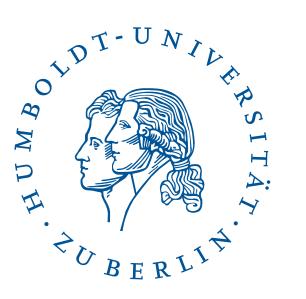
Introduction to Data Analysis

R D Parsons, (+ Orel Gueta, Jakob Nordin)

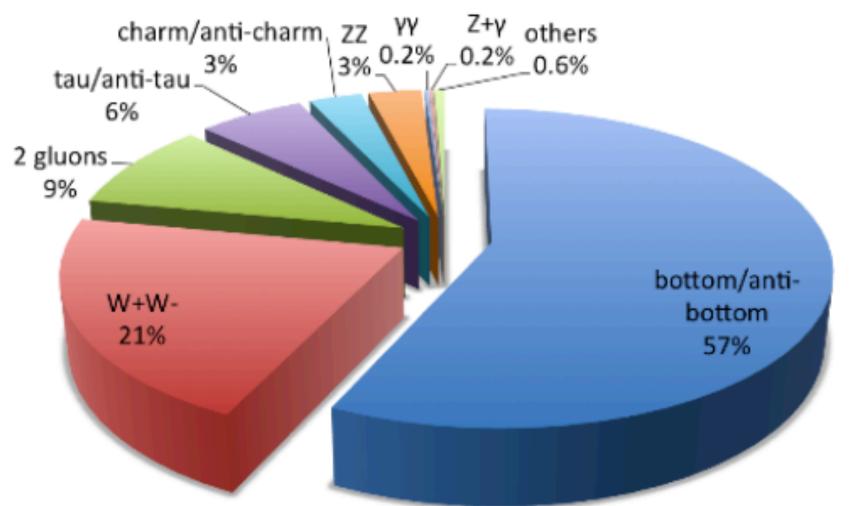


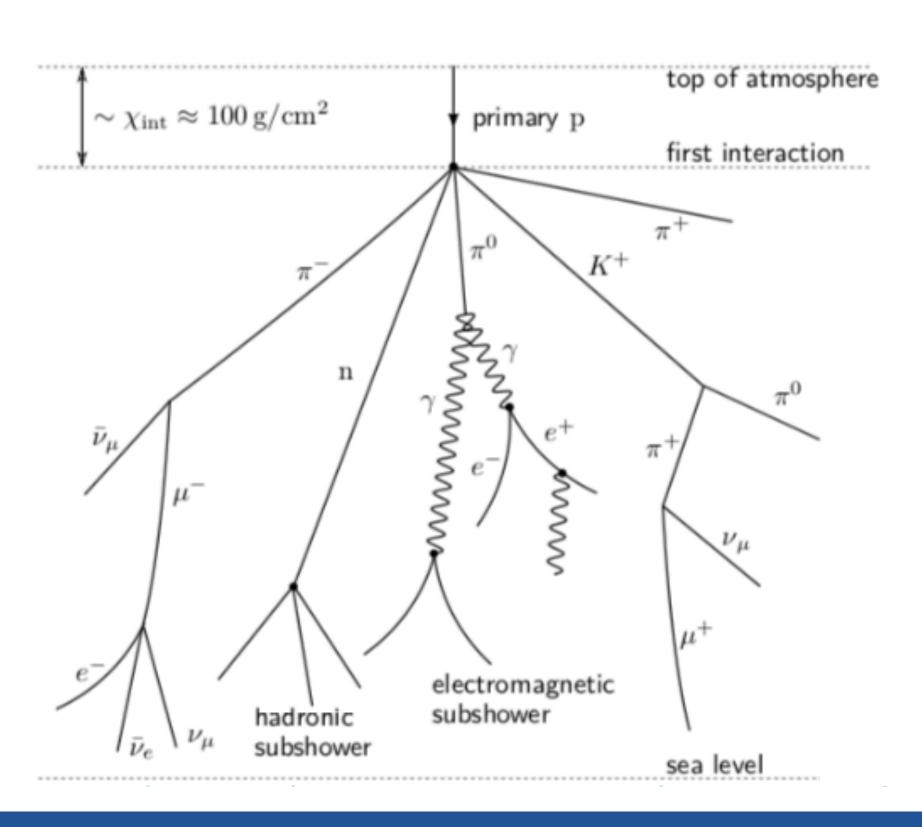
Introduction

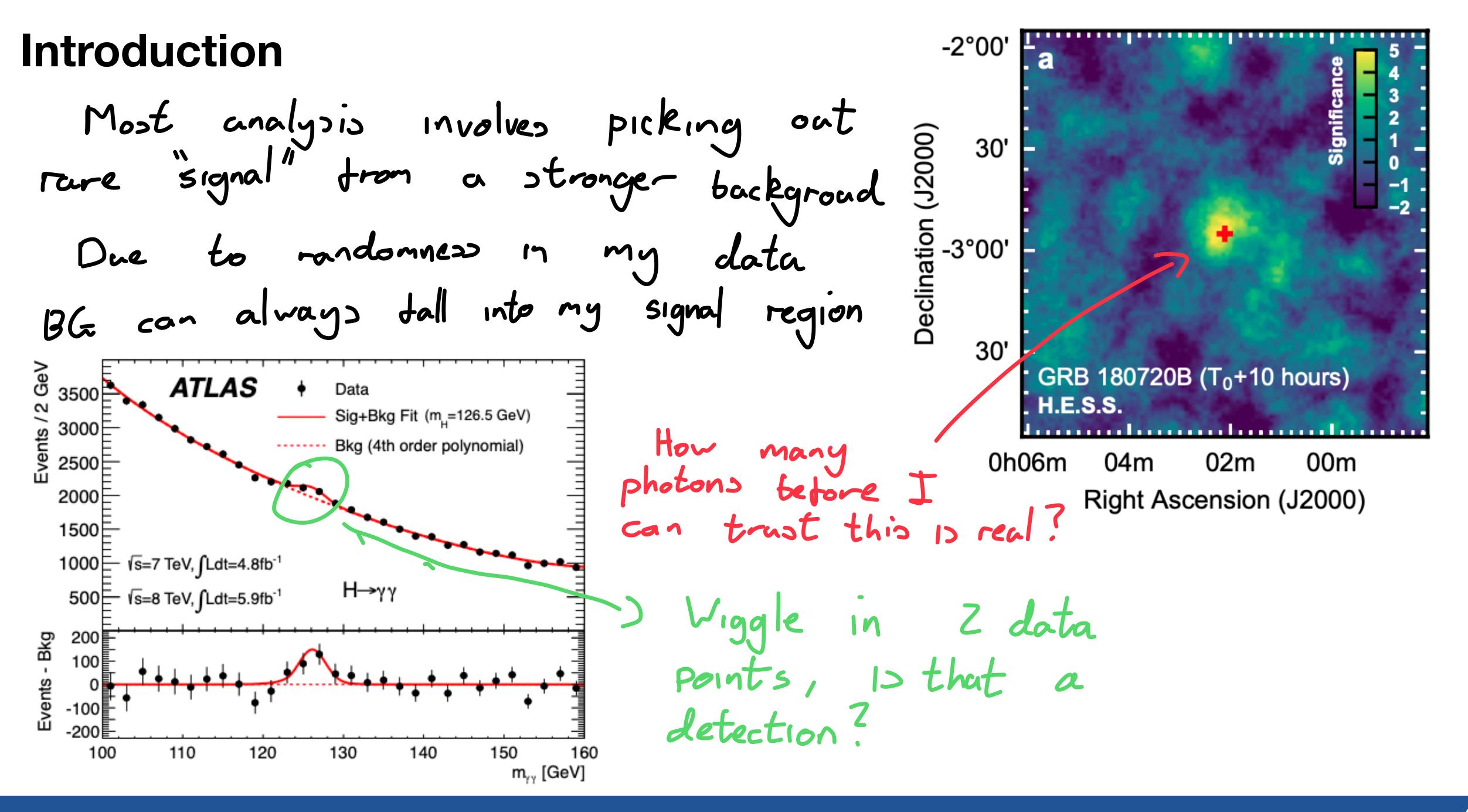
When analysing data randomness appears everywhere Quantum mechanics, particle interactions etc We only know probabilities of occurance

Need to use Statistic> to analyse large volumes of data

Decays of a 125 GeV Standard-Model Higgs boson







A restaurant owner orders 30 rolls every day.

The law in the country states that rolls must weigh **75 grams**

After changing suppliers, the owner suspects that the new baker sells underweight rolls

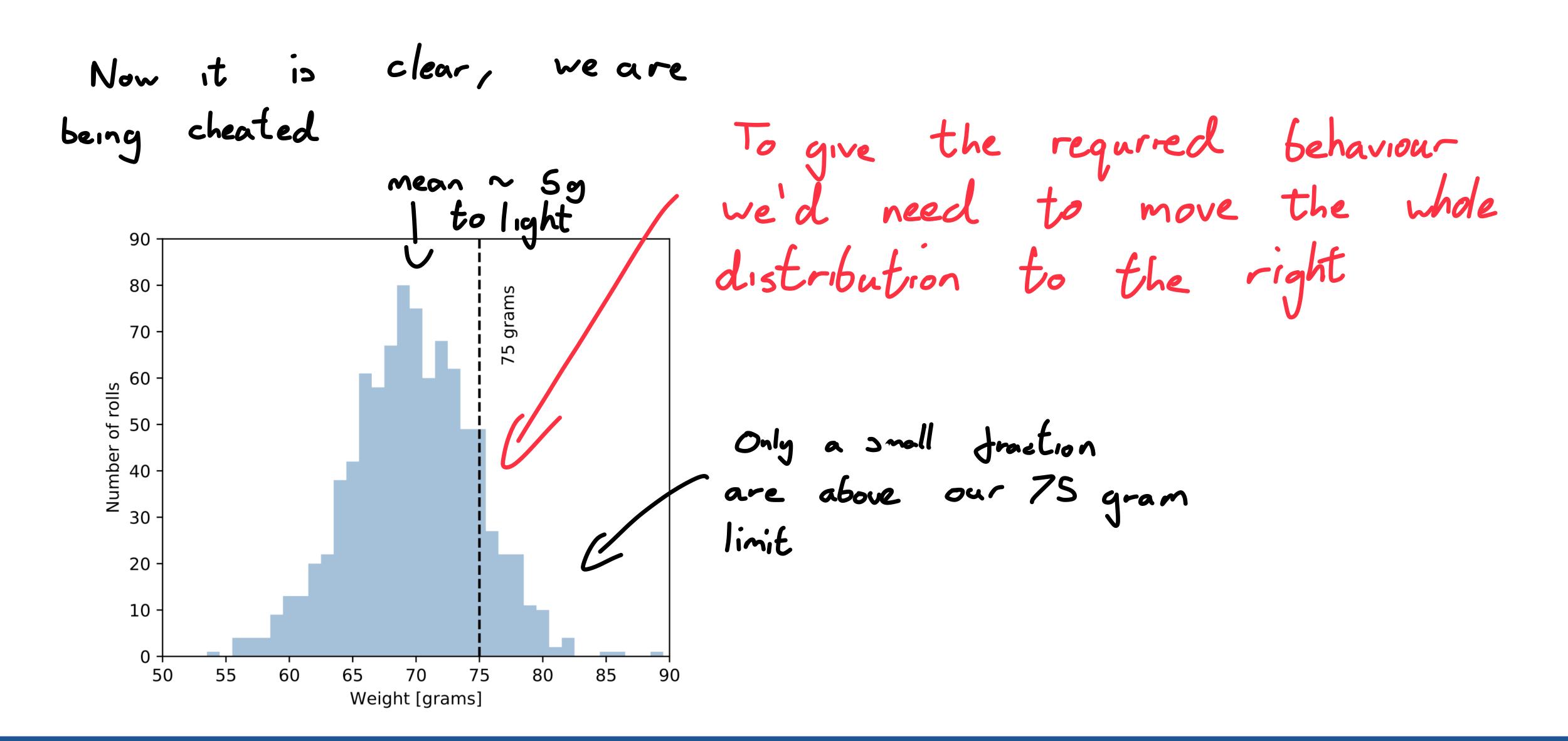
⇒ Investigate! Weigh the rolls (1 gram resolution).

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78, 66, 67, 64, 74, 58, 78, 66, 71, 68, 77, 59, 68, 68, 75, 64, 69, 65, 70, 72
64, 75, 74, 72, 74, 66, 69, 65, 68, 72, 66, 68, 66, 65, 66, 69, 64, 71, 78, 73
69, 65, 66, 78, 70, 66, 70, 80, 70, 73, 71, 68, 64, 68, 68, 72, 74, 74, 71, 74
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70, 62, 73, 68, 70, 76, 71, 71, 71, 66, 74, 77, 73, 74, 65, 65, 62, 76, 68, 76
66, 67, 70, 74, 70, 71, 70, 70, 64, 70, 69, 69, 72, 66, 69, 68, 72, 73, 65, 72
```

Try combining the data by binning with a resolution

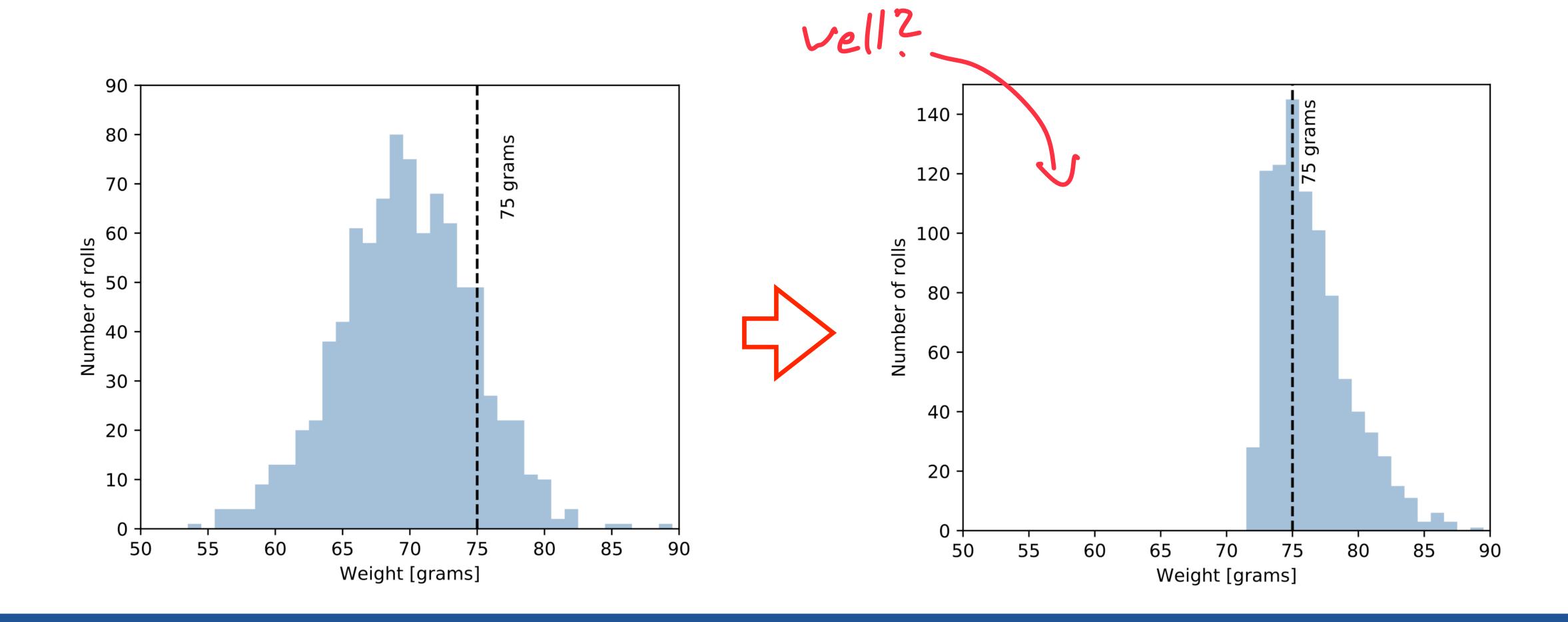
```
A bit more helpful, but difficult to read
```

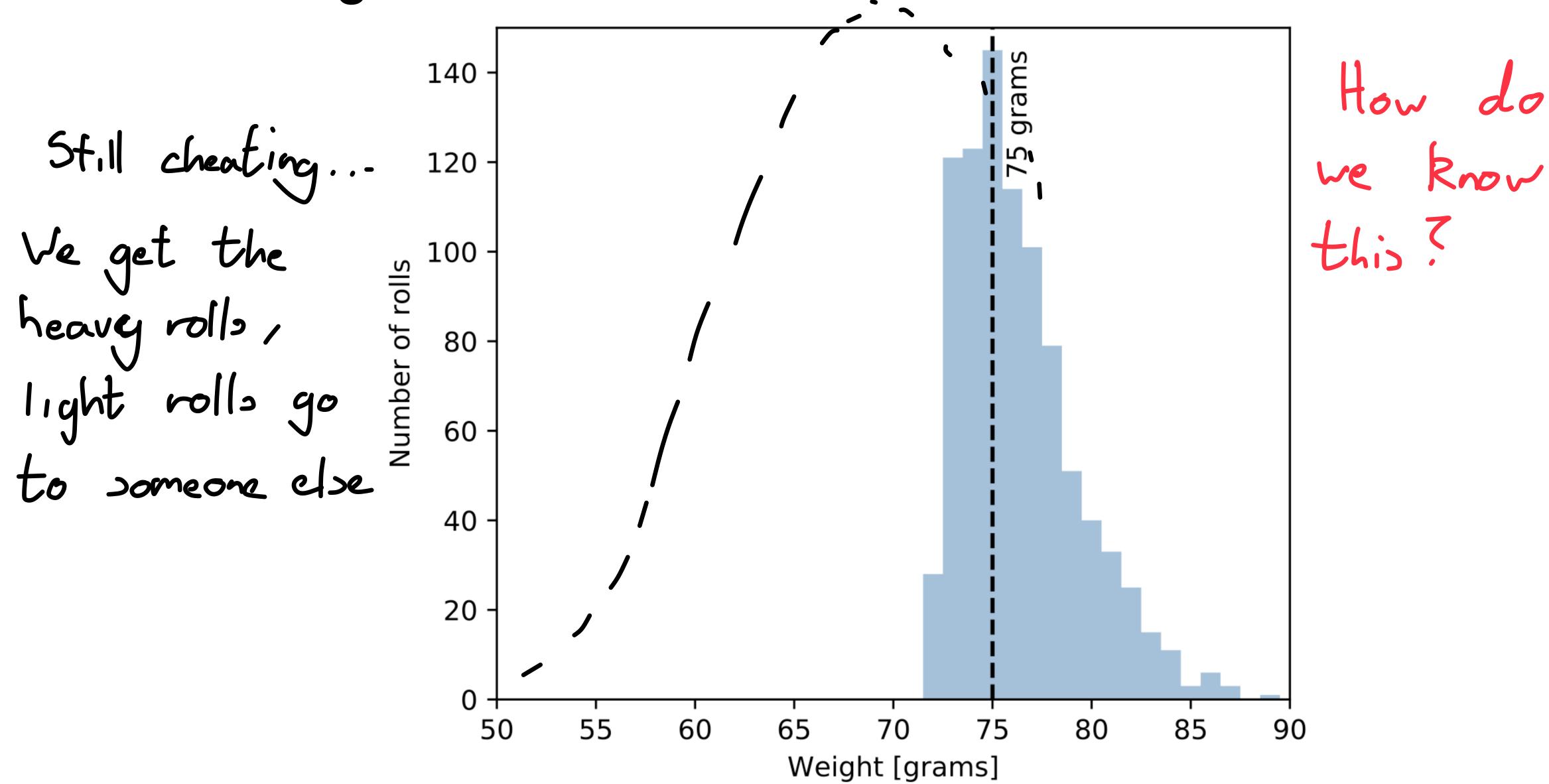
```
Weight[50] = 0 Weight[51] = 0 Weight[52] = 0 Weight[53] = 0 Weight[54] = 1
Weight[55] = 0 Weight[56] = 4 Weight[57] = 4 Weight[58] = 4 Weight[59] = 9
Weight[60] = 13 Weight[61] = 13 Weight[62] = 20 Weight[63] = 22 Weight[64] = 38
Weight[65] = 42 Weight[66] = 61 Weight[67] = 58 Weight[68] = 67 Weight[69] = 80
Weight[70] = 75 Weight[71] = 60 Weight[72] = 68 Weight[73] = 62 Weight[74] = 49
Weight[75] = 49 Weight[76] = 27 Weight[77] = 22 Weight[78] = 22 Weight[79] = 11
Weight[80] = 10 Weight[81] = 2 Weight[82] = 4 Weight[83] = 0 Weight[84] = 0
Weight[85] = 1 Weight[86] = 1 Weight[87] = 0 Weight[88] = 0 Weight[89] = 1
```



He promises to change his ways

Measure again a few weeks later





Probability Distributions

Terminology

An event or random variate is a possible outcome of an experiment governed by a stochastic process.

A population is the set of all possible events. An observation is a realisation of a particular event.

An estimate or a measurement is an attempt to infer properties of the population.

Errors are usually associated with measurements.

A probability distribution function assigns probabilities to events. Can be discrete or continuous.

Cumulative Distribution Function (CDF)

CPF gives us the probability to find a value of x lower than

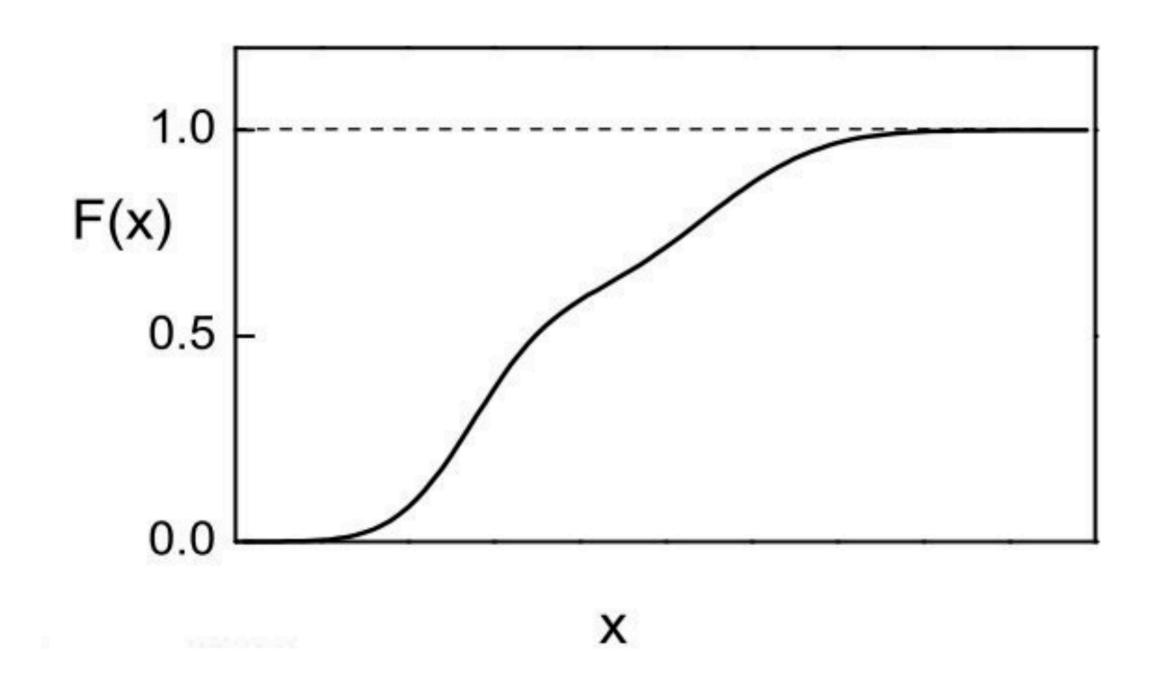
$$F(t) = P\{x < t\}$$
, with $-\infty < t < \infty$.

Axiomatically

F(t) is a non-decreasing function of t,

$$F(-\infty)=0\,,$$

$$F(\infty)=1$$
.



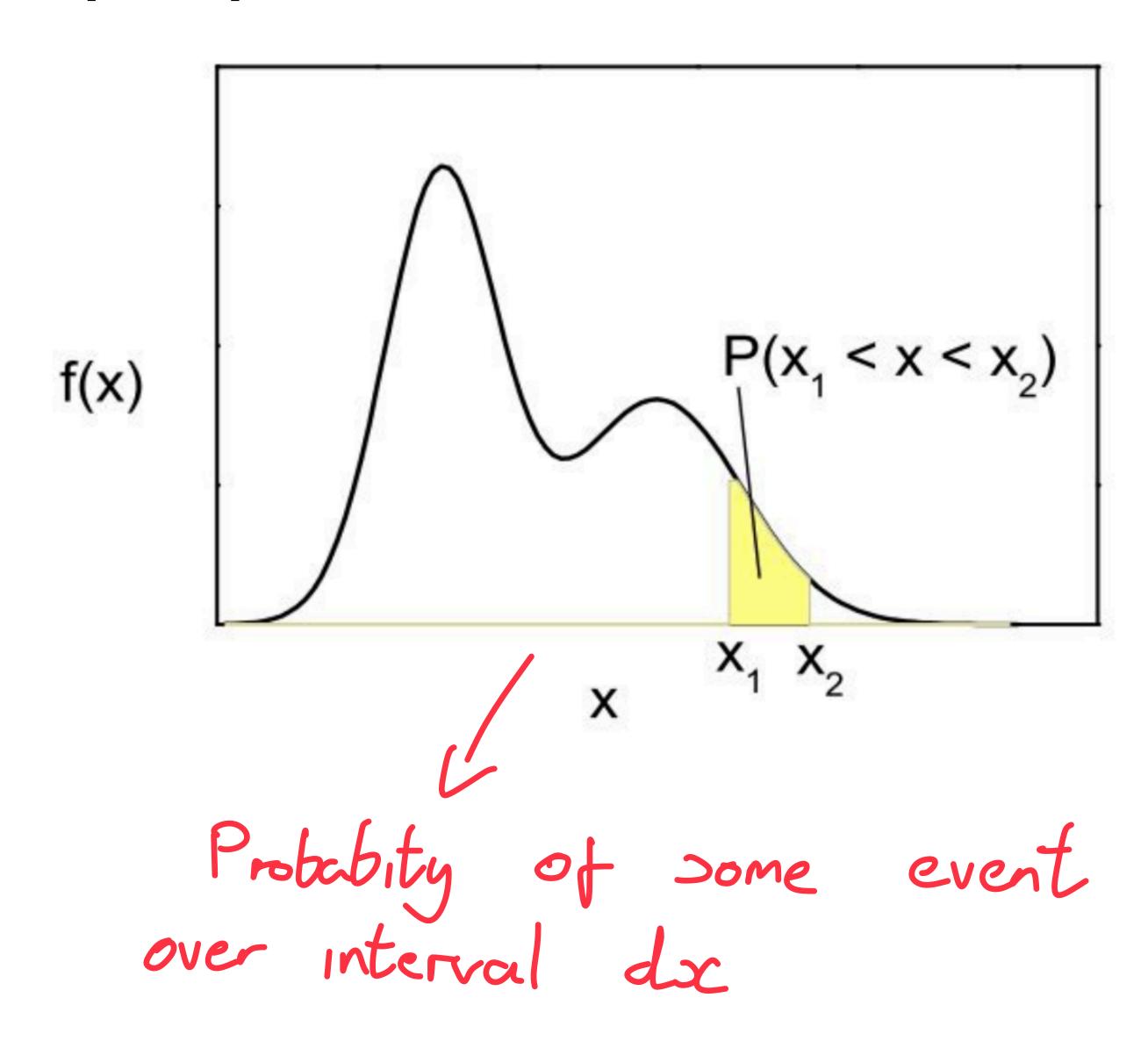
Probability Distribution Function (PDF)

$$f_X(x) = \frac{d}{dx} F_X(x)$$

$$f(-\infty) = f(\infty) = 0$$

$$\int_{-\infty}^{\infty} f(x) dx = 1$$

$$\Pr[a \leq X \leq b] = \int_a^b f_X(x) \, dx.$$



Expectation Value

$$E(u(x)) = \sum_{i=1}^{\infty} u(x_i) p(x_i) \text{ (discrete distribution) },$$

$$E(u(x)) = \int_{-\infty}^{\infty} u(x) f(x) dx \text{ (continuous distribution) }.$$

The expectation value of some quantity u(x) is denoted E(x)

Where \mathbf{x} is randomly distributed according to $\mathbf{f}(\mathbf{x})$,

Can be obtained by taking the average of an infinite number of samples of u(x)

Mean Value

$$E(x) \equiv \langle x \rangle = \mu = \sum_{i=1}^{\infty} x_i p(x_i)$$
 (discrete distribution),
 $E(x) \equiv \langle x \rangle = \mu = \int_{-\infty}^{\infty} x f(x) dx$ (continuous distribution).

The (population) mean, μ =<x>, is the expected value of the variate itself

The sample mean or average based on a finite sample is usually defined as:

$$\overline{x} = \frac{1}{N} \sum_{i} x_{i} .$$

The sample mean is a random variable, or function of the random variate \mathbf{x} , which has the expectation value:

$$\langle \overline{x} \rangle = \frac{1}{N} \sum_{i} \langle x_i \rangle = \langle x \rangle$$

i.e. the mean μ is a property of the unknown, true p.d.f. while the sample mean is a measurement constructed from events or observations. In the limit of large N we expect the sample mean to approach the mean. We can thus use the sample mean to estimate the mean.

Variance

Variance capture the spread (or width of a distribution)

$$var(x) = \sigma^2 = E\left[(x - \mu)^2 \right]$$

$$var(cx) = c^2 var(x)$$
, so if we re-scale to M/σ we get variance of 1

$$\sigma^{2} = E(x^{2} - 2x\mu + \mu^{2})$$

$$= E(x^{2}) - 2\mu^{2} + \mu^{2}$$

$$= E(x^{2}) - \mu^{2}.$$

$$\sigma^2 = \langle (x - \langle x \rangle)^2 \rangle$$

$$\sigma^2 = E(x^2 - 2x\mu + \mu^2)$$

$$= E(x^2) - 2\mu^2 + \mu^2$$

$$= E(x^2) - \mu^2.$$

$$= ((x_1 - \mu_1) + (x_2 - \mu_2))^2 \rangle$$

$$= \langle (x_1 - \mu_1)^2 + (x_2 - \mu_2)^2 + 2(x_1 - \mu_1)(x_2 - \mu_2) \rangle$$

$$= \langle (x_1 - \mu_1)^2 + (x_2 - \mu_2)^2 + 2(x_1 - \mu_1)(x_2 - \mu_2) \rangle$$

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Estimating Variance

Easy to estimate variance if you already know the population mean

$$v_{\mu}^2 = \frac{1}{N} \sum_{i} \left(x_i - \mu \right)^2$$

But usually you don't, so you have to estimate from the sample mean

Using v to estimate the **population variance** would give a biased estimate.

We rather have to use the (well known) correction:

$$v^{2} = \frac{1}{N} \sum_{i} (x_{i} - \overline{x})^{2}$$

$$= \frac{1}{N} \sum_{i} (x_{i}^{2} - 2x_{i}\overline{x} + \overline{x}^{2})$$

$$= \frac{1}{N} \sum_{i} x_{i}^{2} - \overline{x}^{2}.$$

$$\frac{v^2}{N-1} = \frac{\sum_i (x_i - \overline{x})^2}{N(N-1)}$$

Higher Moments...

$$\gamma_1 = E\left[(x - \mu)^3 \right] / \sigma^3$$
 Skewness

$$\beta_2 = E\left[(x - \mu)^4 \right] / \sigma^4$$

$$\gamma_2 = \beta_2 - 3 ,$$

 $\gamma_2 = \beta_2 - 3$,

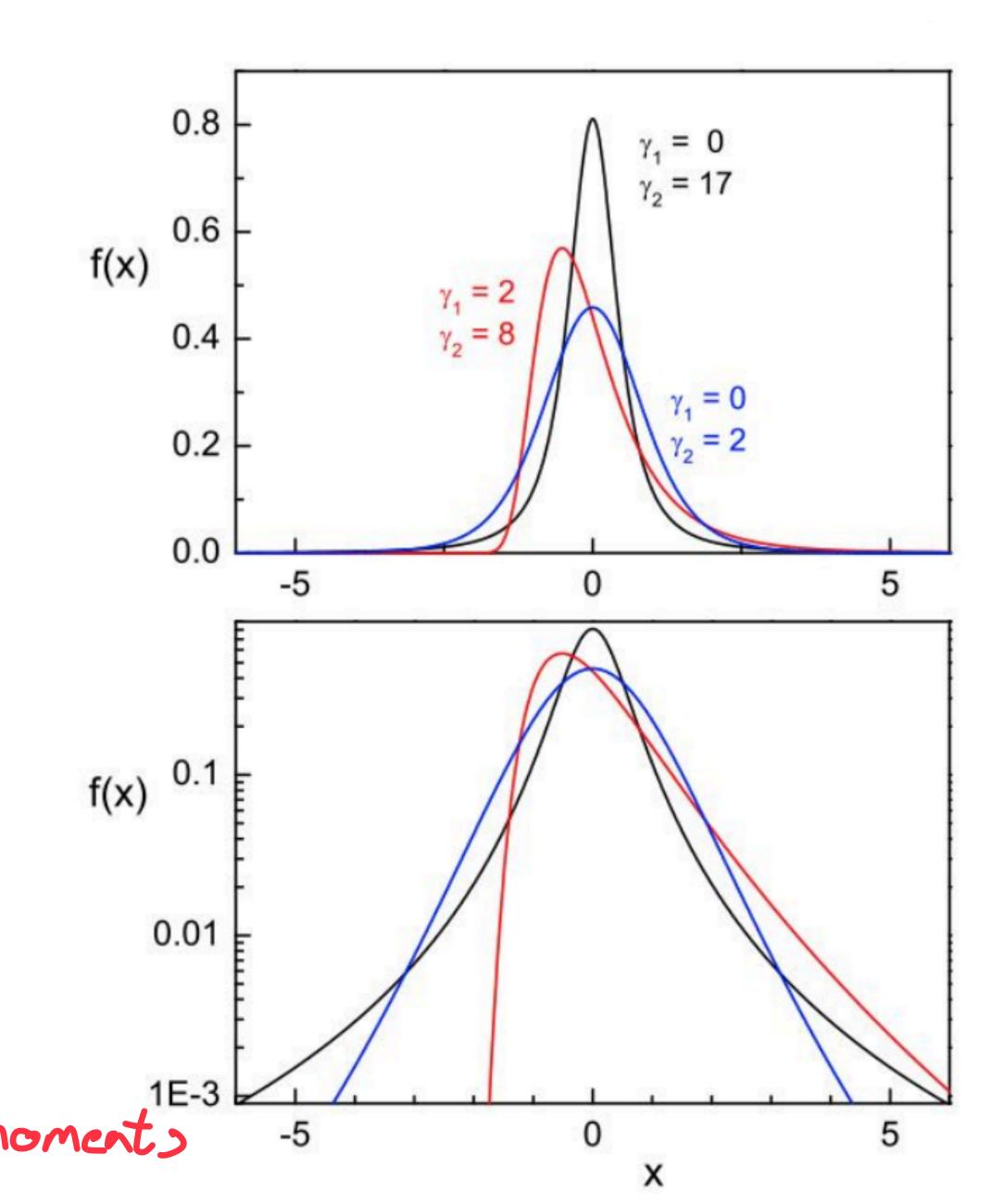
Kurtosis

Central moments

$$\mathcal{U}_{n}' = E((x-\mu)^{2})$$

$$= \int_{-\infty}^{\infty} (x-\mu)^{2} + Cx dx$$

It distributions have the same moments of they are identical



Gaussian (or Normal) Distribution

$$f(x) = \frac{1}{\sigma\sqrt{2\pi}}e^{-\frac{1}{2}(\frac{x-\mu}{\sigma})^2}$$
Probably the best
$$k_{\text{nown}} \text{ PDF}$$
at $x = \mu^2 \sigma$,
$$y = y^{\text{max}}/\sqrt{e}$$

$$\sim 0.606 \text{ yms}$$

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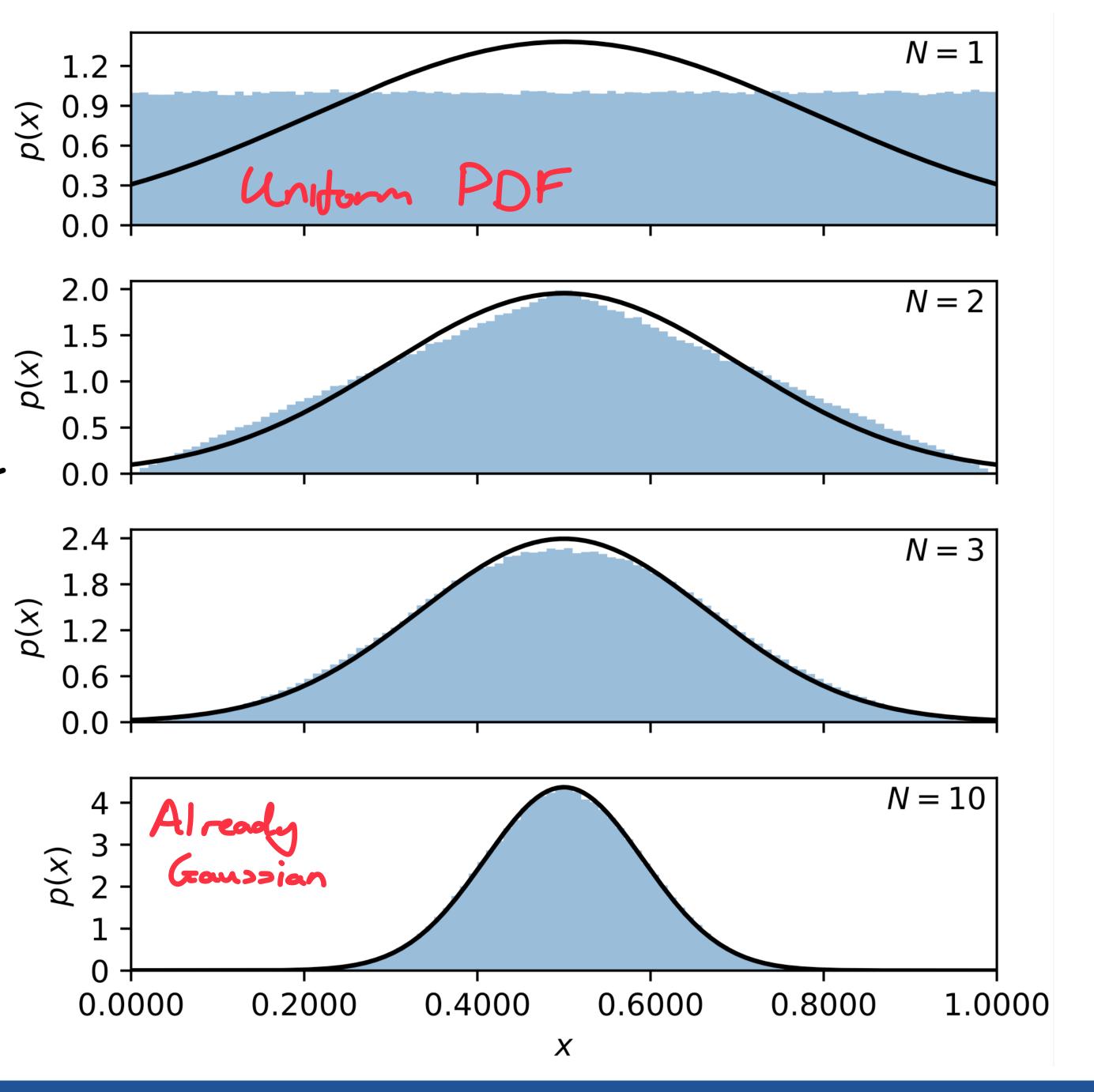
Central Limit Theorem

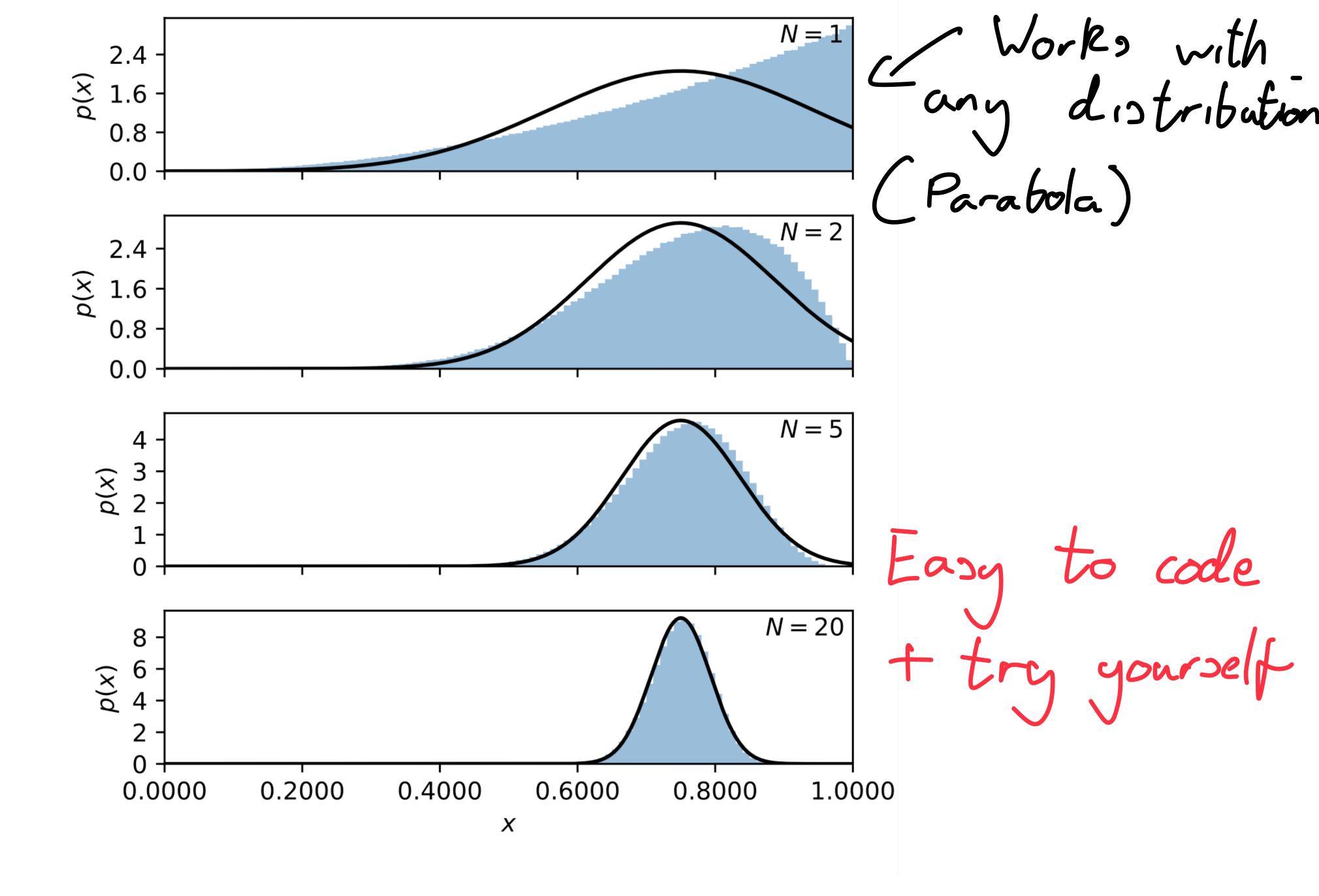
Pick K random variables from any PDF

Repeat N times and take the average (Cor sum)

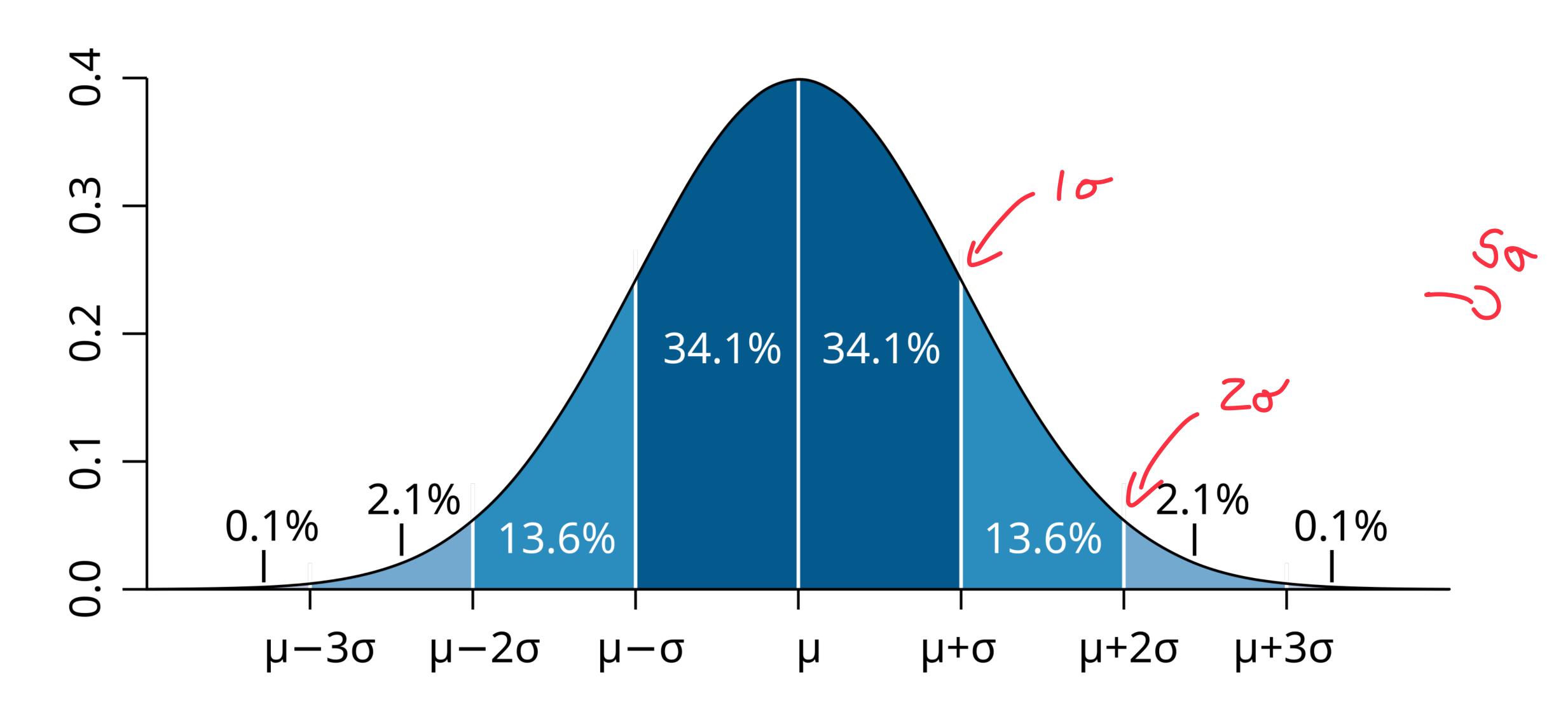
Distribution of mean will follow a Gaussian Cat large enough ND

Larger N for non-uniform distribution





Gaussian (or Normal) Distribution



Binomial Distribution

Probability of events with 2 possible outcomes

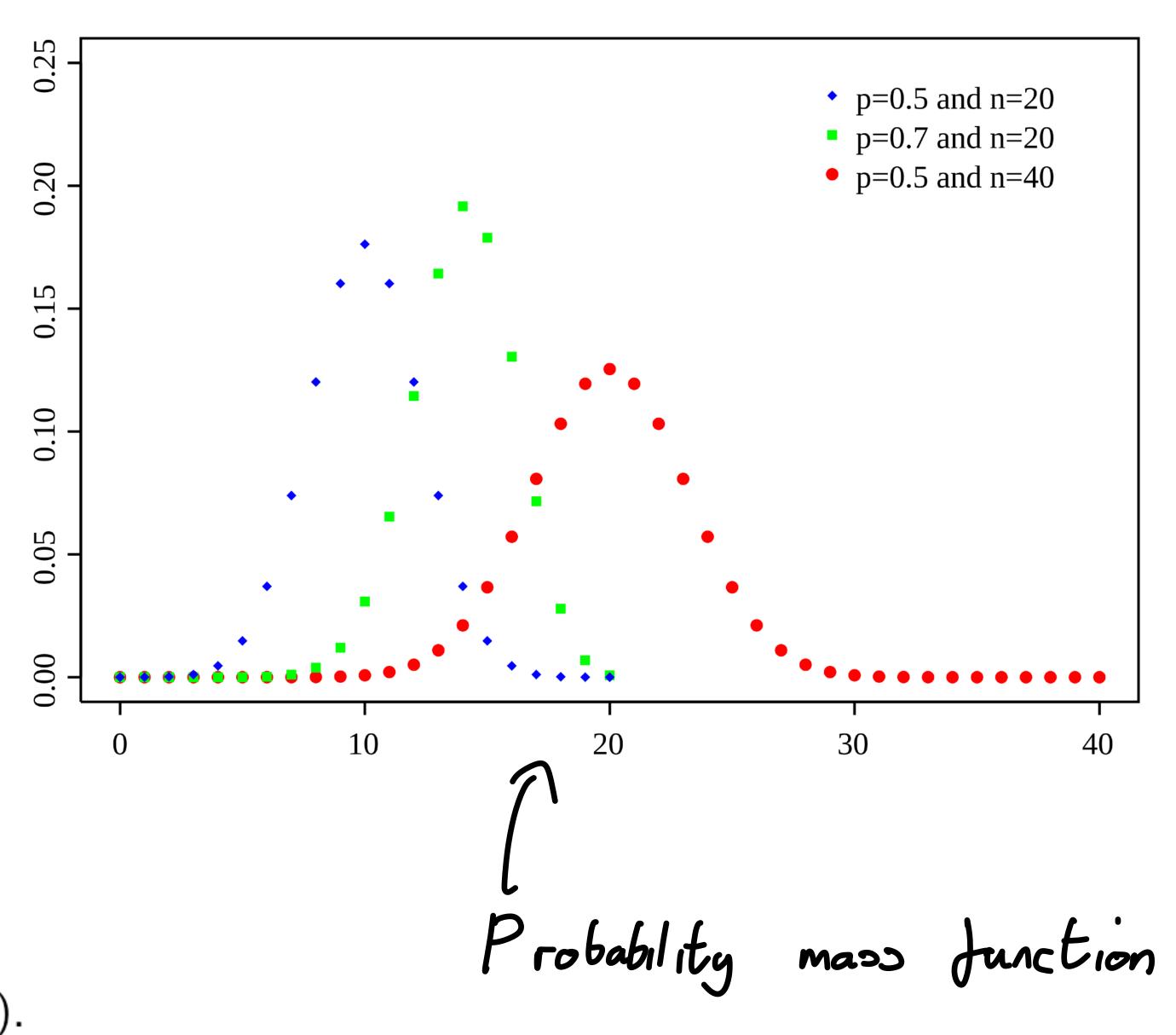
True or False, Heads or tail
Roll a 6 on a dice 34 times in
100 rolls

$$K = 34$$
, $N = 100$, $P = \frac{1}{6}$

$$P(k; p, N) = \frac{k!}{k!(N-k)!} p^{k} (1-p)^{N-k}$$

$$E[k] = \sum_{k} kP(k) = Np.$$

 $E[(k - \langle k \rangle)^{2}] = E[k^{2}] - (E[k])^{2} = Np(1 - p).$

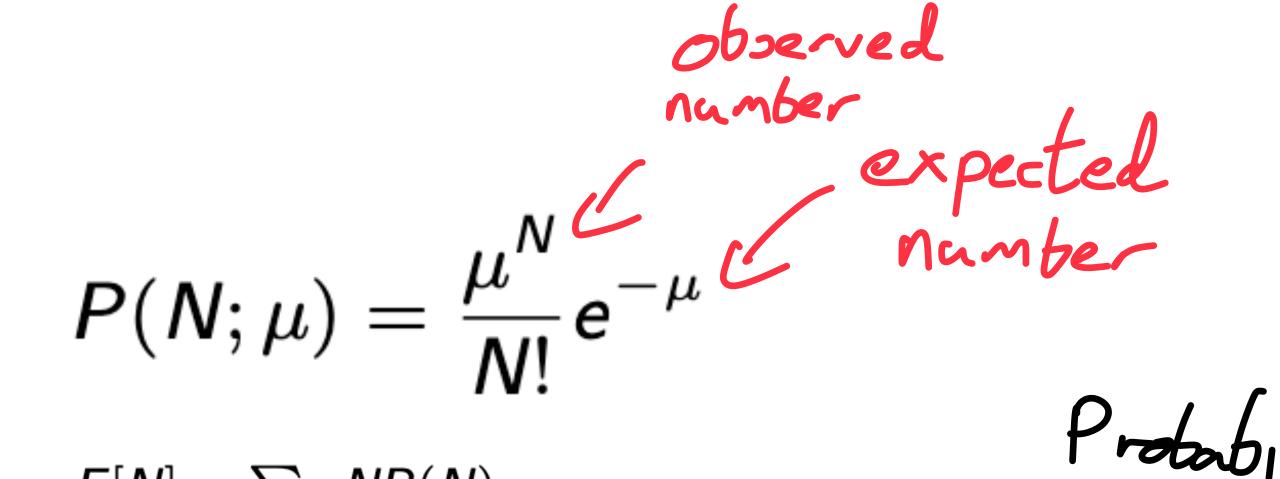


Poisson Distribution

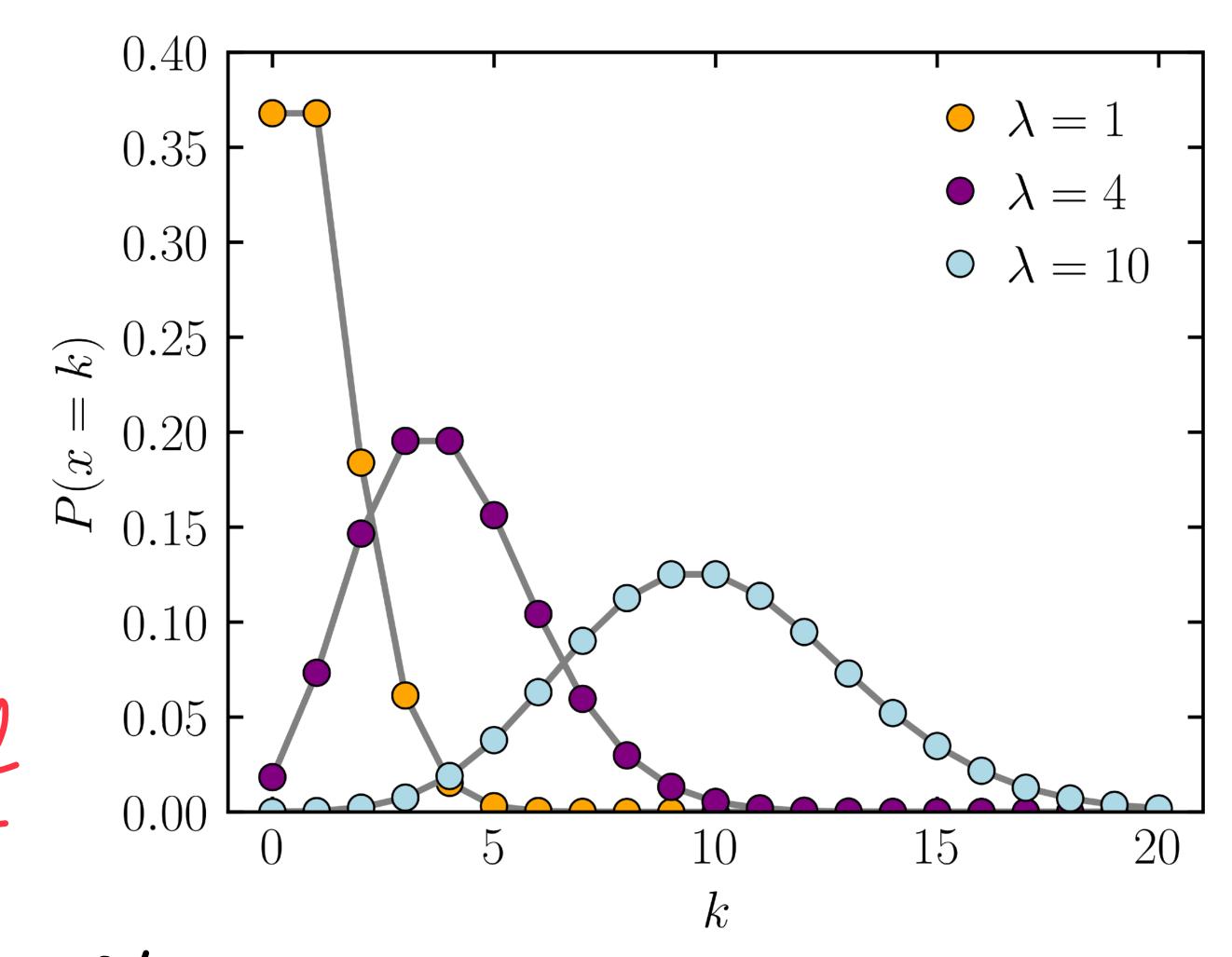
For $N \rightarrow \infty$, $p \rightarrow 0$

Np= const.,

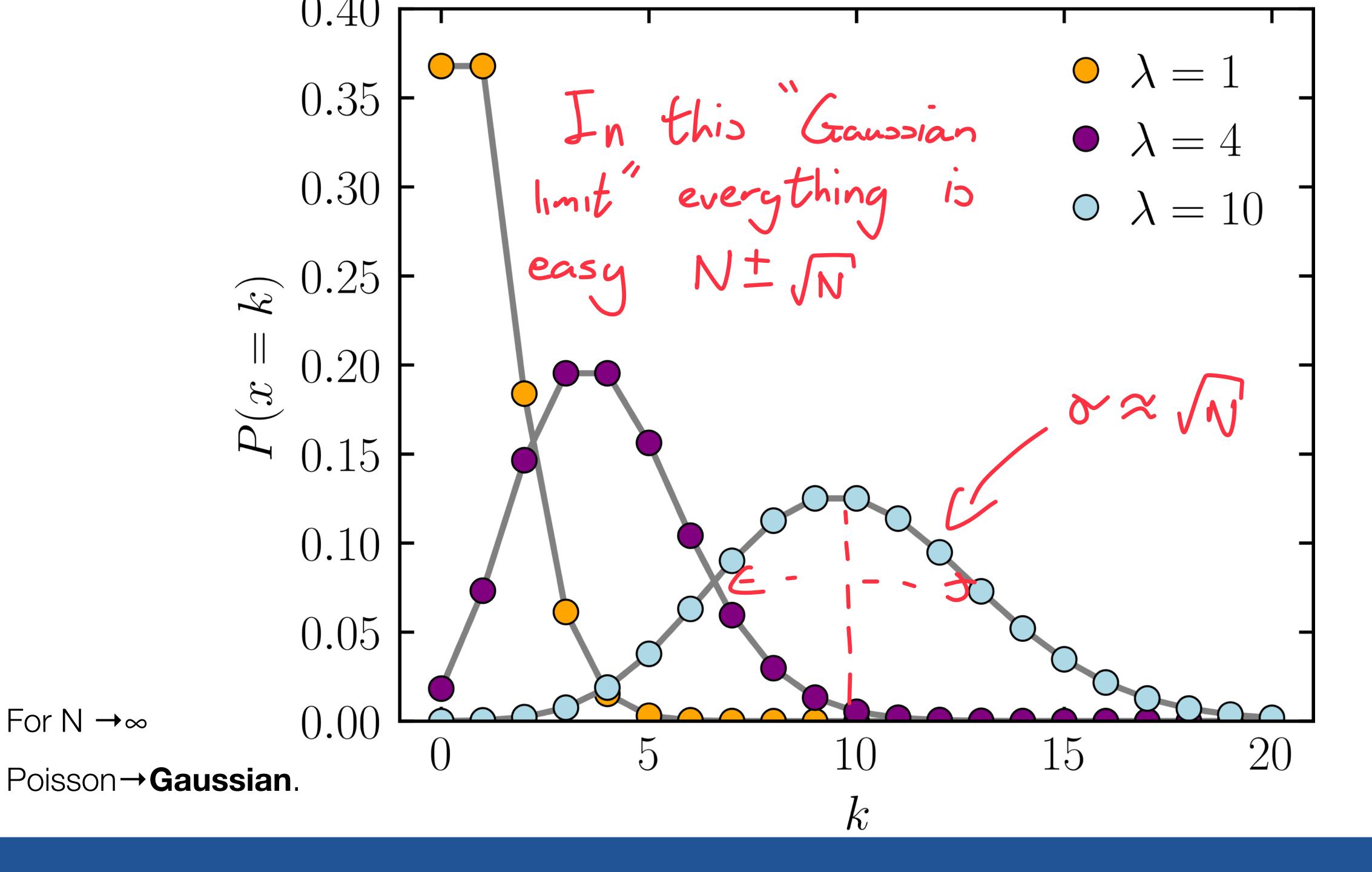
Binomial → Poisson



$$E[N] = \sum_{N} NP(N) = \mu.$$
 $\sigma^2 = \mu.$



Probability of a discrete number of events in a fixed time period



Exponential Distribution

Used typically in processes that have no "memory"

e.g. radioactive decay

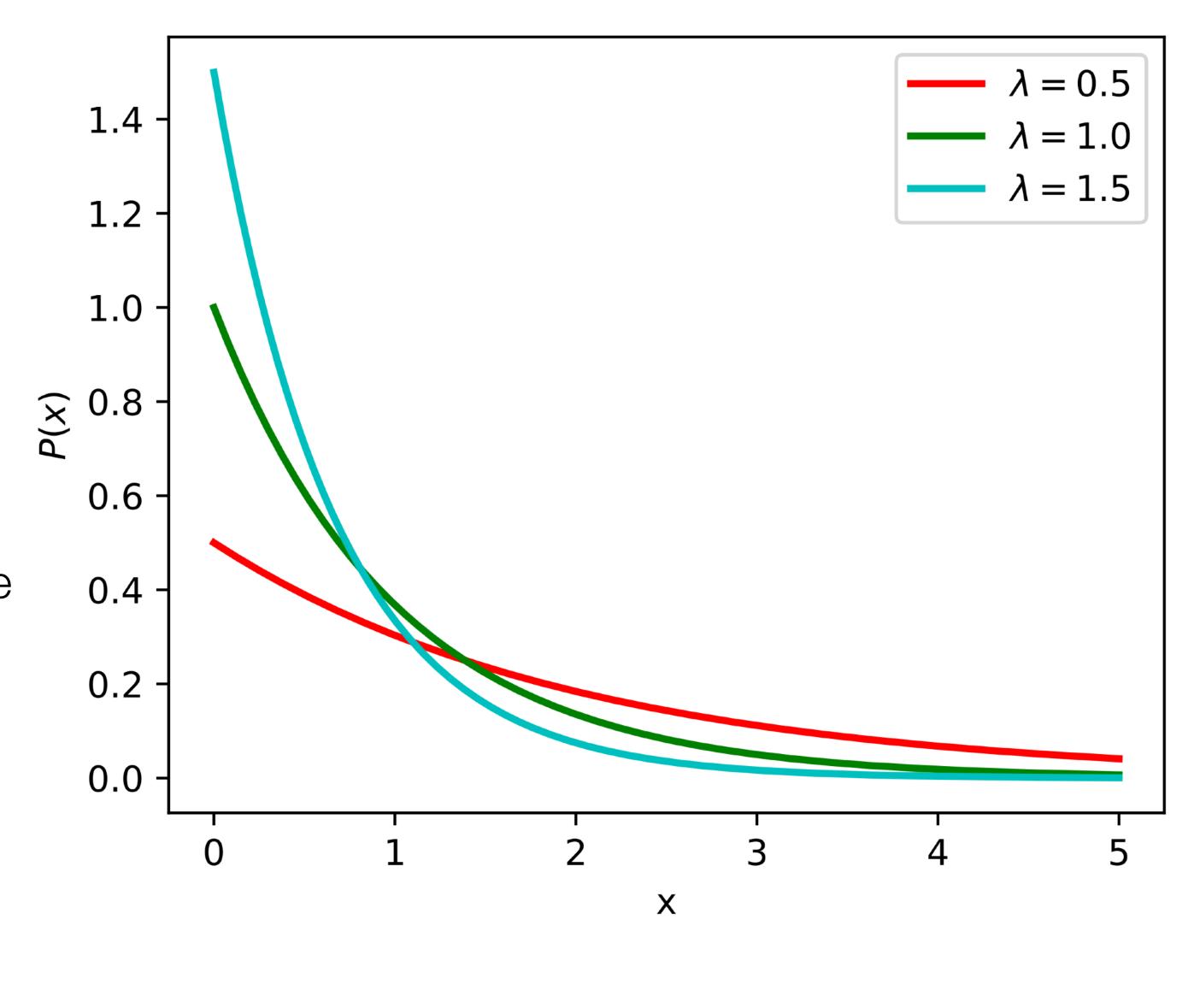
Probability of decay in each time bin **dt** is constant

Lambda>0 and equal to inverse if the lifetime

$$f(t) \equiv f(t|\lambda) = \lambda e^{-\lambda t}$$
 for $t \ge 0$

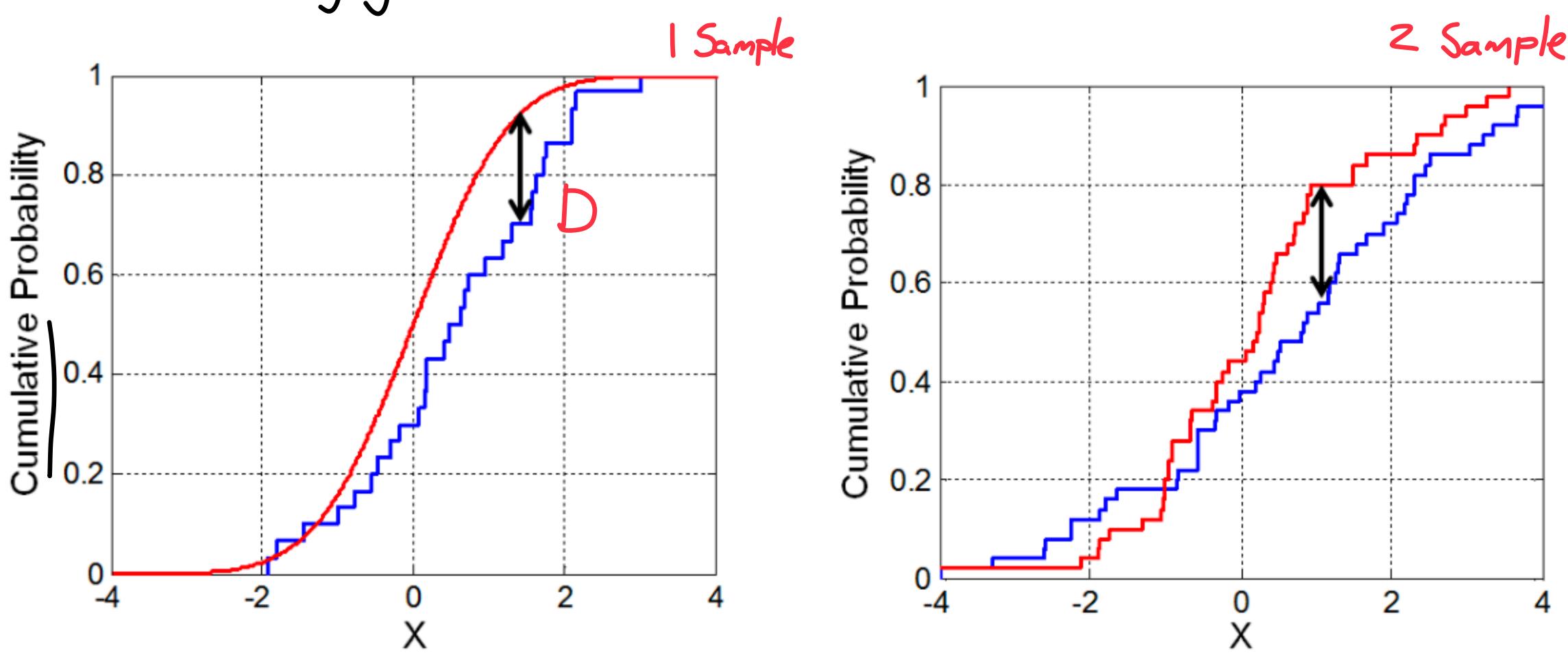
$$F(t) = \int_{-\infty}^{t} f(t')dt' = 1 - e^{-\lambda t}$$

Mean: $\frac{1}{\lambda}$ Variance: $\frac{1}{\lambda^2}$



Kolmogorov Smirnov Test

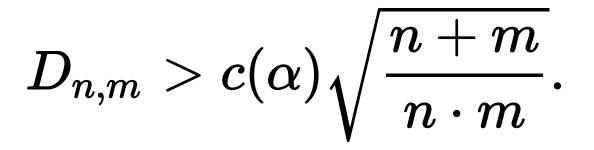
Compare probability that 2 distributions come from same unalying PDF

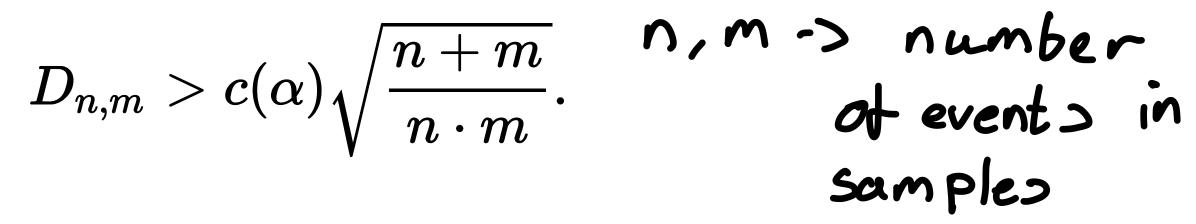


Kolmogorov Smirnov Test

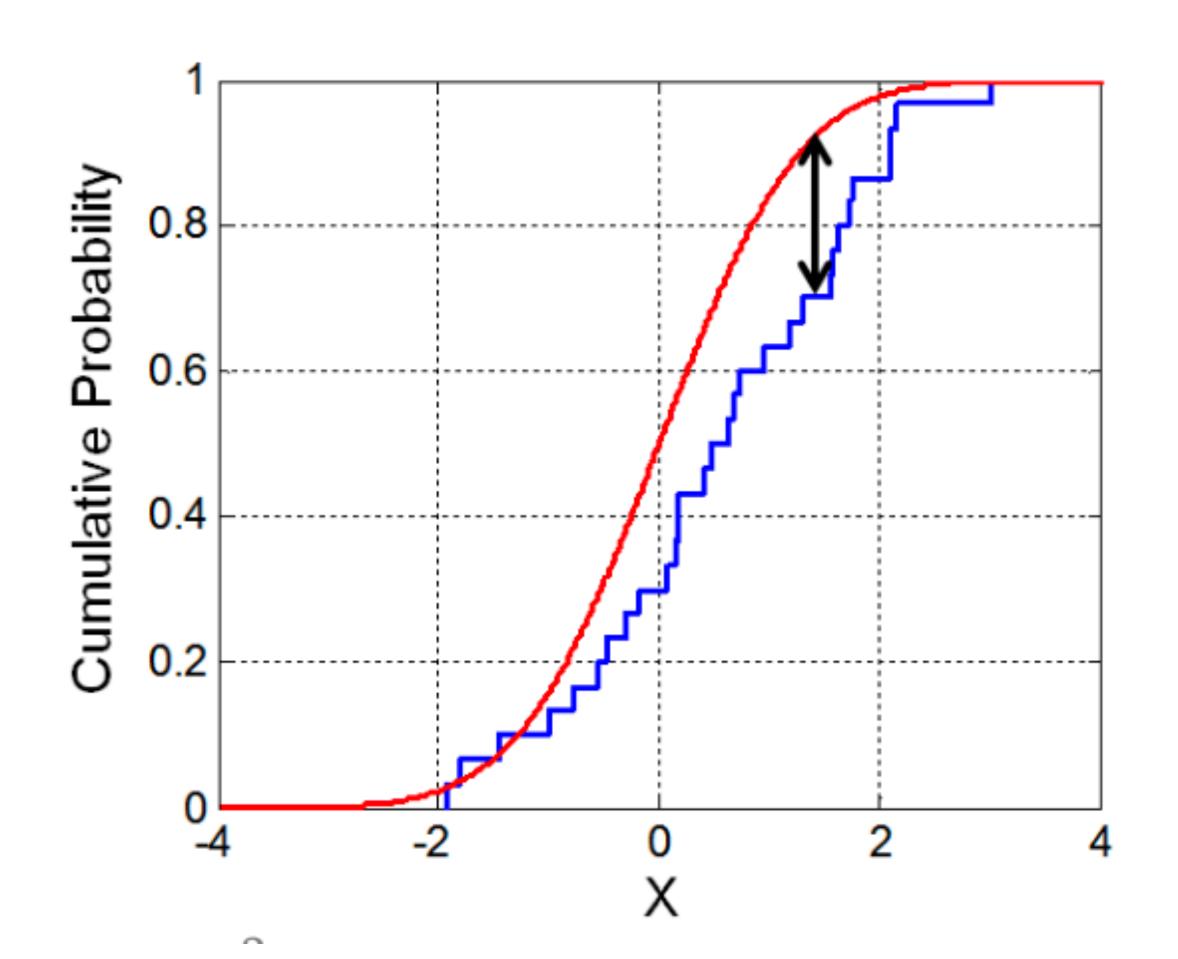
$$D_n = \sup |F_n(x) - F(x)|$$
 Level

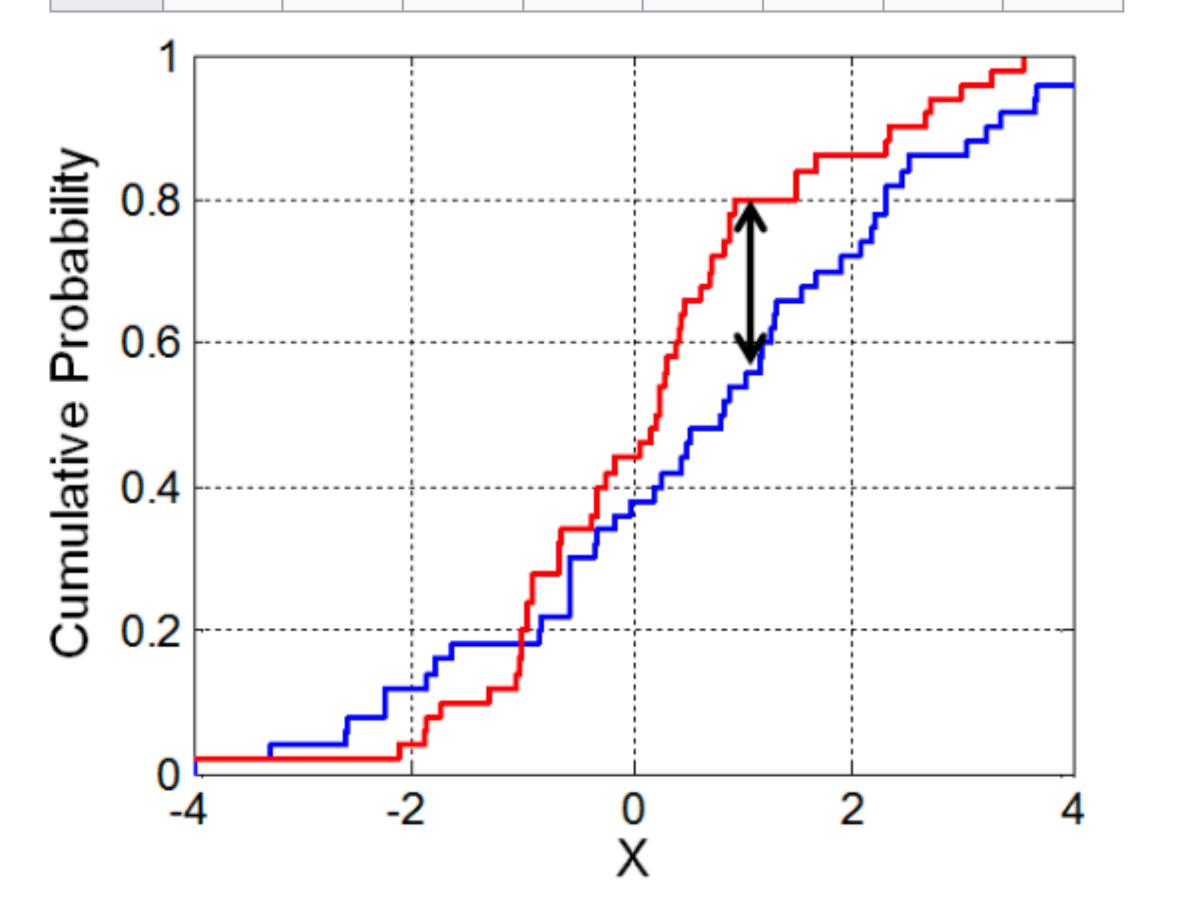






| α | 0.20 | 0.15 | 0.10 | 0.05 | 0.025 | 0.01 | 0.005 | 0.001 |
|----------|-------|-------|-------|-------|-------|-------|-------|-------|
| c(lpha) | 1.073 | 1.138 | 1.224 | 1.358 | 1.48 | 1.628 | 1.731 | 1.949 |



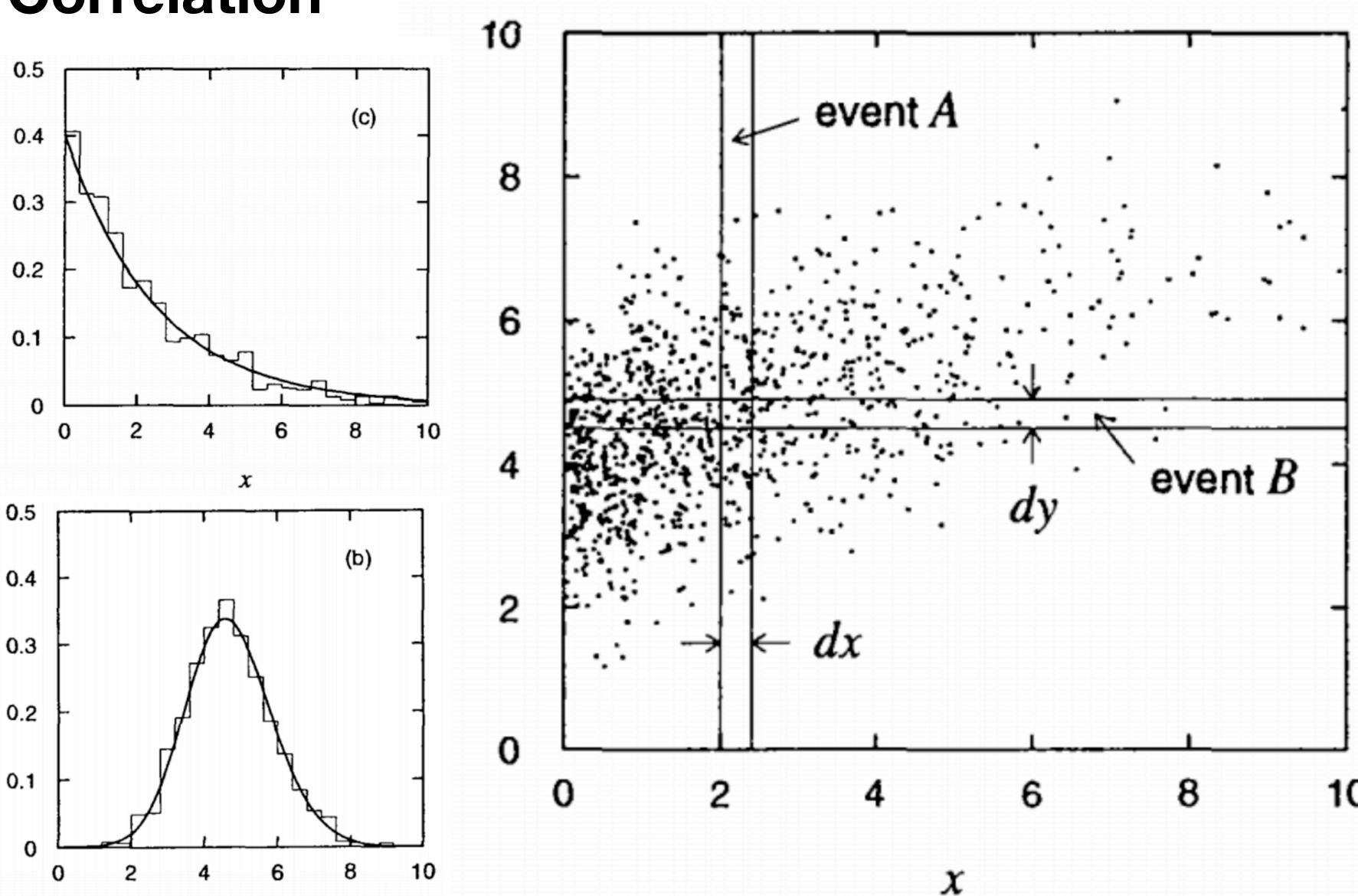


At some point you will come across distributions dependent on several variables

What is the probability for A and B

P(A) = f(x) dec P(B) = f(y) dy =) P(A,B) = f(x,y) decdy

Can project along one axis if you one care about one variable



How correlated are
$$x$$
 and y ?

Covariance

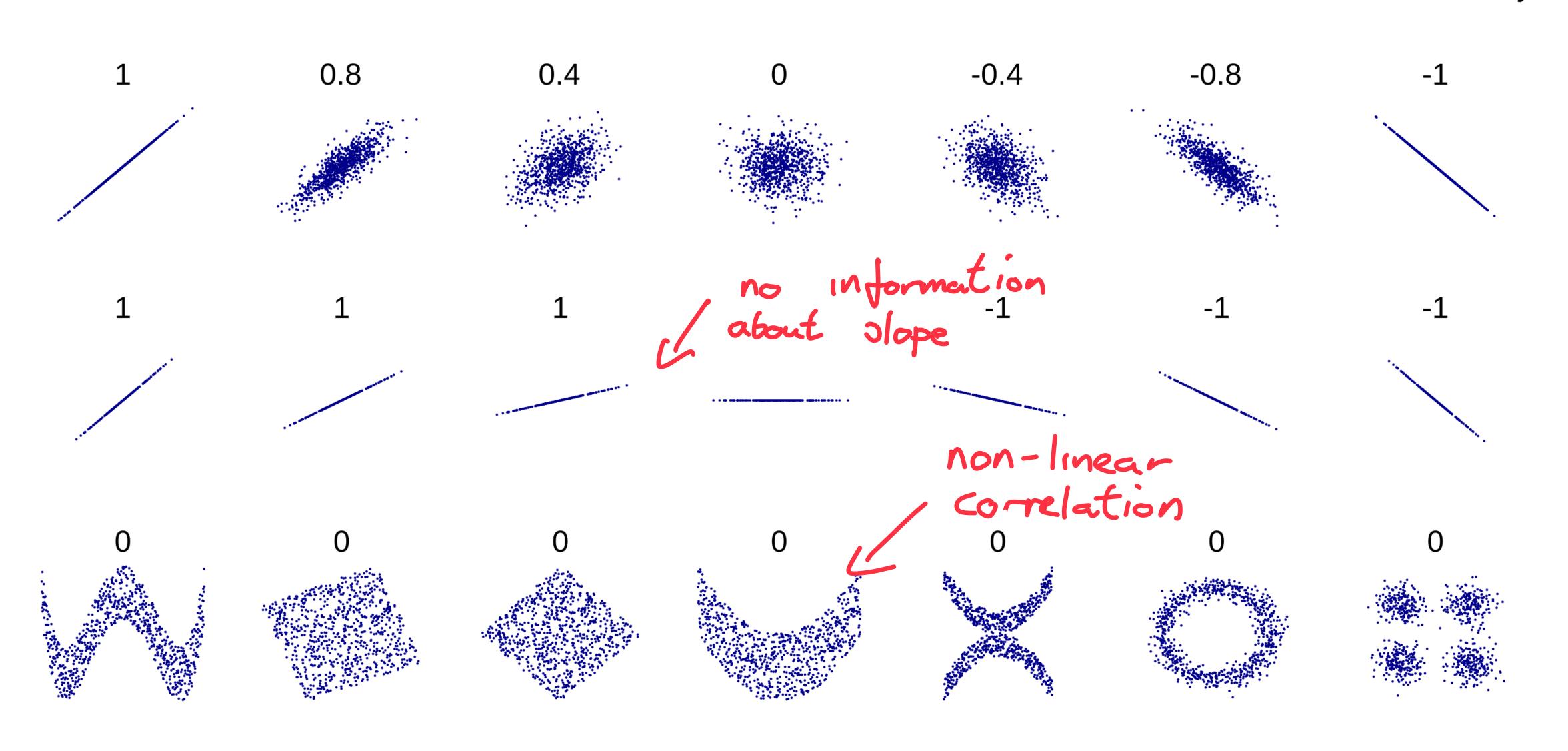
 $V(\infty) = \sigma_x^2 = E[(x - (\infty 7)^2]]$
 $= E[x^2] - (E[x])^2$
 $= E[xy] - E[x]E[y] - 1$

If uncorrelated

 $P(A \wedge B) = P(A) P(B)$

Correlation Coefficient

$$\rho_{x,y} = \frac{C(x,y)}{\sigma_x \sigma_y}$$



$$P(\vec{x}) = \frac{1}{2\pi\sqrt{\det(C)}} \exp\left(-\frac{1}{2}(\vec{x} - \vec{\mu})^T C^{-1}(\vec{x} - \vec{\mu})\right)$$
for $\vec{x} = (x, y) \Rightarrow C = \begin{pmatrix} \sigma_x^2 & \rho \sigma_x \sigma_y \\ \rho \sigma_x \sigma_y & \sigma_y^2 \end{pmatrix} = \begin{pmatrix} \sigma_x^2 & V_{x,y} \\ V_{x,y} & \sigma_y^2 \end{pmatrix}$

$$V_{x,y} = A \log C$$

Parameter Estimation

Least Squares Fit

Best known dit statistic

$$\chi^2 = \sum_{i} \left(\frac{(ax_i + b) - y_i}{\sigma} \right)^2$$

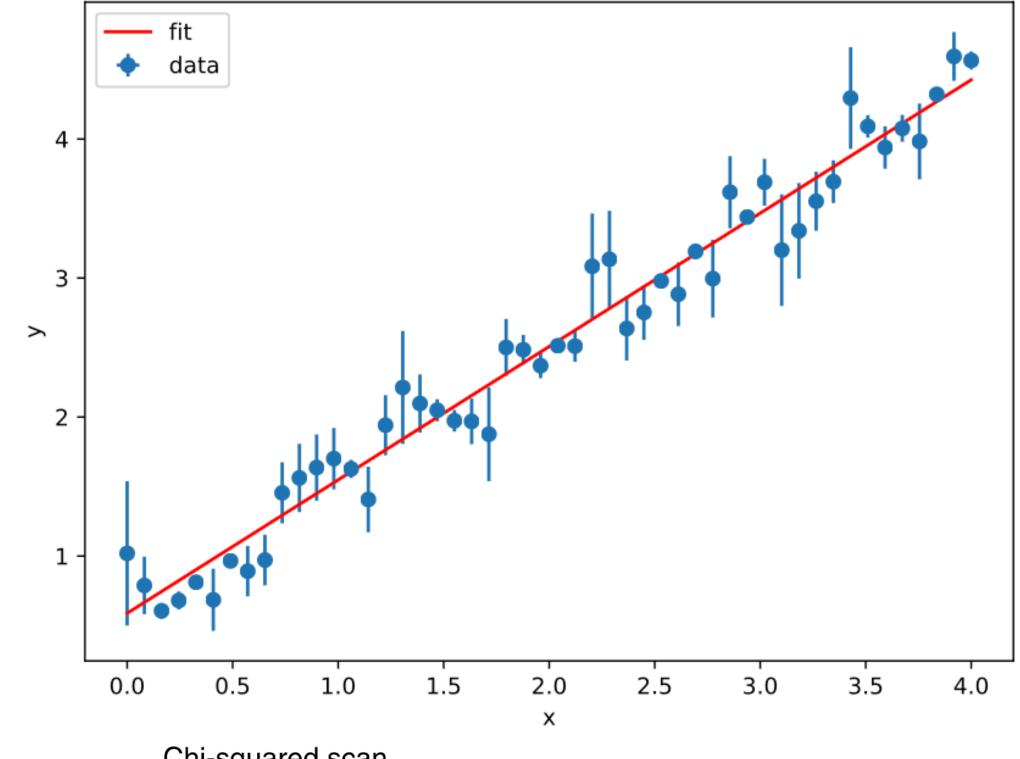
Not Jast good for linear functions

Error bors should be gaussian

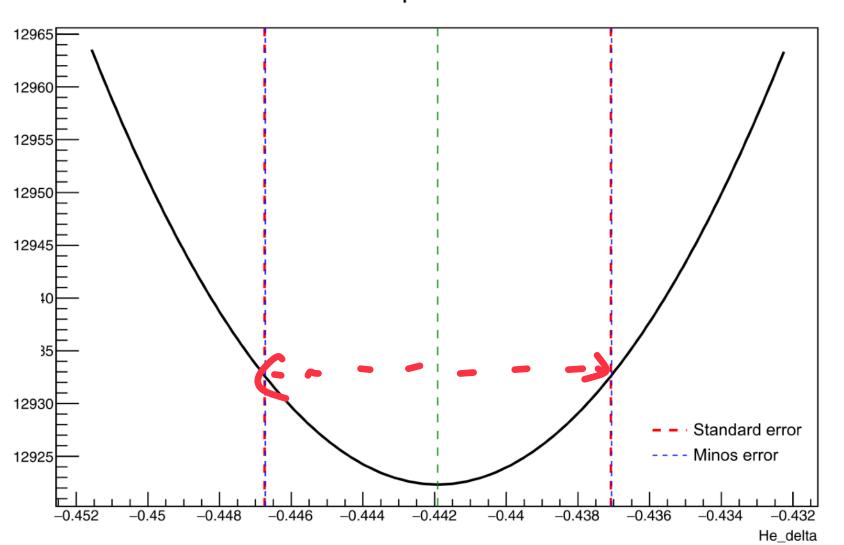
Goodness of fit given by reduced

$$\sigma_{\theta}^2 = 2 / \left(d^2 \chi^2 / d\theta^2 \right)$$

$$\chi^2(\theta) = \chi^2_{\text{min}}(\theta_{\text{best}}) + 1;$$



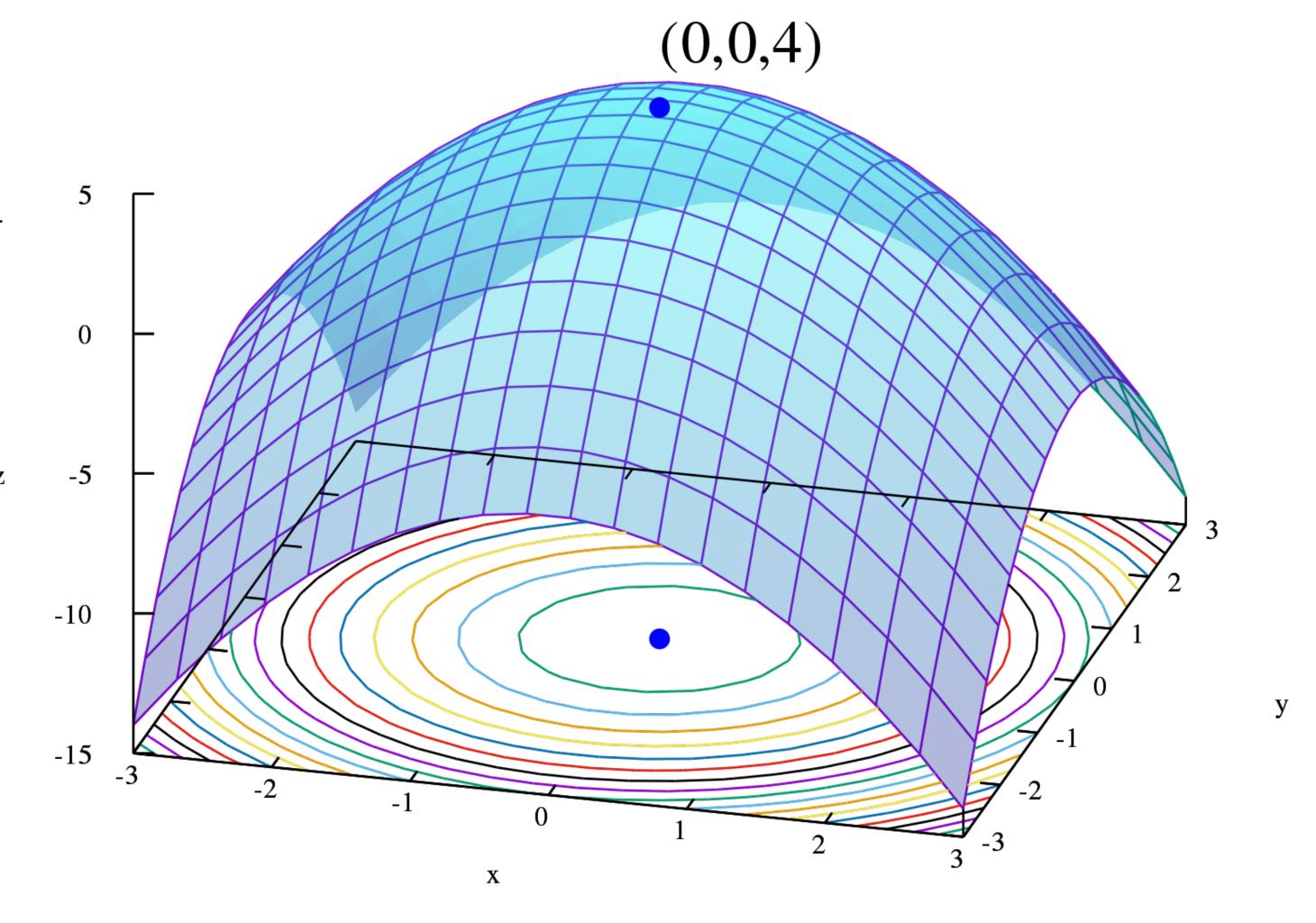
Chi-squared scan



What is Function Minimisation?

"In the more general approach, an optimization problem consists of maximizing or minimizing a real function by systematically choosing input values from within an allowed set and computing the value of the function." -

Wikipedia



https://en.wikipedia.org/wiki/Mathematical_optimization#/media/File:Max_paraboloid.svg

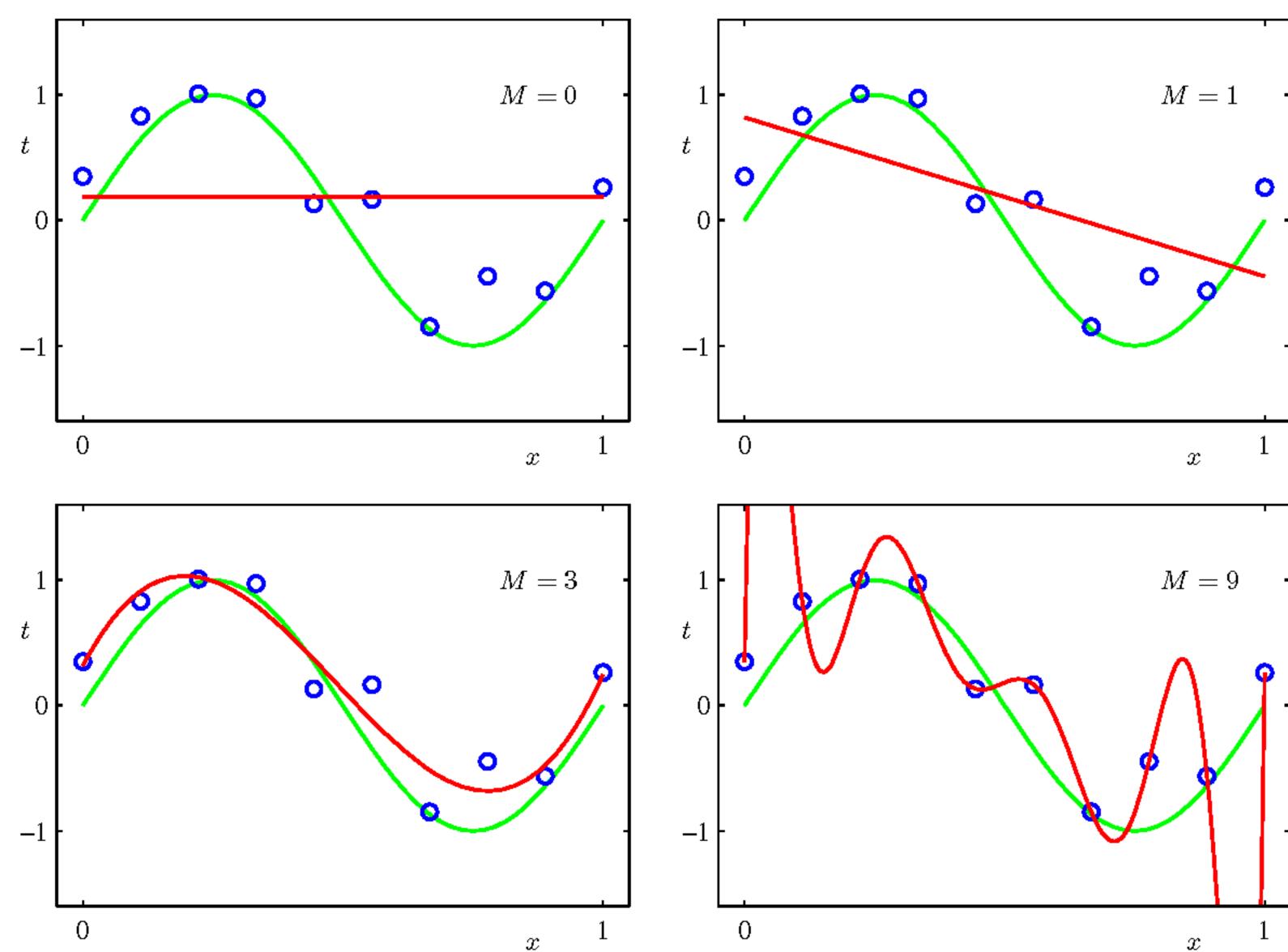
... and why do we want to do it?

Although there are many reasons why you might want to do this

But the most obvious of these is is function fitting

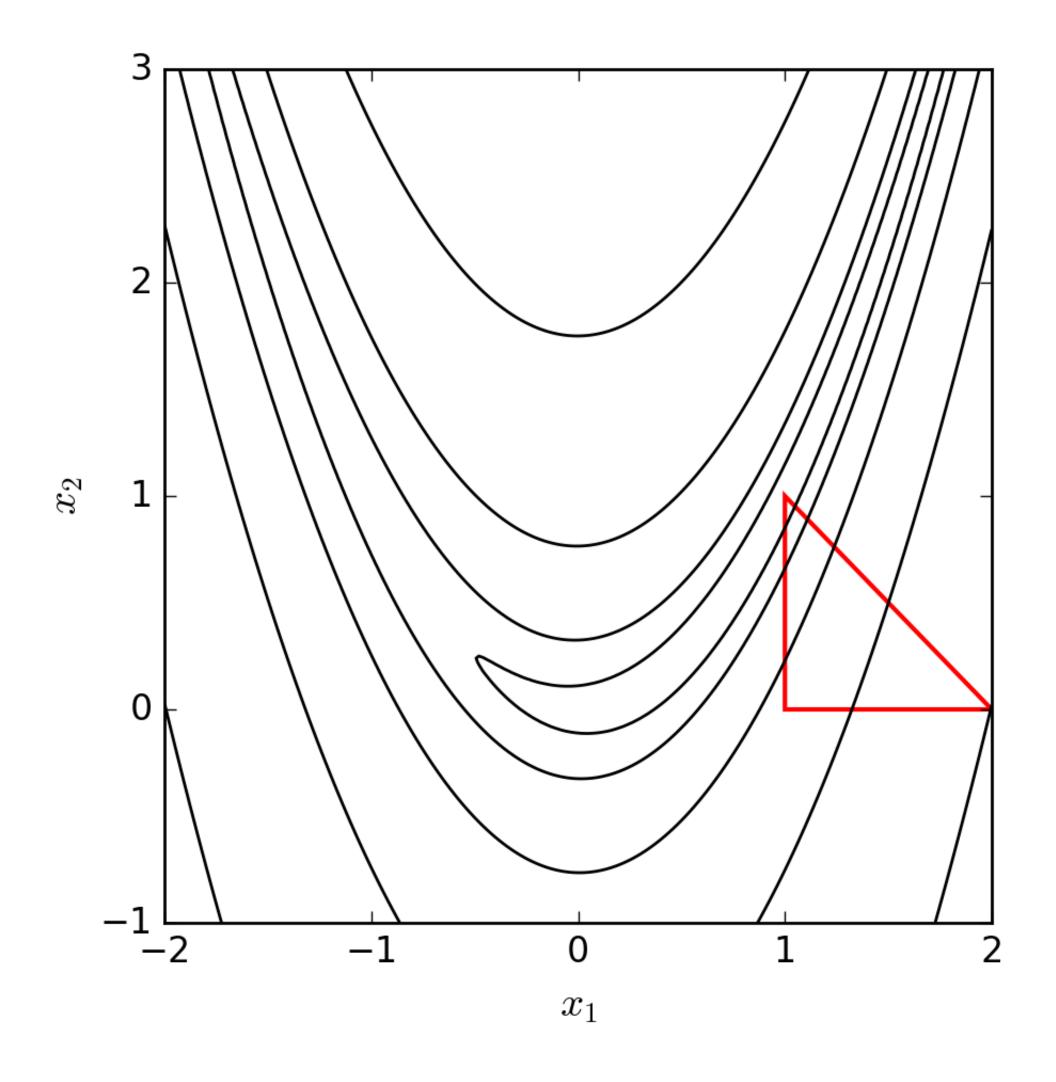
Minimise objective function describing difference between model and data

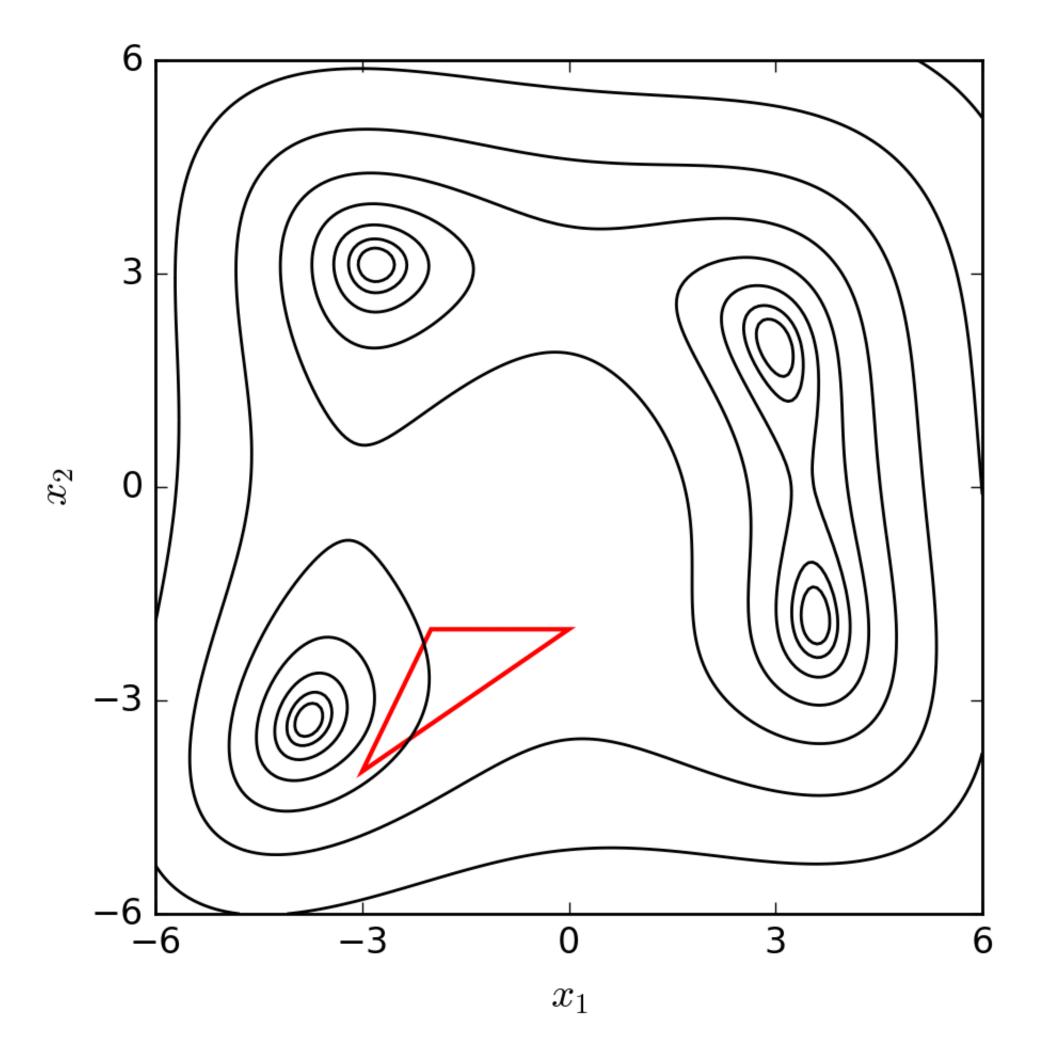
Chi2, likelihood etc



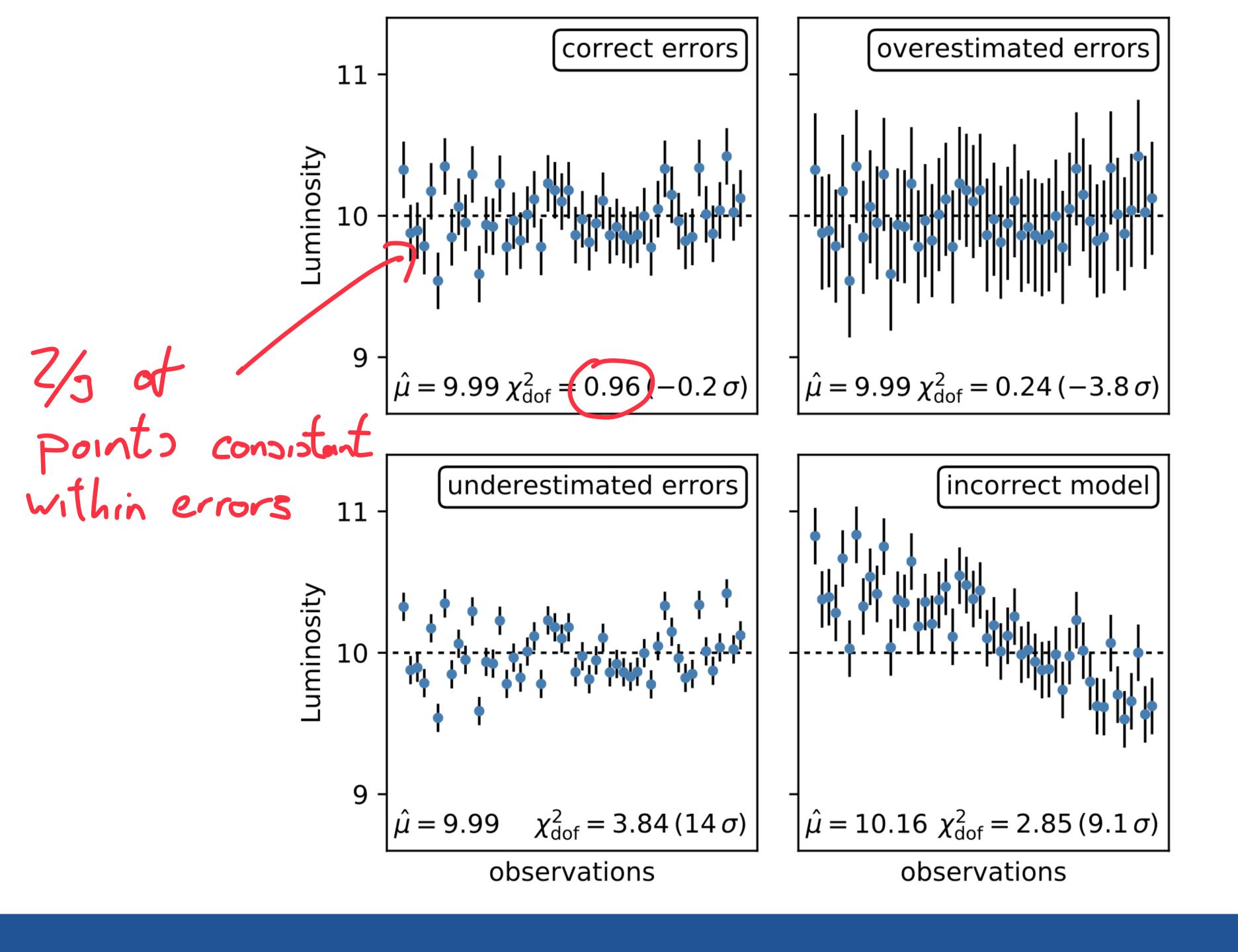
https://www.semanticscholar.org/paper/Introduction-1.1.-Example%3A-Polynomial-Curve-Fitting/001cc52a63e9ed49b7a3c29dc5c1a0dc16406049

Let's start Simple(x)





https://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Himmelblau.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia.org/wiki/Nelder%E2%80%93Mead_method#/media/File:Nelder-Mead_Rosenbrock.gifhttps://en.wikipedia/File:Nelder-Mead_Rosenbrock.gifhttp



Likelihood

PDF constructed with ∞ as $x\mapsto f(x\mid\theta),$

 $\theta\mapsto f(x\mid\theta),$ -> Likelihood asks question, how likely is my data given this model

We can create a likelihood function by assuming a PDF then comparing model to datapoints

 $L(\theta) = \prod_{i}^{N} p(x_{i}, \theta)$ or $\ln L(\theta) = \sum_{i}^{N} \ln p(x_{i}, \theta)$

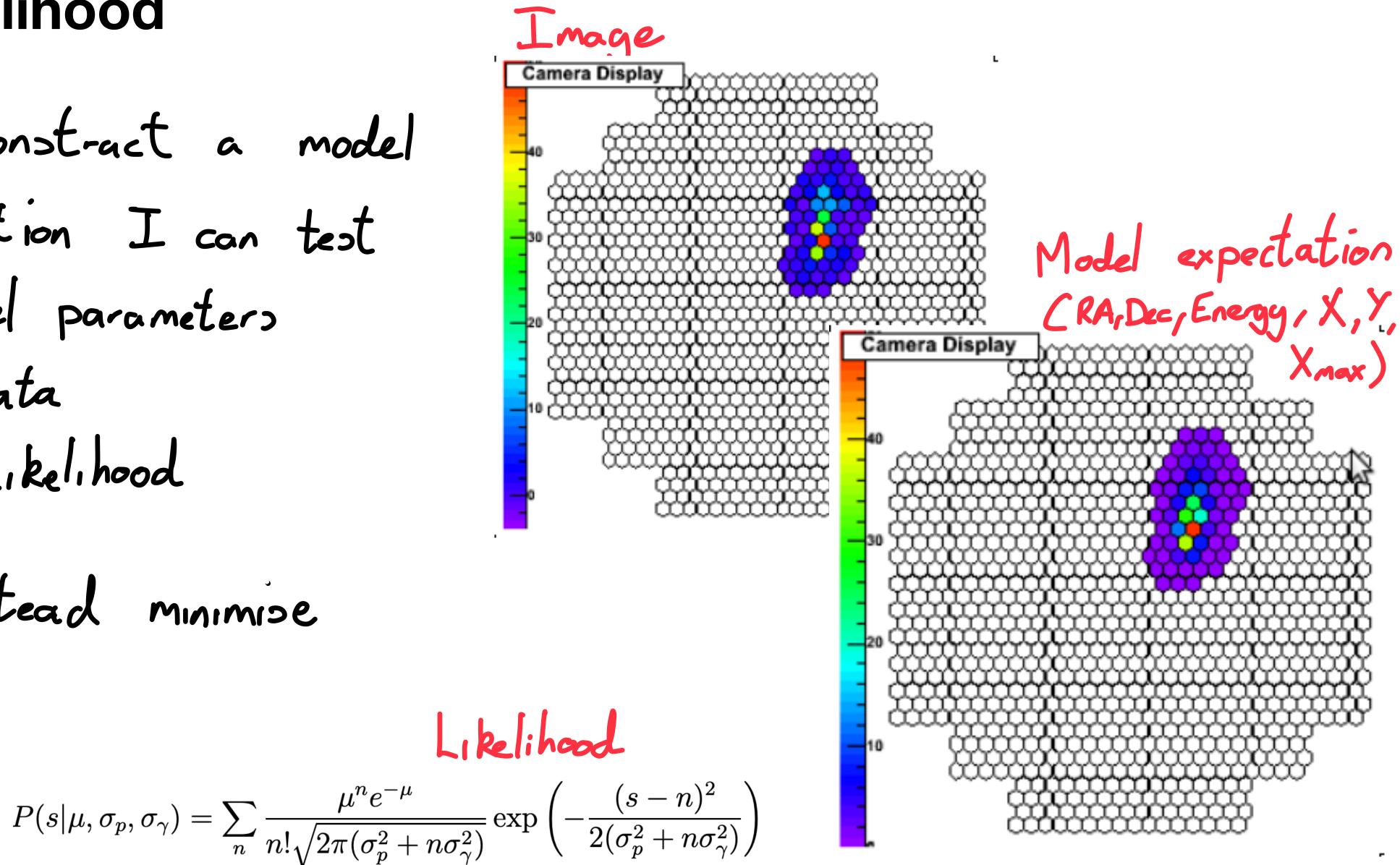
Likelihood

Assume we Hip a coin 3 times, what is the L (HHT/PH=0.5) $= 0.5^3 = 0.125$ But we can test other assumptions about the 6010 LCHHT, PH=1) = IXIXO = 0 L CHHT, PH = 0.25) = 0.25 x 0.25 x 0.75 = 0.045 We can learn something about coin! Most time you will see this as -2 log L= TS

Maximum Likelihood

It I can constract a model + likelihood function I can test find which model parameters best tit my data 1.e. Maximise Likelihood

Usually you instead minimise -2 In L



Maximum Likelihood

We can then search model space using function minimisation like X2 In fact X is a special case of MLE Calculate parameter errors the same, increase of -2 In L of 1

Uncertainties

Now I have best fit values + uncert, how to compare to data

Lets say I get a ratio of Parameter/expectation of 0.97, is this compatible with theory

 0.97 ± 0.005 0.97 ± 0.05 0.97 ± 0.5

Uncertainties

Statistical uncertainties

- Random in nature.
- Fluctuates independently per measurement.
- Unavoidable.
- > Usually, more data → lower uncertainty ($\propto \sqrt{N}$).
- e.g., counting statistics, electronic noise, etc.

Systematic uncertainties

- Usually originate in the instrument.
- > Bias the data by unknown∼constant offset.
- Hard to detect, correct for, estimate.
- > e.g., miscalibration, diff. between data and simulation, simulation statistics, etc.

Normally report these separately

1.0 0.97 ± 0.005 = 0.03 sys

can occur on data + theoretical expectation

Uncertainties

| $z = ax \pm b$ | $\delta z = a \cdot \delta x$ | | |
|---------------------|---|--|--|
| $z = x \pm y$ | $\delta z = [(\delta x)^2 + (\delta y)^2]^{\frac{1}{2}}$ | | |
| z = cxy | $\frac{\delta z}{z} = \left[\left(\frac{\delta x}{x} \right)^2 + \left(\frac{\delta y}{y} \right)^2 \right]^{\frac{1}{2}}$ | | |
| $z = c \frac{x}{y}$ | $\frac{\delta z}{z} = \left[\left(\frac{\delta x}{x} \right)^2 + \left(\frac{\delta y}{y} \right)^2 \right]^{\frac{1}{2}}$ | | |
| $z = cx^a$ | $\frac{\delta z}{z} = a \frac{\delta x}{x}$ | | |
| $z = cx^a y^b$ | $\frac{\delta z}{z} = \left[\left(a \frac{\delta x}{x} \right)^2 + \left(b \frac{\delta y}{y} \right)^2 \right]^{\frac{1}{2}}$ | | |

Errors can be Propagated in the Well known way

Only works it independent and gaussian

It not independent then covariance matrix needed

An example, measuring Higgs properties

Null model Ho, SM only

New physics Hi

Simple hypothesis

Calculate $\mathcal{L}(H_0(\theta))$

- > decide if data is likely for H_0 (p-value).
- > If not, claim discovery (of what?)
- Existence of a particle (Higgs, new particle)
- \rightarrow A new γ -ray source.

Composite hypothesis

compare $\mathcal{L}(H_0(\theta))$ and $\mathcal{L}(H_1(\theta))$.

- > Usually likelihood ratio is used.
- \rightarrow More sensitive to H_1 .
- > Based on *p*-values, which *H_i* is more likely.
- Particle with certain mass, width, coupling constants.
- ightharpoonup Position and spectra of γ -ray source.

| | | True State of Nature | |
|--------------|------------------------------|----------------------|-------------------------|
| | | H_0 is true | H ₀ is false |
| Our Decision | Do not reject H ₀ | Correct decision | Type II error |
| | Reject H ₀ | Type I error | Correct decision |

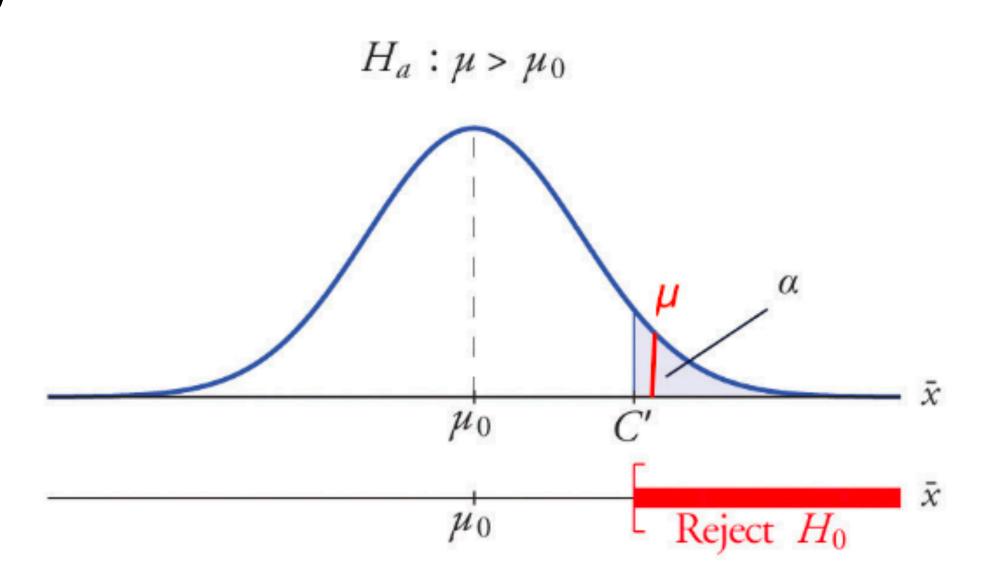
Type 1 error rate (Dignificance X)

$$\alpha = \int_{x \ge x_0} P(x|H_0) dx$$

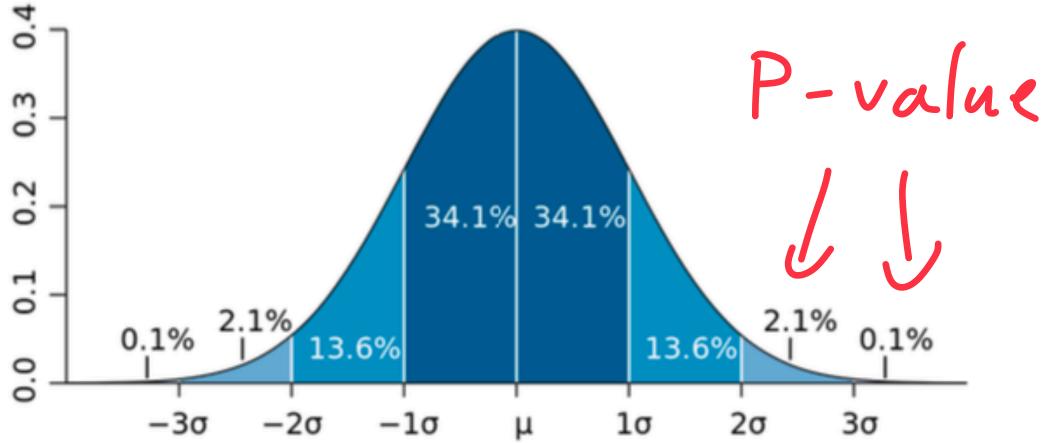
P-value

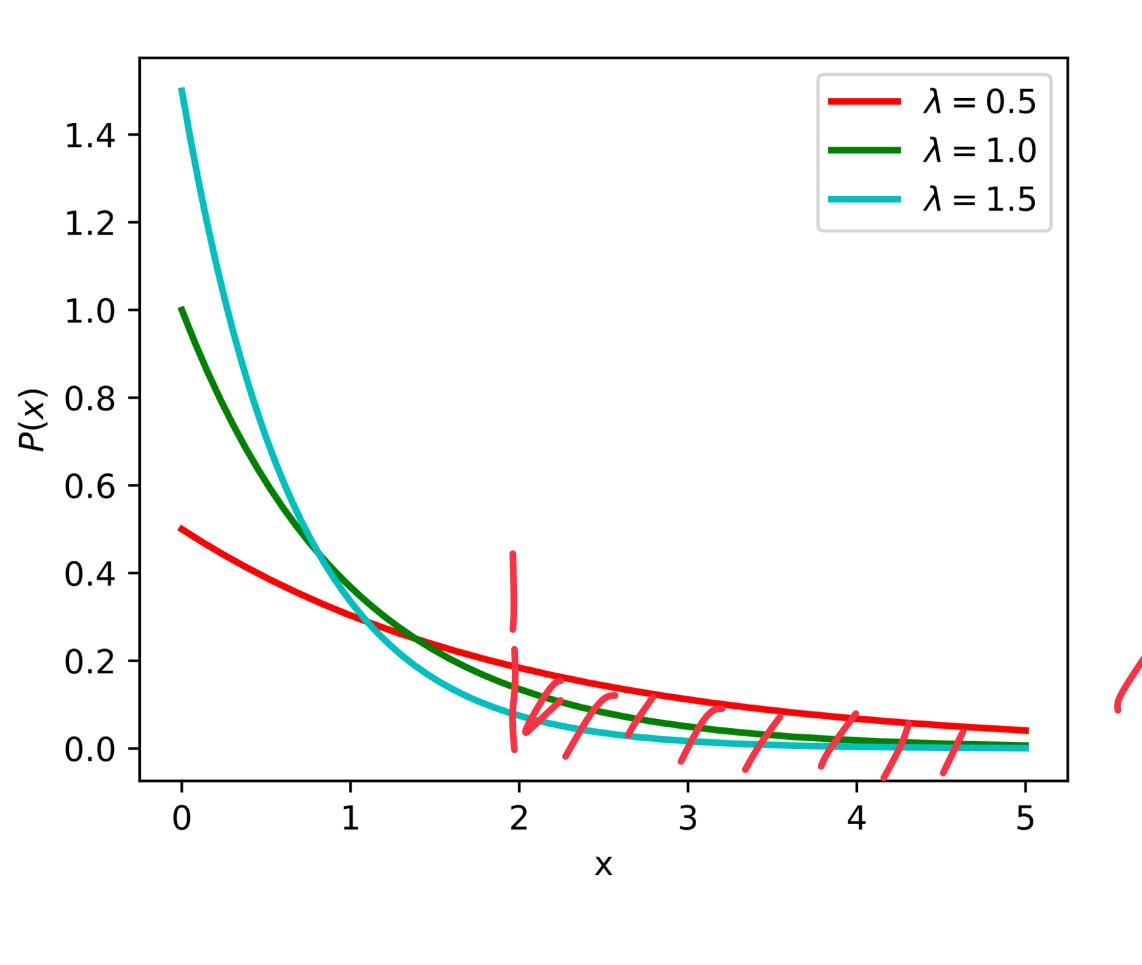
Compare my result to the expected distribution of Ho

It it lies for from expectation we can integrate beyond this to get P-value

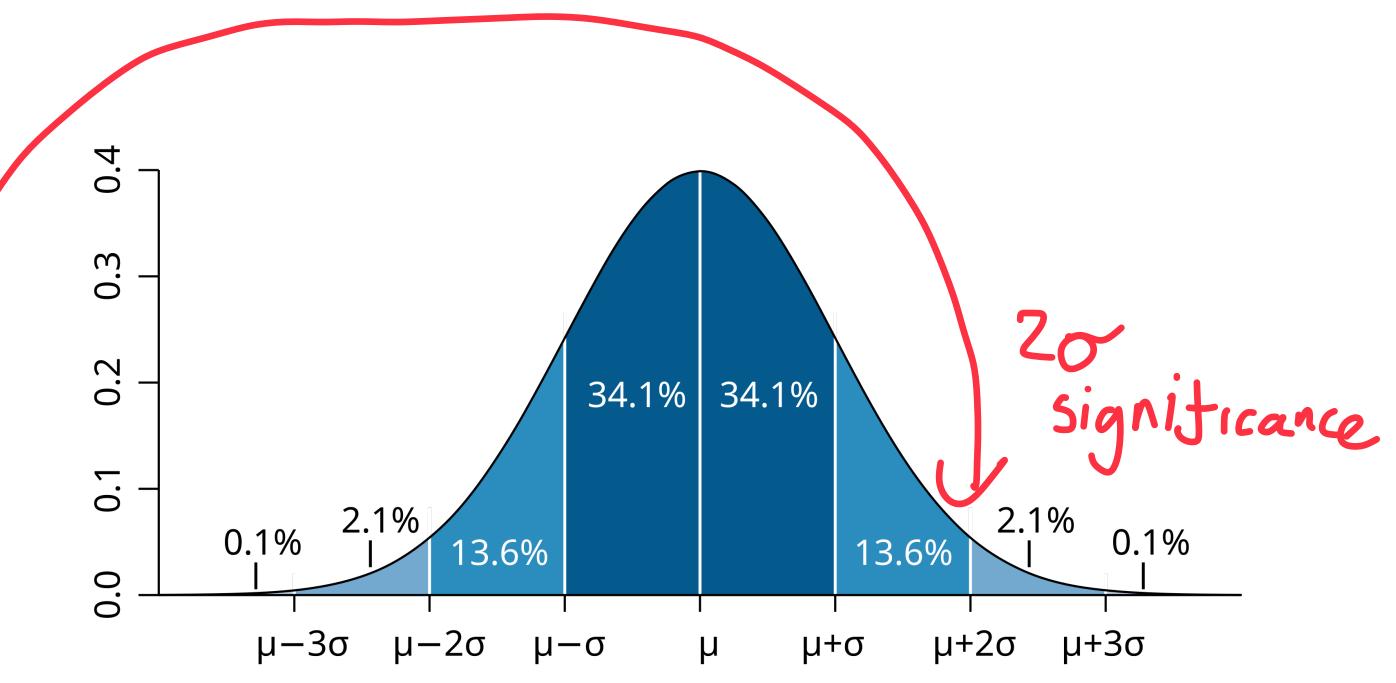


Can be one-sided or two-sided depending on the question I ask





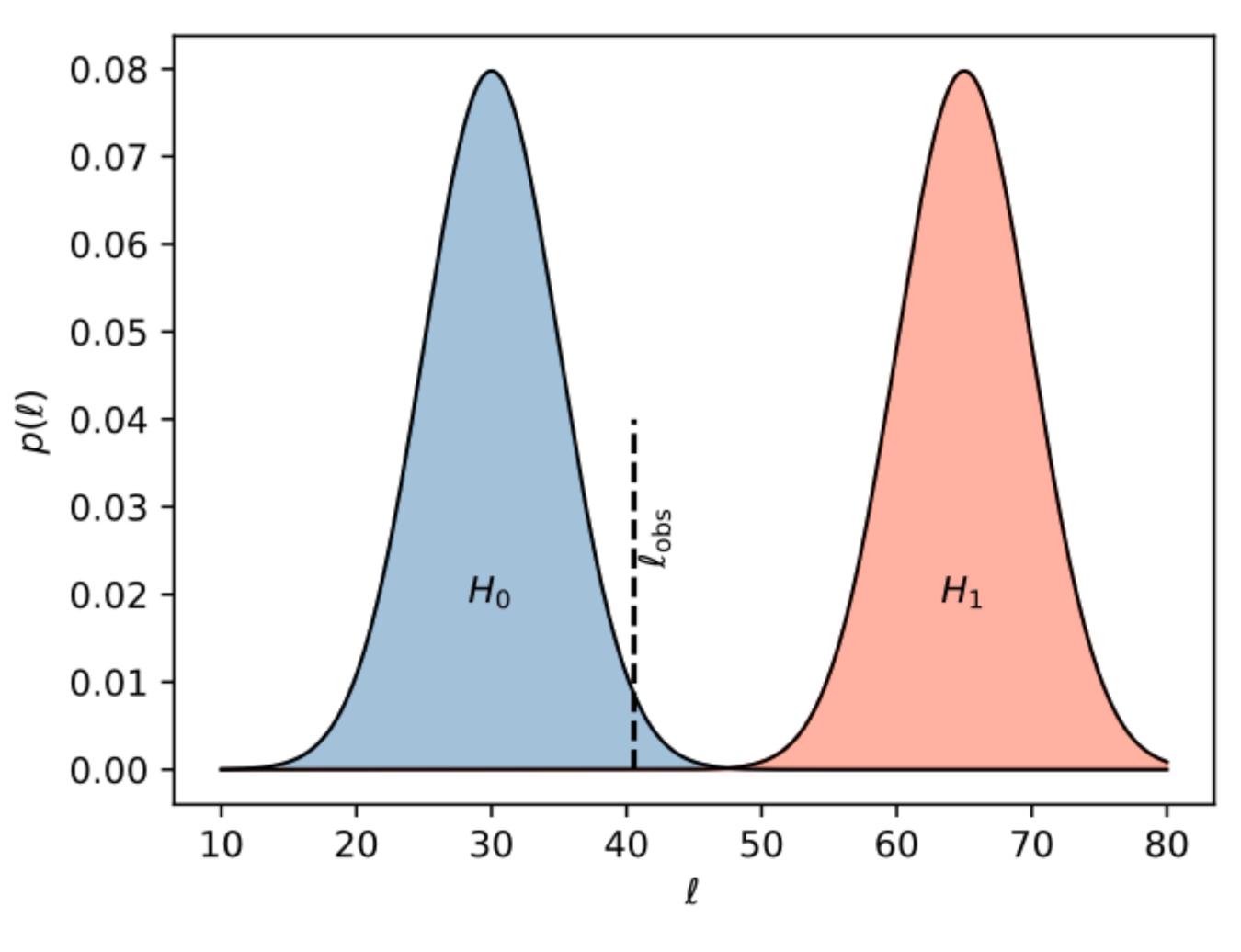
P-values not just for gaussians
We can calculate for any distribution
and convert to equivalent gaussian
number of "sigmas"

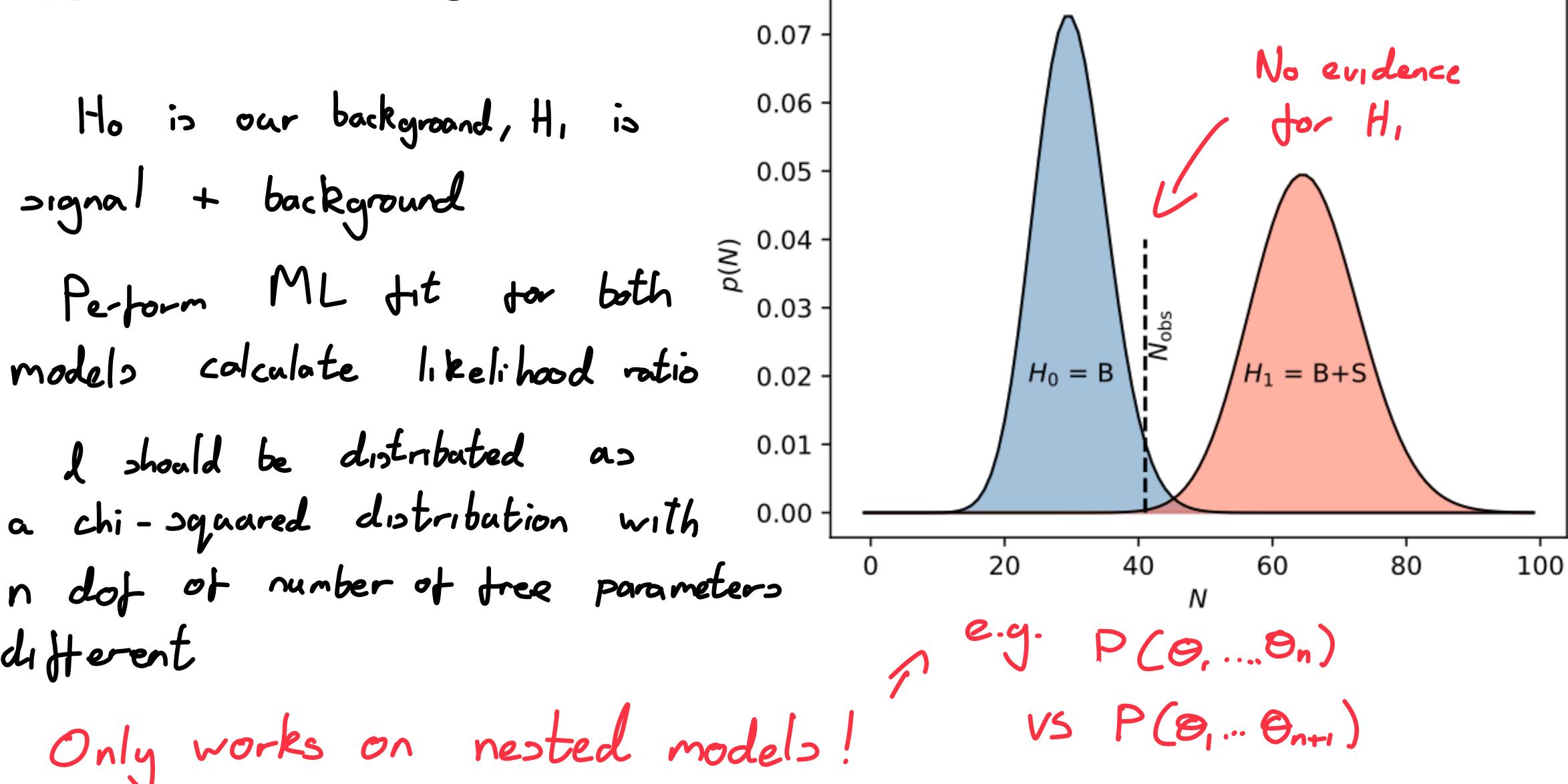


It's often more powerful to compare one hypothesis to another

or Hi?

Liklihood ratio test

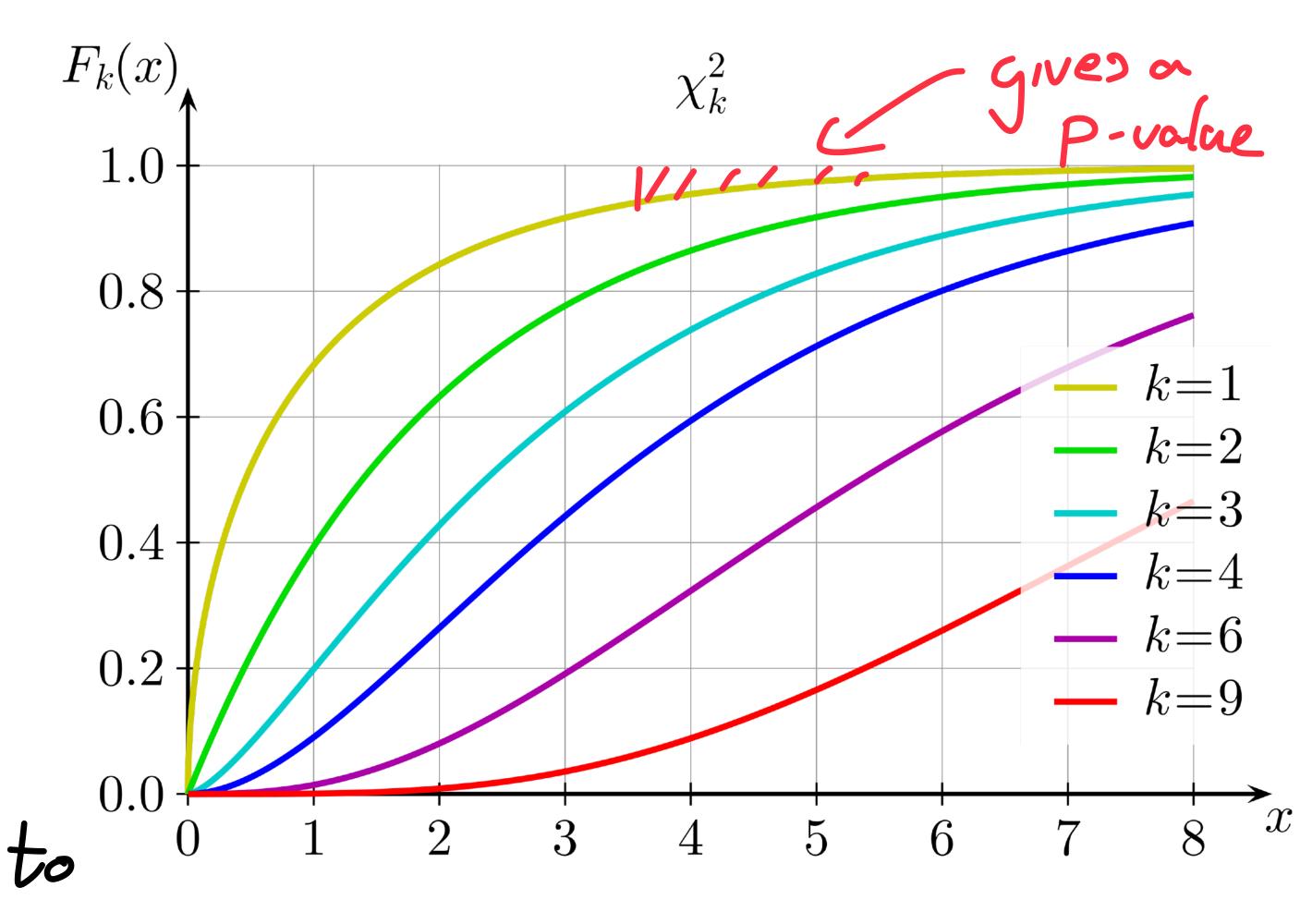


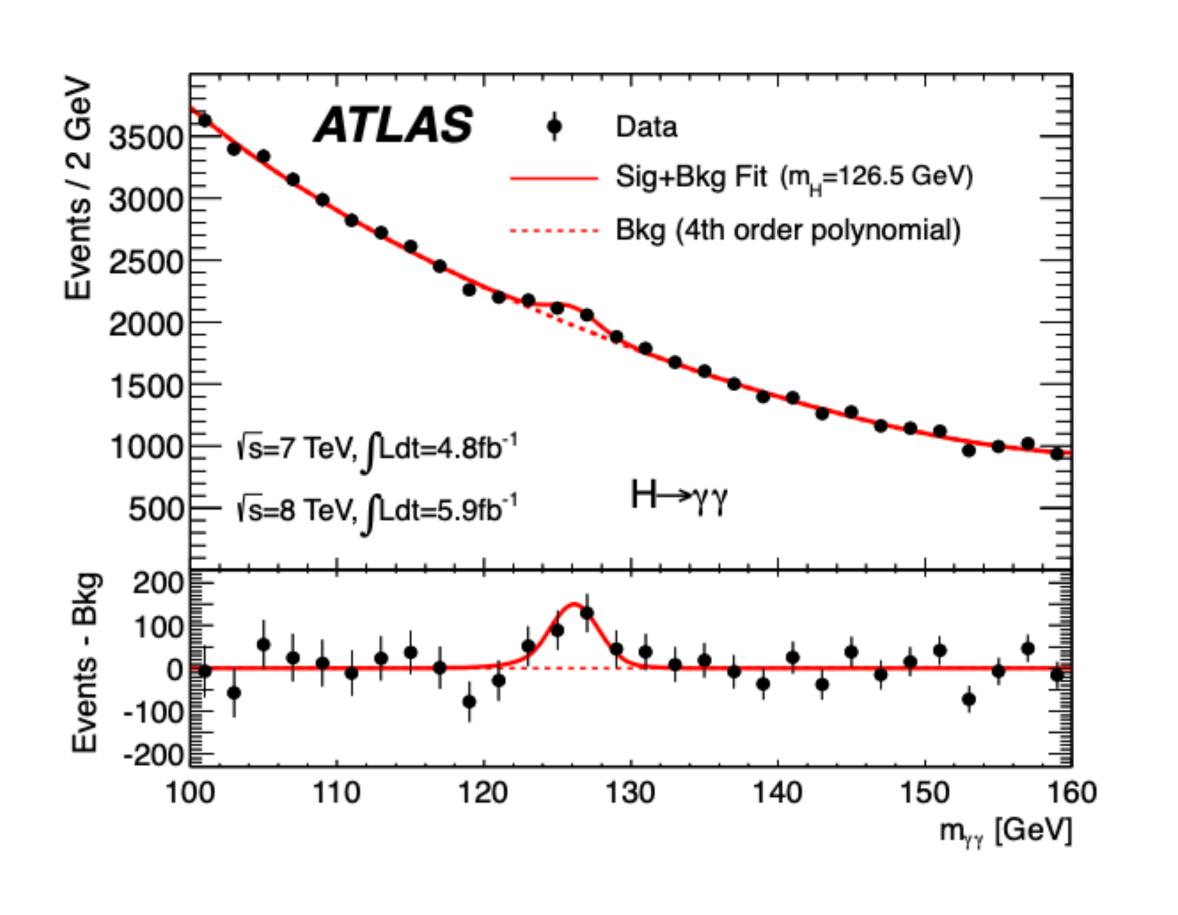


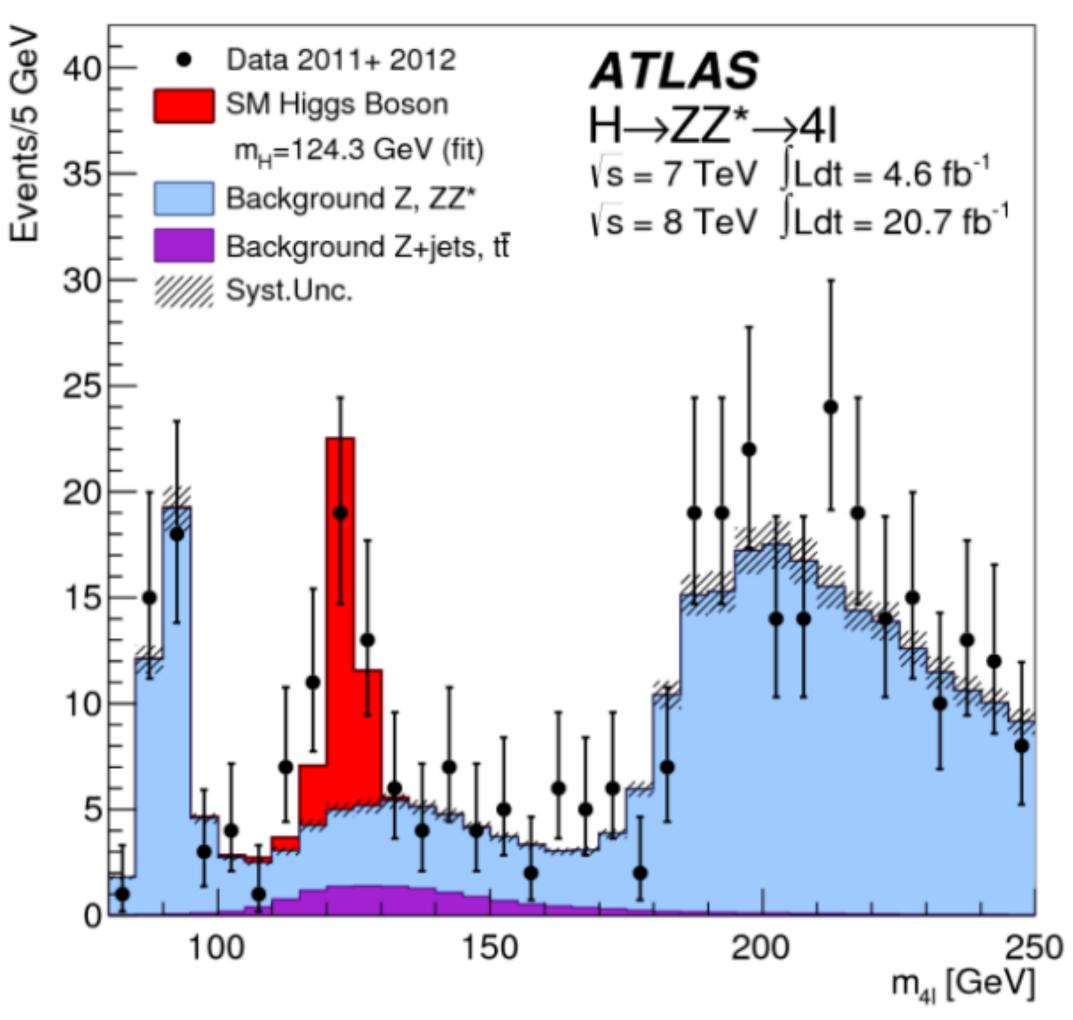
$$\int_{-2}^{2} \ln \left(\frac{L_{H_{\bullet}}}{L_{H_{\bullet}}} \right)$$

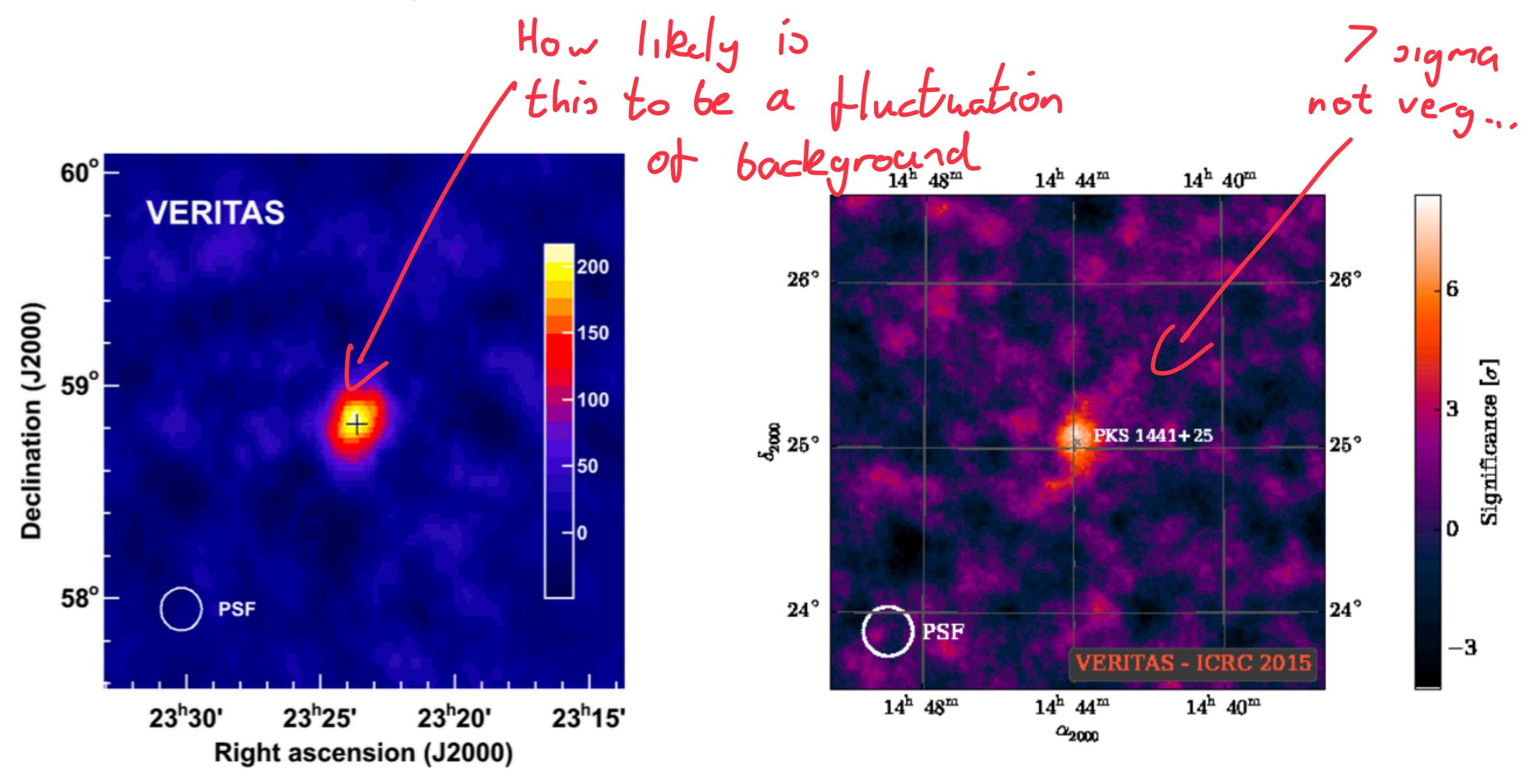
Distributed as χ^2 , K difference in tree parameters

P-value can be converted to gaussian or







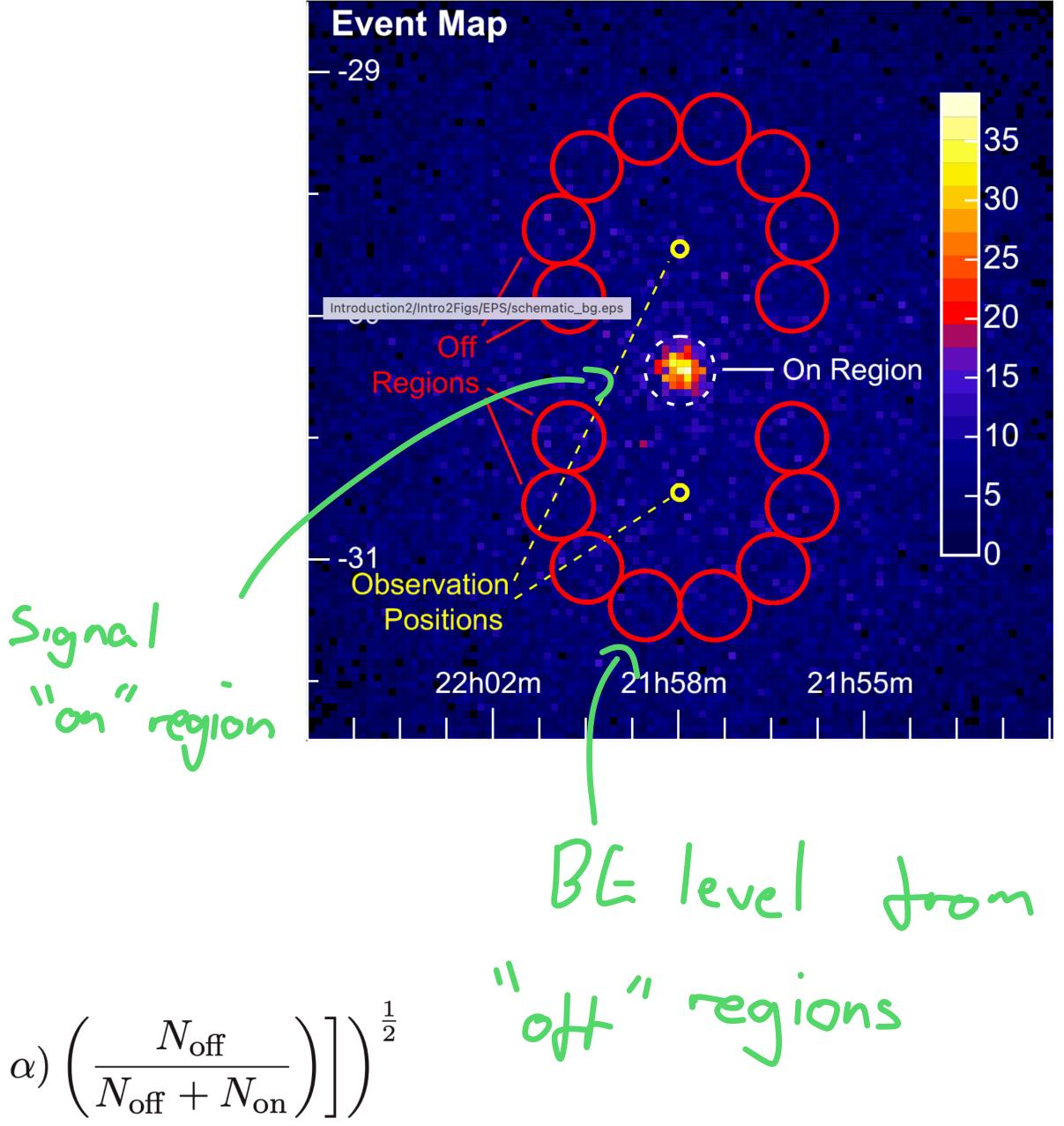


Can estimate source significance by Li and Ma Jornala

Likelihood ratio test of

BG-only hypothesis VS

BG + Signal



$$S = \sqrt{2} \left(N_{\text{on}} \ln \left[\frac{1 + \alpha}{\alpha} \left(\frac{N_{\text{on}}}{N_{\text{off}} + N_{\text{on}}} \right) \right] + N_{\text{off}} \ln \left[(1 + \alpha) \left(\frac{N_{\text{off}}}{N_{\text{off}} + N_{\text{on}}} \right) \right] \right)^{\frac{1}{2}}$$

Look elsewhere effect (trials factor)

HESS galactic 3
Plane scan (660) aprilipe 0
Test for a (72) 1 Test dor a source at every pixel in map Highly unlikely fluctuations 36 34 32 Galactic Longitude (deg)

10^S trials?

Need to modyg

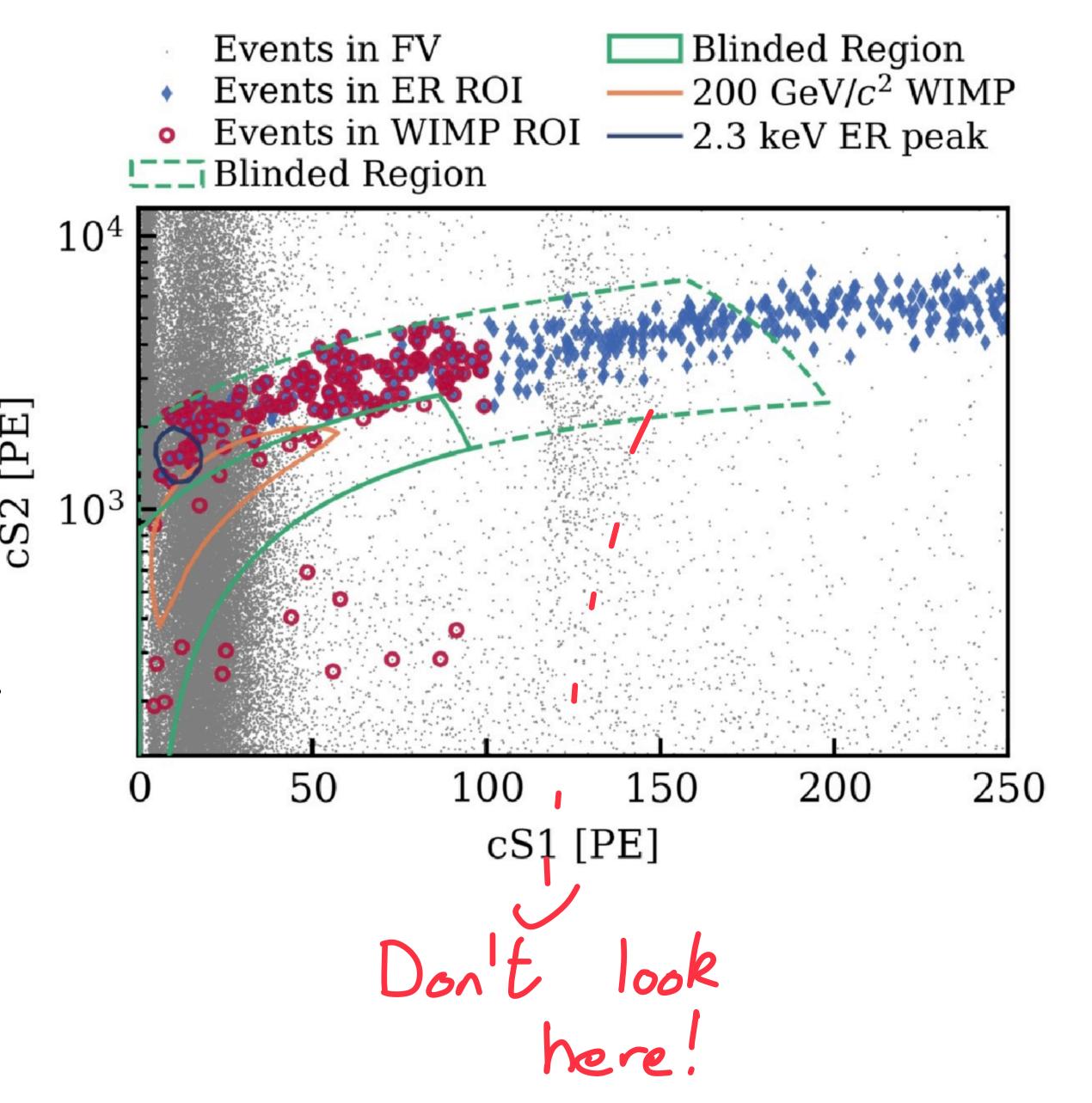
Significance

Blinding of analysis

In order to reduce the number of trials and increase confidence data can be blinded"

Do not allow ourselves to observe expected signal region until analysis is tixed

Could optimise on small fraction of data That data must then be discarded!



Why 5 Sigma?

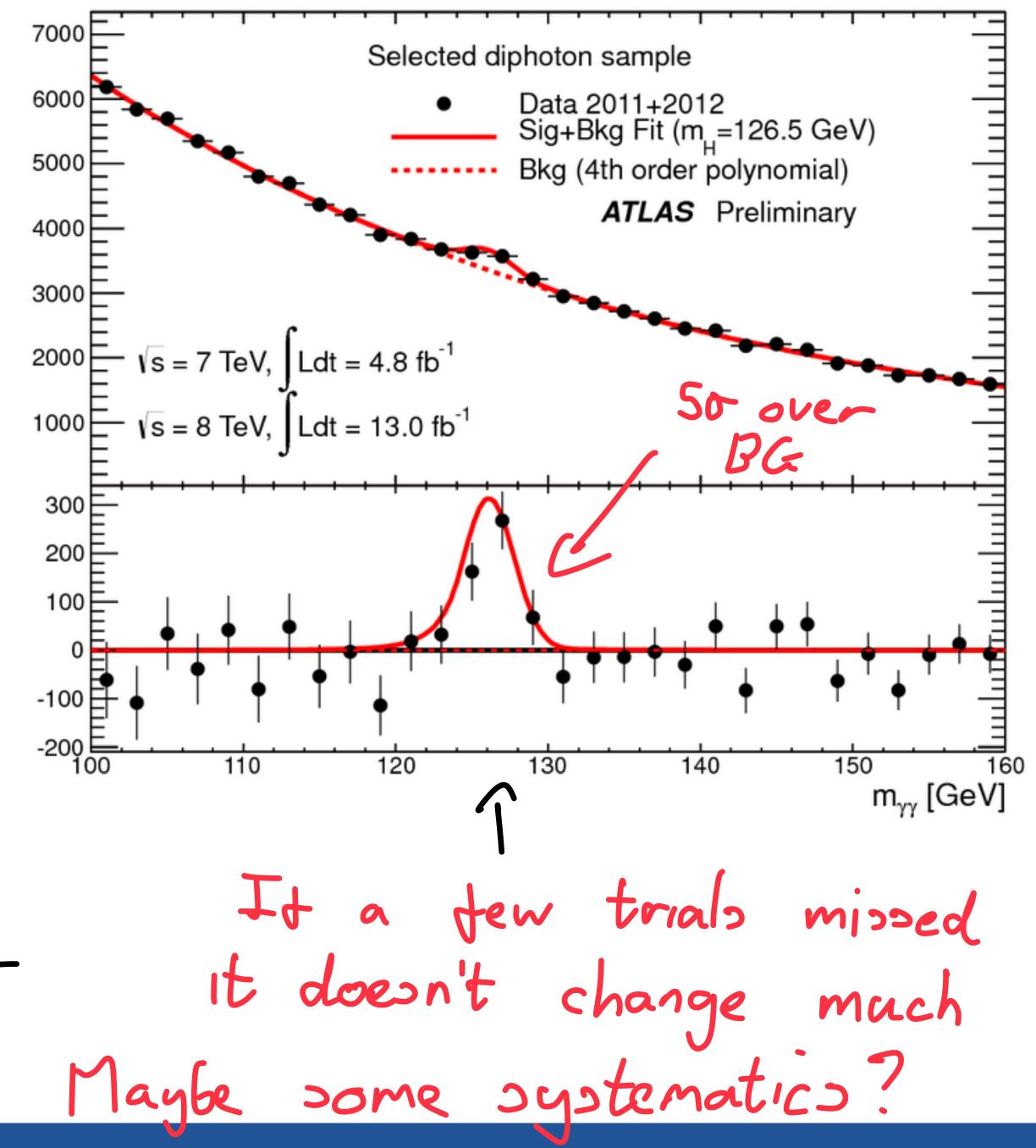
In the world of particle physics (+ astroparticle) 5 or is the "gold standard"

P value < 1e-6, very unlikely

NOT universal

- Astronomy 30
- Brology ~ 20 etc

It its new physics we better be certain



Monte Carlo Techniques

Monte Carlo methods

Algorithm that relies on repeated random sampling Often used with complex non-linear Systems Numerical integration Partrele interactions + decay Uncertainty Estimation

```
from random import random
from math import sqrt, pi
inside=0
n=10000000
i_print = [1, 10, 100, 1000, 10000, 100000, 1000000]
for i in range(0,n+1):
    x=random()
    y=random()
    if sqrt(x*x+y*y)<=1:
        inside+=1
    if i in i_print:
        piNow=4*inside/i
        print ('pi(i=%d) = %.4f, error = %.4f' % (i, piNow, abs(piNow - pi)))</pre>
```

```
pi(i=1) = 4.0000, error = 0.8584
pi(i=10) = 3.6000, error = 0.4584
pi(i=100) = 3.3600, error = 0.2184
pi(i=1000) = 3.1240, error = 0.0176
pi(i=10000) = 3.1264, error = 0.0152
pi(i=100000) = 3.1433, error = 0.0017
pi(i=1000000) = 3.1402, error = 0.0014
```

Monte Carlo integration

Standard computation -> split into sub-shapes + sum area Trapezium rule, Noolc = nde dimensions

Can instead randomly sample functions
Integration error $\propto 1/\sqrt{N}$ faster at

 $f(x) = a \qquad x_1 \qquad x_2 \qquad x_3 \qquad x_4 = b$

