

Hey! I know you are all tired, BUT

I am Bai

I am Bai

From Weizmann Institute

Phenomenologies associate with

Ultralight Field Backgrounds

- Scalars,
- Vectors...

- Dark Matter
- or non-dark Matter

For real...



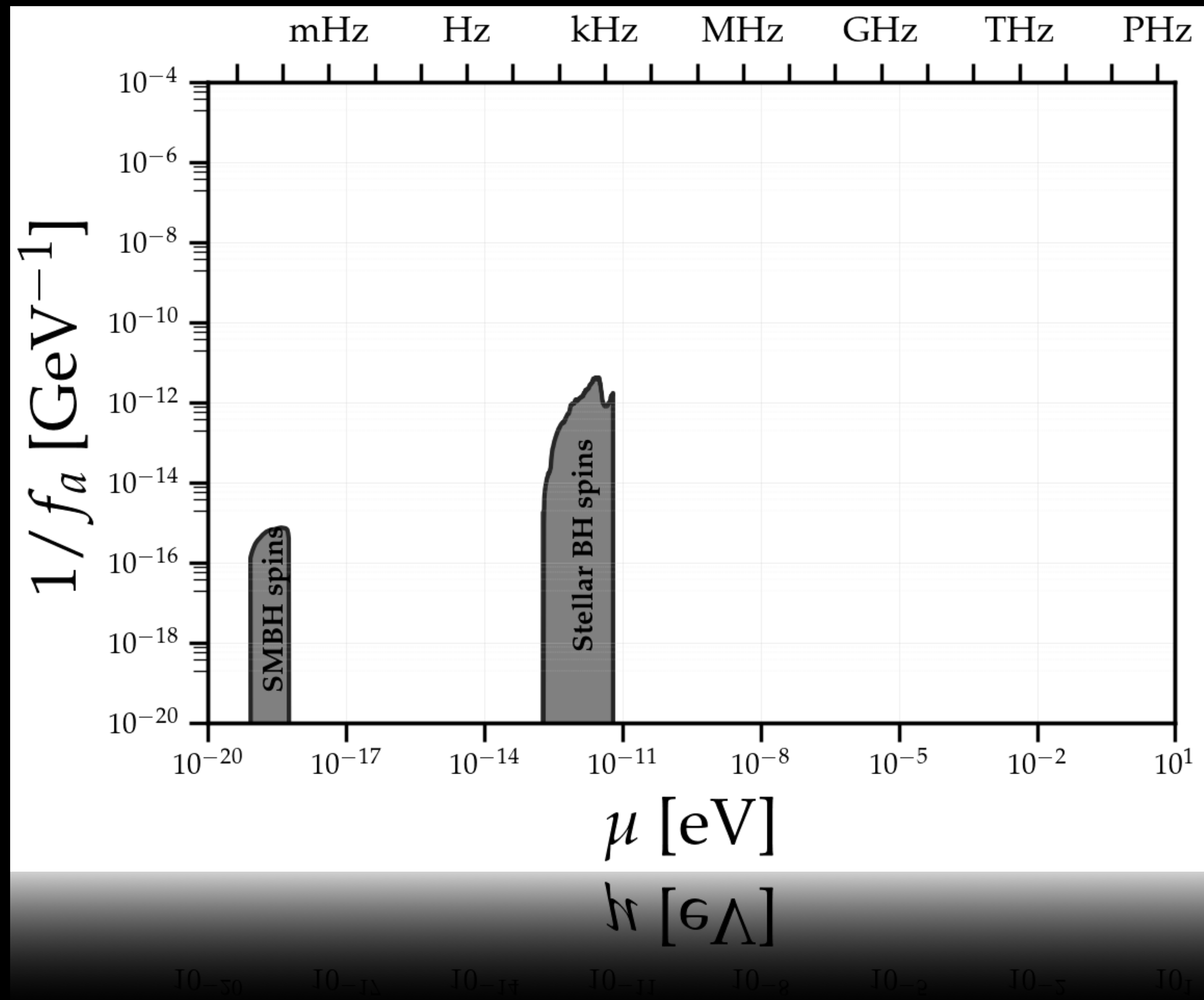


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Superradiance + Spectroscopic Enhanced Phenomenology

Zhaoyu Bai @ Cargese

2025.07.30



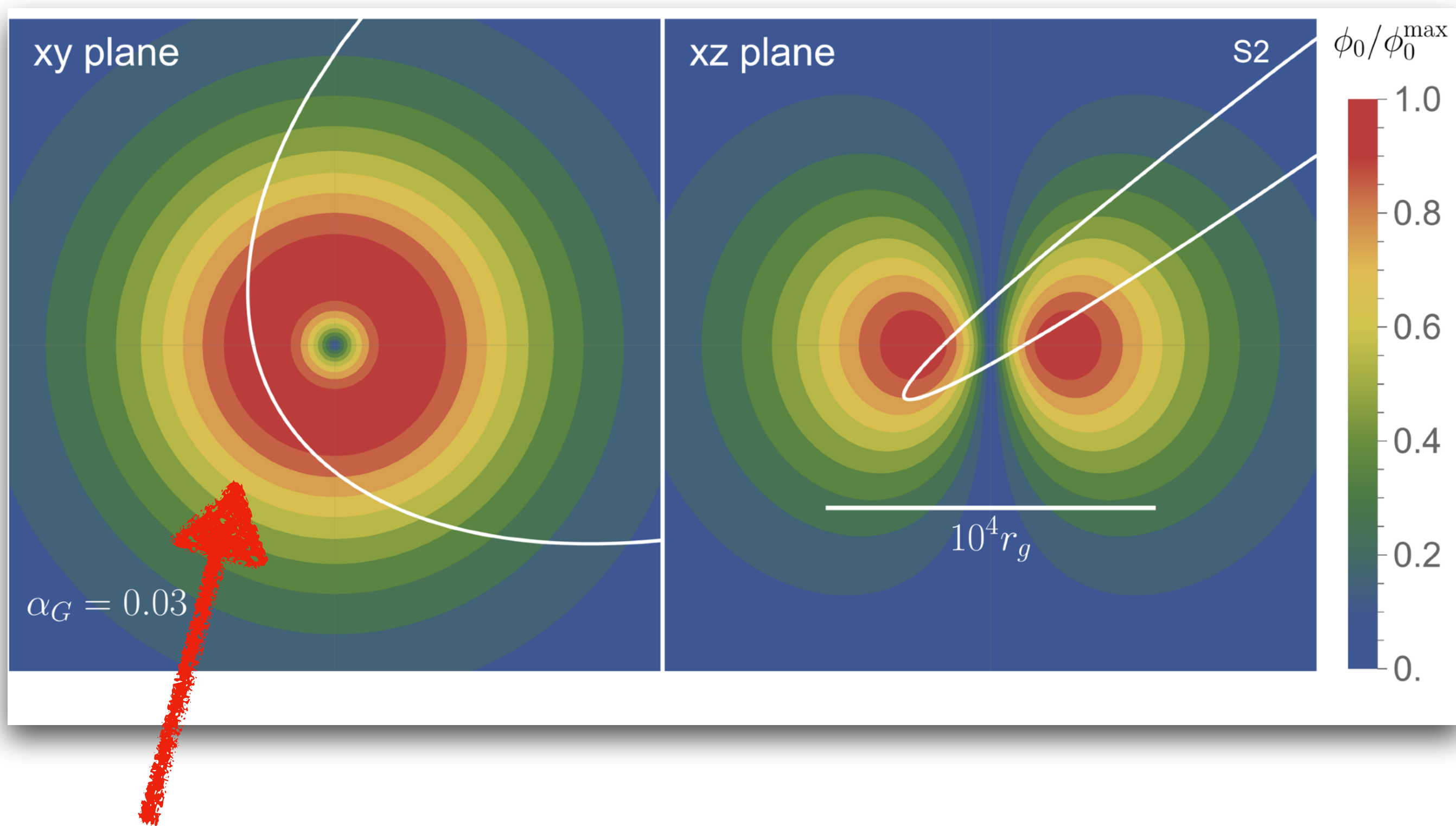
Keep in mind...

**Superradiance forms without the
assumption of dark matter.**

Keep in mind...

And the Amplitude is Huge:

$$\phi_0^{\max} \approx 0.07 \alpha_G^2 M_{\text{pl}} \sqrt{\frac{M_{\text{cloud}}}{M_{\text{BH}}}} \approx 0.5 \alpha_G f_\phi$$



Superradiance ← Scalars backgrounds — Axions/ALP/Generic
Ultralight particles

S-Star Spectroscopy

Existing S2 Star Observation — To Future Prospect

- **Cadence** (exposure time): 10 or 15 mins.
- Over a period of **25 days**, Daily averages of **3 hours** of observations.
- Most Relevant Observation Period (2017/2018): **Periastron**
- The median uncertainty for Hydrogen $\frac{\delta\lambda_j}{\lambda_j}$ is 10^{-4} 10km/s uncertainty of RV.
 $v_{\text{perihelion}} \approx 7,600 \text{ km/s}$
- Future: $\frac{\delta\lambda}{\lambda} \approx 3 \times 10^{-6}$ HISPEC (anticipated in 2026) and MODHIS (2030)
- Future Radial Velocity Precision (with multi-band analysis): $\sim 30 \text{ cm/s}$.

* We primarily focused on the Hydrogen line for S2 Star analysis

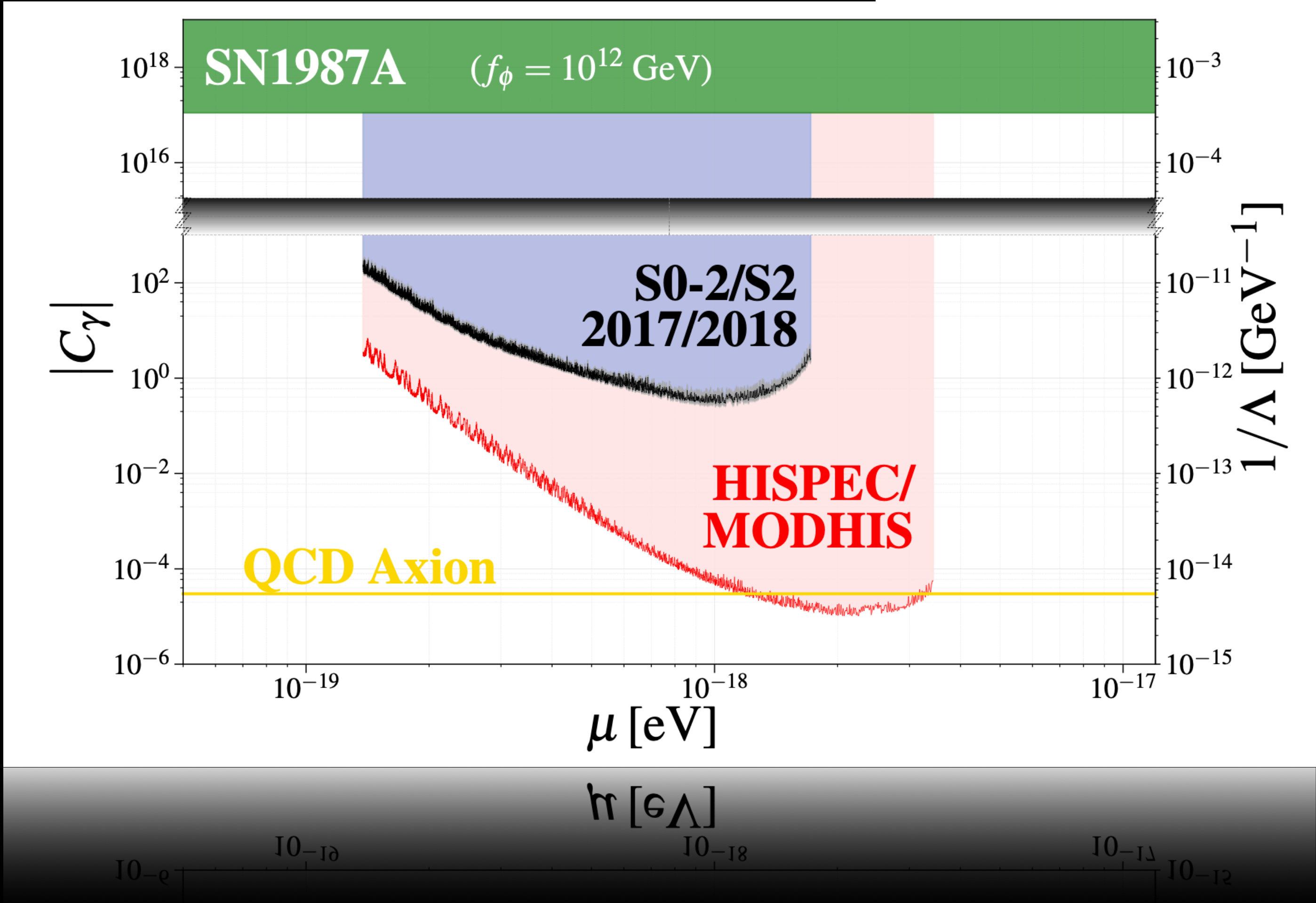
Projected sensitivity on the quadratic scalar–photon coupling

$$\frac{C_\gamma}{4} \frac{\phi^2}{f_\phi^2} F_{\mu\nu} F^{\mu\nu}$$

As a benchmark

$$|C_\gamma| \approx 3 \times 10^{-5}$$

$$\Lambda \equiv f_\phi/|C_\gamma|^{1/2}$$





Thanks

More information (Arxiv operates 24/7):
Call 2507.07482



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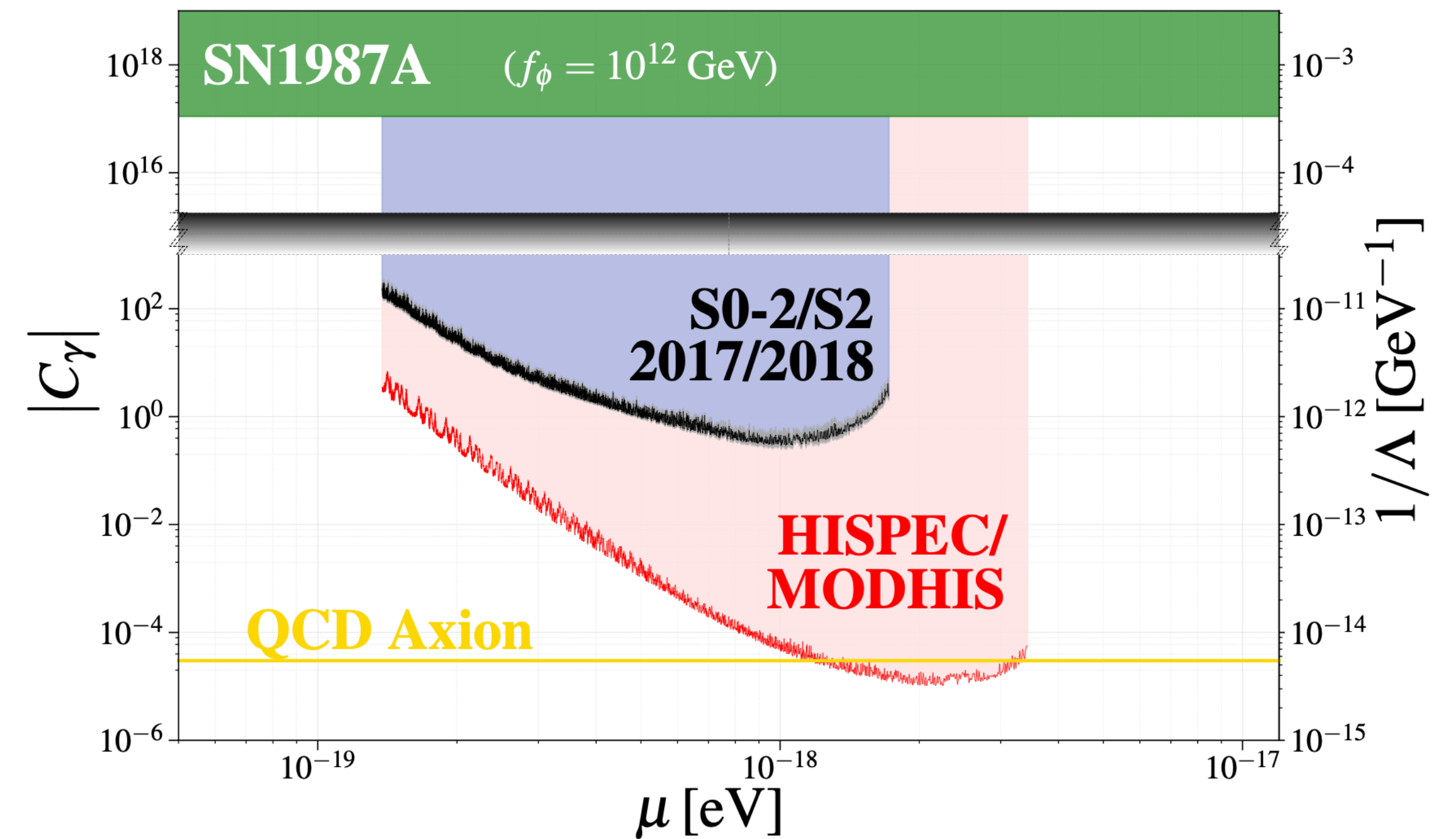
2025.08.02

Results

- **Current data:** Gemini/NIFS over 25 days in 2017–2018.
- **Future prospects:** HISPEC/MODHIS projections incorporating:
 - 10^3 -fold better spectroscopic resolution
 - 1/2 observing cadence
 - 10-fold longer total observation days
 - A late-type star with a semi-major axis 10% that of S2
- Constraint **independent** on f_ϕ

$$\phi_0^{\max} \approx 0.07 \alpha_G^2 M_{\text{Pl}} \sqrt{\frac{M_{\text{cloud}}}{M_{\text{BH}}}} \approx 0.5 \alpha_G f_\phi$$

$$\frac{C_\gamma}{4} \frac{\phi^2}{f_\phi^2} F_{\mu\nu} F^{\mu\nu}$$



Projected constraints on the quadratic scalar–photon coupling

$$\Lambda \equiv f_\phi / |C_\gamma|^{1/2} \quad |C_\gamma| \approx 3 \times 10^{-5}$$

Scalar - Photon Quadratic Couplings

Ultralight Bosons

- **Axions/ALPs:** Nonperturbative QCD or explicit UV effects induce
- **QCD axion:** loop calculations yield $|C_\gamma| \sim 3 \times 10^{-5}$ (UV dependent)
- **Generic Scalar Particles:** With Z2 symmetry.
- **Quadratic Coupling Effect:**
 - Operator $\frac{C_\gamma}{4} \frac{\phi^2}{f_\phi^2} F_{\mu\nu} F^{\mu\nu}$ modifies the lagrangian effectively $-\frac{1}{4} \left[1 - \frac{C_\gamma \phi^2}{f_\phi^2} \right] F_{\mu\nu} F^{\mu\nu}$
 - Fine structure constant $\alpha_{\text{eff}} \approx \alpha \left(1 + \frac{C_\gamma \phi^2}{f_\phi^2} \right)$

S-Star Spectroscopy Summary

- An exposure is around 10 mins.

$$\frac{\delta\alpha_{\text{EM}}}{\alpha_{\text{EM}}} \simeq C_\gamma \frac{\phi^2}{f_\phi^2}$$

- Modulation at Periastron
- Spectral line shift:

$$\frac{\delta\lambda_j}{\lambda_j} \approx -k_{\alpha,j} \frac{\delta\alpha_{\text{EM}}}{\alpha_{\text{EM}}} \approx \frac{\delta v}{c}$$

*The logic to the Soliton Core DM is the same

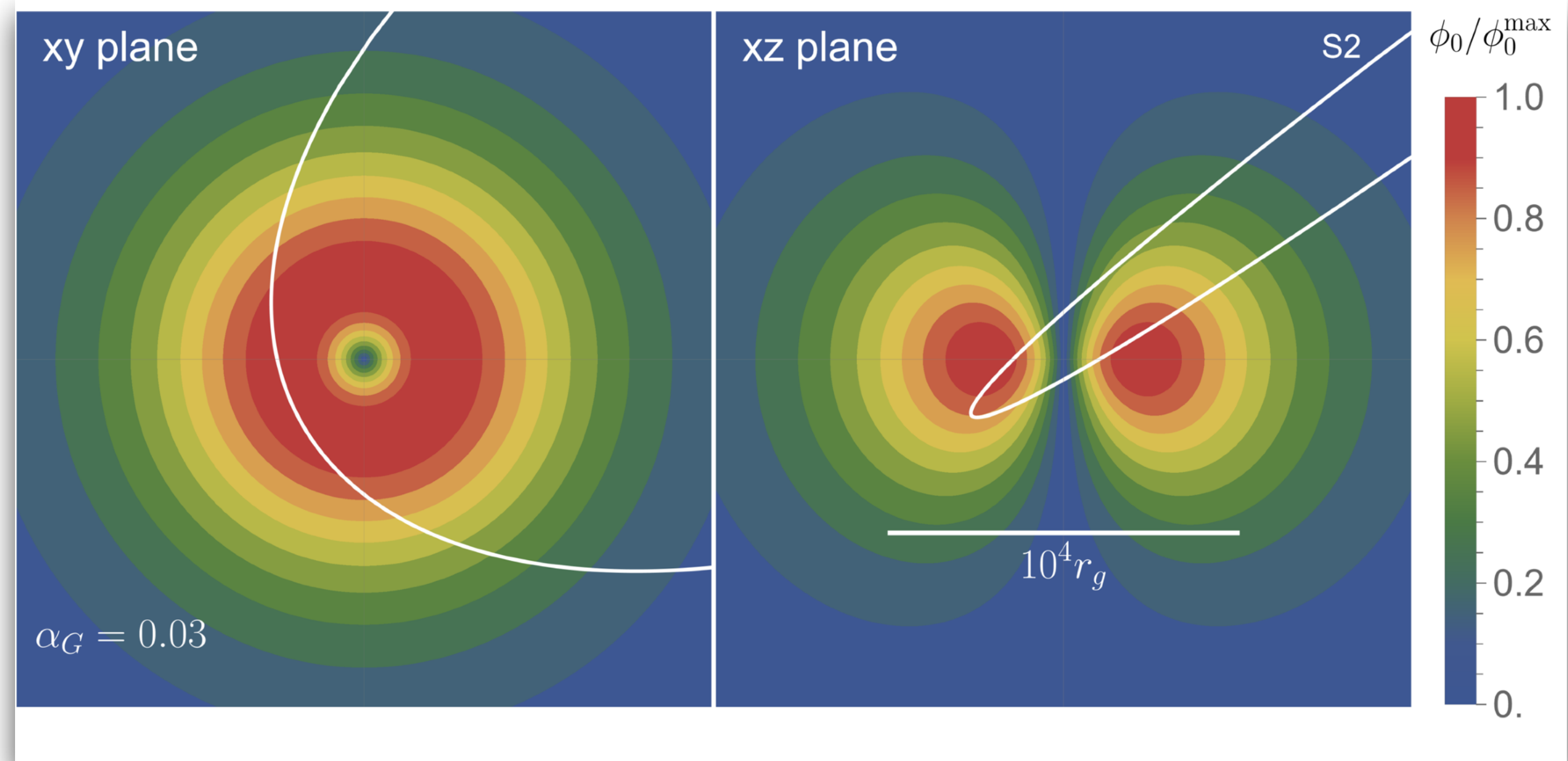


FIG. 1: An example of the normalized axion/scalar cloud density, ϕ_0/ϕ_0^{\max} , for $\alpha_G = 0.03$ on the xy -plane (left) and xz -plane (right) in the BH frame, where the BH spin axis is aligned with the $+z$ direction. The projected orbit of the S2 star is also shown on both planes. The transformation of the S2 orbit into the BH frame adopts fiducial values of $i_{\text{BH}} = 155^\circ$ and $\Omega_{\text{BH}} = 177^\circ$ [66].

