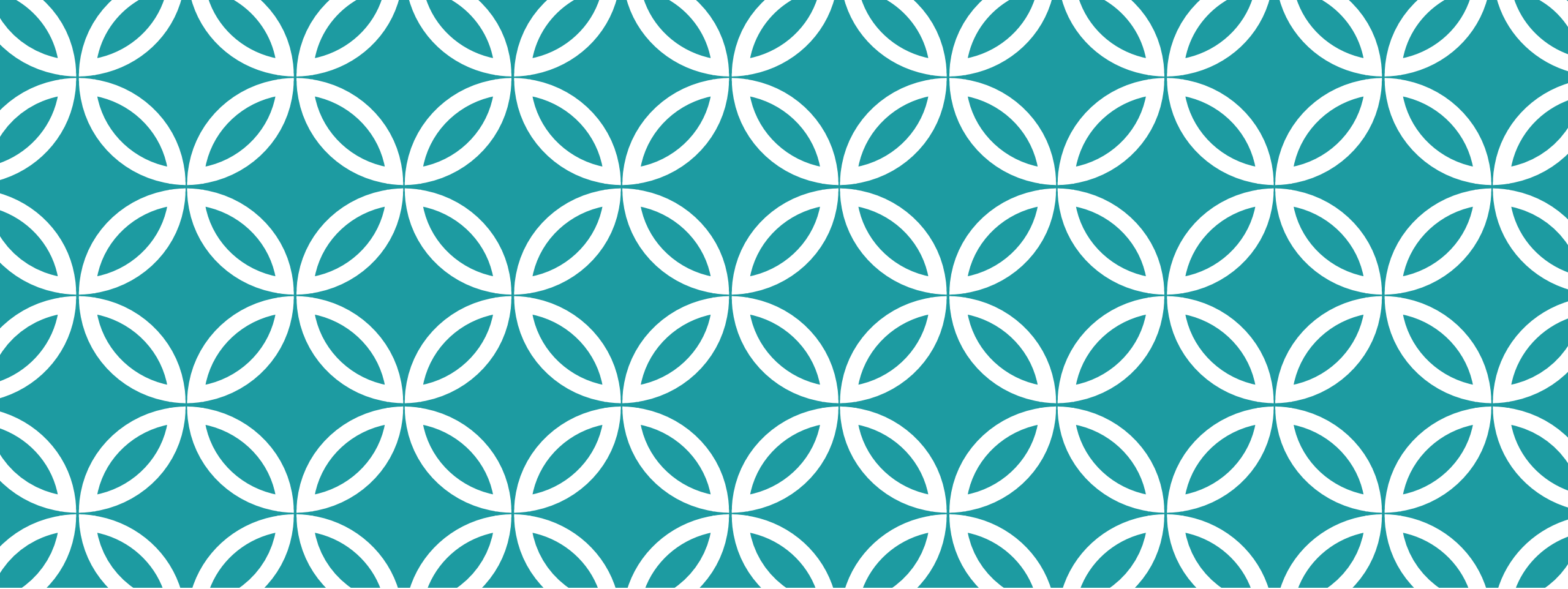


TB22 ANALYSIS

Michal Elad
26 / 09 / 24

OVERVIEW

- Telescope resolution
- Hough Transform (HT) uncertainty
- Calorimetry – prelim

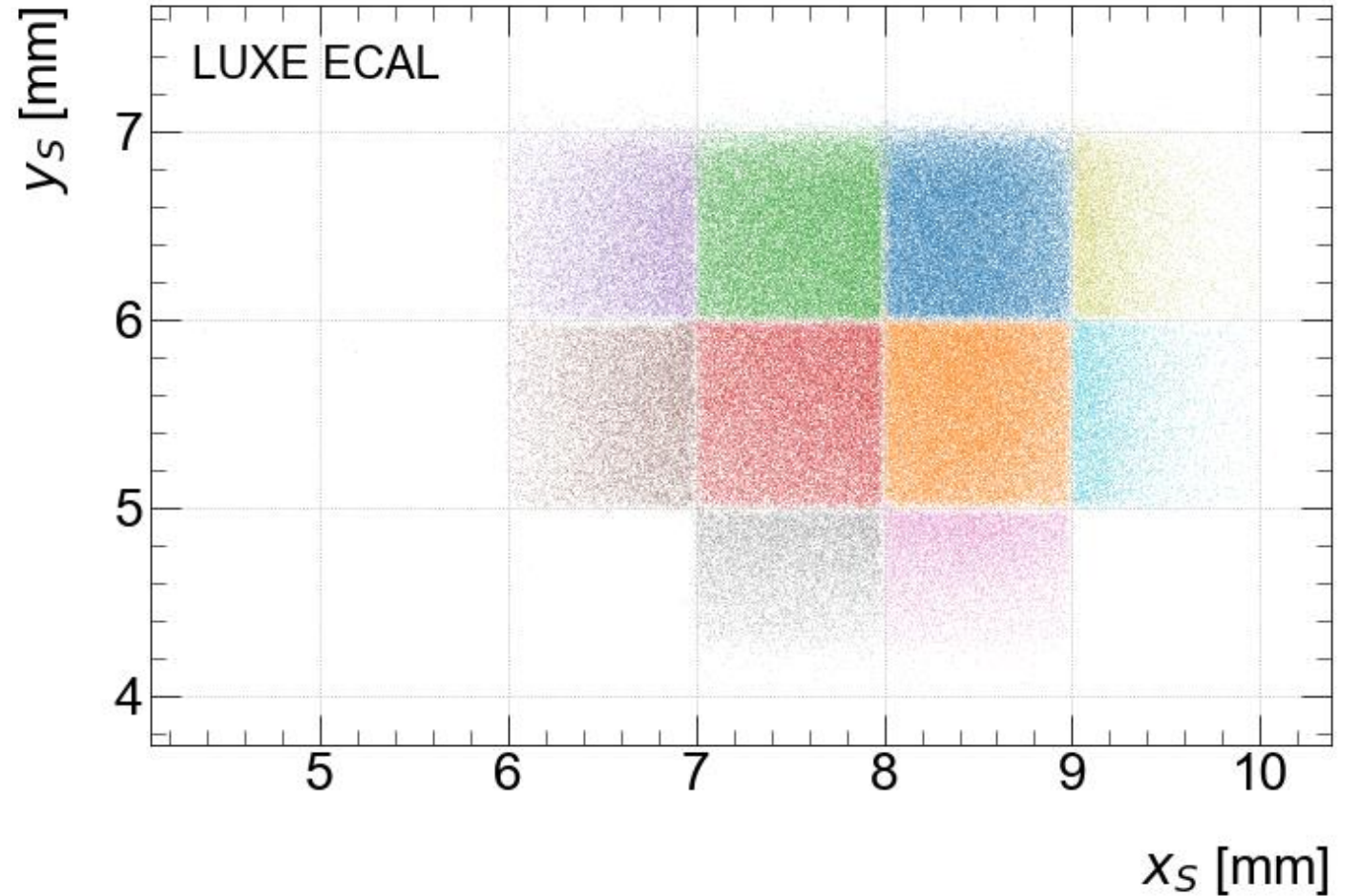


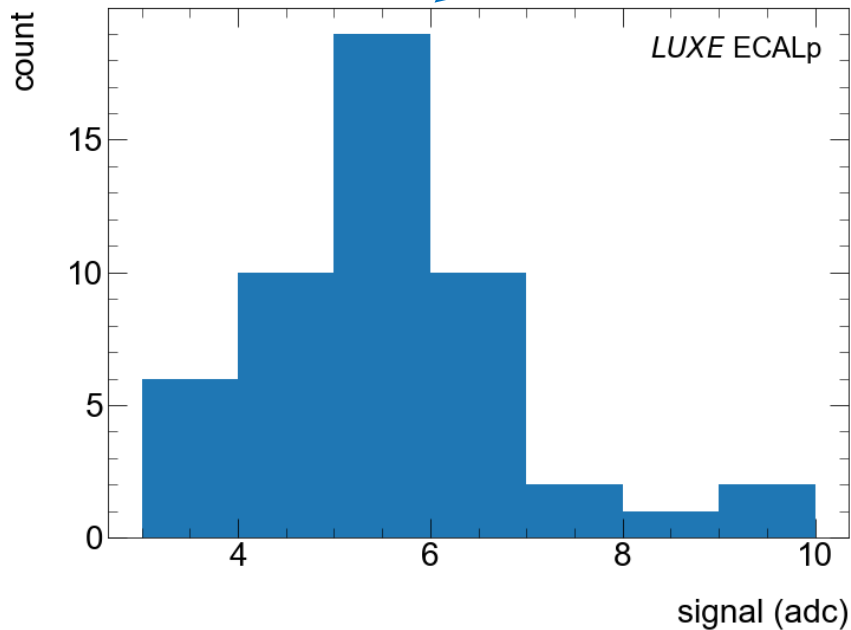
TELESCOPE RESOLUTION

A second attempt

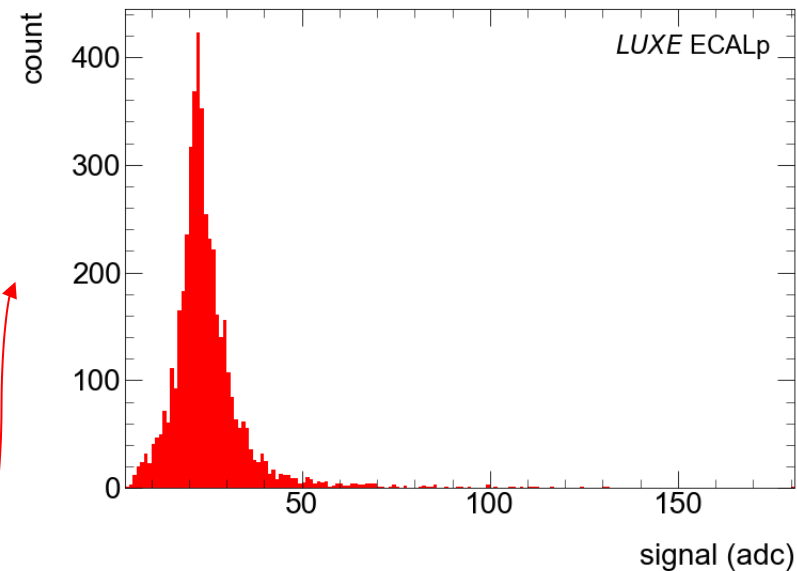
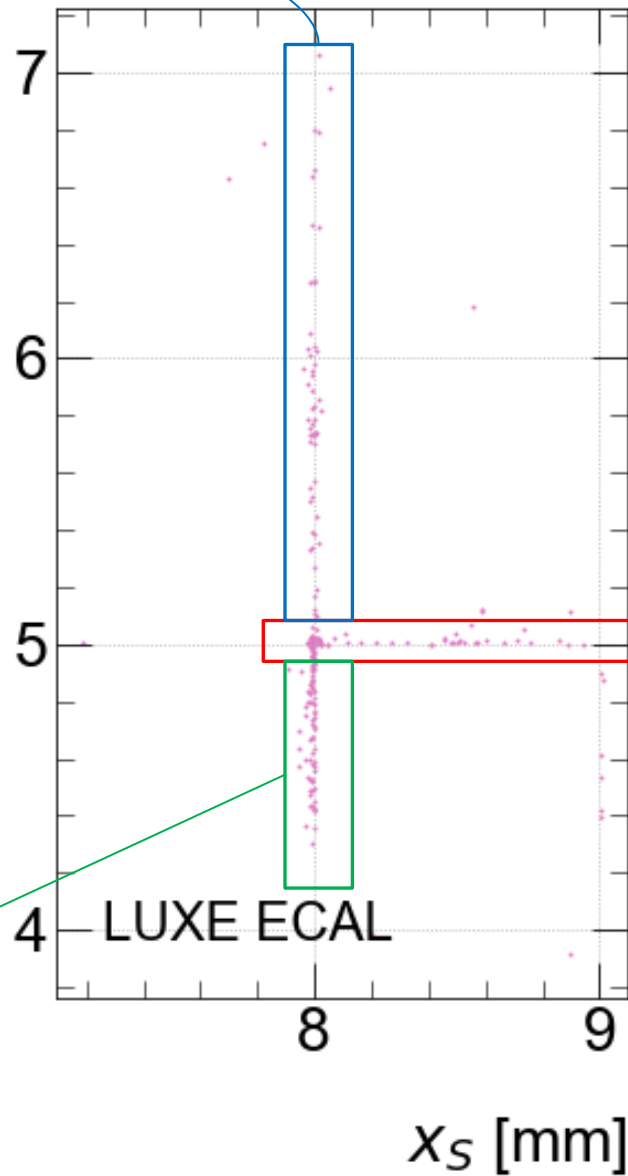
GAAS HITS OUTSIDE PAD RECAP

Events of single track & single pad hit.
Pad-based color assignment.



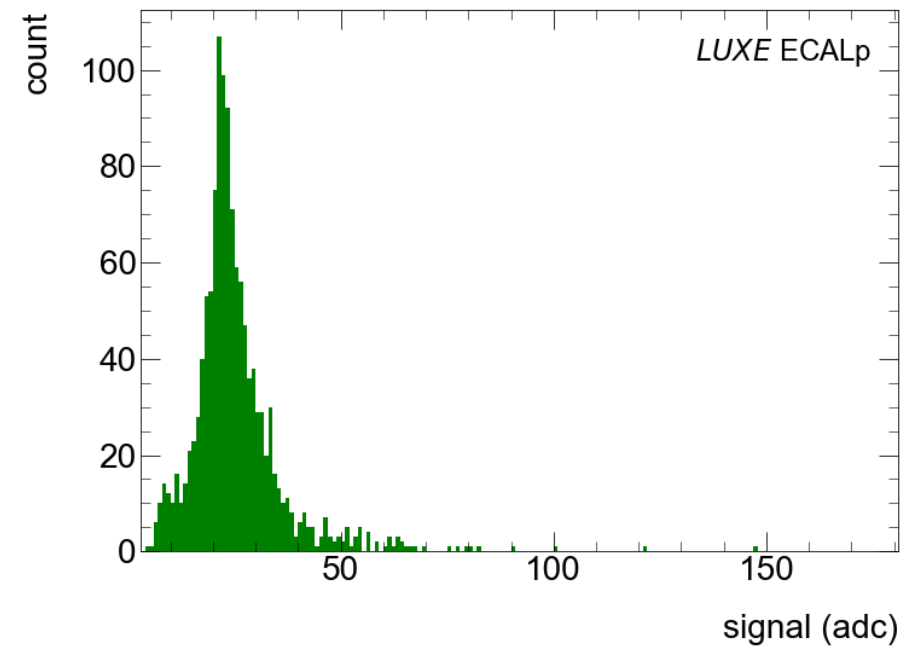


y_S [mm]



- Selected events assigned to pad (8, 4) that fell outside the pad area after alignment.
- Plotted the ADC count histograms of each marked area.

→ ADC count on traces ~ 5

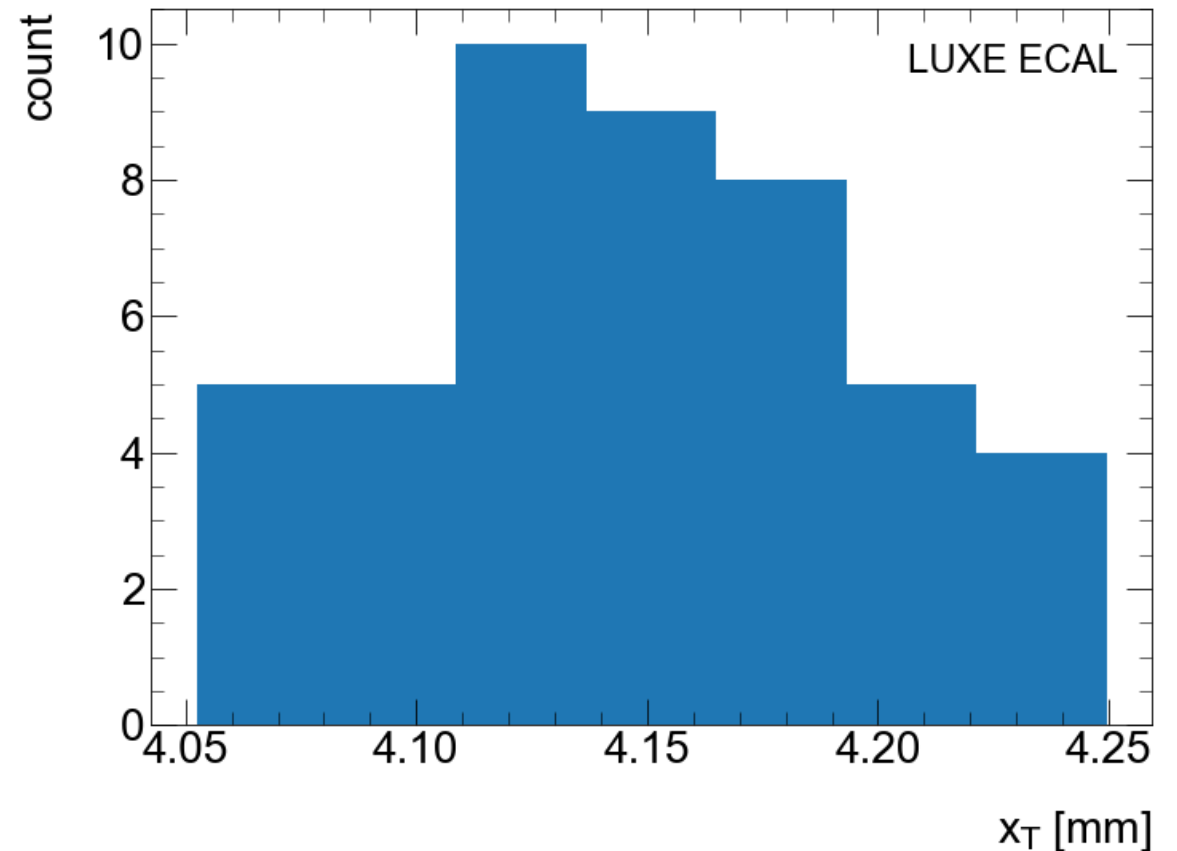


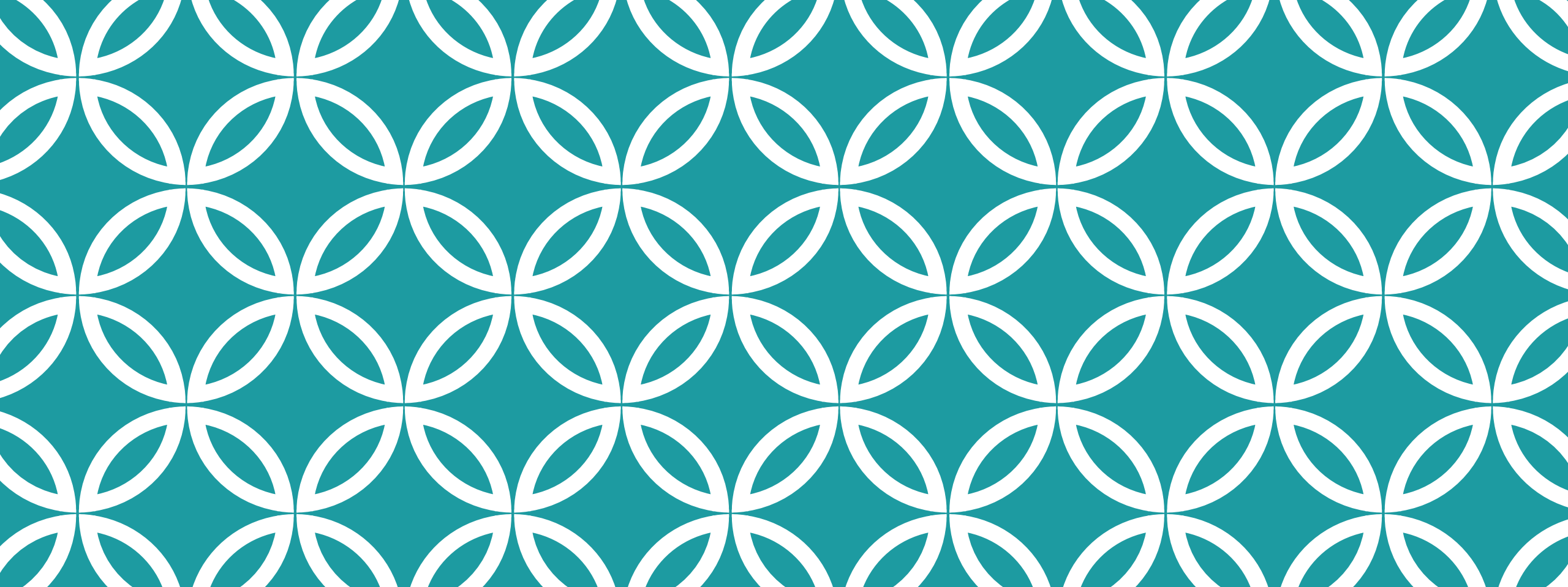
TELESCOPE RESOLUTION

Idea – use hits on the traces to get the telescope resolution.

- Choose events with (x, y) position **far above** the assigned pad.
- Plot a histogram of their x value given by the telescope.
- The standard deviation is the telescope resolution (neglecting the trace width).

The result is ~ 50 microns.

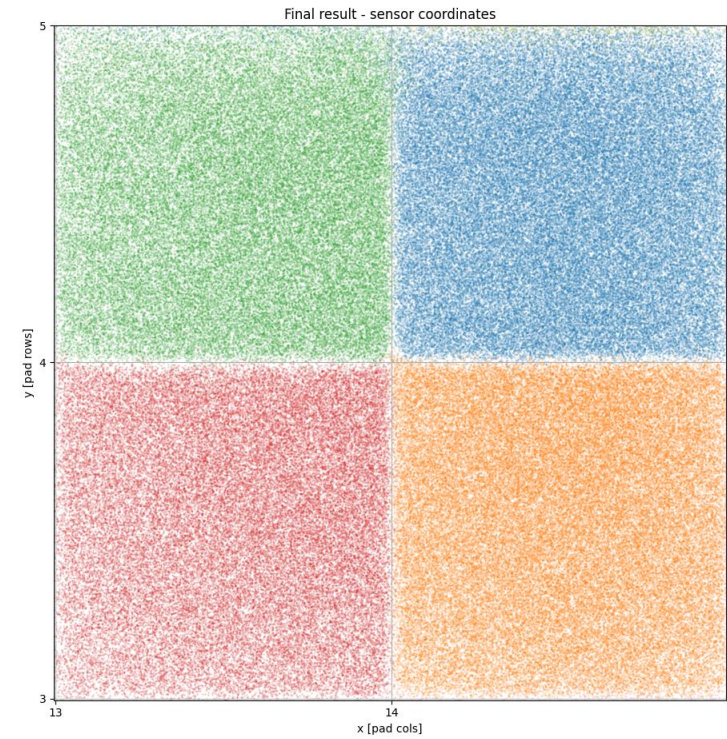
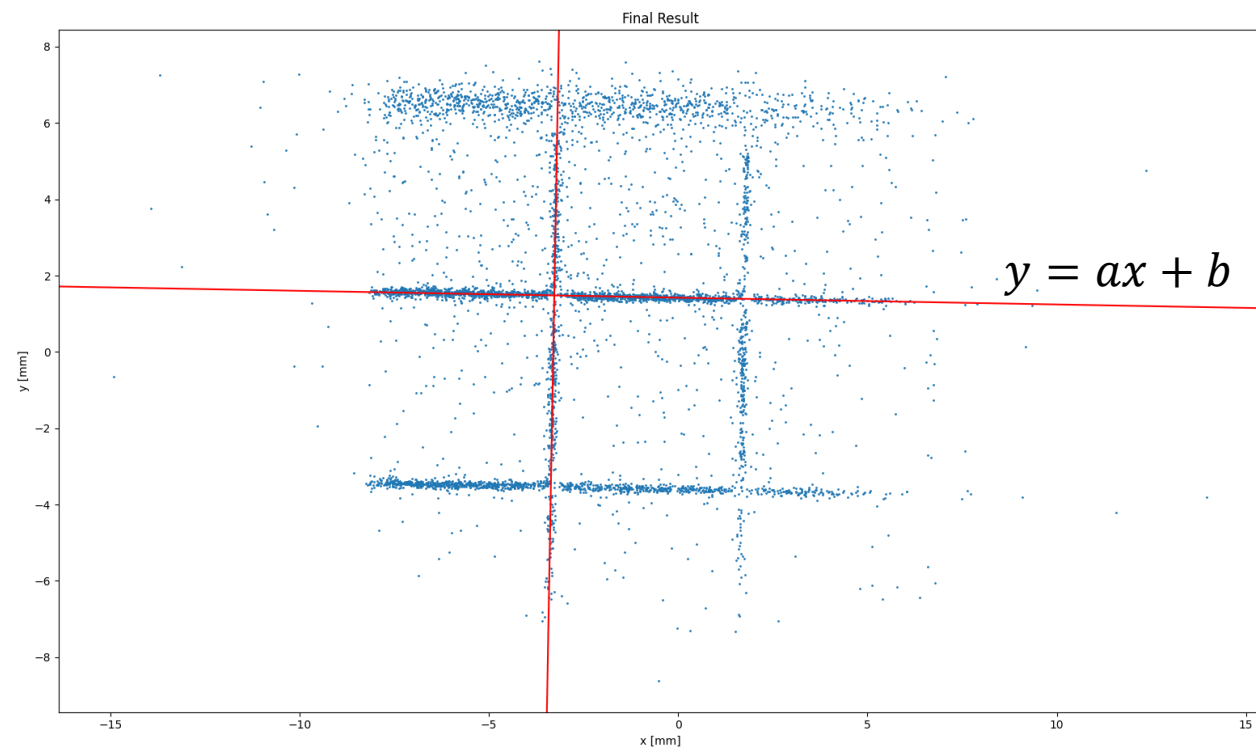




HT UNCERTAINTY

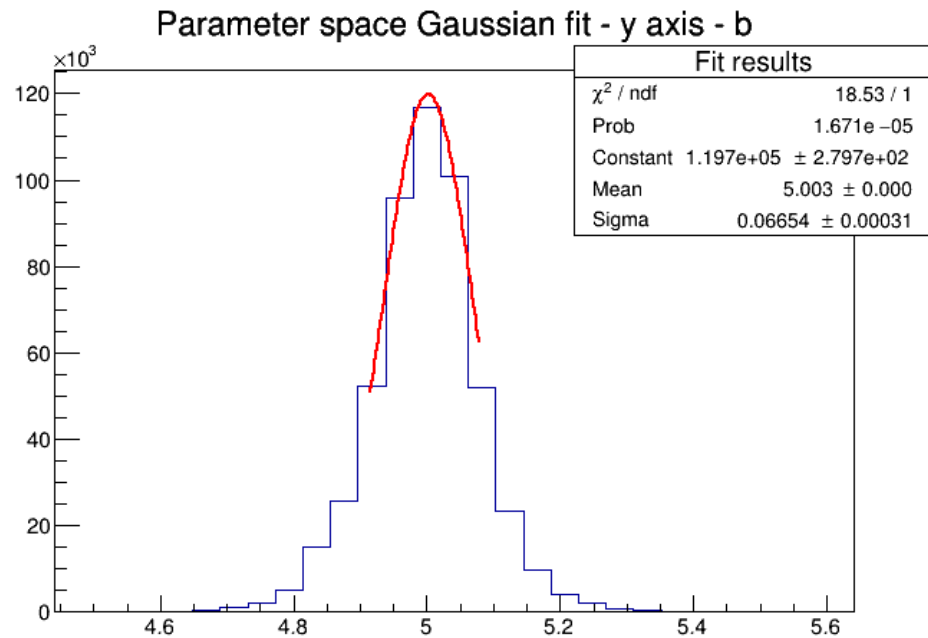
Update & Summary

HT — FIND LINE EQUATION

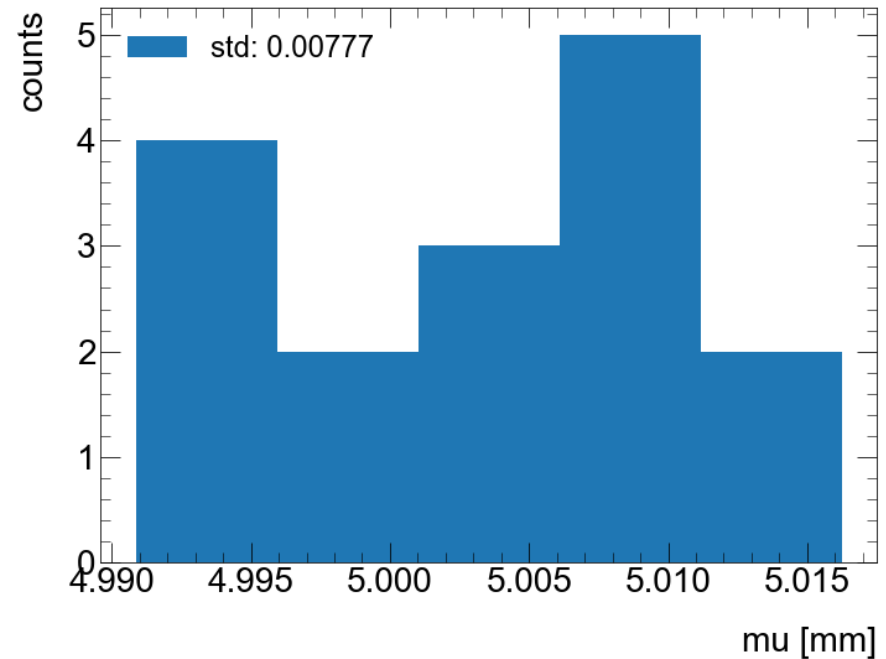


STATISTICAL UNCERTAINTY

Get line parameters distribution
Fit Gaussian → get mpv



Subdivide data to get distribution of mpv
→ std of the mpv distribution is the
algorithm's uncertainty



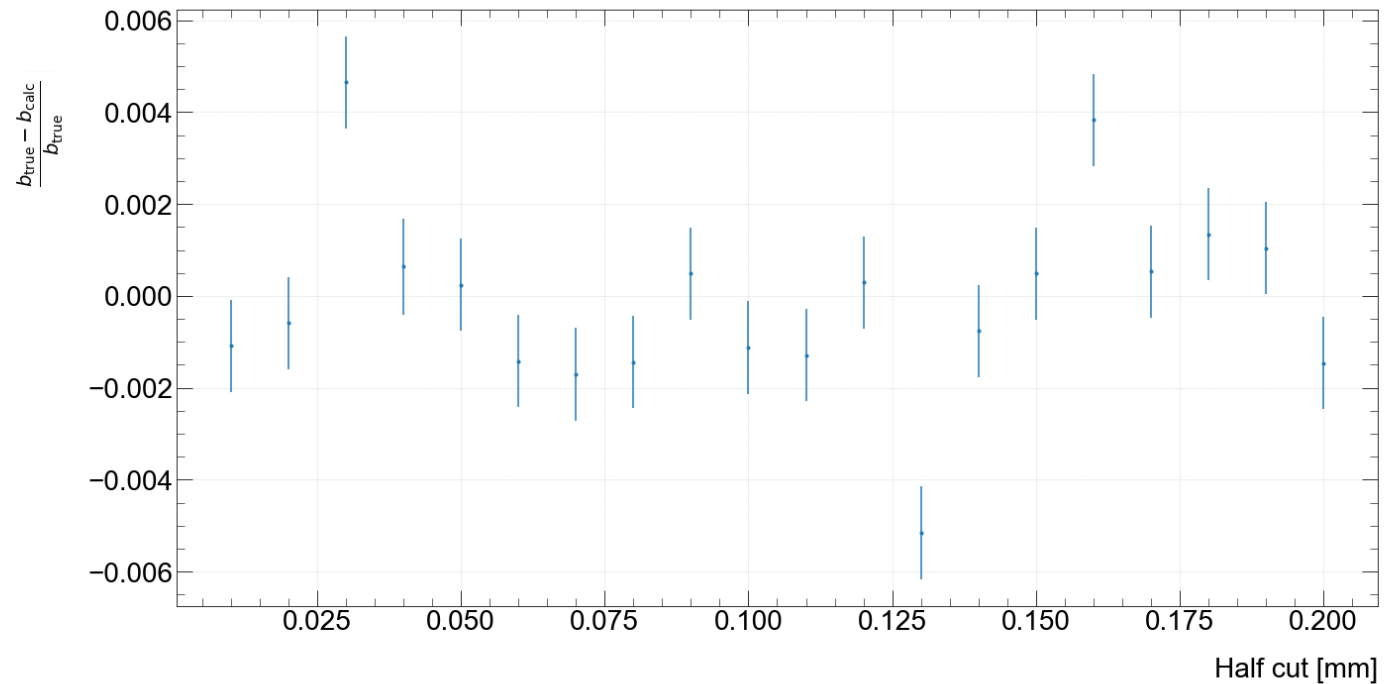
CONTROLLED TEST

- Generated random points from a uniform distribution on a horizontal line.
- Smeared points with σ values of 10, 20, ..., 100 microns.
- Ran the alignment algorithm on the smeared dataset and got the uncertainty.

- In addition, checked the effect of changing the parameters of the algorithm, and statistics.

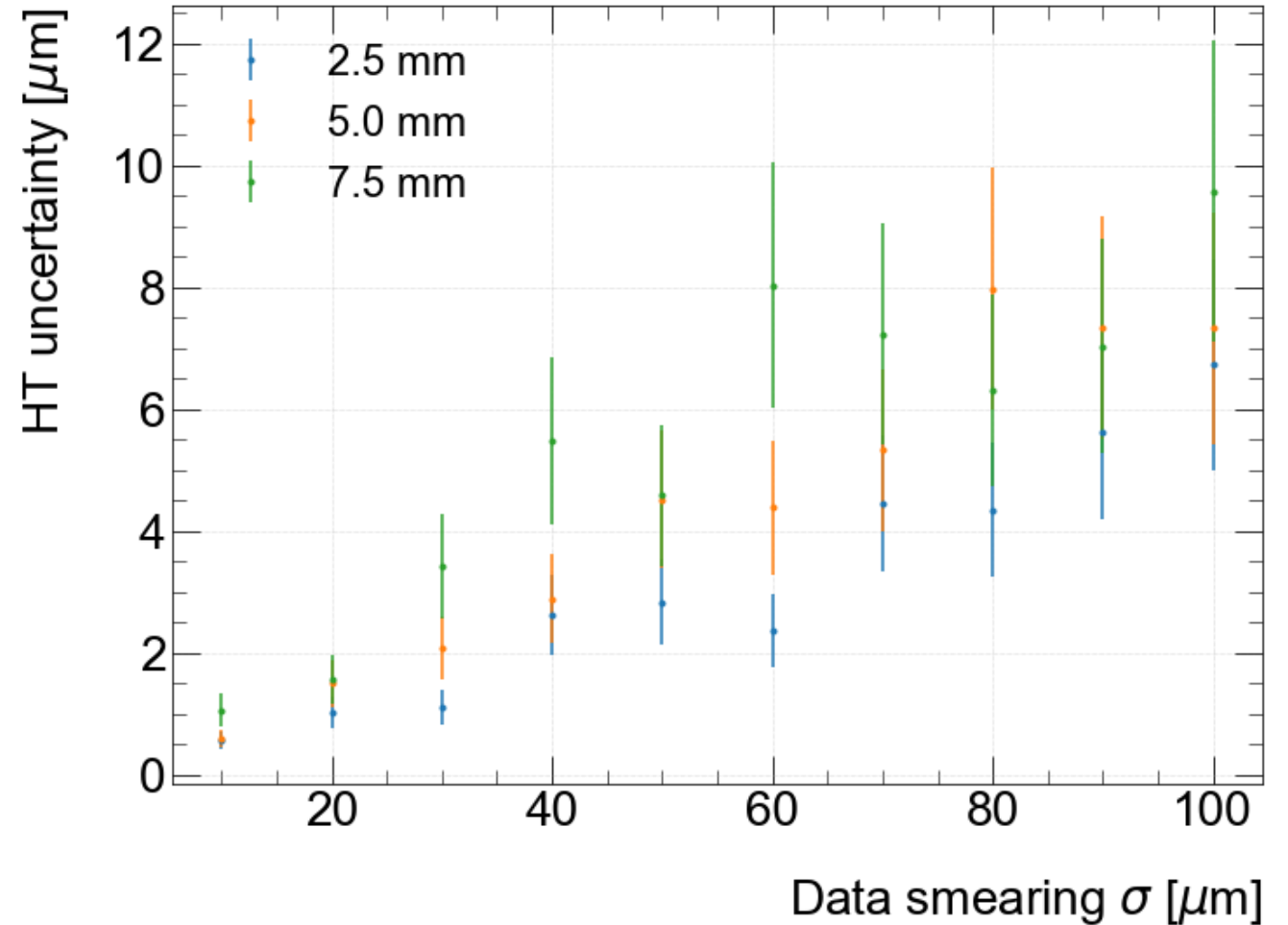
CONTROLLED TEST

Values are nicely distributed around 0



CONTROLLED TEST

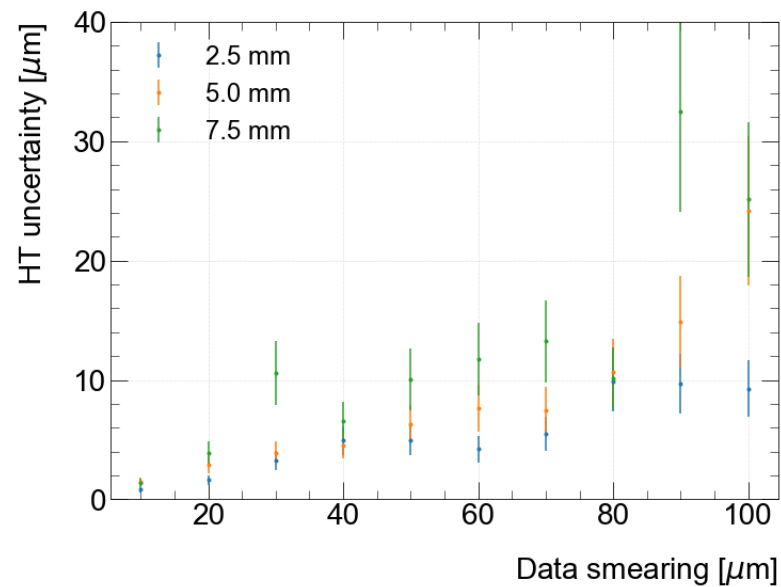
The uncertainty for a given smearing σ slightly varies, but it is always at the level of a few microns and hence negligible.



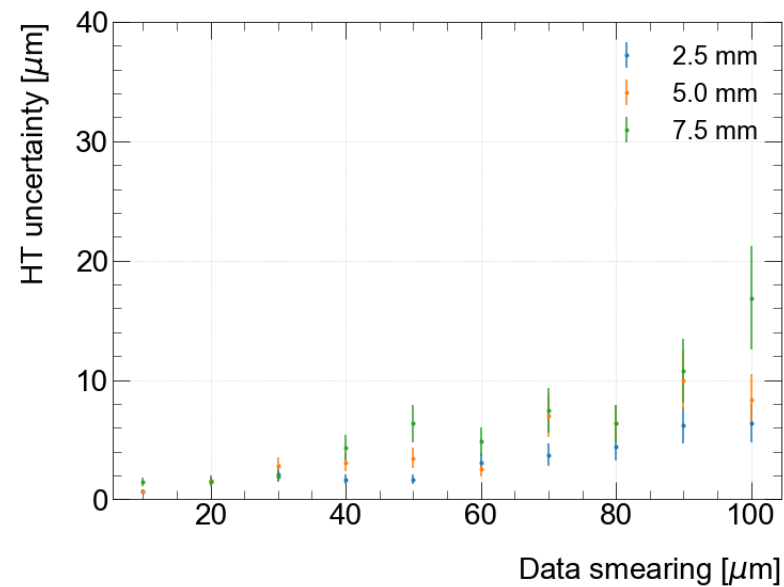
CONTROLLED TEST

Higher statistics give smaller uncertainties, but are still at the order of a few microns

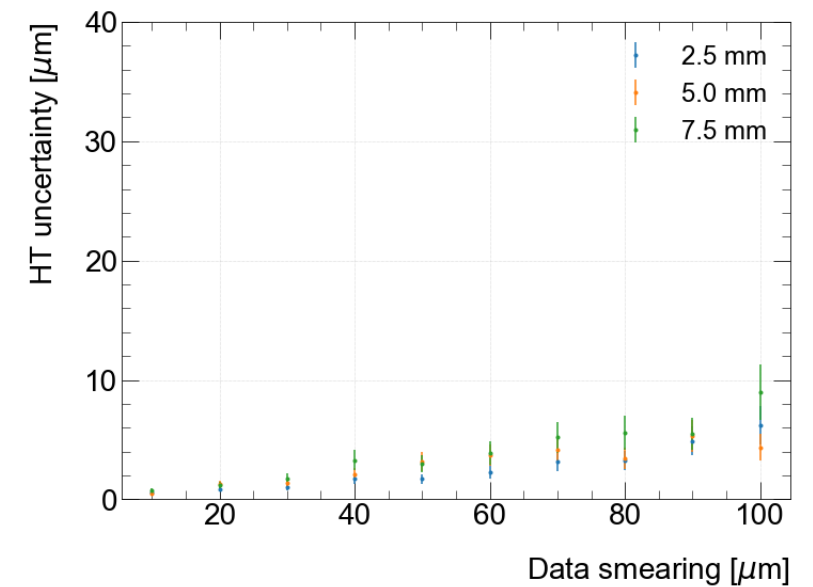
Lower statistics



Mid-range statistics

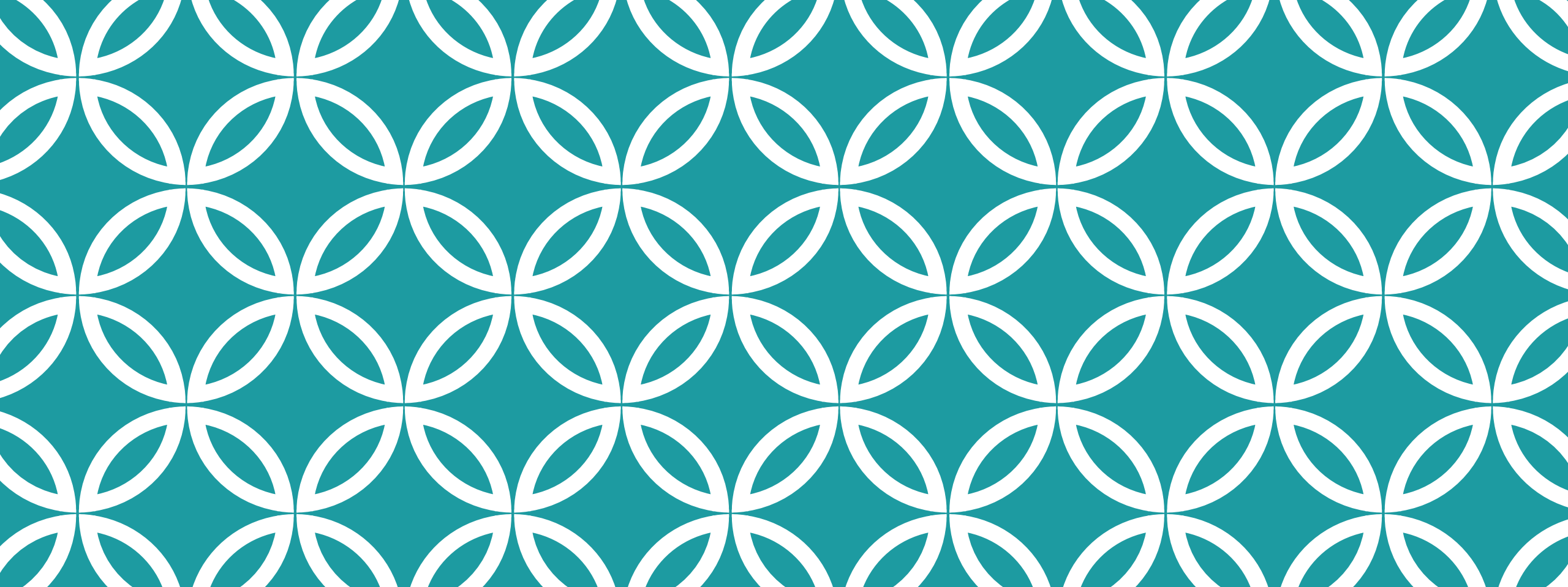


Higher statistics



HT UNCERTAINTY — CONCLUSION

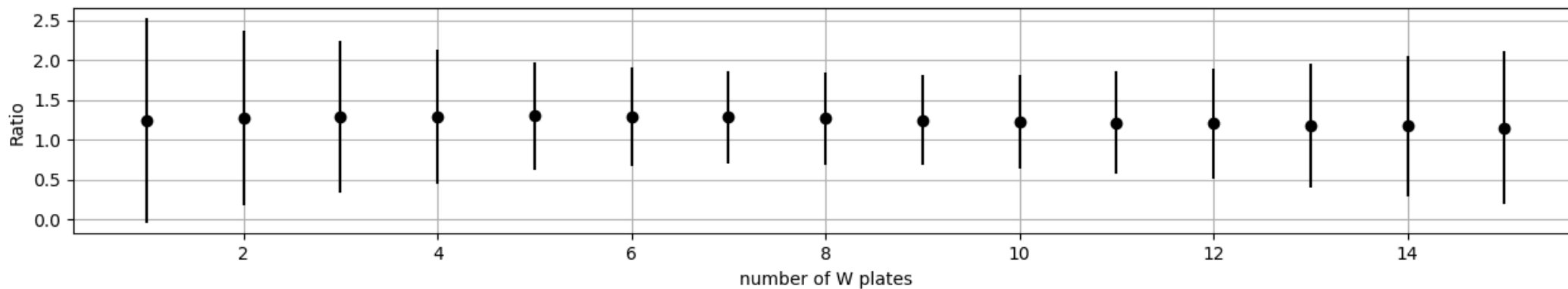
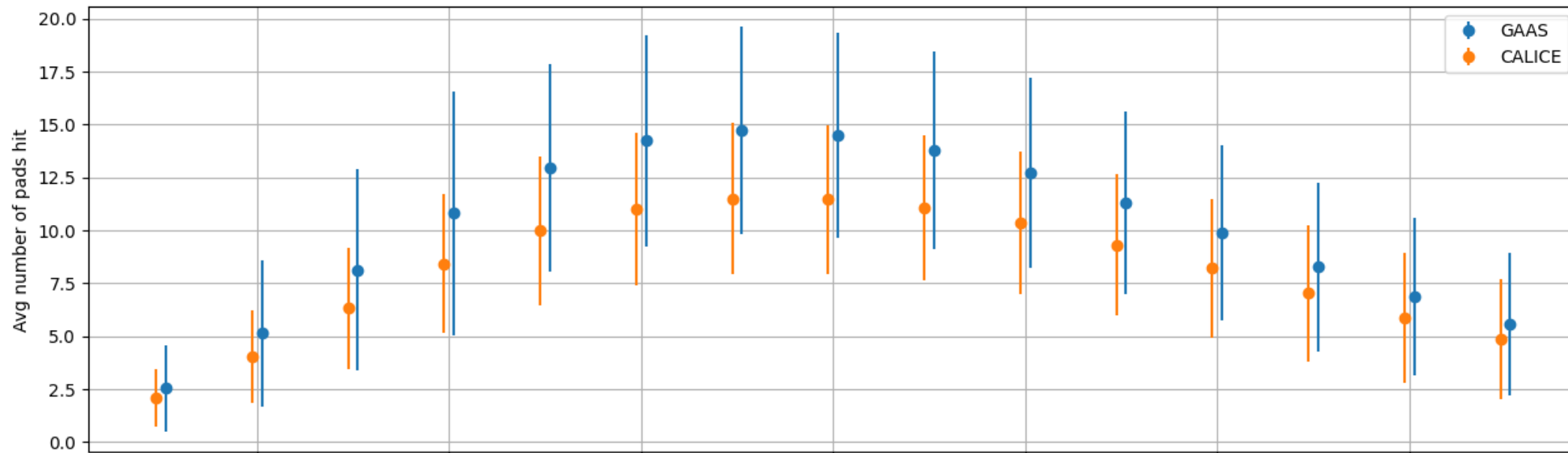
- Statistics is the main variable affecting the algorithm's uncertainty
- The algorithm's uncertainty is at the level of a few microns
- The alignment's uncertainty is dominated by the telescope resolution



CALORIMETRY

Preliminary

NUMBER OF PADS HIT

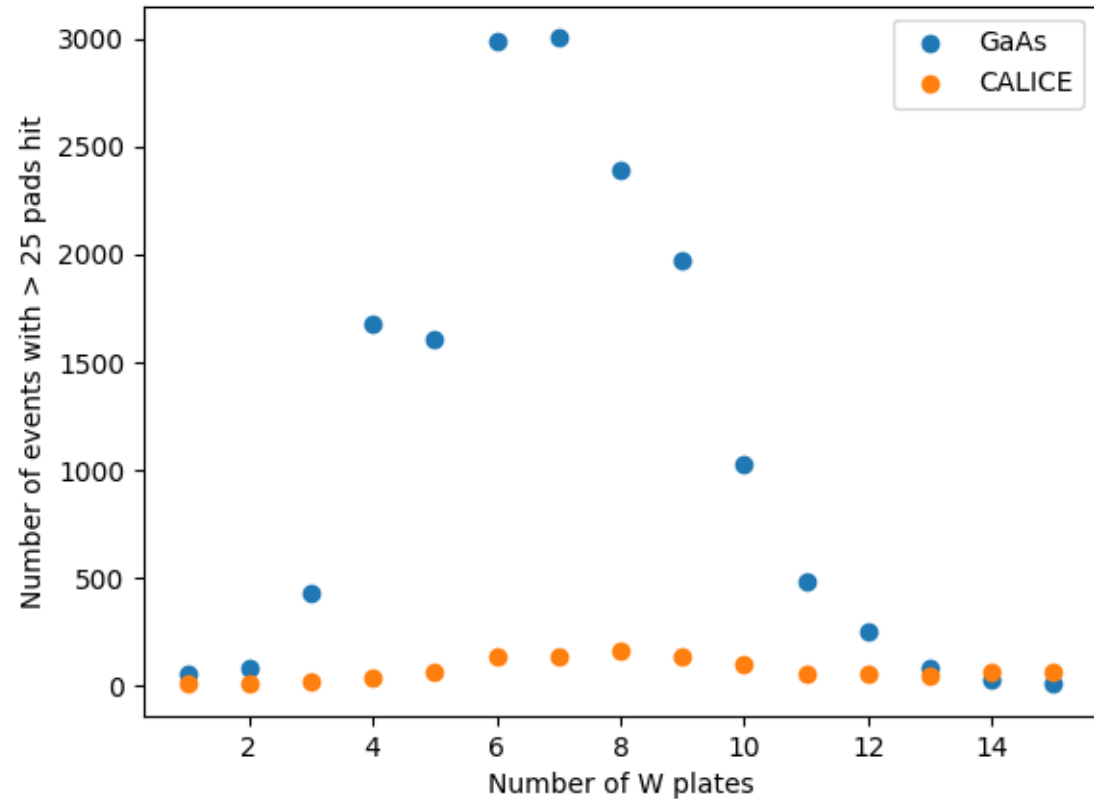


NUMBER OF PADS HIT

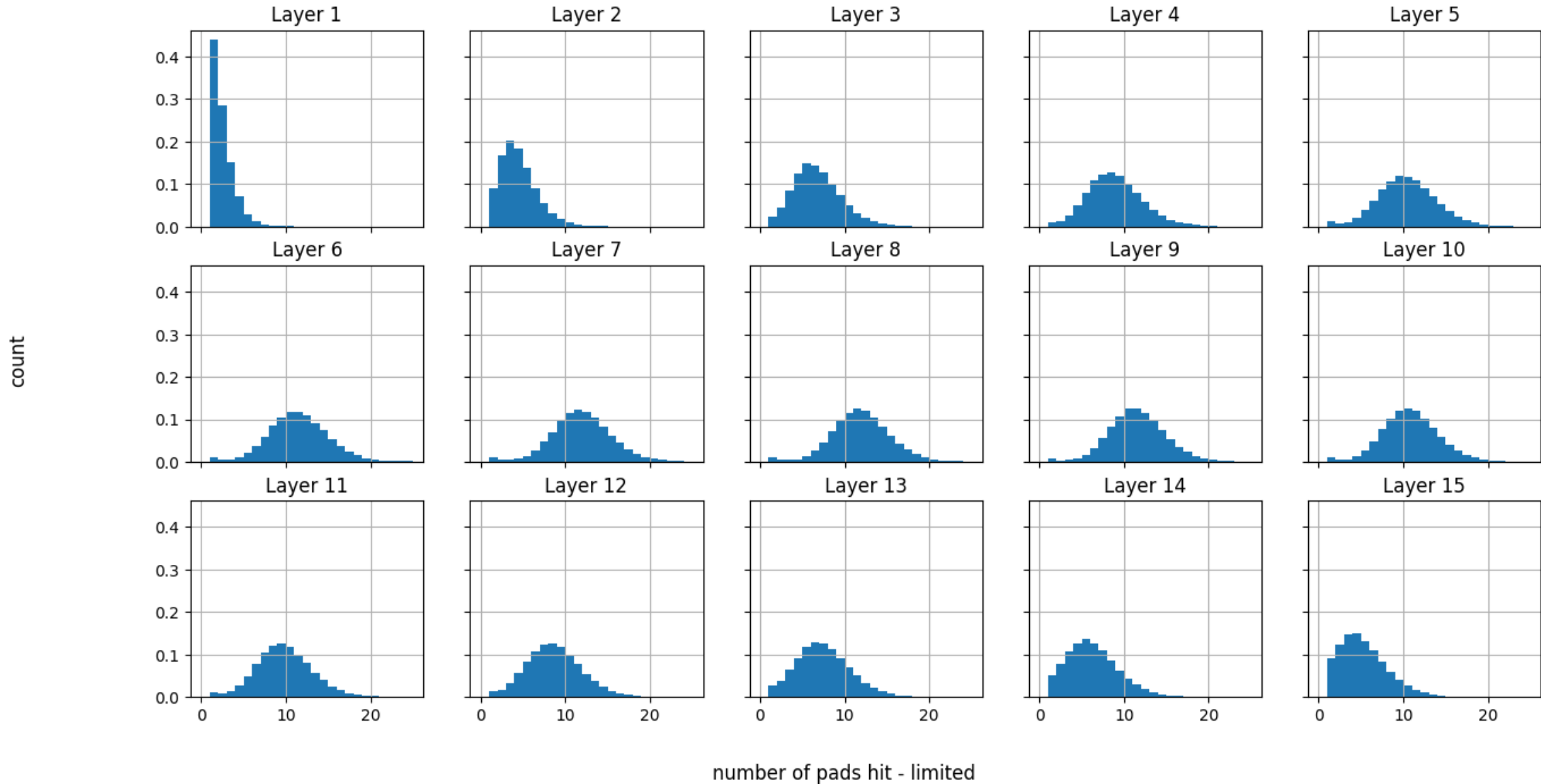
Based on a Molière radius of about 9mm, if the particle hit in the center of a pad, we expect to see about 25 pads hit.

Other hit positions will give a lower number.

Plot to the right shows the number of events with more than 25 pads hit, per layer.

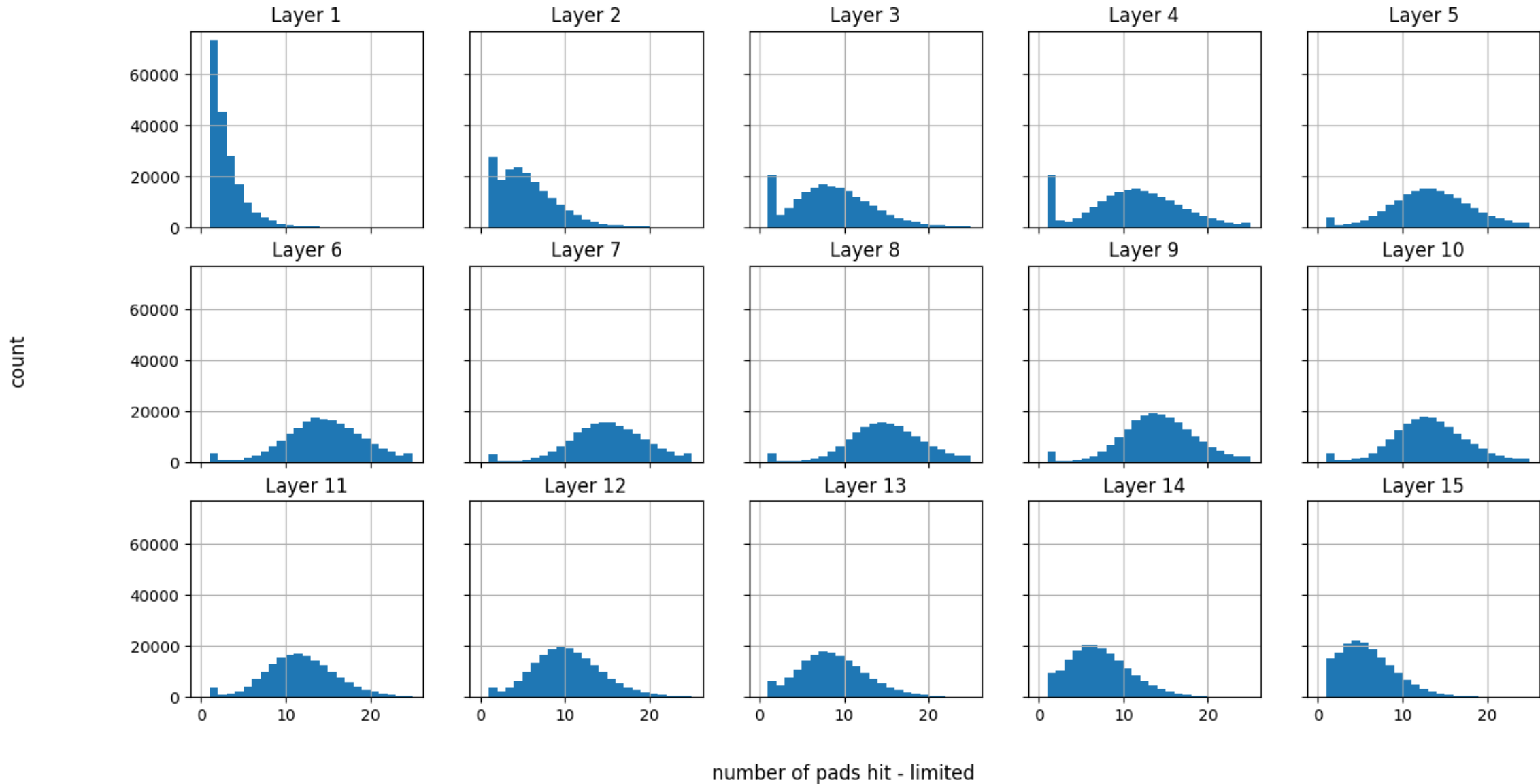


CALICE



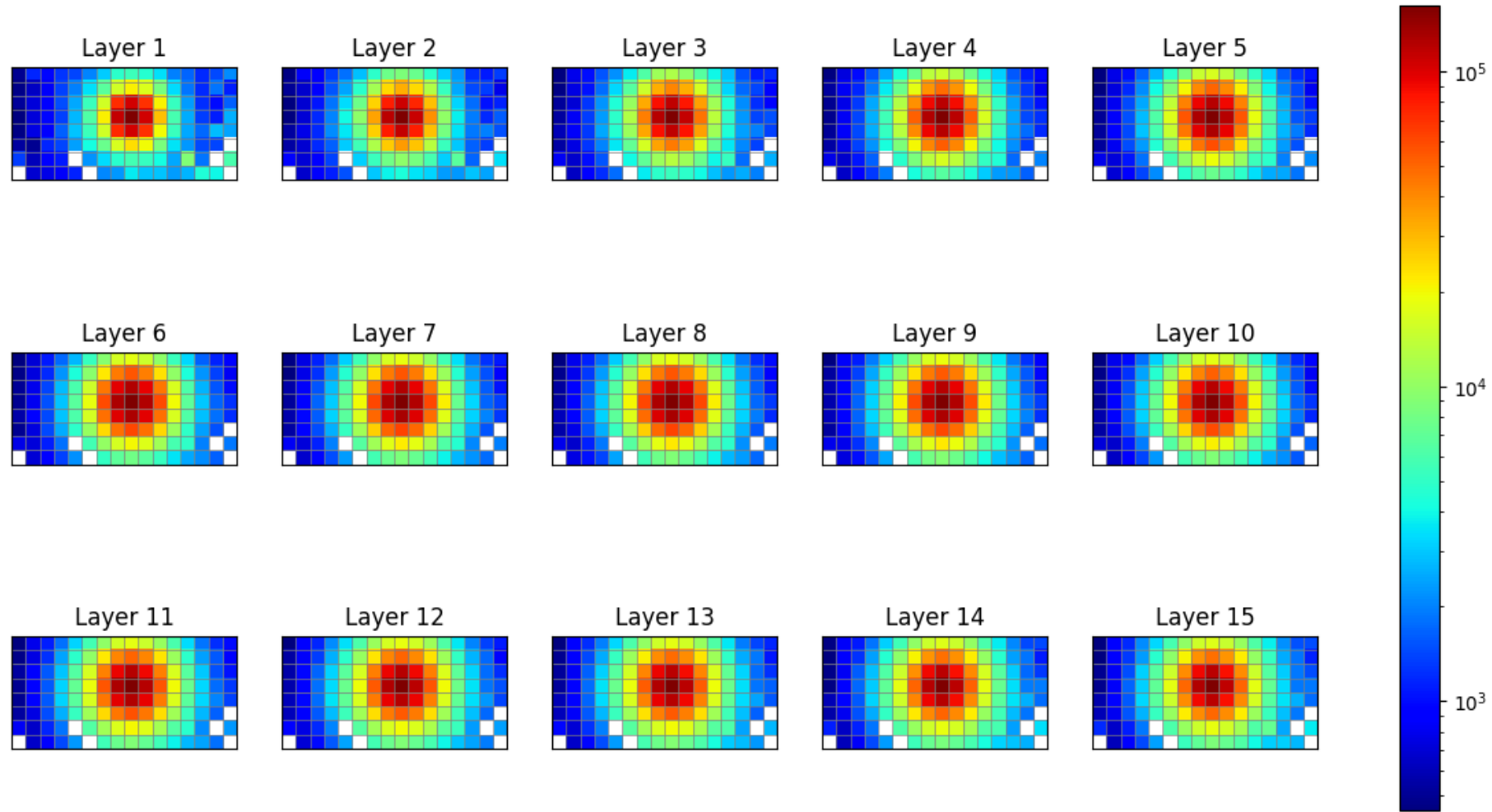
- Events with up to 25 pads
- Full sized plots, layer-by-layer, are available at the end of the presentation

GaAs



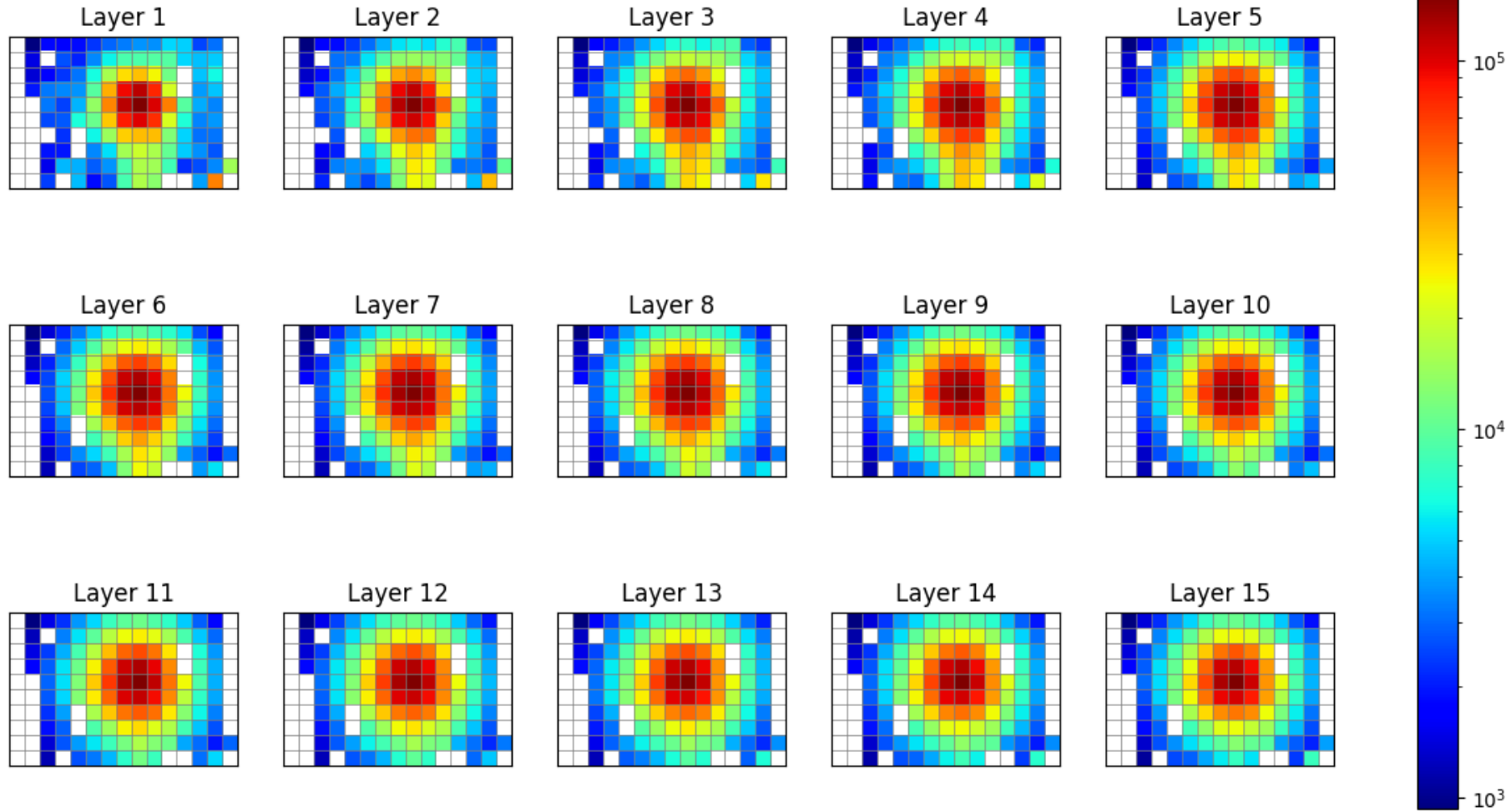
- Events with up to 25 pads
- Full sized plots, layer-by-layer, are available at the end of the presentation

CALICE - number of hits - log scale



Seems
isotropic in
the radial
direction

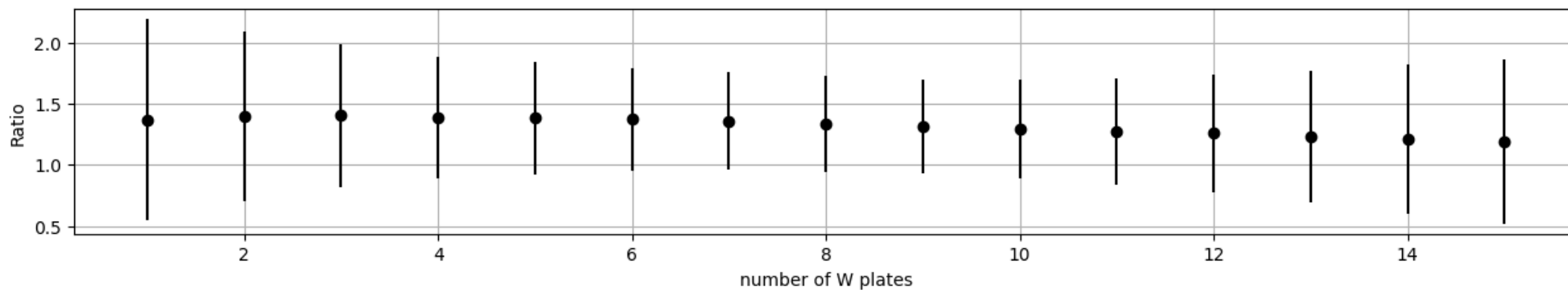
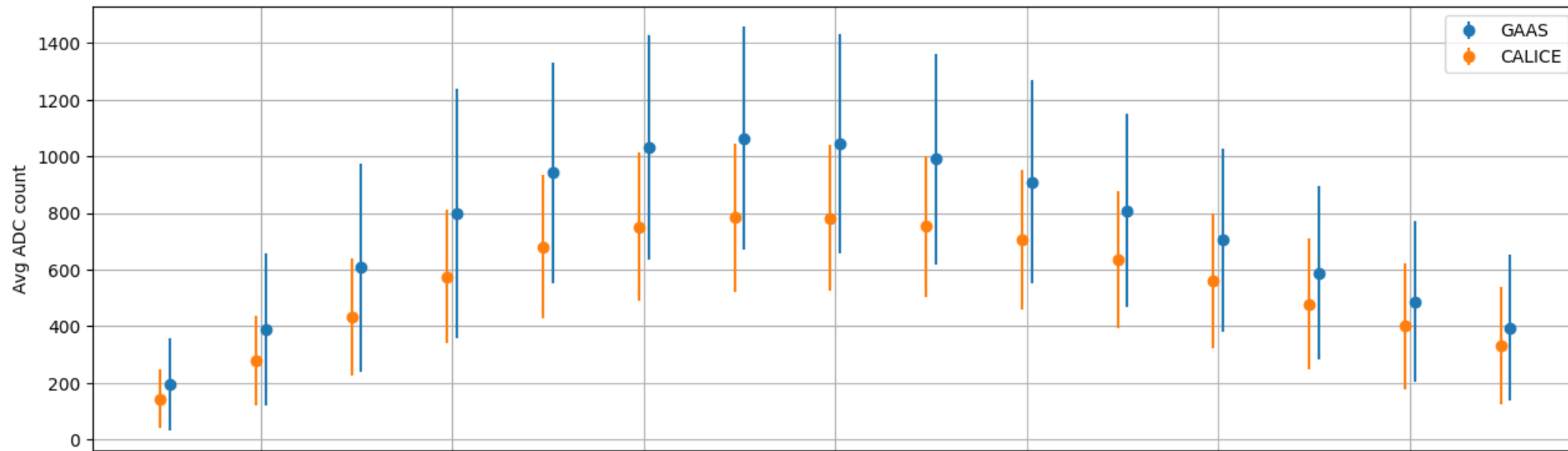
GaAs - number of hits - log scale



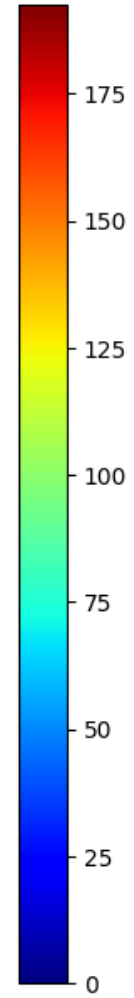
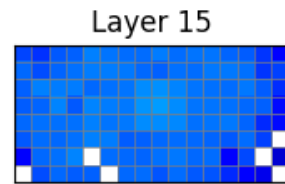
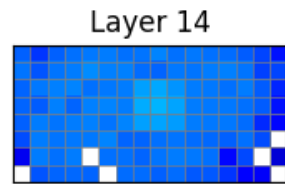
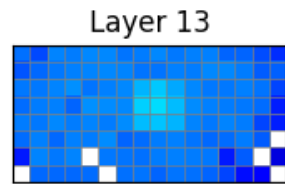
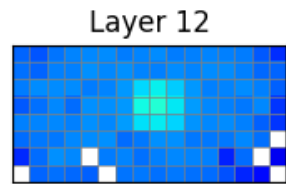
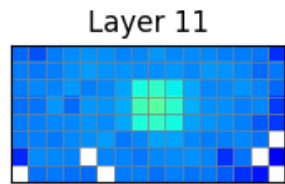
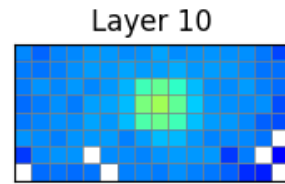
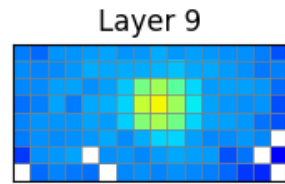
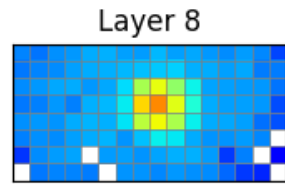
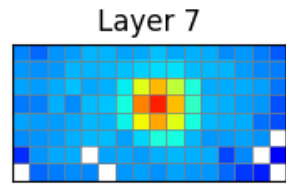
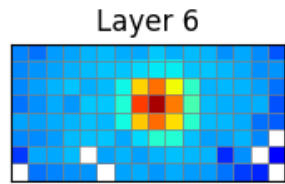
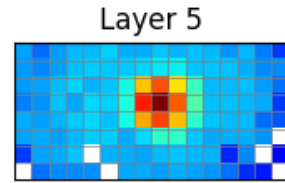
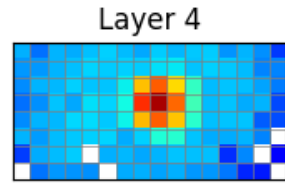
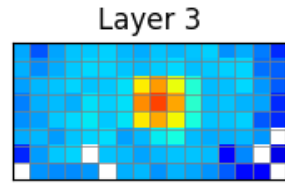
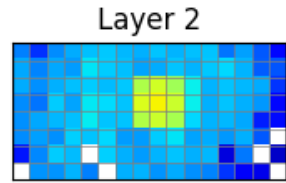
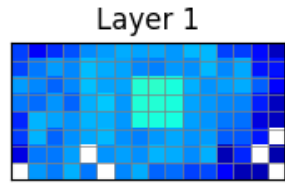
Pads below
the beam
center are
hit more
often than
those above

→ Traces ?

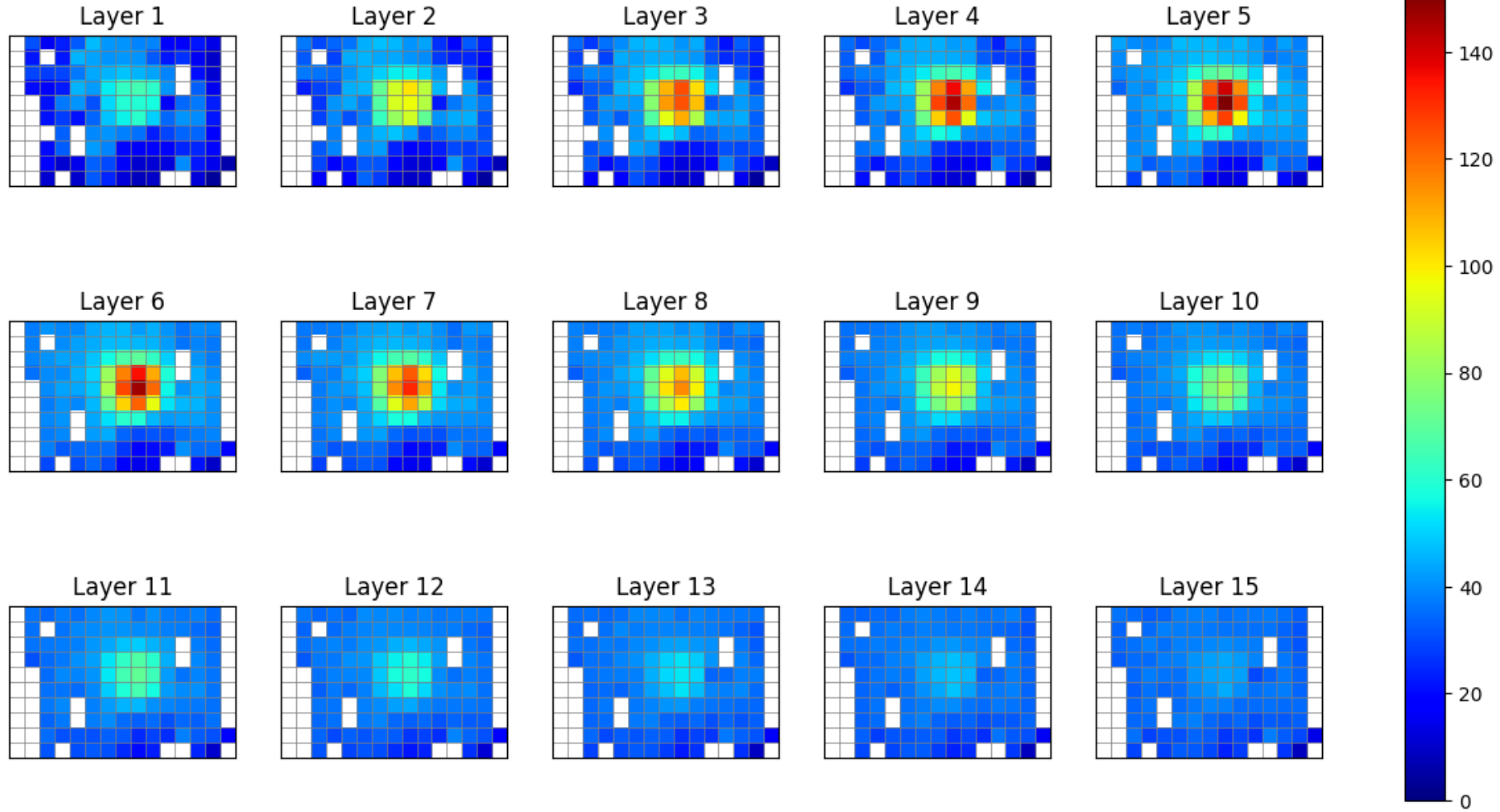
AVERAGE ADC COUNT



CALICE - average ADC count



GaAs - average ADC count



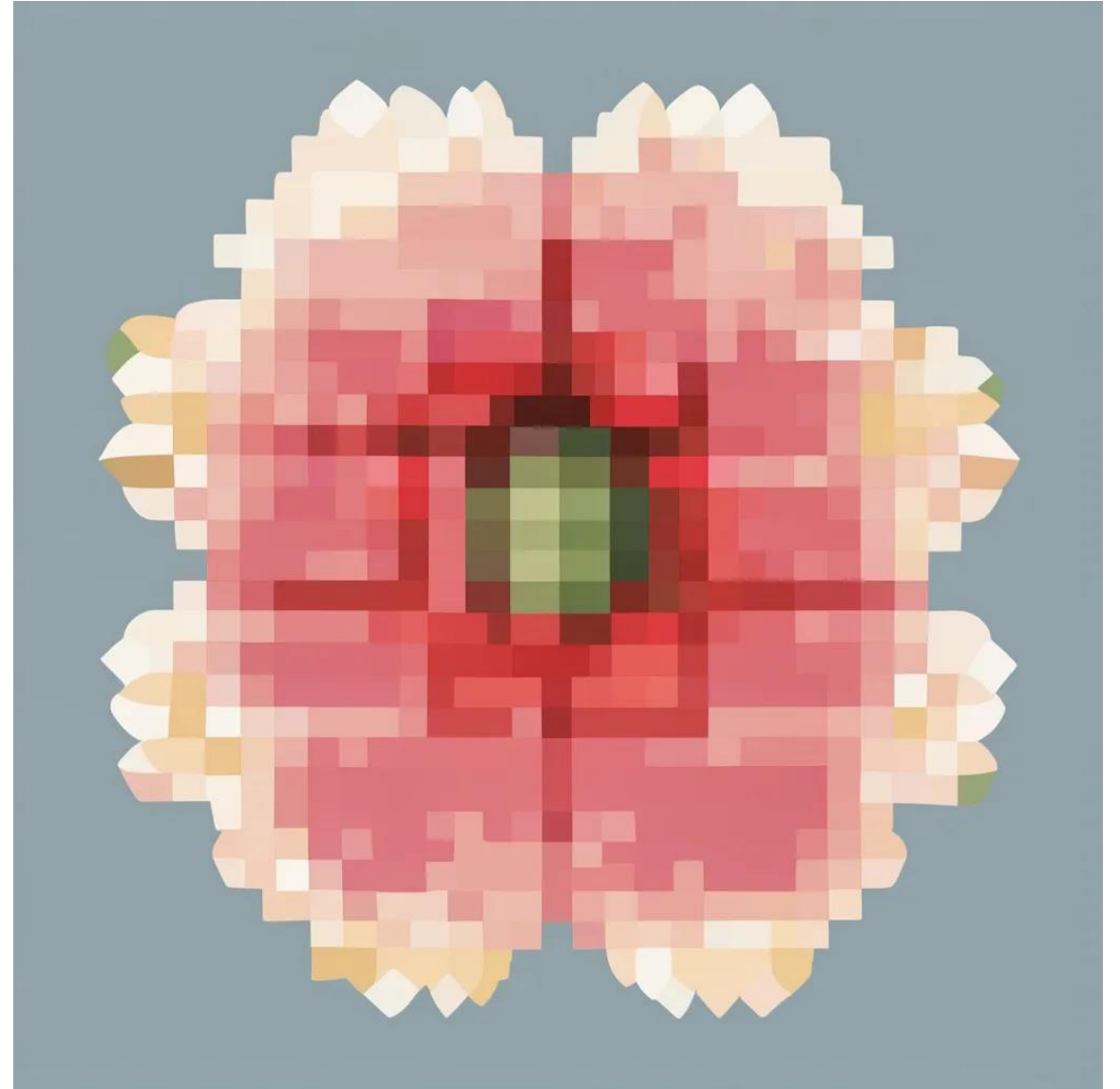
Pads below
the beam
have a lower
average
ADC count

→ Traces ?

SUMMARY

- Traces-based telescope resolution calculation – $50 \mu m$
- Alignment precision is dominated by the telescope resolution
- Calorimetry:
 - Traces seem to cause additional pads hit with lower ADC count.
 - Total ADC count in GaAs is on average higher than CALICE – to be further investigated

THANK YOU!

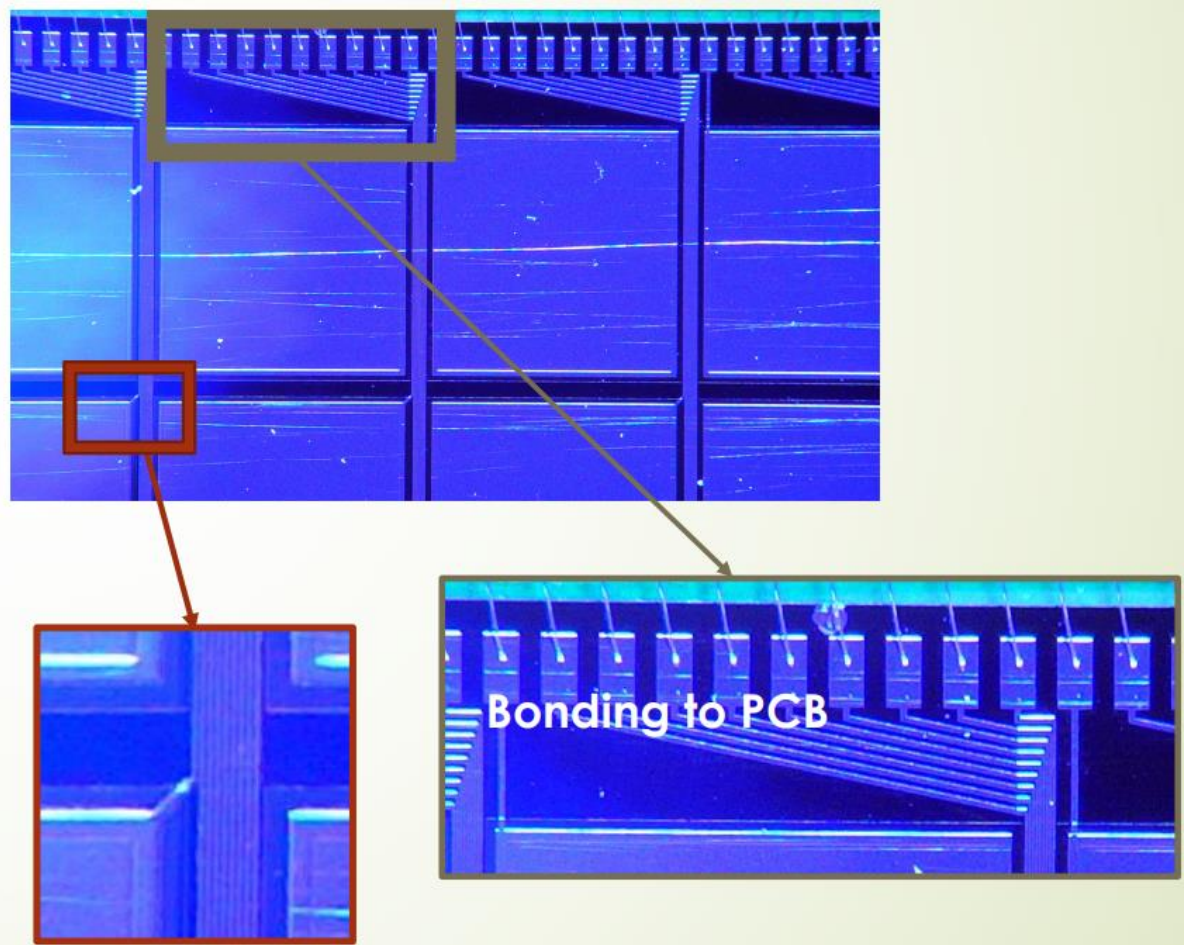


A pixelated flower in reference to the previous slides



BACKUP

TRACES BASED EFFECTS

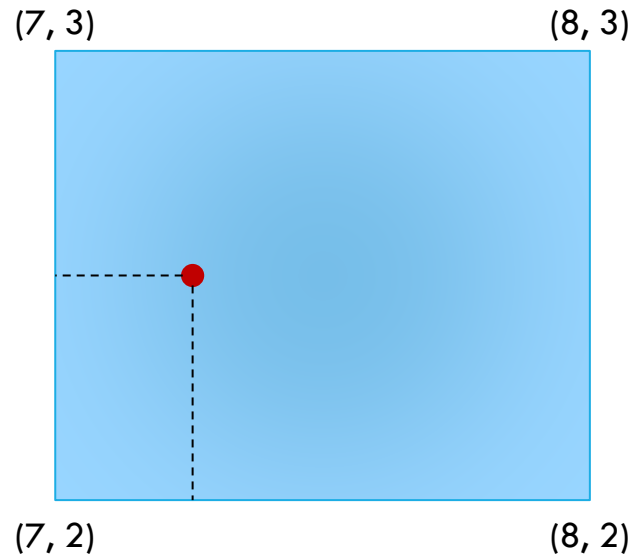


ALIGNMENT GOAL

Translate global position given by the telescope to a local position on the sensor.

Local coordinate system agreed – units of pad length

- e.g. (7.25, 2.5) is located as such pad (7, 2):



ALIGNMENT METHODS

- Max hits in pad

Set a grid and find the position that gives the highest number of hits in the correct pad.

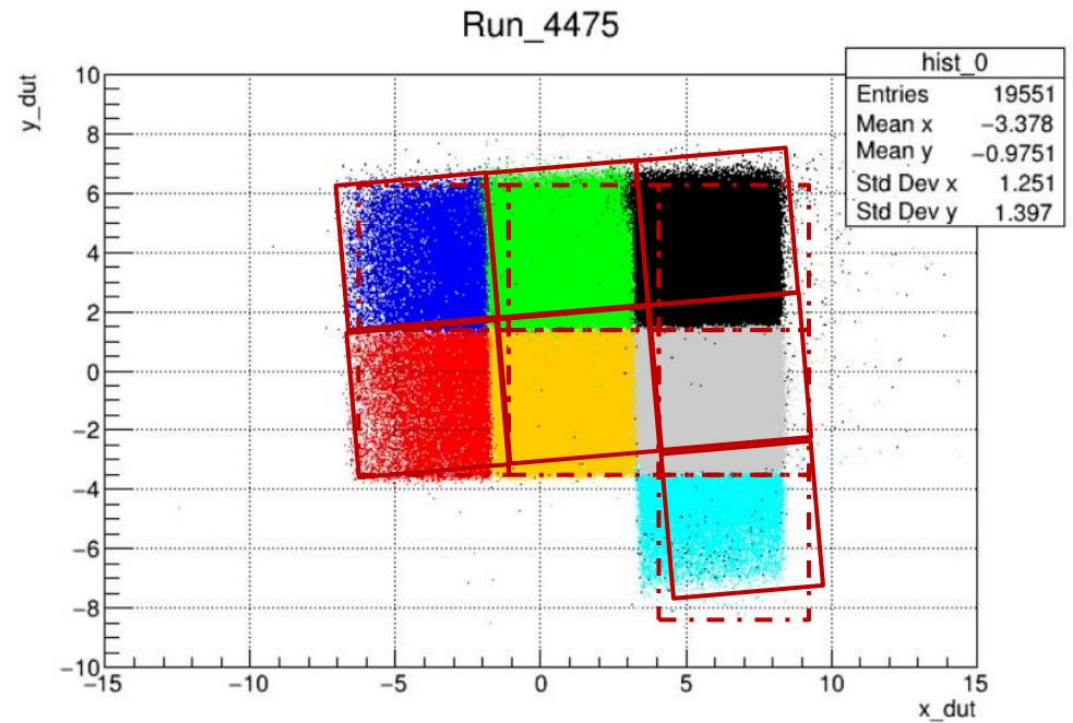
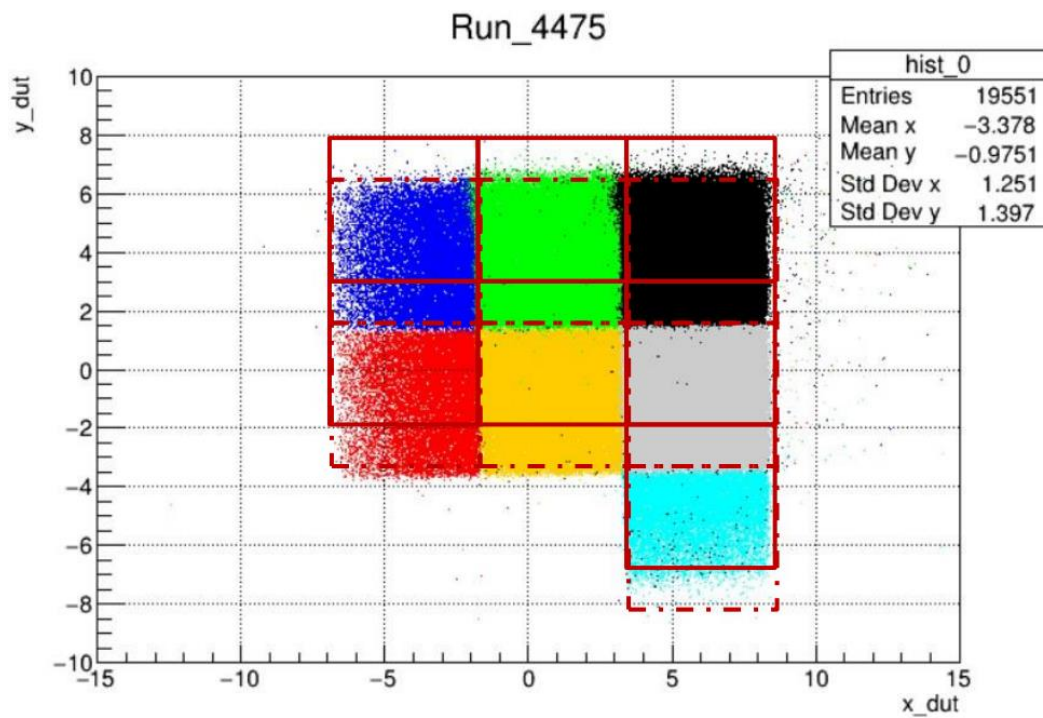
Input data – events of a single electron & a single hit.

- Hough Transform

Used as a line recognition algorithm to find the edges of the pads.

Input data – single electron events that hit in between pads (to get a good “image” of the edges).

MAX HITS



HOUGH TRANSFORM

- Shape detecting framework
- Used a slight variation of the original

FIND RELEVANT AREA

The data is noisy, and we are only interested in a single line.

We don't want to consider all points, so we first find an estimation for the edge position.

Revisited – seems to miss some fine-tuning.

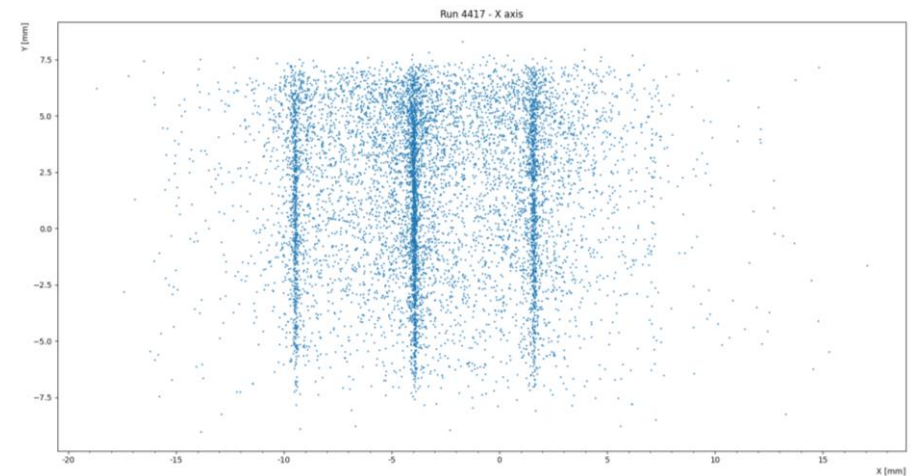
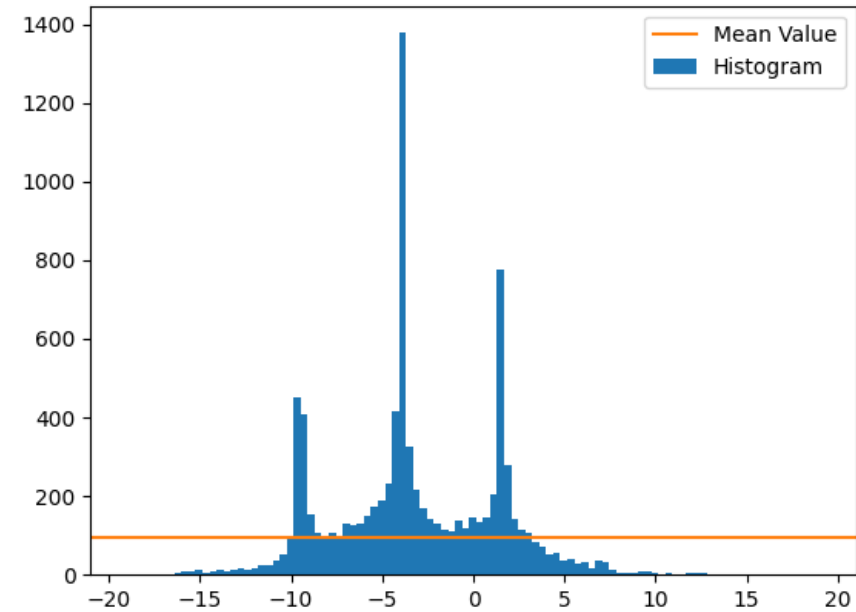


Image and parameter space

$$y = mx + b$$

variables

parameters

$$y - mx = b$$

variables

parameters

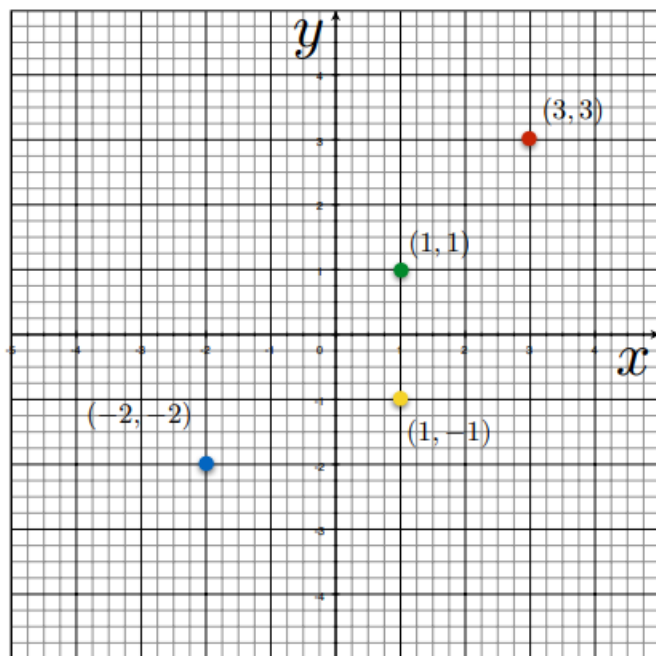
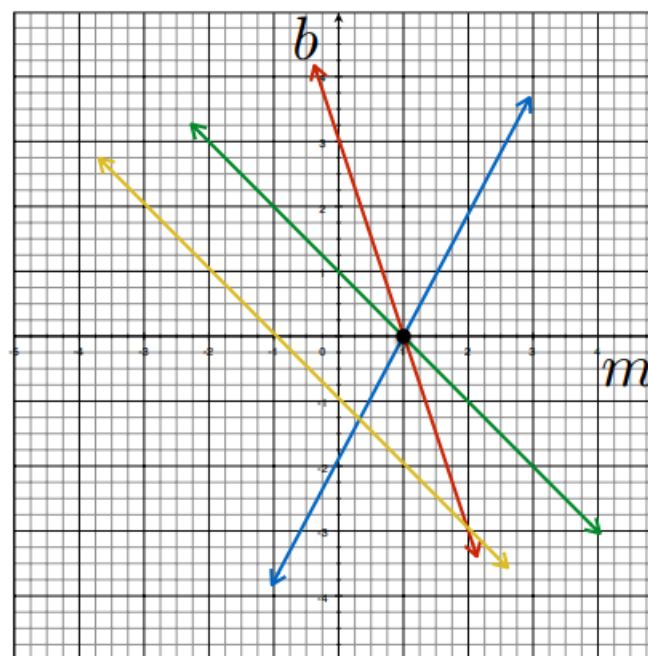


Image space

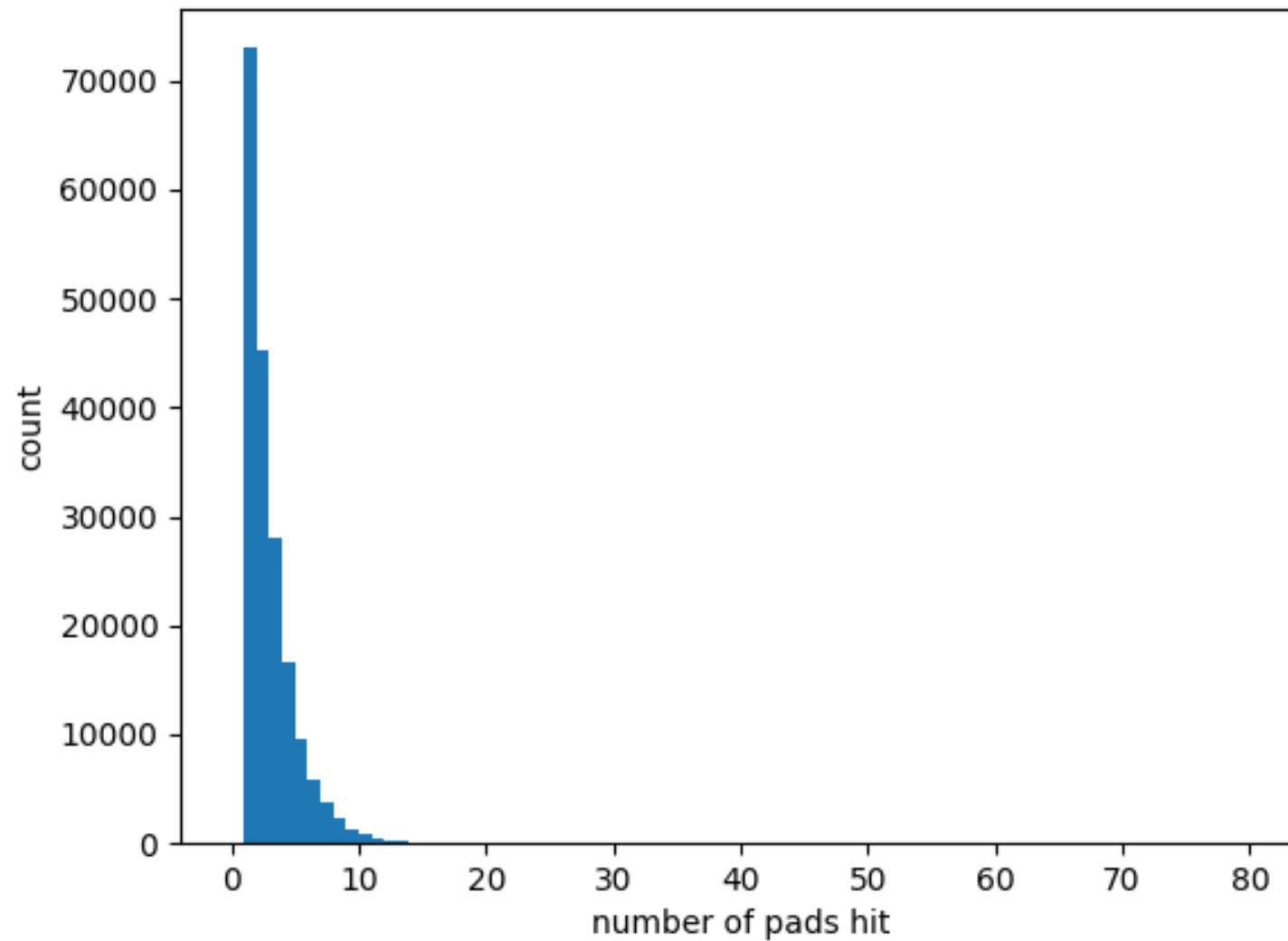
four points
become
?



Parameter space

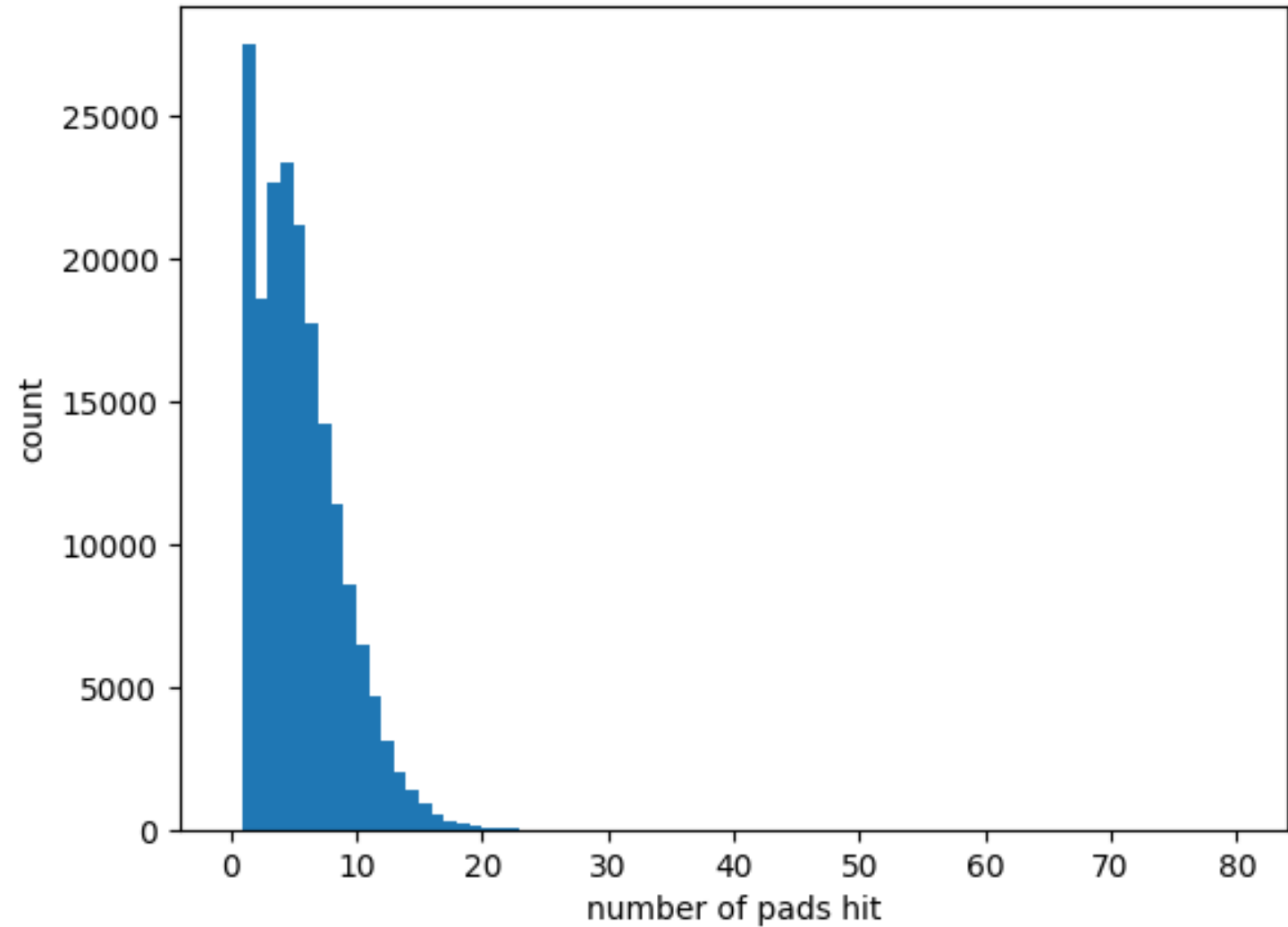
GaAs

Layer 1



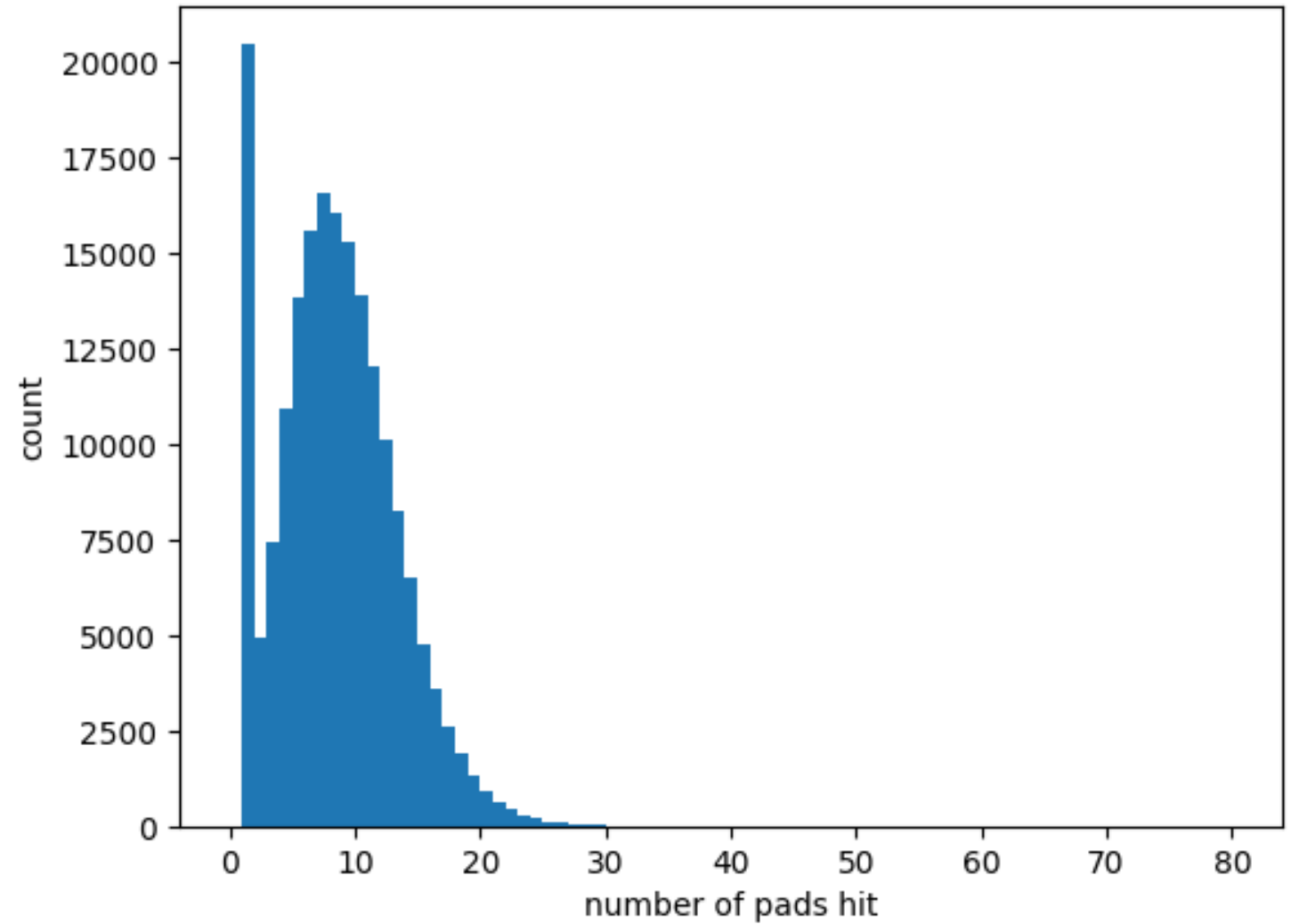
GaAs

Layer 2



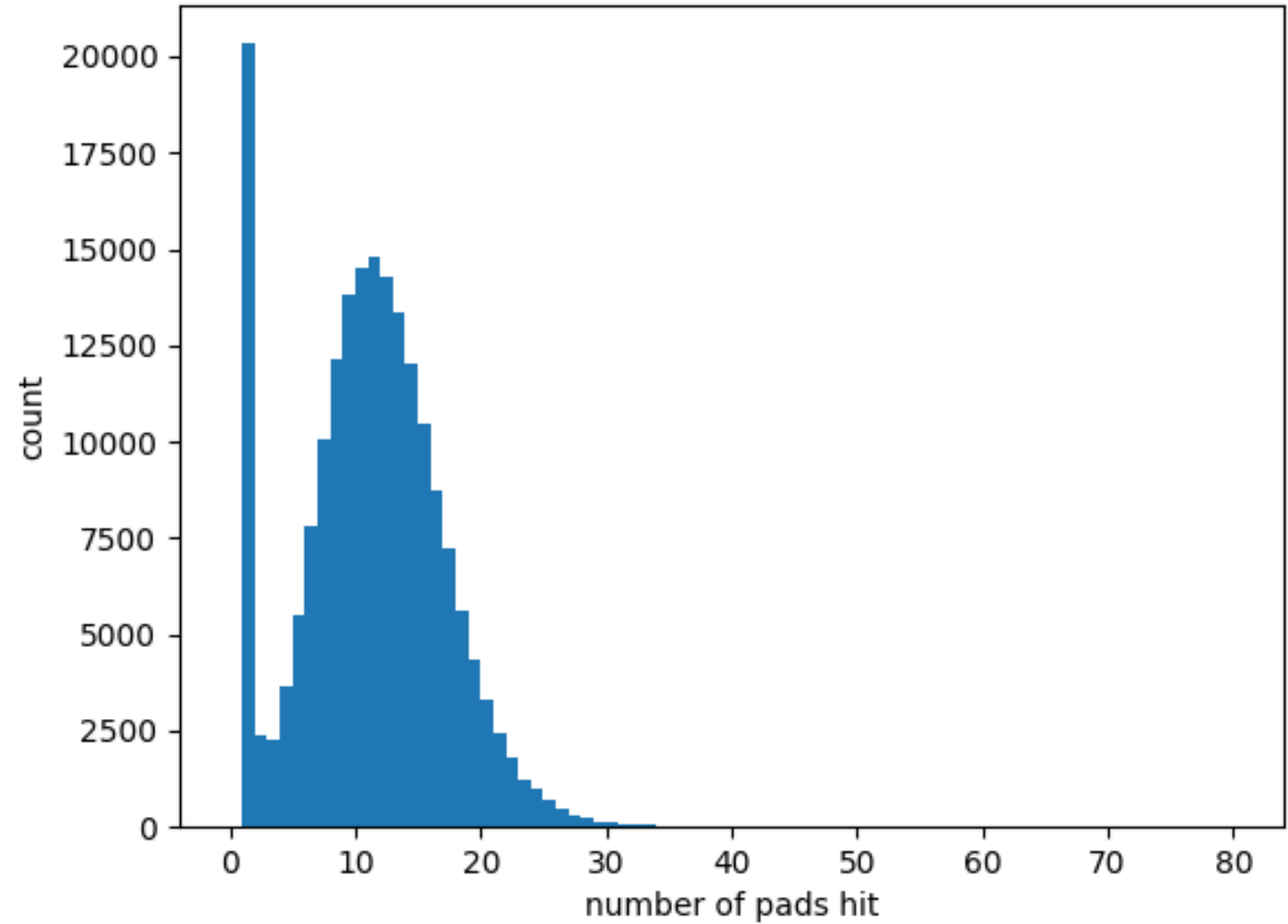
GaAs

Layer 3



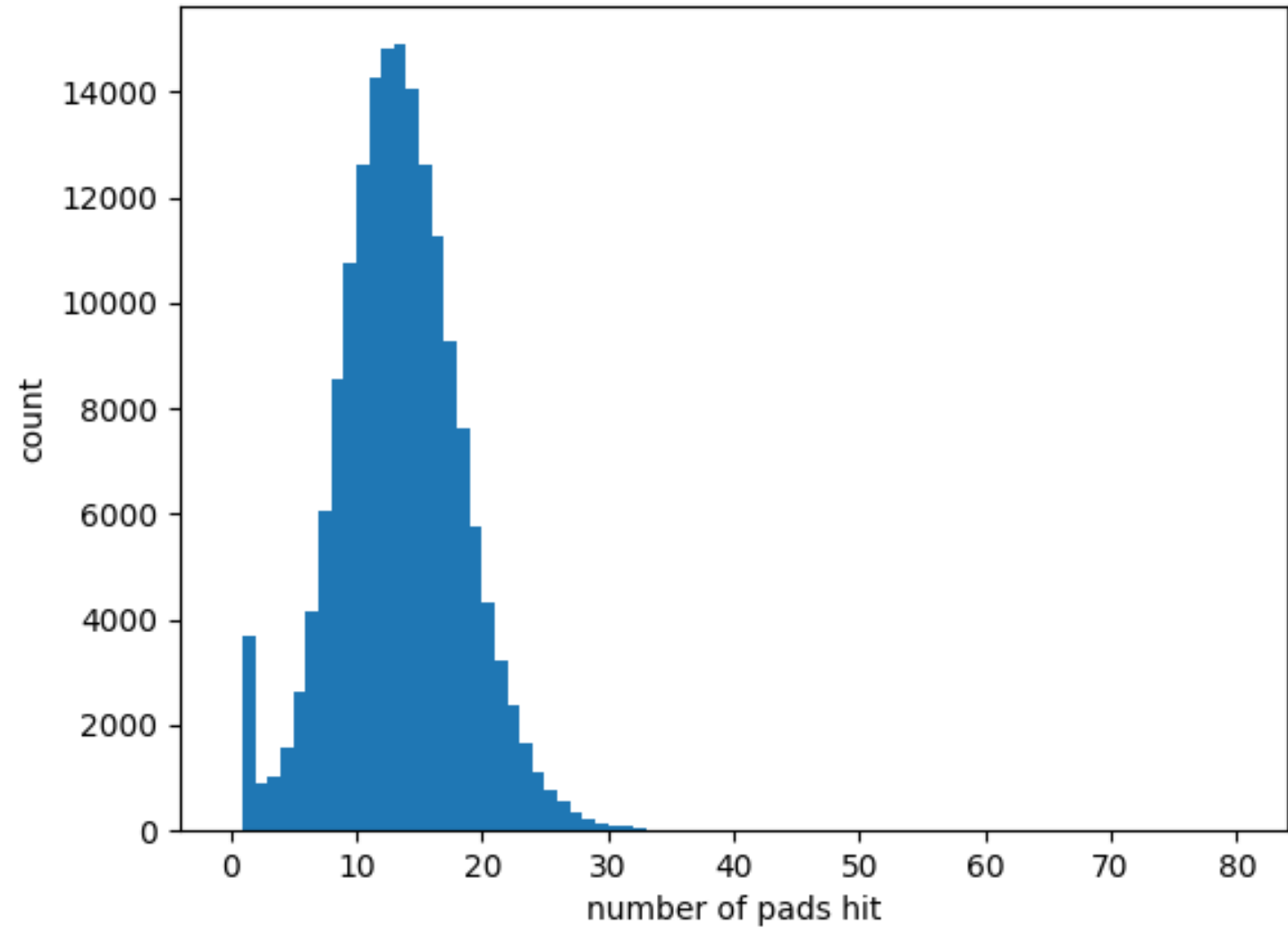
GaAs

Layer 4



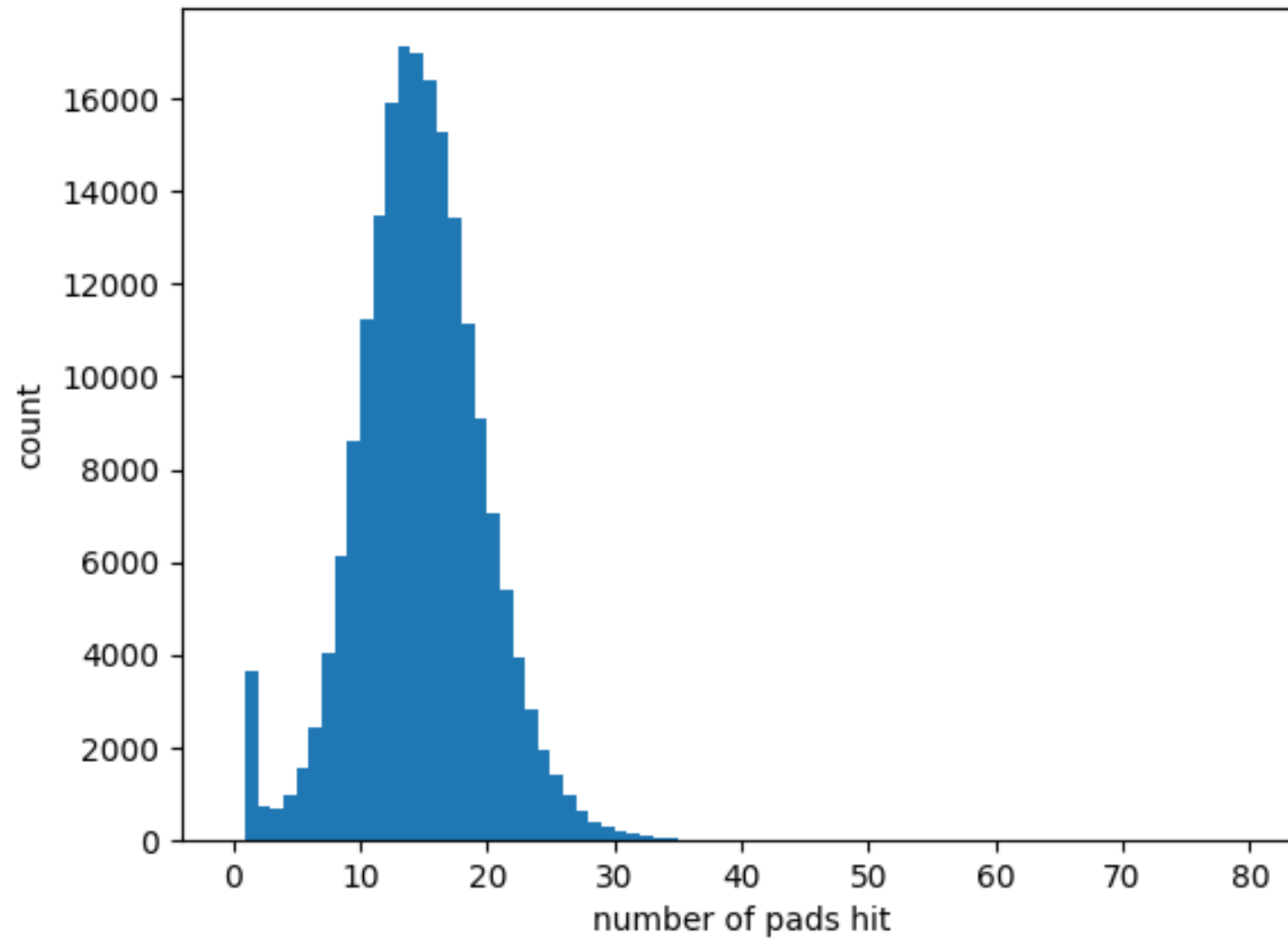
GaAs

Layer 5



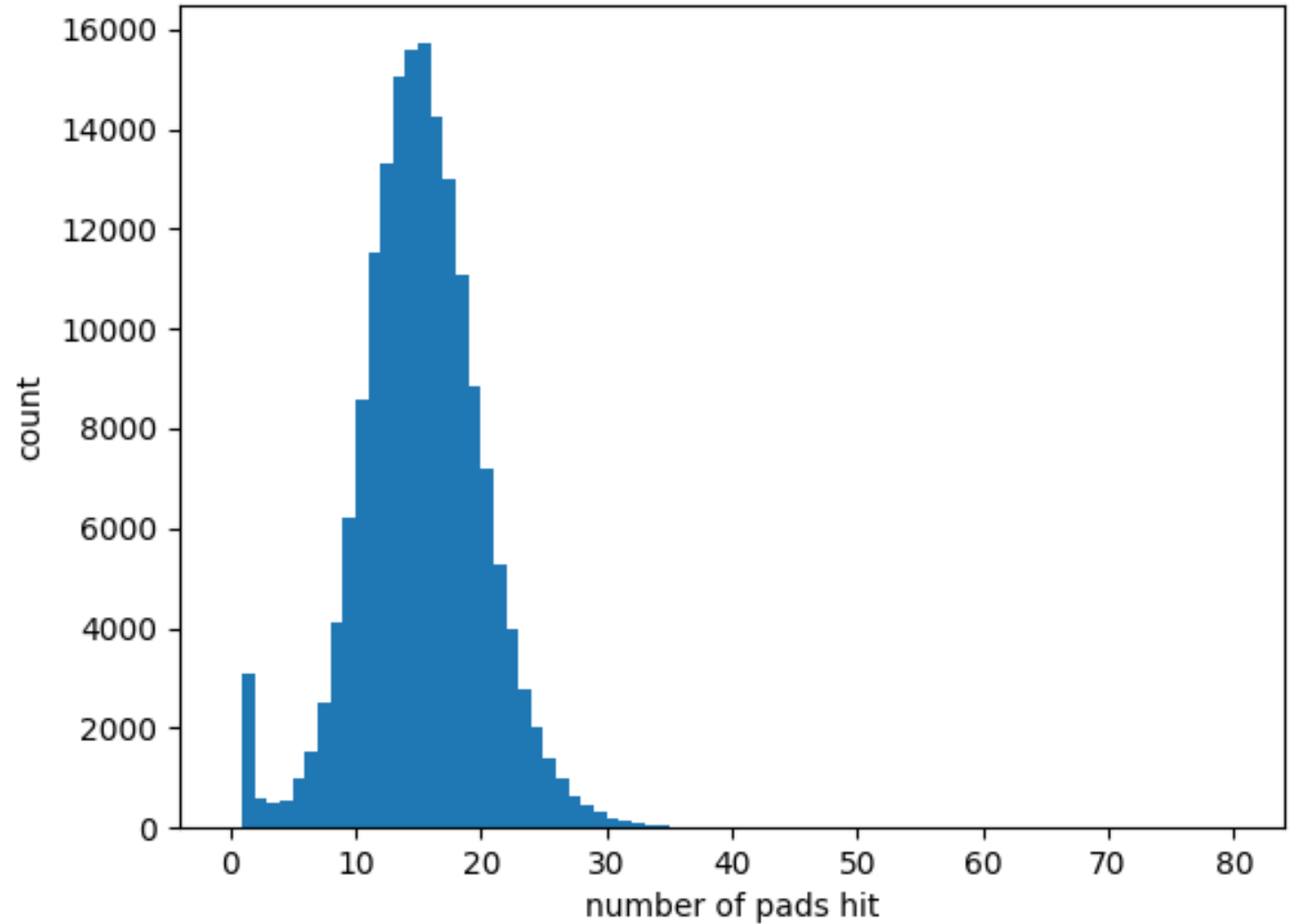
GaAs

Layer 6



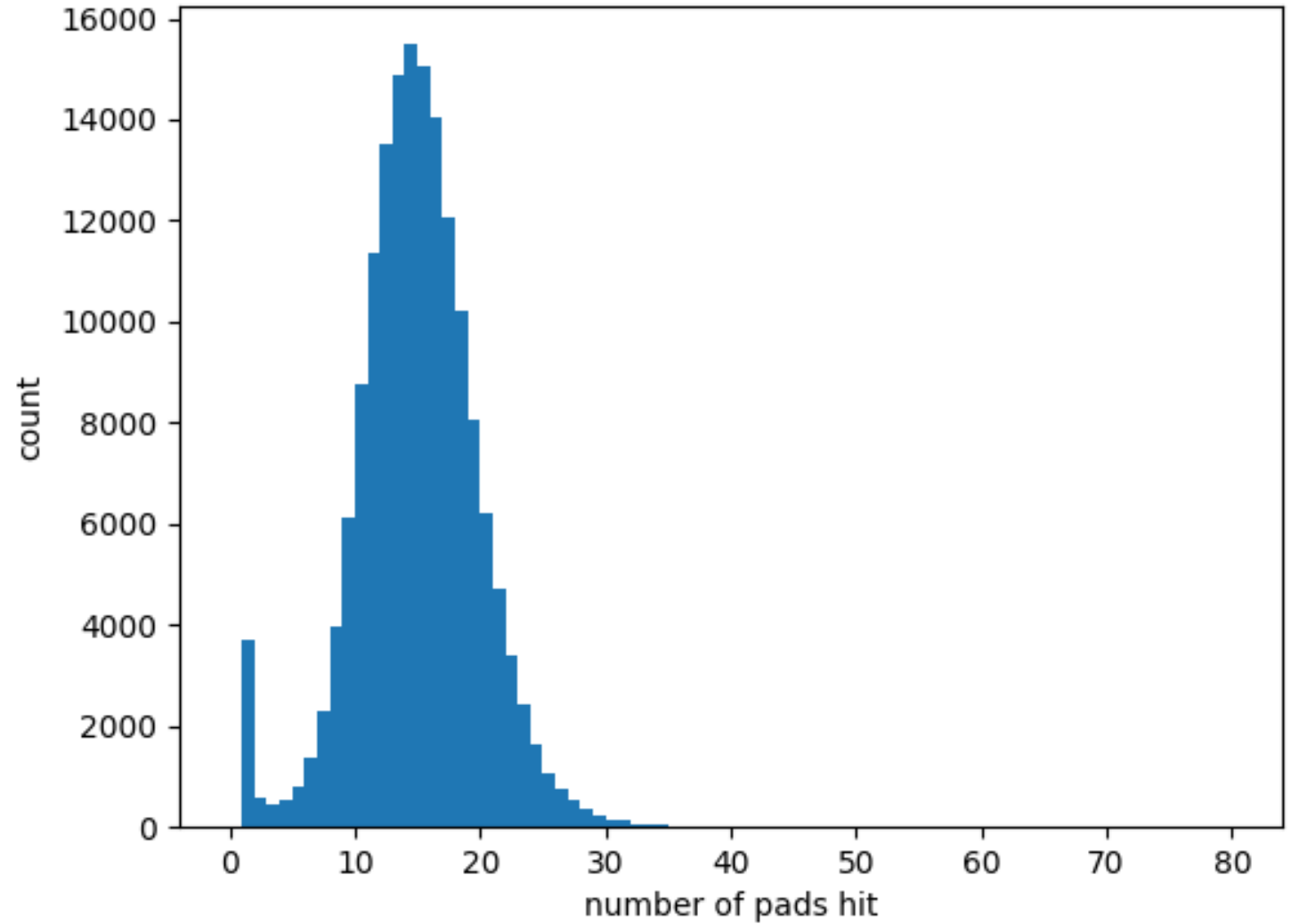
GaAs

Layer 7



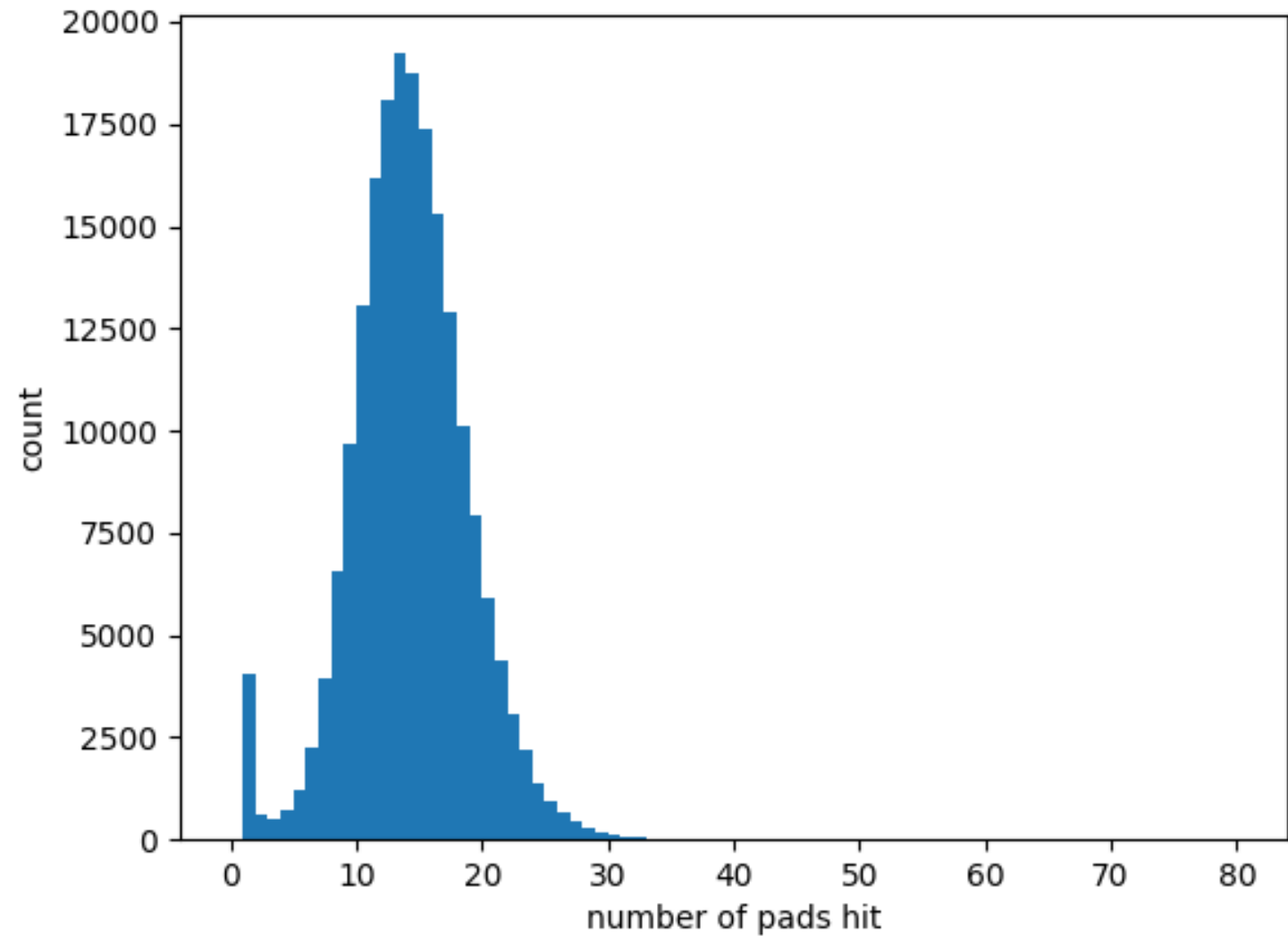
GaAs

Layer 8



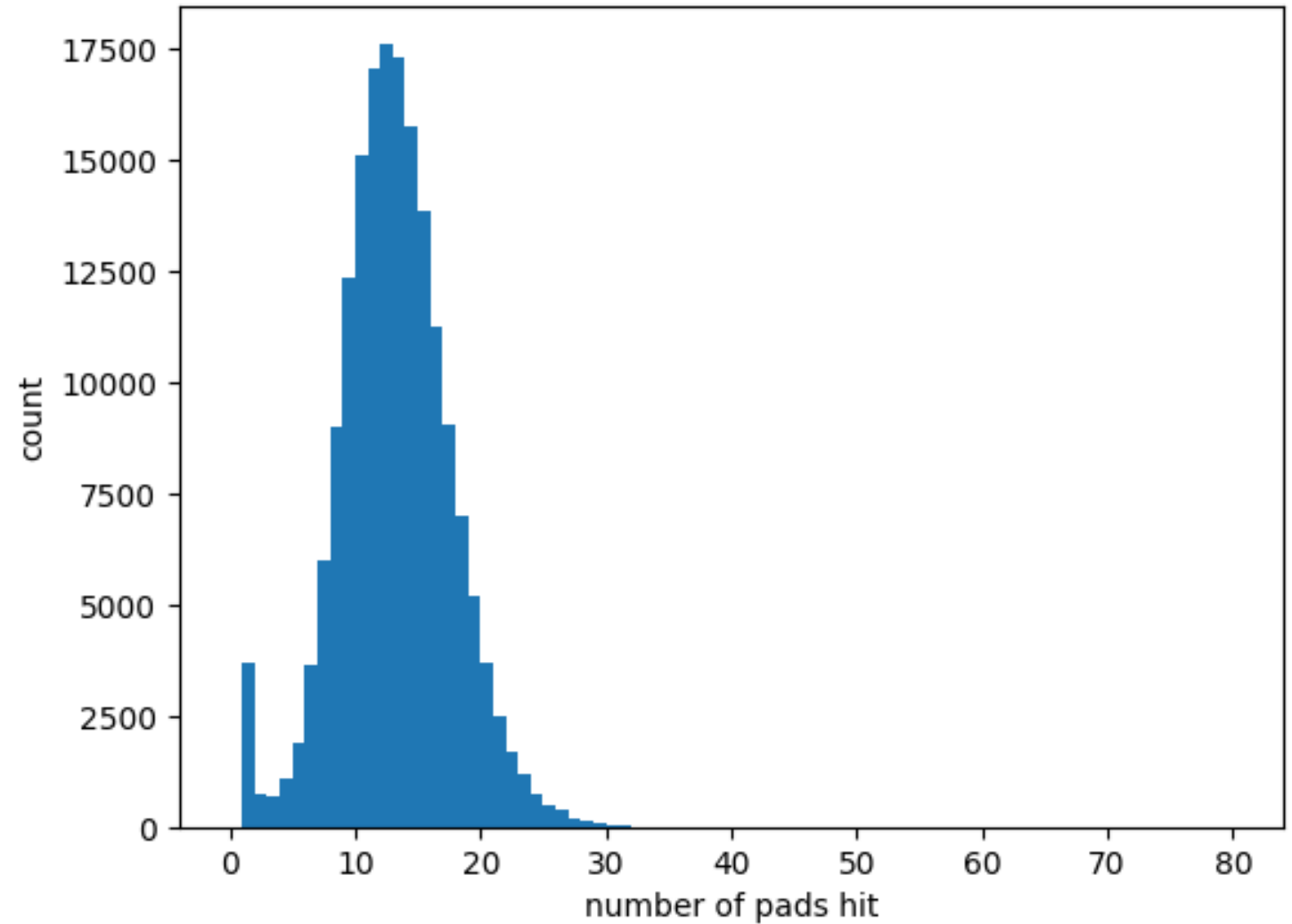
GaAs

Layer 9



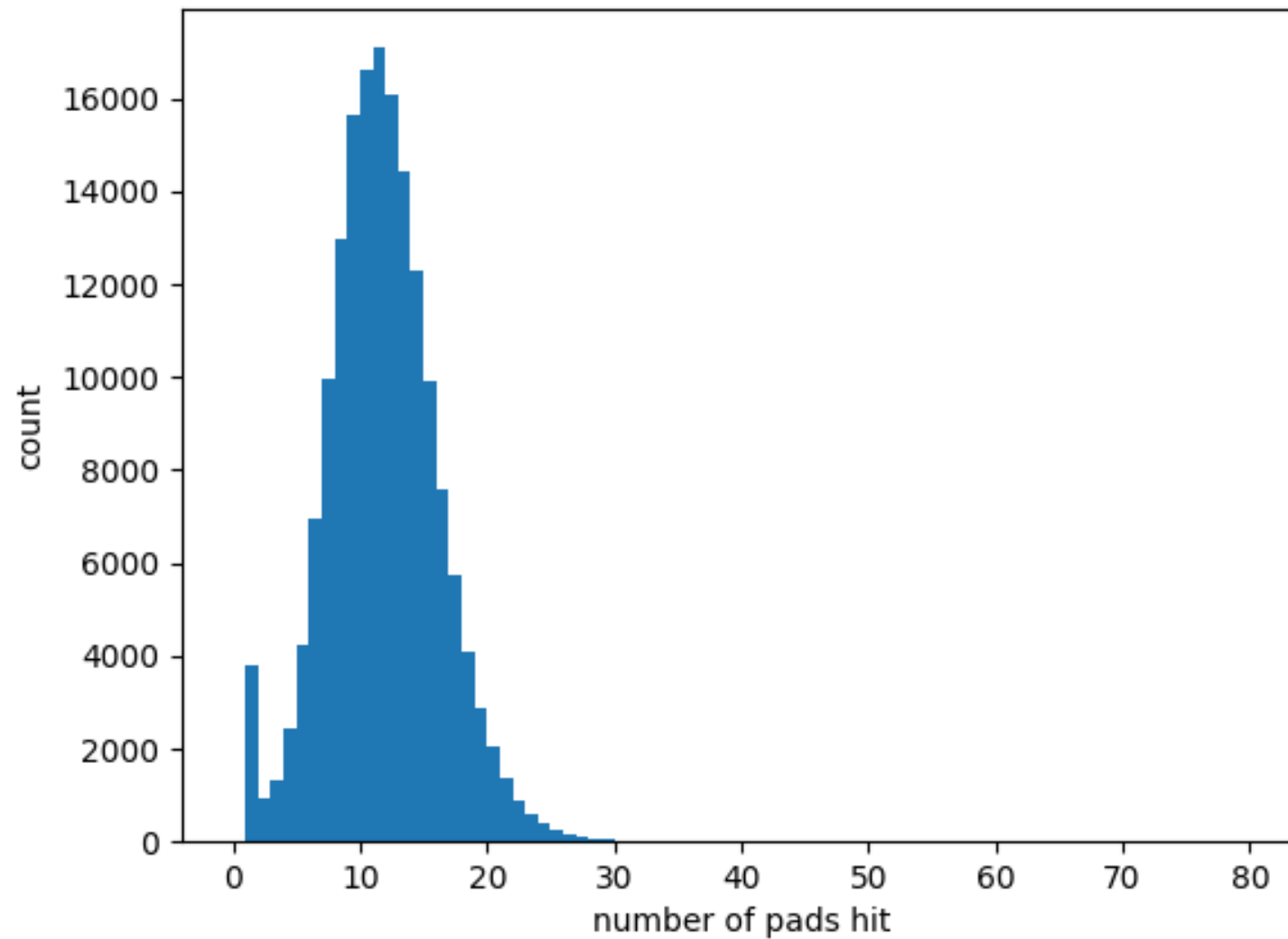
GaAs

Layer 10



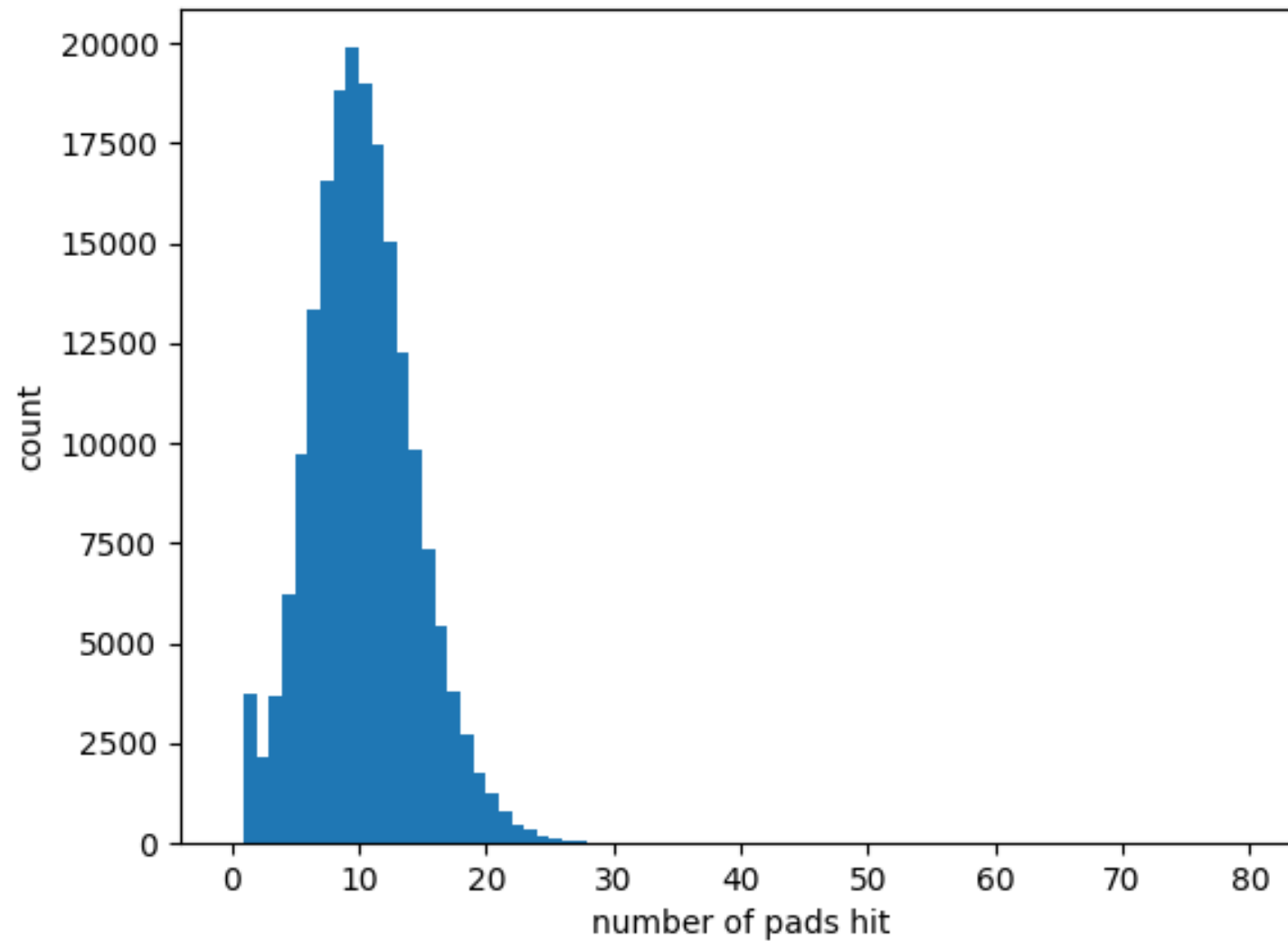
GaAs

Layer 11



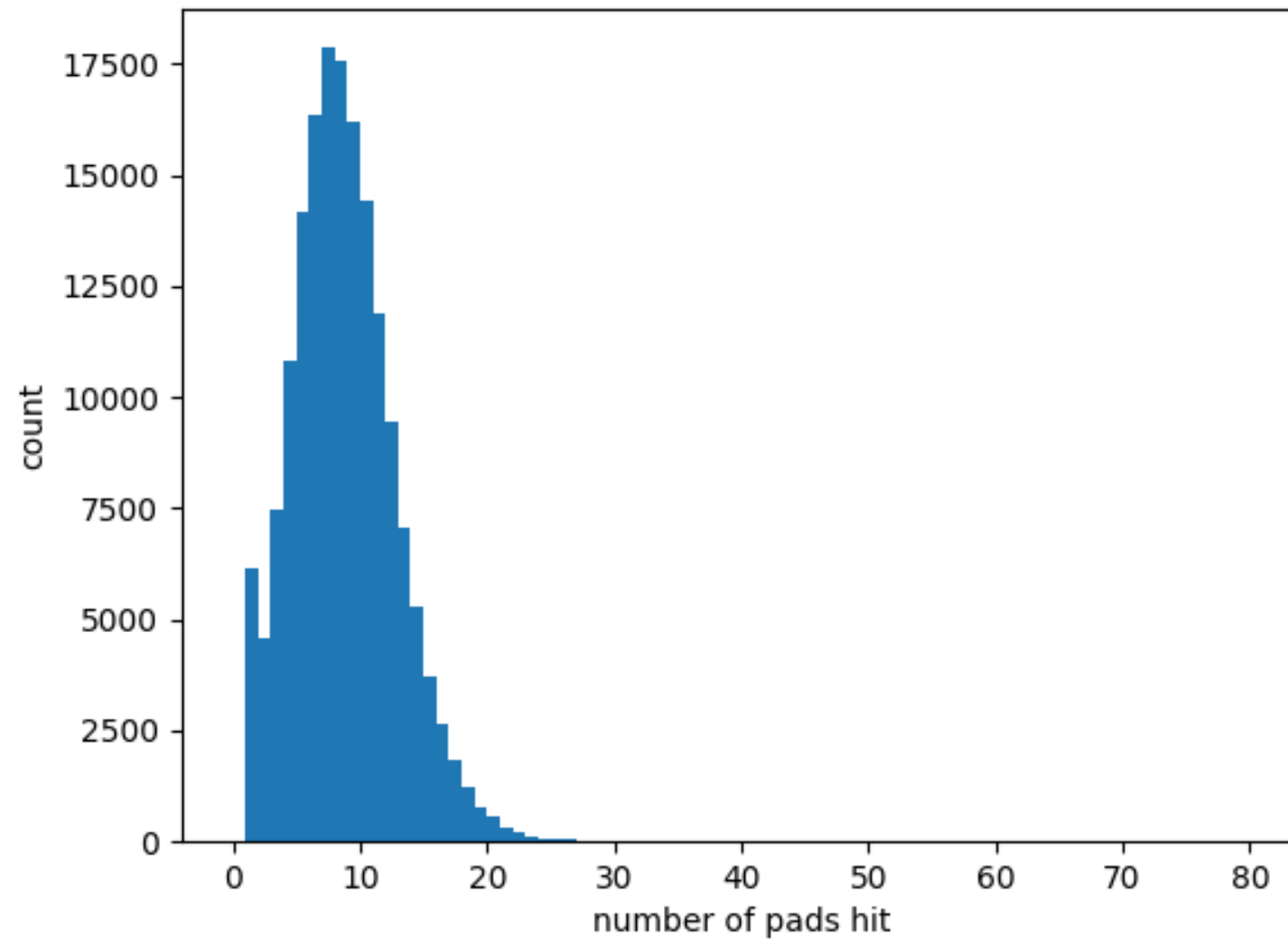
GaAs

Layer 12



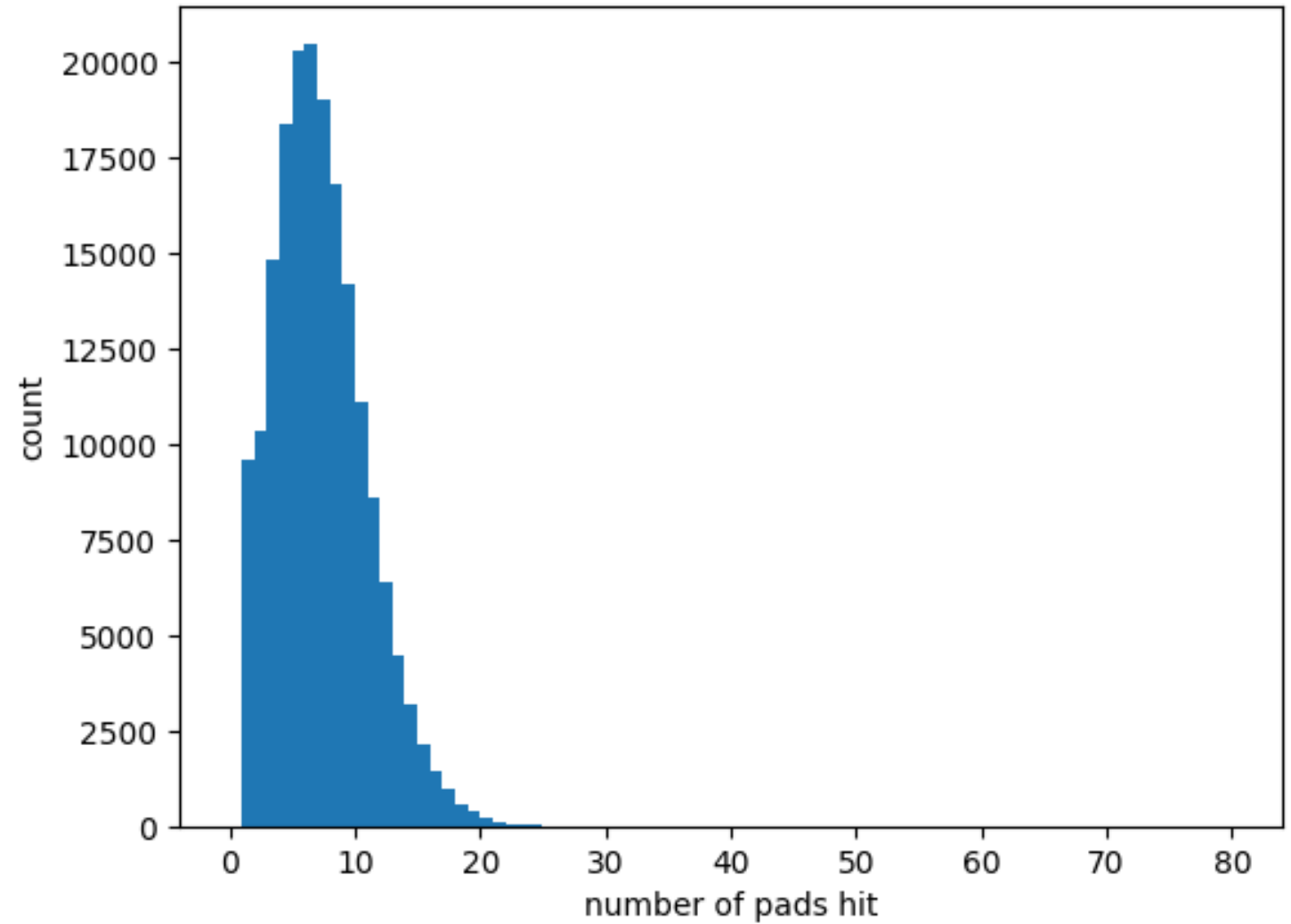
GaAs

Layer 13



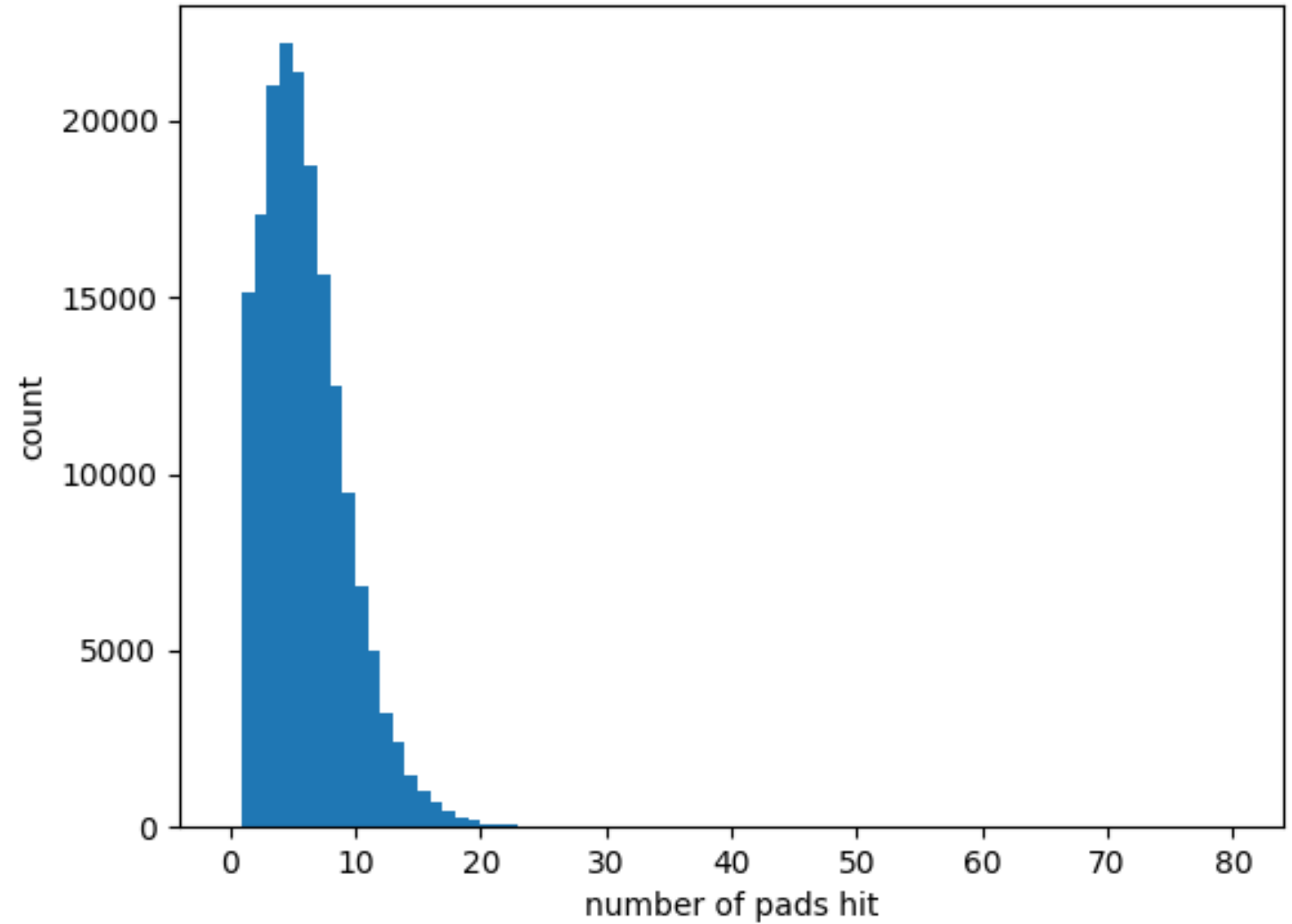
GaAs

Layer 14



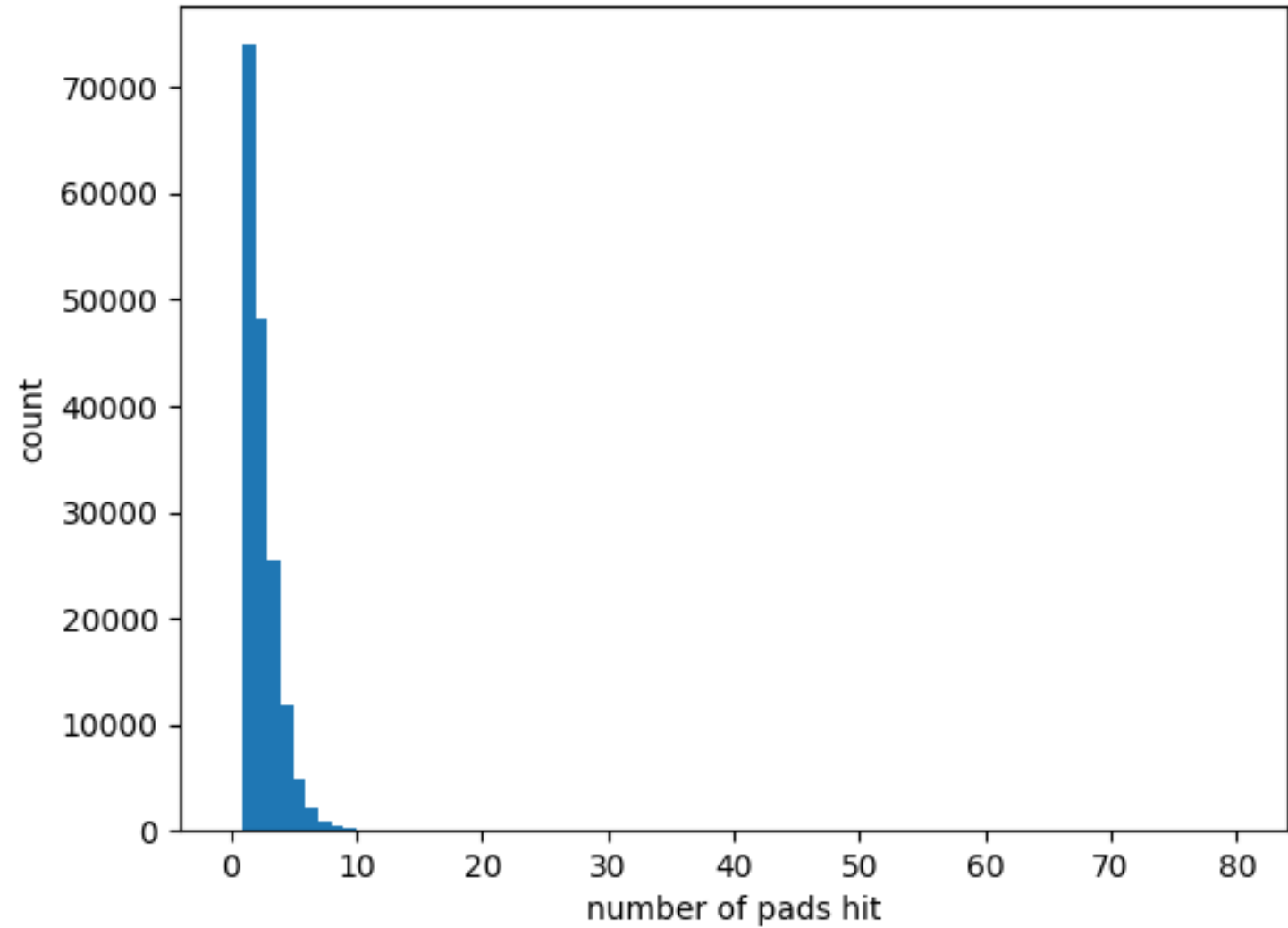
GaAs

Layer 15



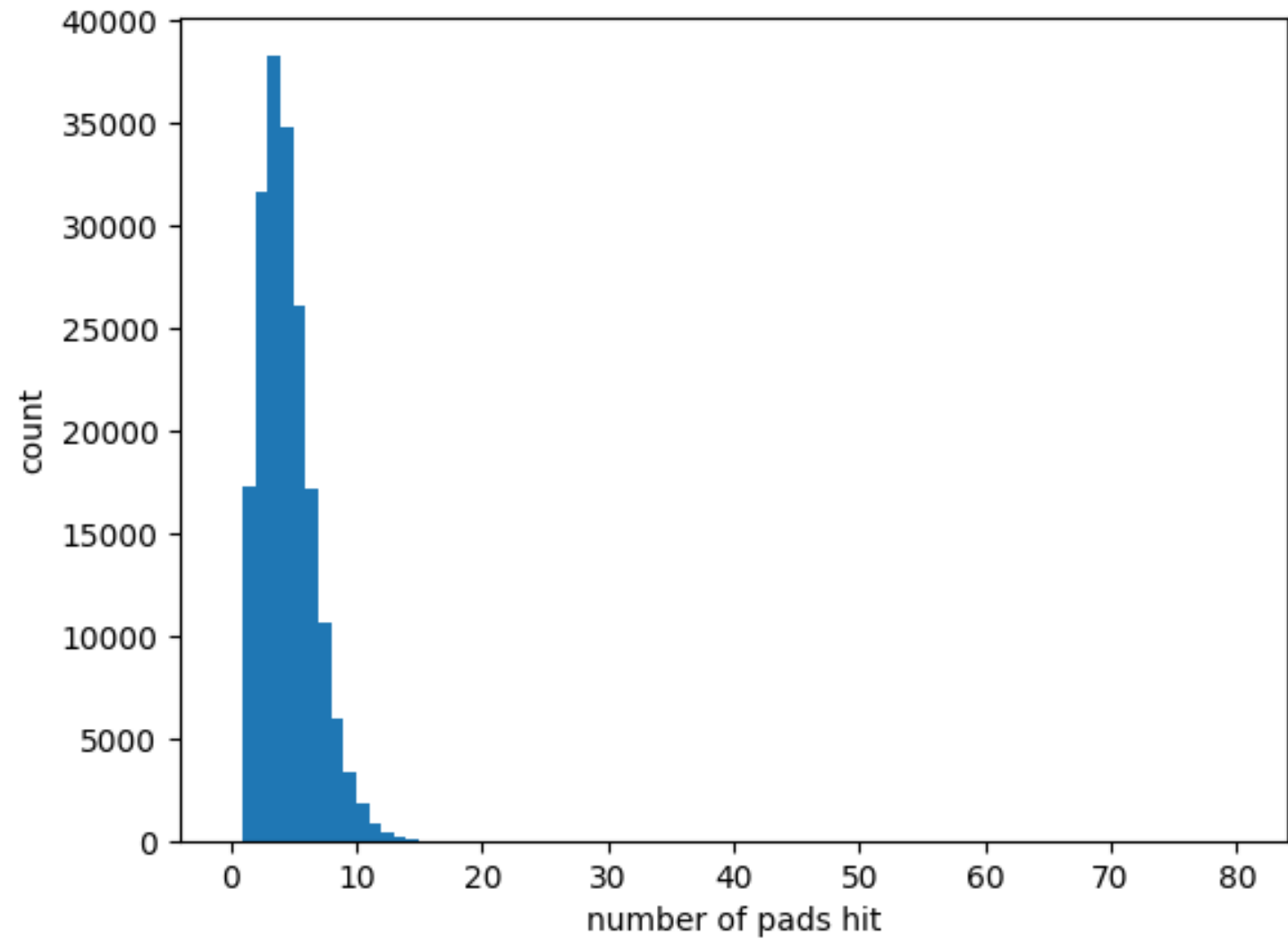
CALICE

Layer 1



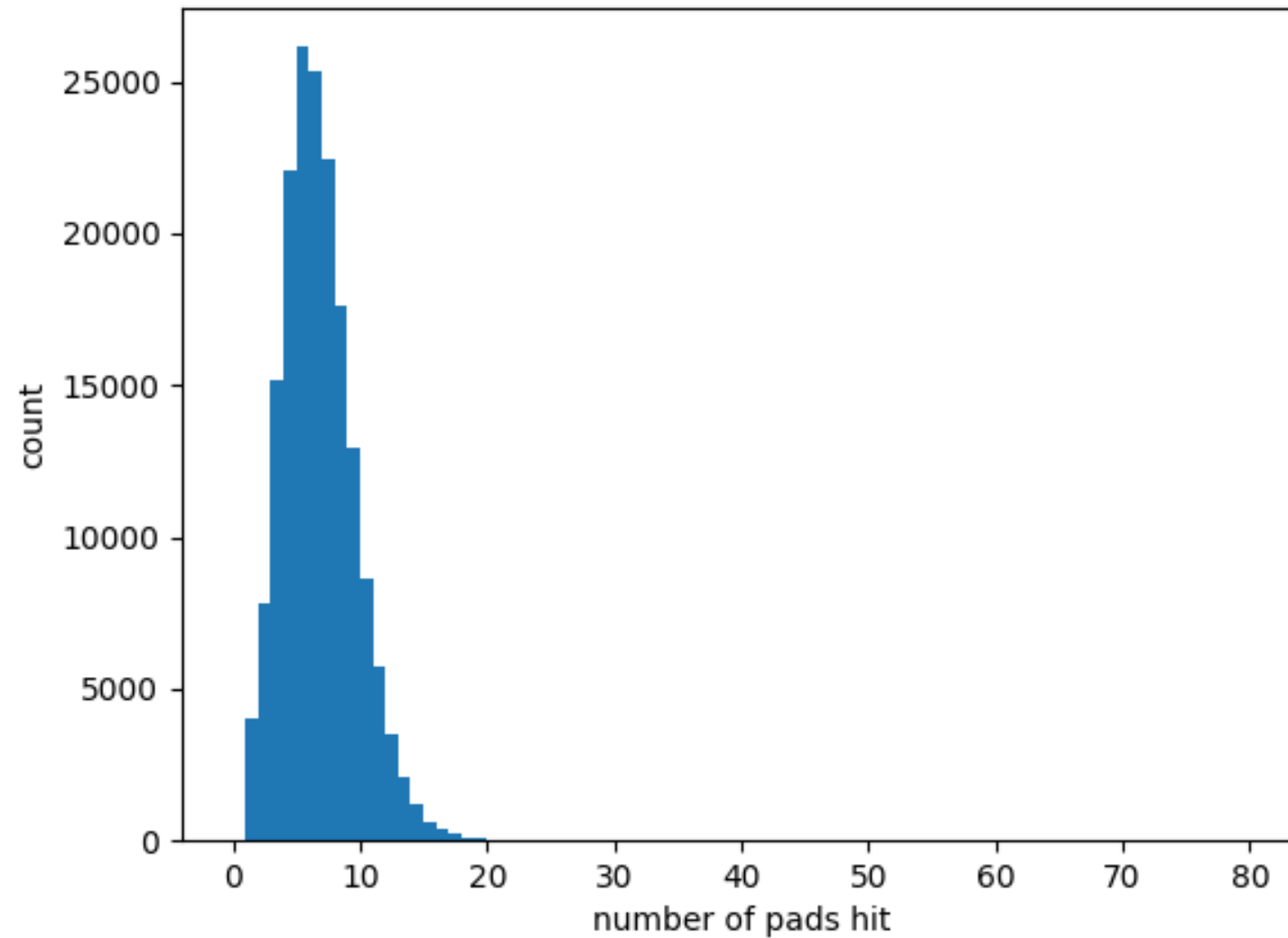
CALICE

Layer 2



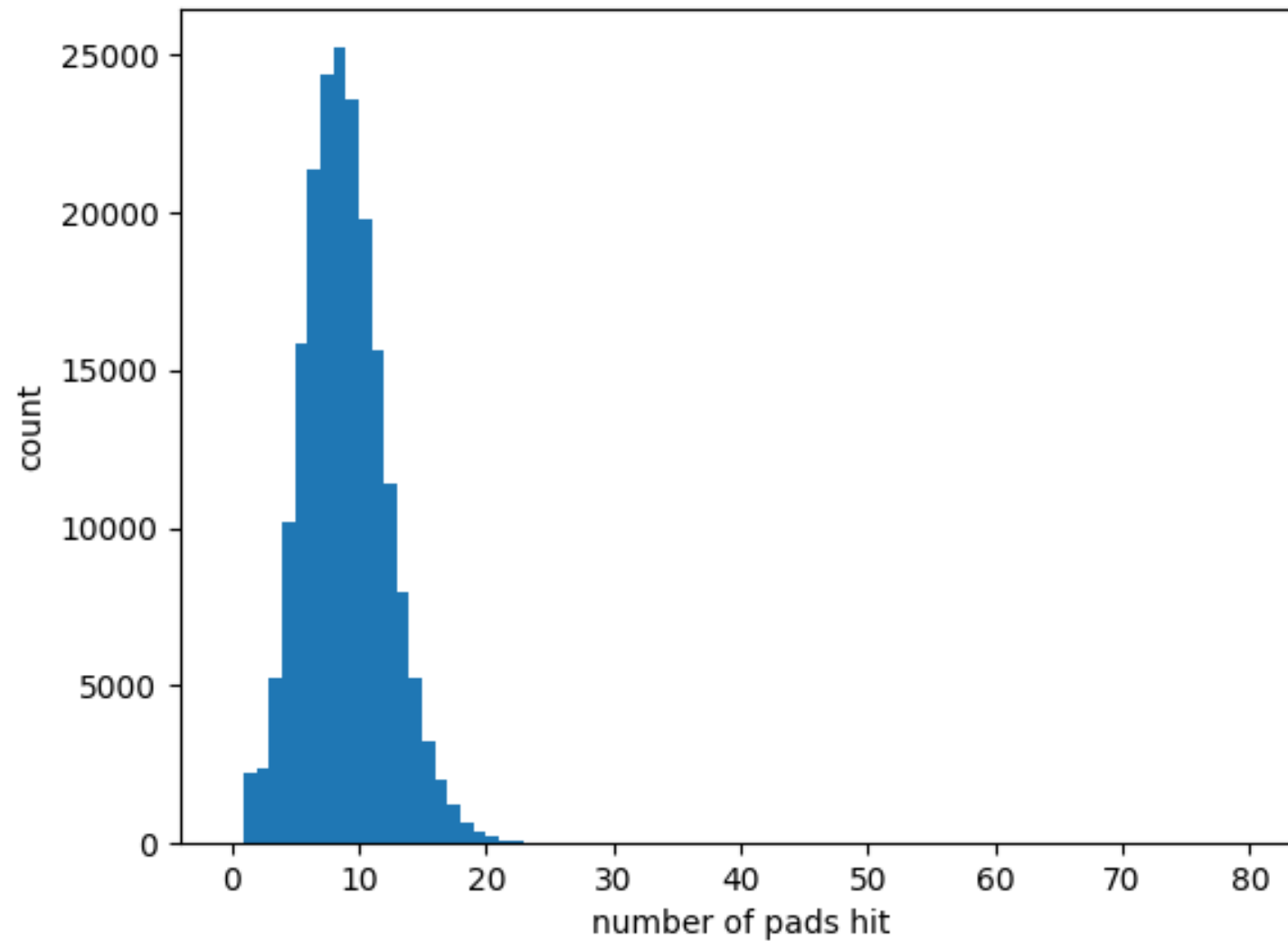
CALICE

Layer 3



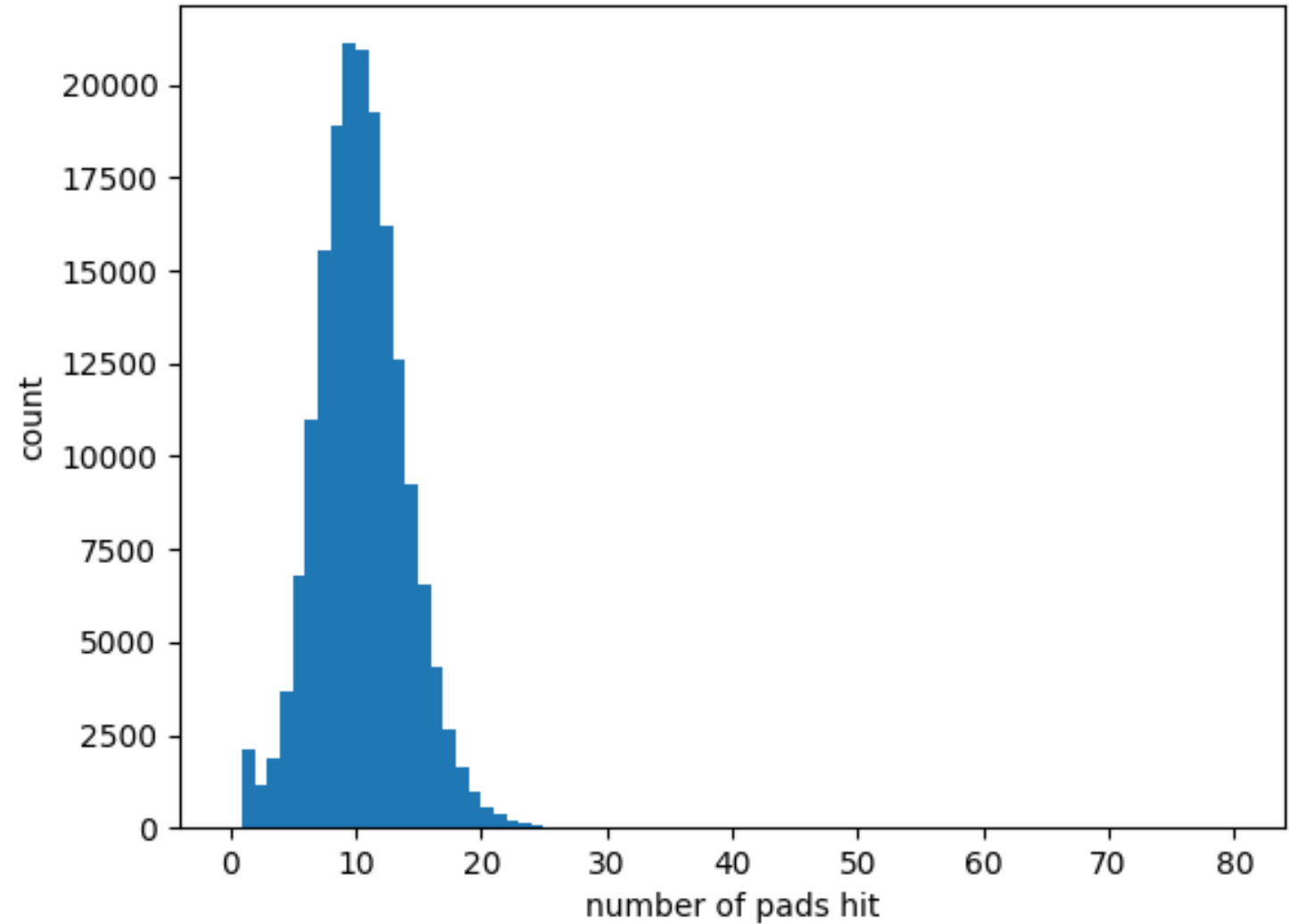
CALICE

Layer 4



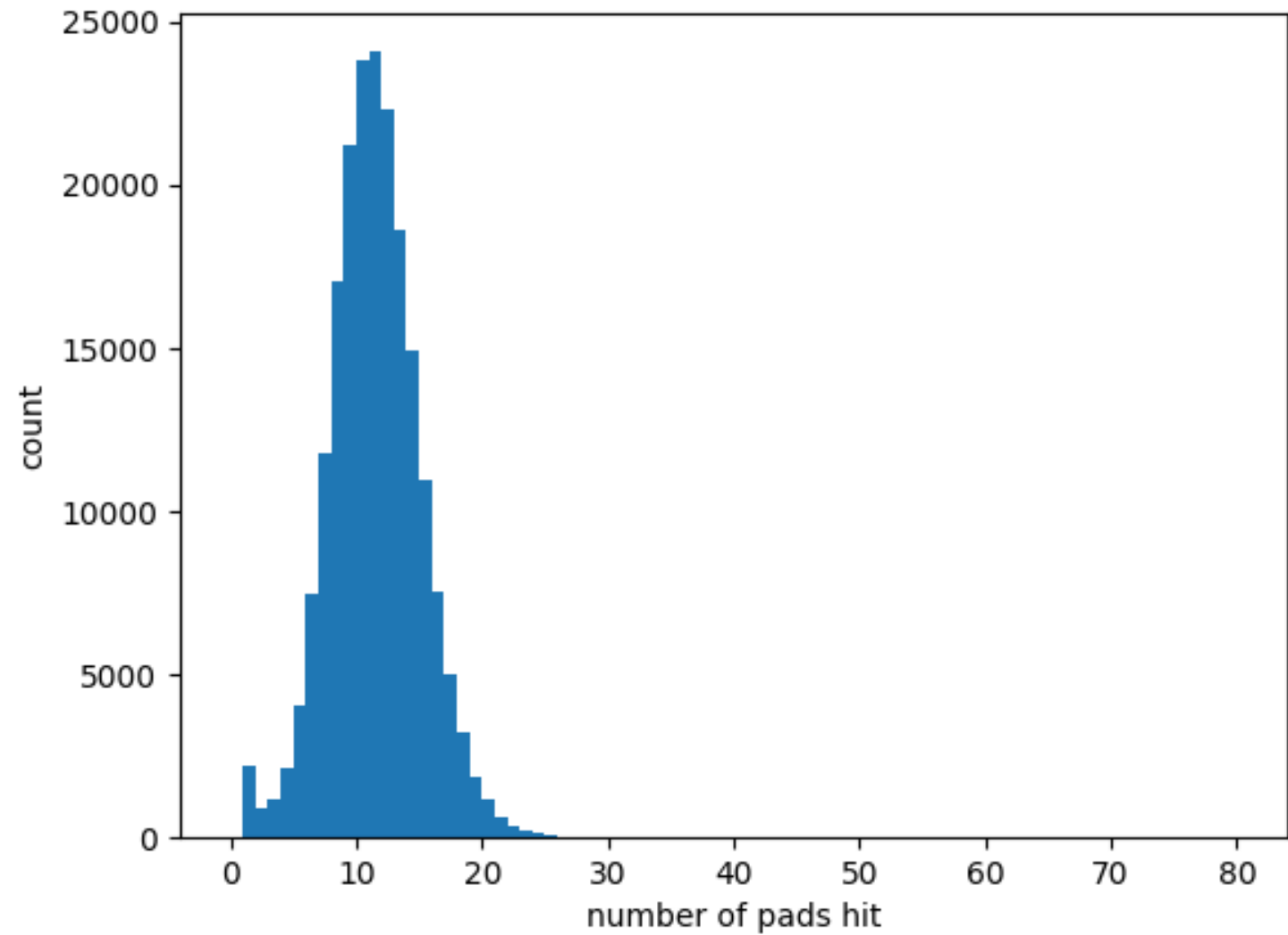
CALICE

Layer 5



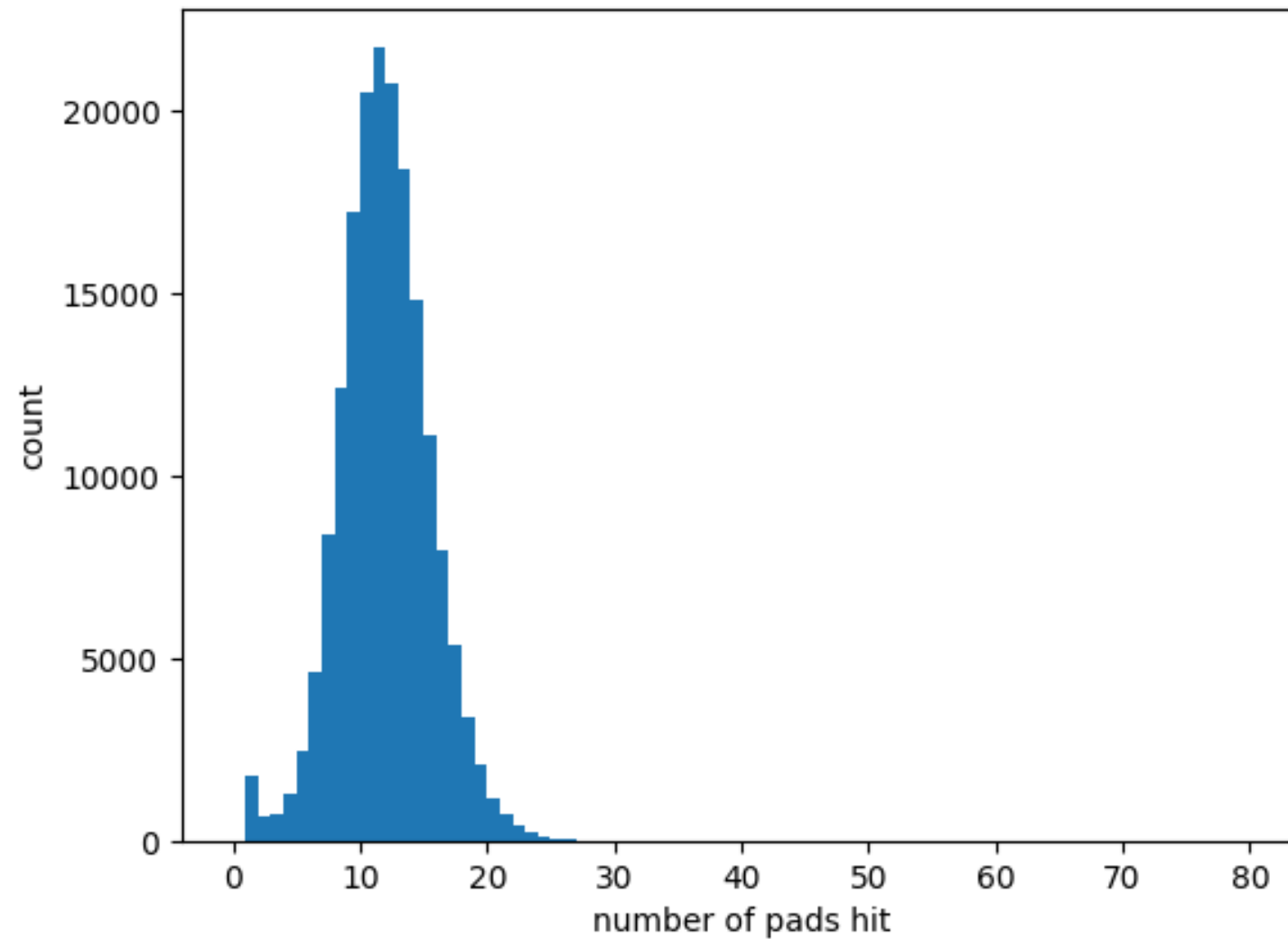
CALICE

Layer 6



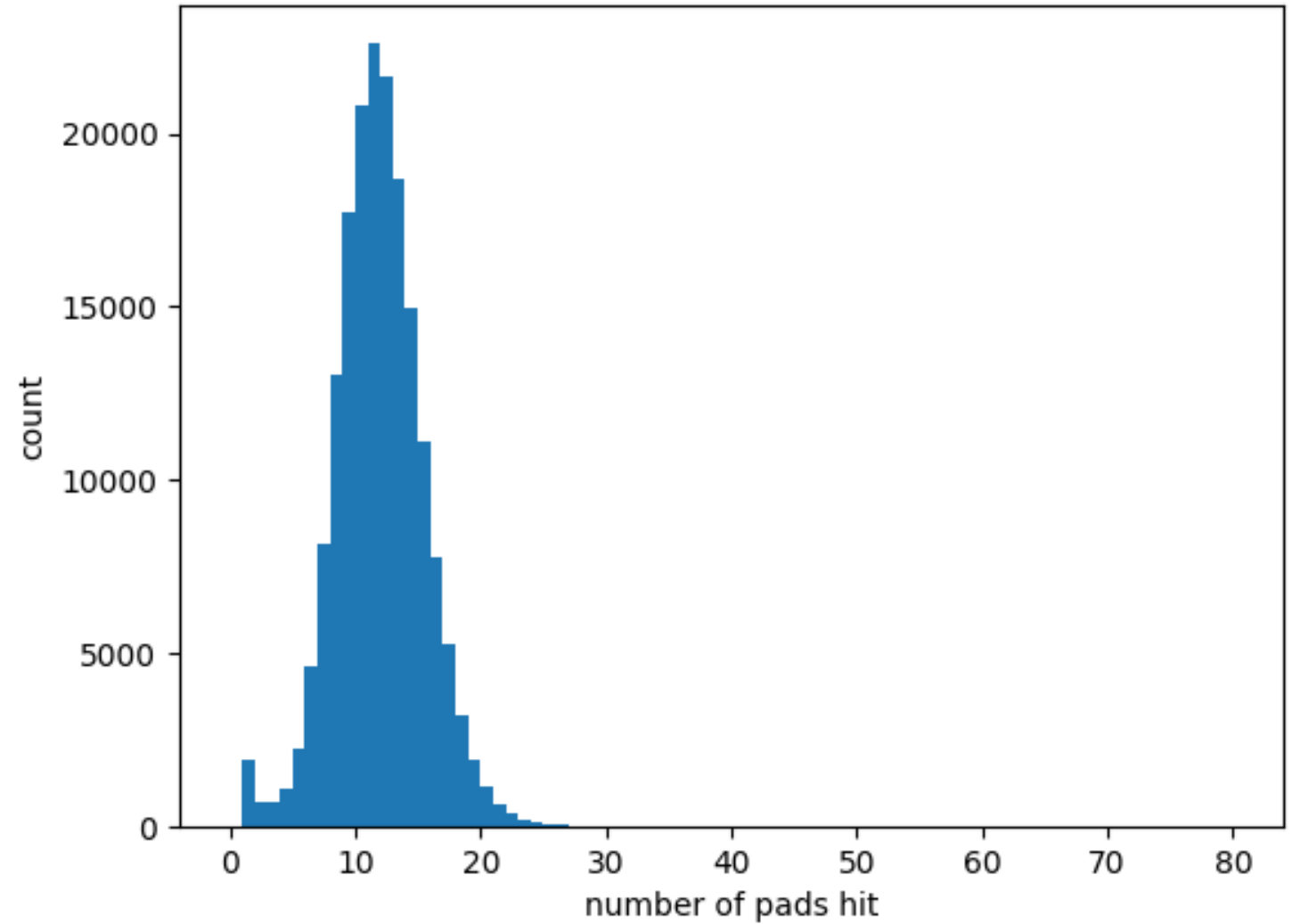
CALICE

Layer 7



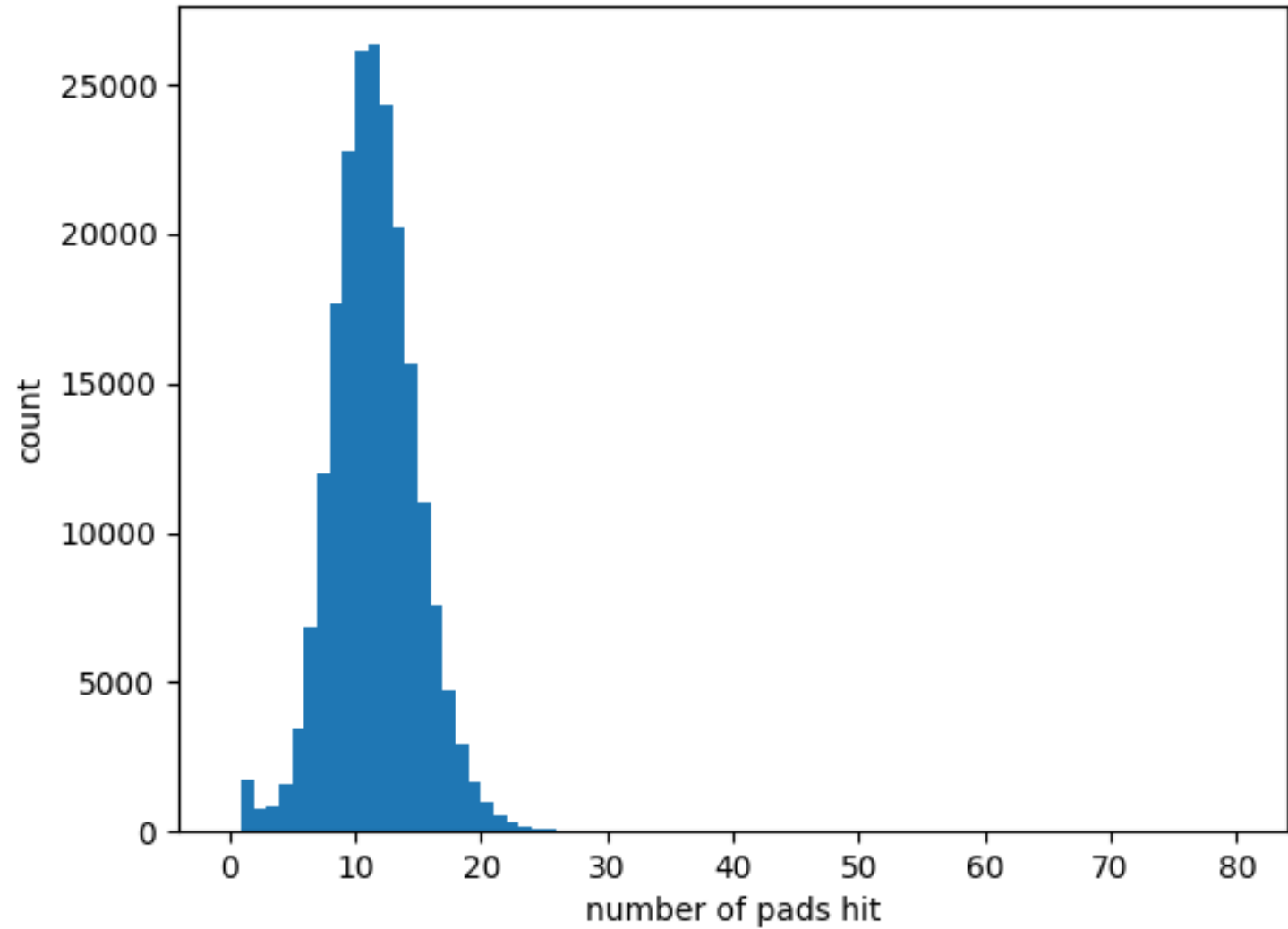
CALICE

Layer 8



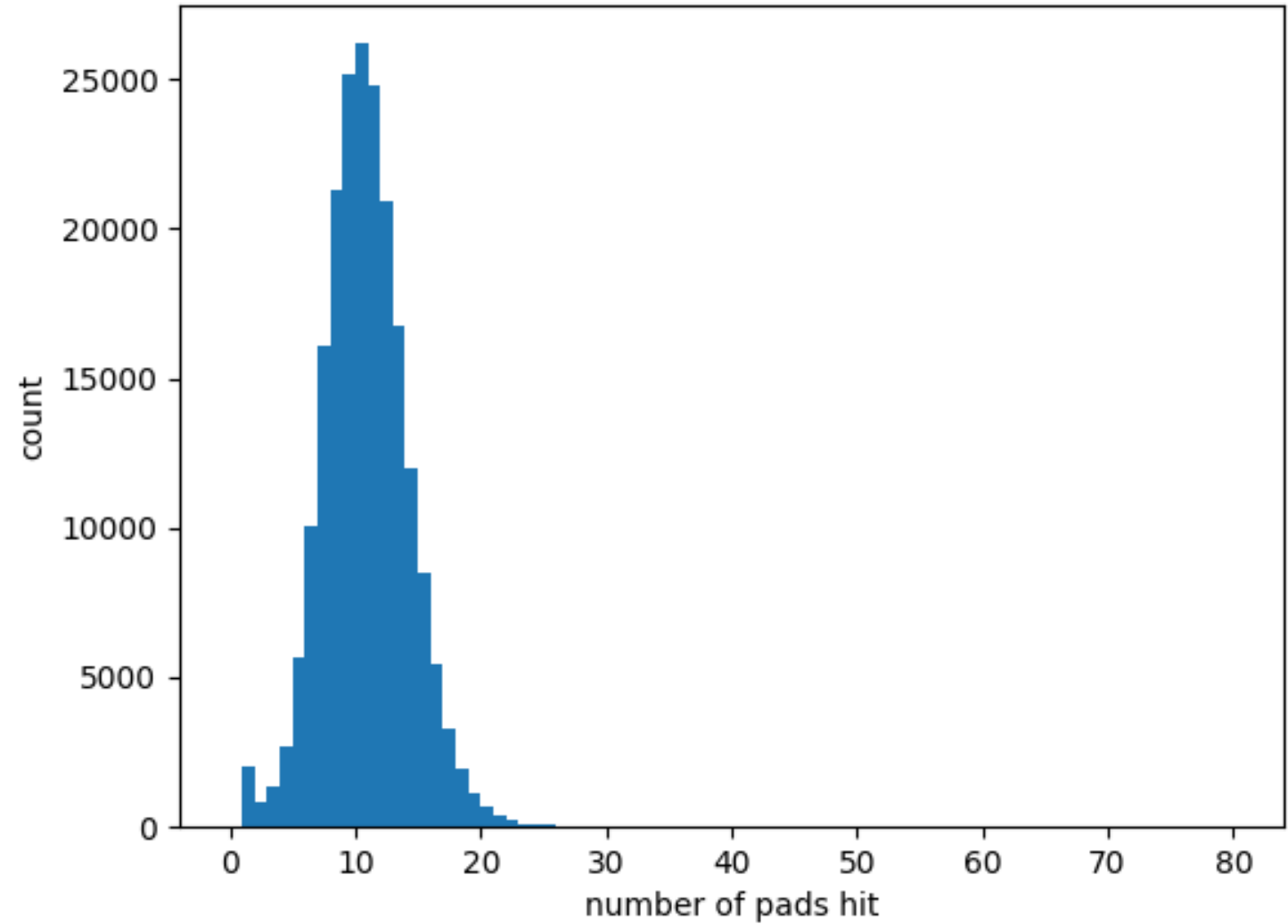
CALICE

Layer 9



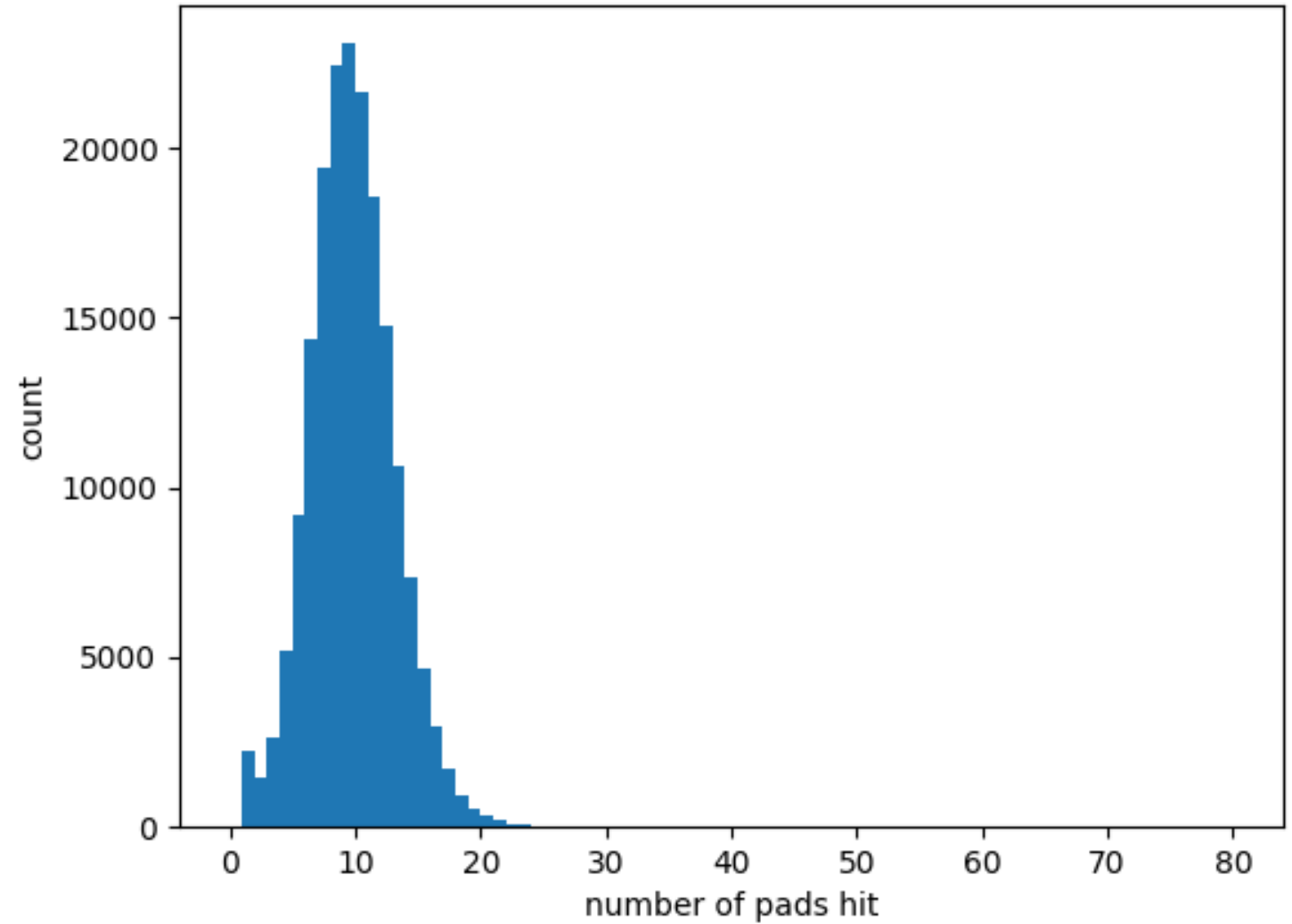
CALICE

Layer 10



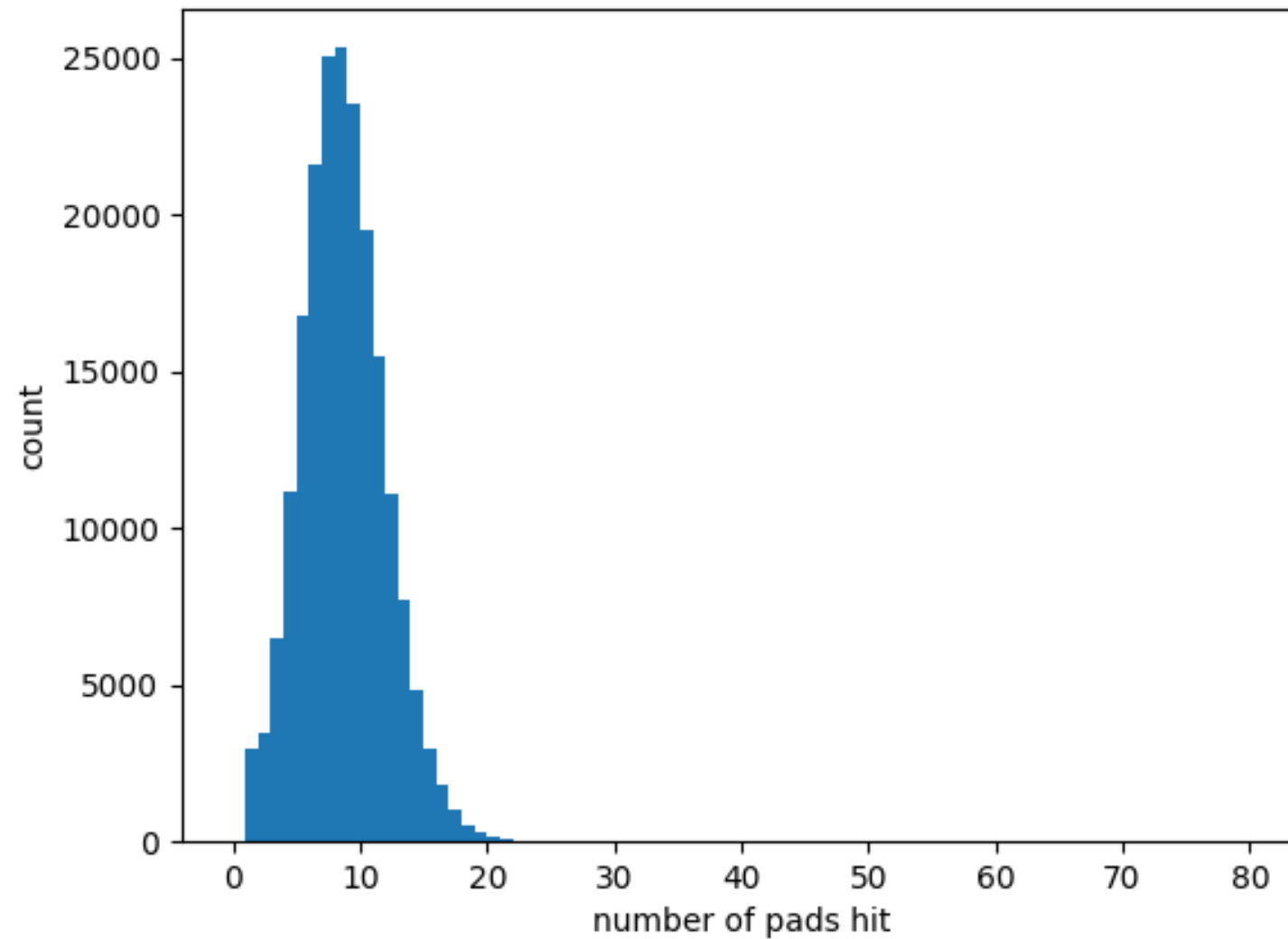
CALICE

Layer 11



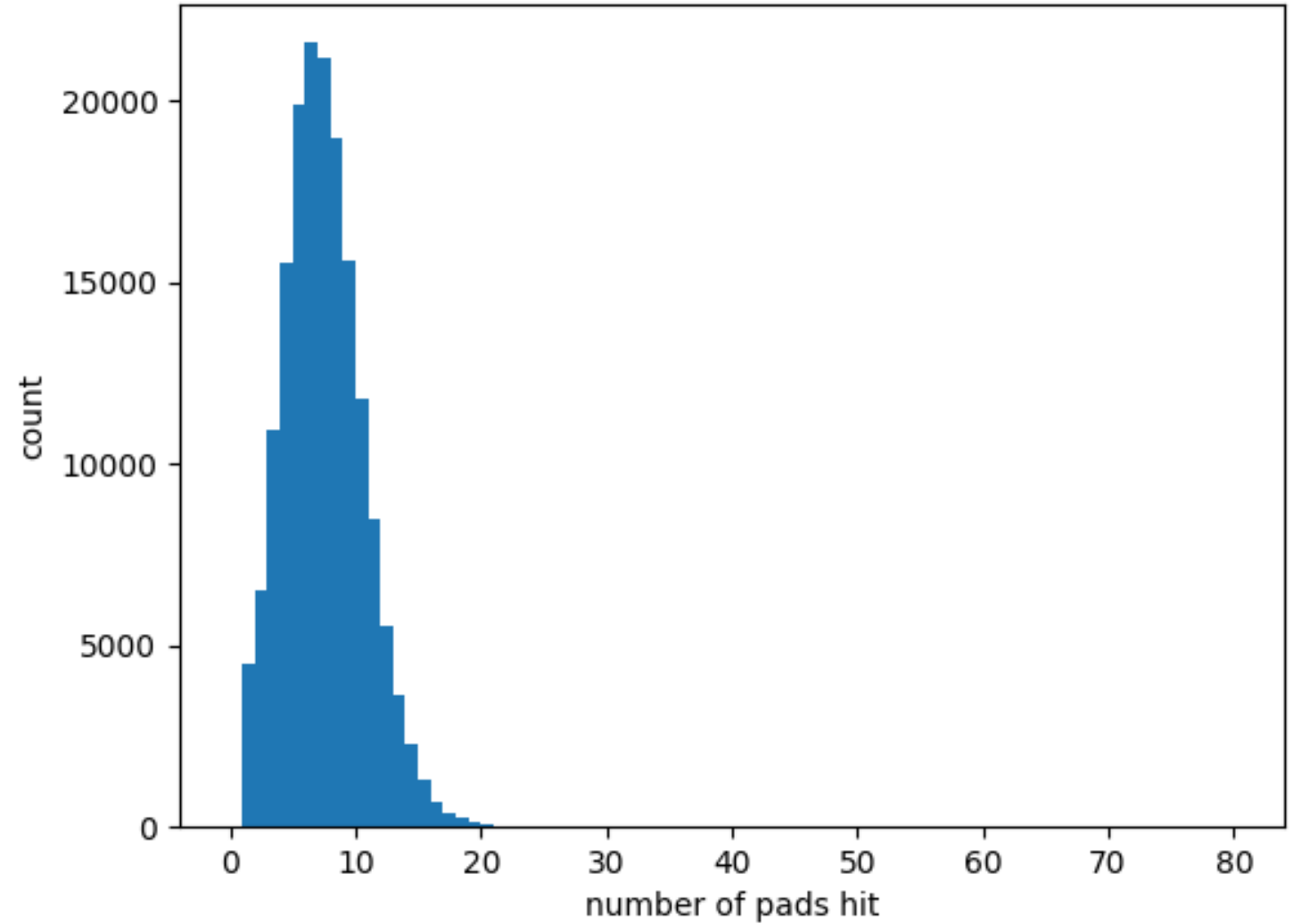
CALICE

Layer 12



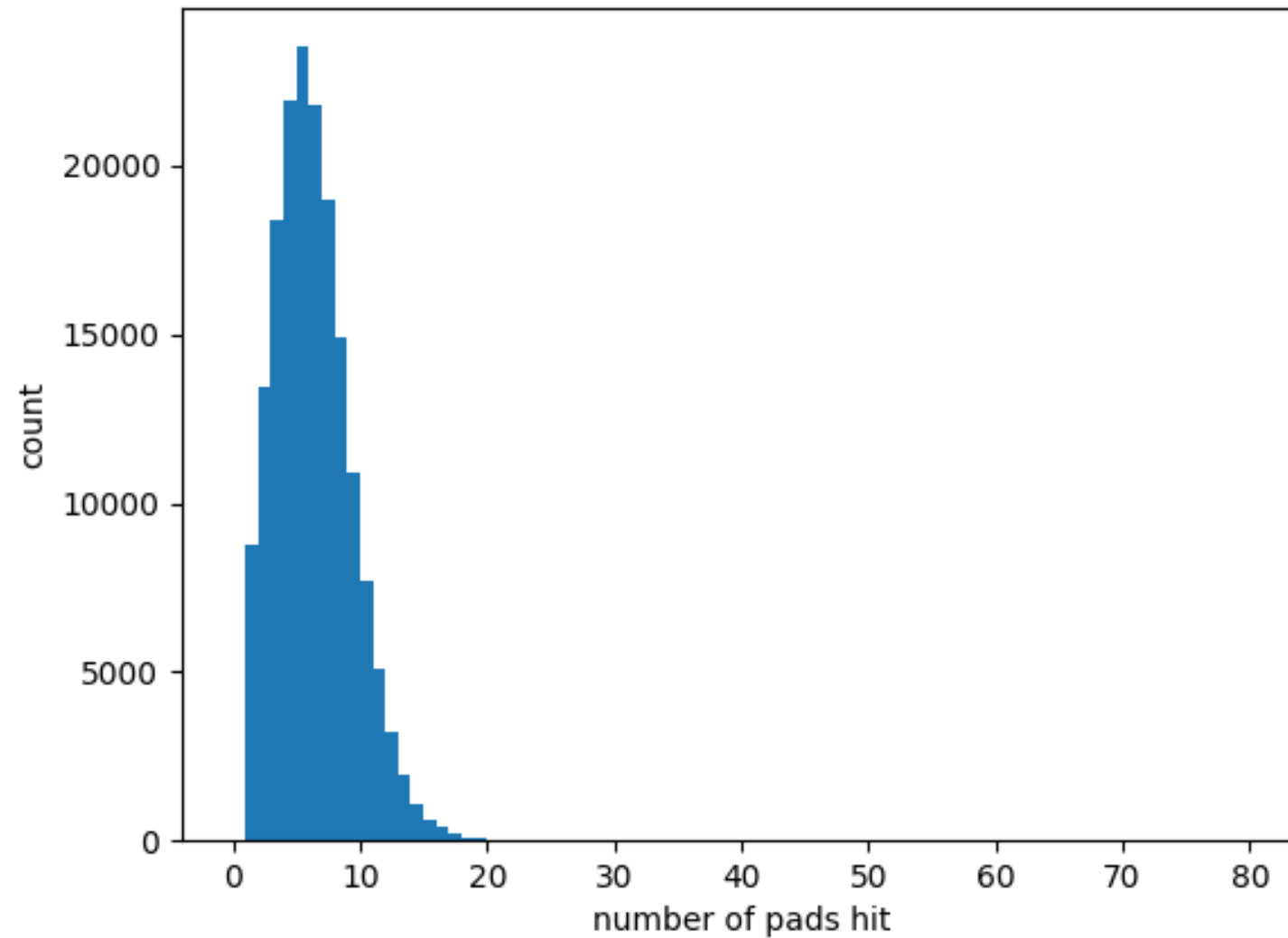
CALICE

Layer 13



CALICE

Layer 14



CALICE

Layer 15

