

May 23rd, 2024

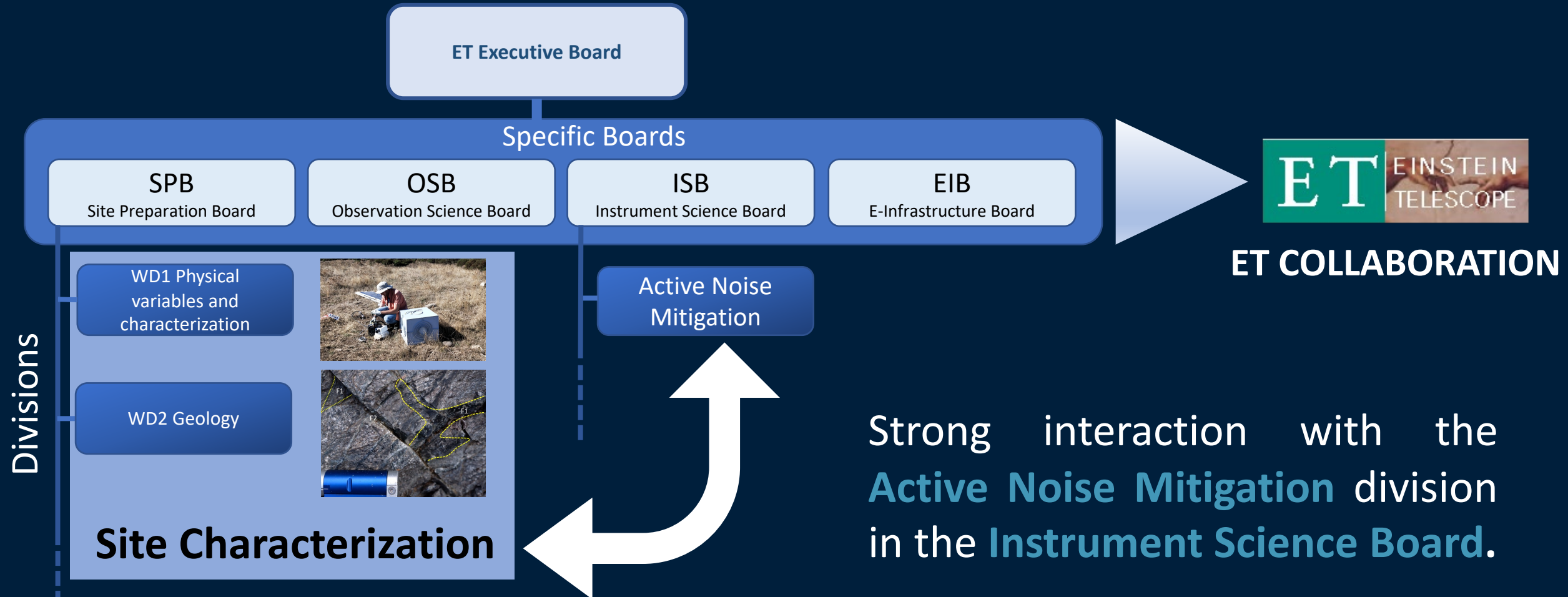
# Status of the Sardinia site

**Davide Rozza**  
on behalf of the Sardinian site characterization team

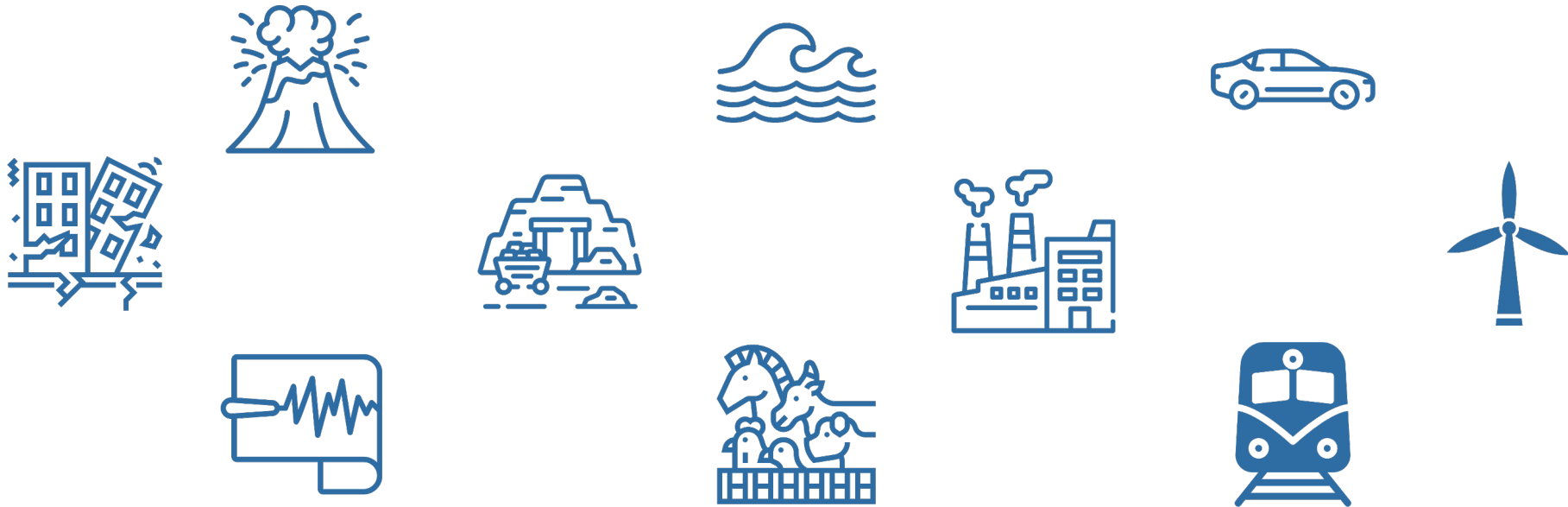




**Site Characterization** at the candidate sites is coordinated in the framework of the ET Collaboration **Site Preparation Board**.



Einstein Telescope will be built underground to operate below 10 Hz.  
Potential **noise sources**, both **natural** and **anthropic** origin,  
**Can affect the Einstein Telescope** measurements.



WE WILL SEE THAT THE ITALIAN CANDIDATE SITE IS CHARACTERISED BY:

Geodynamic quietness

Low Anthropogenic noise

Low E.M. noise<sub>3</sub>

Olbia

Bitti

Onanì

Lula

Olbia

Onanì

Bitti

Lula

Nuoro

100/300 m

10 km

In the **SOS ENATTOS** former mine area, the **SARGRAV laboratory**, a seed of ET, can host:

UNDERGROUND  
EXPERIMENTS

CRYOGENIC  
PAYLOADS

LOW FREQUENCY AND CRYOGENIC  
SENSOR DEVELOPMENT

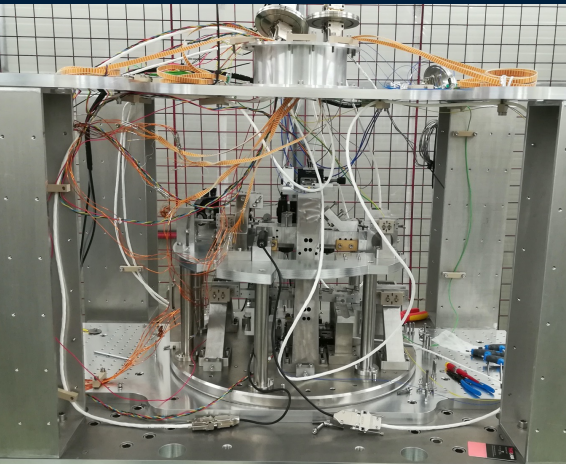
that need **LOW SEISMIC** and **ANTHROPOGENIC NOISE**





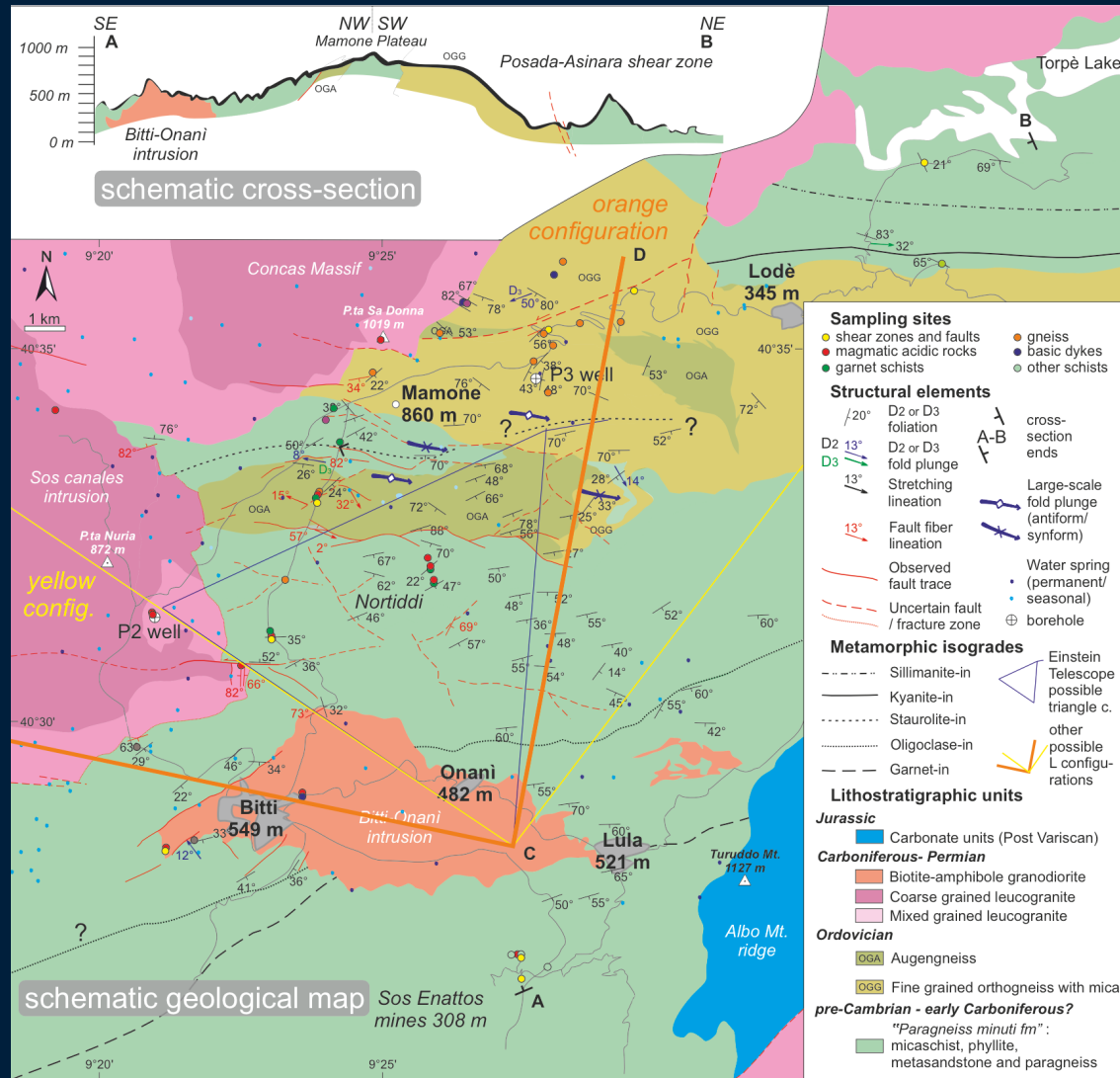
**2010**, first analysis was performed in the Sos Enattos mine.

The mine was not abandoned; therefore, it provided the adequate manpower, infrastructure and experience to start underground characterization studies in **2014**.



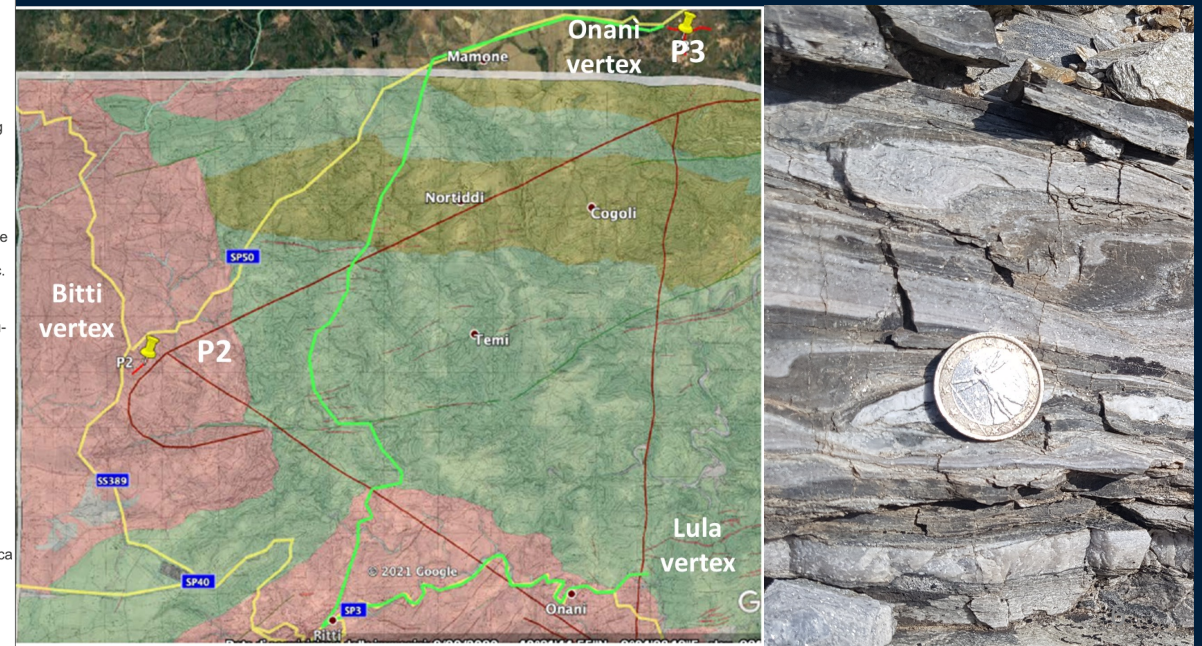
**Today**, the mine is considered a regional heritage site, it hosts the **SARGRAV LABORATORY** and the **ARCHIMEDES EXPERIMENT**.

# The ET Italian candidate site is located in the stable VARISCAN BASEMENT OF SARDINIA



**LITHOLOGIES:** Orthogneiss, granitoids, micaschists.

**P2 and P3 are the borehole locations**  
optimization is ongoing.





# PERMANENT ARRAY since 2019

Since 2019, in Sos Enattos there are:

4 permanent seismic stations for long term studies  
(Trillium 240, 360 and 120 Horizon, Guralp 360)

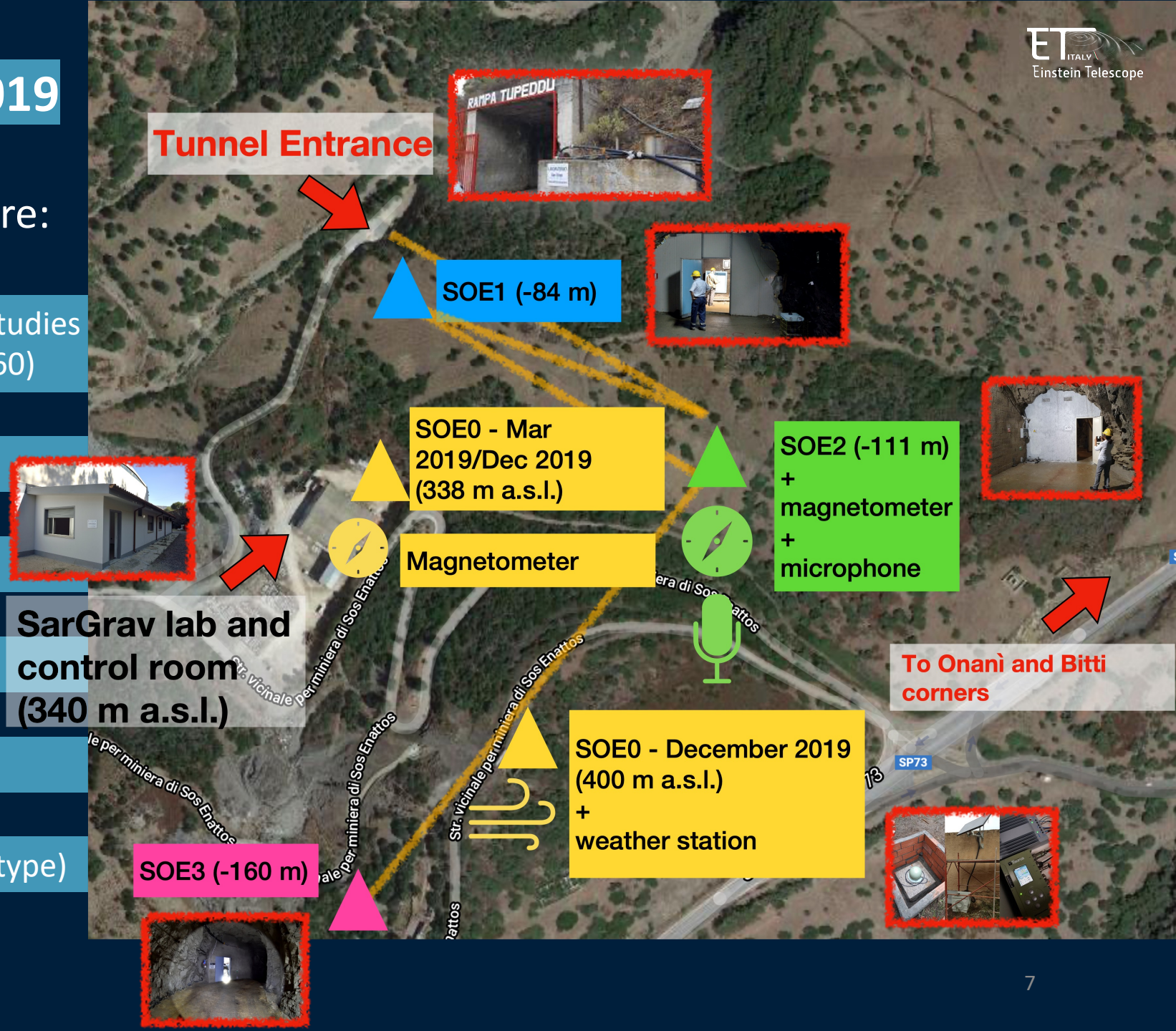
1 weather station

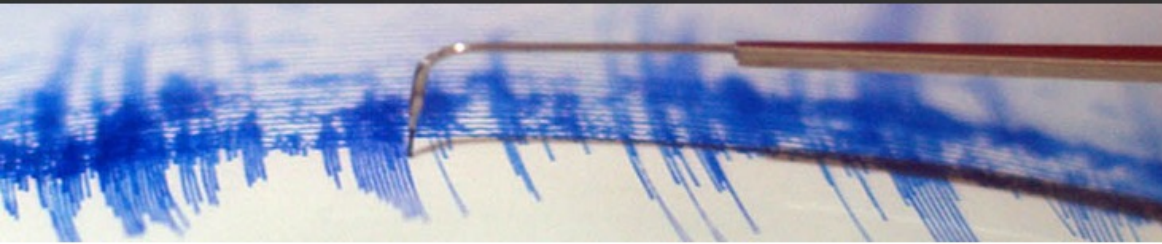
1 microbarometer

3 magnetometers (MF6-06)

2 microphones

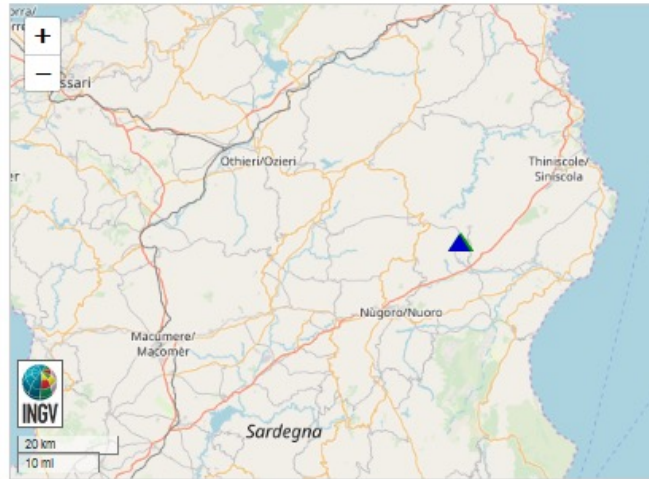
1 high precision tiltmeter (Archimedes prototype)





### Seismic Station SENA Sos Enattos Mine

Network: IV  
Start Date: 2019-10-18T00:00:00  
End Date: --  
Latitude: 40.4444  
Longitude: 9.4566  
Elevation: 338  
[Download StationXML](#)



Number of channels: 3

#### Channel List

Code	Location Code	Start Date	End Date	Data Restriction
HHE		18-10-2019		open
Latitude: 40.4444		Azimuth: 90		
Longitude: 9.4566		Sample Rate: 100		
Elevation: 338		Storage Format: Steim2		
Depth: 111		Sensitivity Value: 478760000		

SOE2 station is integrated into the Italian national seismometer **network of INGV**.  
Station: **SENA**, network:

- IV (Italian National Seismic Network - INSN), 2019-2022/01
- MN (*Mediterranean Very Broadband Seismographic Network*) since 2022/02

<http://cnt.rm.ingv.it/en/instruments/station/SENA>



## PERMANENT ARRAY since 2021

Since 2021, more permanent sensors have been installed at 2 of the proposed vertices (P2, P3)

2 broadband seismometers on surface

2 broadband seismometers in borehole

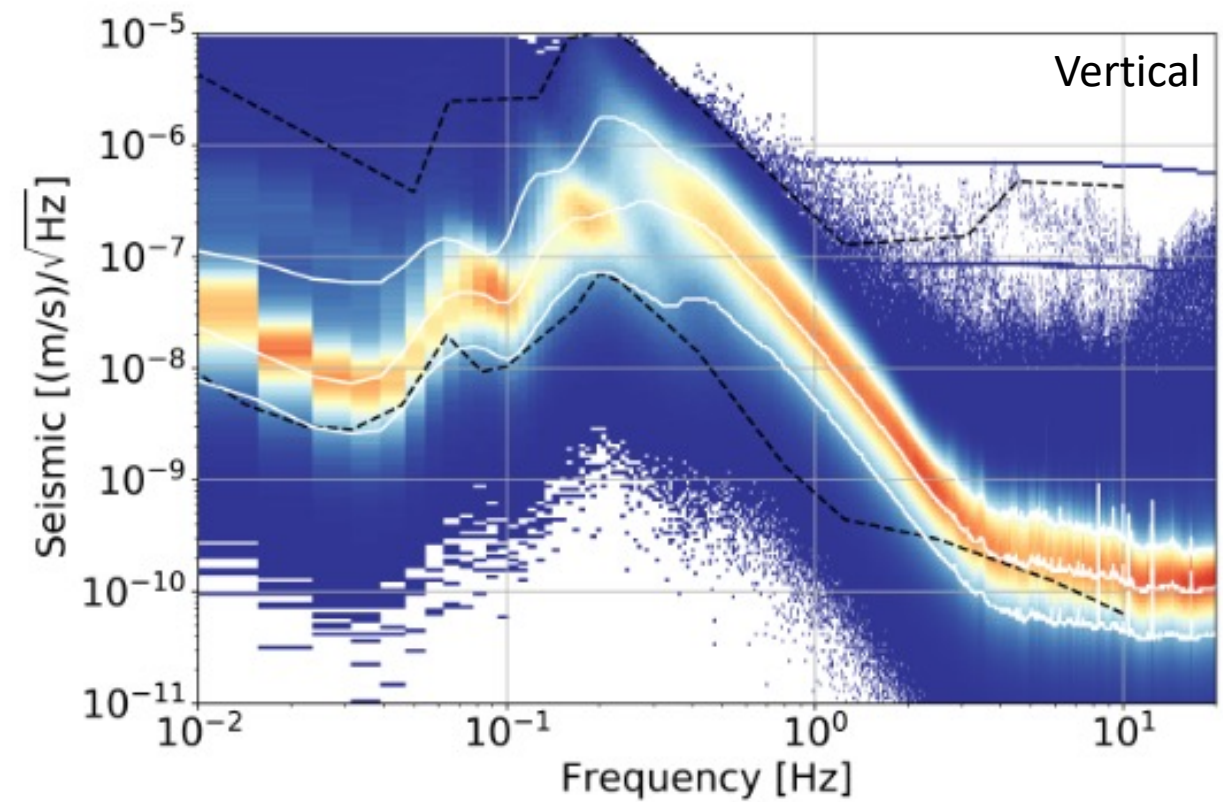
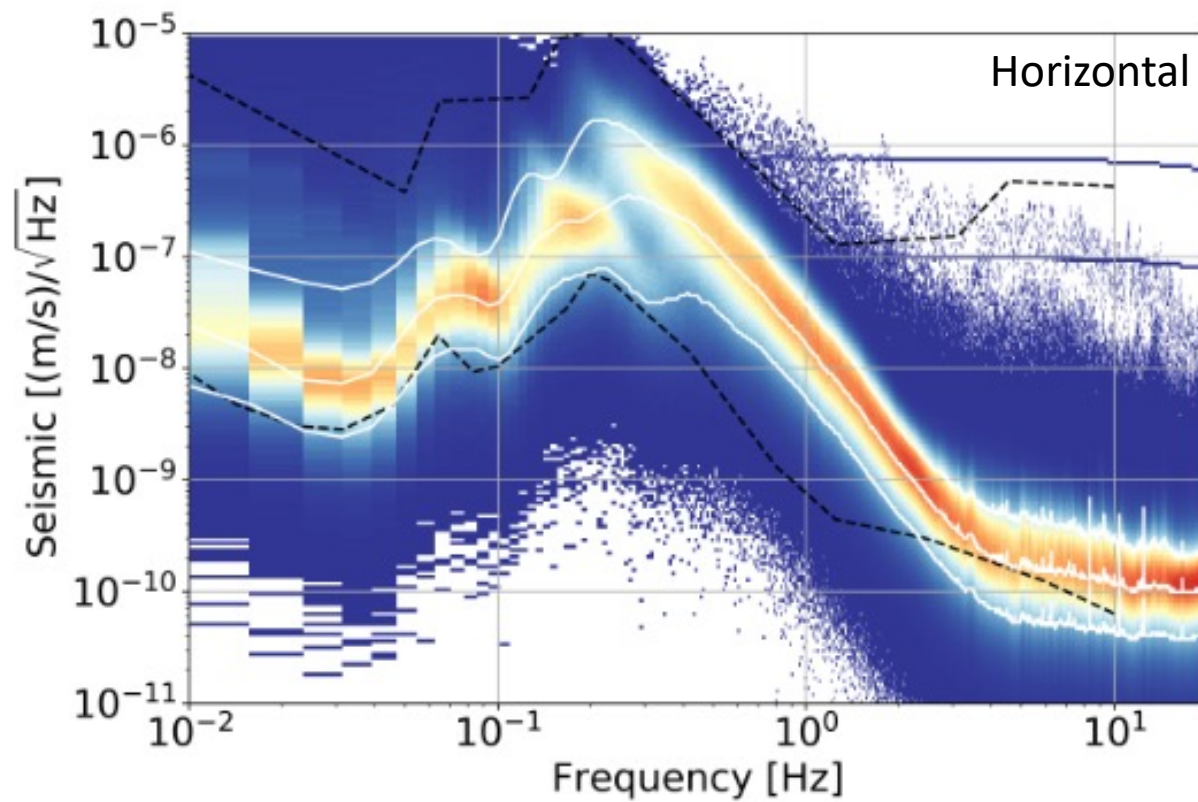
2 magnetometers at P2

Collaboration with  
Germain group (KIT)



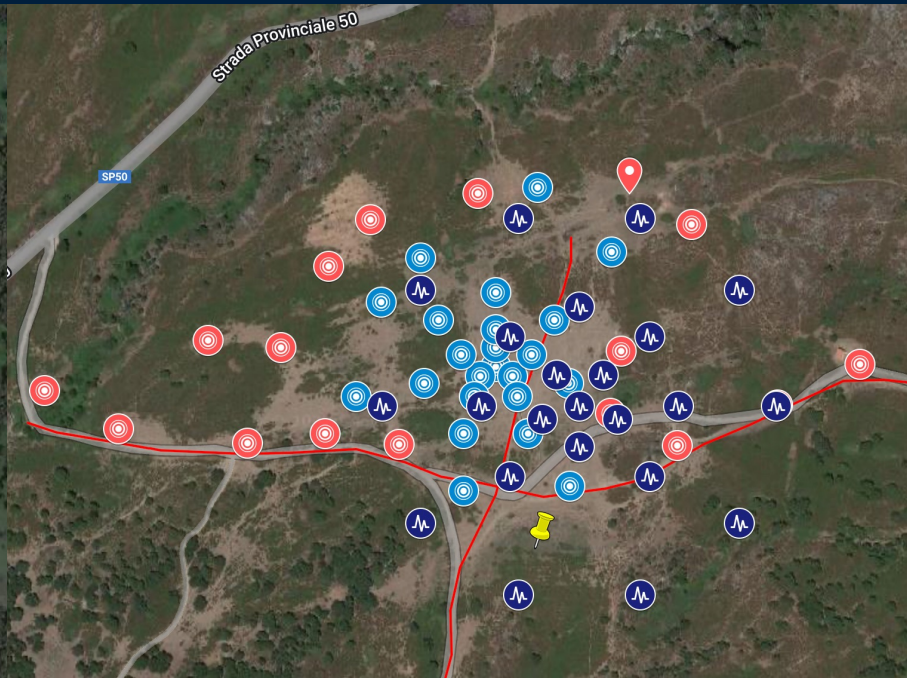


# PPSD - P2 borehole seismometer

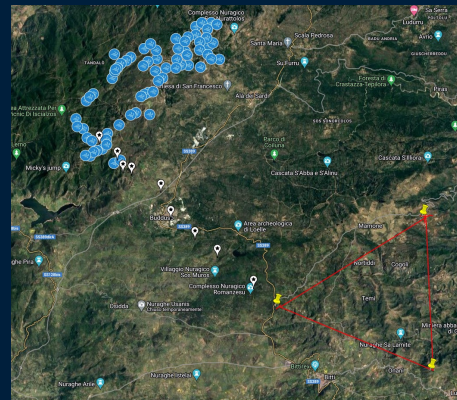


Very low noise background in the **2-10 Hz** band, sometimes even **below** the Peterson's **New Low Noise Model**!





### P3 broadband array + geophones (July & Oct 2021)



## Wind Park broadband array (early 2023)



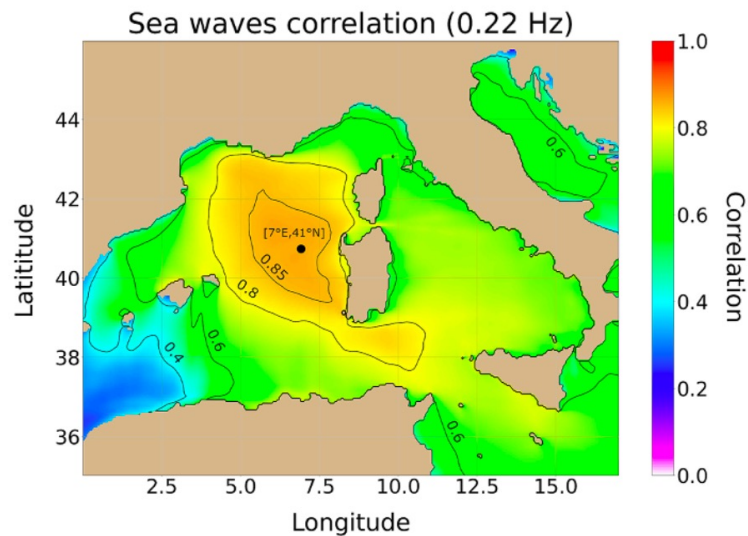
The background image shows the Einstein Telescope facility, a large red steel structure with a solar panel canopy, situated in a valley with green hills and trees in the background. A small white building and a car are visible near the structure.

**IDENTIFICATION** of some  
potential noise sources, both  
of **natural** and **anthropic** origin



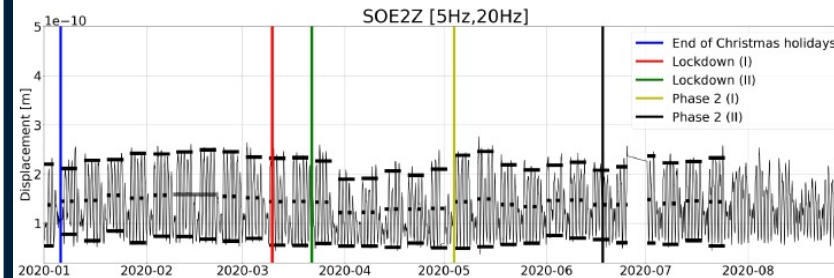
# Hunting the noise sources

## Sea waves



**Microseismic** (0.05-1Hz) coming from **waves** in the Gulf of Lion (~0.22Hz, NW Mediterranean sea) and Atlantic ocean depending on weather conditions and season.

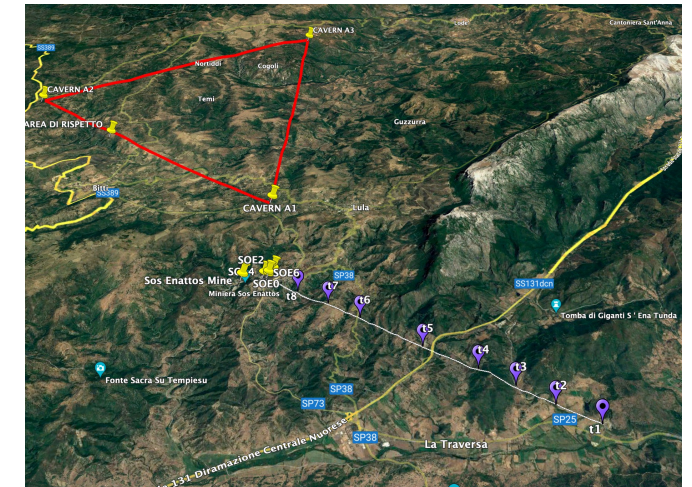
## Human activities



At higher frequencies, **anthropic noise** pattern observed.



Identification of the noise contribution from **two bridges** (peaks at ~4.2 Hz, 4.6 Hz, 6 Hz).





# Hunting the noise sources

The **Budussò Wind Park**: one of the largest wind parks in Italy.

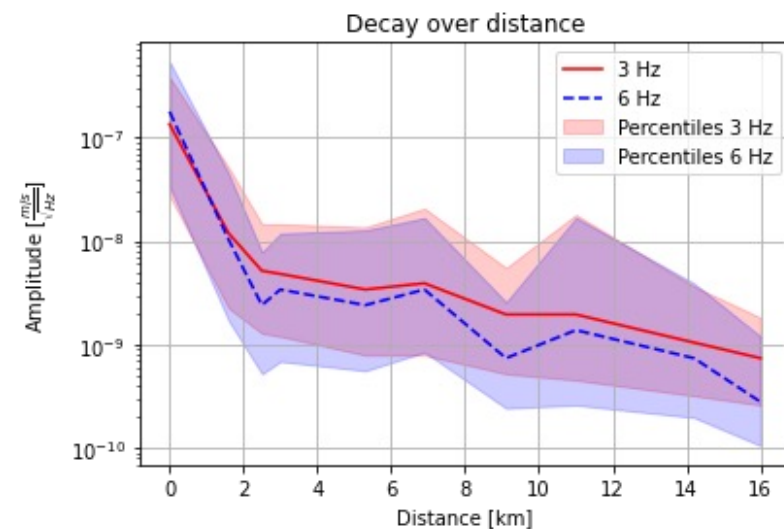
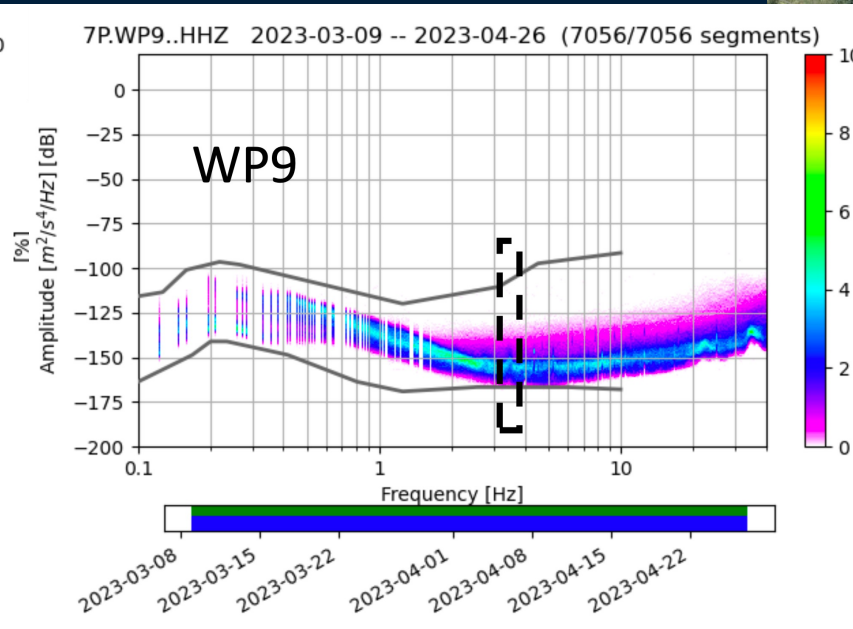
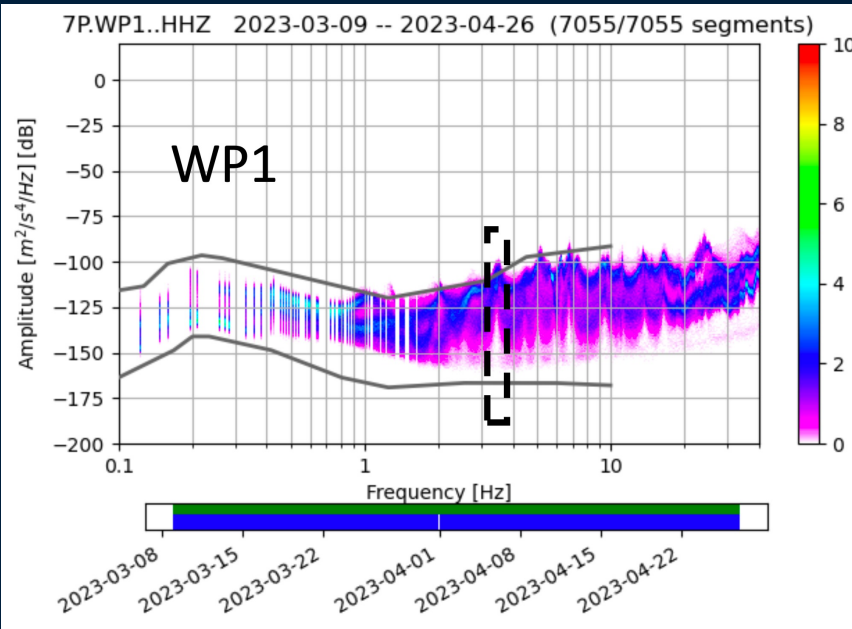
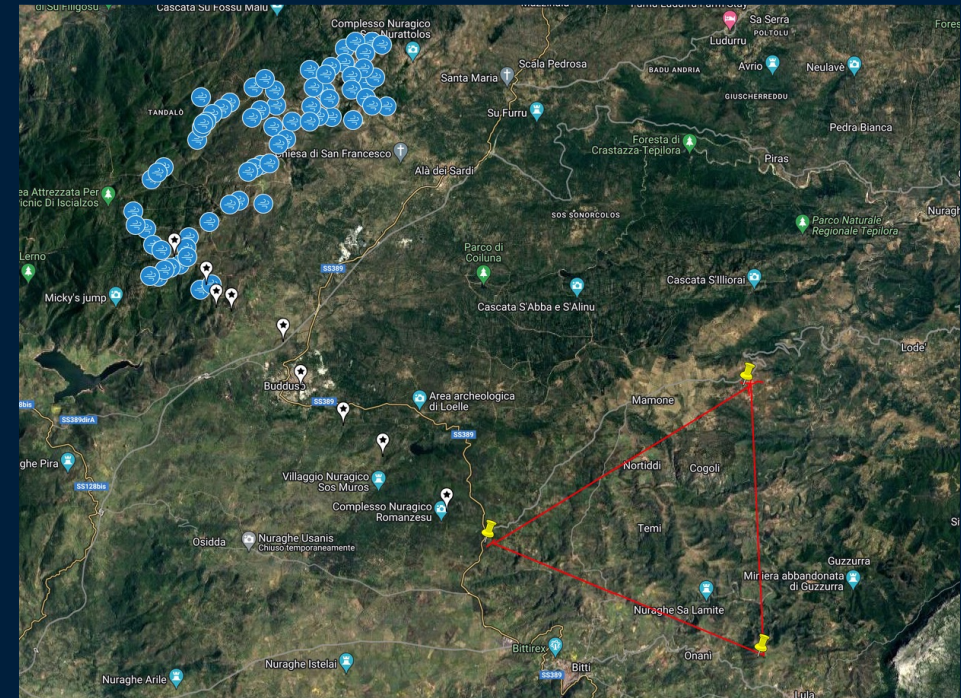
**69 turbines** (~2 MW each).

A total of **130 MW** installed.

Blades motion is **transferred** to tower and to the ground.

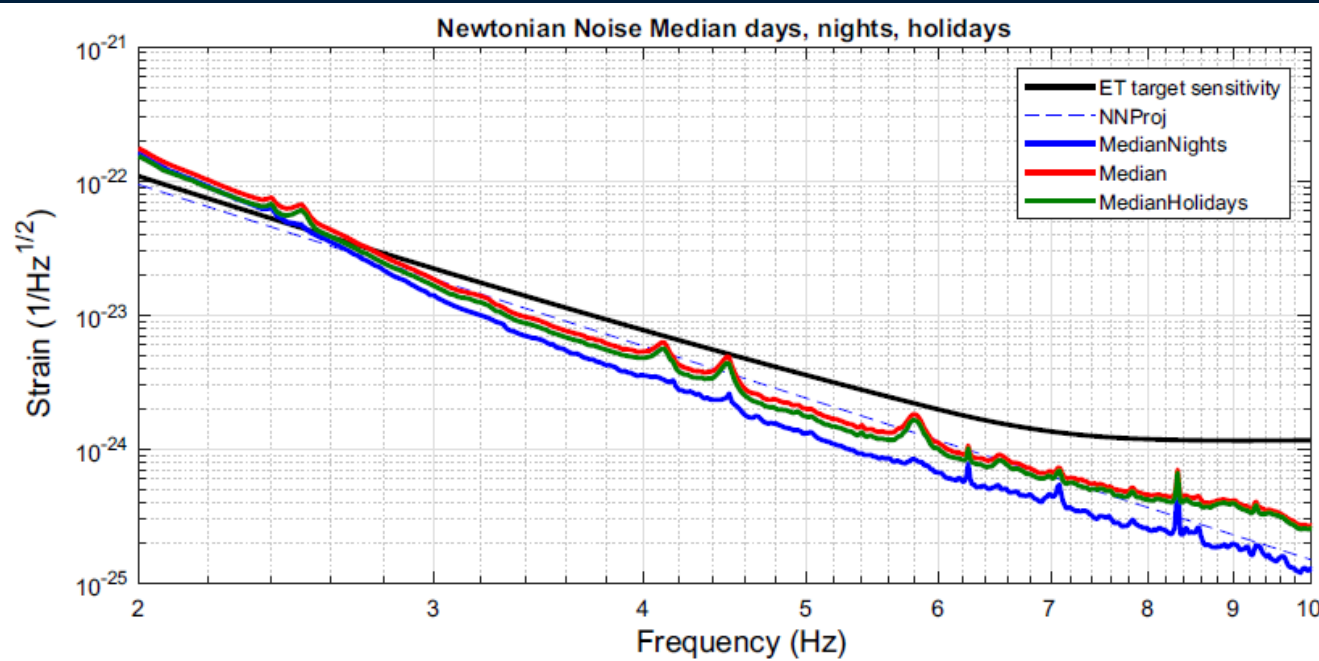
Seismic noise propagates as **surface waves**

Generated noise is found in the **1-10 Hz** frequency band.





# Hunting the noise sources



NEWTONIAN NOISE and seismic glitches  
(based on 2020 data at SOE1)

Defining the Newtonian Noise ASD as:

$$\tilde{h}_{NN}(f) = \frac{4\pi}{3} G \rho_0 \frac{2\sqrt{2}}{L} \frac{1}{(2\pi f)^2} \tilde{x}(f)$$

Defining the Noise-to-Target Ratio of the Newtonian Noise in 1 minute window ( $\sim$ IMBH duration in ET band)

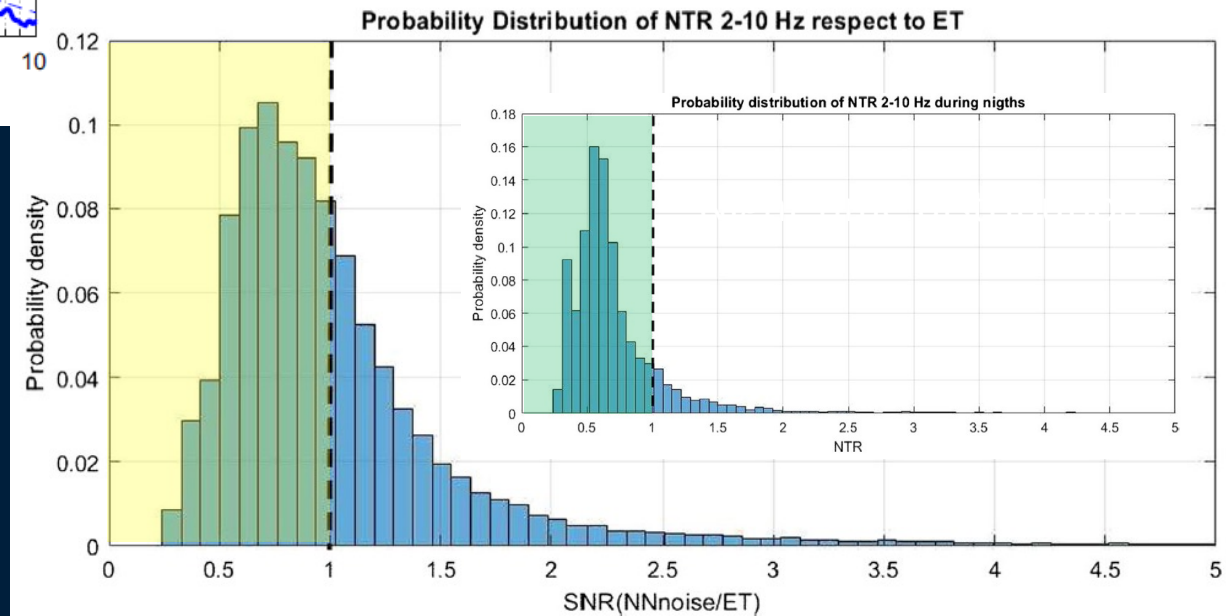
$$\text{NTR} = \sqrt{\frac{1}{\Delta f} \int df \frac{\tilde{N} * \tilde{N}}{S_h}}$$

PSD of NN

PSD of ET sensitivity

$P(\text{NTR} < 1) = 0.6$ , during nights:  $P(\text{NTR} < 1)_n = 0.86$

Need for moderate NN subtraction only for a limited time



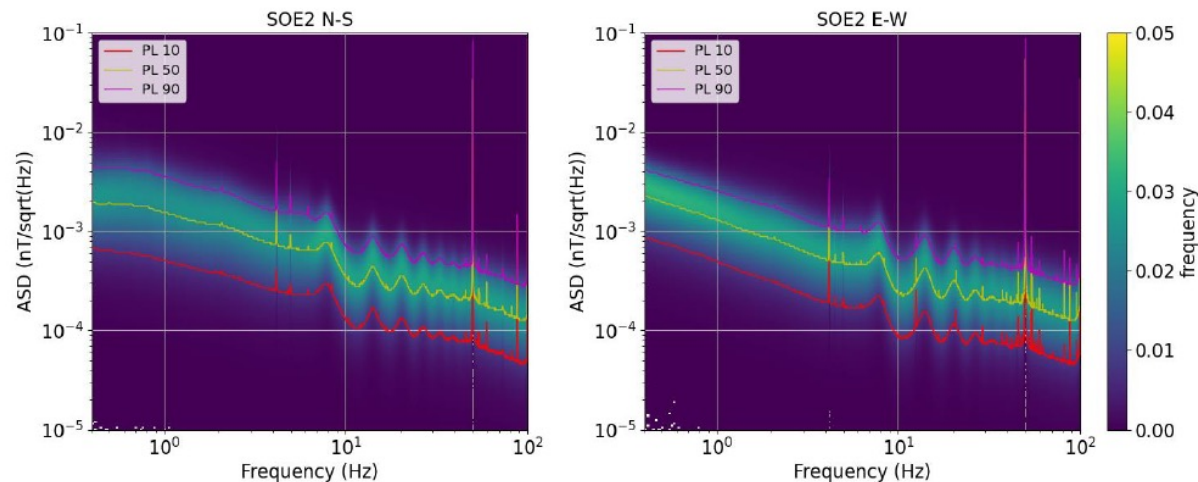
## MAGNETIC NOISE



Most important mechanism in ET-LF:

- **Geomagnetic pulsations Pc1** (0.2-5Hz);
- **Schumann resonances** (5-100Hz)

Artificial LF sources (e.g. 50-60Hz powerlines)

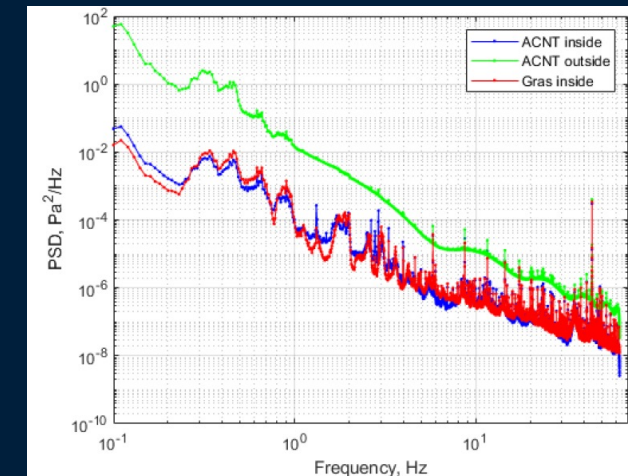


credits: R. De Rosa, R. Romero

## INFRASOUND MEASUREMENTS



**4 microphones** installed along the underground tunnels for long term characterization in a joint Italian-Polish-Hungarian collaboration (*PolGrav-AstroCeNT, Wigner Research Centre*)



credit: T. Bulik



# PUBLICATIONS:

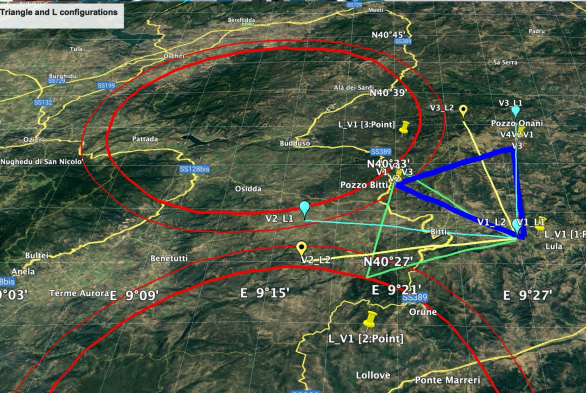
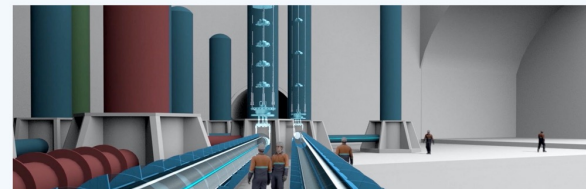
- ❑ L. Naticchioni et al., *Microseismic studies of an underground site for a new interferometric gravitational wave detector*, **CQG**, 2014, <https://doi.org/10.1088/0264-9381/31/10/105016>
- ❑ L. Naticchioni et al., *Characterization of the Sos Enattos site for the Einstein Telescope*, **JPCS** 1468, 2020, <https://doi.org/10.1088/1742-6596/1468/1/012242>
- ❑ M. Di Giovanni et al., *A seismological study of the Sos Enattos Area – the Sardinia Candidate Site for the Einstein Telescope*, **SRL**, 2020 <https://doi.org/10.1785/0220200186>
- ❑ A. Allocca et al., *Seismic glitchness at Sos Enattos site: impact on intermediate black hole binaries detection efficiency*, **EPJP**, 2021 <https://doi.org/10.1140/epjp/s13360-021-01450-8>
- ❑ Allocca et al. *Picoradiant tiltmeter and direct ground tilt measurements at the Sos Enattos site*, *Eur. Phys. J. Plus* **136**, 1069 (2021). <https://doi.org/10.1140/epjp/s13360-021-01993-w>
- ❑ M. Di Giovanni et al., *Temporal variations of the ambient seismic field at the Sardinia candidate site of the Einstein Telescope*, **Geophysical Journal International**, 2023, <https://doi.org/10.1093/gji/ggad178>
- ❑ G. Saccorotti et al., *Array analysis of seismic noise at the Sos Enattos mine, the Italian candidate site for the Einstein Telescope*, 2023, <https://doi.org/10.1140/epjp/s13360-023-04395-2>.
- ❑ L. Naticchioni et al., *Results of the site characterization in Sardinia for the Einstein Telescope*, **PoS Proc. Sci.**, 2023, *accepted for publication*.

+ several internal notes, reports and talks

# NEWS and UPDATES

**EINSTEIN TELESCOPE: CORDATA DI AZIENDE ITALIANE GUIDATA DA ROCKSOIL SI AGGIUDICA LA GARA DA OLTRE 12 MILIONI DI EURO**

Febbraio 5, 2024 · 3 min read



TERABIT NETWORK FOR RESEARCH AND ACADEMIC BIG DATA IN ITALY

Terabit Network for Research and Academic Big Data in Italy ha l'obiettivo di integrare e potenziare tre Infrastrutture digitali di Ricerca strategiche:

• GARR-T, infrastruttura di rete a supporto dell'istruzione e della ricerca in Italia.

• PRACE-Italy, infrastruttura di calcolo ad alte prestazioni (HPC), nodo italiano dell'infrastruttura europea PRACE.

• HPC-BD-AI, infrastruttura di calcolo distribuita di tipo cloud-edge, in grado di gestire risorse di calcolo ad alte prestazioni, big data e applicazioni di intelligenza artificiale.

[Scopri di più](#)

Preliminary **feasibility study** for ET in Sardinia on going (both  $\Delta$  and L).

Preliminary **cost estimation** (excavation).

**Management and strategy** of excavated lands & rocks.

End of activity: June 2025  
Bidbook to end 2025

**Optimization of ET localization**

Sos Enattos area will be reached at 1 TB/s (**TeraBit project**)

**Acoustic campaign** at P2 & P3 borehole areas

**Gravimetric campaign**



# Conclusions

The **geo-physical site characterization** is a task coordinated by SPB of the ET collaboration.

Site characterization is strictly related to the **noise mitigation strategy and detector design**.

**LF sources of noise** (2-10Hz) affected by **seismic** ( $\rightarrow$  **Newtonian**) and **magnetic** noise.

Sardinia is **geologically very quiet**, far from active fault lines, and characterized by low anthropic noise.

Since 2019 (and before in 2010-2014), we installed **permanent and temporary arrays** of sensors and **two instrumented boreholes** (operative since 2021).

Measurements show a peculiar **very low level of seismic noise** in the ET-LF band (2-10Hz), where:

Seismic noise match or goes even below the Peterson's NLNM

The projected (seismic) Newtonian noise is also compatible with  $ET_d$

Electromagnetic noise is very low

Acoustic noise measurement ongoing (also very quiet)

Possible **local sources** of noise (e.g. wind farms) are under study.

**From the geological and physical point of view, Sardinia is an optimal candidate to host the Einstein Telescope, either in  $\Delta$  or in L ( $\rightarrow$  2 sites) configuration!**







## A fruitful collaboration!

The geophysical characterization is a great and fruitful example of **collaboration** between Italian Research Institutes (INFN & INGV), that have brought together **different and complementary skills and expertise** to demonstrate the **extraordinary quality** of the Sardinia candidate site to host ET.

