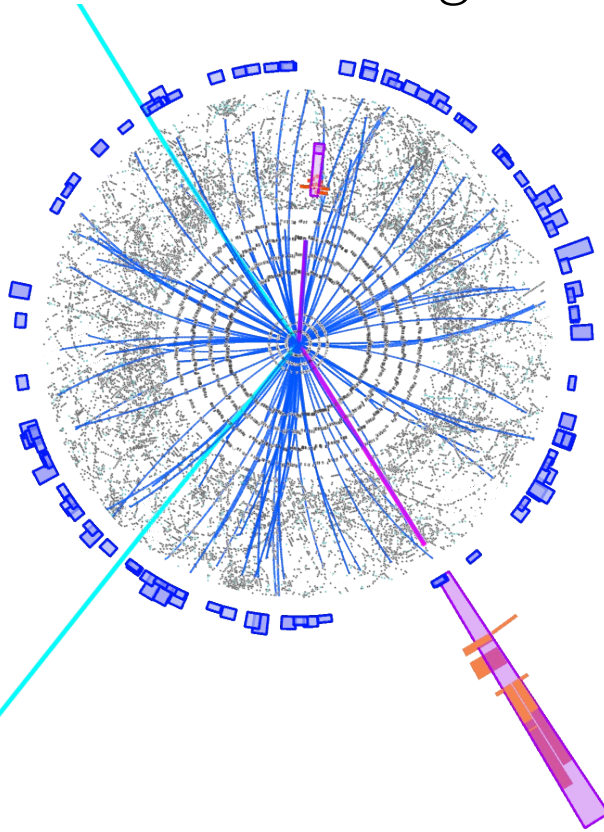


# Detector Physics I

## Particle Detection in High Energy Physics



**Simon Spannagel, DESY**  
Introduction to the Terascale  
17 March 2025

# Table of Contents

- What do we measure?
  - Stable Particles
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  - Semiconductor Detectors
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  - Calorimeters
  - Cherenkov Detectors
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# Prologue: Units

- Elementary charge:  $e = 1.602\,176\,6208 \times 10^{-19} \text{ C}$
- Energy unit eV:  $1 \text{ eV} = 1.602\,176\,6208 \times 10^{-19} \text{ J (CV)}$   
Energy of an electron, accelerated through 1V
- Electron mass:  $m_e = 9.1 \times 10^{-31} \text{ kg}$   
 $= 511 \text{ keV} / c^2$
- Factor  $c^2$  often dropped, units defined with  $c = 1$  (and  $\hbar = 1$ ):  
MeV, GeV, TeV...  
for energy, momentum, (rest-) mass
- Scientific notation: Mantissa plus exponent to Base 10:  $a \times 10^b$

# What do we measure?

Determination of particle properties

# Which particles can we measure?

- Particles have to be long-lived enough to reach the detectors...
  - Many elementary particles have very short life time (Higgs, W, Z...)
  - Measuring their decay products
- Particles have to interact with detector material!
  - Not every particle participates in every force
  - Possibilities of detection depends strongly on forces
- Elementary particles:  $e^{\pm}, \mu^{\pm}, \nu^e, \bar{\nu}^e, \nu^{\mu}, \bar{\nu}^{\mu}, \nu^{\tau}, \bar{\nu}^{\tau}, \gamma$
- Baryons:  $p^{\pm}, n, \Sigma^{\pm}, \Xi_0^{\pm}, \Xi^{\pm}, \Omega^{\pm}$
- Mesons:  $\pi^{\pm}, K^{\pm}, K_0 (K_0^S, K_0^L)$

$$\tau_n \approx 15 \text{ min}$$

$$\tau_{\mu} \approx 2 \cdot 10^{-6} \text{ s}$$

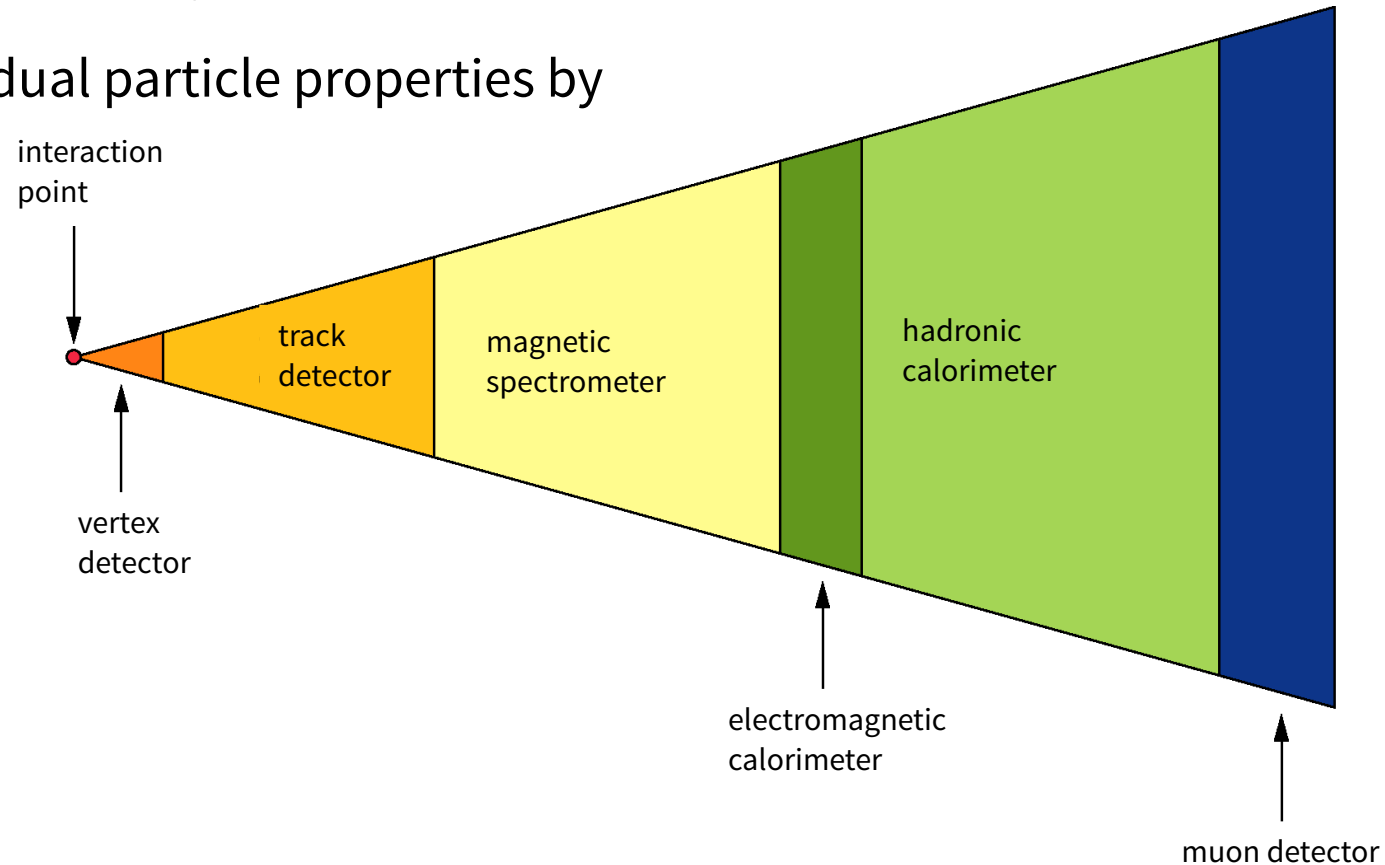
# Observables

- Momentum **p**      Bending radius of track in magnetic field
- Velocity **v**      Time of flight, RICH, etc.
- Charge **Q**      Bending radius in magnetic field
- Lifetime **τ**      Measurement of decay length
- Energy **E**      Absorption in calorimeters
- Rest mass **m**      Indirect measurement e.g. from momentum and energy or velocity

$$E^2 = m^2 c^4 + p^2 c^2 \quad p = \gamma m v = \frac{m v}{\sqrt{1 - v^2/c^2}}$$

# Typical Design of an Experiment

- Combination of complementary detection methods
- Measurement of individual particle properties by separate detectors
- Order is important!  
Some measurements are “destructive”
- Design of many experiments is very similar



# Detector Terminology

- **Dead time:**

*Time period immediately after the detection of a particle during which the detector is not yet ready again to detect another particle.*

- *Non-paralyzable detector: newly occurring event does nothing*
- *Paralyzable detector: newly occurring event extends dead time*

- **Resolution:**

*Achievable uncertainty on the observable*

- **Efficiency:**

*Number of recorded/detected events divided by number of events that occurred*

# Interaction of Radiation with Matter

Energy loss and interaction processes

# Interaction with Matter

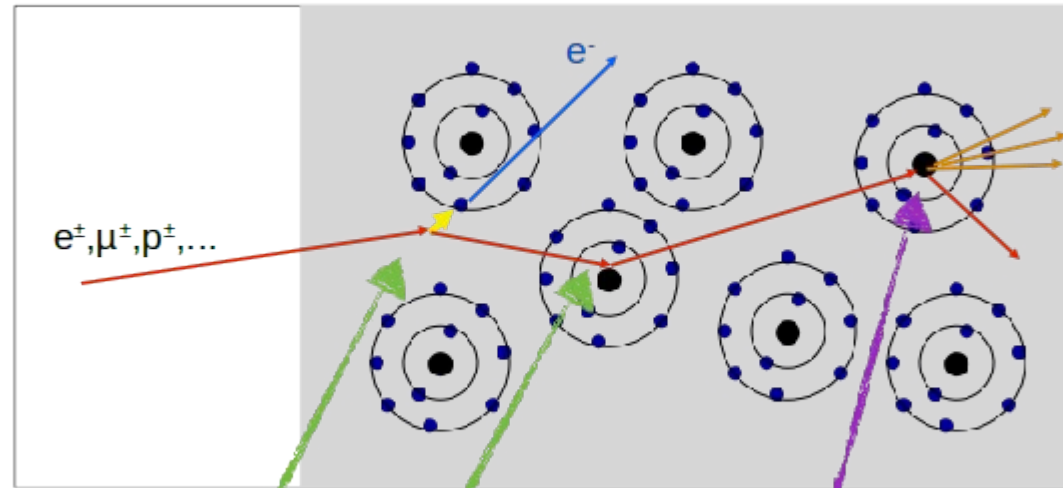
- High energy particles interact via different processes with matter, depending on
  - Particle type
  - Energy
  - Material
- Energy loss of the particles via interaction
  - Energy transfer to material or other (free) particles
  - In Detectors: *Energy loss = signal!*

# Charged Particles

Ionization

**Elastic scattering**  
Recoil from atom  
lattice  $\rightarrow$  photons

**Inelastic scattering**



**Electromagnetic interaction**

Electromagnetic,  
weak or  
strong Interaction

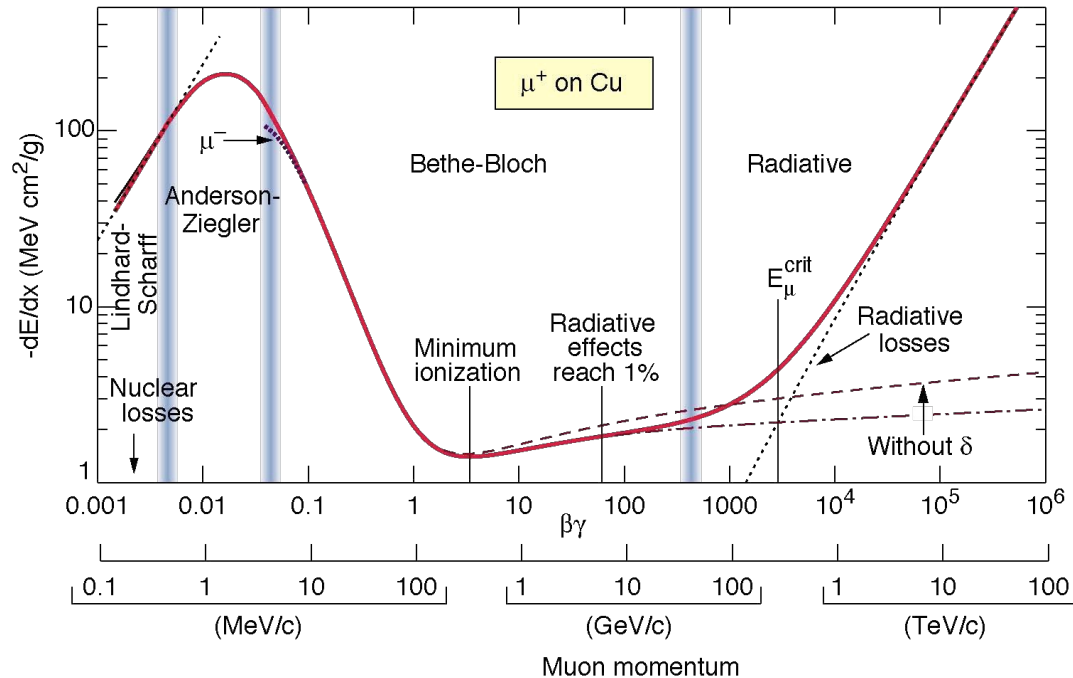
# Mean Energy Loss

- Charged particles interact with electrons in matter

- For heavy charged particles  
**Bethe-Formula**

$$-\left\langle \frac{dE}{dx} \right\rangle \approx K q^2 \frac{1}{\beta^2} \frac{Z}{A} \left[ \ln \left( \frac{2 m_e c^2}{I} \beta^2 \gamma^2 \right) - \beta^2 \right]$$

- Energy loss depends on
  - Properties of projectile: charge, energy
  - Target properties: atomic number, ionization energy, density

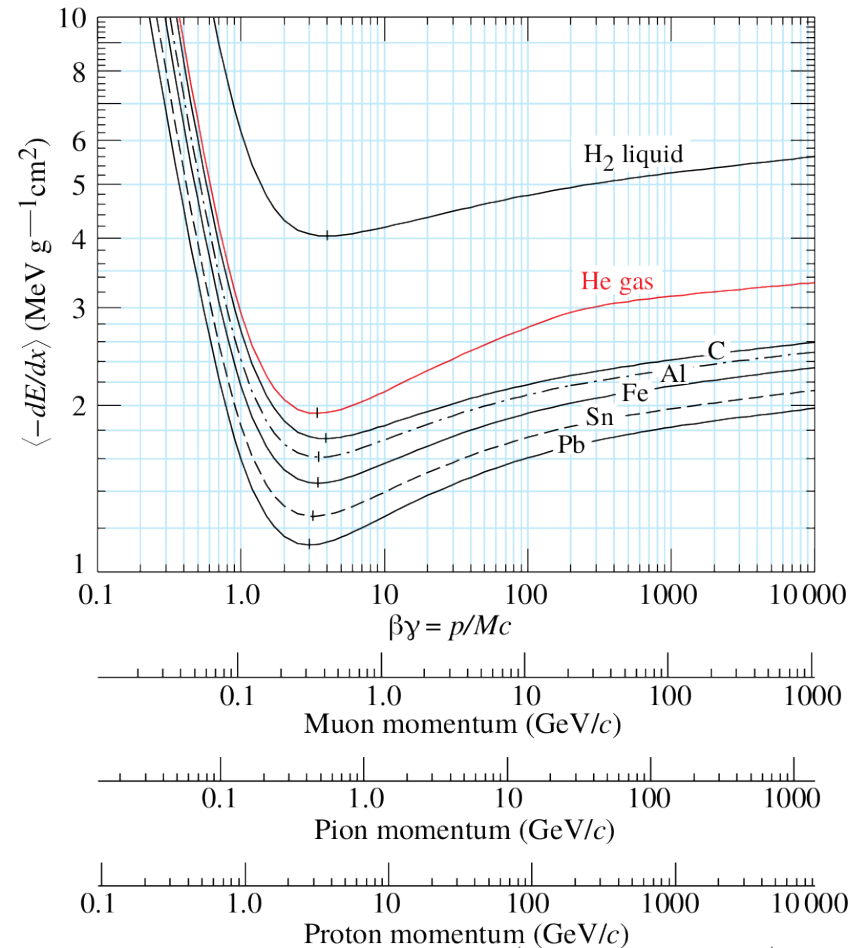


# Mean Energy Loss

- Different components are dominating:
  - At low energies:  $\sim 1/\beta^2$
  - At high energies:  $\sim \ln \gamma$

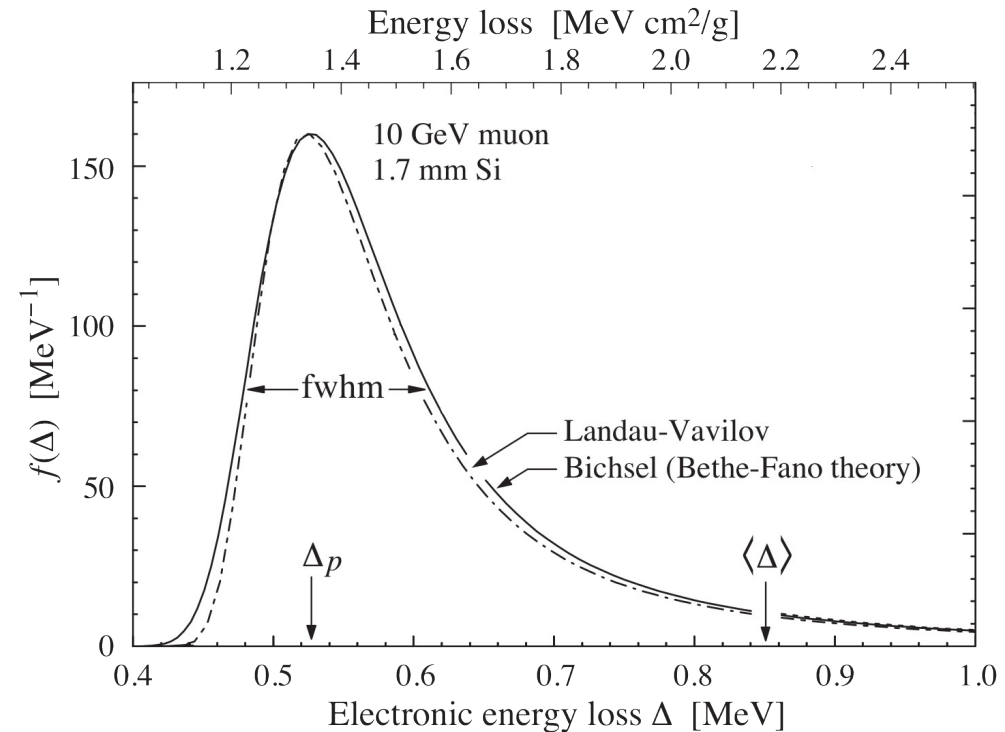
$$-\left\langle \frac{dE}{dx} \right\rangle \approx K q^2 \frac{1}{\beta^2} \frac{Z}{A} \left[ \ln \left( \frac{2 m_e c^2}{I} \beta^2 \gamma^2 \right) - \beta^2 \right]$$

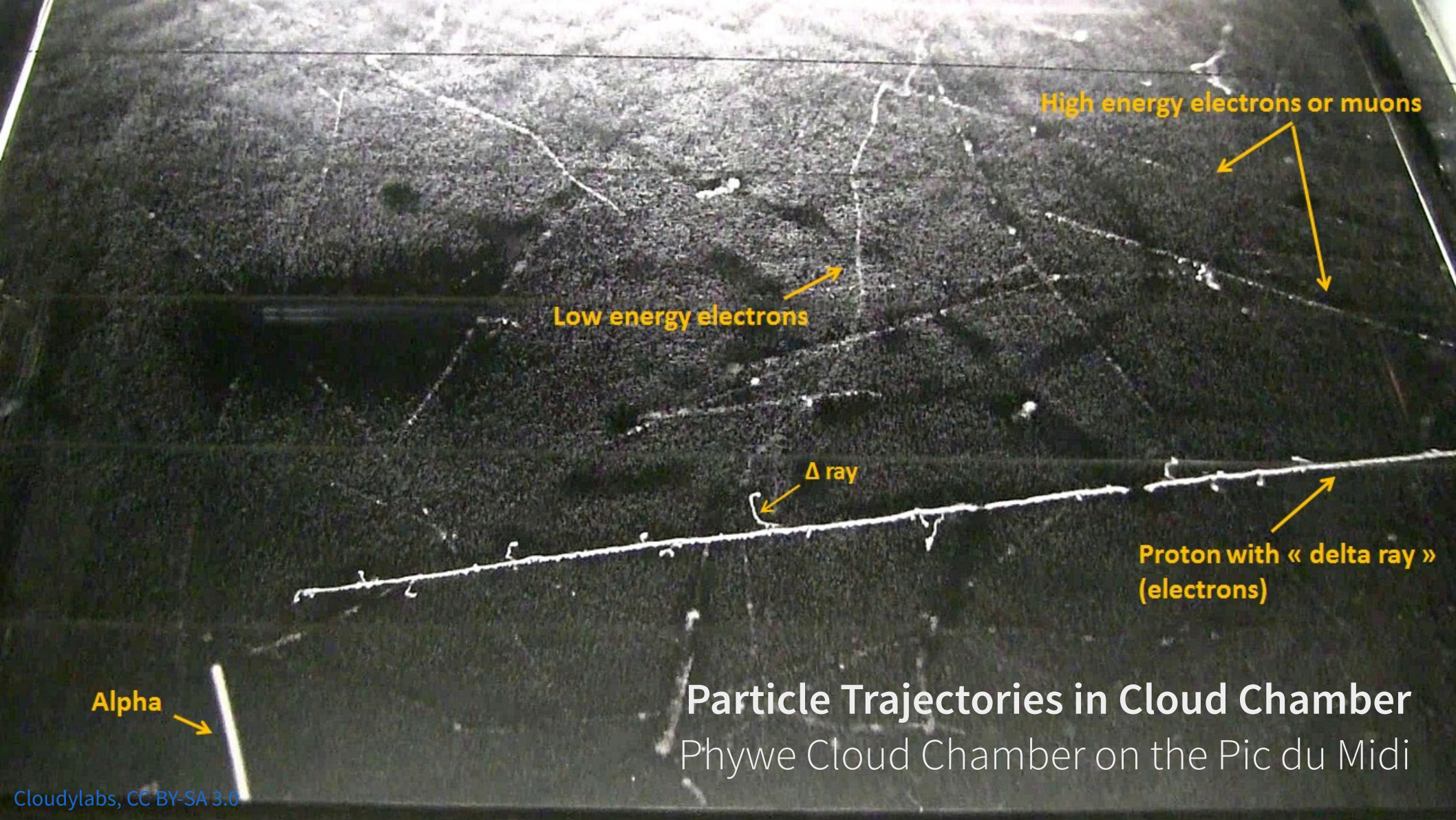
- Minimal energy loss for  $\sim \beta\gamma = 3$   
**MIP: Minimum Ionizing Particle**
- $[dx] = \text{g} / \text{cm}^2 = \text{cm} \times \text{g} / \text{cm}^3$



# Fluctuations in Energy Loss

- Actual energy loss fluctuates around mean value
- **Landau-Vavilov distribution** with long tails to high energies
  - Most probable value  $\ll$  mean value
- Orbital electrons can receive very large energy transfers
  - Creation of delta electrons
  - Delta electrons have enough energy for further ionization





High energy electrons or muons

Low energy electrons

$\Delta$  ray

Proton with « delta ray »  
(electrons)

Alpha

Particle Trajectories in Cloud Chamber  
Phywe Cloud Chamber on the Pic du Midi

# Exception: Electron & Positron

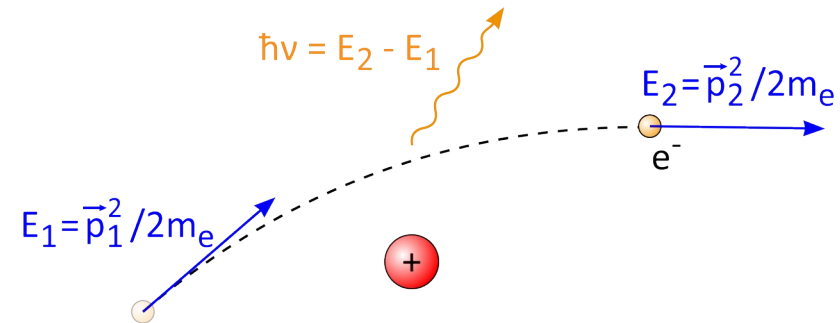
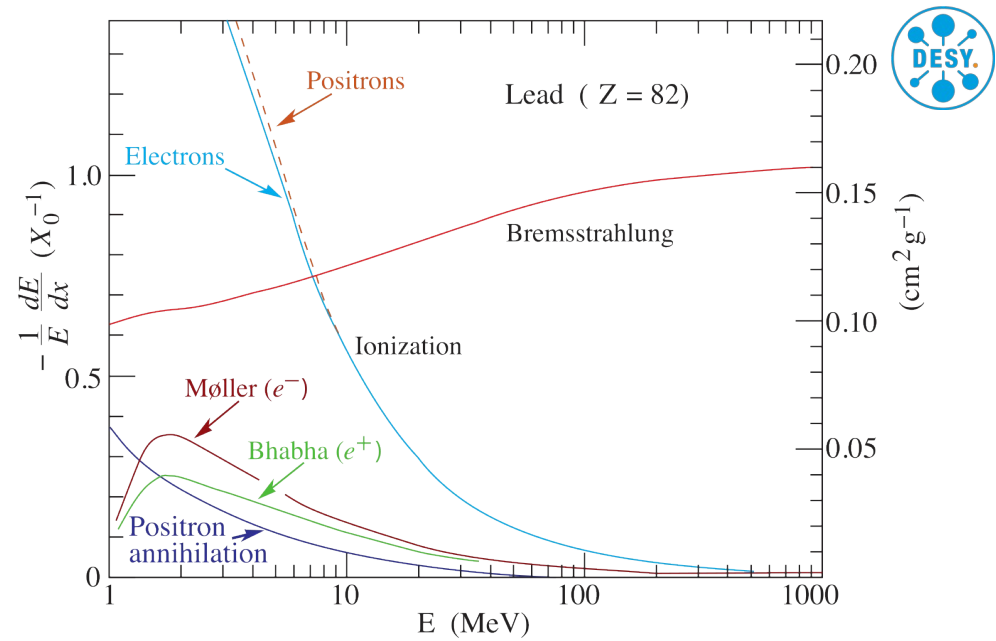
- Special case: low mass  
 $m_e = 0.511 \text{ MeV} / c^2$   
 $m_\mu = 106 \text{ MeV} / c^2 \approx 200 m_e$

- At high energies: Bremsstrahlung

- occurs when the momentum of a charged particle changes, e.g.  
 ...in the Coulomb field of a nucleus  
 ...in the magnetic field: synchrotron radiation

- Here: particles slowed down in matter

- Relevant for electrons:  $-\frac{dE}{dx} \sim E \cdot \frac{1}{m^2}$



# Cherenkov Radiation

- Is emitted, when particle velocity > speed of light

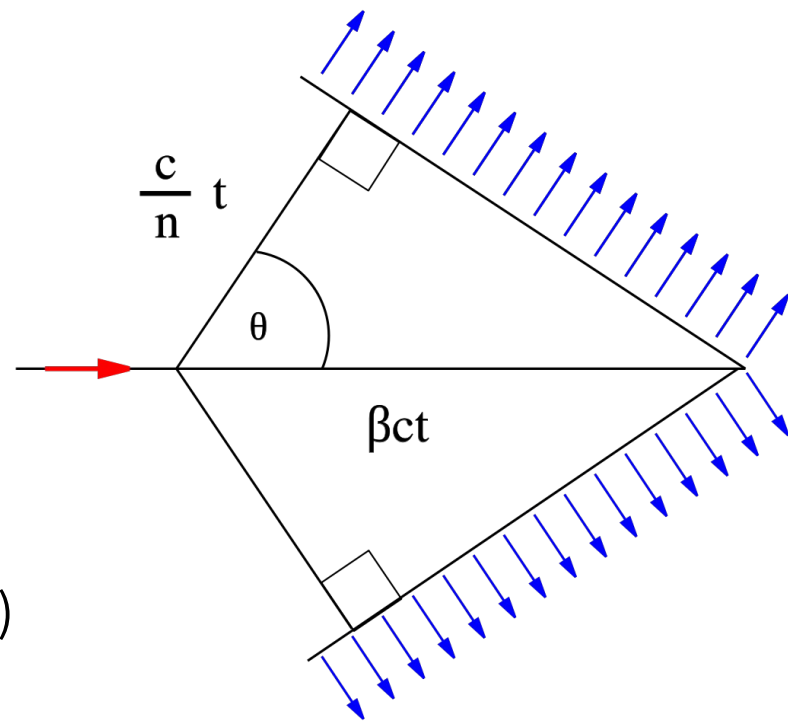
$$c_{\text{Medium}} = \frac{c_{\text{Vacuum}}}{n} < v_{\text{Particle}} < c_{\text{Vacuum}}$$

$n$ : refractive index of medium

- Electromagnetic shockwave with conical shape is emitted under angle  $\theta$ :

$$\cos(\theta) = \frac{1}{\beta n}, \quad \beta = \frac{v}{c}$$

- Very low energy loss (ca. 1% of total energy loss)



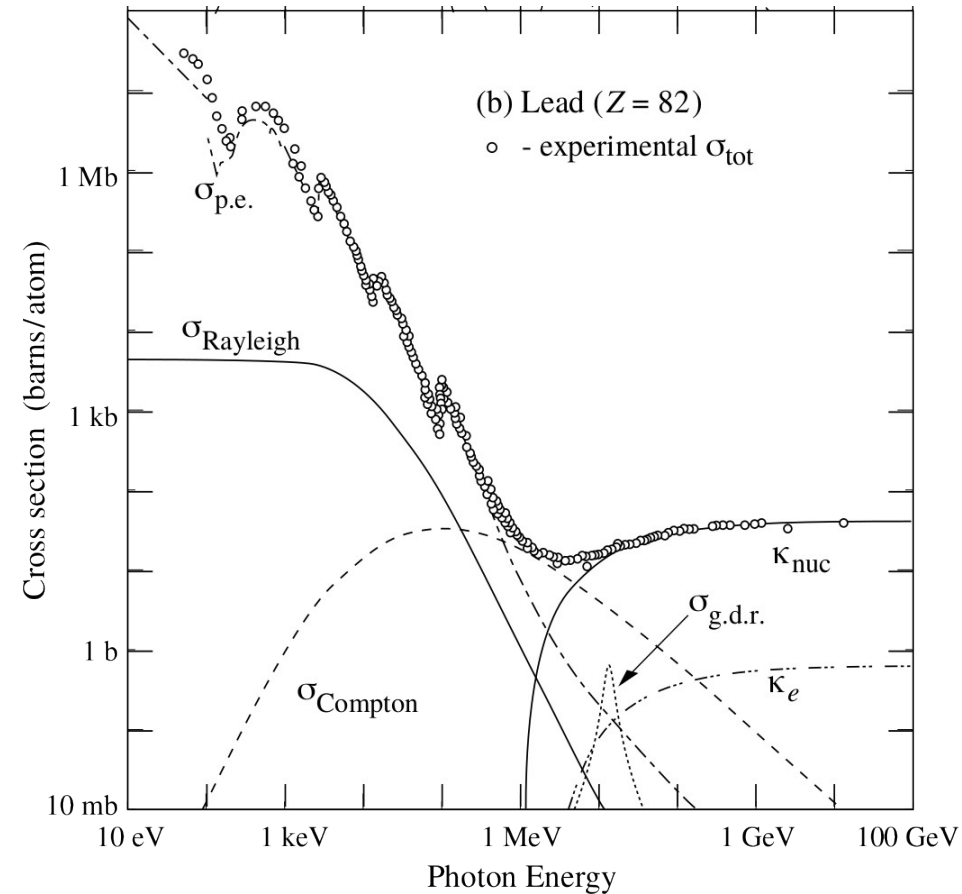


# Cherenkov Radiation of a Nuclear Reactor

Advanced Test Reactor, INL

# Photons

- Electromagnetic interaction
- Different processes dominate, depending on the photon energy:
  - Photo(electric) effect
  - Rayleigh scattering
  - Compton effect
  - Pair production

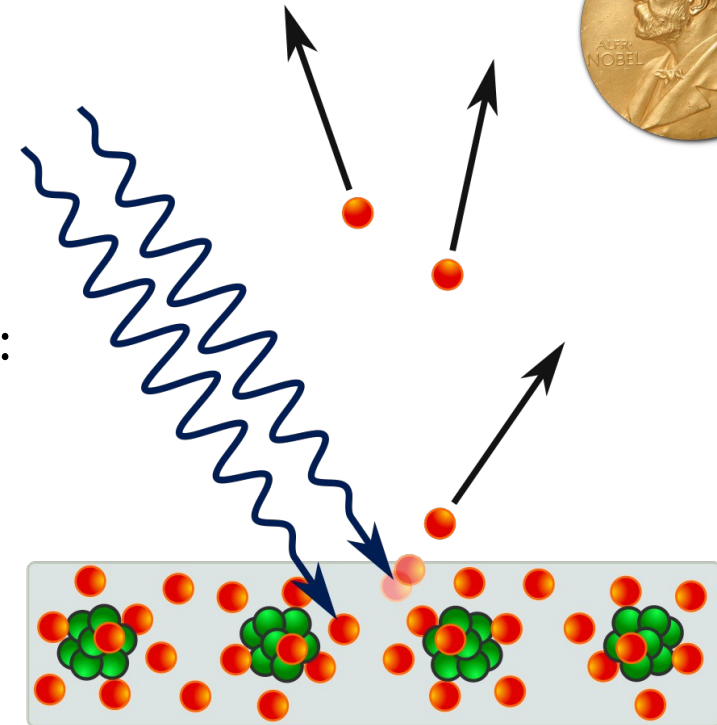


# Photoelectric Effect

- Theoretical description of photo effect:  
Einstein
- Photon is absorbed by electron in atomic shell
- Transferred energy releases/frees electron  

$$\gamma + Atom = e^- + Ion$$
- Process only possible in the field of the nucleus:
  - Momentum conservation
  - Nucleus absorbs recoil
- Cross section of the photoelectric effect shows shell structure of the atom

Nobel price 1921 for Albert



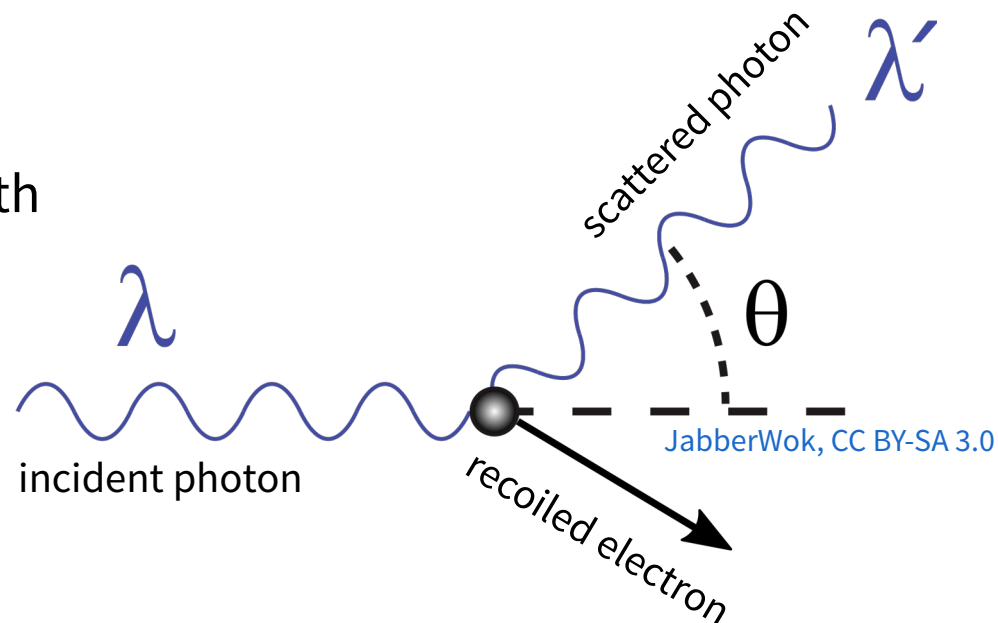
Ponor, CC BY-SA 4.0

# Compton Effect

- Describes scattering of photon on “quasi-free” electron

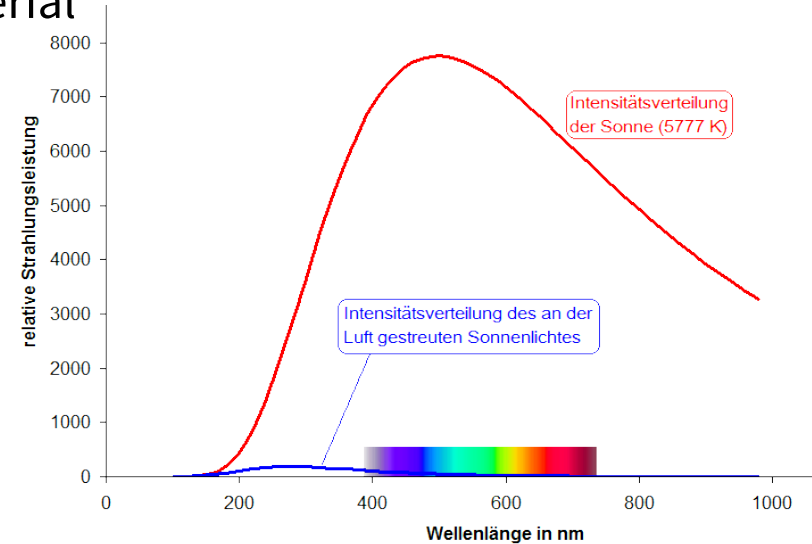
$$\gamma + Atom = \gamma + e^- + Ion$$

- Photon is deflected from its original path
- Wavelength of photon changes through energy transfer to electron



# Photons: Thomson/Rayleigh Scattering

- Elastic scattering: almost no energy transfer to material
- Thomson scattering:  
Photon scattering on free electron
  - Low-energy limit of Compton scattering
- Rayleigh scattering:  
Photon scattering off an entire atom
  - Scattering cross section  $\sigma_{\text{Rayleigh}} \sim f^4$
- Reason for blue / red coloration of the sky depending on the zenith angle
  - Noon: short path through atmosphere, hardly any blue light scattered
  - Morning/evening: long path through atmosphere, much blue light scattered



$$\frac{\sigma_{\text{blue}}}{\sigma_{\text{red}}} = \frac{1/\lambda_{\text{blue}}^4}{1/\lambda_{\text{red}}^4} = \left( \frac{650 \text{ nm}}{450 \text{ nm}} \right)^4 \approx 4.4$$

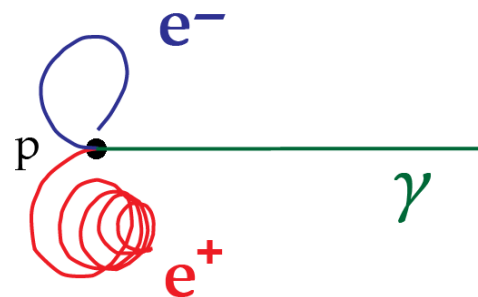
# Photons: Pair Production

- Pair production is the generation of an electron-positron pair by the photon
- Pair production occurs in the field of a partner, which absorbs the recoil (atomic nucleus, but also shell electron)

$$\gamma + p = e^+ + e^- + p$$

- Photon must provide at least rest mass of e<sup>+</sup>e<sup>-</sup> pair plus recoil energy:

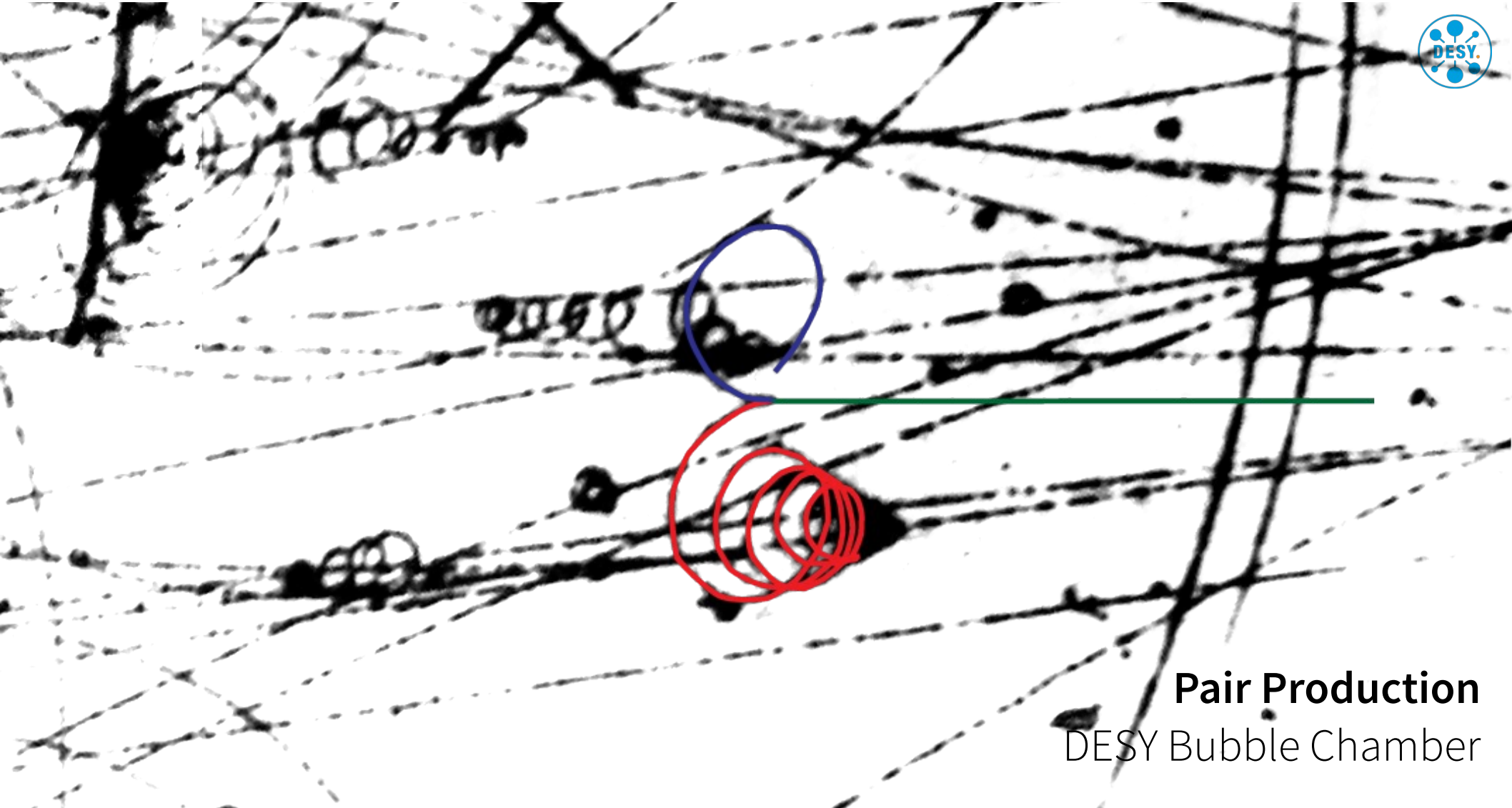
$$E_\gamma \geq 2m_e c^2 \left(1 + \frac{m_e}{M}\right)$$



$$\gamma p \rightarrow e^- e^+ p$$

Ivan Baev, CC BY-SA 3.0

- Recoil energy can often be neglected, e.g. Ge detector:  $\frac{m_e}{M} \approx 7.6 \cdot 10^{-6}$



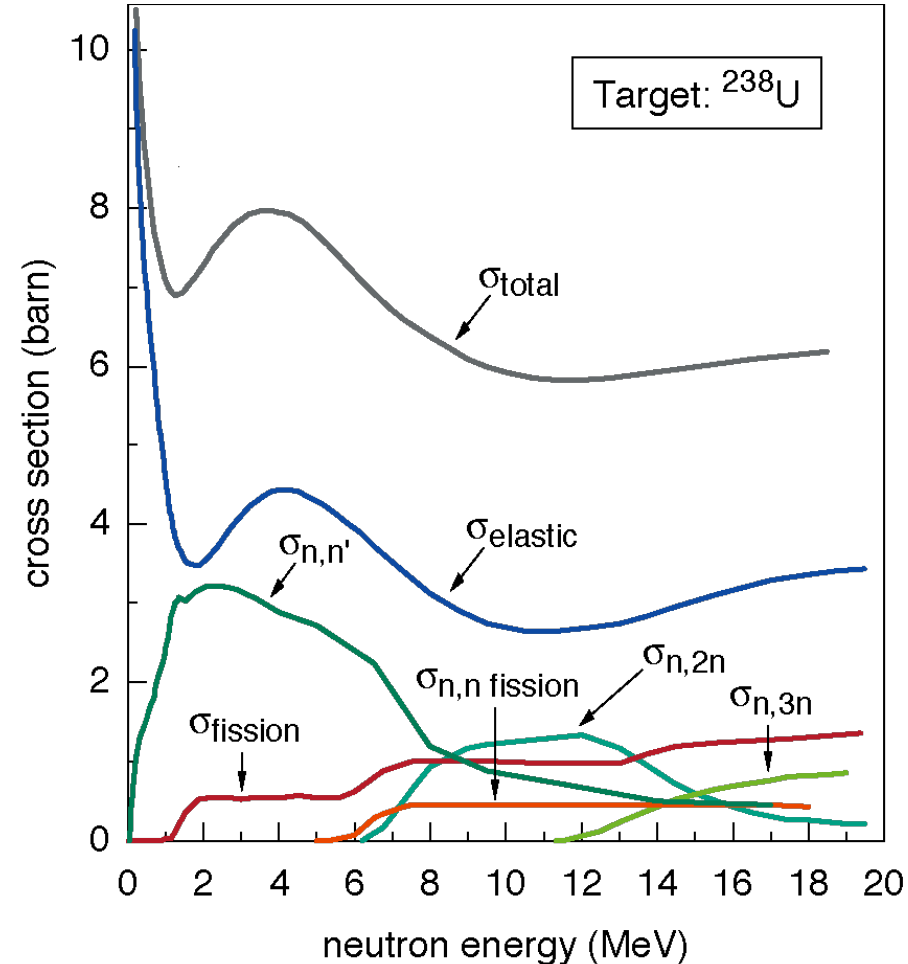
# Pair Production

## DESY Bubble Chamber

# Hadronic Interaction

- Interactions of hadrons with nuclei
- Based on the **strong interaction, QCD**
  - Low range
  - Low probability of hadronic reactions
  - Neutrons can only interact strongly: very penetrating
- Many possible processes (energy dependent)

Elastic, inelastic scattering; neutron capture; reactions with radiation of charged particles; nuclear fission.



# Interaction of Neutrinos

- Neutrinos are exclusively subject to the weak interaction
- Possible Interactions

$$\bar{\nu}_l + p \rightarrow n + l^+$$

$$\nu_l + n \rightarrow p + l^-$$

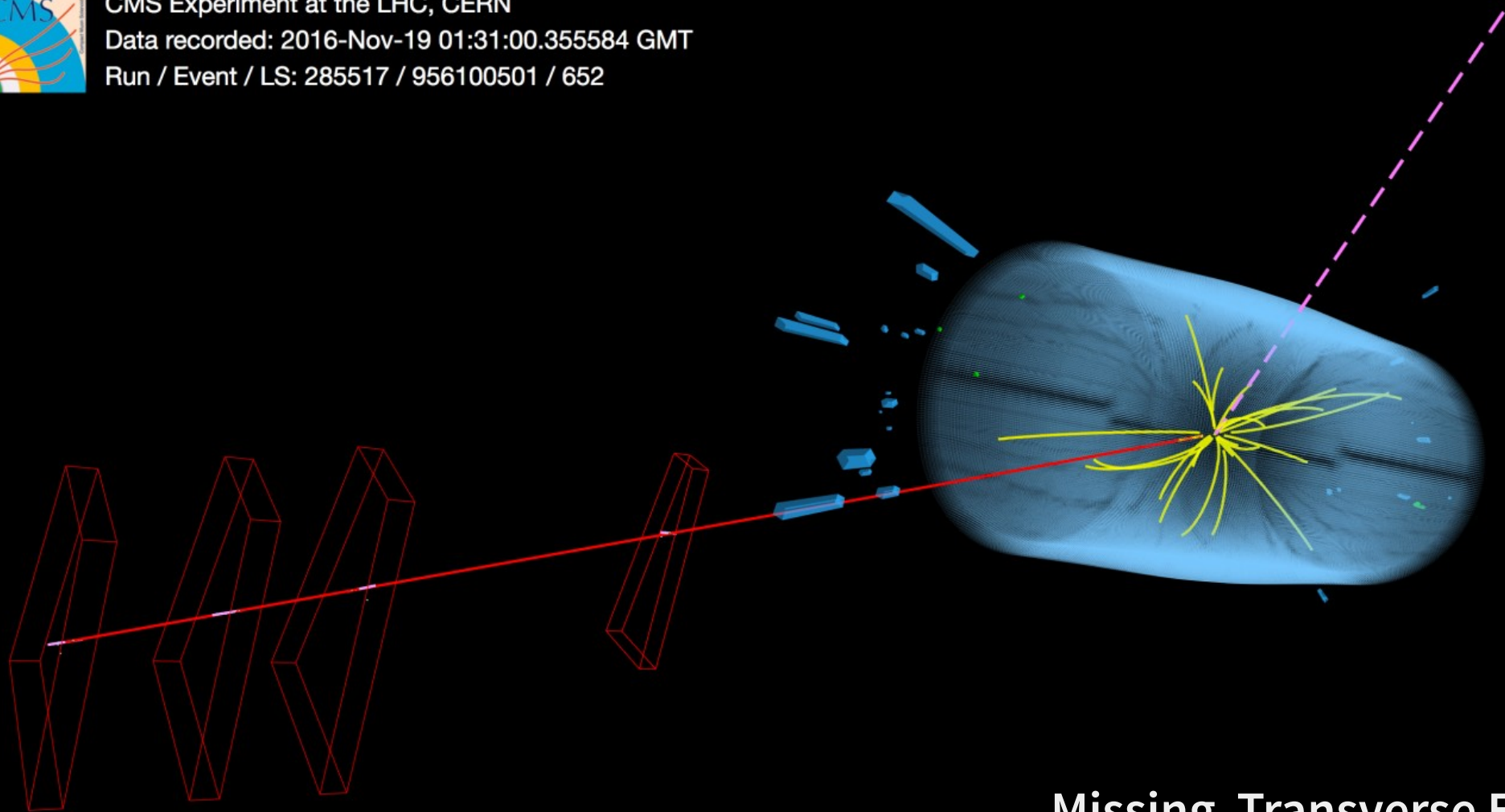
- Neutrino interactions have very (very!) low cross sections
- Detection of neutrinos requires
  - Very large detector and high neutrino fluxes (direct detection) or
  - Hermetic detector for measurement of missing energy (indirect evidence)



CMS Experiment at the LHC, CERN

Data recorded: 2016-Nov-19 01:31:00.355584 GMT

Run / Event / LS: 285517 / 956100501 / 652



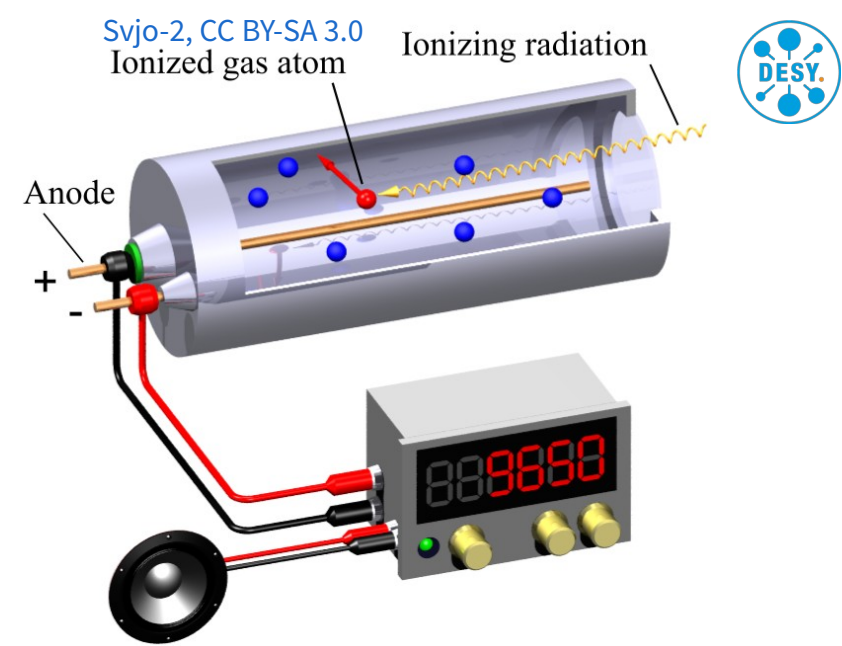
**Missing Transverse Energy**  
Detection of Neutrinos

# Particle Detectors

## Historic Overview

# Geiger-Müller Counter

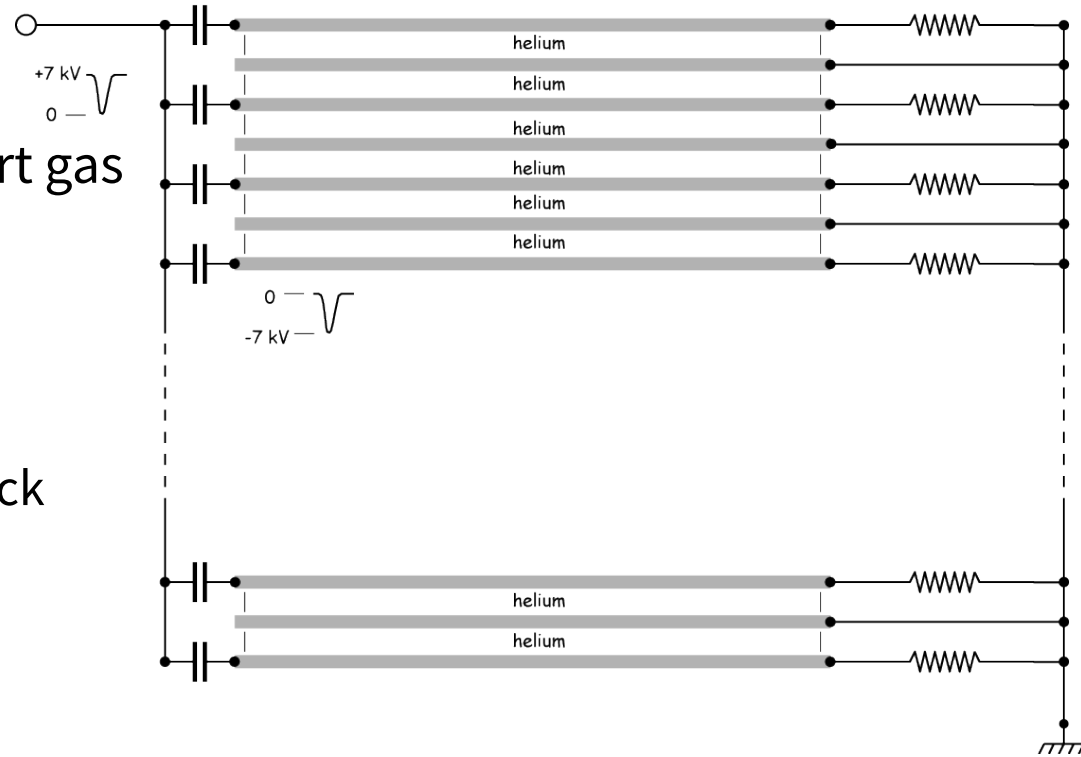
- „Click“ = Particle passing
- „Many Click“ = Many particles
- „A Great Many Click“ = Run...!
- Detector filled with noble gas
  - Charged particles ionize noble gas atoms
  - High voltage applied between electrodes amplifies
- Signal: Current pulse to loudspeaker: Click!



Geiger's lab device  
Science Museum London, CC BY-SA 2.0

# Spark Chamber

- „Many Geiger counters“
- Transparent chamber filled with inert gas and many parallel plates
  - Voltage ( $\sim$  kV) between plates
  - Particles ionize noble gas atoms
  - Small sparks along the particle track
- Analysis by photos or microphones
- Relatively large dead time: quenching time of the avalanche
- Used during the 1930s-1960s





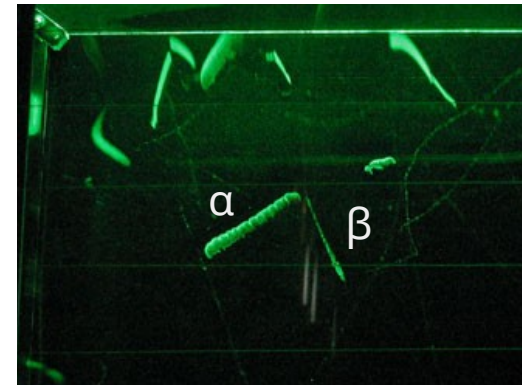
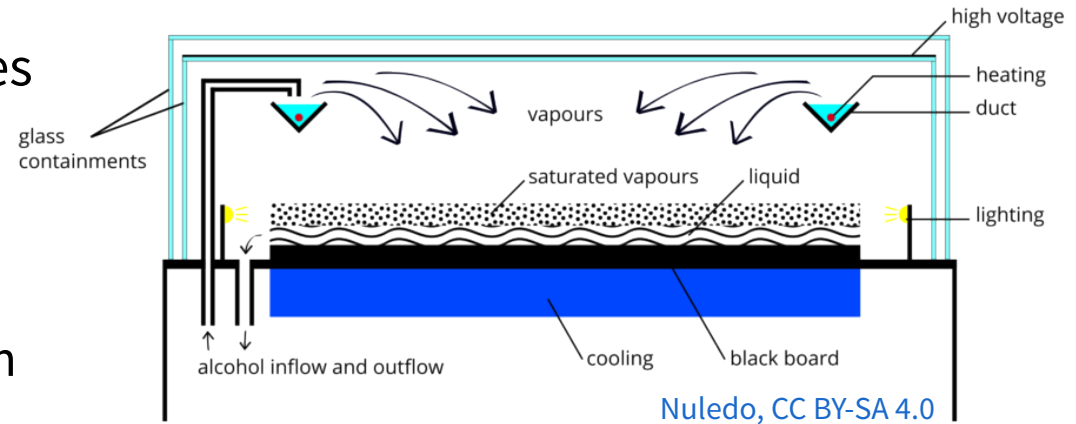
**Spark Chamber**

Melvin Schwartz at BNL AGS

# Cloud Chamber



- Optical detection of charged particles
  - Transparent chamber with supersaturated air-alcohol mixture
  - High-energy charged particle generates ions by impact ionization
  - Ions act as condensation nuclei, droplet formation in the gas mixture
- Type of traces can (sometimes) be identified & assigned to particles
- Nobel Prize 1927 for Charles Thomson Rees Wilson





**Cloud Chamber Image**  
Rendering of particle interaction

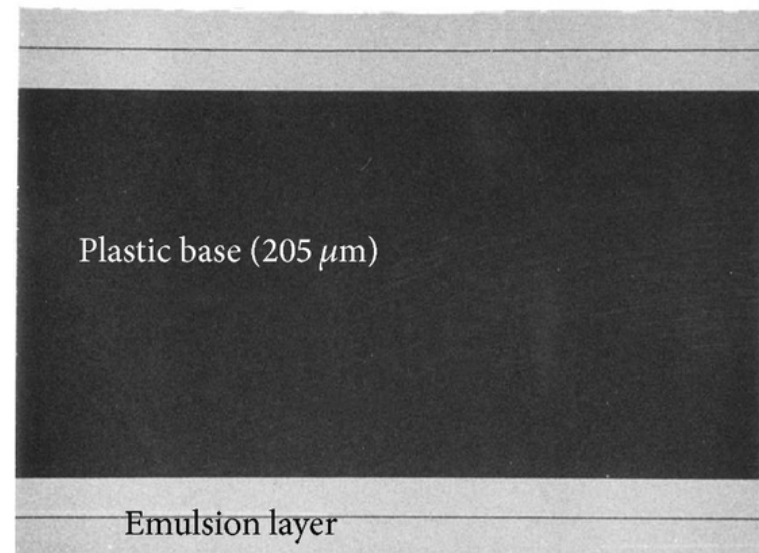


**Detail of droplets**

Cloud chamber tracks after Radon injection

# Photo emulsion / Nuclear emulsion

- Photographic plate with thick sensitive layer and very uniform grain size.
  - Ionizing radiation leaves traces
  - Development of the plate
  - Trajectories of particles (blackened by silver) visible with microscope



OPERA

- Nobel Prize 1950 for Cecil Powell

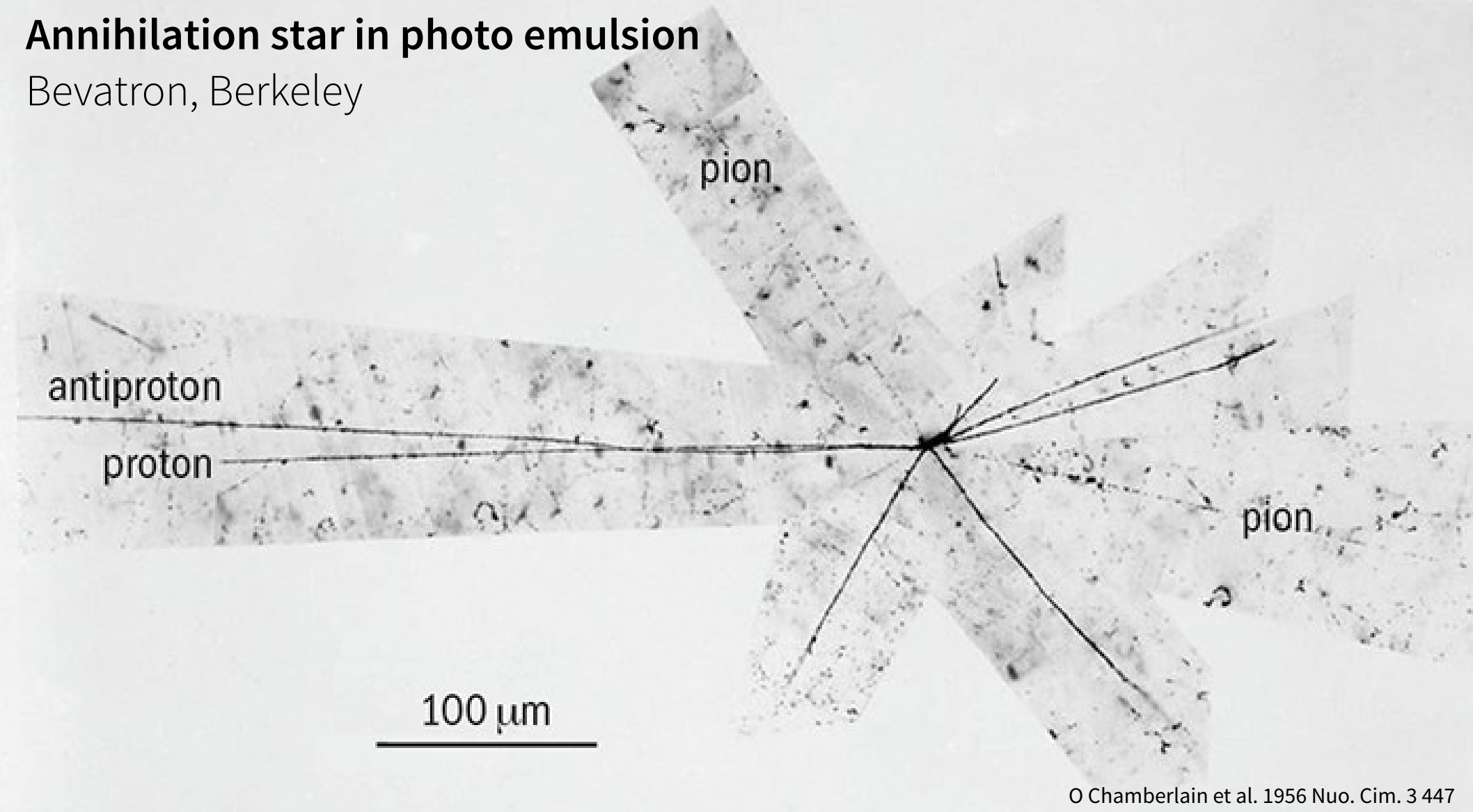


- OPERA experiment in Gran Sasso

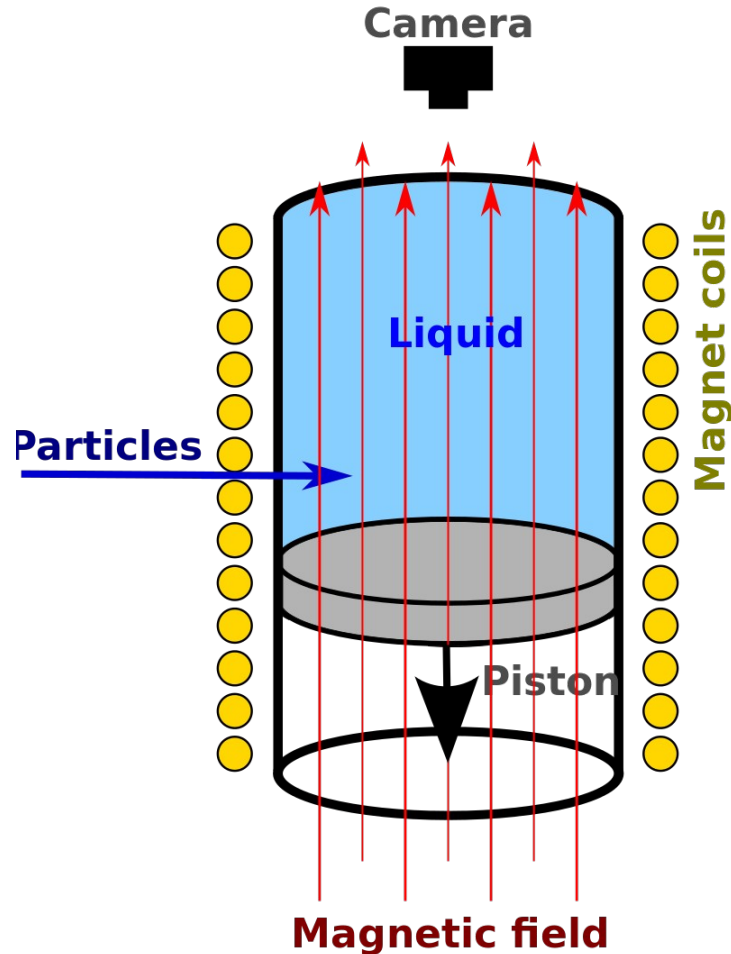
- Photoemulsion plates for particle reconstruction
- AgBr emulsion, cooperation with Fuji Film: 9 million films, each about 10 x 12 cm

# Annihilation star in photo emulsion

Bevatron, Berkeley



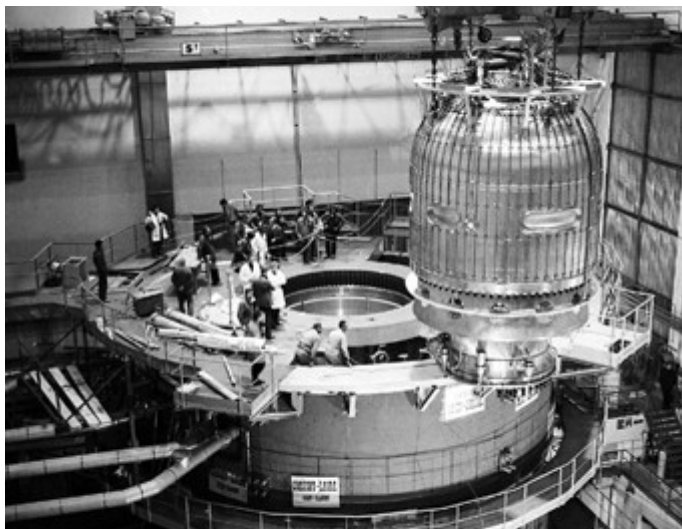
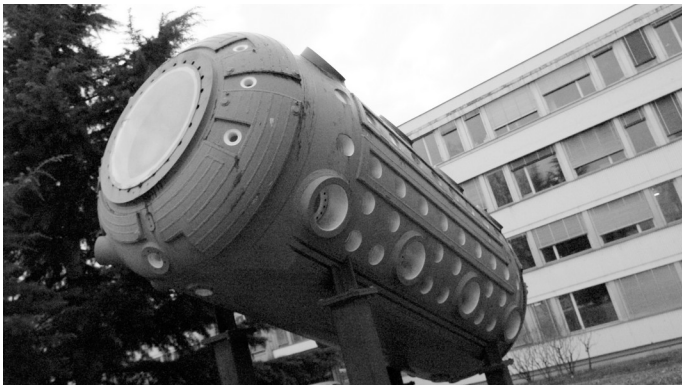
# Bubble Chamber



- Optically transparent chamber filled e.g. with liquid hydrogen
  - Temperature of the liquid close to the boiling point
  - Reduction of chamber pressure with piston
  - Temperature of liquid is now above boiling point
  - Charged particles generate ions along track
  - Ions serve as nuclei for gas bubbles
- Analysis of photos of the traces
- Nobel Prize 1960 for Donald A. Glaser



# Prominent Bubble Chambers



- Most important particle detector type in the 1970s
- Particle source: proton synchrotron
- **Gargamelle** - 1970-1978
  - 4.8 m x 1.88 m, 12 000 l
  - First detection of the Z boson
- **BEBC** - 1971-1984
  - Big European Bubble Chamber
  - 3.5 T magnetic field of superconducting coils
  - Discovery of the D meson

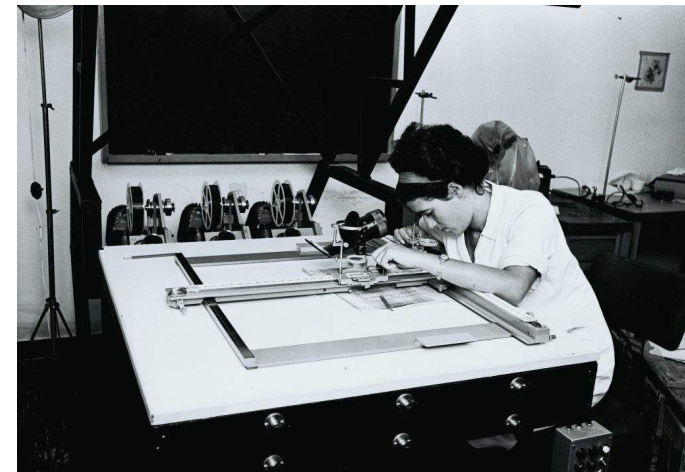
# Bubble Chamber Image

Big European Bubble Chamber @ CERN



# Problems of Early Particle Detectors

- Many based on photographic images
  - Data evaluation complex and only possible by hand
  - Limitation of the amount of data
- Long detector dead times
  - No further measurement possible until gas avalanche has been quenched/bubbles have disappeared
  - Only low particle rates possible
- Only little information in measurement
  - Information about particle location (& momentum), but no energy/time measurement



Weizmann Institute Archives

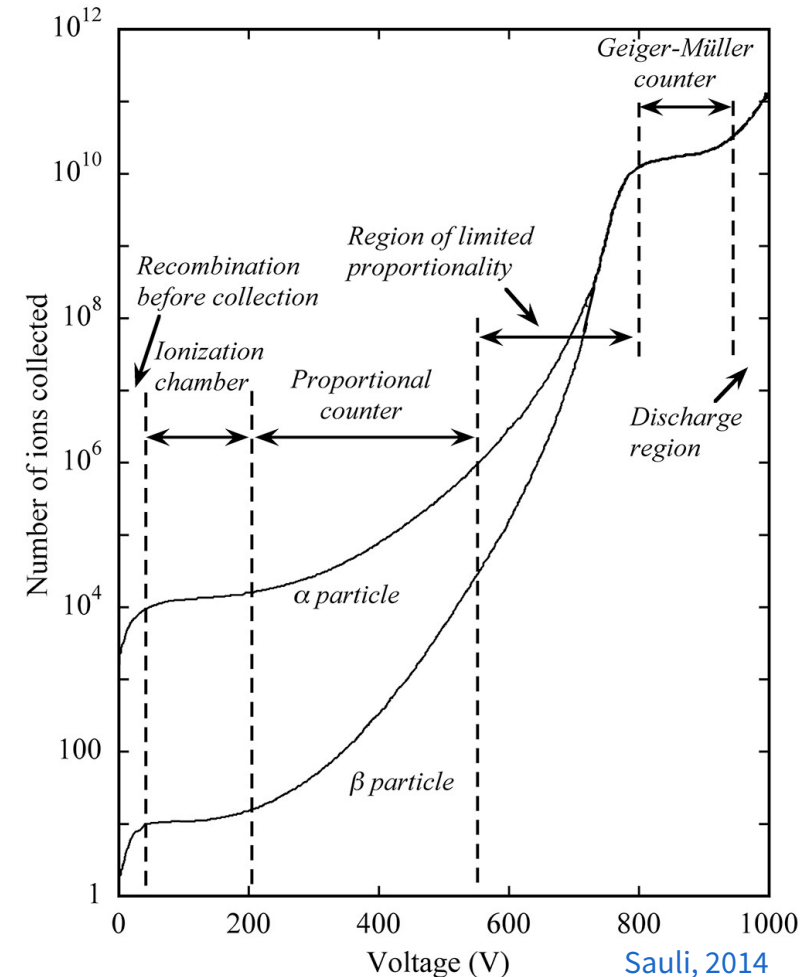


# Gaseous Detectors

Primary & secondary ionization of gas atoms

# Gaseous Detectors: Operating Principle

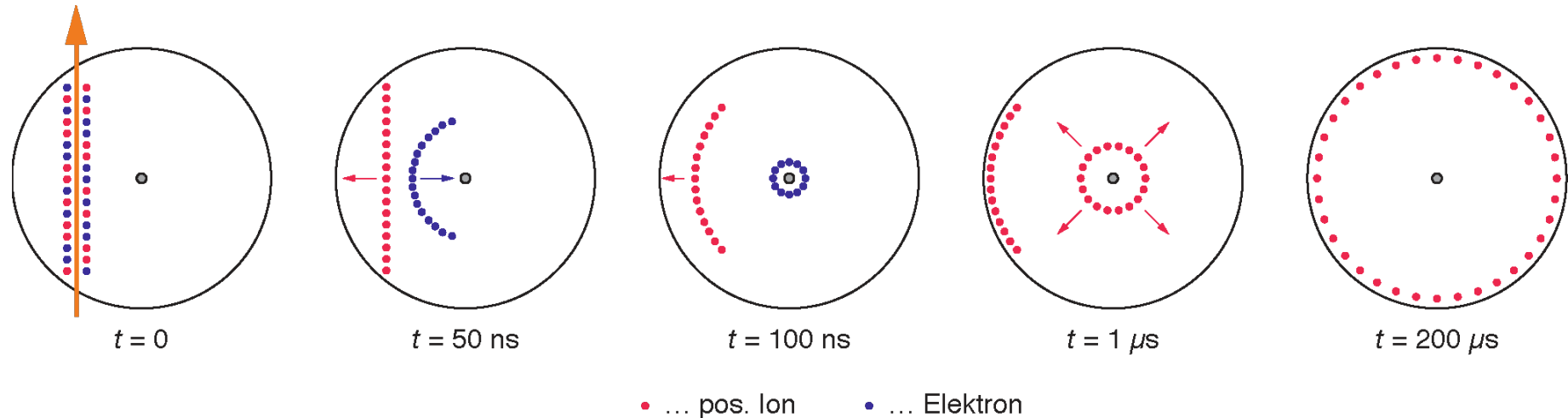
- **Primary signal:** charged particle generates electron-ion pairs by ionization
  - Noble gases: relatively low ionization energy
  - Average energy to generate a pair  $\sim 30$  eV
  - Number proportional to deposited energy
- **Amplification:** different working ranges depending on applied voltage
  - Medium voltages: proportional amplification
  - High voltages: Avalanche formation due to secondary ionizations



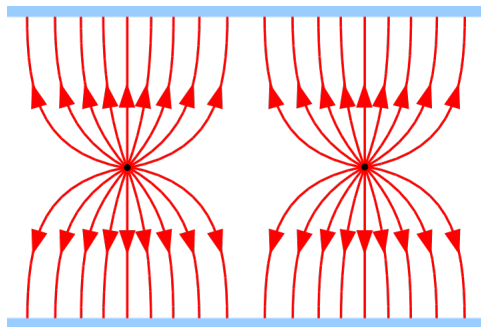
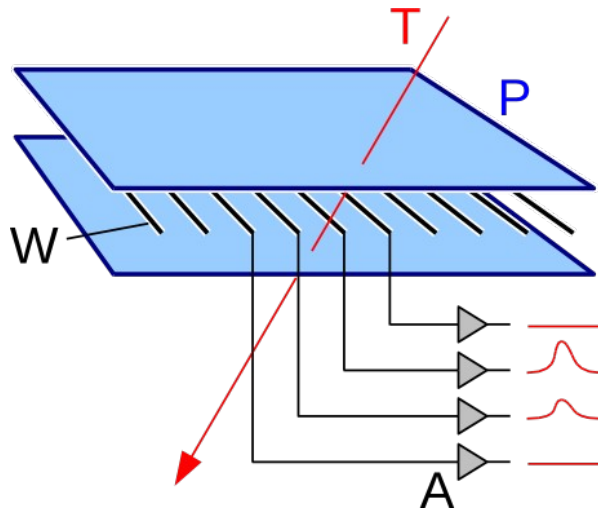
Sauli, 2014

# Proportional Counter

- Very similar to the Geiger counter: anode in the form of thin wire
- High field near wire leads to electron multiplication / signal amplification
- Choice of voltage: proportional range
  - Output signal proportional to original number of ionizations



# Multi Wire Proportional Chamber (MWPC)

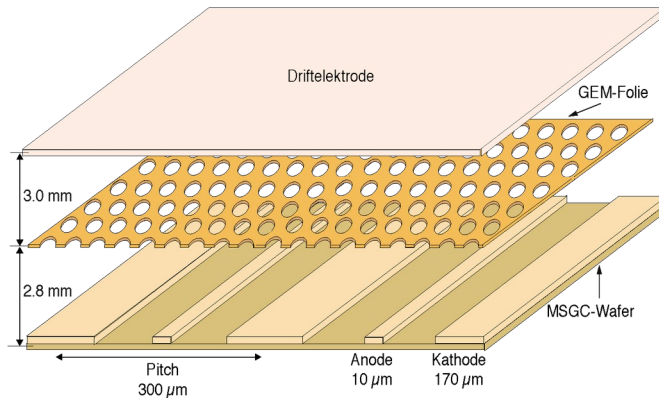


- Essentially many proportional counters next to each other, without separating walls
- Wires spaced a few millimeters apart
  - Good spatial resolution of a traversing particle
  - Large areas possible
  - Electronic selection
- High rates possible: 1000 particles/s  
for comparison, bubble chamber: 1-2 particles/s
- Nobel Prize 1992 for Georges Charpak



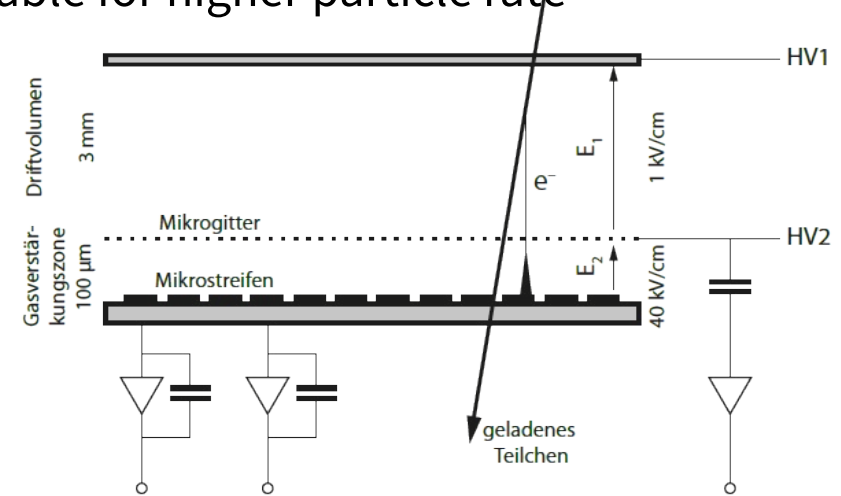
# Micropattern Gas Detectors

- Replacement of fragile wires by micro structures
- Potentially better spatial resolution and applicable for higher particle rate



## Gas Electron Multiplier (GEM)

- Perforated, metalized Kapton foil, High voltage between electrodes
- Strong dipole field in perforation holes: Gas amplification

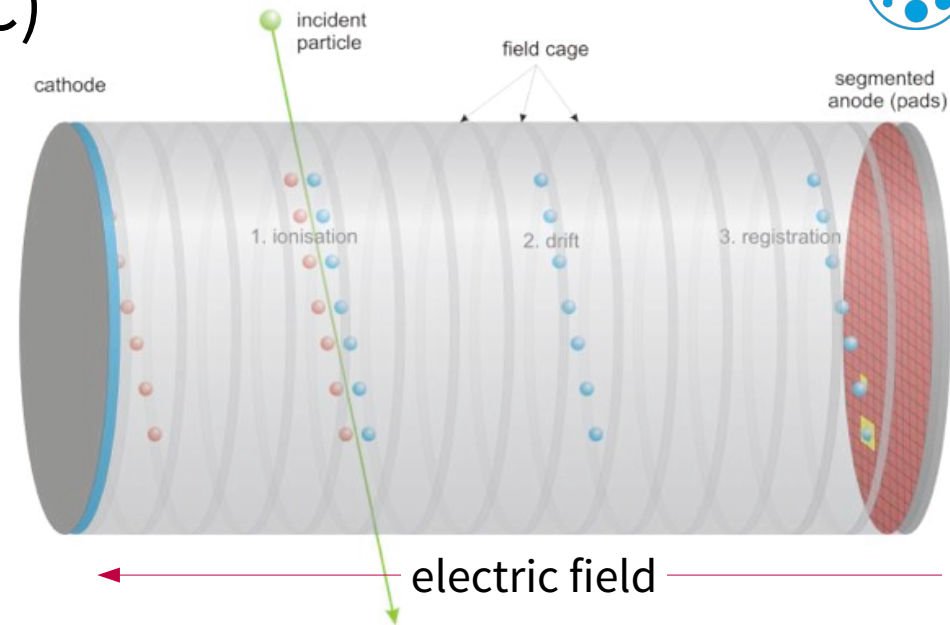


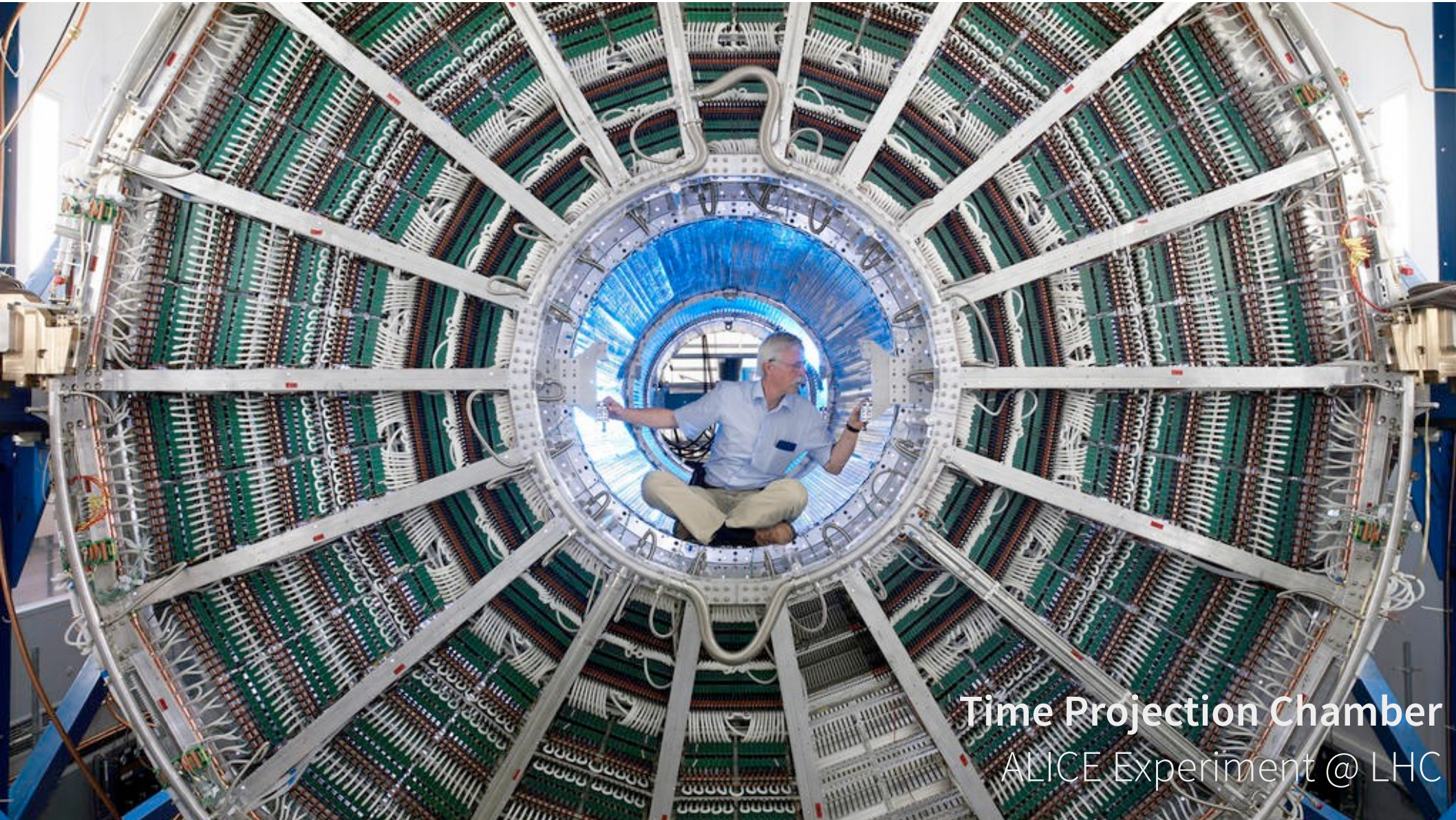
## Micro-Mesh Gas Detectors (Micromegas)

- Metallic micro-grid
- Electron avalanche evolution near the lattice: Gas amplification

# Time Projection Chambers (TPC)

- Large gas detector system
- Ionization along the particle track
  - Electrons and ions drifting in the E-field
  - Segmented anode: 2D information
  - Measurement of drift time: 3D information
- Readout at anode side e.g. via multi-wire proportional chamber, GEMs, ...





Time Projection Chamber  
ALICE Experiment @ LHC

# Summary

- Particle detection
  - Measurement is performed by interaction of the particle to be measured with detector material
  - There is no single detector concept that can detect all particle types / properties
  - One needs several detector technologies and concepts
- Interaction of radiation and matter
  - Ionization and excitation of detector atoms, Bremsstrahlung, Cherenkov radiation
  - Photoelectric effect, Compton effect, pair production,
  - Hadronic interaction, missing energy, ...
- Historical detectors
  - Cloud chamber, photo emulsion, bubble chamber, spark chamber
- Gas detectors
  - Operating principle
  - Proportional counters, multi-wire proportional chamber
  - Micropattern gas detectors
  - Time Projection Chamber

# Outlook

