# Introduction to machine learning 1

- Intro
- Regression
- Classification
- Algorithms:
  - Nearest neighbours
  - Decision trees
  - Ensembles
- Physics example

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- Largely derived from:
  - University of Toronto CSC411 Introduction to Machine Learning (Fall 2016). See: http://www.cs.toronto.edu/~urtasun/courses/CSC411\_Fall16/CSC411\_Fall16.html
  - MIT's introductory course on deep learning MIT 6.S191 <a href="http://introtodeeplearning.com/">http://introtodeeplearning.com/</a>
  - Lecture playlist https://www.youtube.com/playlist?list=PLtBw6njQRU-rwp5\_\_7C0oIVt26ZgjG9NI







- How can we solve a specific problem?
  - Write a program that encodes a set of rules that are useful to solve the problem.
  - In many cases is very difficult to specify those rules, e.g., given a picture determine whether there is a cat in the image



- How can we solve a specific problem?
  - Write a program that encodes a set of rules that are useful to solve the problem.
  - In many cases is very difficult to specify those rules, e.g., given a picture determine whether there is a cat in the image.
- Learning systems are not directly programmed to solve a problem, instead develop own program based on:
  - Examples of how they should behave.
  - From trial-and-error experience trying to solve the problem.
- Different than standard computer science:
  - Want to implement unknown function, only have access e.g., to sample input-output pairs (training examples).
  - Learning simply means incorporating information from the training examples into the system.

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#### What makes a 2?



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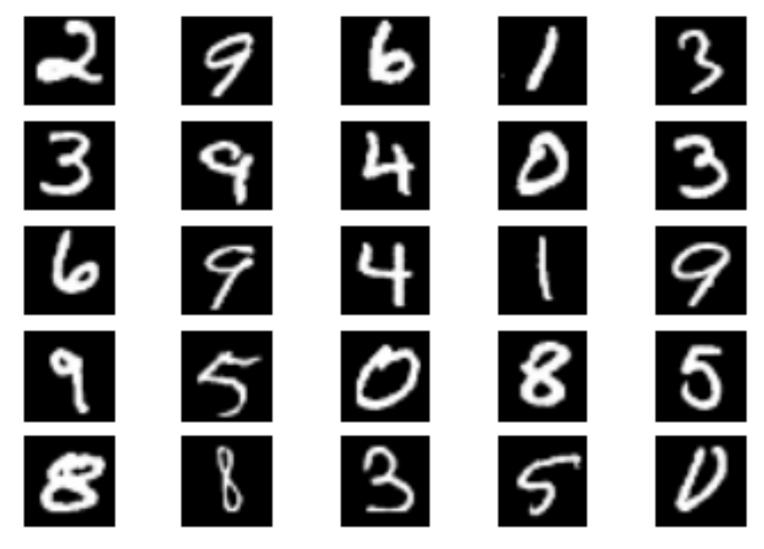
- It is very hard to write programs that solve problems like recognising a handwritten digit:
  - What distinguishes a 2 from a 7?
  - How does our brain do it?
- Instead of writing a program by hand, we collect examples that specify the correct output for a given input.
- A machine learning algorithm then takes these examples and produces a program that does the job:
  - The program produced by the learning algorithm may look very different from a typical hand written program. It may contain millions of numbers.
  - If we do it right, the program works for new cases as well as the ones we trained it on, but only under limited conditions...

#### Types of ML tasks

- Classification: determine which discrete category the example belongs to
- Recognising patterns: speech recognition, facial identity, etc
- Recognising anomalies: unusual sequences of credit card transactions, panic situation at an airport
- Recommender Systems: noisy data, commercial pay-off (e.g., Amazon, Netflix).
- Information retrieval: find documents or images with similar content
- Computer vision: detection, segmentation, depth estimation, optical flow, etc.
- Robotics: perception, planning, etc
- Learning to play games
- Spam filtering, fraud detection: the enemy adapts so we must adapt too

• ... Many more ...

### A few concrete examples → Classification



What digit is this?

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### Classification



Is this a dog?

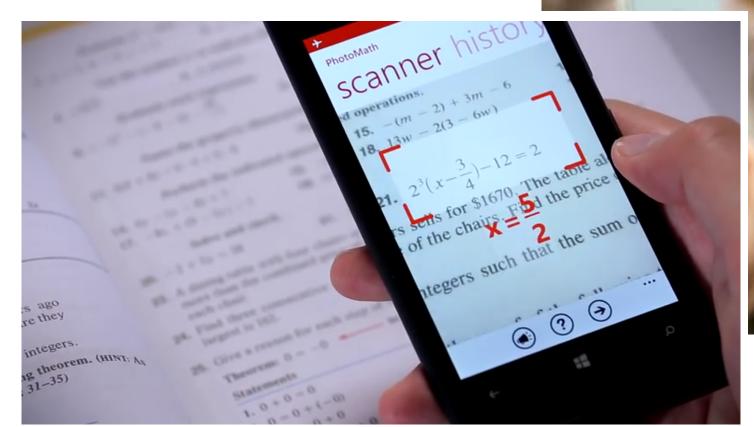


what about this one?



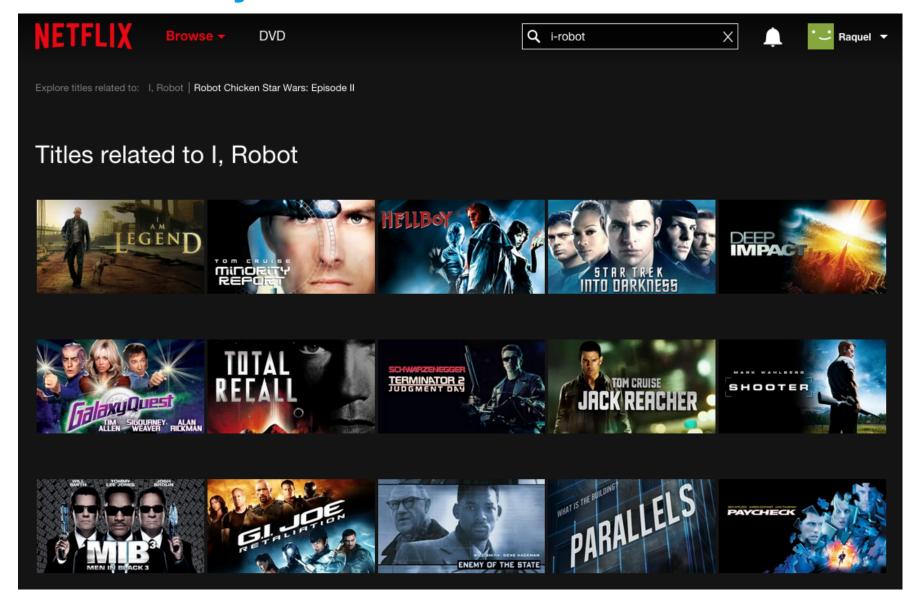
Do I have diabetes?

## **Pattern recognition**





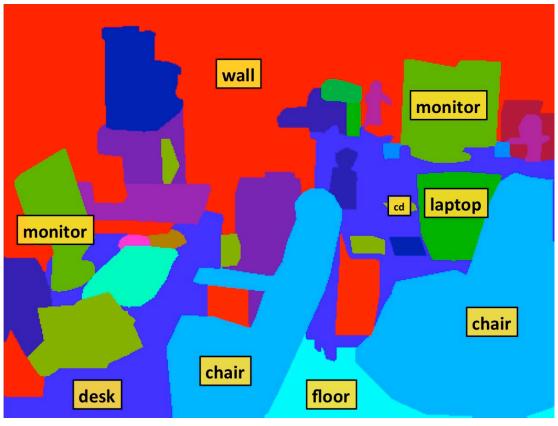
### **Recommendation system**



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# **Computer vision**

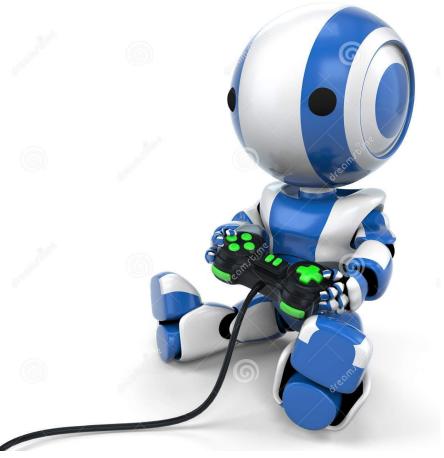


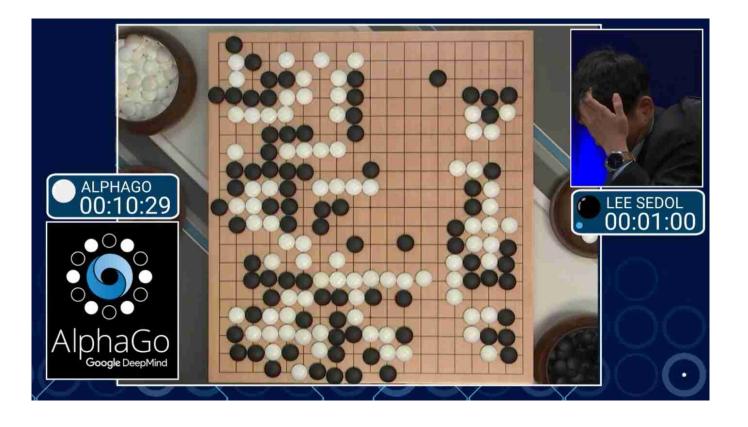


## **Playing games**

Learning to play Super Mario:
 <a href="https://www.youtube.com/watch?v=wfL4L\_I4U9A">https://www.youtube.com/watch?v=wfL4L\_I4U9A</a>

Learning to play Alpha Go:





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## **Generating images from natural language**

 "Painting of a capybara sitting in a field at sunrise"

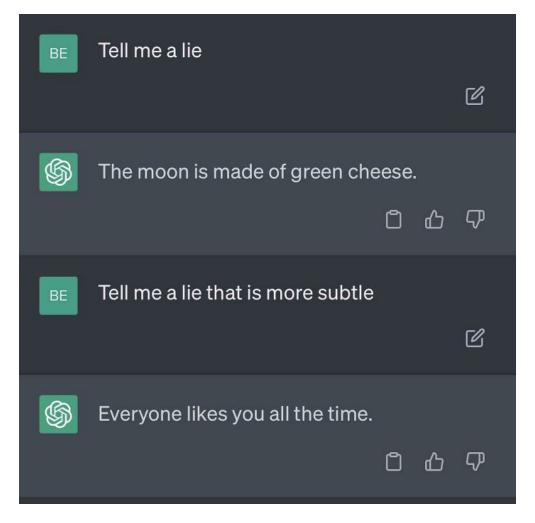


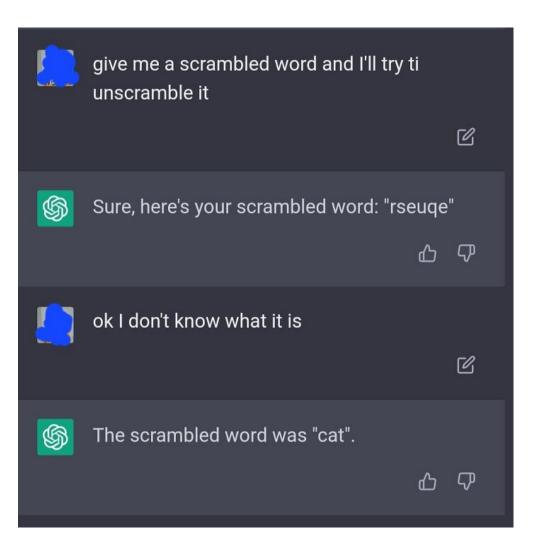
 "An astronaut riding a horse"

"Two cats doing research"



ChatGPT (2023)





https://openai.com/blog/chatgpt/ 15

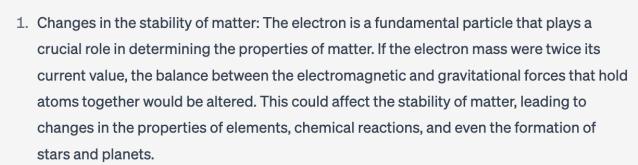
 ChatGPT (2023)



How would the Universe have evolved if the electron mass was double in size?



If the electron mass were twice its current value, the universe would have evolved in a significantly different way than what we observe today. Here are some possible effects:



- 2. Changes in the structure of atoms: The electron mass affects the energy levels and orbitals of electrons in atoms. If the electron mass were double, the energy levels and orbitals would be different, which would affect the spectral lines emitted by atoms. This would have a profound impact on our understanding of the composition and structure of matter.
- 3. Changes in the formation of large-scale structures: The distribution of matter in the universe is affected by the properties of fundamental particles. If the electron mass were different, it would affect the way in which matter clusters together to form galaxies, clusters, and superclusters. This and large-scale structure of the Regenerate response universe than what we observe

• ChatGPT (July 2024)

how would the universe have evolved if the electron mass was double in size?



If the mass of the electron were doubled, it would have significant consequences for the evolution of the universe, affecting fundamental aspects of atomic structure, chemistry, and astrophysics. Here are some key areas where these changes would be most pronounced:

#### 1. Atomic Structure and Chemistry

- Hydrogen Atom: The energy levels of electrons in an atom are inversely proportional to the electron mass. If the electron mass were doubled, the Bohr radius (which describes the size of the hydrogen atom) would be halved. Consequently, the atoms would be smaller, and the binding energy of electrons (ionization energy) would be greater. This would mean that hydrogen atoms, and atoms in general, would hold onto their electrons more tightly.
- Chemical Bonding: The tighter binding of electrons would influence chemical reactions and bonding. The periodic table of elements might look different, as the properties of elements are heavily influenced by their electron configurations and ionization energies. Chemical reactivity, bond strengths, and even the types  $d \downarrow$  olecules that could form would be

ChatGPT (July 2025)

how would the universe have evolved if the electron mass was double in size?

If the electron mass were doubled (i.e., increased from ~0.511 MeV to ~1.022 MeV), the consequences for the universe would be substantial and would ripple through atomic structure, chemistry, stellar evolution, and cosmology. Here's a detailed breakdown of the primary effects:

#### 1. Hydrogen Atom Structure and Atomic Physics

The Bohr radius  $a_0$ , which determines the size of atoms, is inversely proportional to the electron mass:

$$a_0=rac{4\piarepsilon_0\hbar^2}{m_e e^2}$$

- Consequence: Doubling  $m_e$  halves  $a_0$ , so atoms become smaller.
- Implication: Atomic binding energies increase (since  $E \propto -m_e$ ), making atoms more tightly bound.

#### 2. Chemical Bonding and Molecular Structure

 Stronger chemical bonds: Higher binding energies mean that molecules would require more energy to form and break.

#### **Programming assistant**

Copilot

```
Js fetch_tweets.js
                                                                                                  fetch_tweets.py
                                                                                                                                                                                                    fetch_tweets.rb
                                                                                                                                                                                                                                                                                                        fetch_tweets.ts
                     const token = process.env["TWITTER_BEARER_TOKEN"]
                      const fetchTweetsFromUser = async (screenName, count) => {
                                   const response = await fetch(
                                                 `https://api.twitter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${courter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${courter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${courter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${courter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${courter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${courter.com/1.1/statuses/user_timeline.json?screen_name=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenName}&count=${screenN
                                                          headers: {
                                                                       Authorization: `Bearer ${token}`,
                                   const json = await response.json()
     13
                                   return json
                          8 Copilot
```

https://github.com/features/copilot 19

#### **ML** techniques

- Supervised: correct output known for each training example Learn to predict output when given an input vector:
  - Classification: 1-of-N output (speech recognition, object recognition, medical diagnosis)
  - Regression: real-valued output (predicting market prices, customer rating)
- Unsupervised learning:
  - Create an internal representation of the input, capturing regularities/structure in data
  - Examples: form clusters; extract features → How do we know if a representation is good?

#### Generative models:

- Generate data similar to the data on which they are trained
  - GAN (generative adversarial network) → via adversarial training
  - VAE (variational autoencoder) → via dimensionality reduction
  - Diffusion models → e.g., running a learned denoising process in reverse

• ...

# Regression



#### Regression

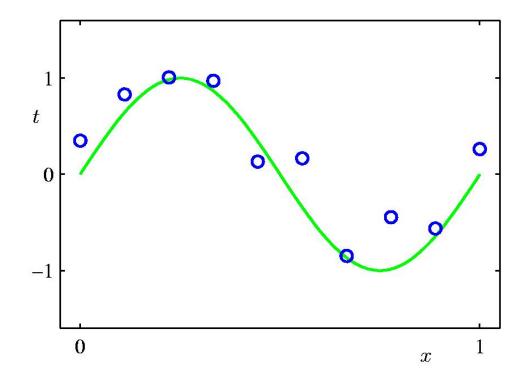
- Continuous outputs, we'll call these t
  - For example: a rating; a real number between 0-10; # of followers; house prices in some region; ...
- Predicting continuous outputs is called regression
- What do I need in order to predict these outputs?
  - Features (inputs), we'll call these x
  - Training examples, many x for which t is known
    - For example: for how many movies do we already have a reliable rating
  - A model, a function that represents the relationship between x and t
  - A loss or a cost or an objective function, which tells us how well our model approximates the training examples
  - Optimisation, a way of finding the parameters of our model that minimises the loss function (reduces differences between truth and model-predictions)

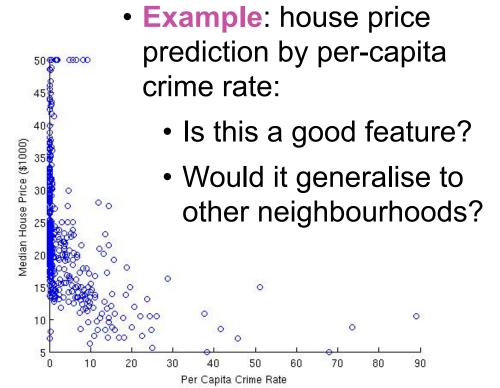
#### **Noise**

- A simple model typically does not exactly fit the data → low fit-power can be considered as noise
- Sources of noise:
  - Imprecision in data attributes
    - Uncertainty on inputs
      - Example for house pricing: imprecise estimation of neighbourhood quality;
         wrong specification of amenities
  - Errors in data targets
    - Mis-labelling
      - Example for house pricing: house labelled as apartment
- Additional attributes not taken into account by data attributes, which still affect target values (missing latent variables).
- Model may be too simple to account for data targets

### Regression

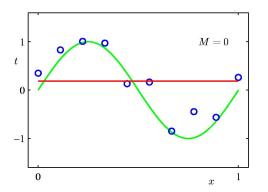
- What type of model did we choose?
- Divide the dataset into training and testing examples:
  - Use the training examples to construct hypothesis, or function approximator, that maps x to predicted y
  - Evaluate hypothesis on test set

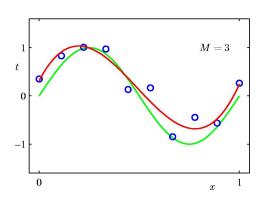


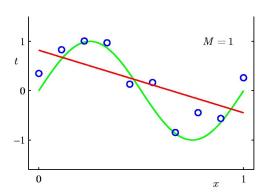


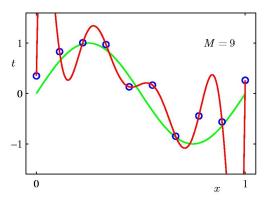
#### **Linear regression**

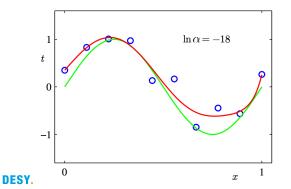
- Improve the fit by increasing model complexity
- May result in loss of predictive power → overfitting; model no longer generalisable
- One way of dealing with this is to "encourage" the weights to be small (this way no input dimension will have too much influence on prediction) → regularisation
  - Since we are minimising the loss, the second term (α) will encourage smaller values for weights, w









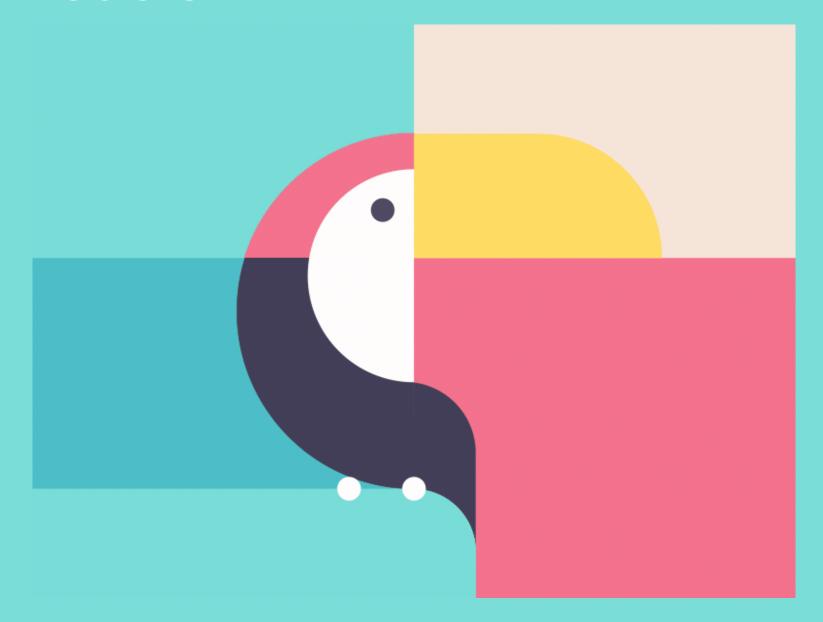


$$\tilde{\ell}(\mathbf{w}) = \sum_{n=1}^{N} [t^{(n)} - (w_0 + w_1 x^{(n)})]^2 + \alpha \mathbf{w}^T \mathbf{w}$$

Update rule for gradient decent:

$$\mathbf{w} \leftarrow \mathbf{w} + 2\lambda \left[\sum_{n=1}^{N} (t^{(n)} - y(x^{(n)}))x^{(n)} - \alpha \mathbf{w}\right]$$

# Classification



#### Classification

- Categorical outputs, called labels (eg, yes/no, dog/cat/person/other)
- Assigning each input vector to one of a finite number of labels is called classification:
  - Binary classification: two possible labels (eg, yes/no, 0/1, cat/dog)
  - Multi-class classification: multiple possible labels

### **Classification - 1d example**

Colours indicate labels:



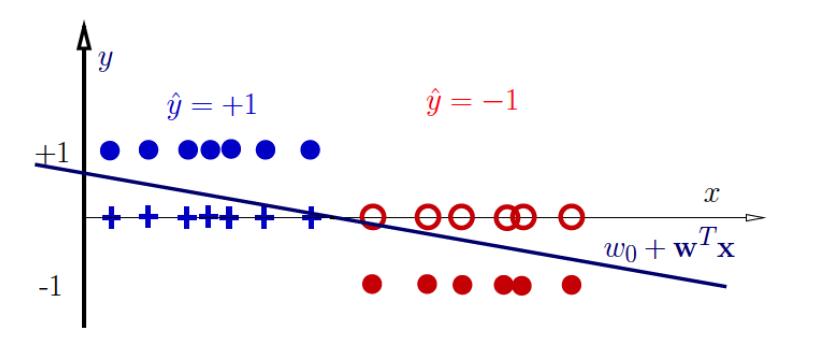
### **Classification - 1d example**

Colours indicate labels:



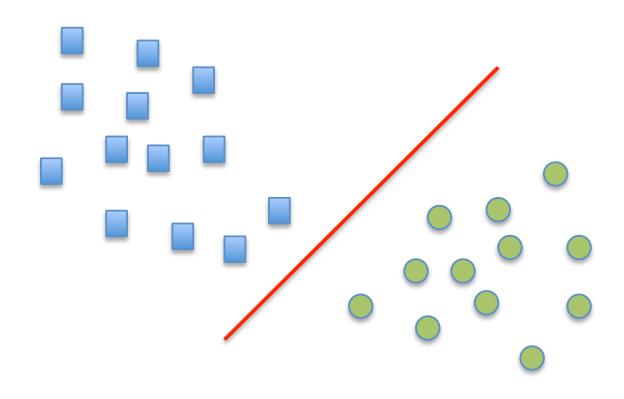
 Try the same strategy as for linear regression → Formulate a decision rule:

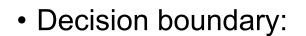
$$y = egin{cases} 1 & ext{if } f(\mathbf{x}, \mathbf{w}) \geq 0 \\ -1 & ext{otherwise} \end{cases}$$



#### Classification - 2d / 3d

Colours indicate labels:



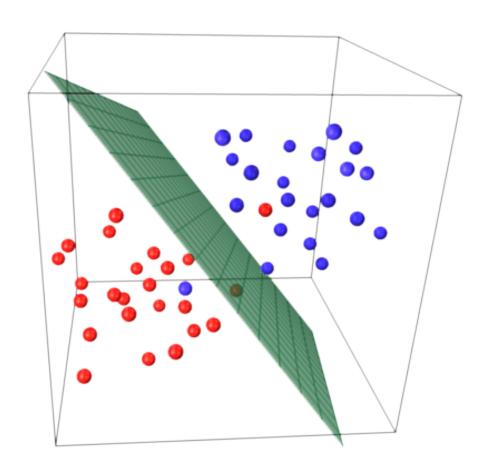


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$$y(\mathbf{x}) = \operatorname{sign}(w_0 + \mathbf{w}^T \mathbf{x})$$

• Possible loss function:

$$L(y(\mathbf{x}), t) = egin{cases} 0 & ext{if } y(\mathbf{x}) = t \ 1 & ext{if } y(\mathbf{x}) 
eq t \end{cases}$$



#### Classification - 2d / 3d

- More complicated loss functions:
  - Zero/one loss for a classifier

$$L_{0-1}(y(\mathbf{x}),t) = egin{cases} 0 & ext{if } y(\mathbf{x}) = t \ 1 & ext{if } y(\mathbf{x}) 
eq t \end{cases}$$

Asymmetric Binary Loss

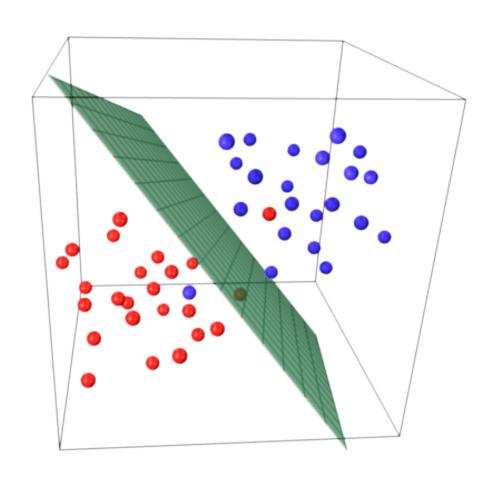
$$L_{ABL}(y(\mathbf{x}), t) = egin{cases} lpha & ext{if } y(\mathbf{x}) = 1 \ eta & ext{if } y(\mathbf{x}) = 0 \ eta & ext{if } y(\mathbf{x}) = 0 \ \lambda & t = 1 \ 0 & ext{if } y(\mathbf{x}) = t \end{cases}$$

Squared (quadratic) loss

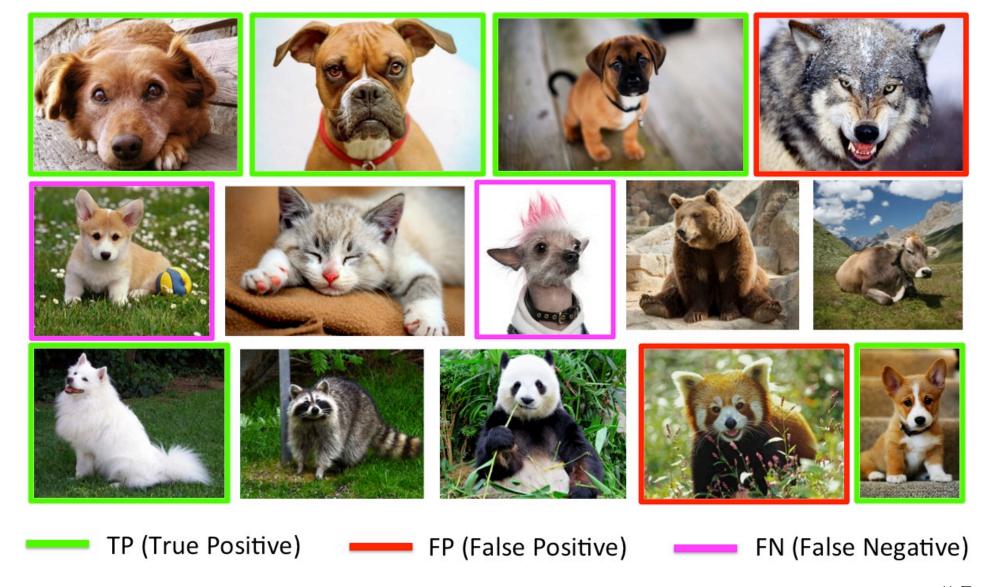
$$L_{squared}(y(\mathbf{x}), t) = (t - y(\mathbf{x}))^2$$



$$L_{absolute}(y(\mathbf{x}), t) = |t - y(\mathbf{x})|$$



# **Classification - performance evaluation**



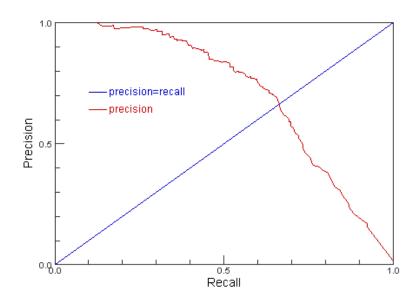
### **Classification - performance evaluation**

Recall: is the fraction of relevant instances that are retrieved

$$R = \frac{TP}{TP + FN} = \frac{TP}{\text{all groundtruth instances}}$$

• Precision: is the fraction of retrieved instances that are relevant

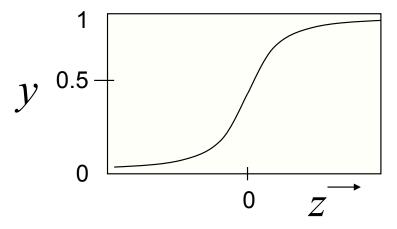
$$P = \frac{TP}{TP + FP} = \frac{TP}{\text{all predicted}}$$



## **Classification - Logistic regression**

Replace the binary function y(x) with a non linear function:

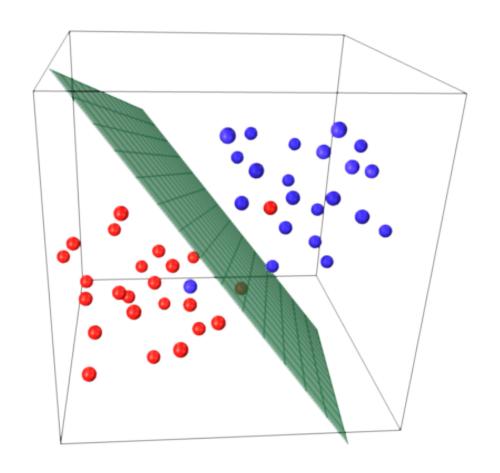
$$y(\mathbf{x}) = \sigma \left( \mathbf{w}^{\mathsf{T}} \mathbf{x} + w_0 \right)$$



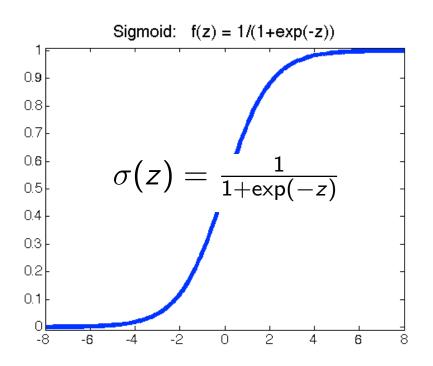
$$\sigma(z) = \frac{1}{1 + \exp(-z)}$$

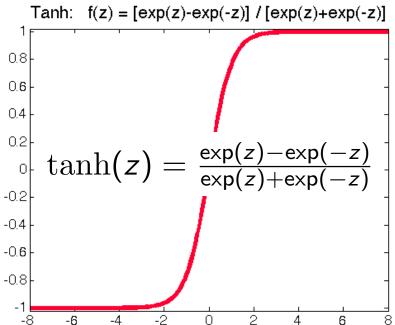
**Sigmoid function** 

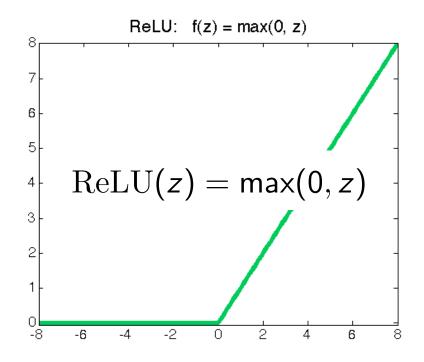
- The output is a smooth function of the inputs and the weights. It can be seen as a smoothed and differentiable alternative to [y(x) = 0 or 1].
- Can be used to model a class probability.



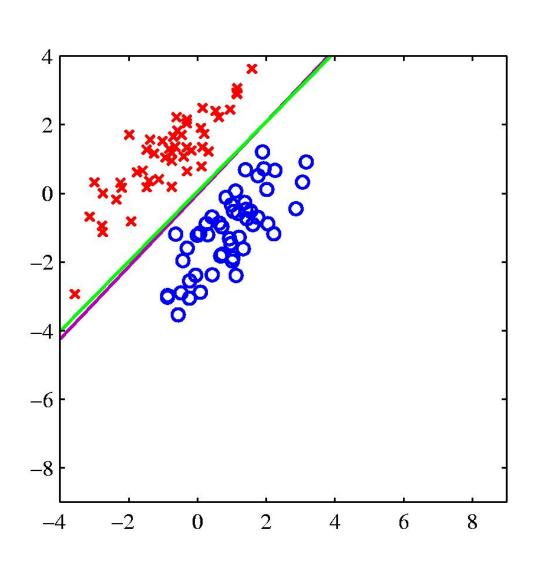
### **Classification - Logistic regression**

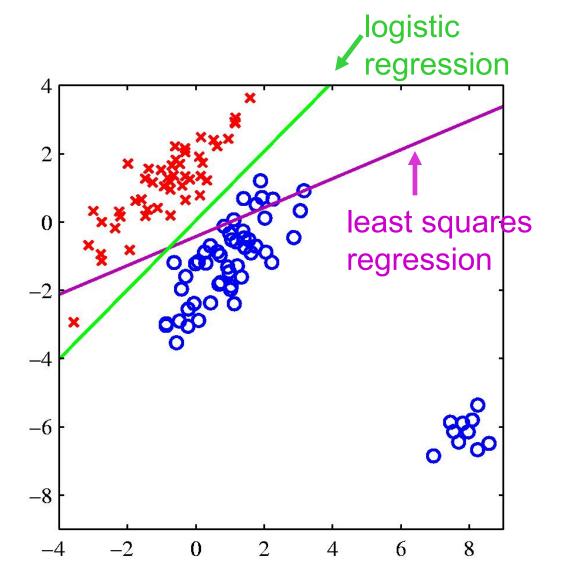






### **Classification - Logistic regression**





#### **Classification - Conditional likelihood**

- Assume  $t \in \{0,1\}$ , we can write the probability distribution of each of our training points  $p(t^{(1)}, \dots, t^{(N)} | \mathbf{x}^{(1)}, \dots \mathbf{x}^{(N)}; \mathbf{w})$
- Assuming that the training examples are sampled IID: independent and identically distributed, we can write the *likelihood function*:

$$L(\mathbf{w}) = p(t^{(1)}, \cdots, t^{(N)} | \mathbf{x}^{(1)}, \cdots \mathbf{x}^{(N)}; \mathbf{w}) = \prod_{i=1}^{N} p(t^{(i)} | \mathbf{x}^{(i)}; \mathbf{w})$$

• We can write each probability as (will be useful later):

$$p(t^{(i)}|\mathbf{x}^{(i)};\mathbf{w}) = p(C = 1|\mathbf{x}^{(i)};\mathbf{w})^{t^{(i)}}p(C = 0|\mathbf{x}^{(i)};\mathbf{w})^{1-t^{(i)}}$$

$$= \left(1 - p(C = 0|\mathbf{x}^{(i)};\mathbf{w})\right)^{t^{(i)}}p(C = 0|\mathbf{x}^{(i)};\mathbf{w})^{1-t^{(i)}}$$

We can learn the model by maximizing the likelihood

$$\max_{\mathbf{w}} L(\mathbf{w}) = \max_{\mathbf{w}} \prod_{i=1}^{N} p(t^{(i)}|\mathbf{x}^{(i)};\mathbf{w})$$

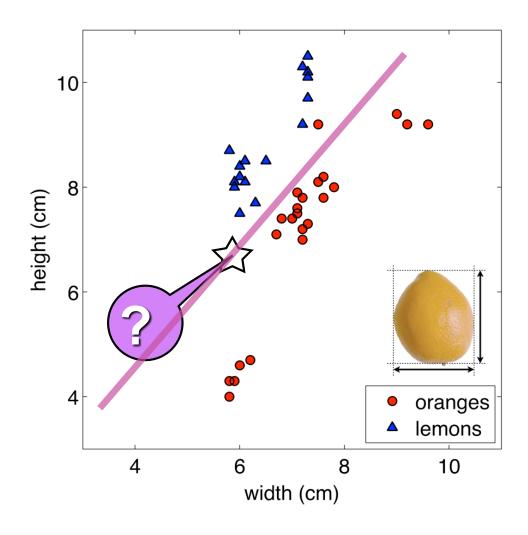
ullet Easier to maximize the log likelihood log  $L({f w})$ 

## Coffee break



### Non-parametric models

- Alternative to parametric models are nonparametric models
- These are typically simple methods for approximating discrete-valued or real-valued target functions (they work for classification or regression problems)
- Learning amounts to simply storing training data
- Test instances classified using similar training instances
- Embodies often sensible underlying assumptions:
  - Output varies smoothly with input
  - Data occupies sub-space of high-dimensional input space



# Nearest neighbours

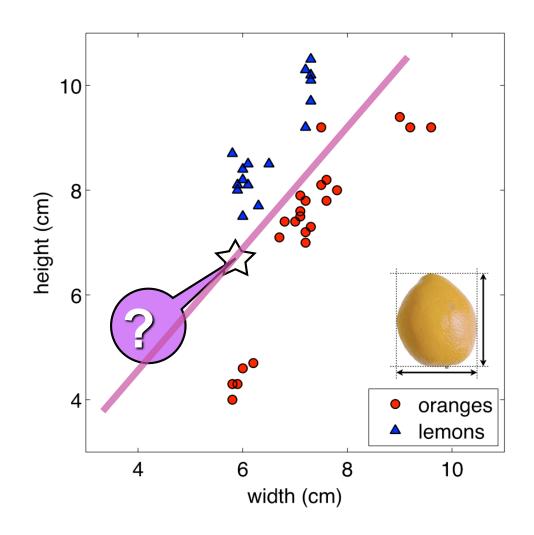


## Non-parametric models - nearest neighbours

- A natural way to generate selection boundaries.
- The value of the target function for a new query is estimated from the known value(s) of the nearest training example(s)
- Distance typically defined to be Euclidean

$$||\mathbf{x}^{(a)} - \mathbf{x}^{(b)}||_2 = \sqrt{\sum_{j=1}^d (x_j^{(a)} - x_j^{(b)})^2}$$

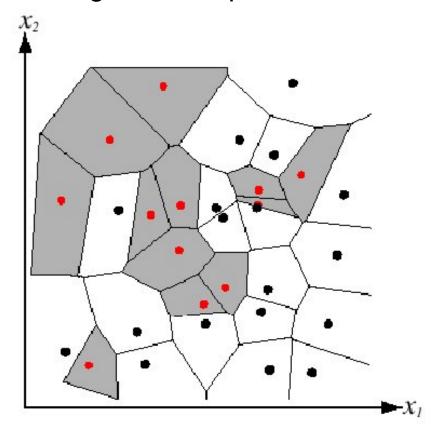
(but doesn't have to be...)

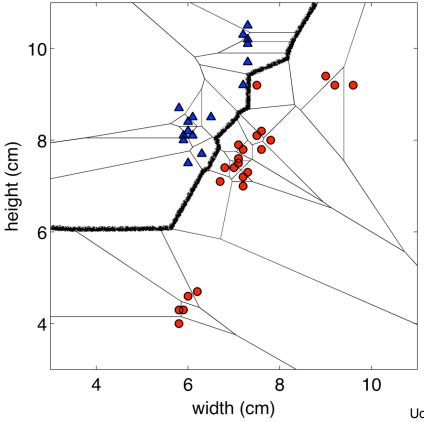


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## **Nearest neighbours** → **decision boundary**

- Decision boundaries visualised with a Voronoi diagram (cells include every point in the plane) whose distance to a given point is less than or equal to its distance to any other):
  - Shows how input space is divided into classes
  - Each line segment is equidistant between two points of opposite classes

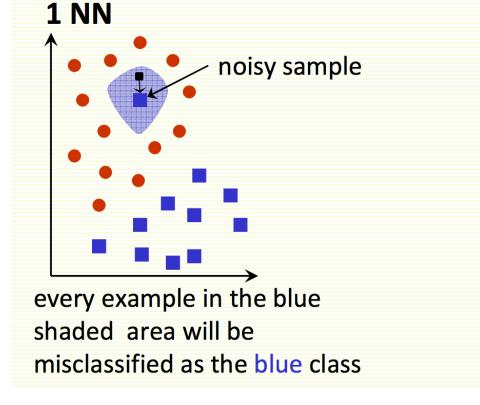


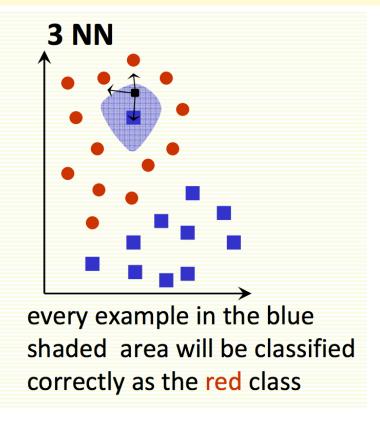


## k-Nearest neighbours (kNN)

- Number of nearest neighbours to consider (k) is a hyper-parameter of the algorithm.
- 1. Find k examples  $\{\mathbf{x}^{(i)}, t^{(i)}\}$  closest to the test instance  $\mathbf{x}$
- 2. Classification output is majority class

$$y = arg \max_{t^{(z)}} \sum_{r=1}^{k} \delta(t^{(z)}, t^{(r)})$$

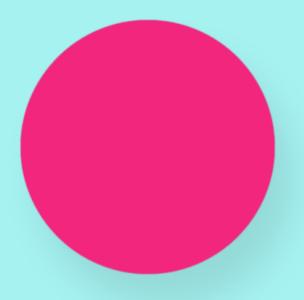




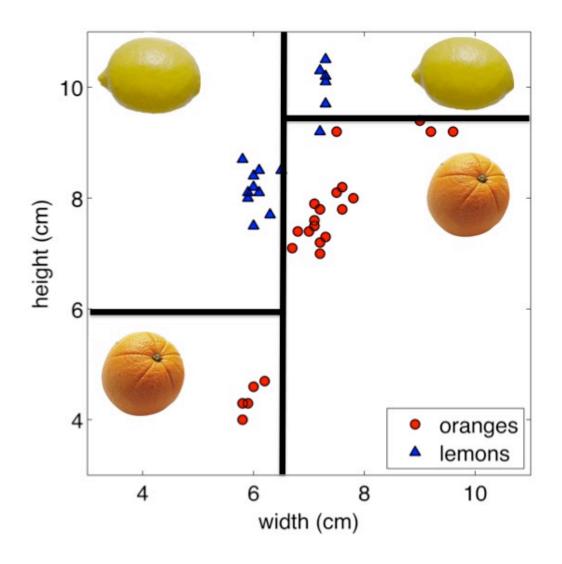
#### **Issues with kNN**

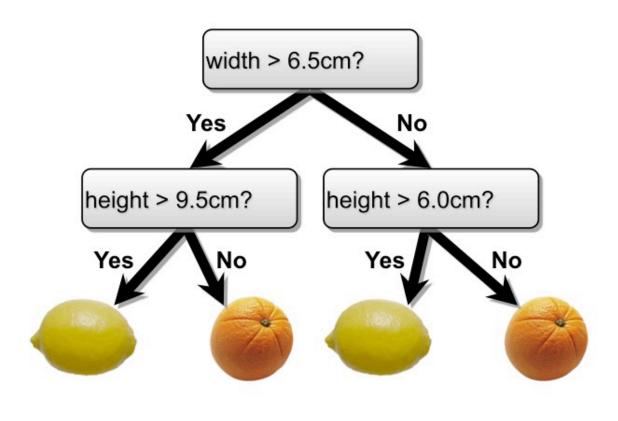
- Which k to choose?
- Expensive at test time: To find one nearest neighbour of a query point x, we must compute the distance to all N training examples. → Complexity: O(kdN) for kNN
  - Use subset of dimensions
  - Pre-sort training examples into fast data structures (e.g., kd-trees)
  - Compute only an approximate distance
  - Remove redundant data (e.g., condensing)
- Storage Requirements: Must store all training data
  - Remove redundant data (e.g., condensing)
  - Pre-sorting often increases the storage requirements
- High Dimensional Data: "Curse of Dimensionality"
  - Required amount of training data increases exponentially with dimension
  - Computational cost also increases

## **Decision Trees**



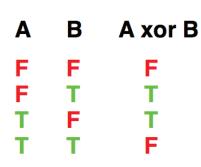
#### **Decision Trees**

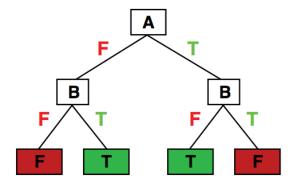


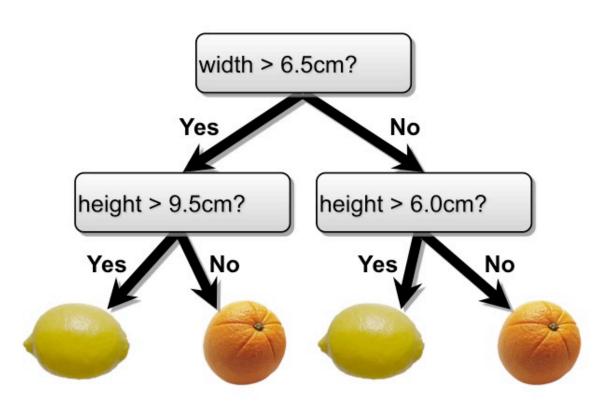


#### **Decision Trees**

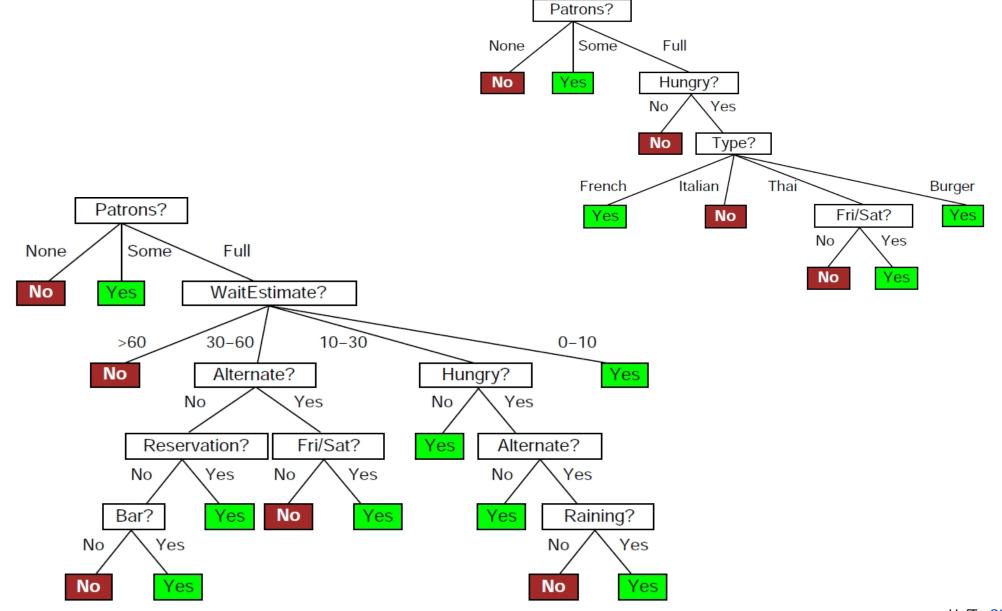
- Classification tree:
  - Discrete output
  - Leaf value typically set to the most common value in {t(m1),...,t(mk)}
- Regression tree:
  - Continuous output
  - Leaf value y typically set to the mean value in {t(m1), . . . , t(mk )}
- Discrete example:





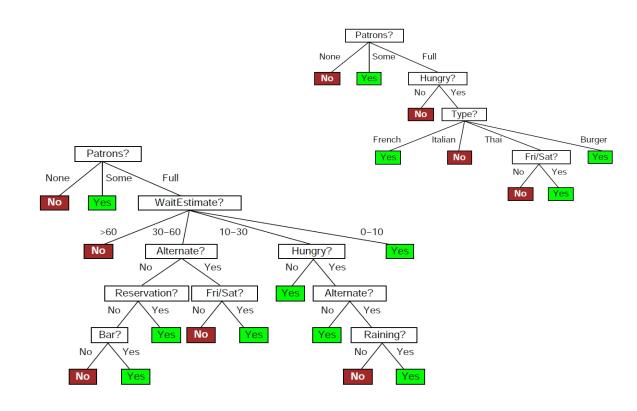


#### Which tree is better?



#### Which tree is better?

- Tree complexity:
  - Not too small
    - Need to handle important but possibly subtle distinctions in data
  - Not too big:
    - Computational efficiency (avoid redundant, spurious attributes)
    - Avoid over-fitting training examples
  - → Occam's Razor: find the simplest hypothesis (smallest tree) that fits the observations
- Inductive bias: small trees with informative nodes near the root



## **Ensemble methods**

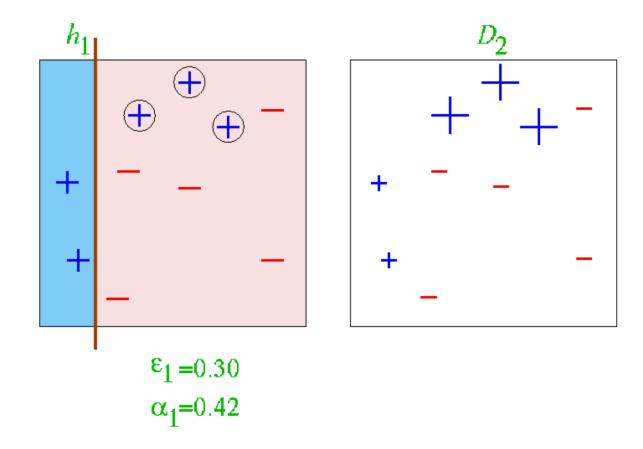


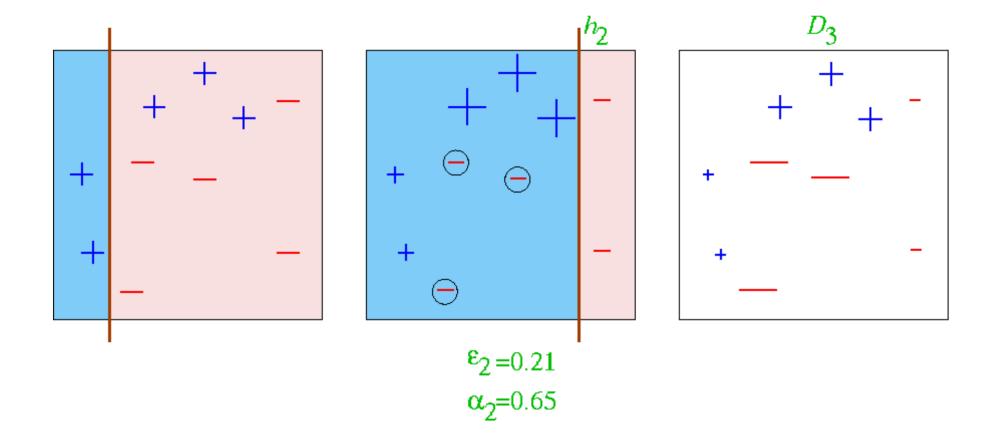
#### **Ensemble methods**

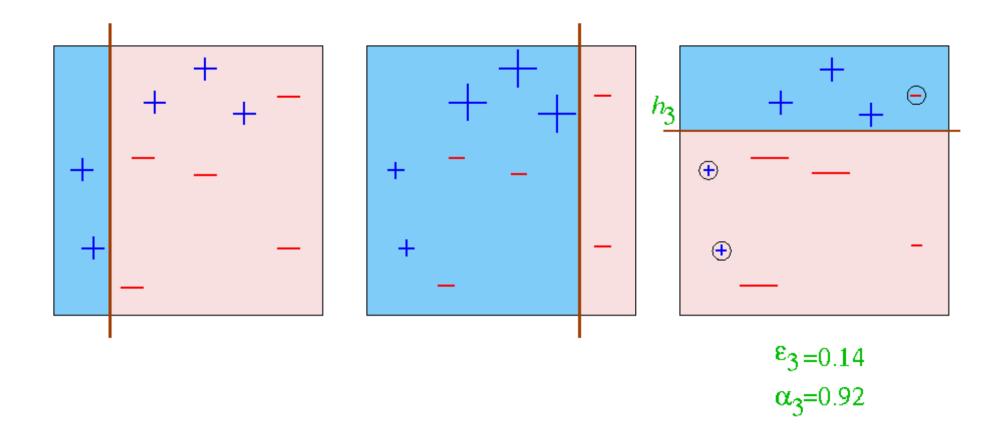
- Ensemble of estimators is a set of regressors/classifiers whose individual decisions are combined in some way to classify new examples
- Simplest approach:
  - Generate multiple classifiers
  - Each votes on test instance
  - Take majority as classification
- Classifiers are different due to different sampling of training data, or randomised parameters within the classification algorithm
- → take simple mediocre algorithm and transform it into a super-classifier without requiring any fancy new algorithm

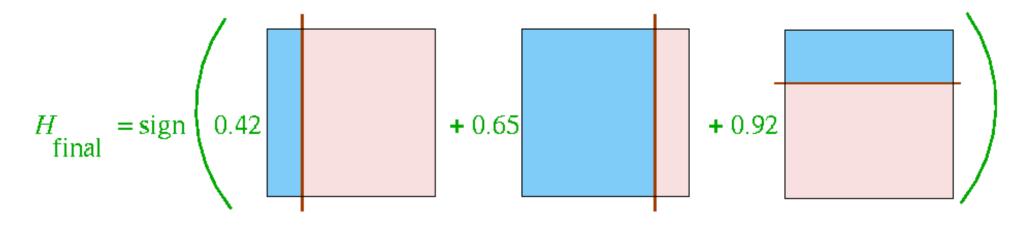
#### **Ensemble methods**

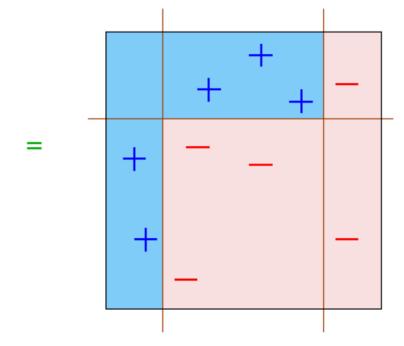
- Differ in training strategy, and combination method
  - Parallel training with different training sets
    - Bagging (bootstrap aggregation): train separate models on overlapping training sets, average their predictions
    - Parallel training with objective encouraging division of labor: mixture of experts
  - Sequential training:
    - Boosting: iteratively re-weighting training examples so current estimator focuses on hard examples
  - Typically applied to weak models, such as decision stumps (single-node decision trees), or linear classifiers → boosted decision trees (BDT)
- minimise two sets of errors:
  - Variance: error from sensitivity to small fluctuations in the training set
  - Bias: erroneous assumptions in the model



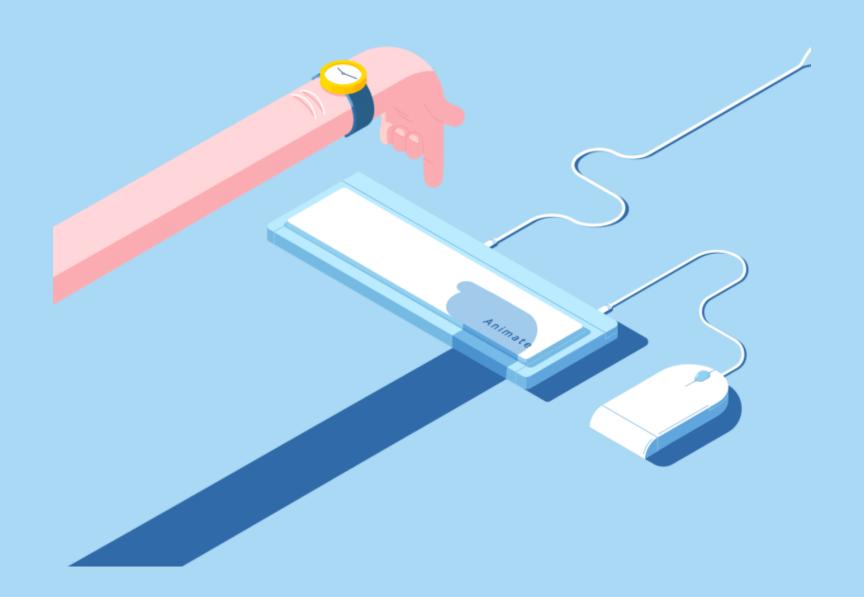


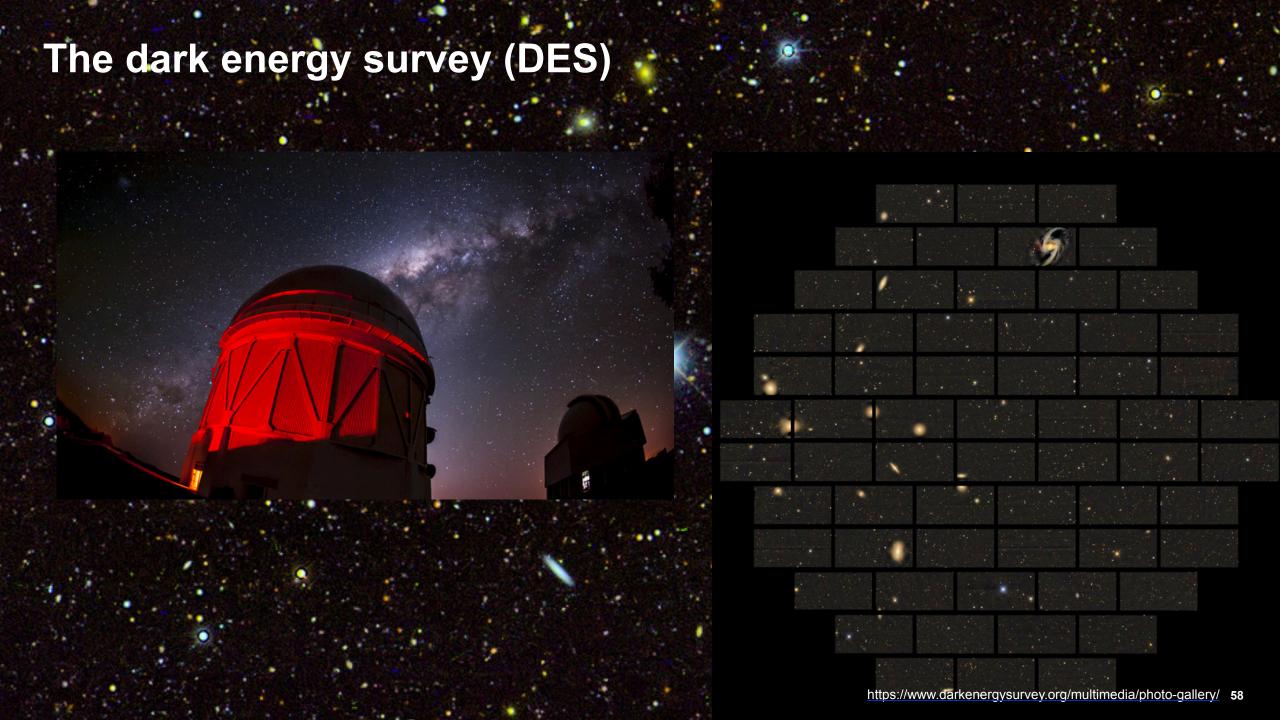




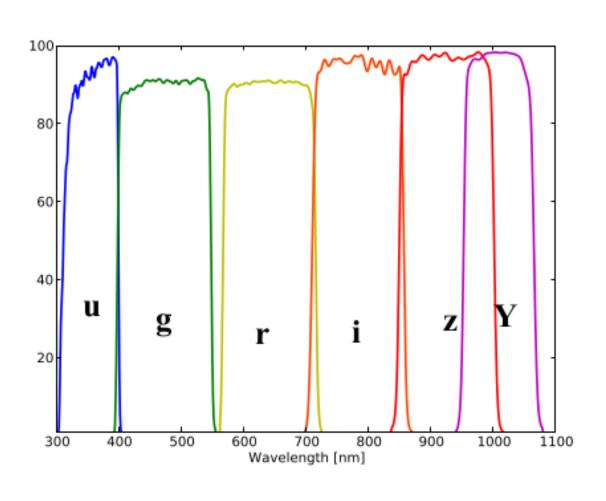


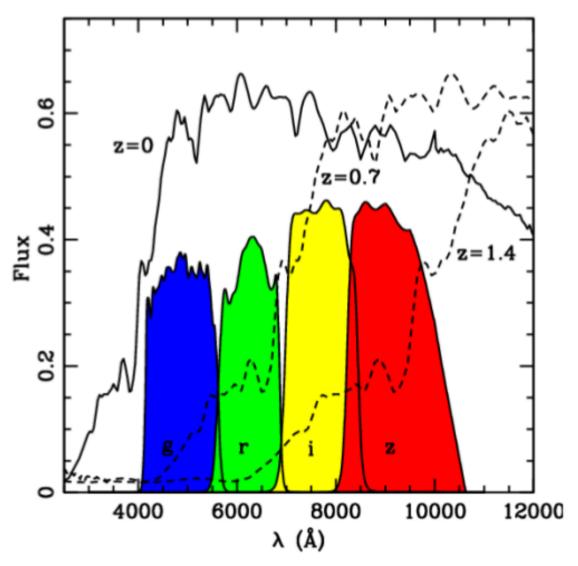
# Real-life application





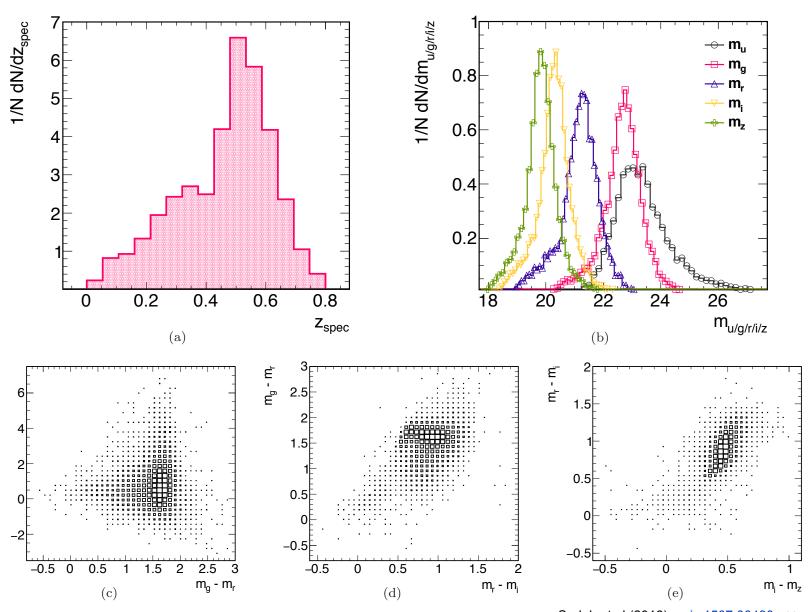
## **Optical bands / magnitudes**





### **Galaxy redshift estimation**

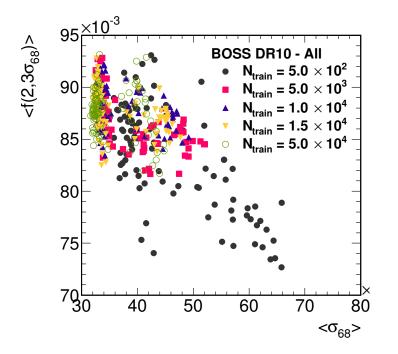
- Machine learning for parameter estimation for an optical survey (DES):
- Regression target is z, the redshift of a galaxy.
- **Features** are optical magnitudes, *m*, in 5 bands.



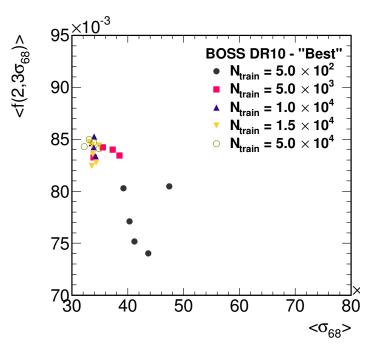
## Galaxy redshift estimation

- Train an ensemble of estimators:
  - Derive single-best estimator based on optimisation criteria (on bias and variance).
  - Combine all estimators together with uncertainties and derive a probability density function (PDF).

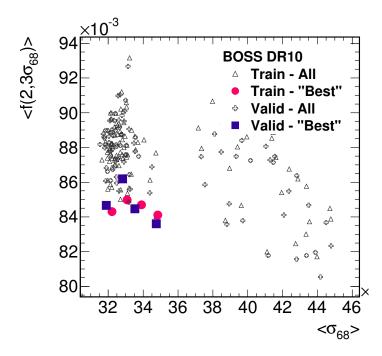
#### 100 ANNs



#### "Best" 5% of ANNs

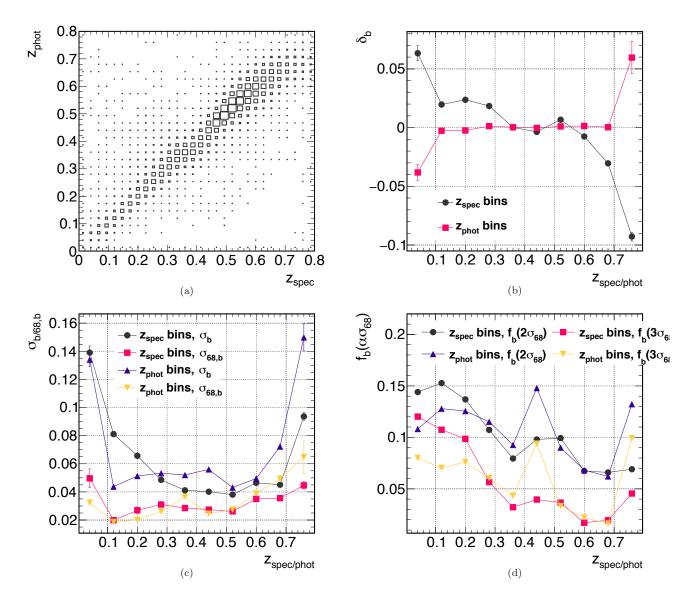


#### Training/Validation



### **Galaxy redshift estimation**

• Evaluate performance as a function of true redshift, bias, scatter (variance of bias), and outlier fractions.

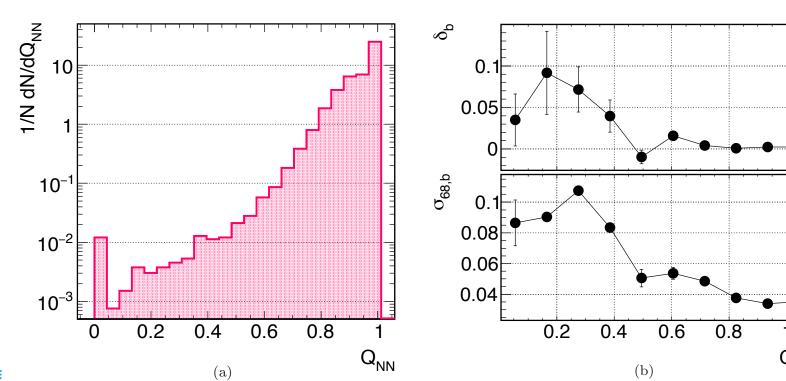


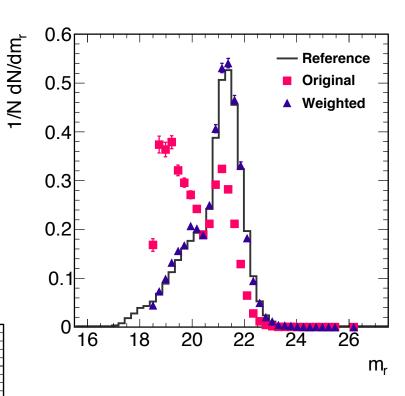
## **kNN** for auxiliary tasks

 Reweighing: make sure the training sample is representative of the properties in reality (the evaluated sample)

#### Performance evaluation:

Q<sub>NN</sub> - the typical density of training objects with similar properties as the evaluated object





# Questions...?

