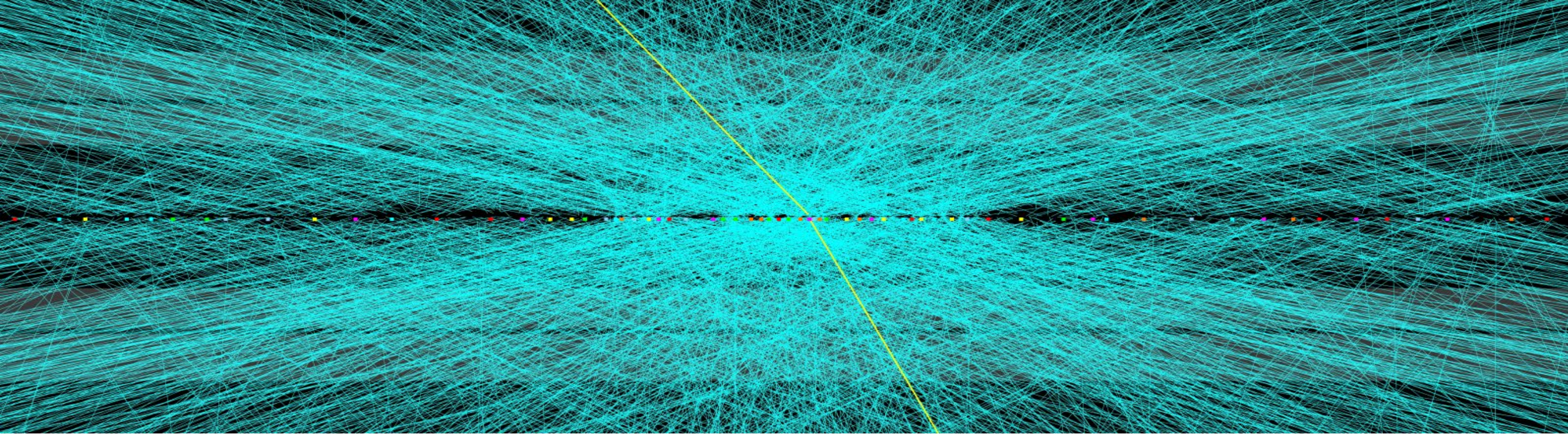




Luminosity studies with Pixel Cluster Counting at the ATLAS experiment

DESY Summer Student Programme

Oleksandra Khrul

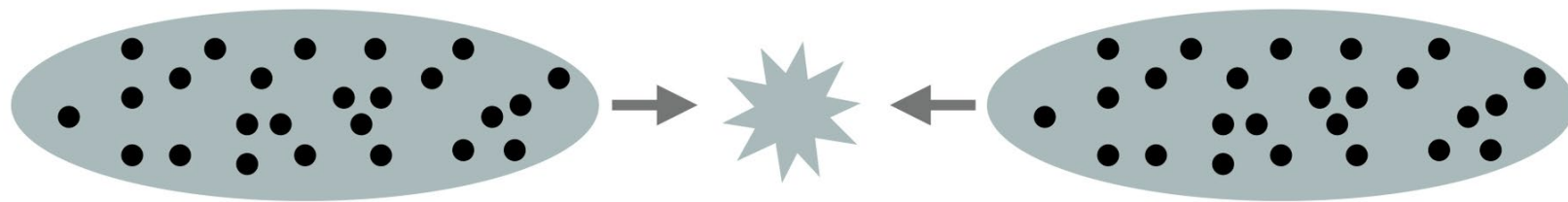
Supervisors: Filippo Dattola, Claudia Seitz



Introduction

Luminosity basics

- Luminosity ($\mathcal{L}$) measures **number of proton collisions** produced by LHC.



- Determining luminosity with high precision is **crucial for cross section measurements**:

$$\sigma(pp \rightarrow X) = \frac{N(pp \rightarrow X)}{\int \mathcal{L} dt}$$

- Luminosity uncertainty dominant in $t\bar{t}$ /W/Z cross section measurements.

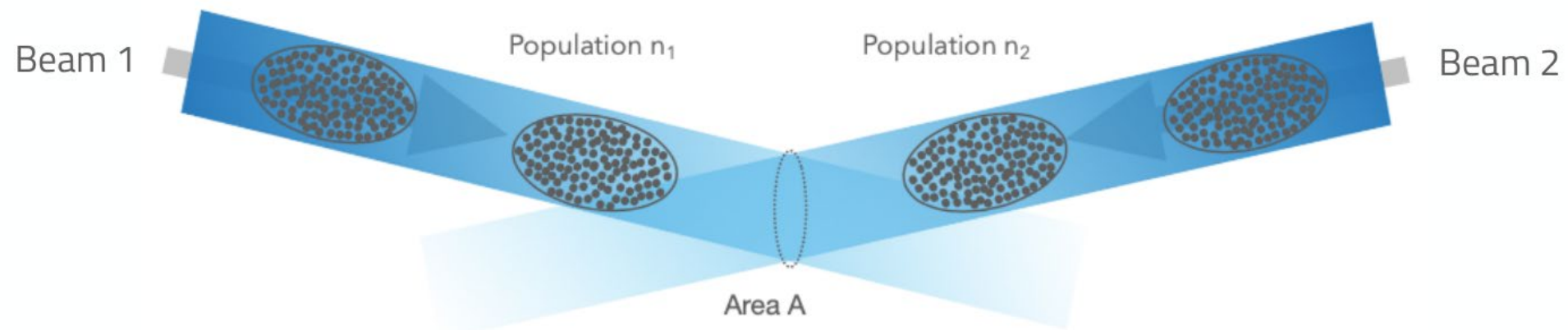
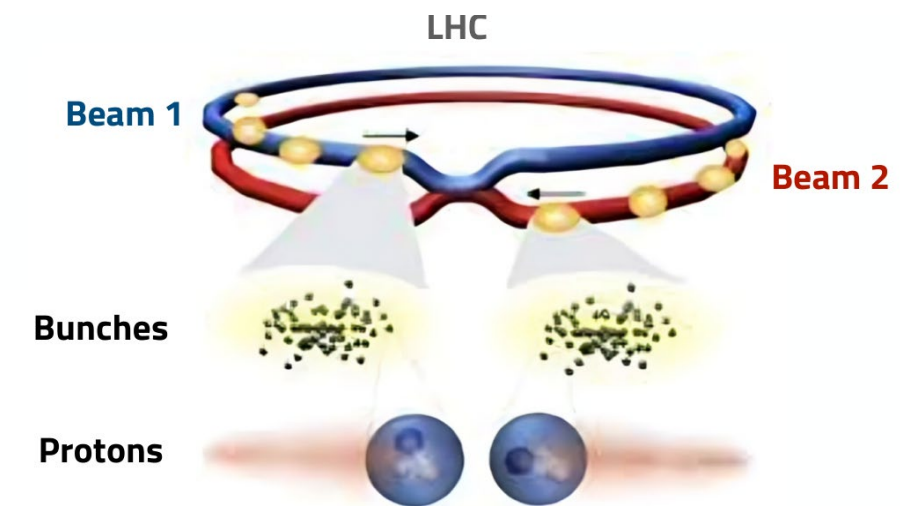
E.g. $\sigma_{t\bar{t}} = 826.4 \pm 3.6 \text{ (stat)} \pm 11.5 \text{ (syst)} \pm 15.7 \text{ (lumi)} \pm 1.9 \text{ (beam) pb}$, [\[Eur.Phys.J.C 80 \(2020\) 528\]](#)

- Record **0.83% uncertainty** achieved in Run 2 [\[Eur.Phys.J.C 83 \(2023\) 10, 982\]](#)

Introduction

Luminosity basics

- How do LHC beams look like?



- **Trains** of filled bunches: several consecutive filled slots.



- **Isolated** bunches: filled bunch with several empty ones before and after.

Introduction

Luminosity basics

- $\mathcal{L}$ is related to the rate of observed events and LHC parameters.

$$\mathcal{L} \propto \frac{n_1 n_2 f_r}{A} = \frac{\mu_{\text{vis}} f_r}{\sigma_{\text{vis}}}$$

LHC beam parameters:

n_i – bunch populations

A – transverse beam size

f_r – LHC revolution frequency

μ_{vis} = visible interaction rate per bunch crossing

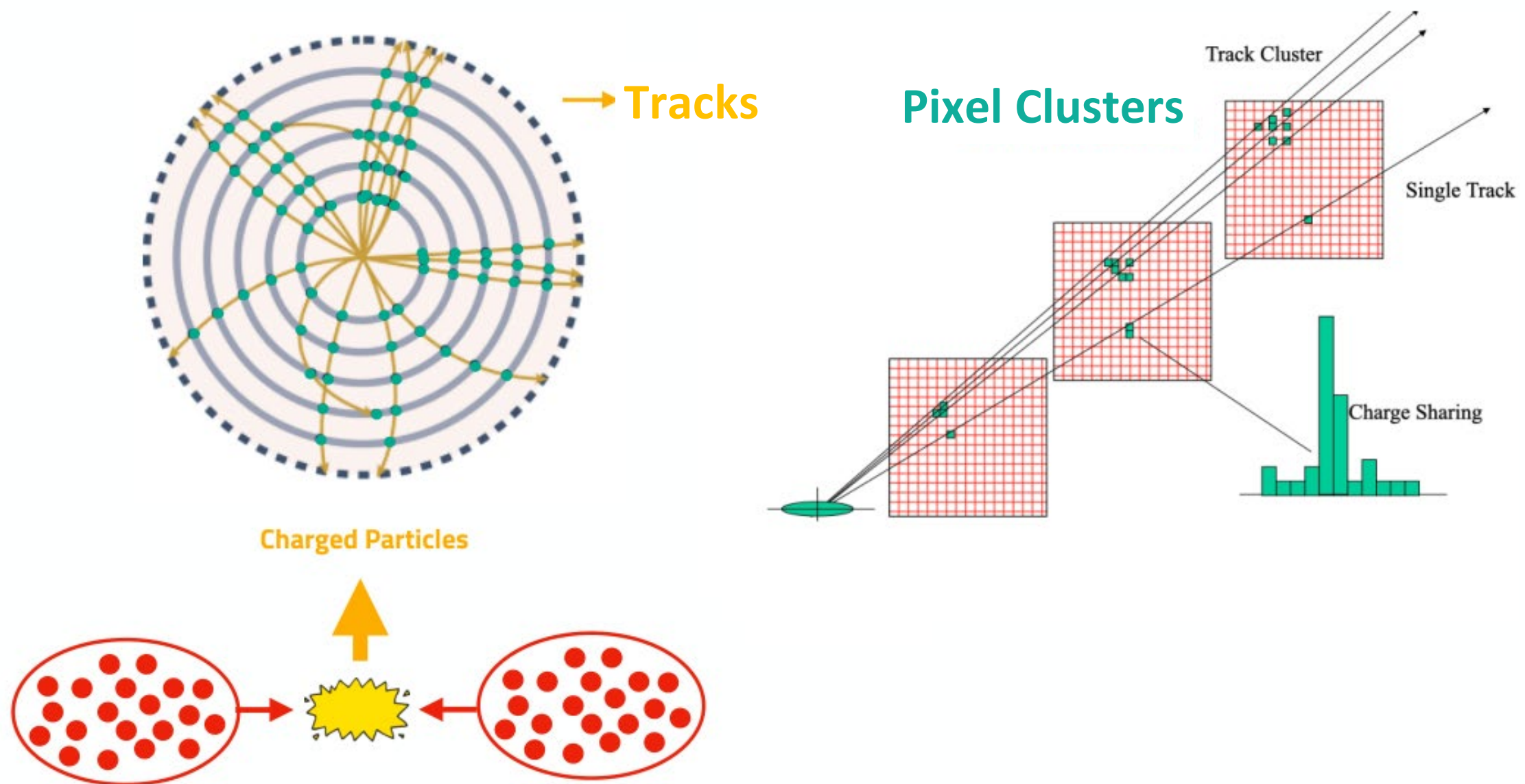
σ_{vis} = calibration constant

⇒ We can measure luminosity by determining μ_{vis} .

Introduction

Luminosity measurements with the ATLAS Inner Detector

- ATLAS monitors luminosity computing μ_{vis} with different detectors and algorithms that are **linear with luminosity and stable over time**.
- For example: **count reconstructed objects produced by charged particles in the Inner Detector**



Introduction

Track Counting and Pixel Cluster Counting

- **Assumption:** number of tracks (clusters) proportional to number of interactions per bunch crossing (μ).
- Measure μ by counting **mean number of tracks (clusters) per event:**

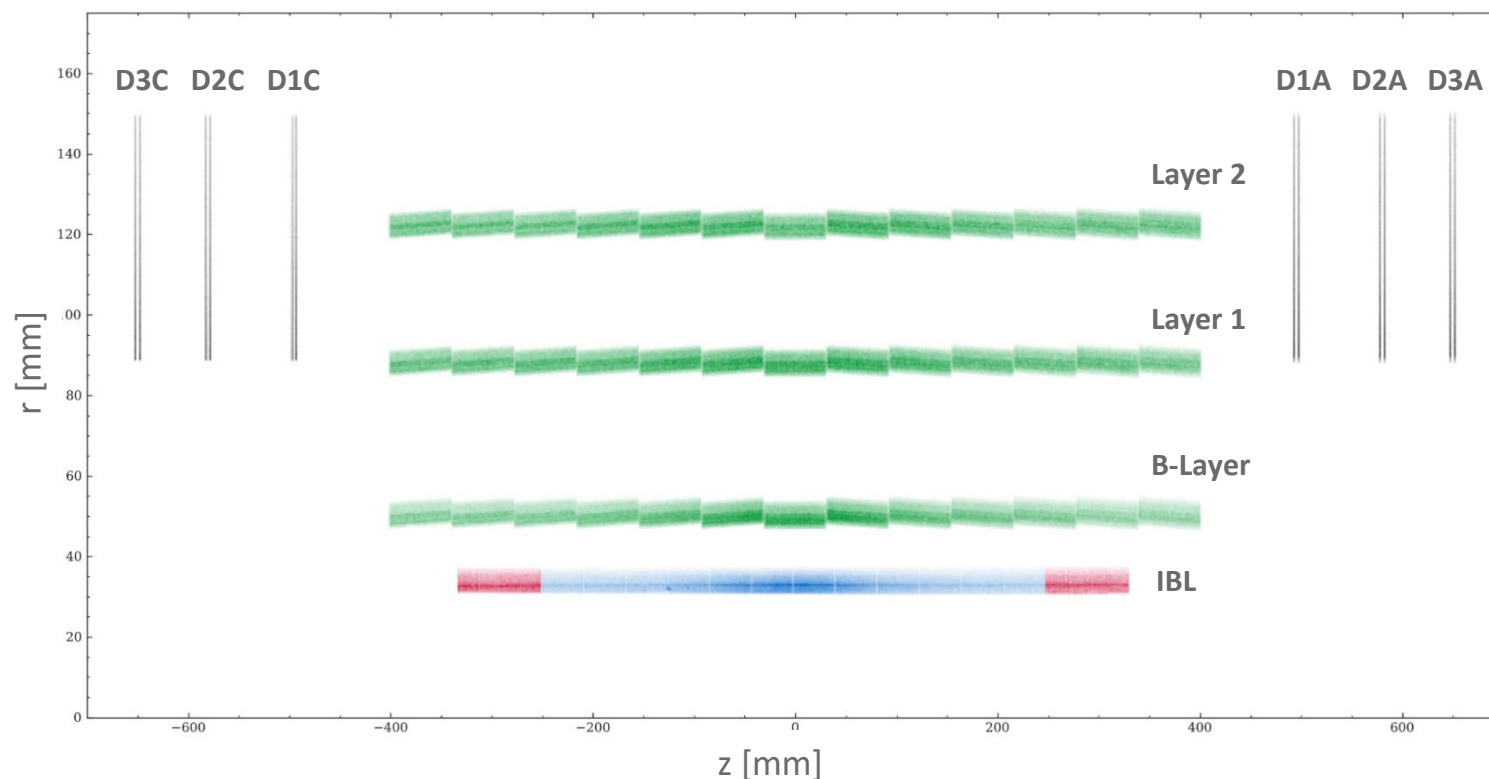
$$\mu_{\text{vis}}^{(\text{TC})\text{PCC}} = \frac{\sum_{i=1}^{N_{\text{events}}} N_{\text{tracks}(\text{clusters})}^i}{N_{\text{events}}}$$

- Track Counting (TC) already in use for ATLAS luminosity measurement.
- **Goal is to add Pixel Cluster Counting (PCC) algorithm in Run 3:**
 - Study cluster properties to identify signal clusters and suppress background;
 - Compare performance of PCC against TC.

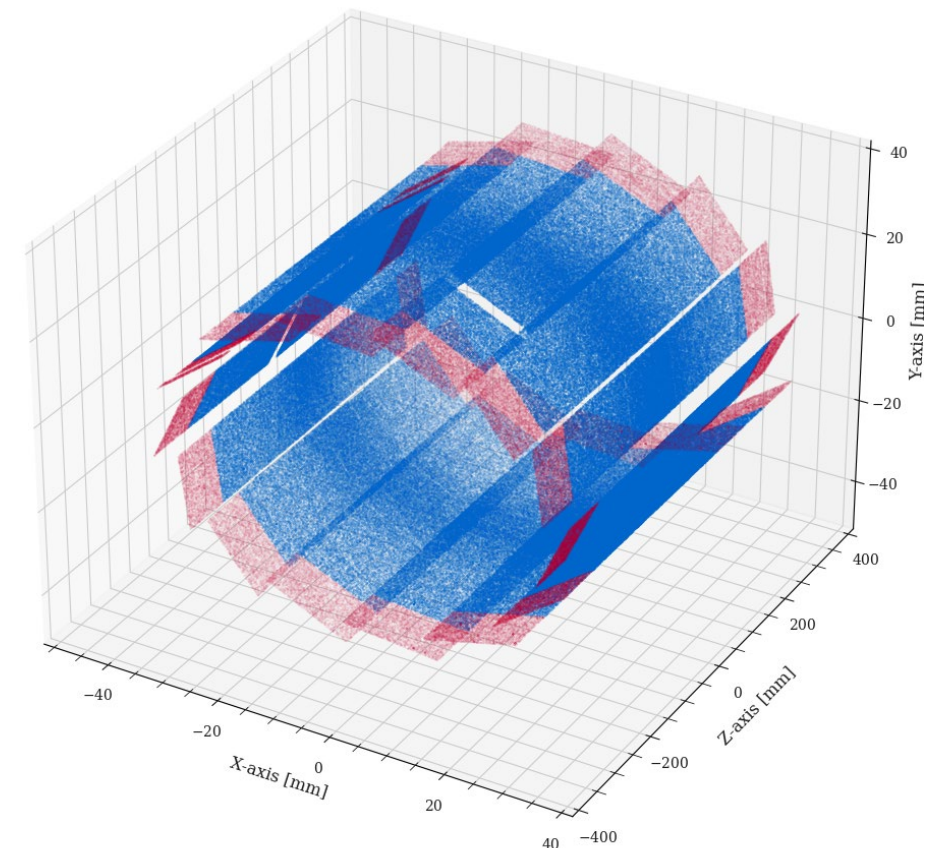
The Pixel Detector

Layout

- The Pixel Detector consists of **4 layers** and **3 disks** of silicon pixel modules.
- **Insertable B-Layer (IBL)** is the closest layer to the interaction point: made of **3D modules** on the sides and **planar modules** in the middle.
- We can visualise the Pixel Detector layout using positions of reconstructed clusters.



Pixel Detector



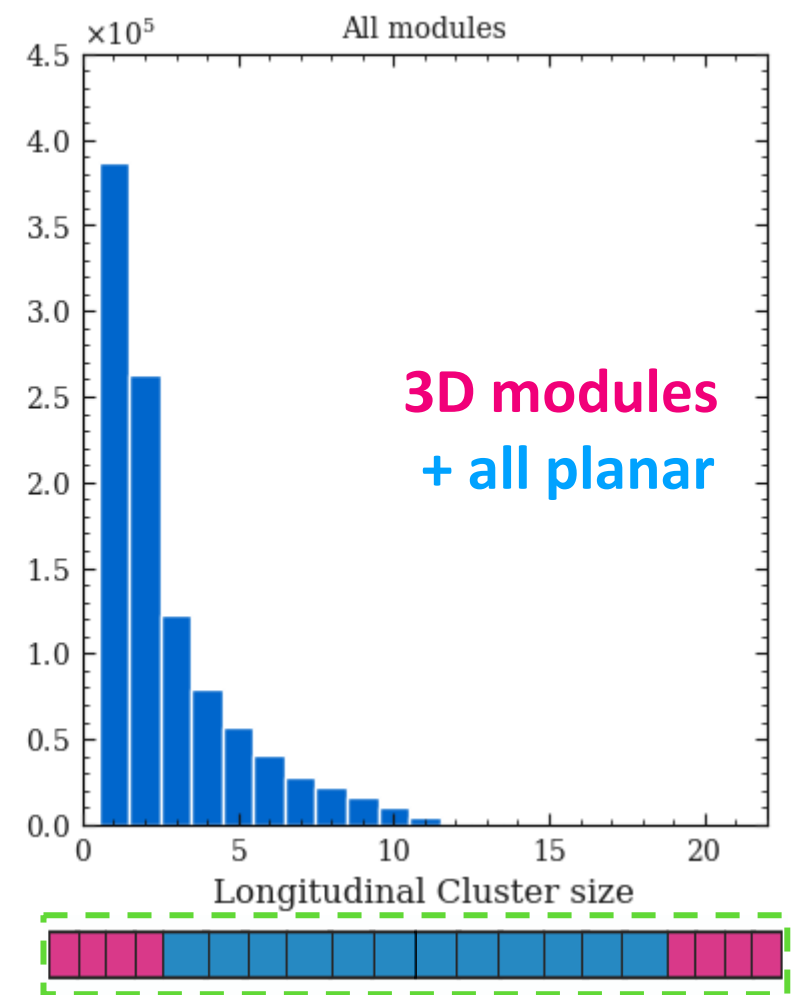
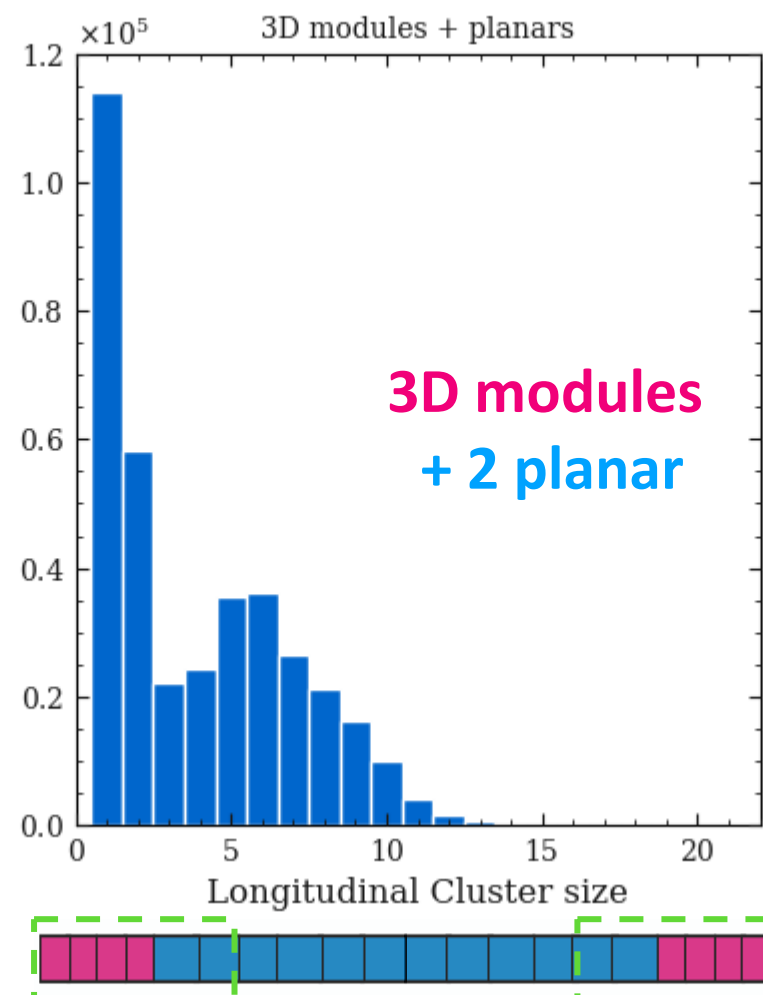
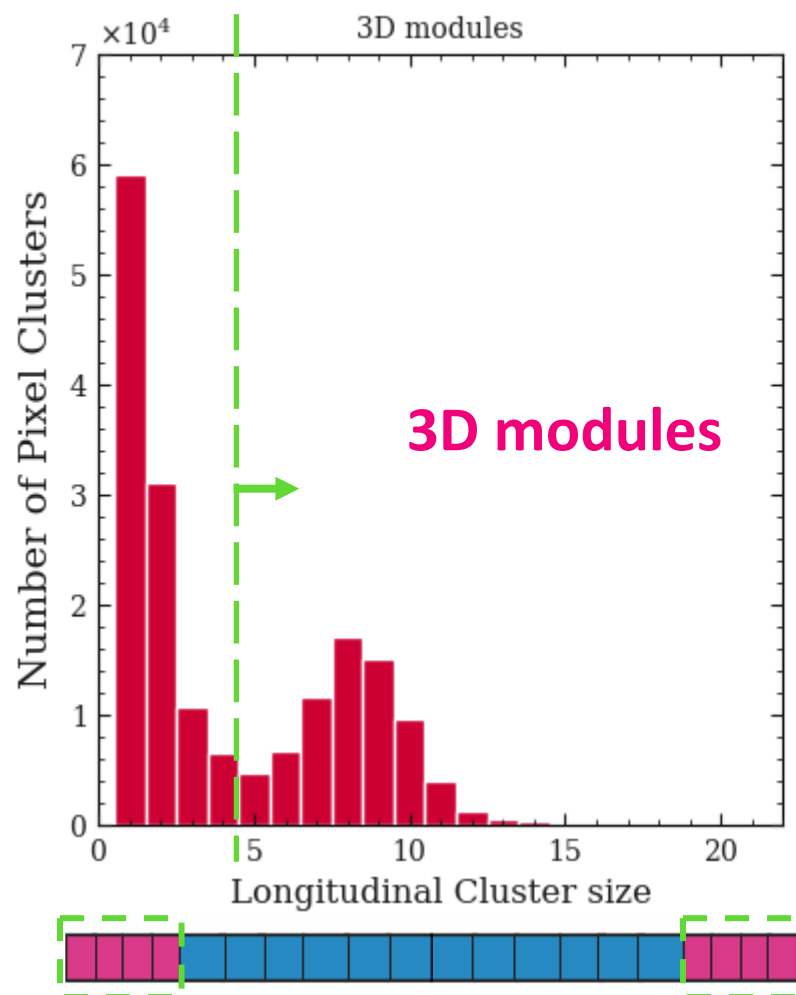
IBL

Pixel Cluster Properties

Longitudinal cluster size

- Primary clusters in **IBL 3D modules** are long in z-direction
⇒ **good signal/background separation.**
- Primary long clusters ($CL_z > 4$) in 3D modules have **Gaussian-like** distribution.
- Background clusters typically short ~ 1 -2 pixels.

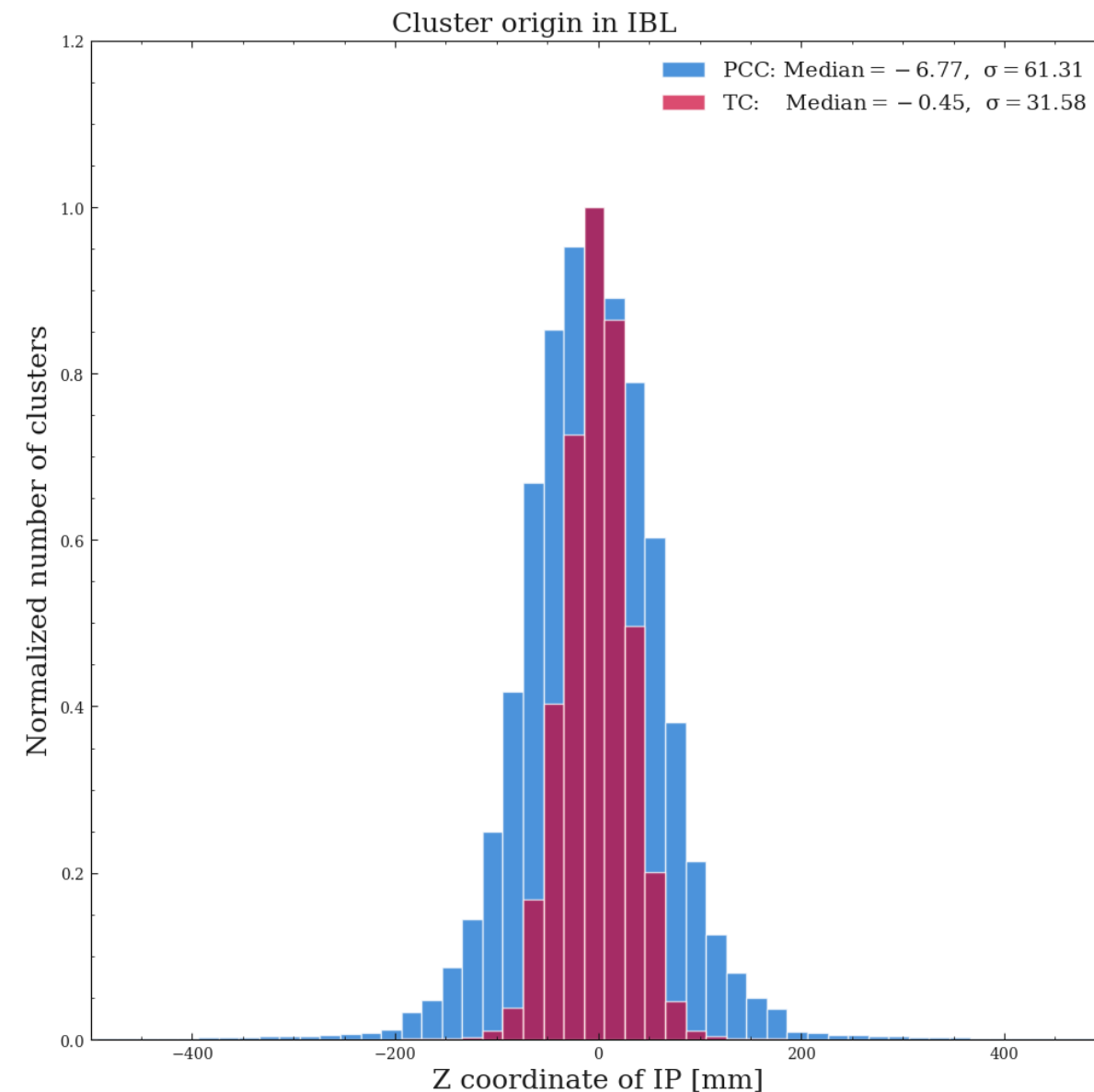
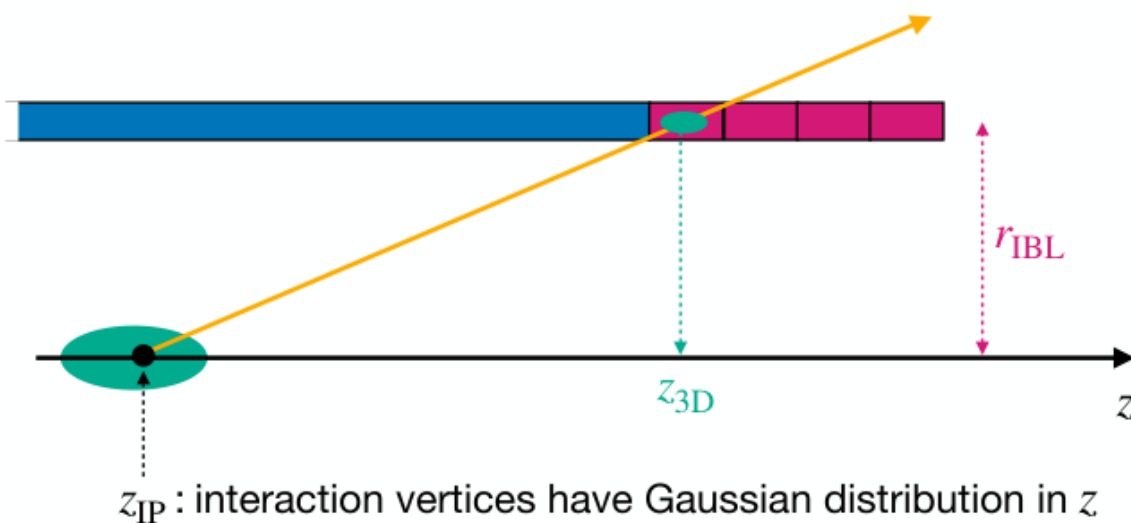
IBL



Pixel Cluster Properties

Longitudinal cluster size $\Rightarrow$ estimation of cluster origin

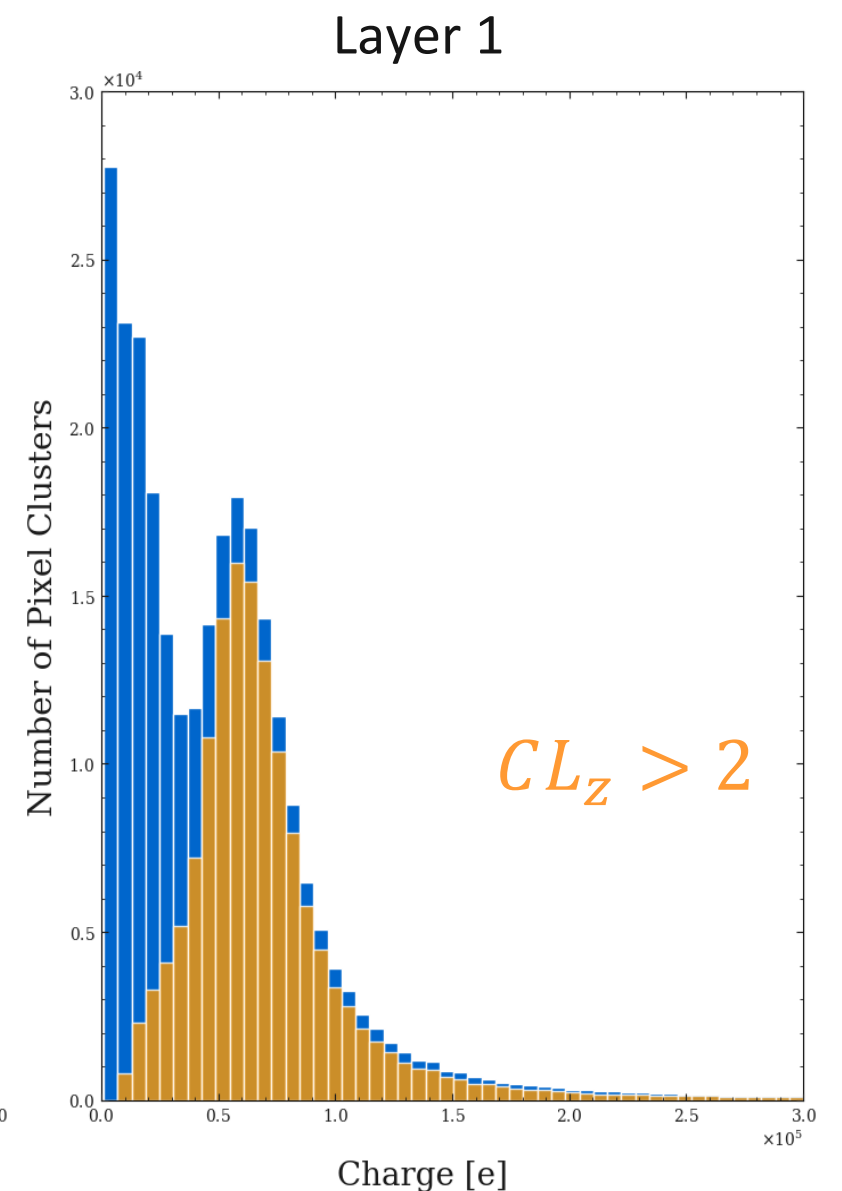
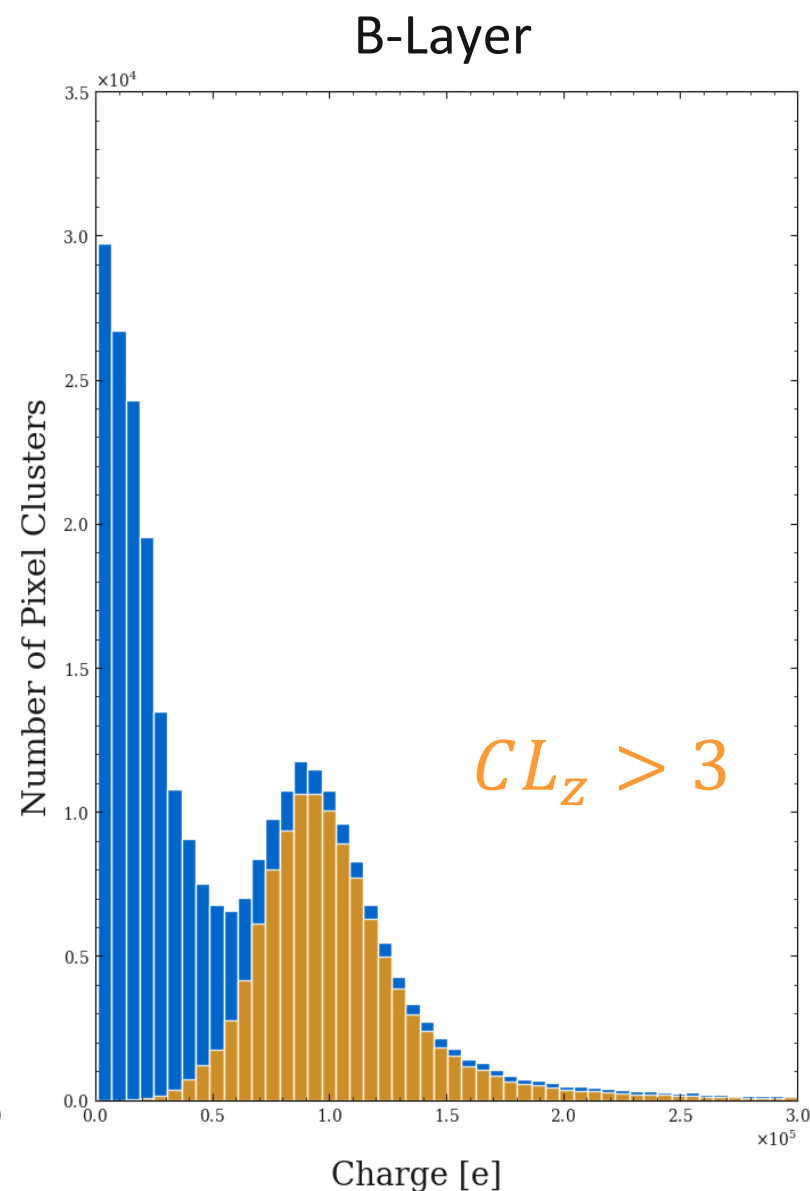
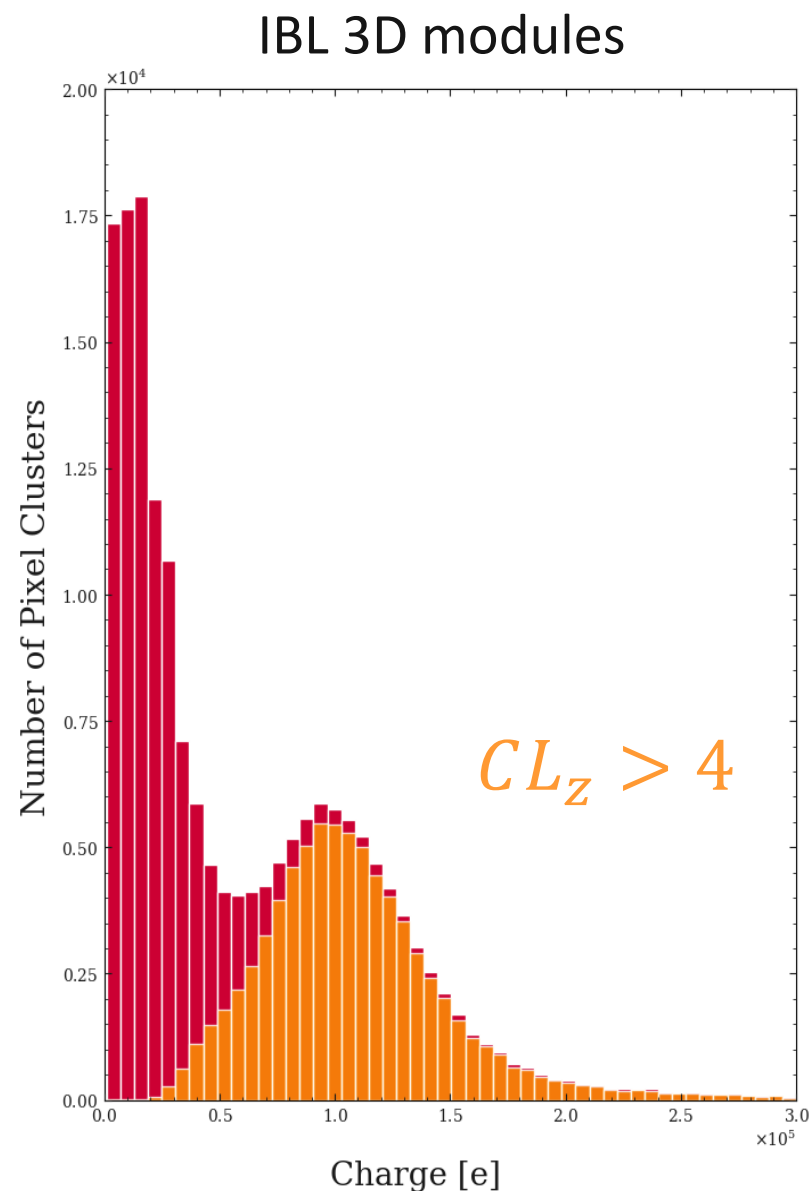
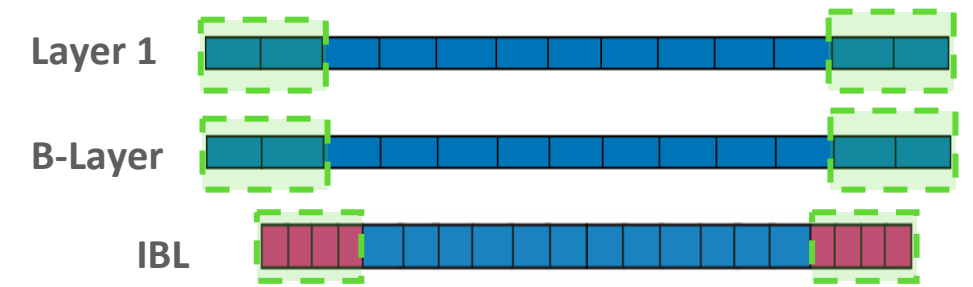
- We can estimate cluster origin (z_{IP}) from $CL_z = \frac{\text{sensor thickness}}{\text{pixel pitch}} \frac{|z_{IP} - z_{3D}|}{r_{IBL}}$
- Selecting **IBL 3D** clusters with $CL_z > 4$ we focus on primary clusters produced by particles originating at the interaction point (IP).
- z_{IP} estimated from clusters CL_z compatible with tracks origin.



Pixel Cluster Properties

Cluster charge

- Charge distribution is **bimodal in all layers** with a Gaussian component.
- Gaussian part due to primary clusters
⇒ cutting on CL_z removes peak at low charge.

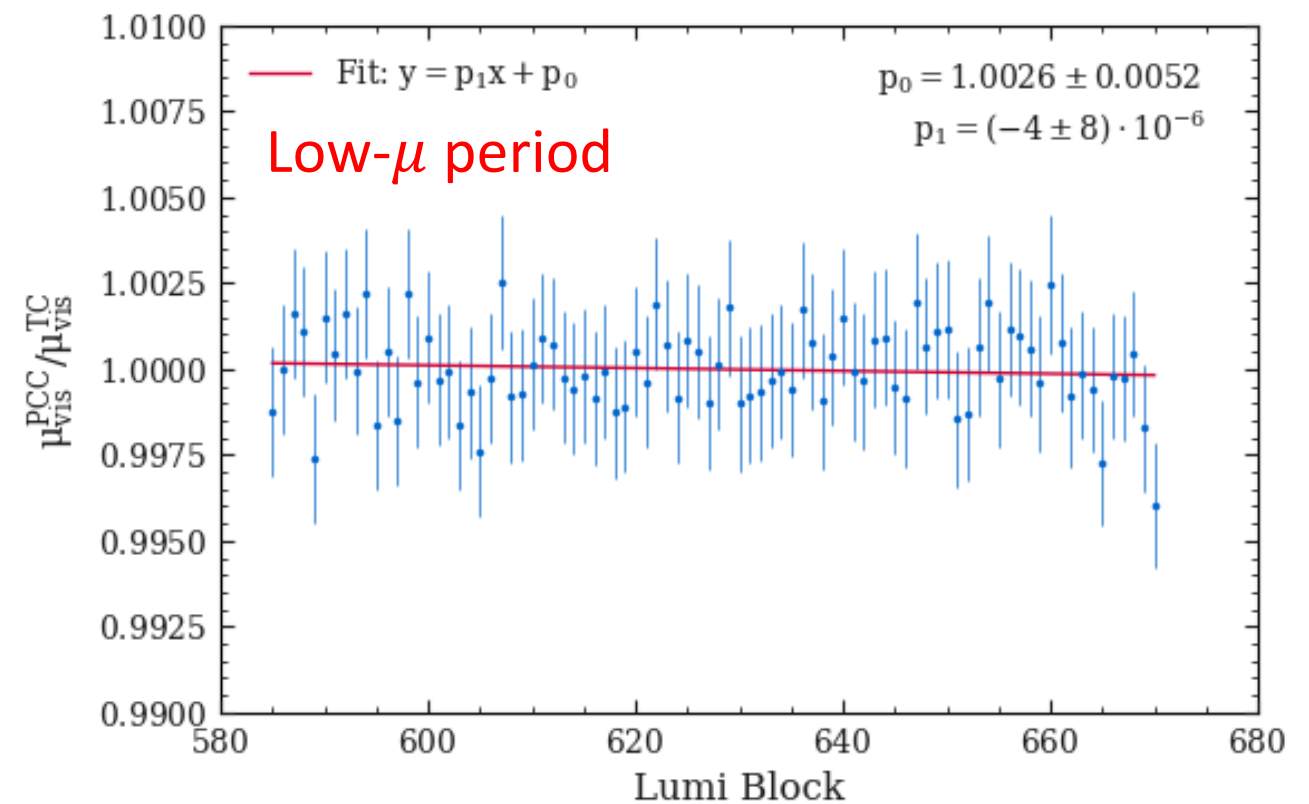
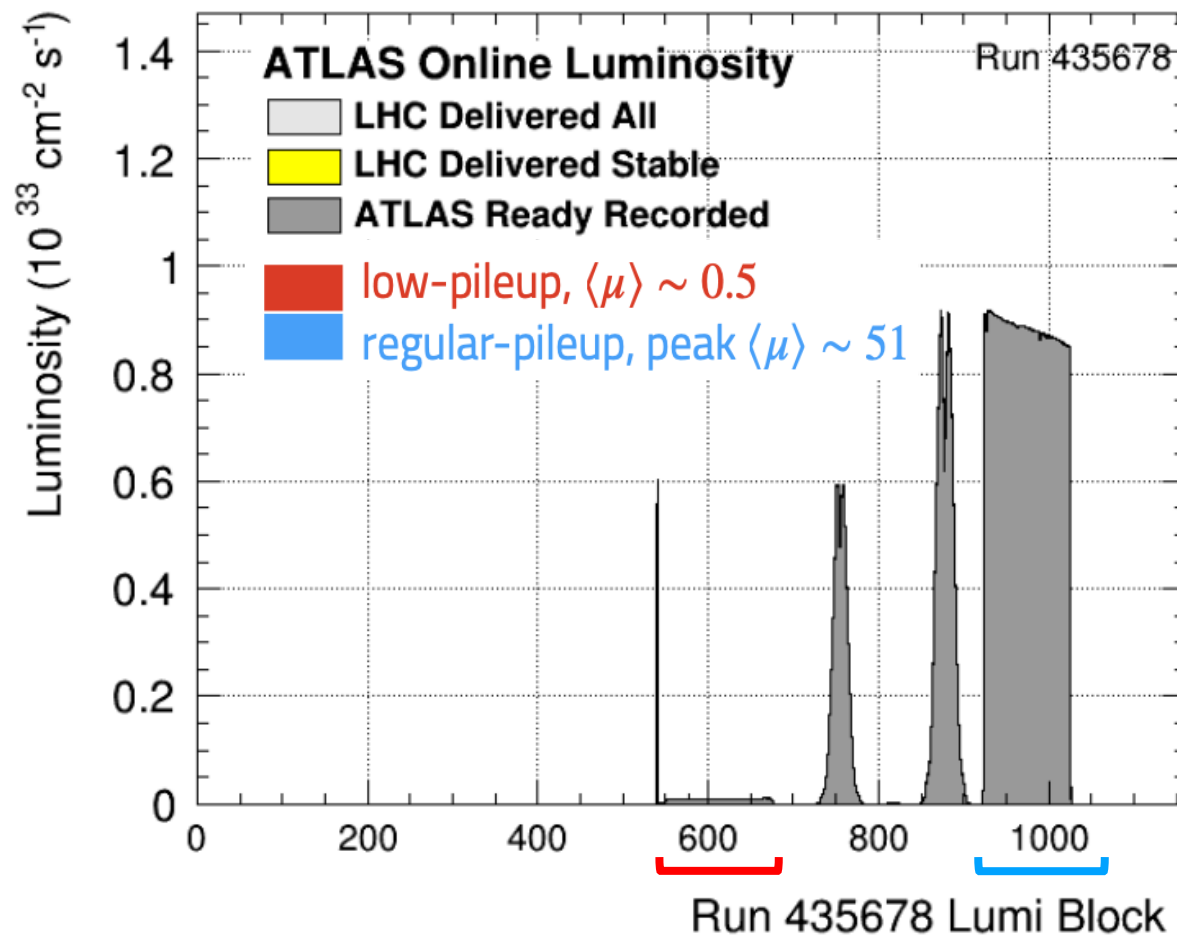


PCC with IBL 3D modules

$$\mu_{\text{vis}}^{\text{PCC(TC)}} = \frac{\sum_{i=1}^{N_{\text{events}}} N_{\text{clusters(tracks)}}^i}{N_{\text{events}}}$$

Linearity checks against TC

- First implement PCC using clusters with $CL_z > 4$ in IBL 3D modules.
- Validate PCC linearity by comparison with TC, which is proven to be linear with μ .
- Use a special run with **isolated bunches** that has both **low- μ** and **regular- μ** periods: normalise $\mu_{\text{vis}}^{\text{PCC}}$ to $\mu_{\text{vis}}^{\text{TC}}$ at **low μ** and check the ratio $\mu_{\text{vis}}^{\text{PCC}} / \mu_{\text{vis}}^{\text{TC}}$ at **regular μ** .



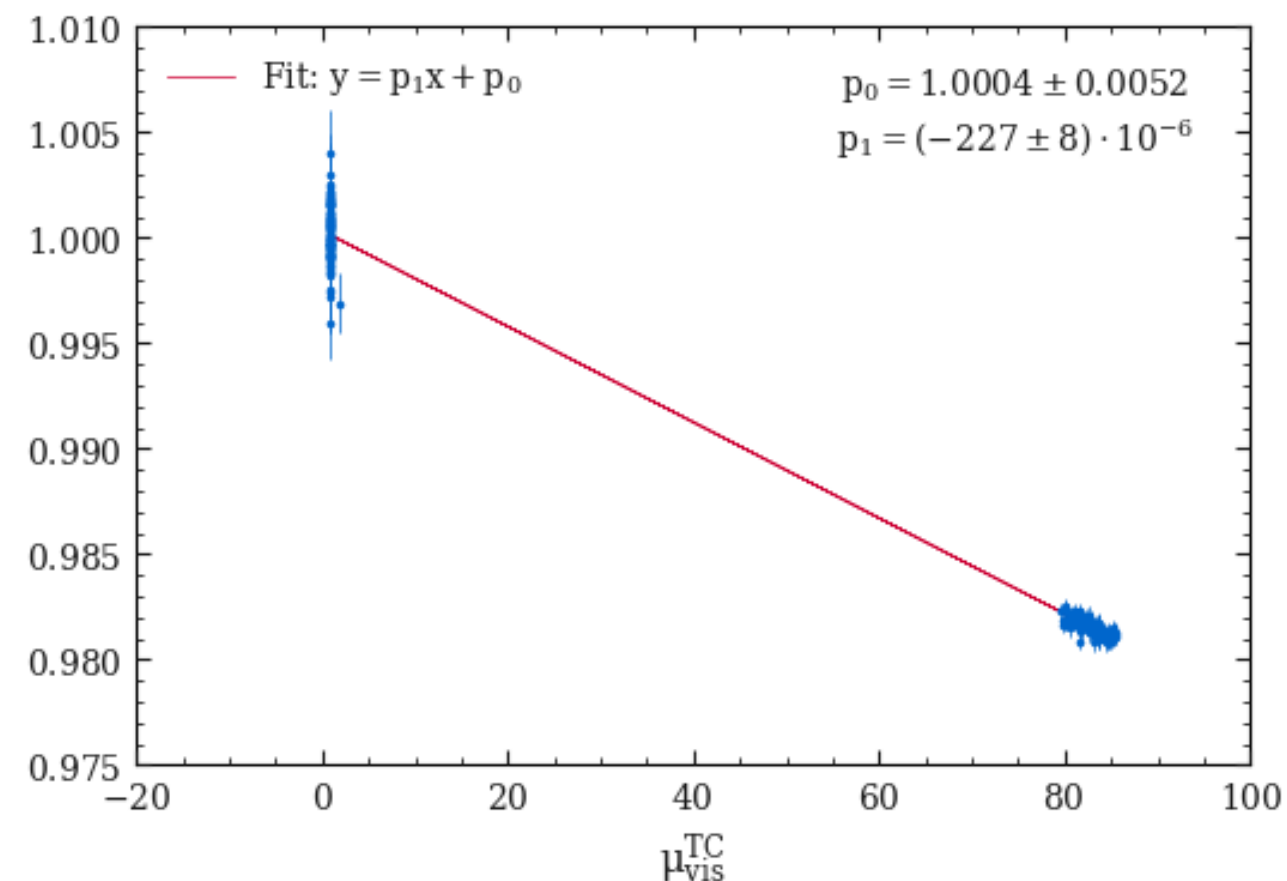
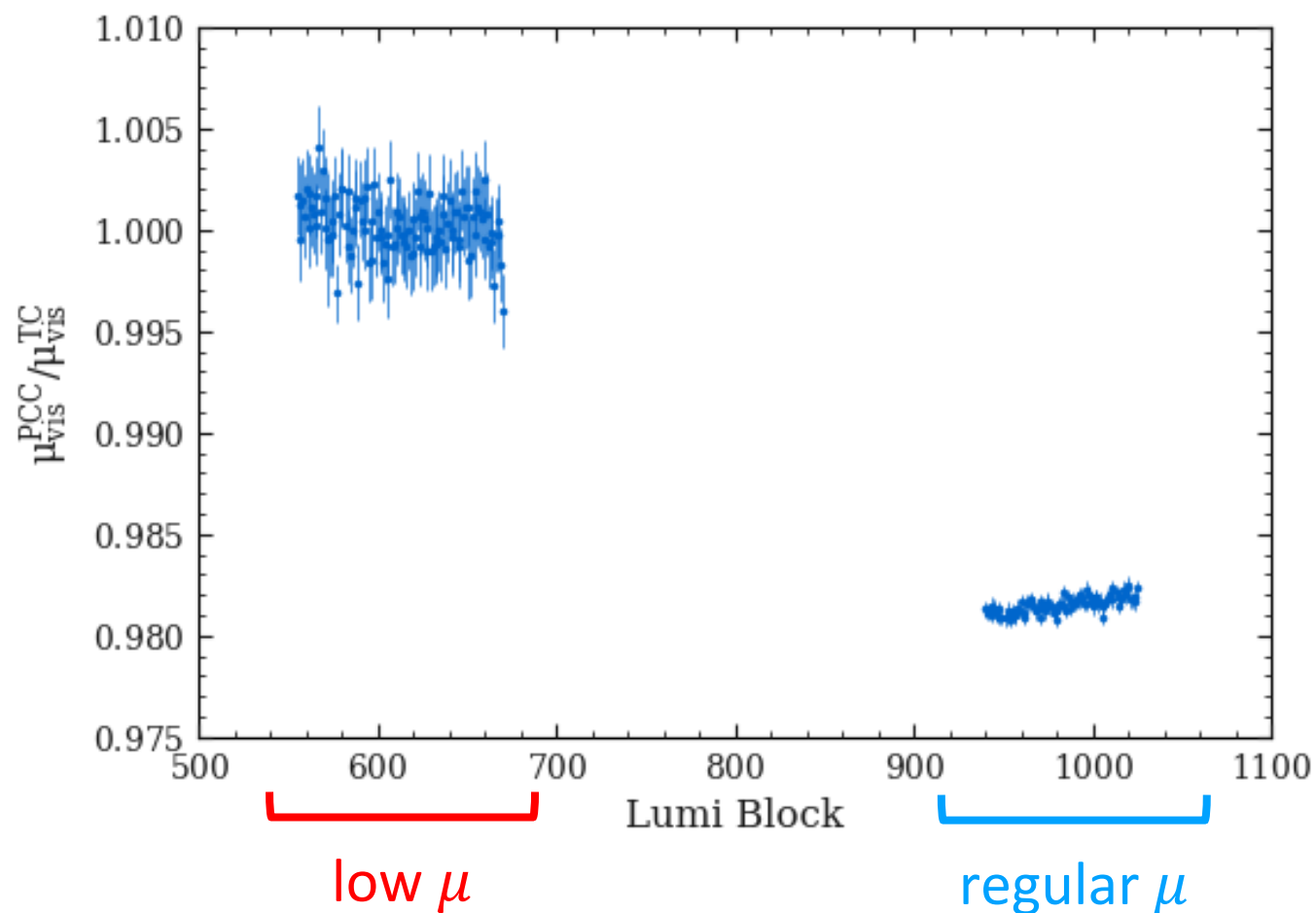
> A **Lumi Block (LB)** is the smallest time unit in a run for luminosity measurements: ~ 60 s in which $\mathcal{L}$ can be considered constant.

PCC with IBL 3D modules

$$\mu_{\text{vis}}^{\text{PCC(TC)}} = \frac{\sum_{i=1}^{N_{\text{events}}} N_{\text{clusters(tracks)}}^i}{N_{\text{events}}}$$

Linearity check at regular pileup

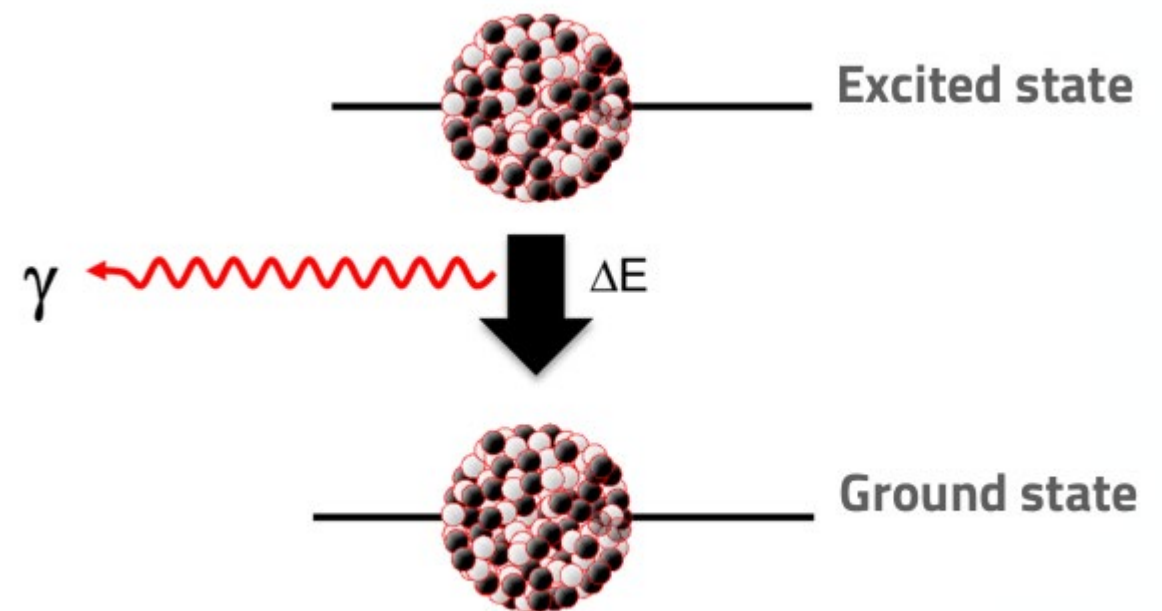
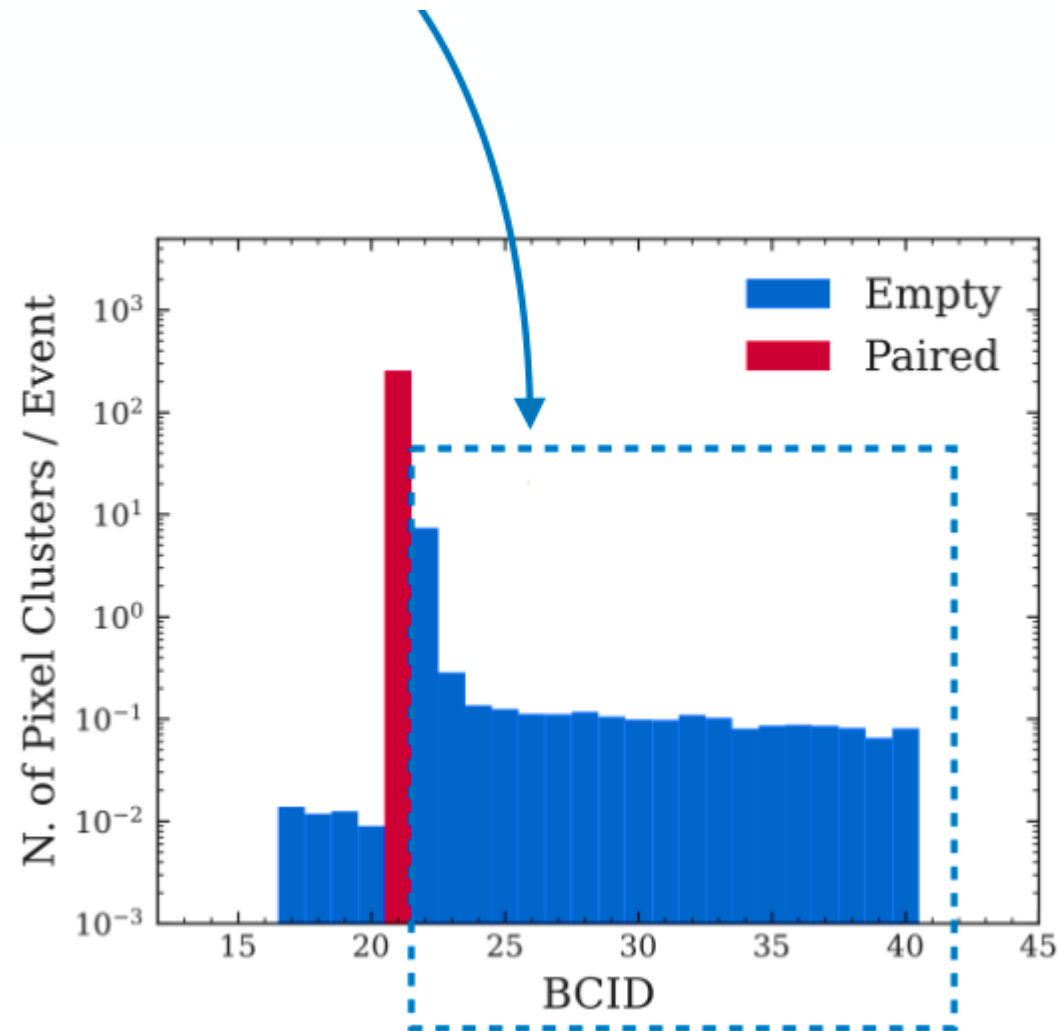
- Variation in $\mu_{\text{vis}}^{\text{PCC}} / \mu_{\text{vis}}^{\text{TC}}$ of $\sim 2\%$ from $\mu_{\text{vis}}^{\text{TC}} = 0.75$ to $\mu_{\text{vis}}^{\text{TC}} = 79$.
- May be due to various sources of **background**, e.g. broken clusters from dead/silent pixels or clusters on module's edge.



Signal vs background separation

Main types of background in normal physics conditions

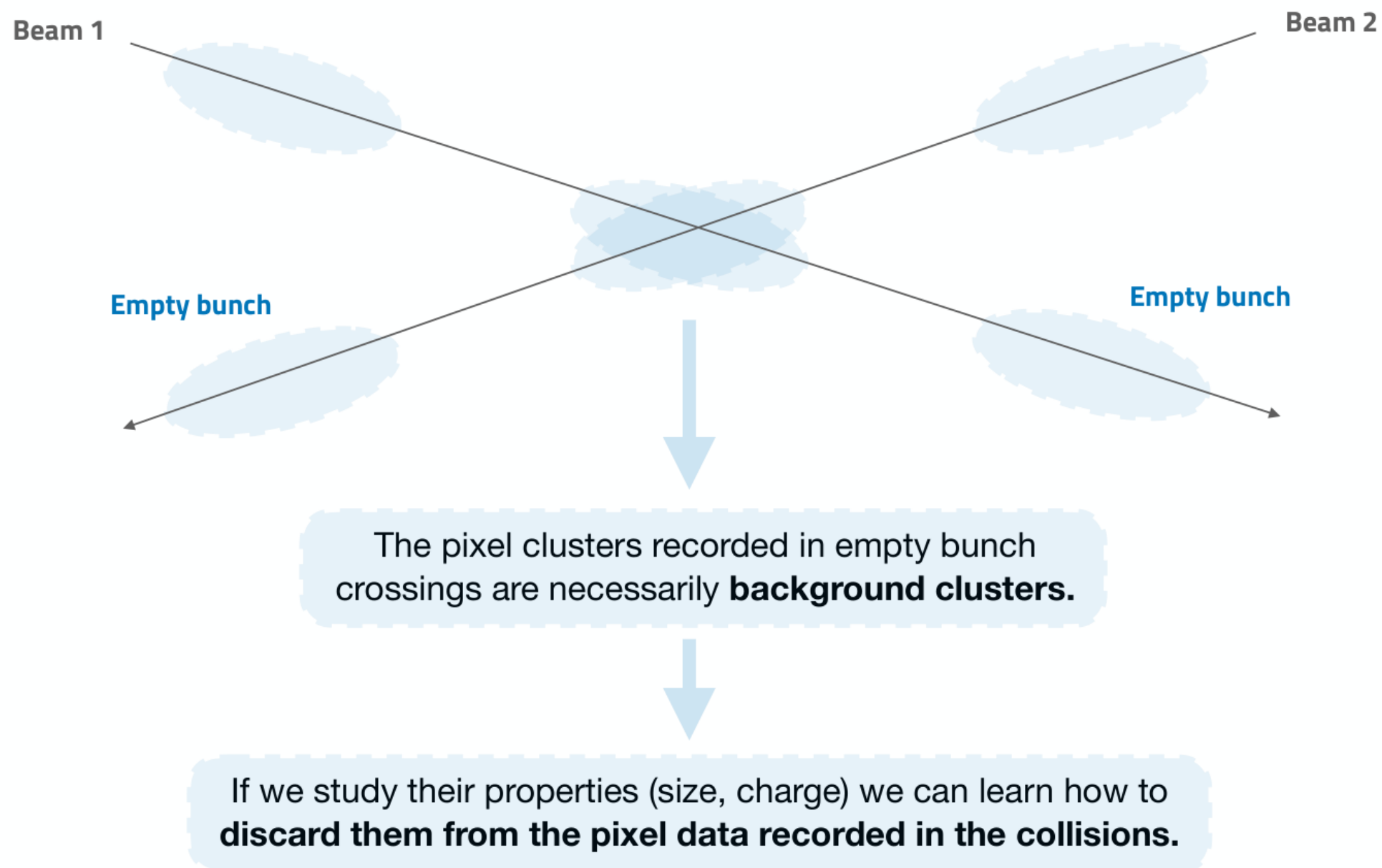
- **Intrinsic background:** noisy/hot pixels firing signal at high rate ~ 1 -pixel clusters.
- **Afterglow** after a collision due to photon emission in nuclear de-excitation of the detector material.



Signal vs background separation

Optimising cluster selections using special runs

- To differentiate between signal and background clusters we can use data recorded with a special random **trigger sampling** also **empty bunch crossings**:



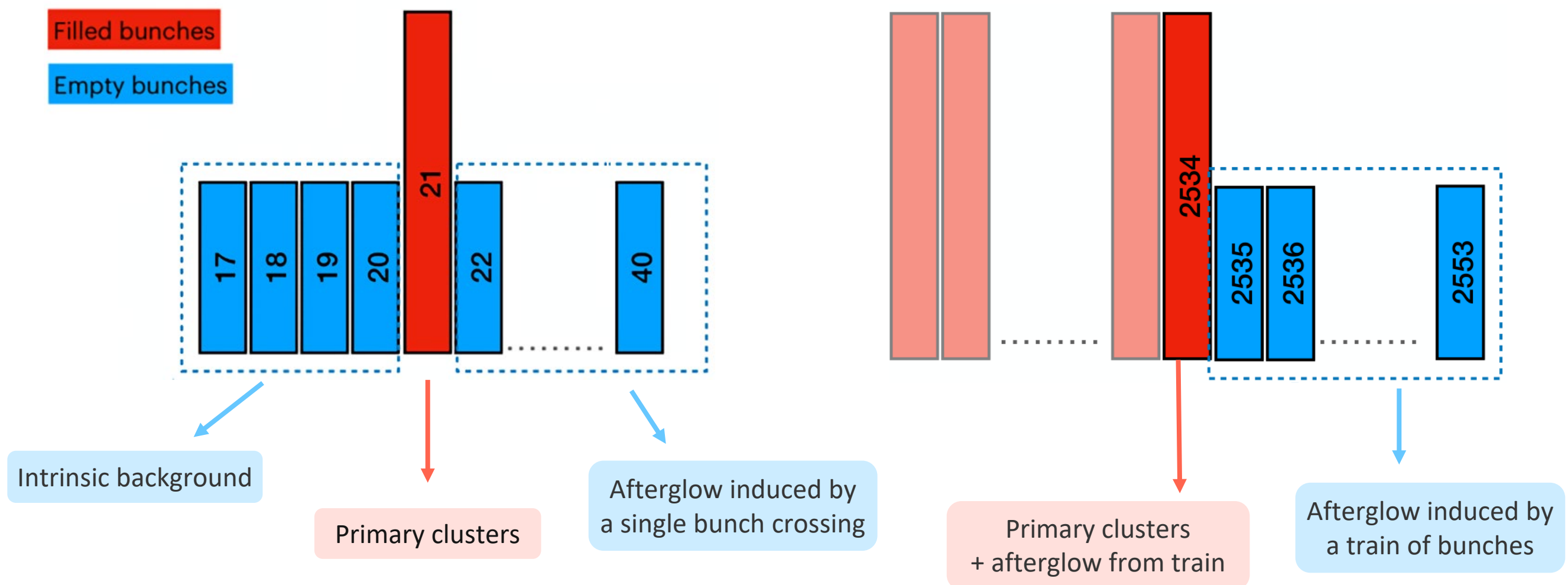
Signal vs background separation

Optimising cluster selections using special runs

- To differentiate between signal and background clusters we can use data recorded with a special random **trigger sampling** also **empty bunch crossings**:

A) Empty bunches around **isolated** filled bunch

B) Empty bunches after the last bunch in a **train** of 12 filled bunches



Signal vs background separation

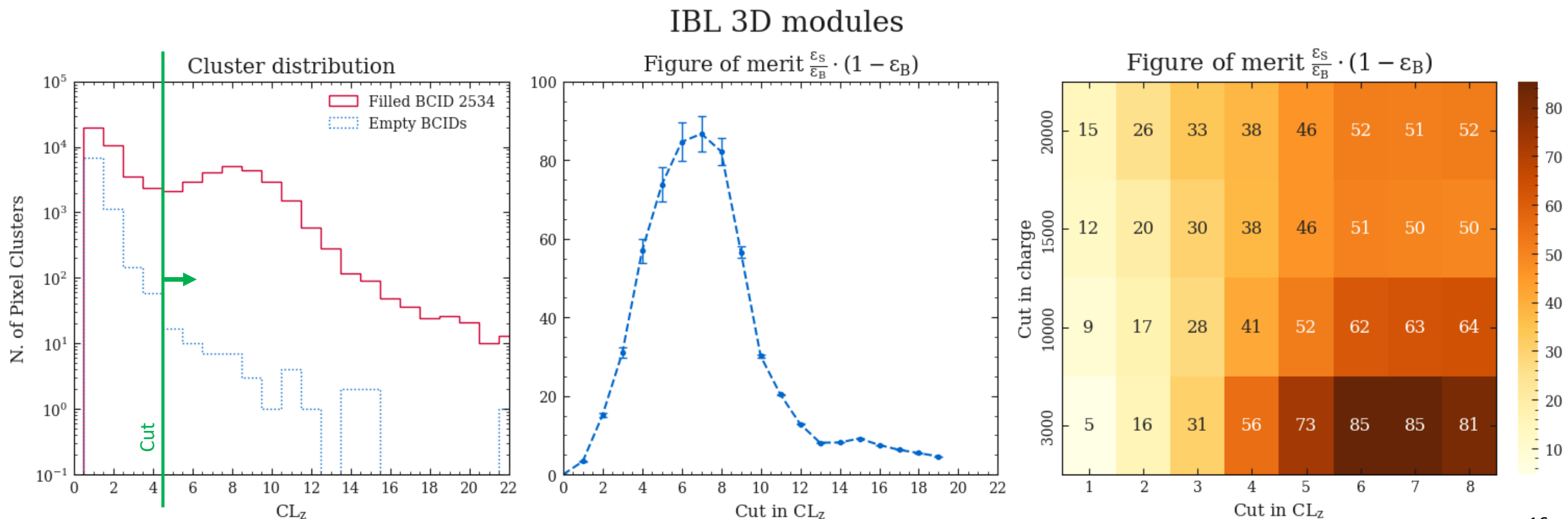


Optimising cluster selections using special runs

- Characterise selections in subregions of the Pixel Detector by monitoring fractions of **signal** (ε_S) and **background** (ε_B) clusters passing **selections**.

$$\varepsilon_S = \frac{N_{sig}(CL_z > cut)}{N_{sig}} \quad \varepsilon_B = \frac{N_{bkg}(CL_z > cut)}{N_{bkg}}$$

- Optimal cuts for each subregion are those maximizing the **figure of merit** $\frac{\varepsilon_S}{\varepsilon_B} (1 - \varepsilon_B)$ in configuration B ($\sim$ normal physics conditions)



Signal vs background separation

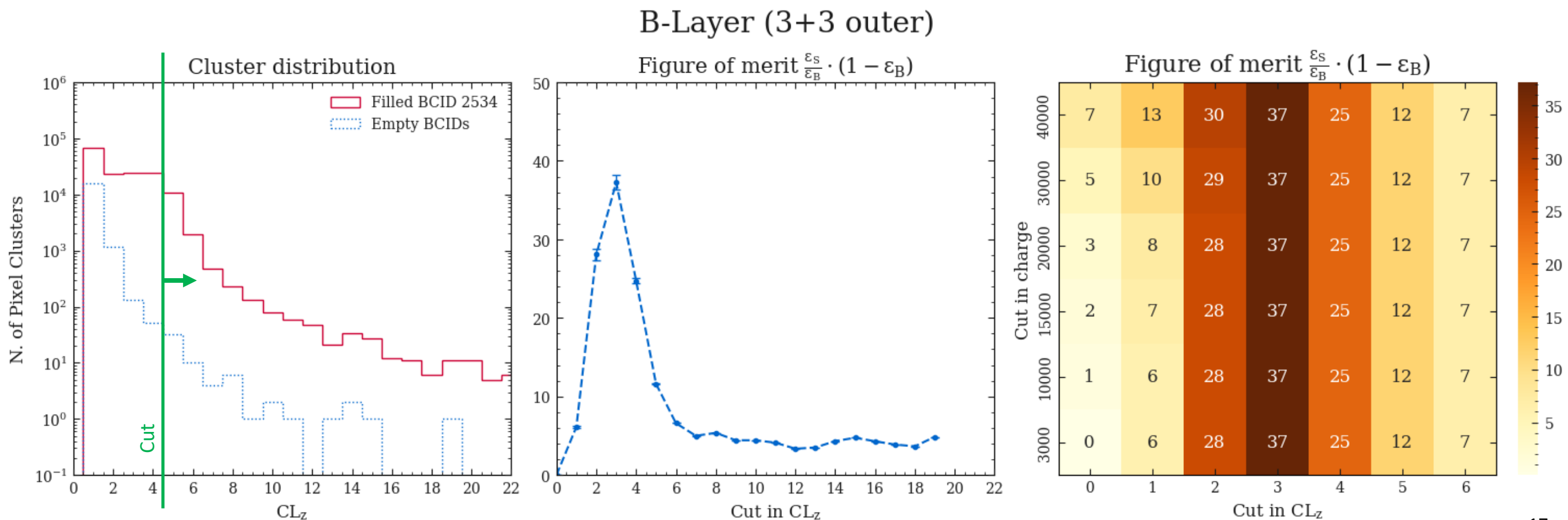


Optimising cluster selections using special runs

- Characterise selections in subregions of the Pixel Detector by monitoring fractions of **signal** (ϵ_S) and **background** (ϵ_B) clusters passing **selections**.

$$\epsilon_S = \frac{N_{sig}(CL_z > cut)}{N_{sig}} \quad \epsilon_B = \frac{N_{bkg}(CL_z > cut)}{N_{bkg}}$$

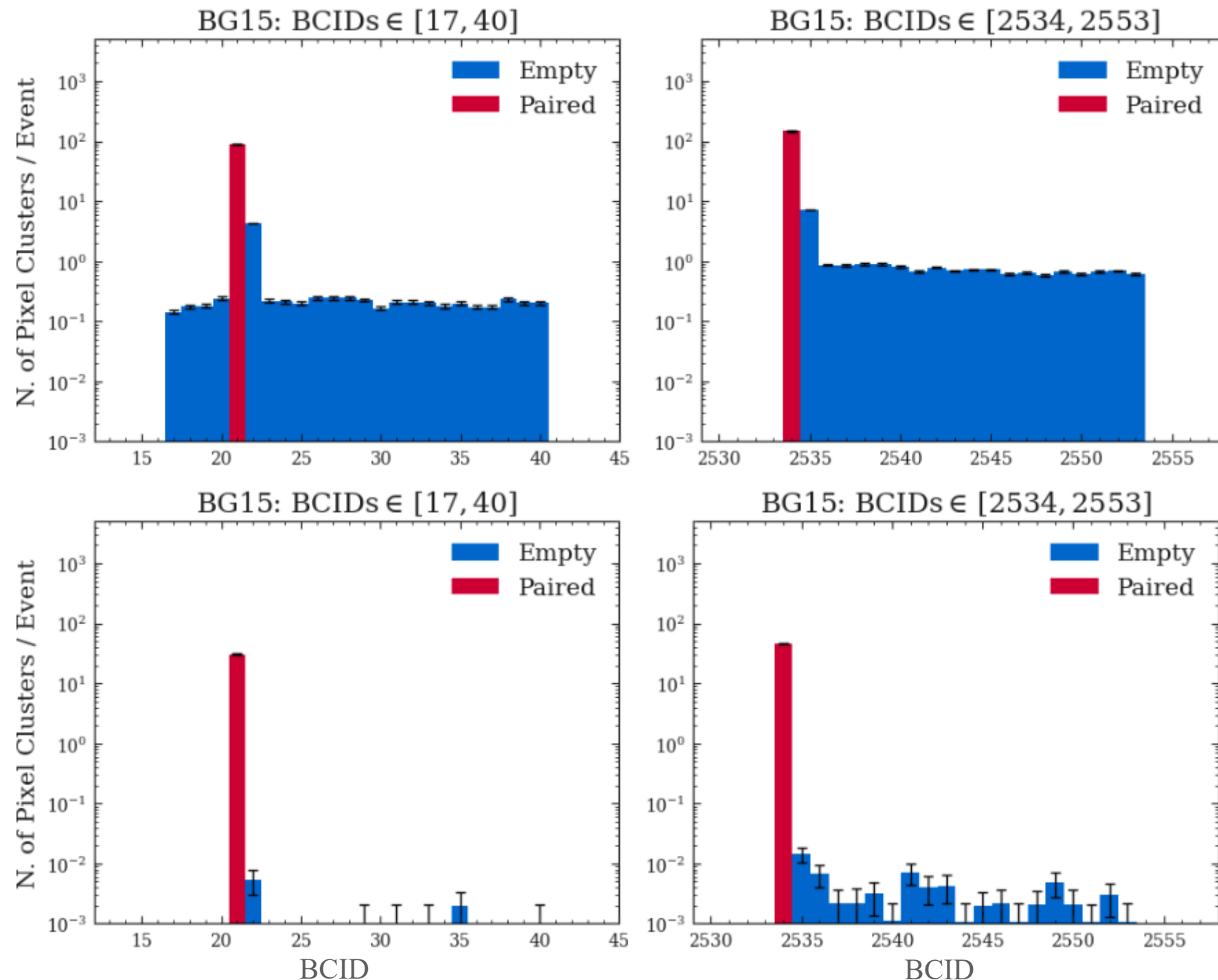
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Signal vs background separation

Applying the optimised selections

IBL 3D modules



Selections:

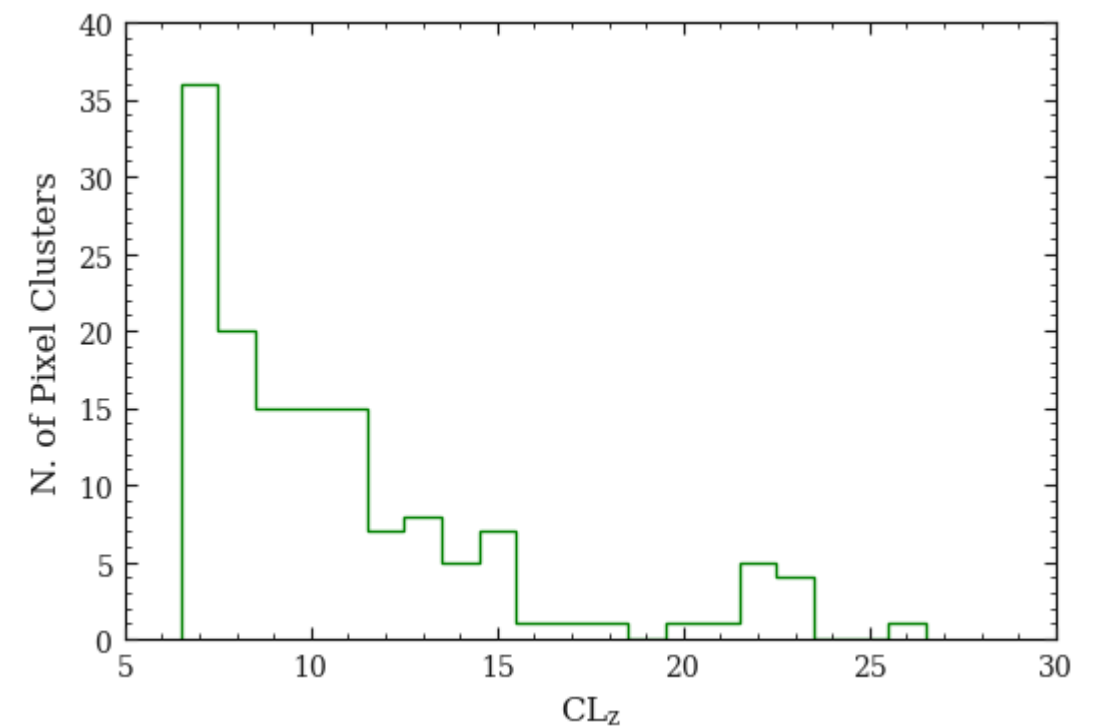
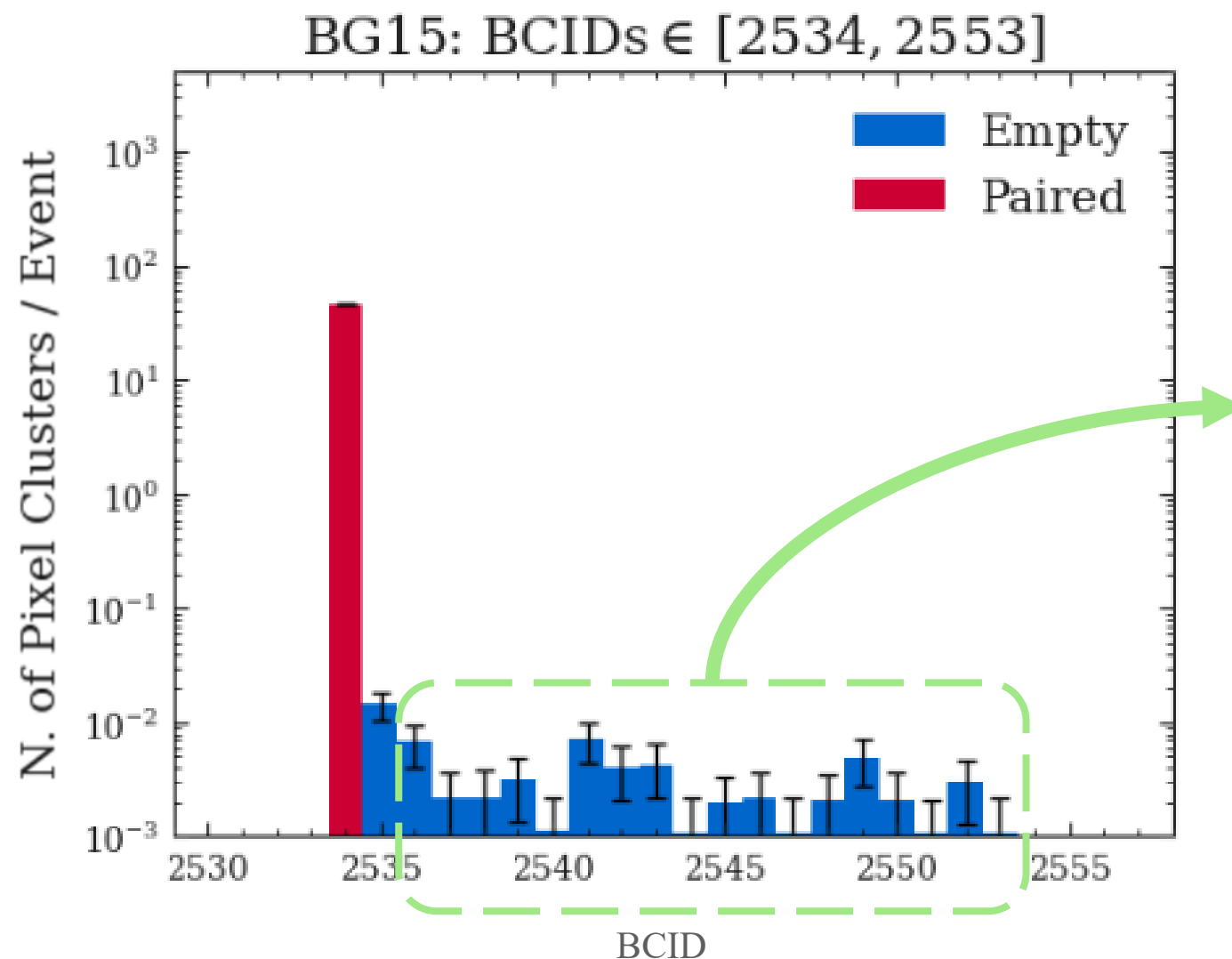
- Charge $> 3000 e$
- $CL_z > 6$

We retain 32% of signal
and 0.36% of background

Signal vs background separation

Residual activity in the empty-BCID tail

- Residual clusters with **large CL_z and CL_ϕ** are still present after cuts in empty bunch crossings, especially in train configuration.
- This could be explained by **other types of background than afterglow**: interactions of protons with residual gas molecules in the beam pipe?



Signal vs background separation

Summary of selections

- Based on the studies a set of selections can be implemented in the online reconstruction for different subregions of the Pixel Detector:

Layer	Subregion	Selection
IBL	3D	$CL_z > 6$, Charge > 3000
	Planar	$CL_z > 3$, Charge > 3000
B-Layer	3A/C side	$CL_z > 3$, Charge > 3000
	7 central	$CL_z > 1$, Charge > 3000
Layer 1	3A/C side	$CL_z > 2$, Charge > 3000
	7 central	$CL_z > 1$, Charge $> 30\,000$
Layer 2	3A/C side	$CL_z > 2$, Charge > 3000
	7 central	$CL_z > 1$, Charge $> 30\,000$
Disks	All	Charge $> 15\,000$

- Selections suppressing most of background and keeping large fraction of signal.

Conclusions

- In this summer project we studied the **Pixel Cluster Counting** method for luminosity measurements at the ATLAS experiment.
- We investigated pixel cluster properties to find **optimal selections** for background rejection in different subregions of the Pixel Detector and validated the **performance of the PCC algorithm** using the IBL 3D modules.
- We identified a set of optimal cuts which allow to **significantly reduce the background** contribution and can be adopted for future ATLAS luminosity measurements with PCC in Run 3.



Thank you!