

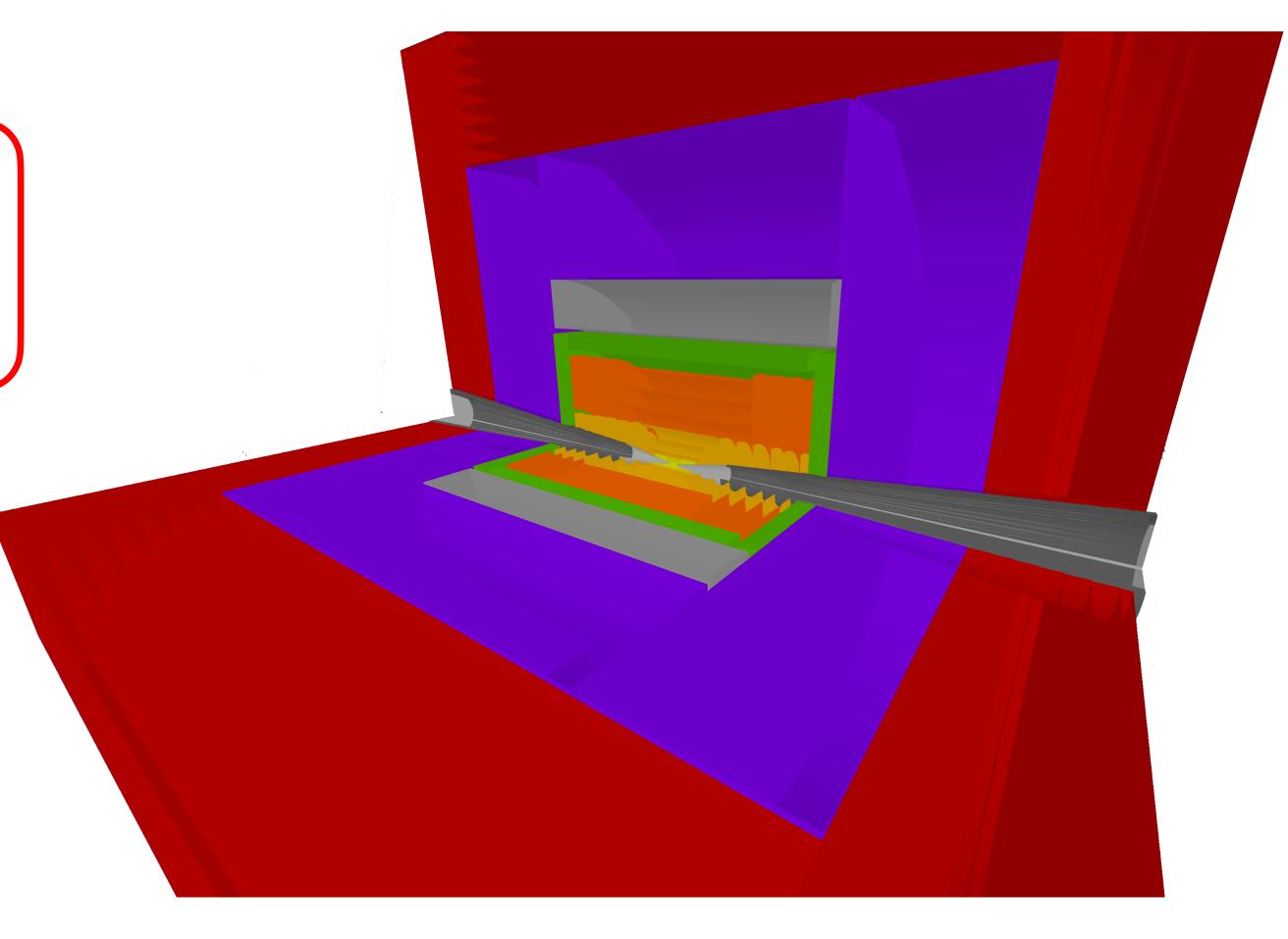






• Important effort in the last year, in particular:

- Understanding the differences between Fluka and MARS frameworks
- Assessment of BIB+IPP impact on detector
- Reconstruction performance for different physics objects
- Question: what did we learn?
- Here: some feedbacks from MUSIC detector point of view

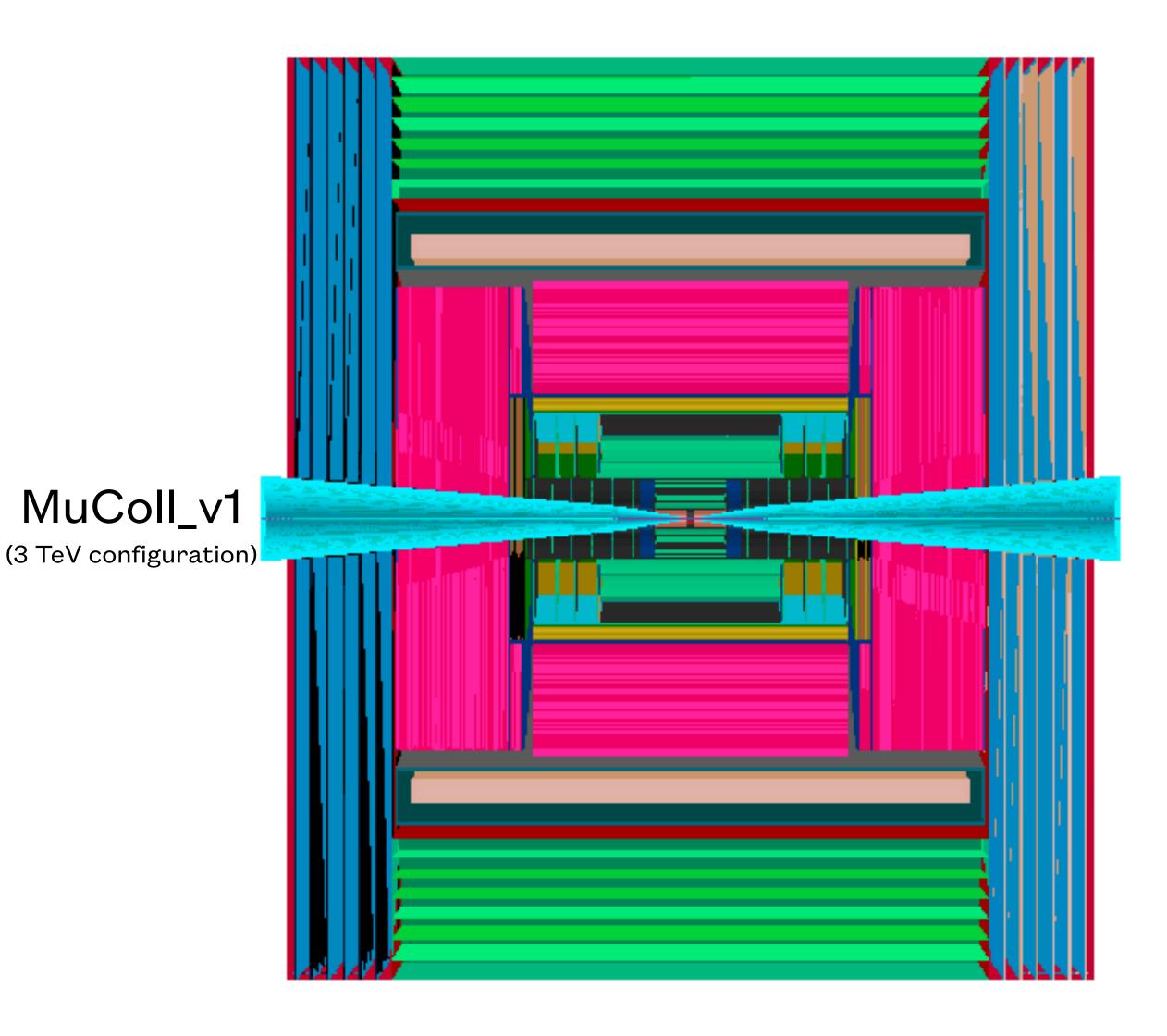


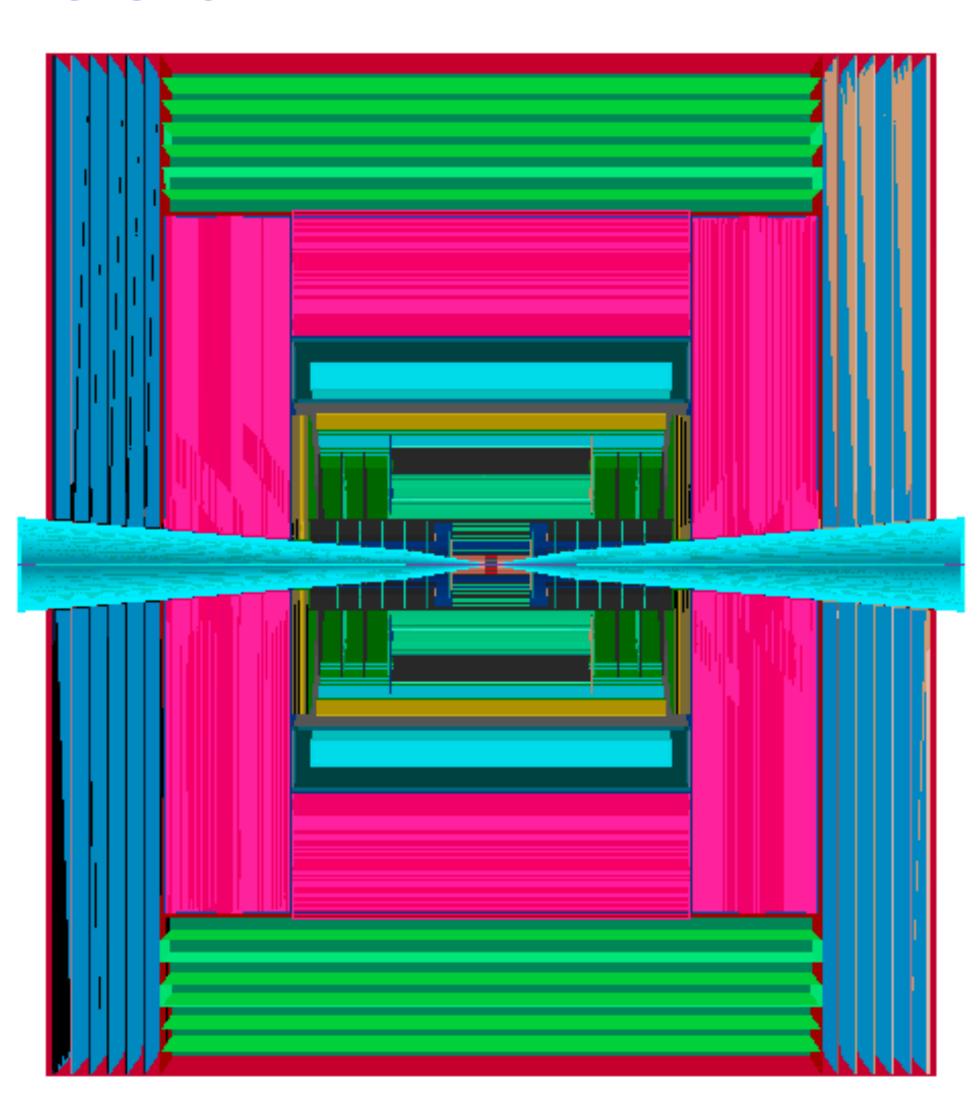






From where we started

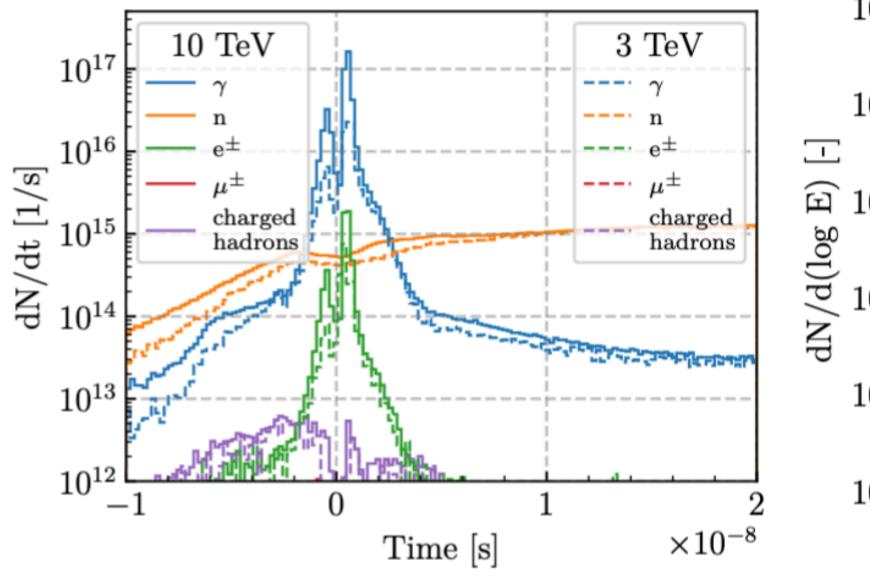


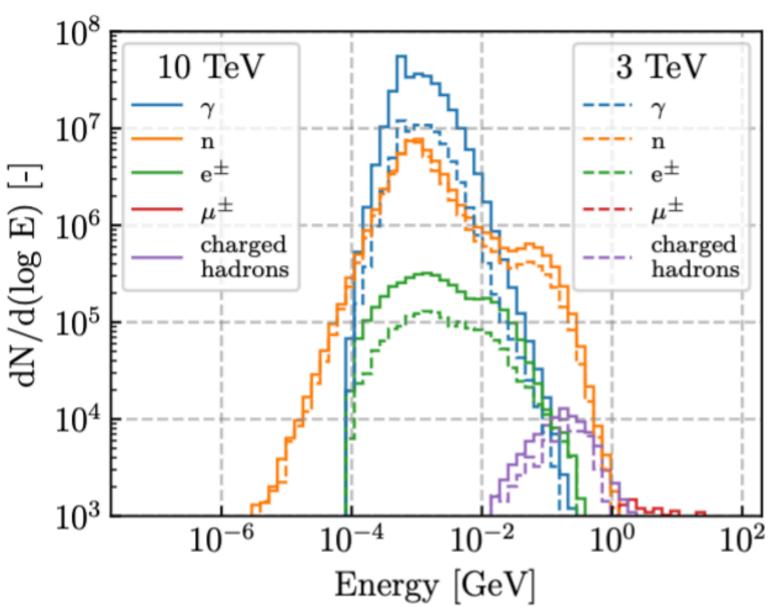


MUSIC

Change from 3 to 10 TeV

- First of all, change from 3 to 10 TeV has been not negligible
- In particular:
 - More energetic photons and electrons
 - More photons and electrons on time w.r.t. bunch crossing
 - Also, IPP now are present and not negligible



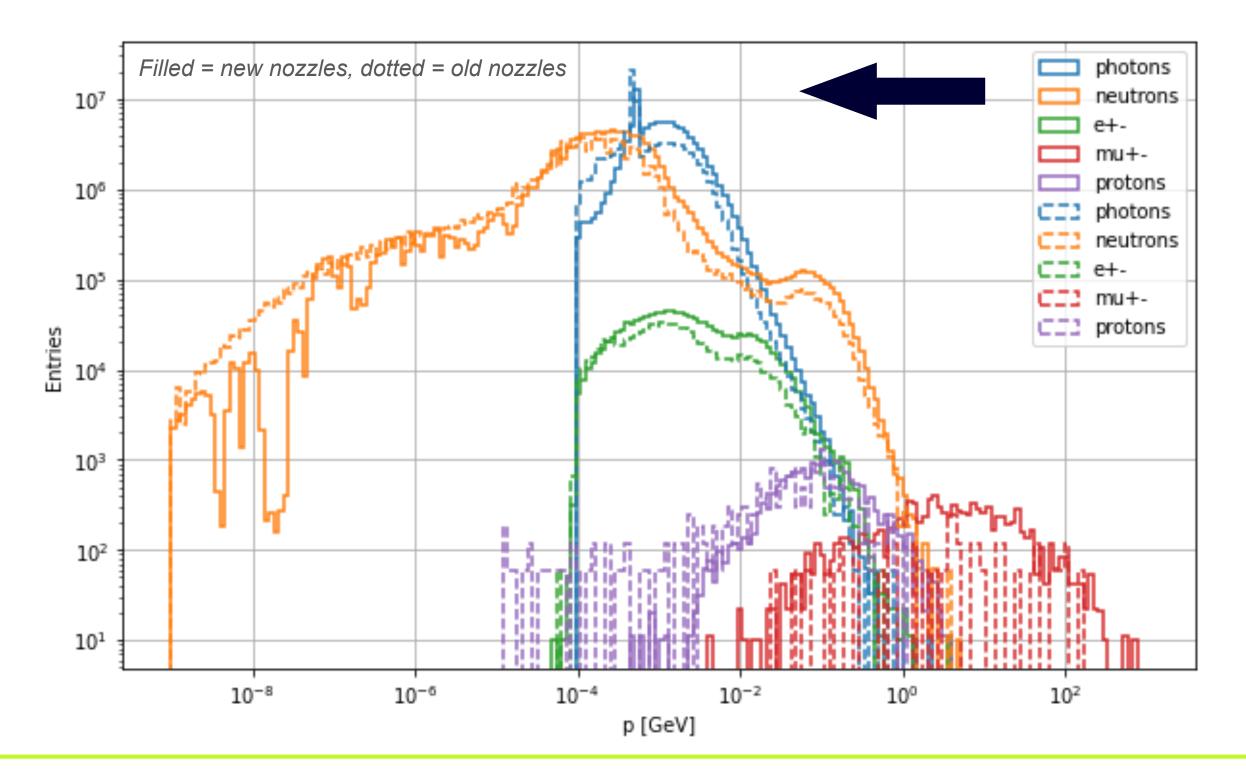


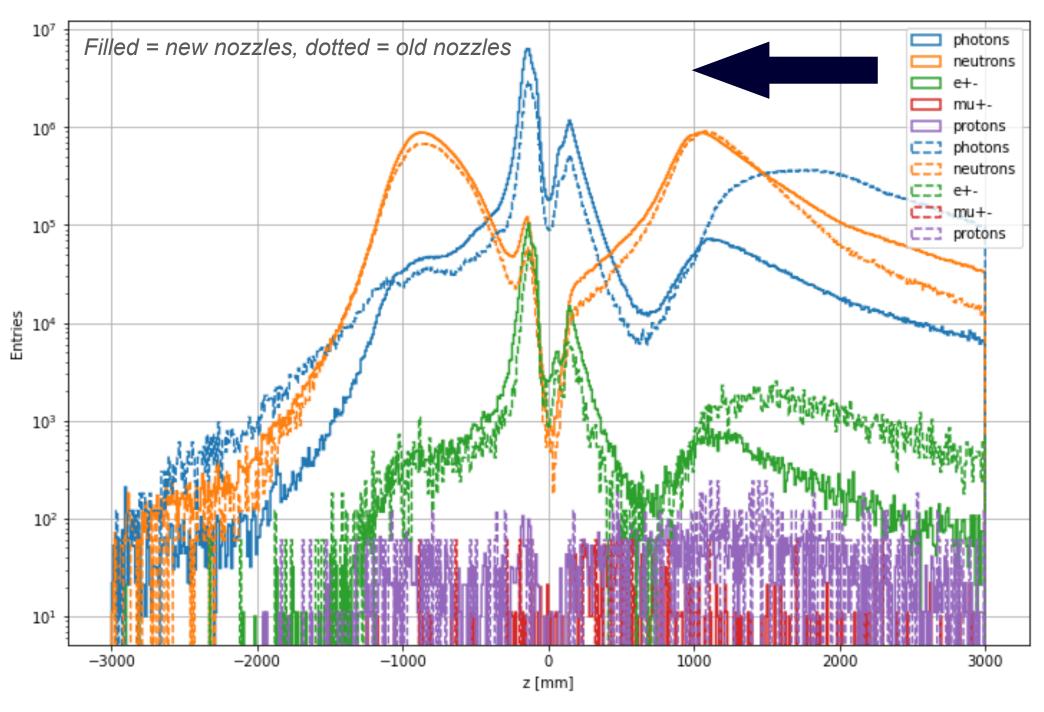




Change from v0.4 to v.08 lattices

- Changes to last lattice version has not been "painless"
 - Increase in number of photons close to interaction point
 - More energetic photons and neutrons



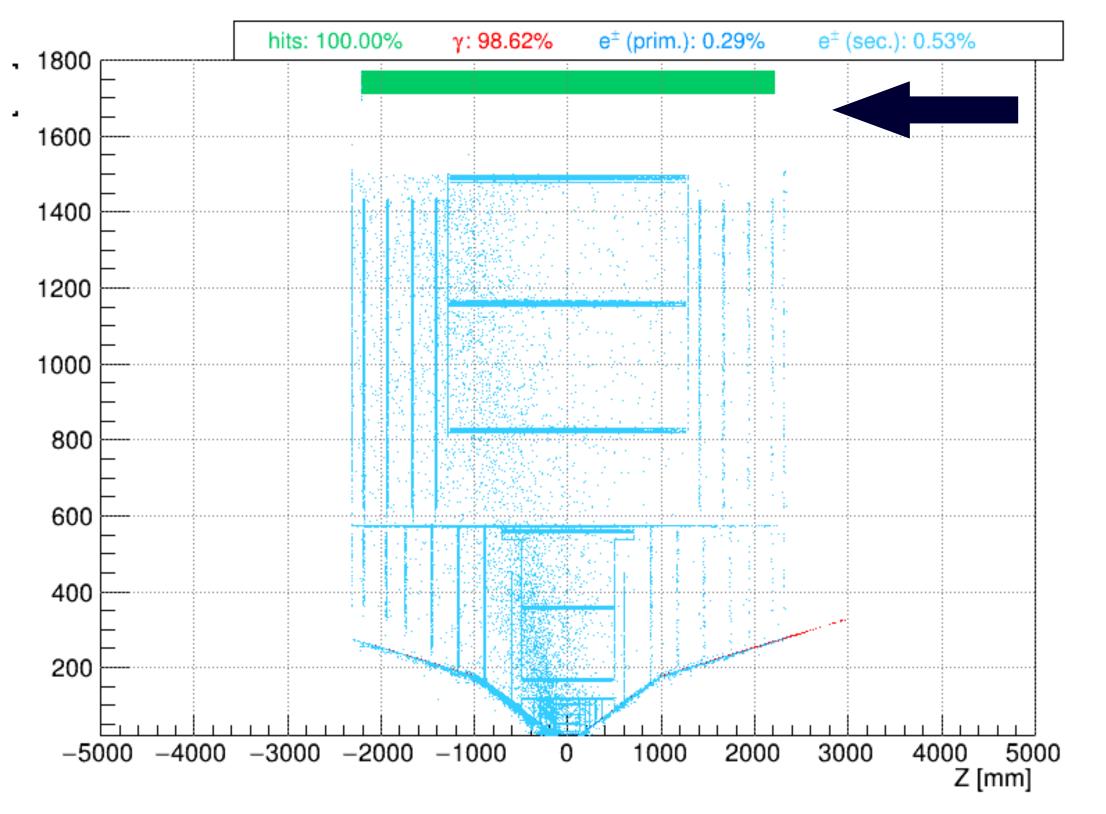


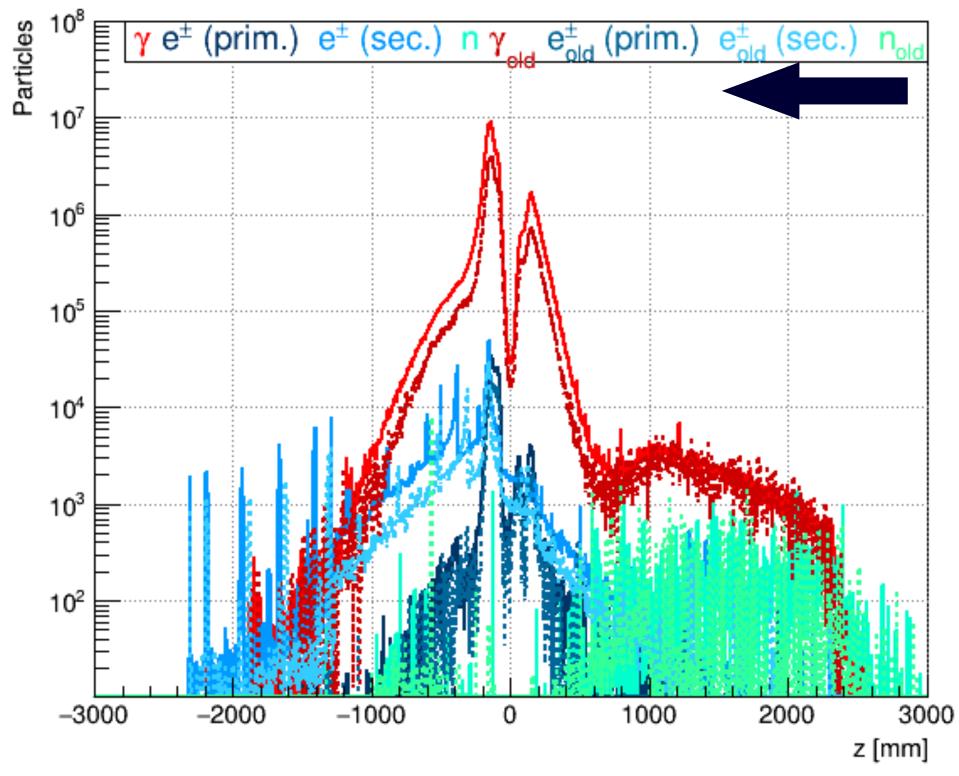




What happens in the detector?

- Let's have a look at particles generating hits in the ECAL
- Mainly primary photons exiting from the tip of the nozzles



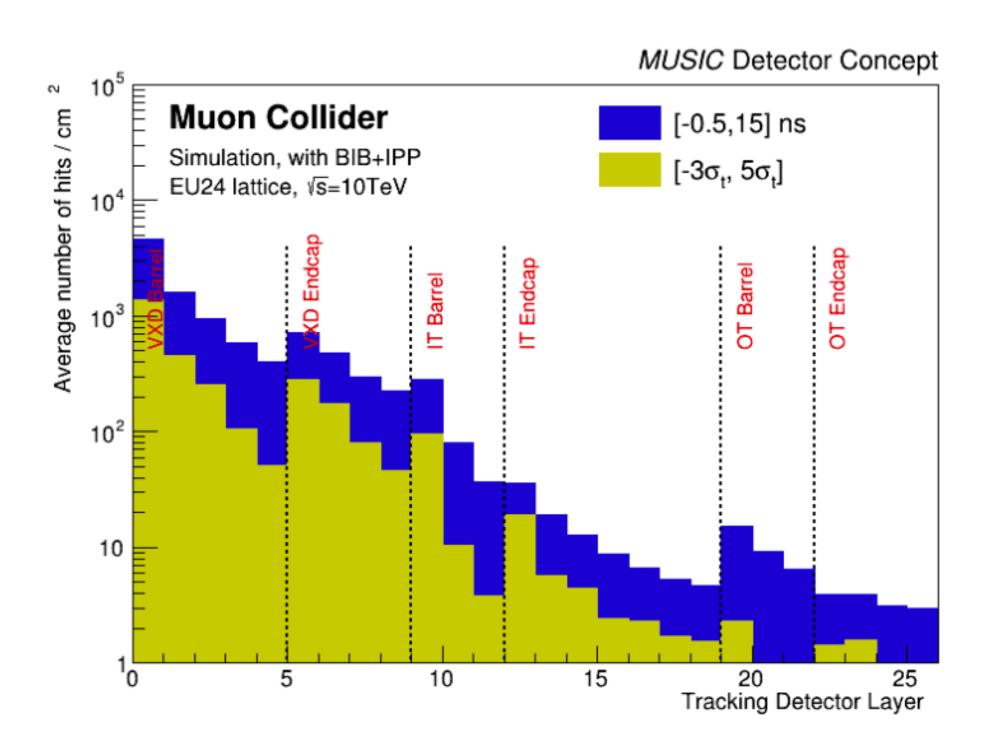


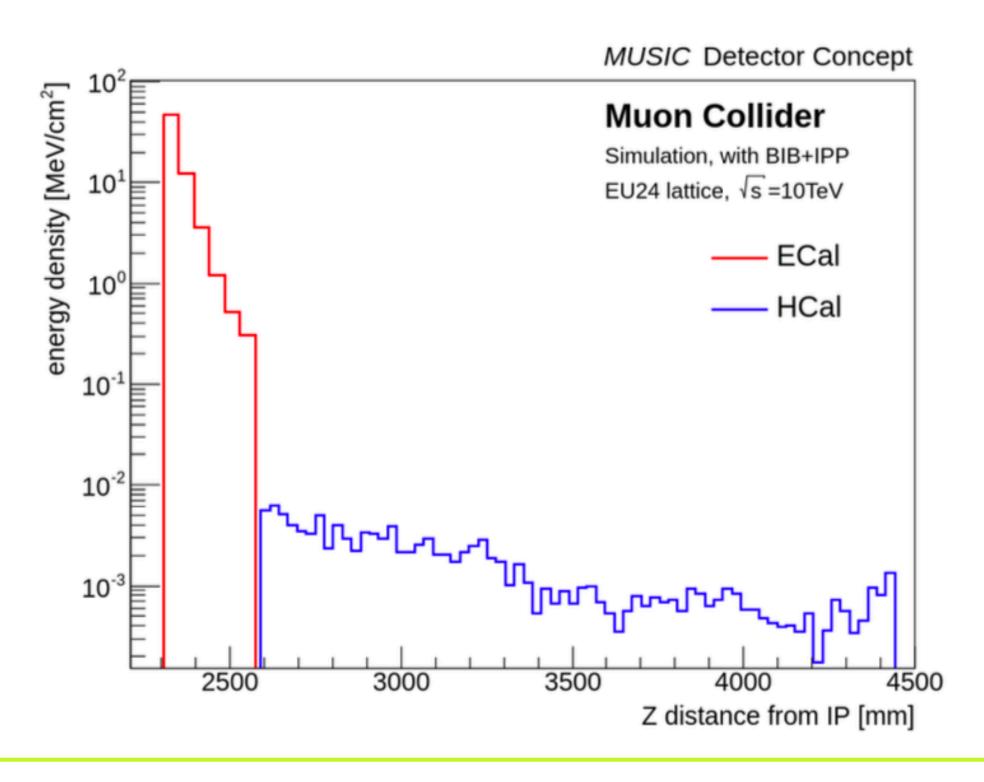




What happens in the detector?

- Very naively:
 - Increase in occupancy in the tracker
 - Higher diffuse background in the calorimeters (particularly ECAL)



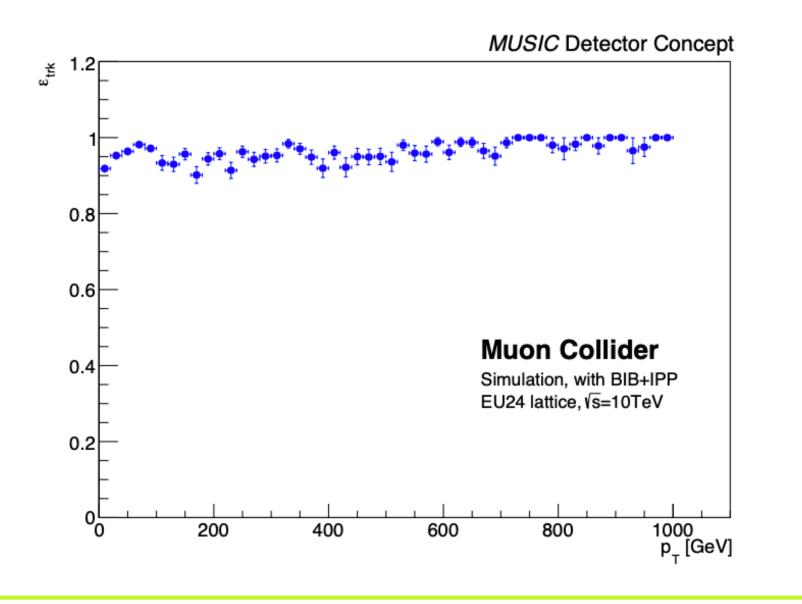


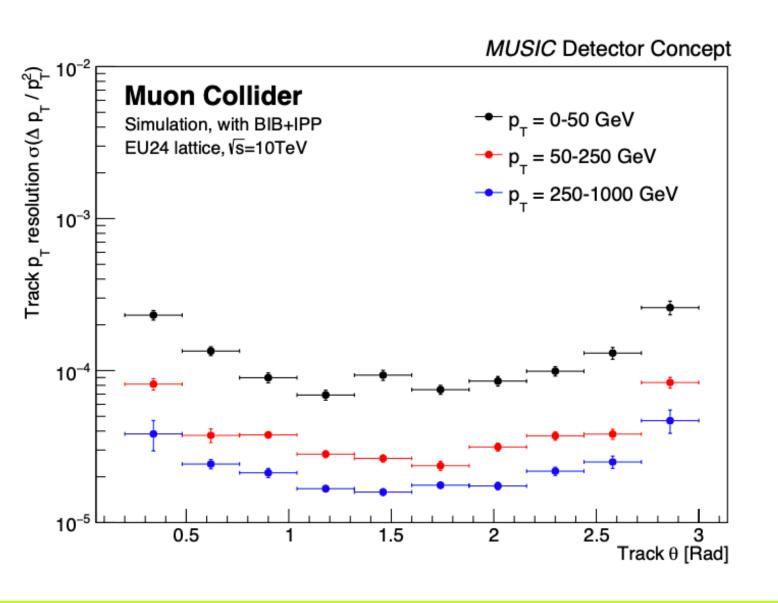


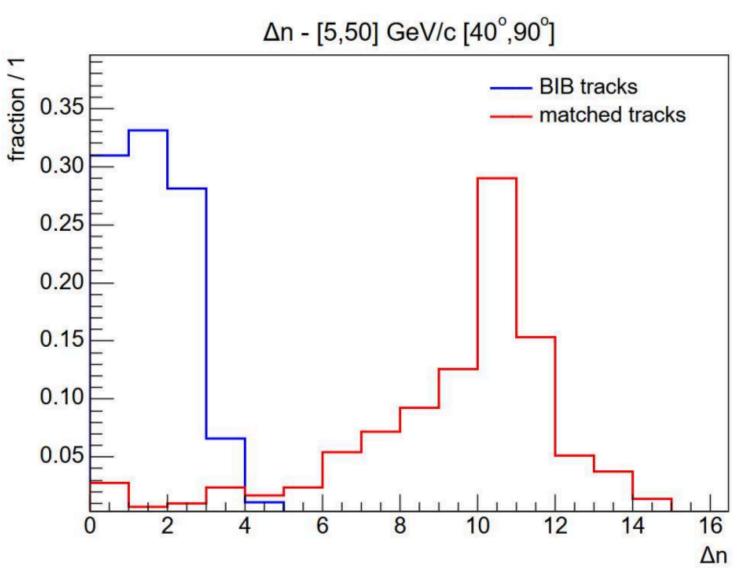


Tracker

- Old ACTS tuning was completely unsustainable
 - Tune ACTS in a better way, getting rid of the first VXD layer to perform seeding
 - Important parallelisation of algorithms Anyway, huge number of tracks per event (~1M)
 - Important to find interesting figures of merit to filter tracks









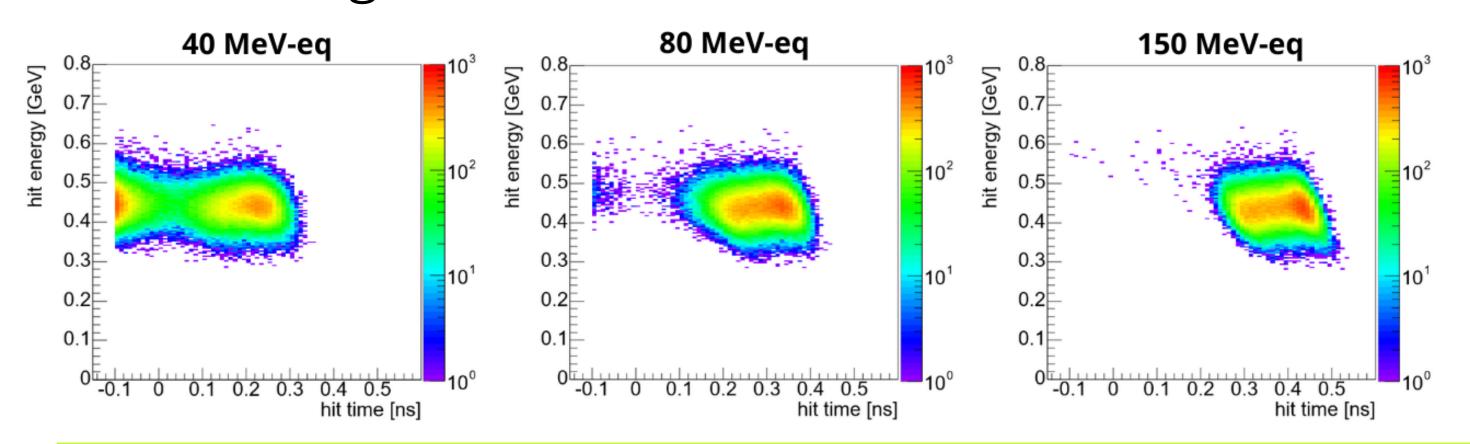


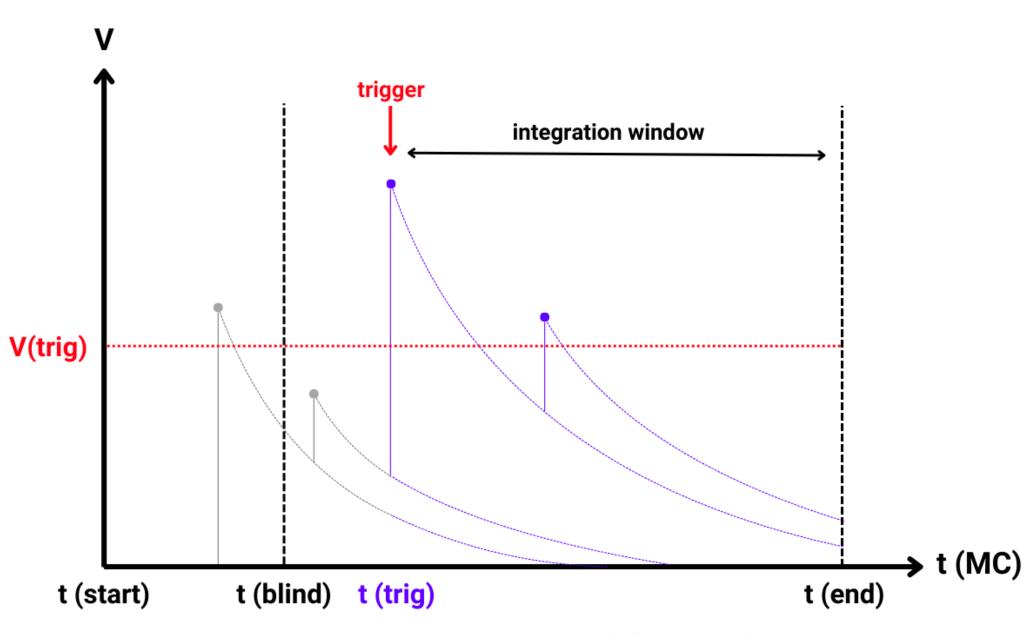


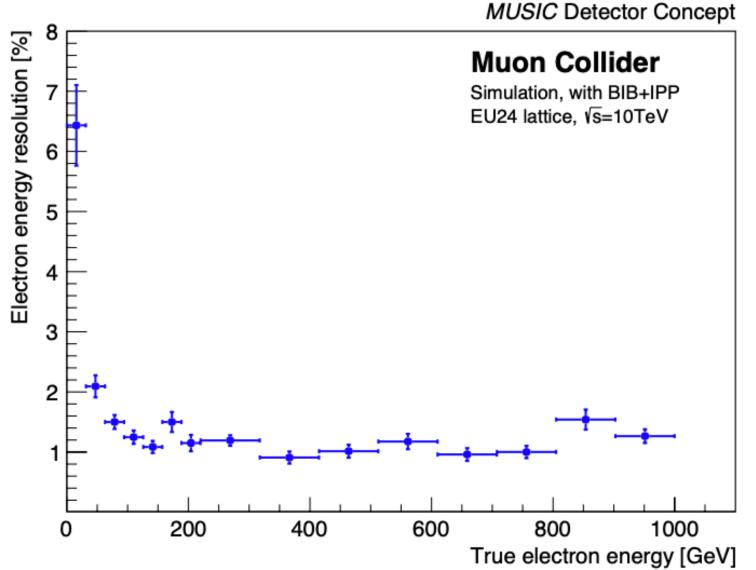


ECAL

- Given the increase in photons, reconstruction in ECAL has become challenging
 - We implemented a signal-threshold digitisation based on hit arrival time and energy
 - Depending also on ECAL region and layer
 - Tuning of hit clustering algorithm by PandoraPFA
- Able to get the same performance that were achieved in 3 TeV configuration











Conclusions

- Going from 3 TeV to 10 TeV, and from v0.4 to v0.8 lattice versions has changed the impact of BIB on detector
- Higher occupancies in the tracker and in ECAL
 - At the beginning it was very difficult to handle files and tools
- New digi/reco techniques and tuning of algorithms (ACTS and PandoraPFA) allowed us to get back to 3 TeV performance (sometimes even slightly better!)
- (personal opinion) Not so much we can do to improve further with present algorithms
- The most "problematic" region is close to the interaction point
- Is there something we can do about it?

