







Engineering and Physical Sciences Research Council

э.

Quantum field theory, quantum reference frames and the type of local algebras

CJ Fewster

Department of Mathematics and YCQT, University of York

DPG-Frühjahrstagung (SMuK), Göttingen April 2025

Comm. Math. Phys. **406**:19 (2025) arXiv:2403.11973 with D Janssen, LD Loveridge, K Rejzner, J Waldron See also a more technical talk in session MP10, 1400 Thursday 3 April

 A remarkable result in mathematical QFT is that all local observable algebras are isomorphic – they are von Neumann factors of the specific hyperfinite type III₁ Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987



з.

- A remarkable result in mathematical QFT is that all local observable algebras are isomorphic – they are von Neumann factors of the specific hyperfinite type III₁ Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987
- Robust result, holding under mild physical conditions even in fixed curved spacetimes.

з.

- A remarkable result in mathematical QFT is that all local observable algebras are isomorphic – they are von Neumann factors of the specific hyperfinite type III₁ Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987
- Robust result, holding under mild physical conditions even in fixed curved spacetimes.
- Recent work has reconsidered the issue, taking into account features of (quantum) gravity, and finding algebras of type II₁ or II_∞.
 Chandrasekaran, Longo, Penington, Witten; Witten; Kudler-Flam, Leutheusser, Satishchandran

= nar

- A remarkable result in mathematical QFT is that all local observable algebras are isomorphic – they are von Neumann factors of the specific hyperfinite type III₁ Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987
- Robust result, holding under mild physical conditions even in fixed curved spacetimes.
- Recent work has reconsidered the issue, taking into account features of (quantum) gravity, and finding algebras of type II₁ or II_∞.
 Chandrasekaran, Longo, Penington, Witten; Witten; Kudler-Flam, Leutheusser, Satishchandran
- These differences sound (and are) technical but have physical consequences, casting light on the degrees of freedom available in theories of quantum gravity and bearing on the interaction between measurement and symmetry.

- A remarkable result in mathematical QFT is that all local observable algebras are isomorphic – they are von Neumann factors of the specific hyperfinite type III₁ Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987
- Robust result, holding under mild physical conditions even in fixed curved spacetimes.
- Recent work has reconsidered the issue, taking into account features of (quantum) gravity, and finding algebras of type II₁ or II_∞.
 Chandrasekaran, Longo, Penington, Witten; Witten; Kudler-Flam, Leutheusser, Satishchandran
- These differences sound (and are) technical but have physical consequences, casting light on the degrees of freedom available in theories of quantum gravity and bearing on the interaction between measurement and symmetry.
- This talk: describe background, demystify (factors? types? quantum reference frames?) for these results, and briefly present our extension of the above results.

In this talk, QFT is described in the framework of algebraic QFT.

'Algebra first' approach

think Heisenberg rather than Schrödinger



In this talk, QFT is described in the framework of algebraic QFT.

'Algebra first' approach

think Heisenberg rather than Schrödinger

 Axiomatic rather than constructive; not Lagrange-centric theory is assumed ready-made, can describe non-Lagrangian theories

In this talk, QFT is described in the framework of algebraic QFT.

'Algebra first' approach

think Heisenberg rather than Schrödinger

- Axiomatic rather than constructive; not Lagrange-centric theory is assumed ready-made, can describe non-Lagrangian theories
- Algebra, states and representations separated as different levels of structure cf conventional QFT; facilitates treatment of unitarily inequivalent representations

з.

In this talk, QFT is described in the framework of algebraic QFT.

'Algebra first' approach

think Heisenberg rather than Schrödinger

- Axiomatic rather than constructive; not Lagrange-centric theory is assumed ready-made, can describe non-Lagrangian theories
- Algebra, states and representations separated as different levels of structure cf conventional QFT; facilitates treatment of unitarily inequivalent representations
- Structural, conceptual and partly operational viewpoint

з.

In this talk, QFT is described in the framework of algebraic QFT.

'Algebra first' approach

think Heisenberg rather than Schrödinger

- Axiomatic rather than constructive; not Lagrange-centric theory is assumed ready-made, can describe non-Lagrangian theories
- Algebra, states and representations separated as different levels of structure cf conventional QFT; facilitates treatment of unitarily inequivalent representations
- Structural, conceptual and partly operational viewpoint
- Technicalities: operator algebras

In this talk, QFT is described in the framework of algebraic QFT.

'Algebra first' approach

think Heisenberg rather than Schrödinger

- Axiomatic rather than constructive; not Lagrange-centric theory is assumed ready-made, can describe non-Lagrangian theories
- Algebra, states and representations separated as different levels of structure cf conventional QFT; facilitates treatment of unitarily inequivalent representations
- Structural, conceptual and partly operational viewpoint
- Technicalities: operator algebras

Nonetheless: specific theories can be constructed in AQFT framework, AQFT has led to new constructive techniques Lechner, Bostelmann, Cadamuro, Tanimoto,... interacting theories can be constructed in pAQFT Brunetti, Fredenhagen, Hollands, Wald, Duetsch, Rejzner...

Vibe Why algebra? AQFT

The role of algebra



CJ Fewster QFT, QRF and types



Vibe Why algebra? AQFT

The role of algebra



・ロト・日本・山田・ 山田・ 山田・



The role of algebra



$$\langle C \rangle = p \langle A \rangle + (1-p) \langle B \rangle \longrightarrow C = p \langle A + (1-p) B \rangle$$

CJ Fewster QFT, QRF and types

а.

The set of observables should support functional calculus and convex combinations, from which it follows that also linear combinations are supported.





The set of observables should support functional calculus and convex combinations, from which it follows that also linear combinations are supported.

We can construct a symmetric Jordan product using these operations

$$A \circ B := \frac{(A+B)^2 - (A-B)^2}{4}$$

5/23

The set of observables should support functional calculus and convex combinations, from which it follows that also linear combinations are supported.

We can construct a symmetric Jordan product using these operations

$$A \circ B := rac{(A+B)^2 - (A-B)^2}{4}$$

Given an antisymmetric commutator as well, we define an algebra product

$$AB := A \circ B + \frac{1}{2}[A, B].$$

Ż

A *-operation is natural if we allow complex-valued measurement outcomes.

The set of observables should support functional calculus and convex combinations, from which it follows that also linear combinations are supported.

We can construct a symmetric Jordan product using these operations

$$A \circ B := rac{(A+B)^2 - (A-B)^2}{4}$$

Given an antisymmetric commutator as well, we define an algebra product

$$AB := A \circ B + \frac{1}{2}[A, B].$$



э.

A *-operation is natural if we allow complex-valued measurement outcomes.

Conclusion Quantum observables naturally generate a unital *-algebra.

Additional structure permits varying strength of functional calculus.

-algebra A complex vector space, with product (associative & distributive) and an antilinear involution $A \mapsto A^$, and a unit **1**.

$$\lambda A$$
, $A + B$, AB , A^*
 $A\mathbf{1} = A$, $(A + B)C = AC + BC$, $(AB)^* = B^*A^*$, ...



▲□▶▲□▶▲□▶▲□▶ □ の00

Additional structure permits varying strength of functional calculus.

*-algebra supports polynomial functions





Additional structure permits varying strength of functional calculus.

- *-algebra supports polynomial functions
- ► C*-algebra A *-algebra with a complete norm || || obeying

 $||AB|| \leq ||A|| ||B||, \qquad ||A^*A|| = ||A||^2$

Equivalently, a norm-closed self-adjoint subalg. of $\mathcal{B}(\mathcal{H})$ for some Hilbert space \mathcal{H} .

Additional structure permits varying strength of functional calculus.

- *-algebra supports polynomial functions
- C*-algebra supports continuous functions



Additional structure permits varying strength of functional calculus.

- *-algebra supports polynomial functions
- C*-algebra supports continuous functions
- ▶ von Neumann algebra/ W^* -algebra A *-subalgebra $\mathcal{M} \subset \mathcal{B}(\mathcal{H})$ such that

$$\mathcal{M} = \mathcal{M}'', \quad \text{where} \quad \mathcal{N}' = \{A \in \mathcal{B}(\mathcal{H}) : [A, B] = 0 \ \forall B \in \mathcal{N}\}$$

Equivalently, a weakly closed unital *-subalg. of $\mathcal{B}(\mathcal{H})$.

= nar

Additional structure permits varying strength of functional calculus.

- *-algebra supports polynomial functions
- C*-algebra supports continuous functions
- von Neumann algebra/W*-algebra supports bounded functions

= nar

An AQFT on spacetime **M** comprises...

Μ

• a C^* -algebra $\mathcal{A}(\mathbf{M})$ for the whole spacetime



э

An AQFT on spacetime \boldsymbol{M} comprises...



- ▶ a C^* -algebra $\mathcal{A}(\mathbf{M})$ for the whole spacetime
- ▶ to each [suitable] subregion $O \subset M$ a C^* -subalgebra

 $\mathcal{A}(\boldsymbol{M}; \boldsymbol{O}) \subset \mathcal{A}(\boldsymbol{M})$

Interpretation

Elements of $\mathcal{A}(\mathbf{M}; O)$ with $A^* = A$ are observable in O

Prototypical elements are bounded continuous functions of quantum fields averaged within O.

An AQFT on spacetime *M* comprises...



- ▶ a C^* -algebra $\mathcal{A}(\mathbf{M})$ for the whole spacetime
- ▶ to each [suitable] subregion $O \subset M$ a C^* -subalgebra

 $\mathcal{A}(\boldsymbol{M}; \boldsymbol{O}) \subset \mathcal{A}(\boldsymbol{M})$

•
$$O_1 \subset O_2 \implies \mathcal{A}(\boldsymbol{M}; O_1) \subset \mathcal{A}(\boldsymbol{M}; O_2)$$

< ∃ >

ъ

Isotony

An AQFT on spacetime \boldsymbol{M} comprises...



- \blacktriangleright a $\mathit{C^*}\text{-algebra}\ \mathcal{A}(\textit{\textbf{M}})$ for the whole spacetime
- ▶ to each [suitable] subregion $O \subset M$ a C^* -subalgebra

 $\mathcal{A}(\boldsymbol{M}; \boldsymbol{O}) \subset \mathcal{A}(\boldsymbol{M})$

$$\bullet \ O_1 \subset O_2 \implies \mathcal{A}(\mathbf{M}; O_1) \subset \mathcal{A}(\mathbf{M}; O_2)$$
 Isotony

 $\blacktriangleright O_1 \perp O_2 \implies [\mathcal{A}(\boldsymbol{M}; O_1), \mathcal{A}(\boldsymbol{M}; O_2)] = \{\mathbf{0}\} \text{ Einstein causality}$

An AQFT on spacetime \boldsymbol{M} comprises...



- a C^* -algebra $\mathcal{A}(\mathbf{M})$ for the whole spacetime
- ▶ to each [suitable] subregion $O \subset M$ a C^* -subalgebra

 $\mathcal{A}(\boldsymbol{M}; \boldsymbol{O}) \subset \mathcal{A}(\boldsymbol{M})$

- $\bullet \ O_1 \subset O_2 \implies \mathcal{A}(\boldsymbol{M}; O_1) \subset \mathcal{A}(\boldsymbol{M}; O_2)$ Isotony
- $\blacktriangleright O_1 \perp O_2 \implies \left[\mathcal{A}(\textbf{\textit{M}}; O_1), \mathcal{A}(\textbf{\textit{M}}; O_2) \right] = \{ 0 \} \text{ Einstein causality}$
- if O_1 contains a Cauchy surface of O_2 then

 $\mathcal{A}(\boldsymbol{M}; O_1) = \mathcal{A}(\boldsymbol{M}; O_2)$ Timeslice

An AQFT on spacetime *M* comprises...



- \blacktriangleright a $\mathit{C^*}\text{-algebra}\ \mathcal{A}(\textit{\textbf{M}})$ for the whole spacetime
- ▶ to each [suitable] subregion $O \subset M$ a C^* -subalgebra

 $\mathcal{A}(\boldsymbol{M}; \boldsymbol{O}) \subset \mathcal{A}(\boldsymbol{M})$

- $\bullet \ O_1 \subset O_2 \implies \mathcal{A}(\boldsymbol{M}; O_1) \subset \mathcal{A}(\boldsymbol{M}; O_2)$ Isotony
- $\blacktriangleright O_1 \perp O_2 \implies [\mathcal{A}(\boldsymbol{M}; O_1), \mathcal{A}(\boldsymbol{M}; O_2)] = \{\mathbf{0}\} \text{ Einstein causality}$
- if O_1 contains a Cauchy surface of O_2 then

 $\mathcal{A}(\boldsymbol{M}; O_1) = \mathcal{A}(\boldsymbol{M}; O_2)$ Timeslice

▶ a group homomorphism $\alpha : Sym(\mathbf{M}) \rightarrow Aut(\mathcal{A}(\mathbf{M}))$ s.t.

 $\alpha(\tau)(\mathcal{A}(\boldsymbol{M}; \boldsymbol{O})) = \mathcal{A}(\boldsymbol{M}; \tau(\boldsymbol{O}))$

) Symmetry

7 / 23

Consider a representation π of $\mathcal{A}(\mathbf{M})$ on Hilbert space \mathcal{H} (e.g., a GNS rep.).

- For each spacetime region O,
 - $\pi(\mathcal{A}(\boldsymbol{M}; O)) \subset \mathcal{B}(\mathcal{H})$ contains at most the same info as $\mathcal{A}(\boldsymbol{M}; O)$



→

э

M Consider a representation π of $\mathcal{A}(\mathbf{M})$ on Hilbert space \mathcal{H} (e.g., a GNS rep.).

For each spacetime region O,

- $\pi(\mathcal{A}(\mathbf{M}; O)) \subset \mathcal{B}(\mathcal{H})$ contains at most the same info as $\mathcal{A}(\mathbf{M}; O)$
- but the von Neumann algebra

$$\mathcal{M}(O) = \pi(\mathcal{A}(\boldsymbol{M};O))''$$

can contain extra elements incorporating information about the representation.



→

э

Consider a representation π of $\mathcal{A}(\mathbf{M})$ on Hilbert space \mathcal{H} (e.g., a GNS rep.).

For each spacetime region O,

- $\pi(\mathcal{A}(\boldsymbol{M}; \mathcal{O})) \subset \mathcal{B}(\mathcal{H})$ contains at most the same info as $\mathcal{A}(\boldsymbol{M}; \mathcal{O})$
- but the von Neumann algebra

$$\mathcal{M}(O) = \pi(\mathcal{A}(\boldsymbol{M};O))''$$

can contain extra elements incorporating information about the representation.

▶ Representations that are 'not too different' give isomorphic von Neumann algebras

イロト イボト イヨト イヨト



Consider a representation π of $\mathcal{A}(\mathbf{M})$ on Hilbert space \mathcal{H} (e.g., a GNS rep.).

For each spacetime region O,

- $\pi(\mathcal{A}(\boldsymbol{M}; \mathcal{O})) \subset \mathcal{B}(\mathcal{H})$ contains at most the same info as $\mathcal{A}(\boldsymbol{M}; \mathcal{O})$
- but the von Neumann algebra

$$\mathcal{M}(O) = \pi(\mathcal{A}(\boldsymbol{M};O))''$$

can contain extra elements incorporating information about the representation.

- Representations that are 'not too different' give isomorphic von Neumann algebras
- ▶ The vN algebras can encapsulate the QFT under good physical conditions



Classification of von Neumann algebras Murray & von Neumann... ... Connes

 \blacktriangleright A vN algebra $\mathcal M$ is a factor if $\mathcal M \cap \mathcal M' = \mathbb C \mathbf 1$


- \blacktriangleright A vN algebra $\mathcal M$ is a factor if $\mathcal M \cap \mathcal M' = \mathbb C 1$
- ▶ All vN factors fall into three types: I, II, III with subclassification
 - I_n, $n \in \mathbb{N} \cup \{\infty\}$
 - II_1 , II_∞
 - III_{λ} , $\lambda \in [0, 1]$

= nar

- \blacktriangleright A vN algebra $\mathcal M$ is a factor if $\mathcal M \cap \mathcal M' = \mathbb C \mathbf 1$
- ▶ All vN factors fall into three types: I, II, III with subclassification
 - I_n, $n \in \mathbb{N} \cup \{\infty\}$
 - II_1 , II_∞
 - III_{λ} , $\lambda \in [0, 1]$
- Type I_n : $\mathcal{M} \cong \mathcal{B}(\mathcal{H})$ for Hilbert space \mathcal{H} of dimension n
 - for $n < \infty$ this is matrix theory every element has a finite trace
 - ▶ for $n = \infty$ trace class operators are weakly dense in M semifinite trace

Trace = faithful, normal convex map $\tau : \mathcal{M}_+ \to [0, \infty]$, s.t. $\tau(A^*A) = \tau(AA^*)$ for all A.

▶ ▲ 臣 ▶ ▲ 臣 ▶ ○ 臣 ● の Q @

- \blacktriangleright A vN algebra $\mathcal M$ is a factor if $\mathcal M \cap \mathcal M' = \mathbb C \mathbf 1$
- ▶ All vN factors fall into three types: I, II, III with subclassification
 - I_n, $n \in \mathbb{N} \cup \{\infty\}$
 - ▶ II_1 , II_∞
 - III_{λ} , $\lambda \in [0, 1]$
- Type I_n : $\mathcal{M} \cong \mathcal{B}(\mathcal{H})$ for Hilbert space \mathcal{H} of dimension n
 - for $n < \infty$ this is matrix theory every element has a finite trace
 - for $n = \infty$ trace class operators are weakly dense in \mathcal{M} semifinite trace
- ▶ Type II₁ factors have a finite trace (but are not type I_n)

 $\mathsf{Trace} = \mathsf{faithful}, \ \mathsf{normal} \ \mathsf{convex} \ \mathsf{map} \ \tau : \mathcal{M}_+ \to [0,\infty], \ \mathsf{s.t.} \ \tau(\mathcal{A}^*\mathcal{A}) = \tau(\mathcal{A}\mathcal{A}^*) \ \mathsf{for} \ \mathsf{all} \ \mathcal{A}.$

- \blacktriangleright A vN algebra $\mathcal M$ is a factor if $\mathcal M \cap \mathcal M' = \mathbb C \mathbf 1$
- ▶ All vN factors fall into three types: I, II, III with subclassification
 - I_n, $n \in \mathbb{N} \cup \{\infty\}$
 - ▶ II_1 , II_∞
 - III_{λ} , $\lambda \in [0, 1]$
- Type I_n : $\mathcal{M} \cong \mathcal{B}(\mathcal{H})$ for Hilbert space \mathcal{H} of dimension n
 - for $n < \infty$ this is matrix theory every element has a finite trace
 - for $n = \infty$ trace class operators are weakly dense in \mathcal{M} semifinite trace
- Type II₁ factors have a finite trace (but are not type I_n)
- \blacktriangleright Type II_∞ factors have a semifinite trace (but are not type $I_\infty)$

Trace = faithful, normal convex map $\tau : \mathcal{M}_+ \to [0, \infty]$, s.t. $\tau(A^*A) = \tau(AA^*)$ for all A.

- \blacktriangleright A vN algebra $\mathcal M$ is a factor if $\mathcal M\cap \mathcal M'=\mathbb C 1$
- ▶ All vN factors fall into three types: I, II, III with subclassification
 - I_n, $n \in \mathbb{N} \cup \{\infty\}$
 - ▶ II_1 , II_∞
 - III_{λ} , $\lambda \in [0, 1]$
- Type I_n : $\mathcal{M} \cong \mathcal{B}(\mathcal{H})$ for Hilbert space \mathcal{H} of dimension n
 - for $n < \infty$ this is matrix theory every element has a finite trace
 - for $n = \infty$ trace class operators are weakly dense in \mathcal{M} semifinite trace
- Type II₁ factors have a finite trace (but are not type I_n)
- \blacktriangleright Type II_∞ factors have a semifinite trace (but are not type $I_\infty)$
- Type III factors do not admit any (semi)finite trace.

 τ(*A*) ∈ {0, ∞} for every element *A* ∈ *M*₊.

 $\mathsf{Trace} = \mathsf{faithful}, \ \mathsf{normal} \ \mathsf{convex} \ \mathsf{map} \ \tau : \mathcal{M}_+ \to [0,\infty], \ \mathsf{s.t.} \ \tau(\mathcal{A}^*\mathcal{A}) = \tau(\mathcal{A}\mathcal{A}^*) \ \mathsf{for} \ \mathsf{all} \ \mathcal{A}.$



▲日 > ▲母 > ▲目 > ▲目 > ▲目 > ④ < @

CJ Fewster QFT, QRF and types

10/23



ヨト・モラト

= nar

Intro AQET Types of vN factors Type change Summary Classification Type II Universal type Recap Traces and entropy
A Type III example Two spin chains, Hilbert space
$$\mathcal{H} = \bigotimes_{r=1}^{\infty} (\mathbb{C}^2 \otimes \mathbb{C}^2)$$
 (suitably defined)
 $\Psi = \frac{|00\rangle + |11\rangle}{\sqrt{2}} \otimes \cdots$
Generate \mathcal{M} from red chain observables $X = \bigotimes_{r=1}^{\infty} (A_r \otimes \mathbf{1})$ with $A_r = \mathbf{1}$ for suff. large r .
Then $X \mapsto \langle \Psi | X \Psi \rangle = \lim_{R \to \infty} \frac{1}{2^R} \prod_{r=1}^R \operatorname{Tr} A_r$ extends to a finite trace on \mathcal{M}
so \mathcal{M} is of Type II₁.

The universal type of local algebras in QFT

Theorem Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987 For a W^* -AQFT $O \mapsto \mathcal{M}(O)$ obeying reasonable conditions (nuclearity, scaling limits)

each $\mathcal{M}(O)$ is isomorphic to the unique hyperfinite factor of type III₁.

Haagerup, 1987



The universal type of local algebras in QFT

Theorem Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987 For a W^* -AQFT $O \mapsto \mathcal{M}(O)$ obeying reasonable conditions (nuclearity, scaling limits) each $\mathcal{M}(O)$ is isomorphic to the unique hyperfinite factor of type III₁.

Many consequences – distinguishing QFT from QM – including

Distinction between theories is due to 'relative position' of $\mathcal{M}(\mathcal{O})$'s in $\mathcal{B}(\mathcal{H})$.

I na∩

Theorem Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987 For a W^* -AQFT $O \mapsto \mathcal{M}(O)$ obeying reasonable conditions (nuclearity, scaling limits) each $\mathcal{M}(O)$ is isomorphic to the unique hyperfinite factor of type III₁.

Many consequences – distinguishing QFT from QM – including

Distinction between theories is due to 'relative position' of $\mathcal{M}(\mathcal{O})$'s in $\mathcal{B}(\mathcal{H})$.

Infinite degeneracy No $\mathcal{M}(O)$ contains a nonzero finite rank projection

The universal type of local algebras in QFT

Theorem Fredenhagen 1985; Buchholz, D'Antoni, Fredenhagen 1987 For a W^* -AQFT $O \mapsto \mathcal{M}(O)$ obeying reasonable conditions (nuclearity, scaling limits) each $\mathcal{M}(O)$ is isomorphic to the unique hyperfinite factor of type III₁.

Many consequences – distinguishing QFT from QM – including

Distinction between theories is due to 'relative position' of $\mathcal{M}(\mathcal{O})$'s in $\mathcal{B}(\mathcal{H})$.

Infinite degeneracy No $\mathcal{M}(O)$ contains a nonzero finite rank projection

Local states are impure No normal state restricts to any $\mathcal{M}(O)$ as a pure state.

See Yngvason The role of Type III factors in quantum field theory for more.

= nar

Recap

Different vN algebra types correspond roughly to different physical settings:

- ▶ Type I_n (qubits)
- Type I_{∞} (quantum mechanics)
- ▶ Type II₁ (some quantum stat. mech. models)
- Type II_{∞} (ditto)
- ▶ Type III₁ (quantum field theory) Buchholz, D'Antoni, Fredenhagen
- Type ??? (quantum gravity)

I na∩

Recap

Different vN algebra types correspond roughly to different physical settings:

- Type I_n (qubits) \leftarrow finite trace
- ▶ Type I_{∞} (quantum mechanics) ← semifinite trace
- ▶ Type II₁ (some quantum stat. mech. models) \leftarrow finite trace
- Type II_{∞} (ditto) \leftarrow semifinite trace
- ▶ Type III₁ (quantum field theory) \leftarrow no (semi)finite trace
- Type ??? (quantum gravity)

= nar

What the AI thinks... Image credits: Google Gemini



くして 前 ふかく ボット 日本 くりく

Lost without a trace?

A finite trace τ makes \mathcal{M} into a pre-Hilbert space w.r.t. $(A, B) \mapsto \tau(A^*B)$.



Lost without a trace?

A finite trace τ makes \mathcal{M} into a pre-Hilbert space w.r.t. $(A, B) \mapsto \tau(A^*B)$.

Any vector state $\psi \in \mathcal{H}$ can be written

 $\langle \psi | \pi(A)\psi \rangle = \tau(\rho\pi(A)),$ for unique 'density matrix' $\rho \in \mathcal{M}$ (Riesz rep. + a bit more) with entropy $S = -\tau(\rho \log \rho).$

Lost without a trace?

A finite trace τ makes \mathcal{M} into a pre-Hilbert space w.r.t. $(A, B) \mapsto \tau(A^*B)$.

Any vector state $\psi \in \mathcal{H}$ can be written

 $\langle \psi | \pi(A)\psi \rangle = \tau(\rho\pi(A)),$ for unique 'density matrix' $\rho \in \mathcal{M}$ (Riesz rep. + a bit more) with entropy $S = -\tau(\rho \log \rho).$

> Evidently this is impossible for $\mathcal{M}(O)$ in AQFT. "I really apologise for that." – Klaus Fredenhagen

= nar

QFT vs Gravity

The description of (A)QFT is focussed on algebras associated with spacetime subsets.

QFT lives on manifolds



з.

QFT vs Gravity

The description of (A)QFT is focussed on algebras associated with spacetime subsets.

QFT lives on manifolds

However, Einstein's hole argument shows that

General Relativity is pointless

due to diffeomorphism invariance - it lives on manifolds modulo diffeos.

(4) E (4) E (4) E (4)

QFT vs Gravity

The description of (A)QFT is focussed on algebras associated with spacetime subsets.

QFT lives on manifolds

However, Einstein's hole argument shows that

General Relativity is pointless

due to diffeomorphism invariance - it lives on manifolds modulo diffeos.

What is the description of quantum observables allowing for dynamical gravity?

 \longrightarrow Investigate in de Sitter spacetime.

・ 同下 ・ ヨト ・ ヨト

QFT in de Sitter spacetime

n-dimensional de Sitter spacetime is the hyperboloid

$$(X^0)^2 - \sum_{r=1}^n (X^r)^2 = -a^2$$

in (n + 1)-dim Minkowski, with the induced metric.



Image: A match a ma

< ∃⇒

э

QFT in de Sitter spacetime

n-dimensional de Sitter spacetime is the hyperboloid

$$(X^0)^2 - \sum_{r=1}^n (X^r)^2 = -a^2$$

in (n + 1)-dim Minkowski, with the induced metric.

Any timelike geodesic has an associated static patch consisting of all points with which an observer on the curve can communicate.



Image: A math a math

QFT in de Sitter spacetime

n-dimensional de Sitter spacetime is the hyperboloid

$$(X^0)^2 - \sum_{r=1}^n (X^r)^2 = -a^2$$

in (n + 1)-dim Minkowski, with the induced metric.

Any timelike geodesic has an associated static patch consisting of all points with which an observer on the curve can communicate.

The boundary of the static patch is the observer's cosmological horizon.

Spacetime isometry subgroup $\mathbb{R}\times \mathrm{SO}(3)$ preserves the patch.







Consider QFT representing 'weak quantum gravity' around a de Sitter background, in a de Sitter invariant vacuum state. Consider a timelike geodesic and its static patch.

▶ The static patch observable algebra has Type III₁





- ▶ The static patch observable algebra has Type III₁
- As the background is fixed the diffeo ambiguity is reduced to ℝ × SO(3).



- ▶ The static patch observable algebra has Type III₁
- As the background is fixed the diffeo ambiguity is reduced to $\mathbb{R} \times SO(3)$.
- Idea Introduce a clock system for the observer, covariant w.r.t. the time translations.



- The static patch observable algebra has Type III₁
- As the background is fixed the diffeo ambiguity is reduced to $\mathbb{R} \times \mathrm{SO}(3)$.
- Idea Introduce a clock system for the observer, covariant w.r.t. the time translations.
- Treat the total QFT+clock Hamiltonian as a constraint.



Consider QFT representing 'weak quantum gravity' around a de Sitter background, in a de Sitter invariant vacuum state. Consider a timelike geodesic and its static patch.

- The static patch observable algebra has Type III₁
- As the background is fixed the diffeo ambiguity is reduced to $\mathbb{R} \times \mathrm{SO}(3)$.
- Idea Introduce a clock system for the observer, covariant w.r.t. the time translations.
- Treat the total QFT+clock Hamiltonian as a constraint.
- The physical observables are those joint clock–QFT observables that are invariant under the time translations.



A B A B A
 A
 B
 A
 A
 B
 A
 A
 B
 A
 A
 B
 A
 A
 B
 A
 A
 B
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A
 A

Consider QFT representing 'weak quantum gravity' around a de Sitter background, in a de Sitter invariant vacuum state. Consider a timelike geodesic and its static patch.

- ▶ The static patch observable algebra has Type III₁
- As the background is fixed the diffeo ambiguity is reduced to $\mathbb{R} \times \mathrm{SO}(3)$.
- Idea Introduce a clock system for the observer, covariant w.r.t. the time translations.
- Treat the total QFT+clock Hamiltonian as a constraint.
- The physical observables are those joint clock–QFT observables that are invariant under the time translations.
- The resulting algebra has type II₁.

Uses Takesaki's theory of vN crossed products.



▲□▶▲御▶★臣▶★臣▶ 臣 のの

The CLPW model (continued)

The II₁ factor has a finite trace (unique up to scale, fixed by $\tau(\mathbf{1}) = \mathbf{1}$). The dS invariant Euclidean vacuum state corresponds to density matrix $\rho_{dS} = \mathbf{1}$, so

$$\mathcal{S}(
ho_{\mathsf{dS}}) = \mathsf{0},$$
 and in fact $\mathcal{S}(
ho) \leqslant \mathsf{0}$ for general states

з.

The CLPW model (continued)

The II₁ factor has a finite trace (unique up to scale, fixed by $\tau(\mathbf{1}) = \mathbf{1}$). The dS invariant Euclidean vacuum state corresponds to density matrix $\rho_{dS} = \mathbf{1}$, so

$$S(
ho_{\mathsf{dS}}) = 0,$$
 and in fact $S(
ho) \leqslant 0$ for general states

Compare with the expectations from gravity

$$S_{
m tot} \leqslant S_{
m Gibbons-Hawking} = rac{A}{4G}.$$

NB the trace is only fixed up to scale, so S is only fixed up to a constant. In this sense, agreement is obtained.

э.

The CLPW model (continued)

The II₁ factor has a finite trace (unique up to scale, fixed by $\tau(\mathbf{1}) = \mathbf{1}$). The dS invariant Euclidean vacuum state corresponds to density matrix $\rho_{dS} = \mathbf{1}$, so

$$S(
ho_{\mathsf{dS}}) = 0,$$
 and in fact $S(
ho) \leqslant 0$ for general states

Compare with the expectations from gravity

$$S_{
m tot} \leqslant S_{
m Gibbons-Hawking} = rac{A}{4G}.$$

NB the trace is only fixed up to scale, so S is only fixed up to a constant. In this sense, agreement is obtained.

Q How fine-tuned is all this? Other spacetimes? Other symmetries? Why that clock?

э.

(Quantum) reference frames

We cannot measure absolute positions or times - hardwired into microphysical laws.

No absolute position



ъ

A B >
 A B >
 A

(Quantum) reference frames

We cannot measure absolute positions or times – hardwired into microphysical laws. Instead, we measure relative to the position of objects or events – physical systems constituting a reference frame.


(Quantum) reference frames

We cannot measure absolute positions or times – hardwired into microphysical laws. Instead, we measure relative to the position of objects or events – physical systems constituting a reference frame.

Natural to assume that these systems are governed by quantum theory – quantum reference frames. Idea goes back to Aharonov and beyond that to Eddington.

Reference system



(Quantum) reference frames

We cannot measure absolute positions or times – hardwired into microphysical laws. Instead, we measure relative to the position of objects or events – physical systems constituting a reference frame.

Natural to assume that these systems are governed by quantum theory – quantum reference frames. Idea goes back to Aharonov and beyond that to Eddington.

There are several recent formalisations

- operational [our choice] Busch, Loveridge, Miyadera...
- perspective neutral Höhn, Giacomini, Bruckner...
- information-based Bartlett, Spekkens...

CJ Fewster QFT, QRF and types

Given a quantum system \mathcal{H}_S with symmetry group *G* represented unitarily by U_S , a QRF is another quantum system \mathcal{H}_R , with *G* represented unitarily by U_R , and a *G*-covariant observable *P*, technically, a POVM on value space Σ acted on by *G*.



= nar

Given a quantum system \mathcal{H}_S with symmetry group G represented unitarily by U_S , a QRF is another quantum system \mathcal{H}_R , with G represented unitarily by U_R , and a G-covariant observable P, technically, a POVM on value space Σ acted on by G.

• $\langle P(\Delta) \rangle_{\psi}$ = probability that the QRF 'reads' a value in $\Delta \subset \Sigma$ in state $\psi \in \mathcal{H}_R$.

= nar

Given a quantum system \mathcal{H}_S with symmetry group G represented unitarily by U_S , a QRF is another quantum system \mathcal{H}_R , with G represented unitarily by U_R , and a G-covariant observable P, technically, a POVM on value space Σ acted on by G.

- $\langle P(\Delta) \rangle_{\psi}$ = probability that the QRF 'reads' a value in $\Delta \subset \Sigma$ in state $\psi \in \mathcal{H}_R$.
- G-covariance means $U_R(g)P(\Delta)U_R(g)^{-1} = P(g.\Delta)$

Given a quantum system \mathcal{H}_S with symmetry group *G* represented unitarily by U_S , a QRF is another quantum system \mathcal{H}_R , with *G* represented unitarily by U_R , and a *G*-covariant observable *P*, technically, a POVM on value space Σ acted on by *G*.

- $\langle P(\Delta) \rangle_{\psi}$ = probability that the QRF 'reads' a value in $\Delta \subset \Sigma$ in state $\psi \in \mathcal{H}_R$.
- *G*-covariance means $U_R(g)P(\Delta)U_R(g)^{-1} = P(g.\Delta)$
- The physical observables are the joint system–QRF observables that are invariant under the *G*-action on $\mathcal{H}_S \otimes \mathcal{H}_R$.

E.g., relative position of two quantum particles on the line.

Given a quantum system \mathcal{H}_S with symmetry group G represented unitarily by U_S , a QRF is another quantum system \mathcal{H}_R , with G represented unitarily by U_R , and a G-covariant observable P, technically, a POVM on value space Σ acted on by G.

Observation The CLPW clock defines a QRF for time translations $G = \mathbb{R}$, $\Sigma = \mathbb{R}$ $\mathcal{H}_R = L^2(\mathbb{R}^+, dE)$, $H_R = E$, $U_R(t) = e^{iH_R t}$,

$$P(\Delta) = W^* \chi_{\Delta} W, \qquad (W\psi)(t) = rac{1}{\sqrt{2\pi}} \int_{\mathbb{R}^+} dE \, e^{-iEt} \psi(E),$$

for $\Delta \subset \mathbb{R}$, and one has $U_R(t)P(\Delta)U_R(t)^* = P(\Delta + t)$, i.e., P is a time observable.

Given a quantum system \mathcal{H}_S with symmetry group G represented unitarily by U_S , a QRF is another quantum system \mathcal{H}_R , with G represented unitarily by U_R , and a G-covariant observable P, technically, a POVM on value space Σ acted on by G.

Observation The CLPW clock defines a QRF for time translations $G = \mathbb{R}$, $\Sigma = \mathbb{R}$ $\mathcal{H}_R = L^2(\mathbb{R}^+, dE)$, $H_R = E$, $U_R(t) = e^{iH_R t}$,

$${\cal P}(\Delta) = W^* \chi_{\Delta} W, \qquad (W\psi)(t) = rac{1}{\sqrt{2\pi}} \int_{\mathbb{R}^+} dE \, e^{-iEt} \psi(E),$$

for $\Delta \subset \mathbb{R}$, and one has $U_R(t)P(\Delta)U_R(t)^* = P(\Delta + t)$, i.e., P is a time observable.

P is a POVM, but not a PVM.

Consistent with Pauli's theorem on nonexistence of time observables if H is semibounded.

The CLPW observer isn't coupled to the system – what (if anything) is it observing? Our approach: start from a description of measurement theory in QFT CJF & Verch, CMP 2020



з.

The CLPW observer isn't coupled to the system – what (if anything) is it observing? Our approach: start from a description of measurement theory in QFT CJF & Verch, CMP 2020

The QFT is measured by locally coupling to a probe theory

= nar

The CLPW observer isn't coupled to the system – what (if anything) is it observing? Our approach: start from a description of measurement theory in QFT CJF & Verch, CMP 2020

- The QFT is measured by locally coupling to a probe theory
- Isometrically related couplings should be indistinguishable

The CLPW observer isn't coupled to the system – what (if anything) is it observing? Our approach: start from a description of measurement theory in QFT CJF & Verch, CMP 2020

- The QFT is measured by locally coupling to a probe theory
- Isometrically related couplings should be indistinguishable
- Ambiguity is resolved by introducing a QRF Interpretation: the QRF triggers the measurement

The CLPW observer isn't coupled to the system – what (if anything) is it observing? Our approach: start from a description of measurement theory in QFT CJF & Verch. CMP 2020

- The QFT is measured by locally coupling to a probe theory
- Isometrically related couplings should be indistinguishable
- Ambiguity is resolved by introducing a QRF Interpretation: the QRF triggers the measurement
- Invariant elements of the joint QFT-QRF algebra are physical observables.

The CLPW observer isn't coupled to the system – what (if anything) is it observing? Our approach: start from a description of measurement theory in QFT CJF & Verch. CMP 2020

- The QFT is measured by locally coupling to a probe theory
- Isometrically related couplings should be indistinguishable
- Ambiguity is resolved by introducing a QRF Interpretation: the QRF triggers the measurement
- Invariant elements of the joint QFT-QRF algebra are physical observables.
- Investigate in generality what properties are required for change of algebra type.

Techniques include Mackey theory, and dilation of nonsharp QRFs to sharp ones.

I na∩

Consider situations with a spacetime isometry group $G = \mathbb{R} \times H$ with compact H so that the \mathbb{R} action on the (type III) QFT vN algebra is the modular flow for a β -KMS state (includes CLPW).

Consider situations with a spacetime isometry group $G = \mathbb{R} \times H$ with compact H so that the \mathbb{R} action on the (type III) QFT vN algebra is the modular flow for a β -KMS state (includes CLPW).

Theorem The algebra of QFT–QRF invariants A_{inv} is always semifinite (not type III).

Consider situations with a spacetime isometry group $G = \mathbb{R} \times H$ with compact H so that the \mathbb{R} action on the (type III) QFT vN algebra is the modular flow for a β -KMS state (includes CLPW).

Theorem The algebra of QFT–QRF invariants A_{inv} is always semifinite (not type III). Furthermore, A_{inv} is finite if the QRF admits a particular well-behaved β -KMS weight

(effectively constrains the growth of # QRF d.o.f. with energy).

I na∩

Consider situations with a spacetime isometry group $G = \mathbb{R} \times H$ with compact H so that the \mathbb{R} action on the (type III) QFT vN algebra is the modular flow for a β -KMS state (includes CLPW).

Theorem The algebra of QFT–QRF invariants A_{inv} is always semifinite (not type III). Furthermore, A_{inv} is finite if the QRF admits a particular well-behaved β -KMS weight (effectively constrains the growth of # QRF d.o.f. with energy).

Summary: CLPW is generalised within a broader, operationally motivated, setting, elucidating the roles of good thermodynamic behaviour for the QRF and gravity.

For more detail, see Thursday's talk.

Summary

- Explanation of the role of algebras in QFT
- Explanation of the definition and significance of vN algebra type
- Universal type of local algebras in QFT
- The CLPW model and change of type
- CLPW generalised and more operationally understood.



 $\begin{array}{l} Hyperfinite \ type \ III_1 \ factor \\ according \ to \ Google \ Gemini \end{array}$









QFT, QRF and types

23 / 23