

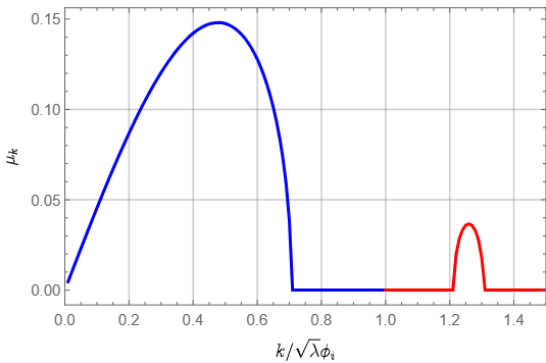
First year PhD student at **DESY Hamburg** working on **Axion phenomenology in the early universe**

Supervisor: **Geraldine Servant**

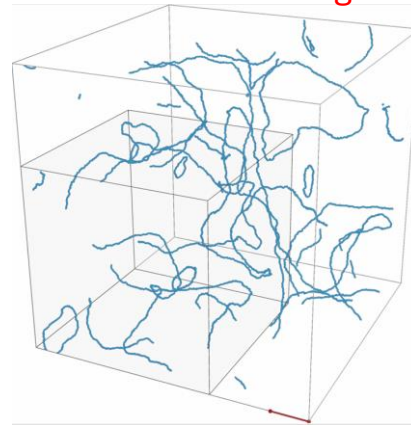


## Scientific Interests:

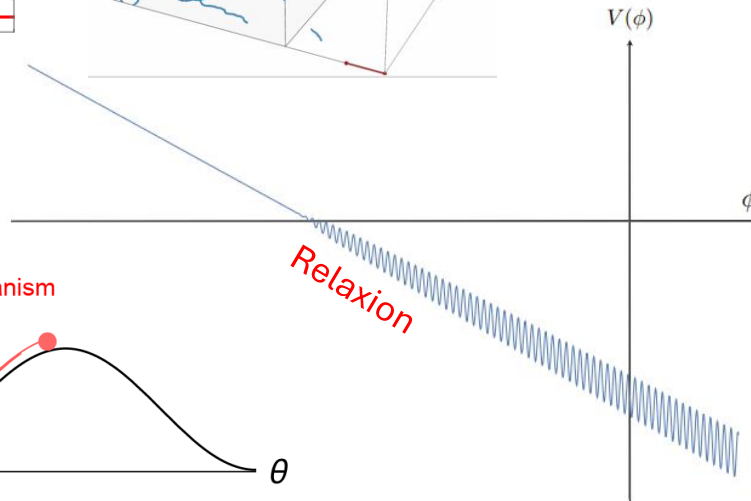
### Parametric Resonance



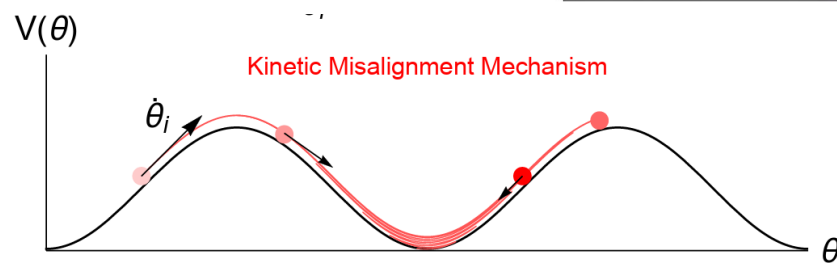
### Strings



### Relaxion



### Kinetic Misalignment Mechanism



## Hobbies:

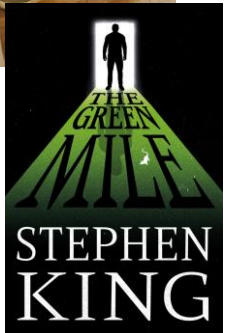
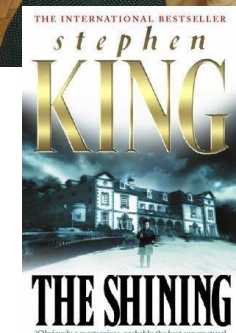
Jazz dances



Guitar



Horror books



# Kinetic Axion from non-minimally coupled PQ field

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Enrico Morgante, Riccardo Natale    2508.XXXXX



# Misalignment Mechanism

PQ symmetry restored at high energies

Symmetry breaking around the scale  $f_a$

Explicit symmetry breaking at low energies via non perturbative effects

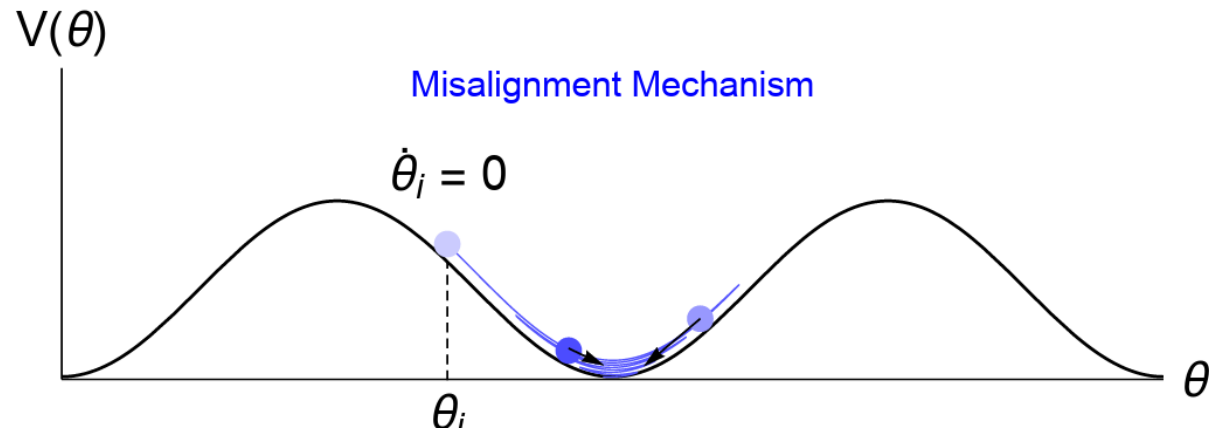
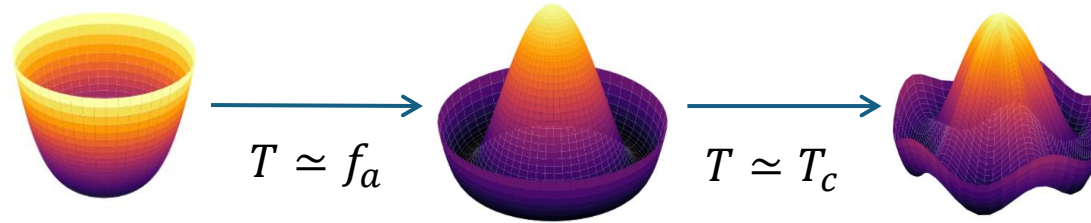


Fig. by «Harigaya et al, Axion Kinetic Misalignment Mechanism»

# Misalignment Mechanism

PQ symmetry restored at high energies

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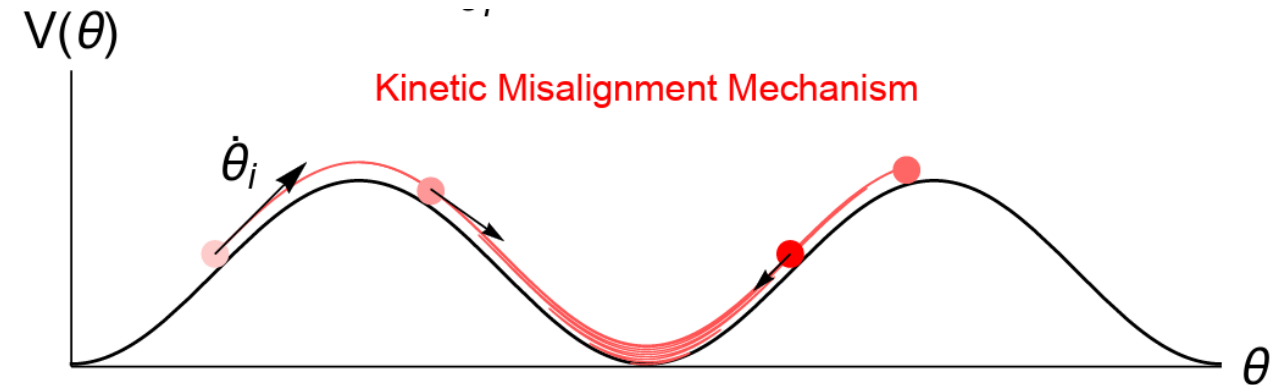
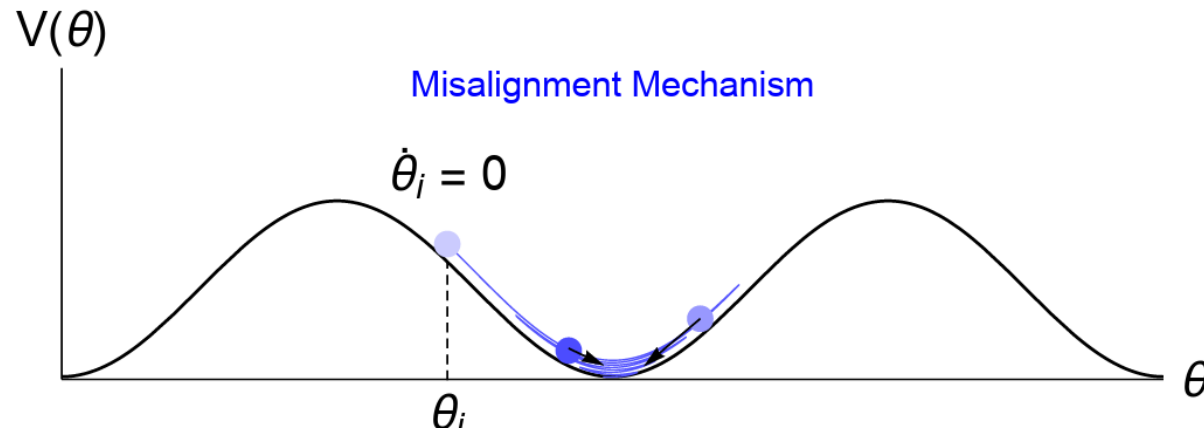
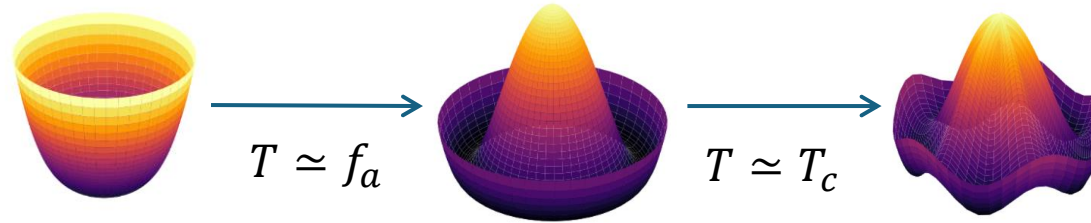


Fig. by «Harigaya et al, Axion Kinetic Misalignment Mechanism»

# How to generate a large axion velocity

Kinetic Misalignment can be generated by **high-order operators** producing an **explicit breaking of U(1) symmetry**

$$V_{\mathcal{PQ}} = 2^{\frac{n}{2}} \frac{A\Phi^n}{nM_{\text{Pl}}^{n-3}} + \text{h.c.}$$

In order to make this operators relevant the radial mode has to reach **high field values**

- **Quantum Fluctuations** during inflation
- **Hubble-Induced mass**
- **Coupling with the Inflaton**

2408.17013 Lee, Menkara, Seong, Song

2408.08355 Eröncel, Sato, Servant, Sørensen

2312.17730 Co, Yamada

2310.17710 Lee, Menkara, Seong, Song

2004.00629 Co, Hall, Harigaya, Olive, Verner

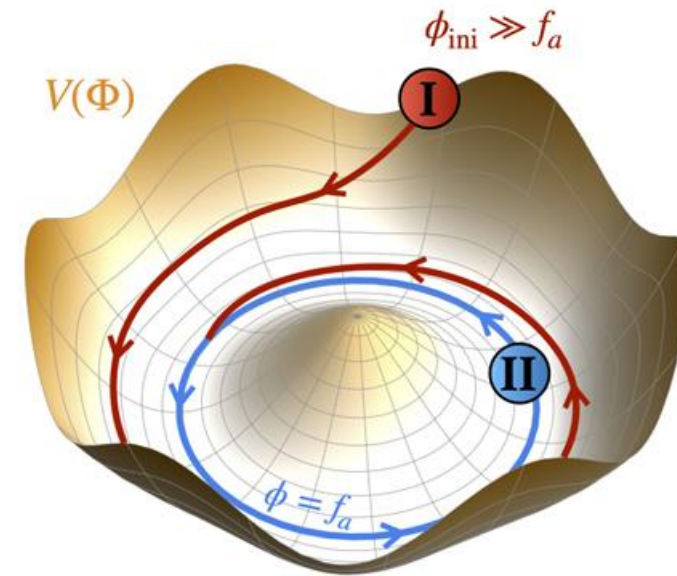


Fig. by Gouttenoire et al, «Kination Cosmology from scalar fields and gravitational waves signatures»

# Kinetic Axion with NMC PQ field

$$S = \int d^4x \sqrt{-g} \left[ \frac{M_{\text{Pl}}^2}{2} R + \mathcal{L}_{\text{inf}} - g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi^\dagger - \xi R |\Phi|^2 - V_{\text{PQ}}(|\Phi|) - V_{\text{PQ}}(\Phi, \Phi^\dagger) \right]$$

- A **non-minimal coupling**
- A **stiff era**

This transition generates a **tachyonic instability**

$$R = 3(1 - 3\omega)H^2$$

$$m^2 \approx 12\xi H_I^2$$

During Inflation

$$m^2 \propto -H^2$$

After Inflation

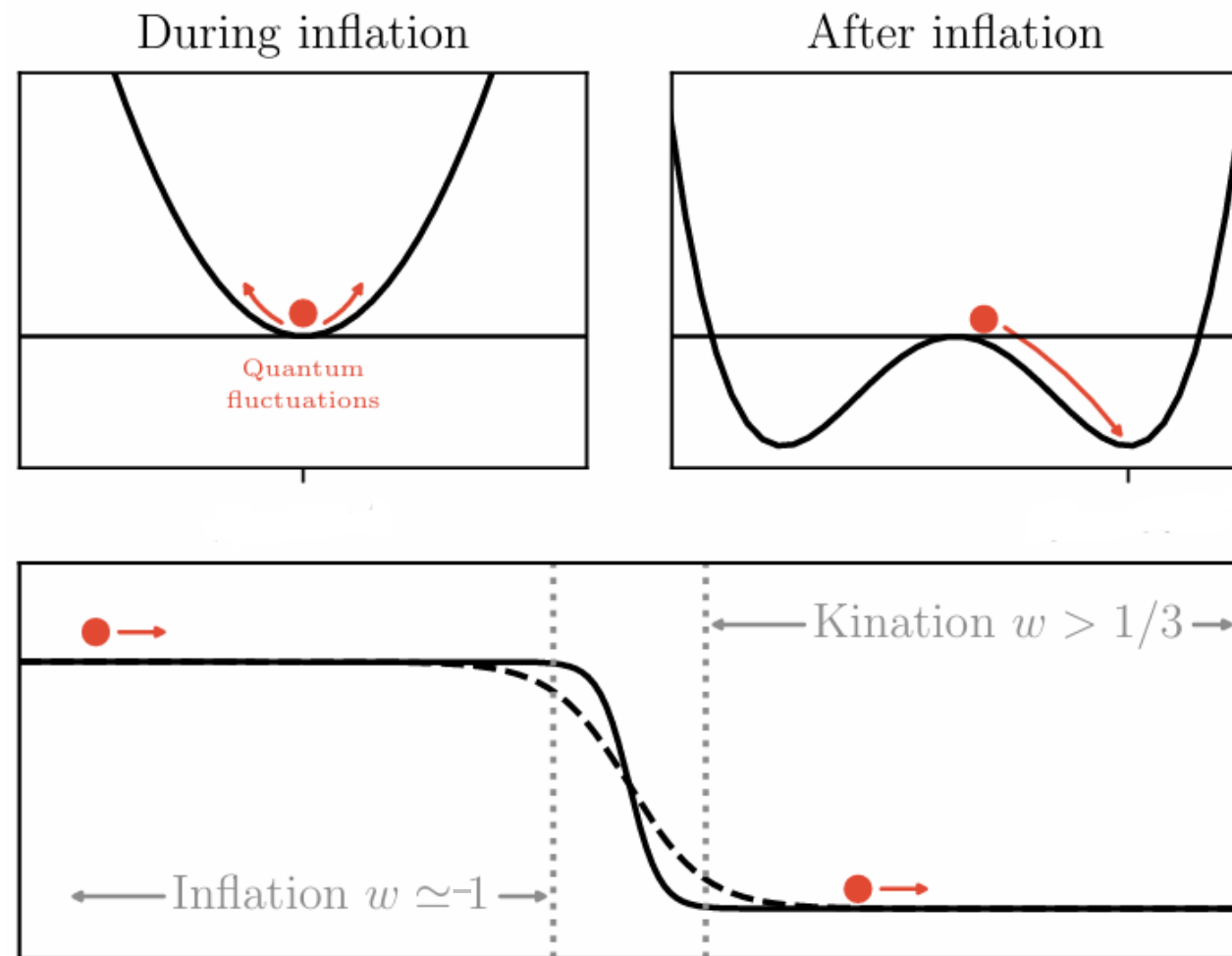
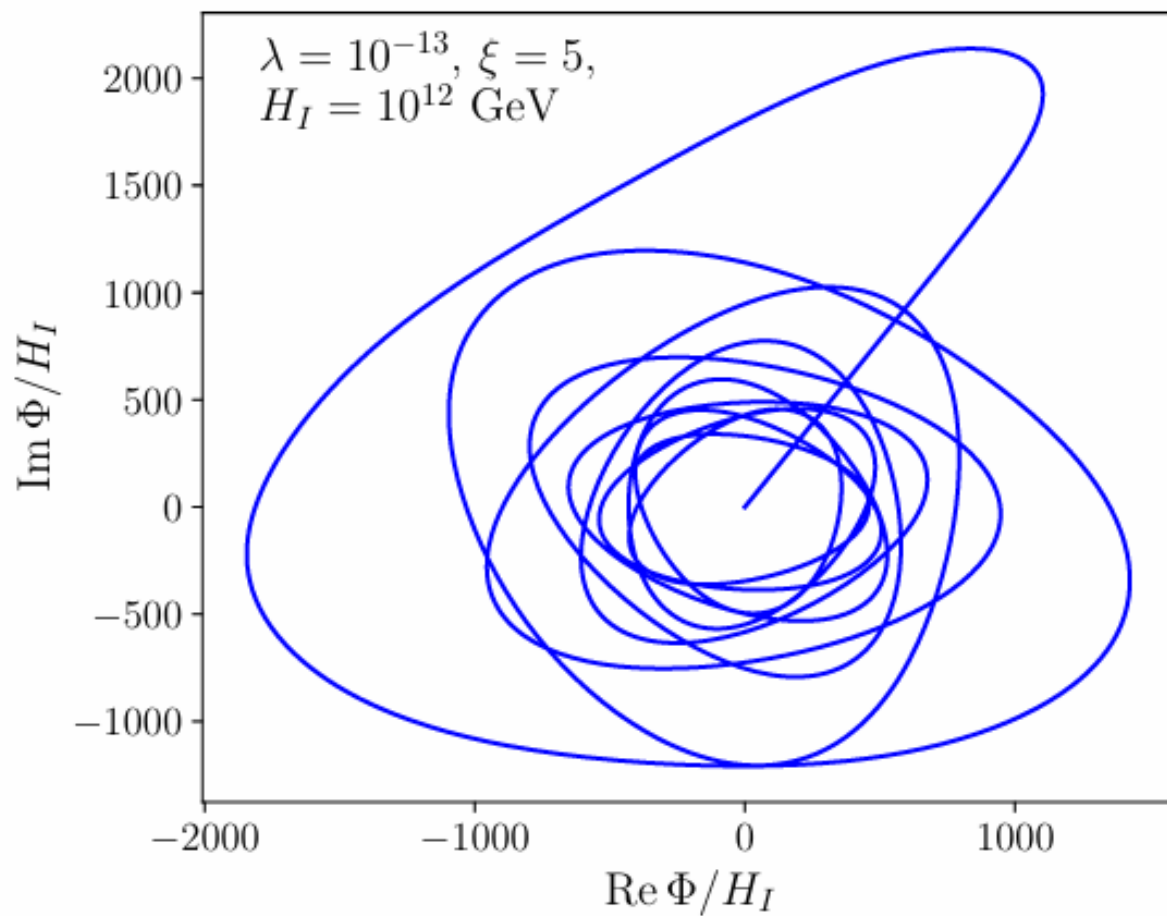
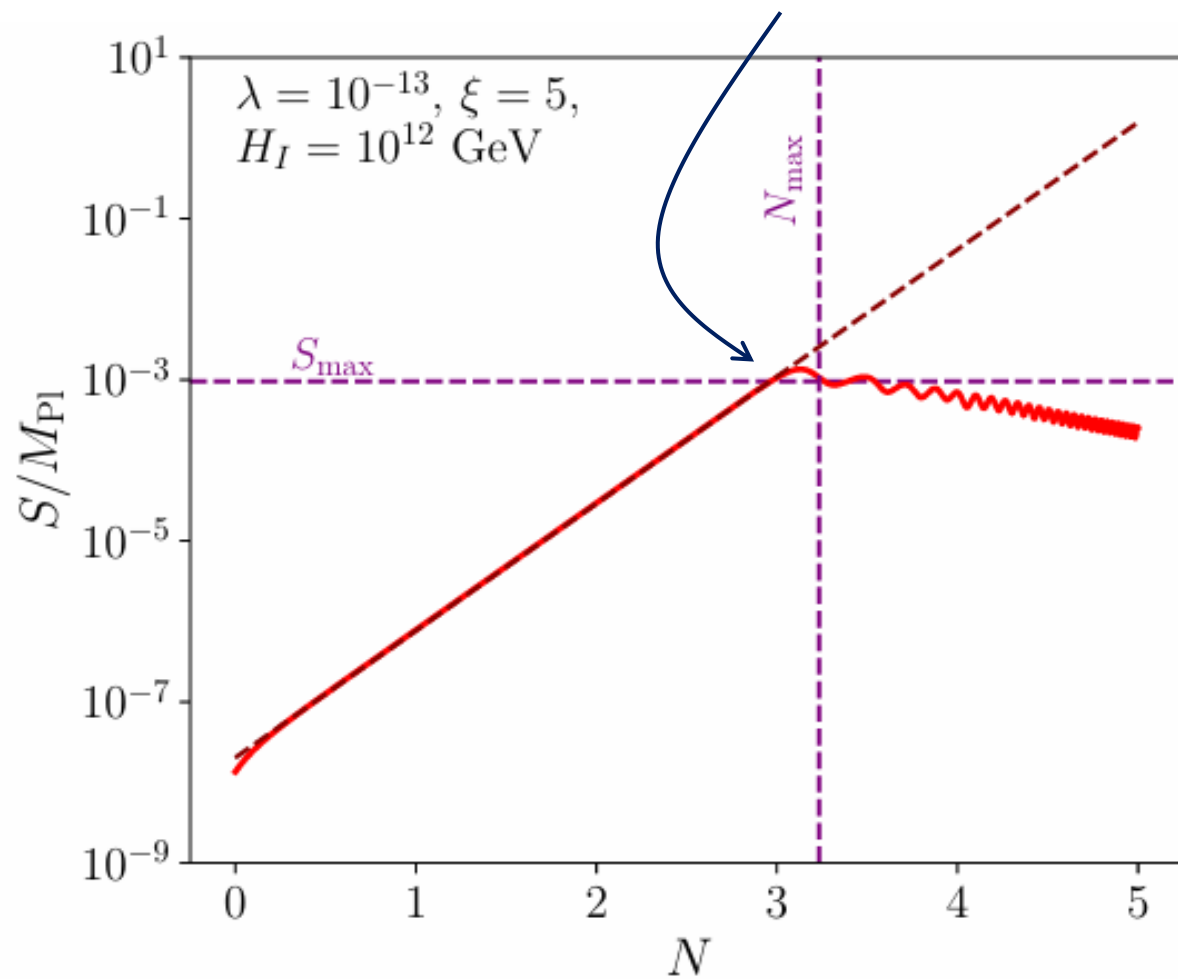


Fig. by Figueroa et al «Ricci Reheating on the Lattice»

# PQ Field Dynamics

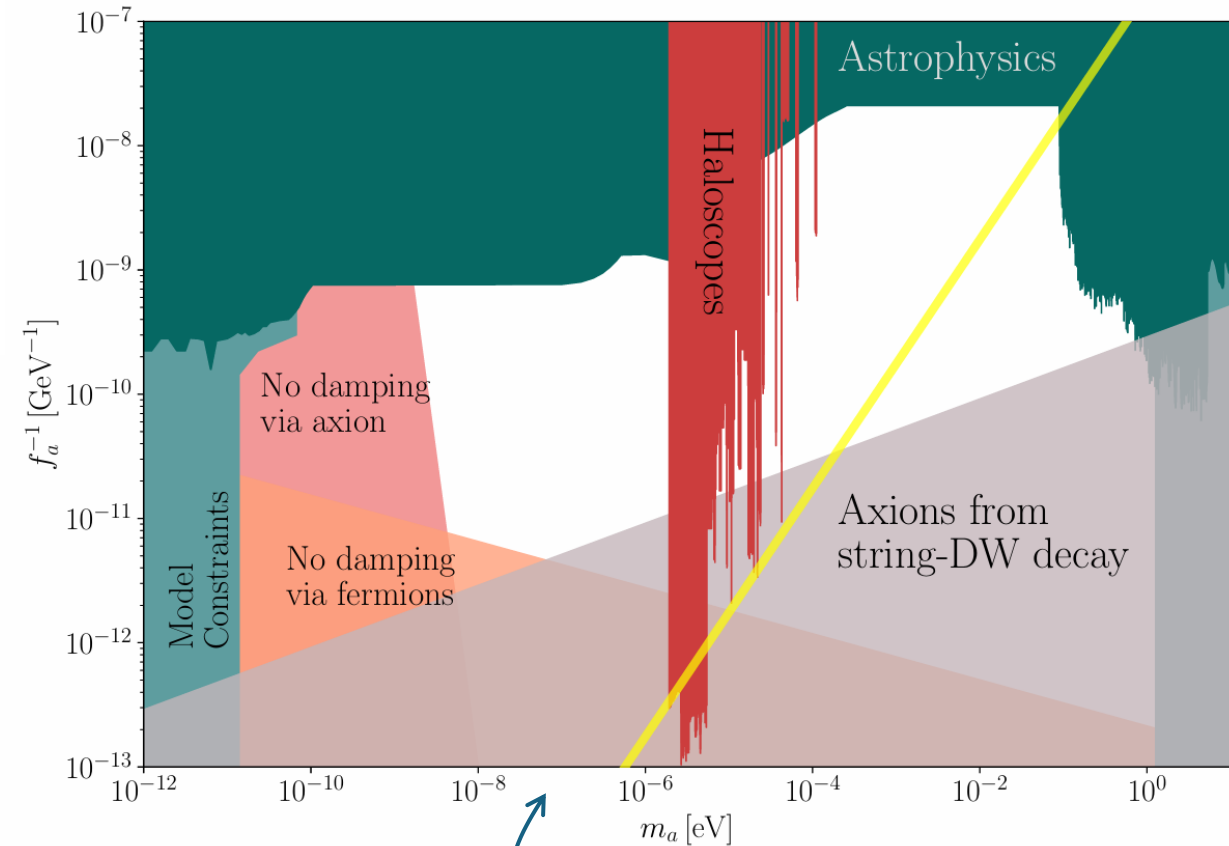
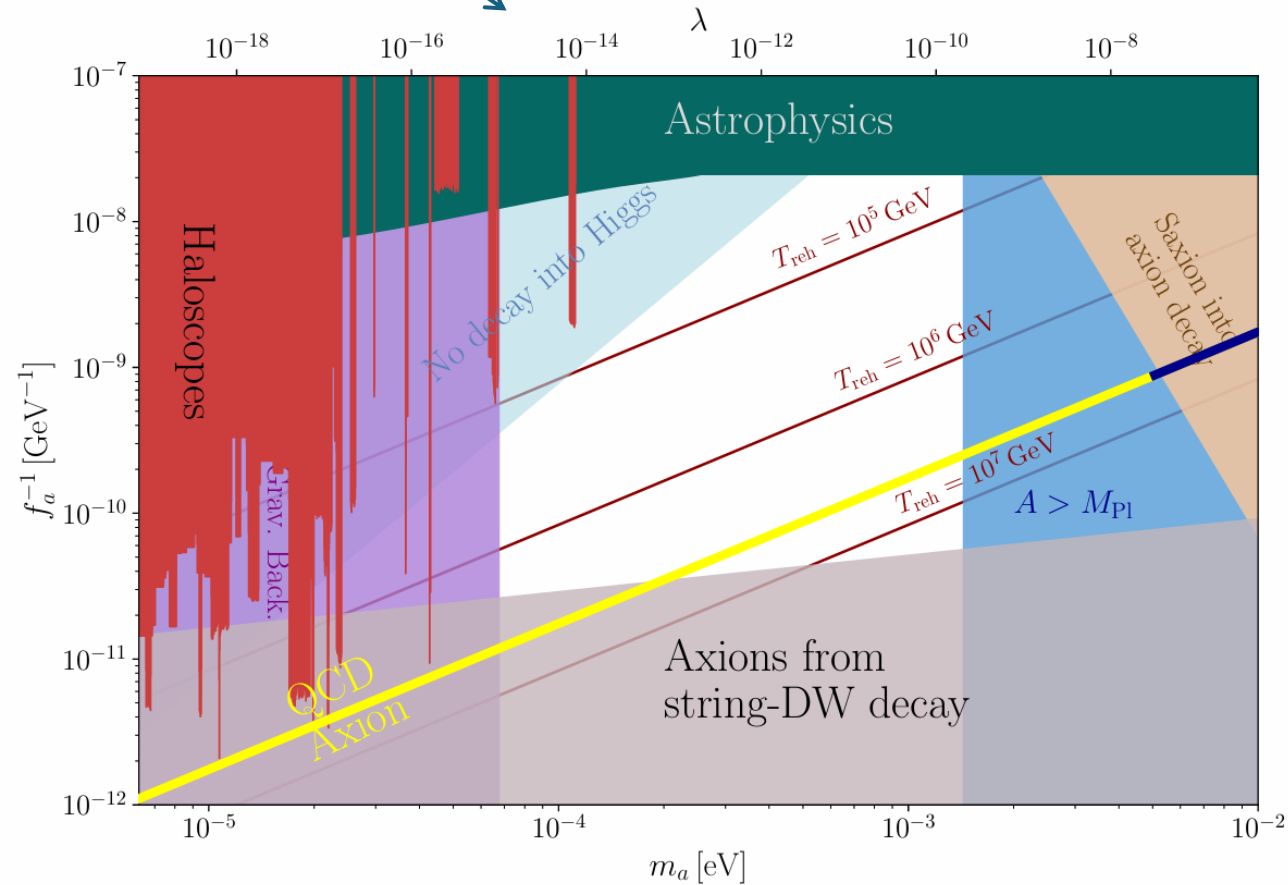
Here the field starts rotating!





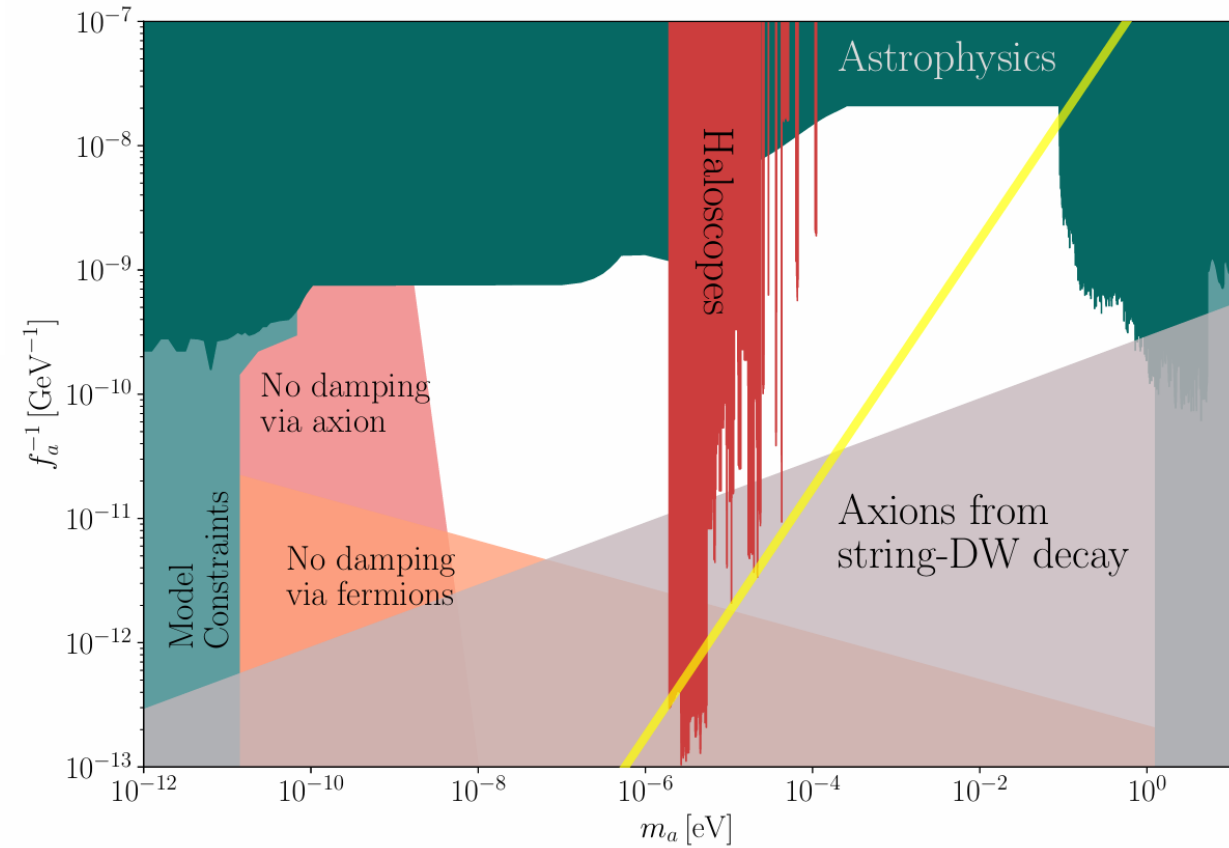
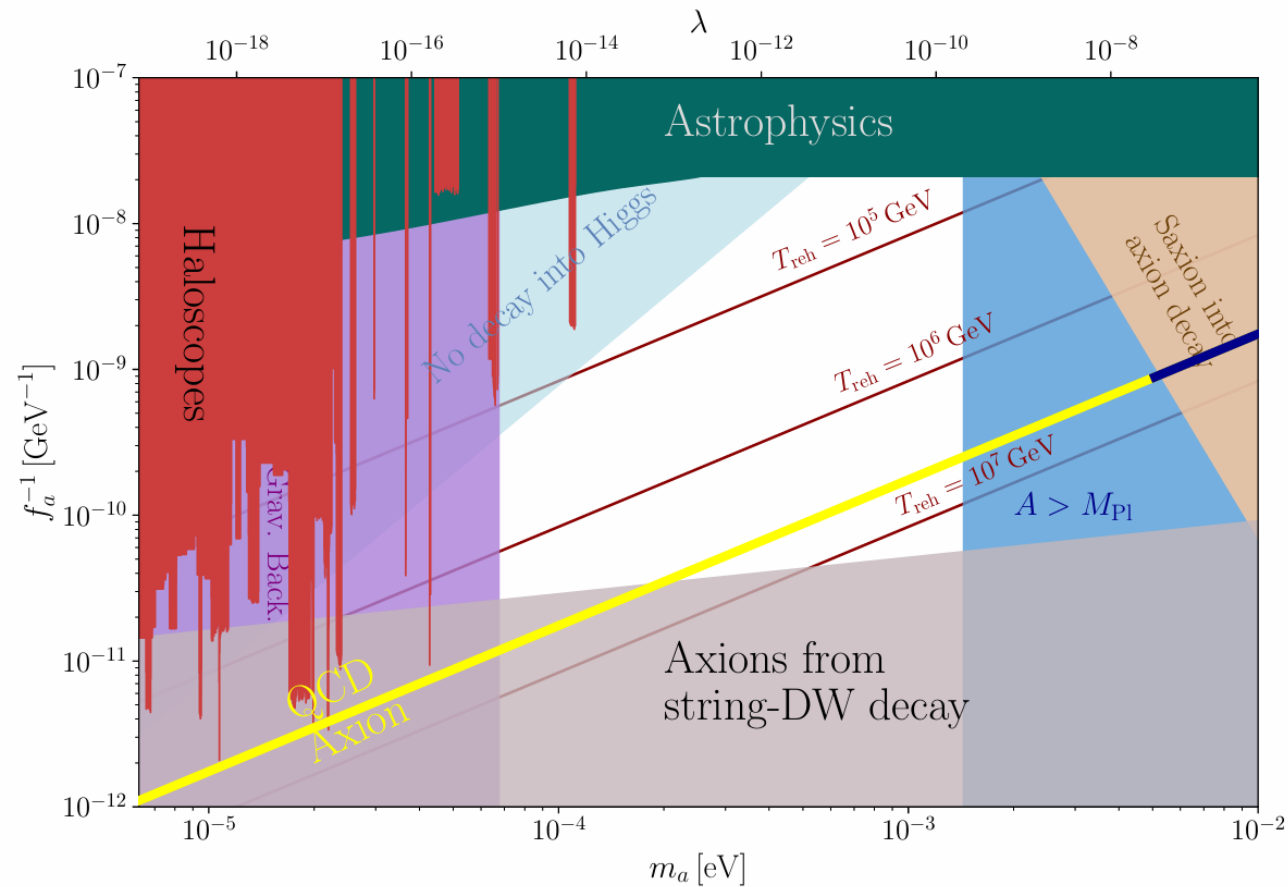
# Axion Dark Matter Abundance

If the radial mode is also responsible for reheating





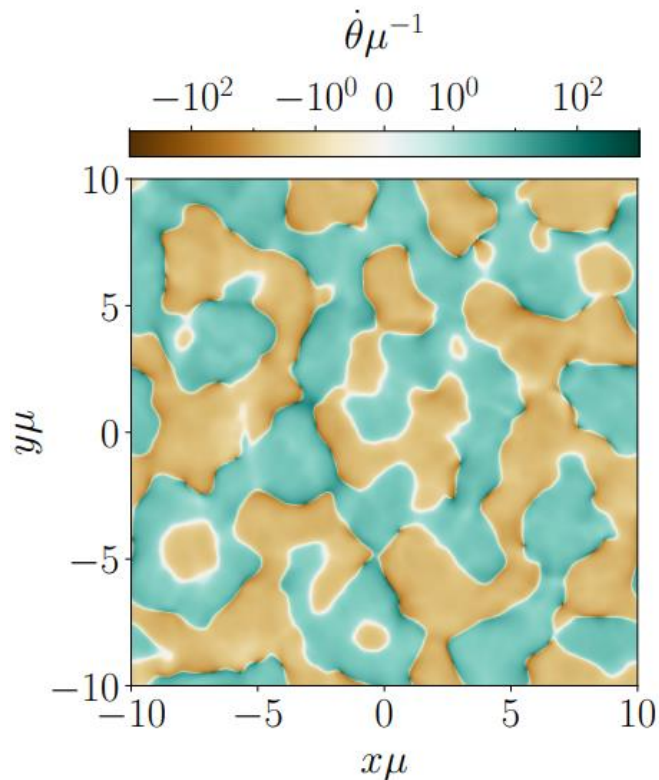
*Thank you!*



# Backup Slide

Because spontaneous symmetry breaking happens **after inflation**, the axion field takes on **different values** in separate regions of the universe.

$$\dot{\theta}_{\max} = \beta \sin(n\theta) \frac{AS_{\max}^{n-2}}{M_{\text{Pl}}^{n-3} H_{\max}}$$



Once the kick is applied, the universe becomes **divided into multiple domains**, each carrying a U(1) charge of either **positive** or **negative** sign.

