

Effects of PQ symmetry breaking on the axion cosmological production through misalignment

Upcoming work with Luca di Luzio

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Background



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The QCD axion potential and VEV

- > Axion T-dependence: QCD generates a potential of the form (Lattice fit: Borsanyi et al. 2016)

$$V_{\text{QCD}} \approx -m_a^2(T) f_a^2 \cos\left(\frac{a}{f_a}\right) \quad \text{where} \quad m_a^2(T) \approx m_a^2 \max\left[\left(\frac{T}{T_{\text{QCD}}}\right)^{-2b}, 1\right]$$

$$\text{where } b \approx 3.92 \quad \text{and} \quad T_{\text{QCD}} \approx 150 \text{ MeV}$$

- > PQV sensitivity: QCD preserves PQ - ensured by Vafa-Witten
Other contributions in general violate it:

$$V(\theta) \approx \underbrace{\Lambda_{\cancel{PQ}}^4 \theta}_{\sim \text{anything else}} + \frac{1}{2} \underbrace{\Lambda_{\text{PQ}}^4 \theta_{\text{eff}}^2}_{\supset \text{QCD}} + \mathcal{O}(\theta^3)$$

The VEV must be small today:

$$\theta_{\text{eff}} \simeq -\frac{\Lambda_{\cancel{PQ}}^4}{\Lambda_{\text{PQ}}^4} \lesssim 10^{-10} \quad \text{from nEDM}$$

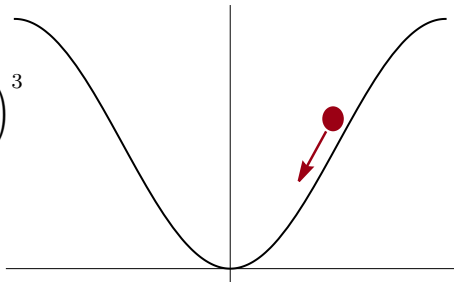
Dark matter from misalignment

- > Axion initially frozen at a random $\theta_{\text{ini}} \in -\pi, \pi$
- > Field starts oscillating when $m_a(T) > 3H(T)$
- > Generates a DM relic density:

$$\rho_{a, \text{today}} \approx \frac{1}{2} \underbrace{m_a^2 f_a^2}_{\approx (0.75 \text{ MeV})^4} \theta_{\text{ini}}^2 \frac{m_a(T_{\text{osc}})}{m_a} \frac{g_{*s}(T_0)}{g_{*s}(T_{\text{osc}})} \left(\frac{T_0}{T_{\text{osc}}} \right)^3$$

- > Matches the observed DM relic only for a unique oscillation temperature:

$$T_{\text{osc}} \approx 900 \text{ MeV} \left(\frac{\rho_{a, \text{today}}}{\rho_{\text{DM, today}}} \right)^{-1/7} \theta_{\text{ini}}^{2/7}.$$



PQ violation at higher temperatures?

Large hierarchy at T_{osc} :

$$\frac{V_{\text{QCD}}(900 \text{ MeV})}{V_{\text{QCD}}} \sim 8 \times 10^{-7}$$

- > PQ violation more competitive at higher temp.
- > However, constant temperature PQV can't dominate at $T_{\text{osc}} \approx 900 \text{ MeV}$
- > Higher $m_a \rightarrow$ higher $T_{\text{osc}} \rightarrow$ possibility of PQV domination:

$$V_{\cancel{PQ}} \gtrsim V_{\text{QCD}}(T_{\text{osc}}) \quad \text{possible for} \quad m_a \gtrsim 5 \times 10^{-3} \text{ eV}$$

Some hope of interesting phenomenology?

Constant temperature PQ violation



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Constant temperature PQ violation

Possible origins:

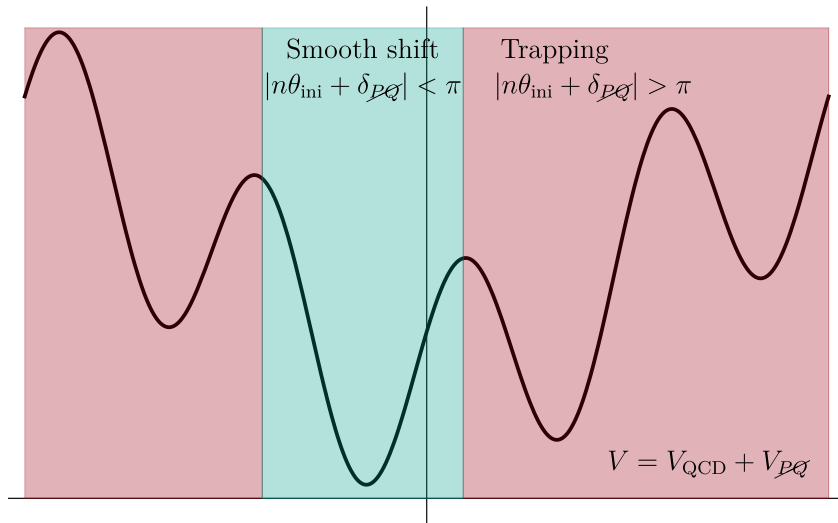
- > Planck suppressed operators
- > Any BSM physics far above $T \sim 1$ GeV

General form:

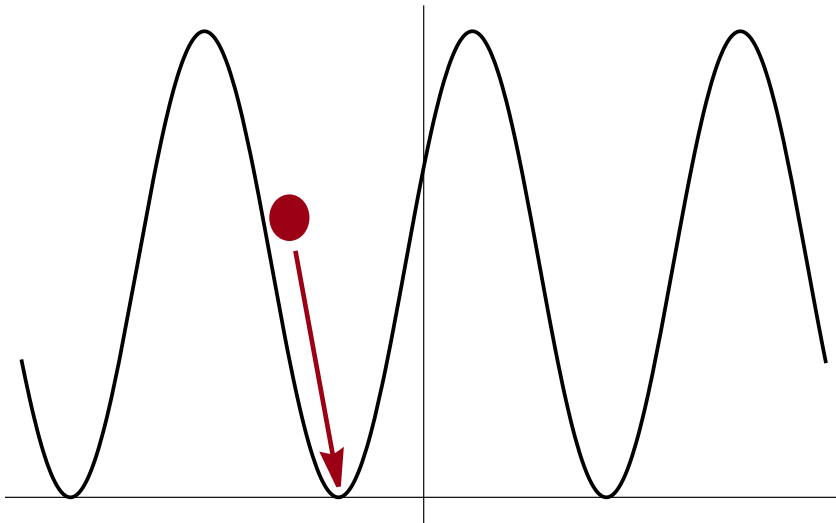
$$V_{PQ} = -\Lambda_{PQ}^4 \cos(n\theta + \delta_{PQ})$$

- > Already studied in literature: (Jeong, Matsukawa, Nakagawa, Takahashi, 2022)
- > Two distinct regimes

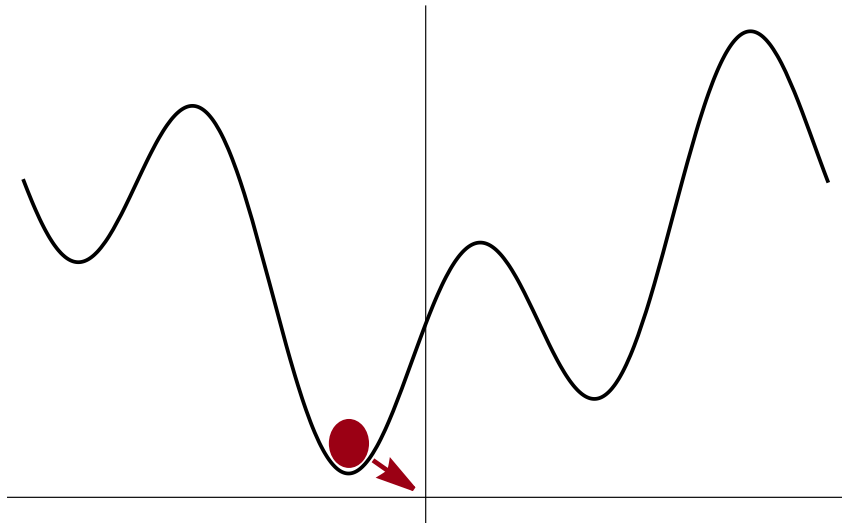
The two regimes identified by Jeong et al.



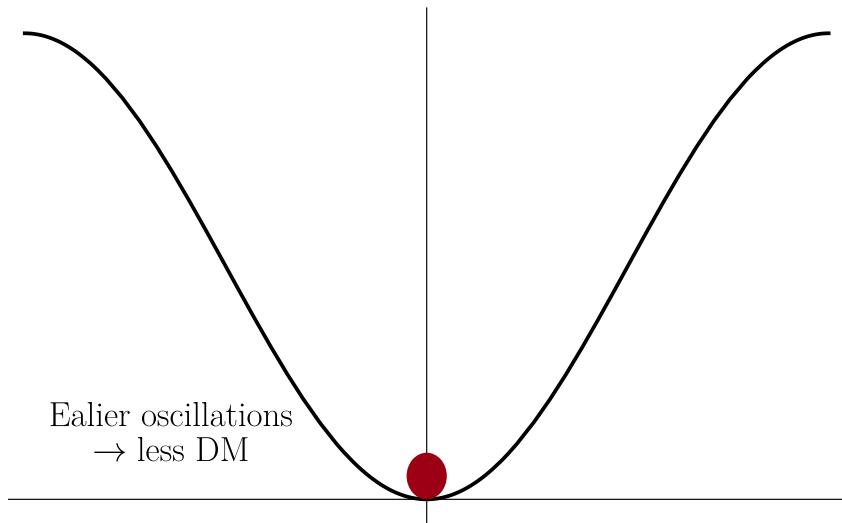
Smooth shift regime



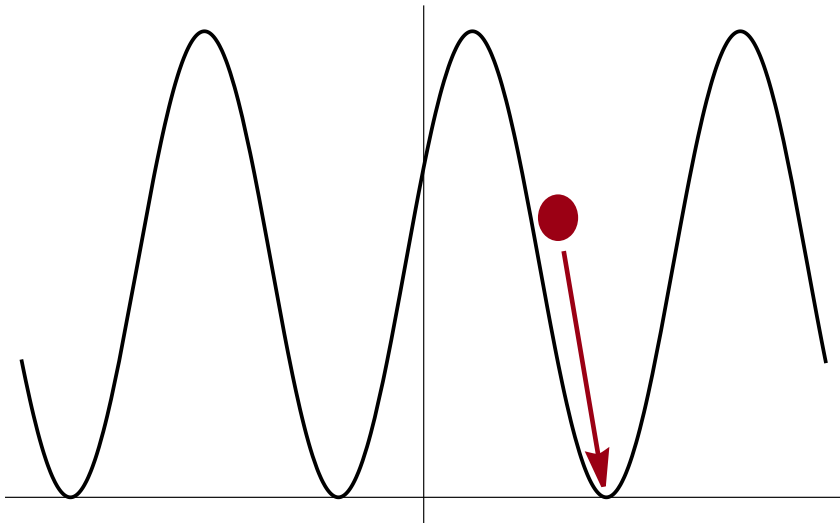
Smooth shift regime



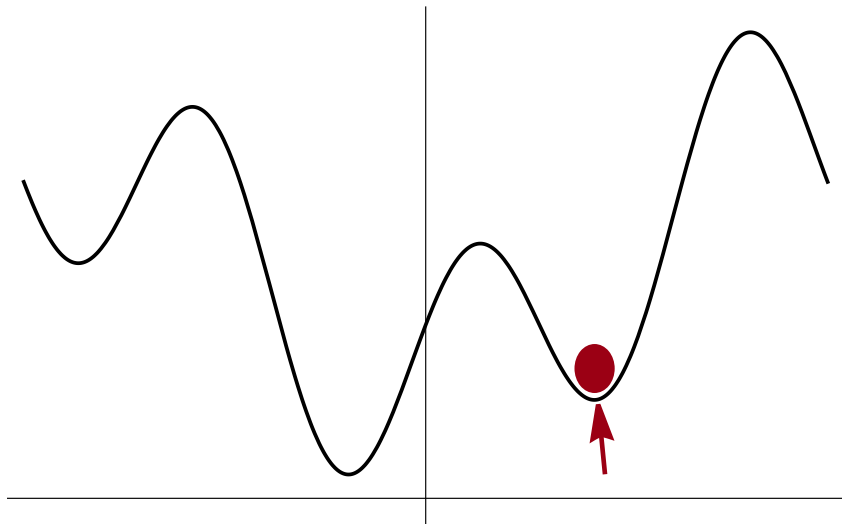
Smooth shift regime



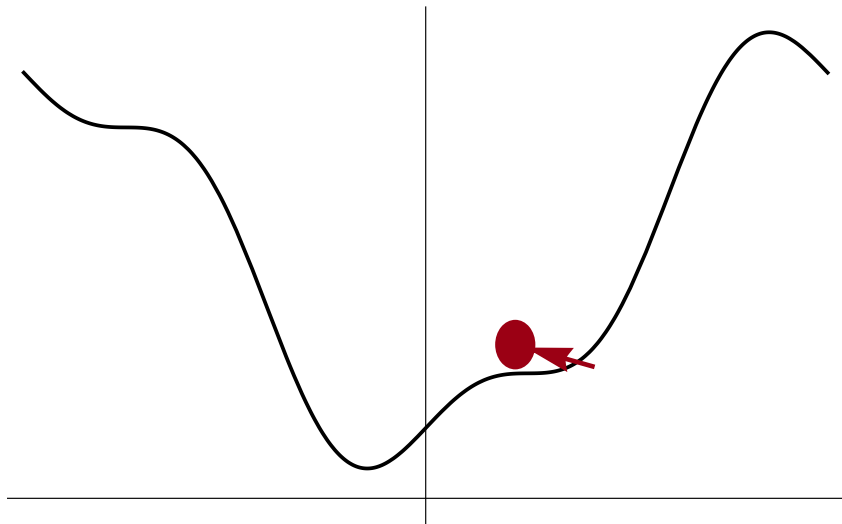
Trapping regime



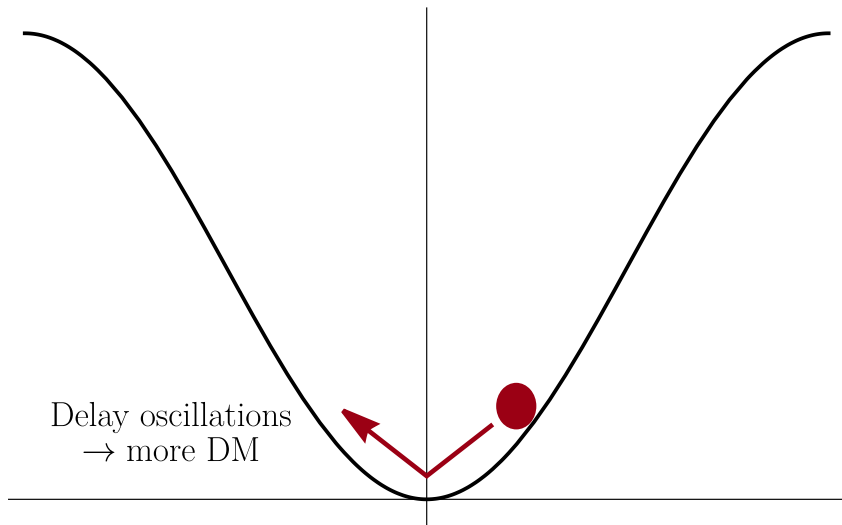
Trapping regime



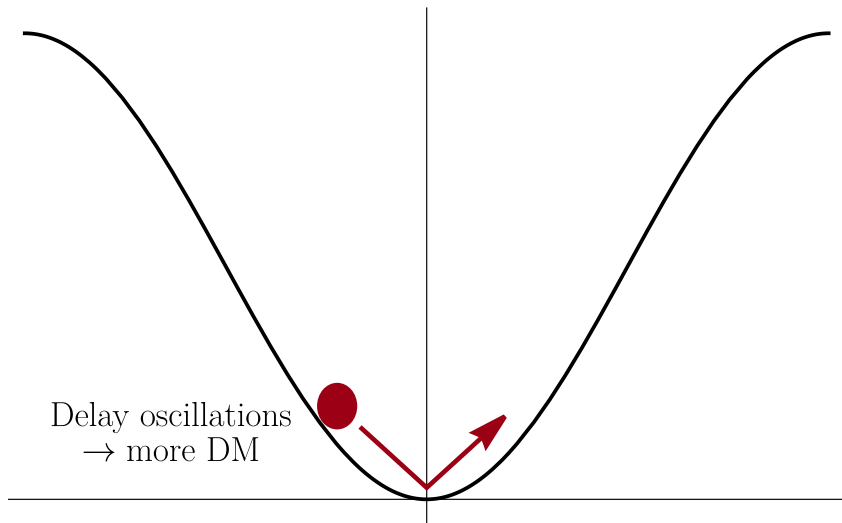
Trapping regime



Trapping regime



Trapping regime



Trapped misalignment - analytic solution

- > General arguments: $m_a > 5 \times 10^{-3} \text{ eV}$
- > DM under-production $\rightarrow$ Trapped misalignment more interesting
- > Trapped misalignment release condition:

$$\frac{\partial V}{\partial \theta} = 0 \quad \text{and} \quad \frac{\partial^2 V}{\partial \theta^2} = 0,$$

Simplifies to

$$\tan \theta_{\text{trap}} = \frac{1}{n} \tan(n\theta_{\text{trap}} + \delta_{PQ}) \quad \rightarrow \quad (n-1) \mathcal{O}(1) \text{ release angles}$$

$$T_{\text{trap}} \approx 1.3 \text{ GeV} \left(\frac{n\Lambda_{PQ}}{10^{-3} \text{ GeV}} \right)^{-0.13}.$$

- > Freedom to choose $T_{\text{trap}} \rightarrow$ choose DM relic!

Constraints from nEDM

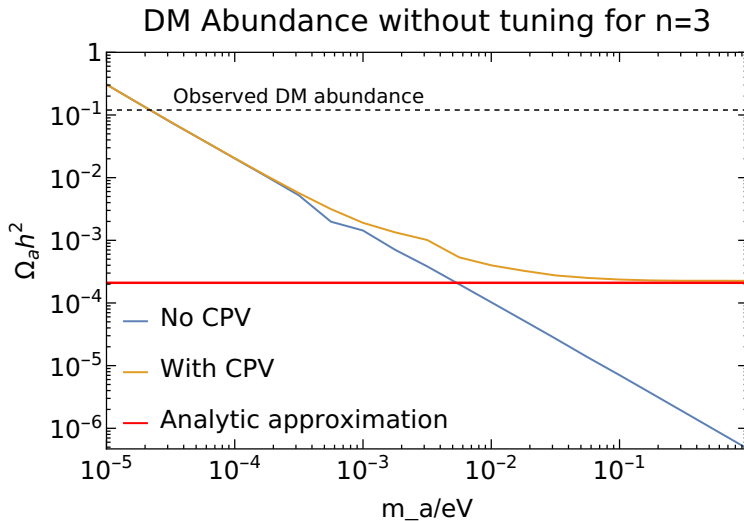
VEV constrained by nEDM:

$$|\theta_{\text{eff}}| \approx \frac{n\Lambda_{PQ}^4}{m_a^2 f_a^2} \sin \delta_{PQ}, \quad \text{must satisfy} \quad |\theta_{\text{eff}}| < 10^{-10}$$

Saturating the nEDM bounds yields best-case DM relic:

$$\frac{\rho_{a, \text{today}}}{\rho_{\text{DM, today}}} \lesssim 5 \times 10^{-4} \frac{\theta_{\text{ini}}^2}{\sin^{0.88}(\delta_{PQ})}, \quad \text{for} \quad m_a \gg 10^{-3}$$

Trapped misalignment - numeric solution



Trapped misalignment - tuned regime

- > Results in the tuned regime from Jeong, Matsukawa. Nakagawa, Takahashi, 2022:

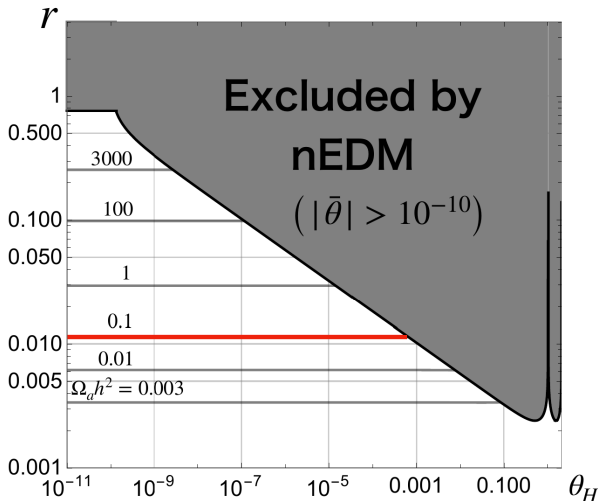
- > Notation:

$$\theta_H = \delta_{PQ}$$

$$r = \frac{\Lambda_{PQ}^4}{m_a^2 f_a^2}$$

- > 100% DM requires tuning:

$$\delta_{PQ} < 10^{-3}$$



Temperature-dependent PQ violation



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Temperature-dependent PQ violation

To improve, we now go to temperature-dependent potentials:

$$V_{\text{th}} = -\Lambda_{PQ}^4(T) \times \cos(n\theta + \delta_{PQ})$$

Y. Zhang (2305.15495) proposed to generate such potentials from

$$\mathcal{L} \supset \left(\frac{\phi}{\Lambda_{PQ}} \right)^n \times \mathcal{O}_{\text{SM}}$$

- > Proposed as a solution to DW problem
- > Zhang considered only $n = 1$
- > We generalize to $n > 1$, update constraints, and evaluate DM

Which SM fields can we couple to?

- > Relevant temperature: $T \gtrsim \text{GeV}$: After EWSB, before QCD phase transition
- > Only fields with $m \lesssim \text{GeV}$
- > Candidate fermion fields: e, μ, u, d, s :

$$V_{\text{th},f} \approx -\frac{n_{\text{col}}}{6} \left(\frac{\frac{1}{\sqrt{2}} f_a}{\Lambda_{PQ}} \right)^n m_f^2 T^2 \cos(n\theta + \delta_{PQ}) \quad \text{for } T < v_{\text{EW}}$$

- > Gauge boson candidates: Gluons and photons:

$$V_{\text{th},GG} \approx \text{loop factor} \times 2\pi\alpha_s^2 \left(\frac{\frac{1}{\sqrt{2}} f_a}{\Lambda_{PQ}} \right)^n T^4 \cos(n\theta + \delta_{PQ})$$

What about photons? See our paper!

Misalignment impact of T-dependent PQV

Previous solutions can be straight-forwardly generalized:

$$\Lambda_{PQ}^4 \rightarrow \Lambda_{PQ}^4(T) = \lambda_{PQ}^4 \left(\frac{\frac{1}{\sqrt{2}} f_a}{\Lambda_{PQ}} \right)^n T^q$$

General solution:

$$T_{\text{trap}} \approx \left(2^{\frac{n}{4b}} n^{-\frac{1}{2b}} \lambda_{PQ}^{-2/b} m_a^{1/b} f_a^{-\frac{n-2}{2b}} T_{\text{QCD}} \Lambda_{PQ}^{\frac{n}{2b}} \right)^{\frac{2b}{2b+q}}$$

Fix to target of $T_{\text{trap}} \approx 900 \text{ MeV} \rightarrow \text{DM solution for } \Lambda_{PQ}$

Understanding the evolution

The evolution can be understood with effective masses:

$$m_{a,\text{th}}^2 \equiv \frac{1}{a} \frac{\partial V_{\text{th}}}{\partial a} \Big|_{\text{min}}$$

Allows for easier interpretation of evolution:

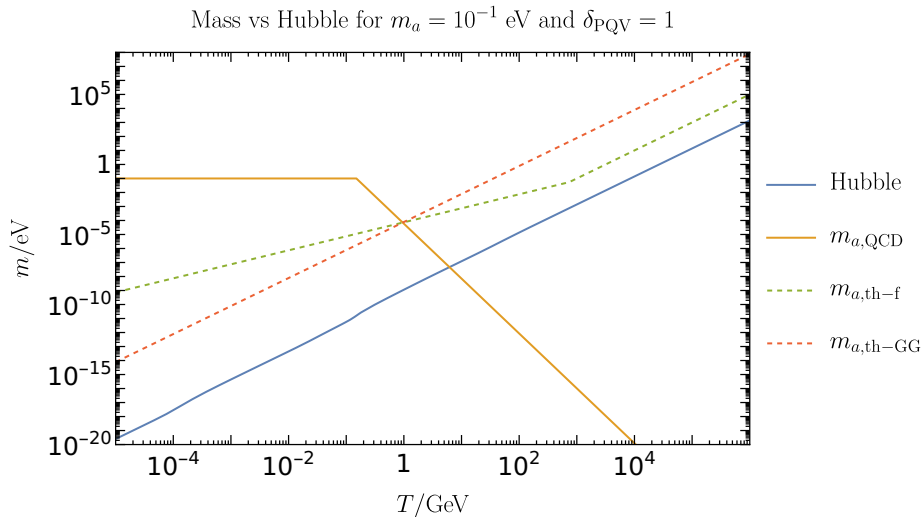
> Release condition:

$$m_a(T_{\text{trap}}) \approx m_{a,\text{th}}$$

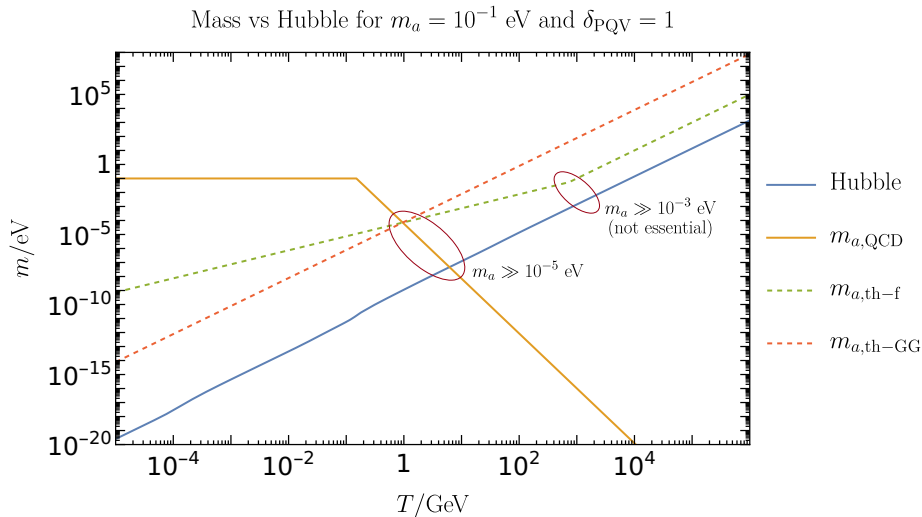
> Hubble domination:

$$m_{a,\text{th}} > 3H$$

Understanding the evolution



Understanding the evolution



Constraints on Fermion Yukawas

The scenario to a number of constraints, investigated by Zhang (2209.09429).

> Coleman-weinberg driven VEV:

$$\theta_{\text{eff}} \approx \frac{nn_{\text{col}}}{4\pi^2} \frac{m_f^4}{m_a^2 f_a^2} \left(\frac{f_a}{\sqrt{2}\Lambda_{\cancel{PQ}}} \right)^n \left[\ln \left(\frac{m_f^2}{\mu^2} \right) - 1 \right] \sin \delta_{\cancel{PQ}} \ll 10^{-10}$$

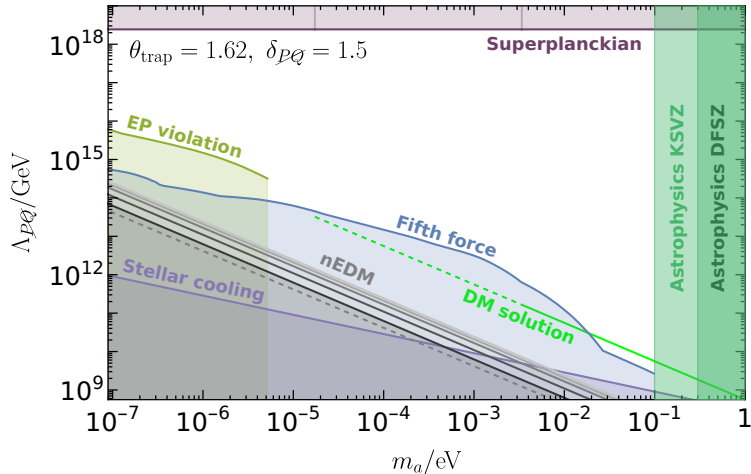
> Long range interactions:

$$\mathcal{L} \supset g_{a\bar{f}f} a \bar{f} f \quad \text{where} \quad g_{a\bar{f}f} = \frac{m_f}{f_a} \left(\frac{f_a}{\sqrt{2}\Lambda_{\cancel{PQ}}} \right)^n \sin \delta_{\cancel{PQ}}$$
$$\rightarrow V_{\text{Yukawa}} = -g_{a\bar{f}f}^2 \frac{e^{-m_a r}}{4\pi r}$$

> Stellar cooling bound on $g_{a\bar{f}f}$

Example solution: Testable with fifth force

Electron: Value of $\Lambda_{p\varphi}$ which yields observed DM density for $n=2$



Constraints on gluon interactions

- > For GG , The VEV $\langle GG \rangle$ dominates generates a zero-temperature PQV potential:

$$V_{\langle GG \rangle} = -2\alpha_s \langle GG \rangle \left(\frac{f_a}{\sqrt{2}\Lambda_{PQ}} \right)^n \cos(n\theta + \delta_{PQ}) \quad \text{where} \quad \left\langle \frac{\alpha_s}{\pi} GG \right\rangle \approx (330 \text{ MeV})^4$$

- > As before, $V_{\langle GG \rangle} \rightarrow \theta_{\text{eff}} \rightarrow$ upper bound on $\Lambda_{PQ} \leftarrow$ dominates long range forces
- > Viable solutions for $m_a > 10^{-5} \text{ eV}$: Example:

$$\frac{\text{DM solution for } \Lambda_{PQ}}{\text{Upper bound on } \Lambda_{PQ}} \approx 1.5 \times \left(\frac{10^{-10}}{\theta_{\text{eff}}} \right)^{1/2} \quad \text{for} \quad \delta_{PQ} = 3, \quad n = 2$$

- > All non-tuned solutions are close to this nEDM bound
- > Warning! Assumes free QCD close to Λ_{QCD} !

Summary

- > We expanded on existing literature by considering misalignment impact of temperature-dependent PQ violating potentials
- > Take-away points:
 - PQ-violating potentials can impact misalignment without violating nEDM
 - If V_{PQ} is constant in T near $T \sim \text{GeV}$, and there is no tuning, then
 - V_{PQ} can have an impact for $m_a \gg 10^{-3} \text{ eV}$
 - $\frac{\rho_{a, \text{today}}}{\rho_{\text{DM, today}}}$ cannot be raised further than $\text{few} \times 10^{-4}$
 - If V_{PQ} is T-dependent, is generated by $(\phi/\Lambda_{\text{QCD}})^n \times \mathcal{O}_{\text{SM}}$, and there is no tuning, then
 - axion DM can motivated by trapped misalignment across the $m_a \sim 10^{-5} \text{ eV} \rightarrow 10^{-1} \text{ eV}$ range
 - all DM solutions found here have clear signals in either fifth-force or nEDM
- > What to expect from our paper: Better overview of viable parameters, monopole-dipole forces, and conclusions on FF .

Thank you!

Questions?

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