega_0) = \sqrt{\frac{\omega_0}{\omega}} \int_{-\infty}^{+\infty} du \exp\left(i\omega_0 u + i\frac{\omega}{\omega_0} e^{-\alpha u}\right)$$

$$\beta(\omega, \omega_0) = \sqrt{\frac{\omega_0}{\omega}} F(-\omega, \omega_0) = \sqrt{\frac{\omega_0}{\omega}} \int_{-\infty}^{+\infty} du \exp\left(i\omega_0 u - i\frac{\omega}{\omega_0} e^{-\alpha u}\right)$$

Property It can be shown

$$\begin{aligned} |\beta(\omega, \omega_0)|^2 &= \frac{\sqrt{2}}{\omega} |F(-\omega, \omega_0)|^2 \\ &= e^{-2\pi\omega/\omega_0} |\alpha(\omega, \omega_0)|^2 \end{aligned}$$

This can be proven by:

$$F(\omega, \omega_0) := \frac{1}{2\pi\omega_0} \exp\left(i\frac{\omega_0}{\omega} \ln\left(\frac{\omega}{\omega_0}\right) + \frac{\pi i\omega_0}{2\omega}\right)$$

•  $\Gamma\left(-i\frac{\omega_0}{\omega}\right) \quad \omega > 0 \quad \omega_0 > 0$

Cl<sub>21m</sub> (UNRUNN EFFECT)

$$\begin{aligned} \langle N(\omega_0) \rangle &= \langle 0 | b_{\omega_0}^\dagger b_{\omega_0} | 0 \rangle_{M_2} = \int d\omega d\omega' \beta^*(\omega_0, \omega') \beta(\omega_0, \omega) \\ &\quad \cdot \langle \hat{a}_{\omega}, \hat{a}_{\omega'}^\dagger \rangle \sim \delta(\omega - \omega') \\ &= \int d\omega |\beta(\omega_0, \omega)|^2 \sim \frac{V}{e^{\frac{\omega_0}{kT}} - 1} \end{aligned}$$

$$\omega/T = \frac{\alpha}{2\pi}$$

Proof.

$$\int (\alpha(\omega, \omega) \alpha^*(\omega, \omega') - \beta(\omega, \omega) \beta^*(\omega, \omega')) d\omega = \delta(\omega - \omega')$$

Considering  $\omega \sim \omega'$  and quantization in a box, we have:

$$\int d\omega (|\alpha(\omega, \omega)|^2 - |\beta(\omega, \omega)|^2) \sim V$$

But we invoke the property stated below:

$$\int d\omega (e^{2\pi/a\omega} - 1) |\beta(\omega, \omega)|^2 \sim V$$

$$\langle N_{\omega} \rangle \sim \frac{V}{e^{2\pi/a\omega} - 1}$$

i.e. an accelerated observer sees the Minkowski vacuum as a thermal bath w/ a Bose-Einstein type of radiation w/  $T = a/2\pi$ .