

Neutrino physics overview

Julia Gehrlein

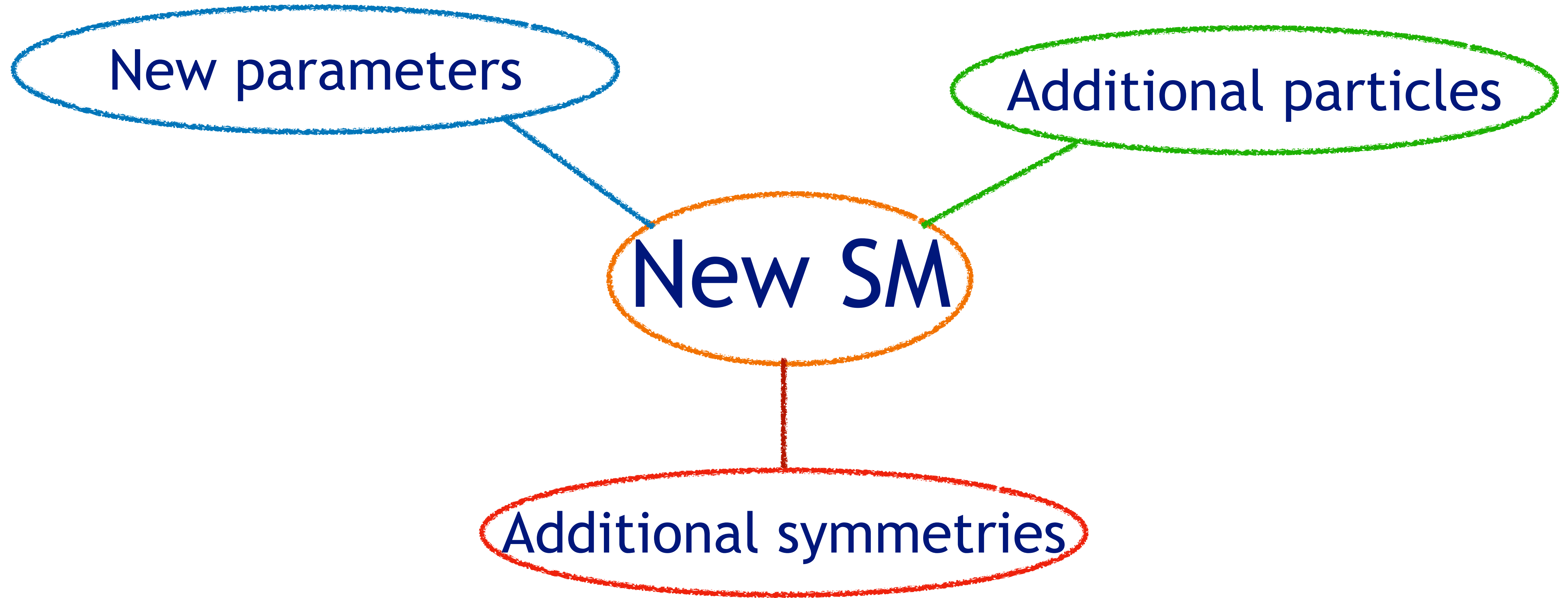
Physics Department
Colorado State University

DESY Theory workshop 2025

September 26 2025



**COLORADO STATE
UNIVERSITY**

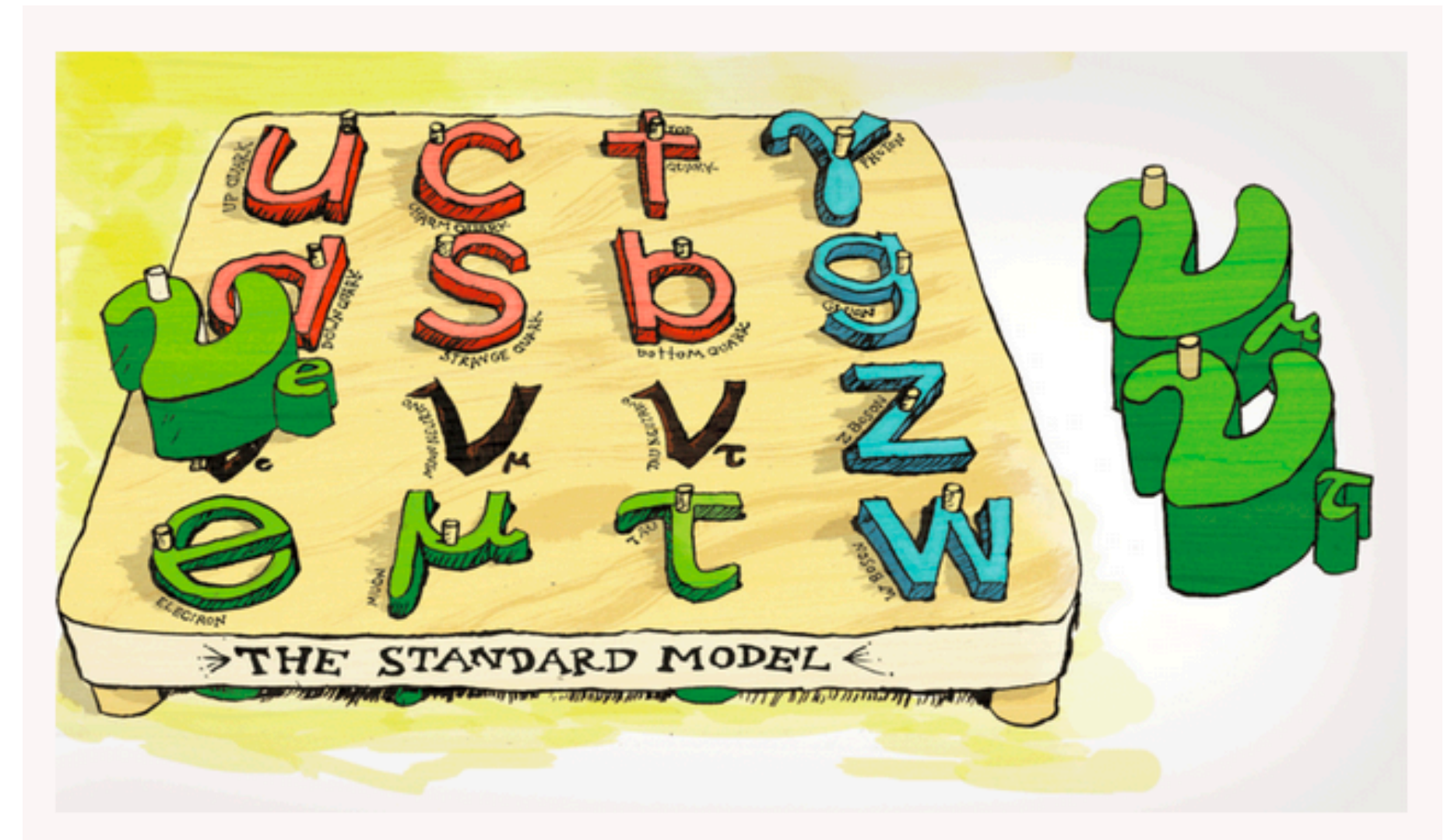


Neutrinos can provide guidance

New parameters

Observation of neutrino oscillations:

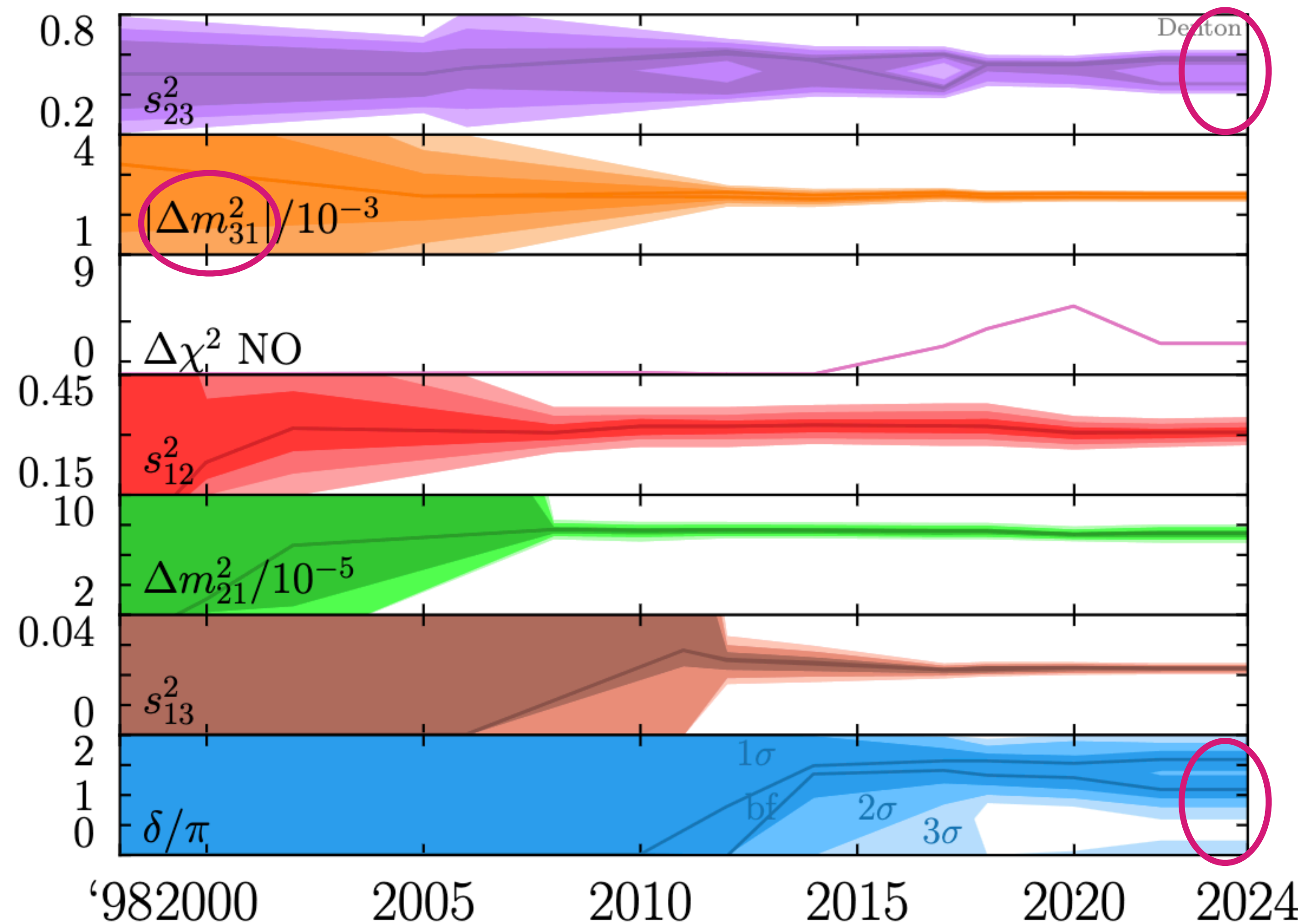
- **Strong** evidence of physics beyond the SM
- introduced **more parameters** to the model (3 angles, at least one phase, 3 masses)
- **need** to introduce neutrino mass mechanism



Neutrino parameters are an **essential** part of the new SM!

New parameters

Where do we stand in measuring the parameters?



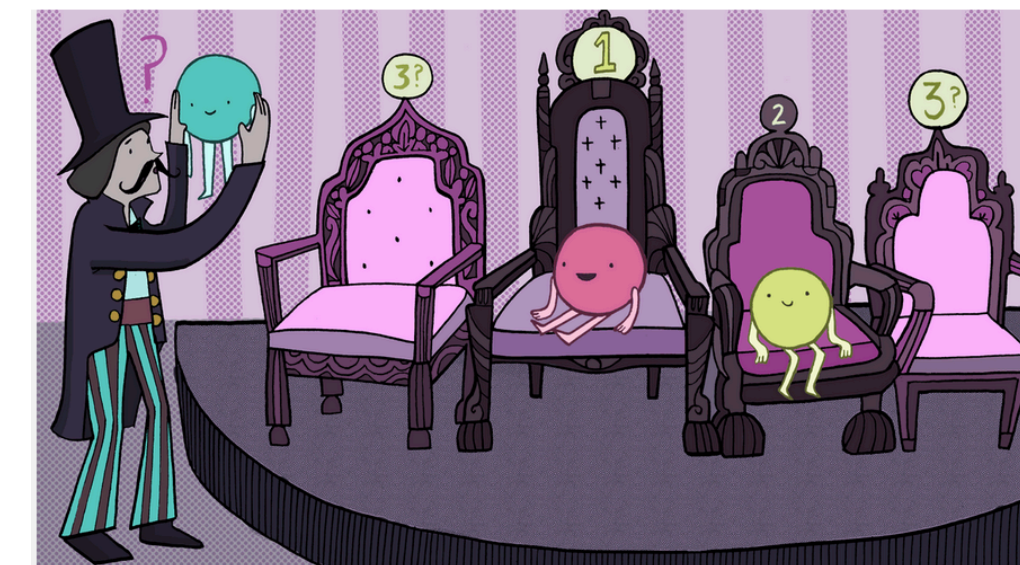
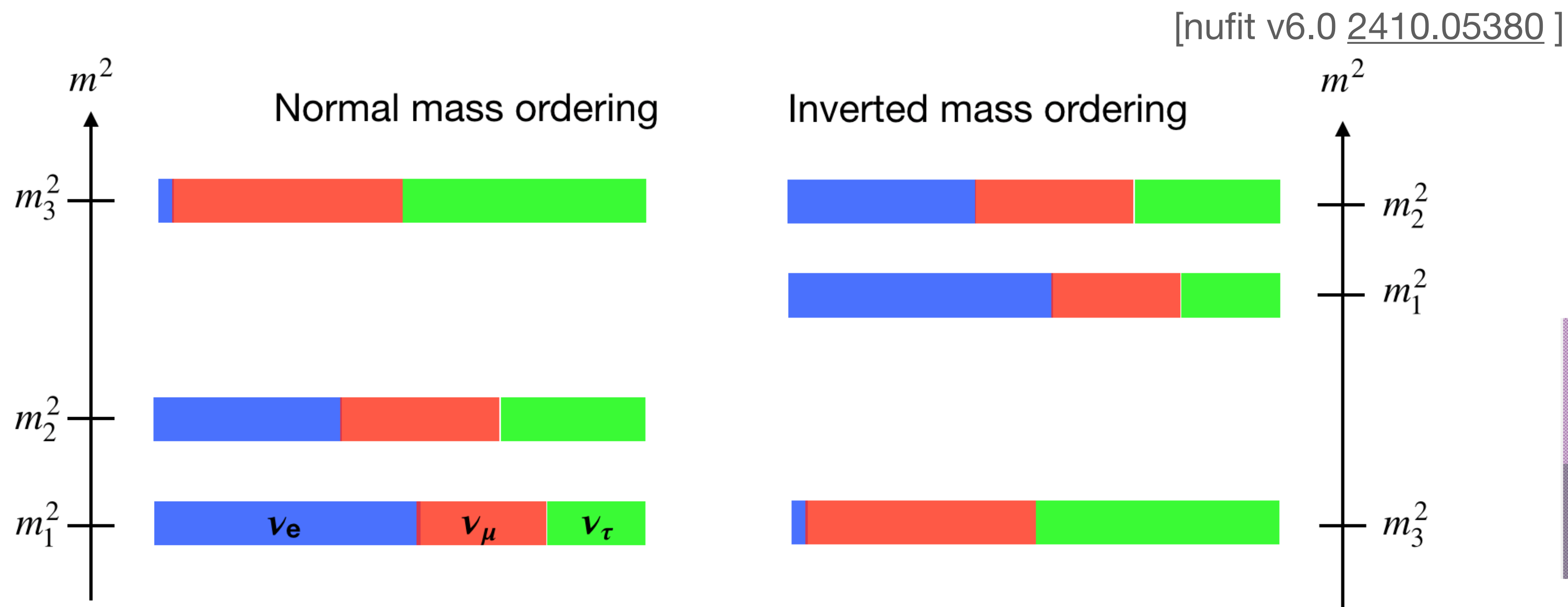
$$\Delta m^2_{ij} = m_i^2 - m_j^2$$

[update from
Denton et al '22]

New parameters

Where do we stand in measuring the parameters?

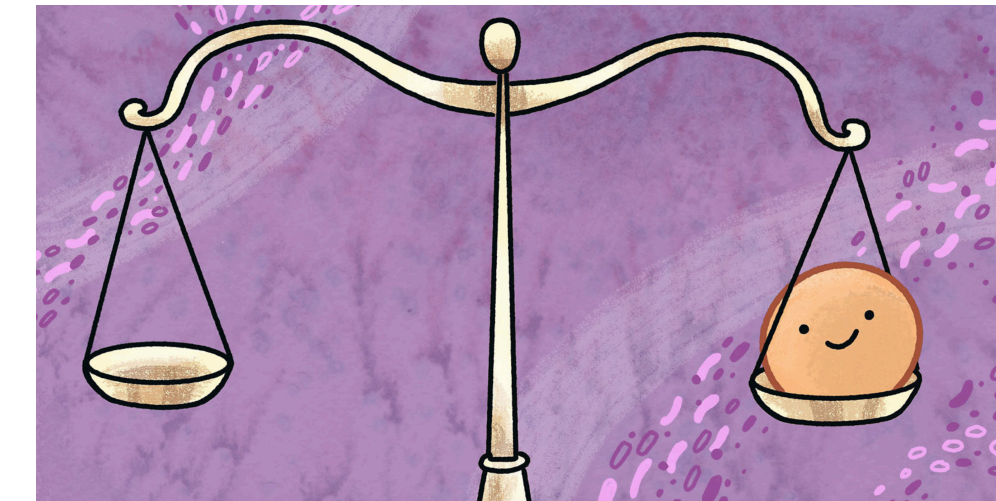
No strong preference for mass ordering from oscillation experiments
(Depending on data sets combined)



New parameters

Where do we stand in measuring the parameters?

What is the absolute neutrino mass scale?



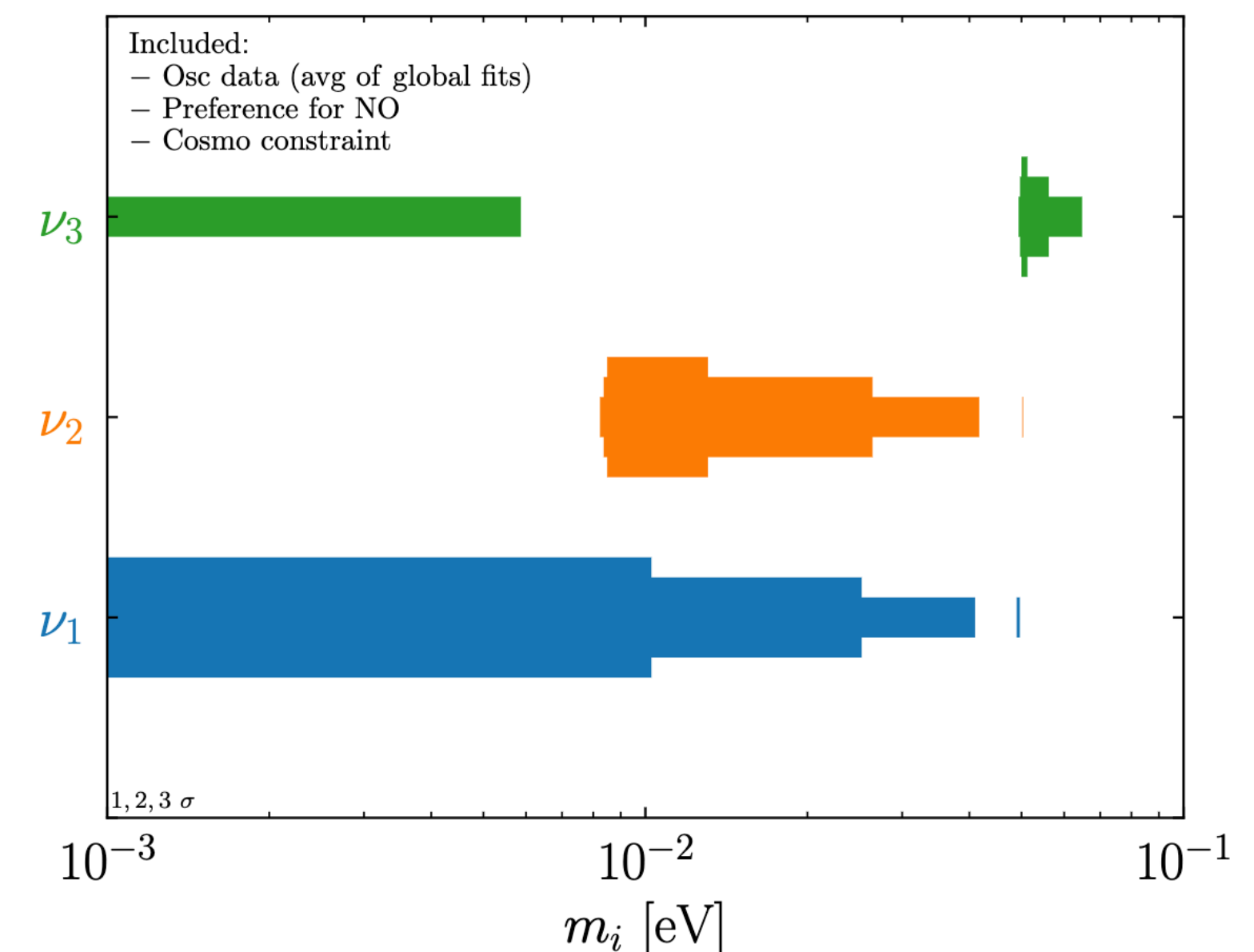
[Denton, JG '23]

Strong bound from cosmology on sum of neutrino masses:

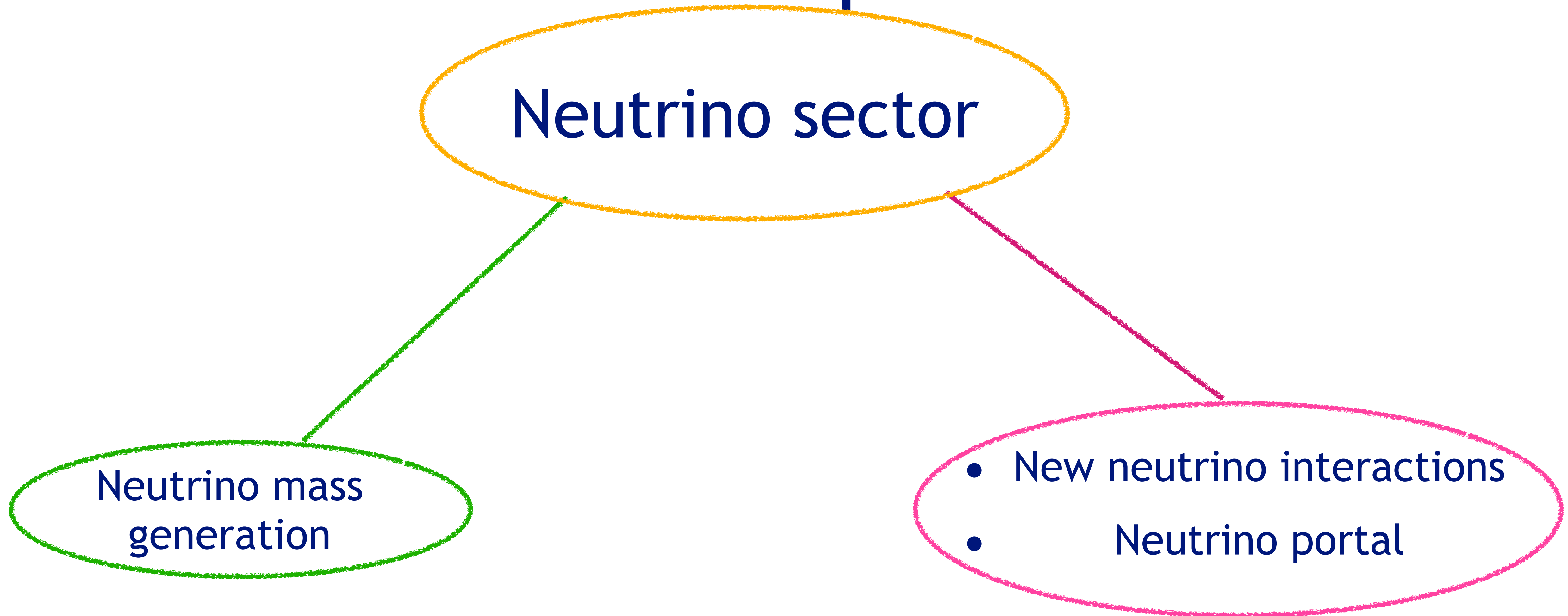
$$\sum m_\nu \lesssim 0.072 \text{ eV}$$

[DESI, '24]

(Depends on data sets combined)



Additional particles



Additional particles

What is the neutrino mass generation mechanism?

Dirac neutrinos

$$m_D = y_\nu \bar{\nu}_L \tilde{H} N_R$$

like other SM fermions

Majorana neutrinos

$$m_M = M_N \bar{\nu}_L \nu_L^c$$

Neutrinos are the **only** SM particles
that could have such a mass term

Term **not** gauge invariant!

In any case need **new particles!**

→ observation of neutrino oscillations **predicts new particles!**

Additional particles

Neutrino mass generation mechanism

Majorana mass term arises from **higher dimension operators**

→ explains **smallness** of neutrino mass due to suppression by high scale

Majorana neutrino mass operators occur at **odd dimension**

[Kobach '16]

$$\mathcal{O} \propto (LLHH)(H^\dagger H)^n$$
$$\rightarrow m_\nu \propto c \frac{v^2}{\Lambda} \left(\frac{v}{\Lambda} \right)^{d-5}$$

Additional particles

Neutrino mass generation mechanism

Dimension 5 operator

$$\mathcal{L}_5 = c_5 \frac{(\bar{L}^c \tilde{H}^*)(\tilde{H}^\dagger L_L)}{\Lambda}$$

Only dim-5 operator that can
be build with SM fields alone

Expect first signs of
new physics from lowest SMEFT operator

→ **neutrino mass**

Dimension 7

$$\mathcal{L}_7 \propto c_7 \frac{(LLHH)(H^\dagger H)}{\Lambda^3}$$

Dimension 9

$$\mathcal{L}_9 \propto c_9 \frac{(LLHH)(H^\dagger H)(H^\dagger H)}{\Lambda^5}$$

...

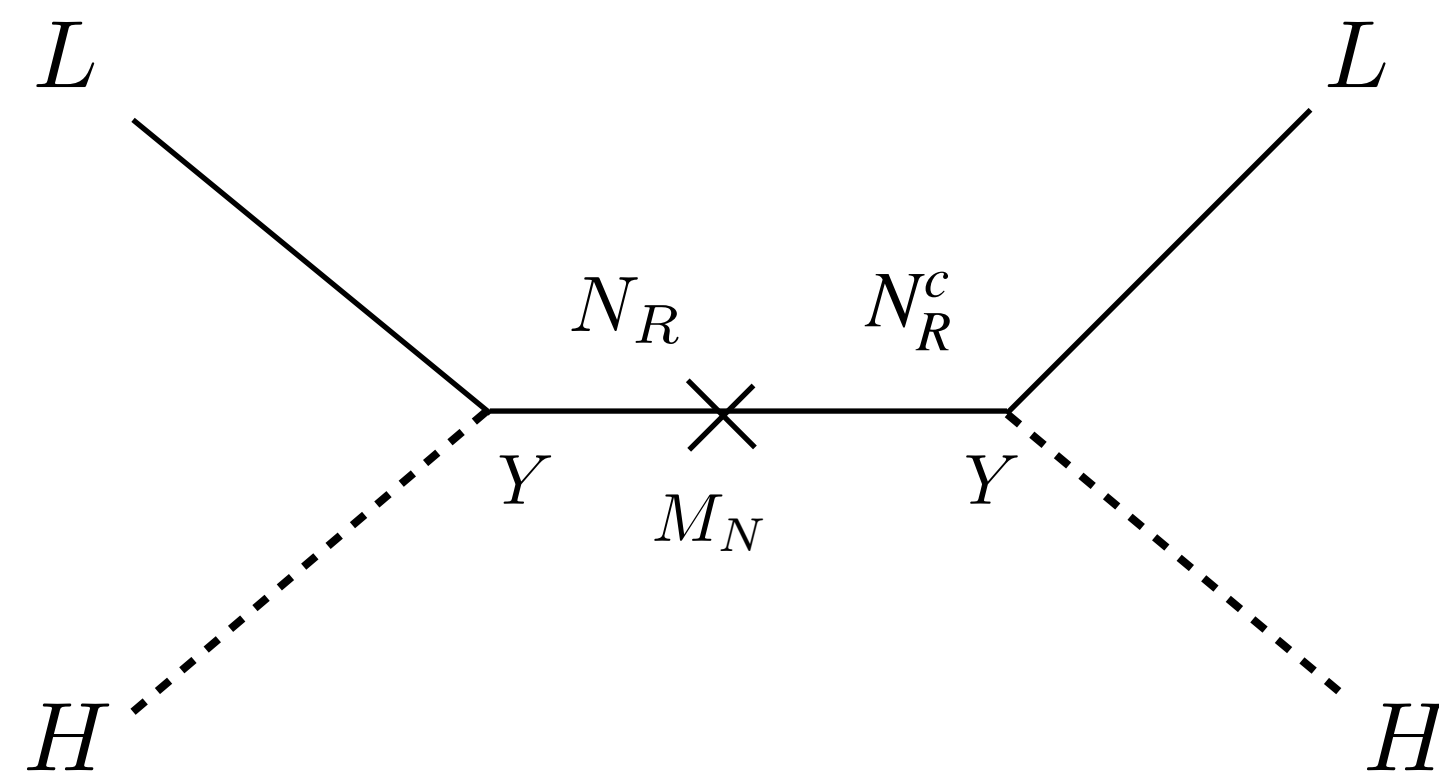
Additional particles

Neutrino mass generation mechanism

Identify **UV complete models** for higher dimensional operators

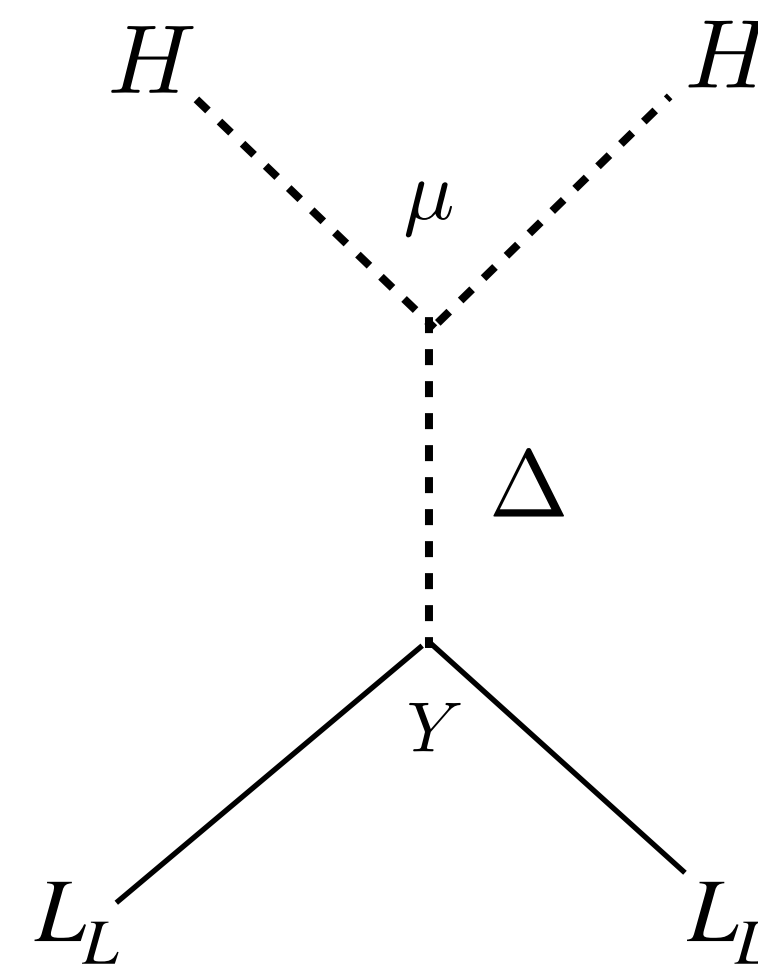
Example: realizations of dim-5 operator

Type I seesaw



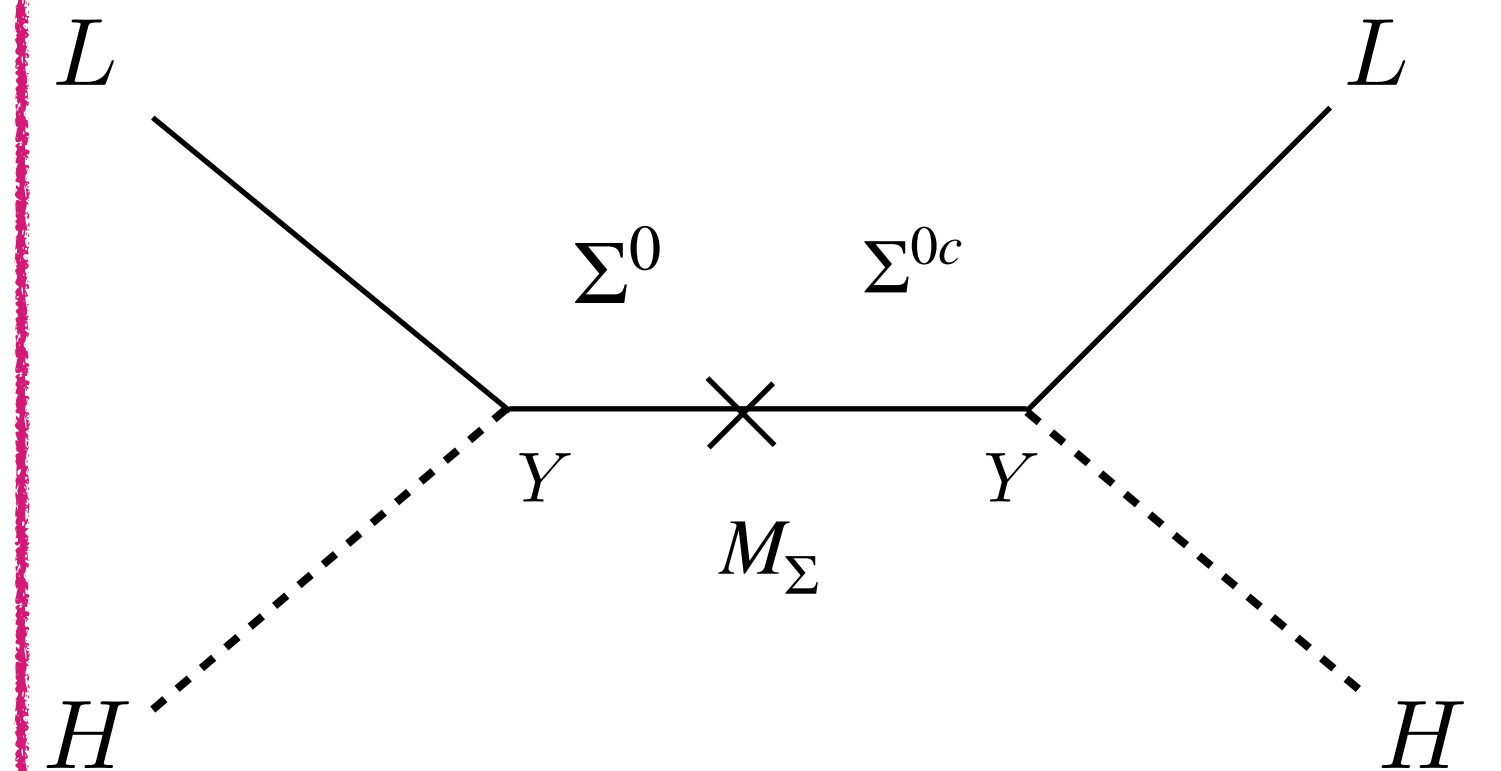
N_R : SM singlet
“Right-handed neutrino”

Type II seesaw



Δ : $SU(2)_L$ scalar triplet

Type III seesaw



Σ : $SU(2)_L$ fermion triplet

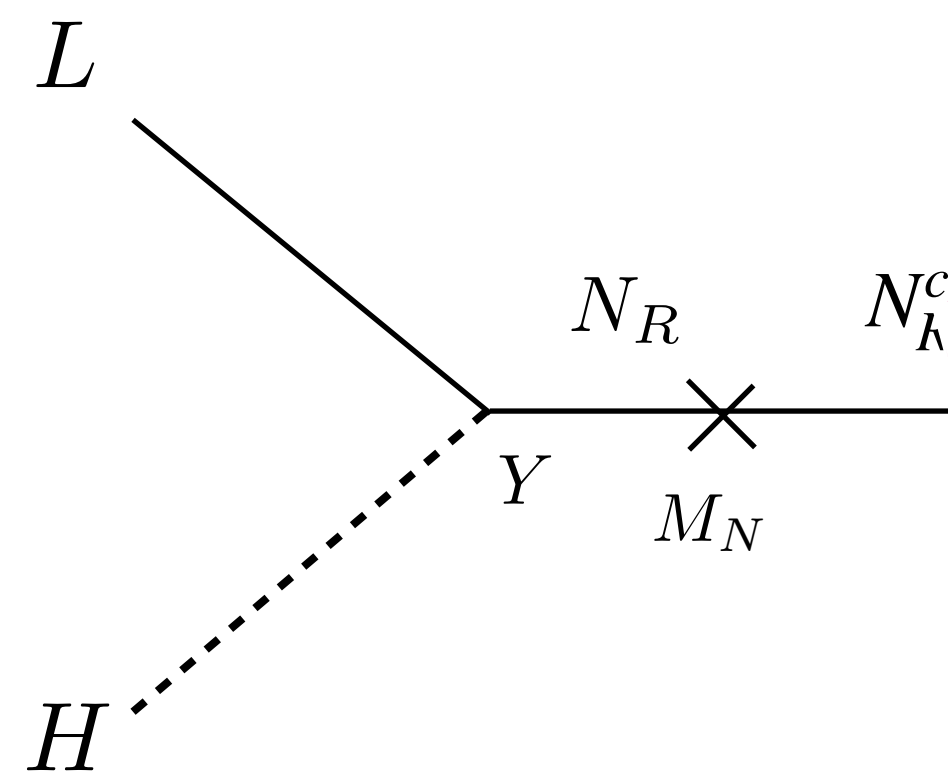
Additional particles

Neutrino mass generation mechanism

Identify **UV complete models** for higher dimensional operators

Example: realizations of dim-5 operator

Type I seesaw



N_R : SM singlet
“Right-handed neutrino”

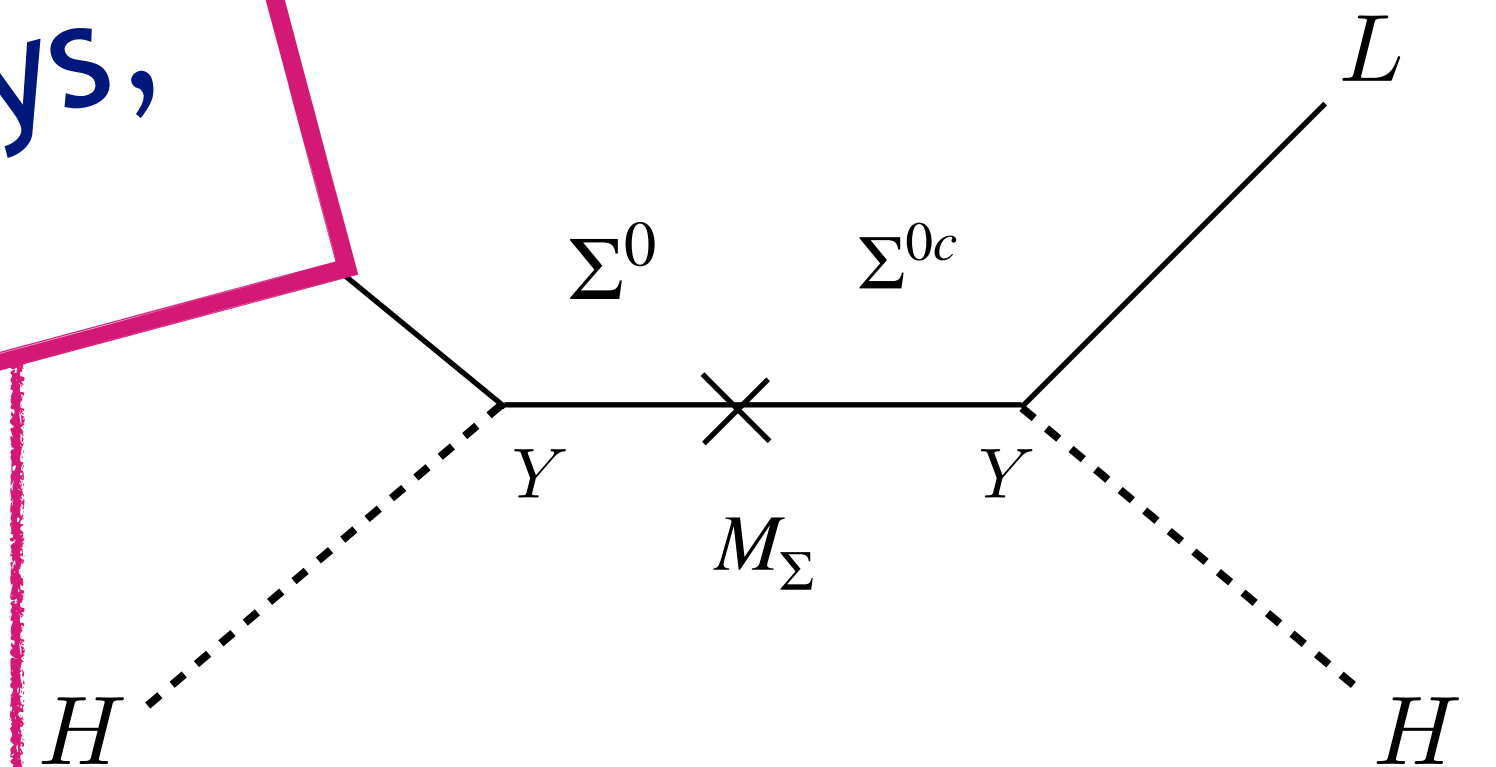
Type II seesaw

H



Δ : $SU(2)_L$ scalar triplet

Type III seesaw



Σ : $SU(2)_L$ fermion triplet

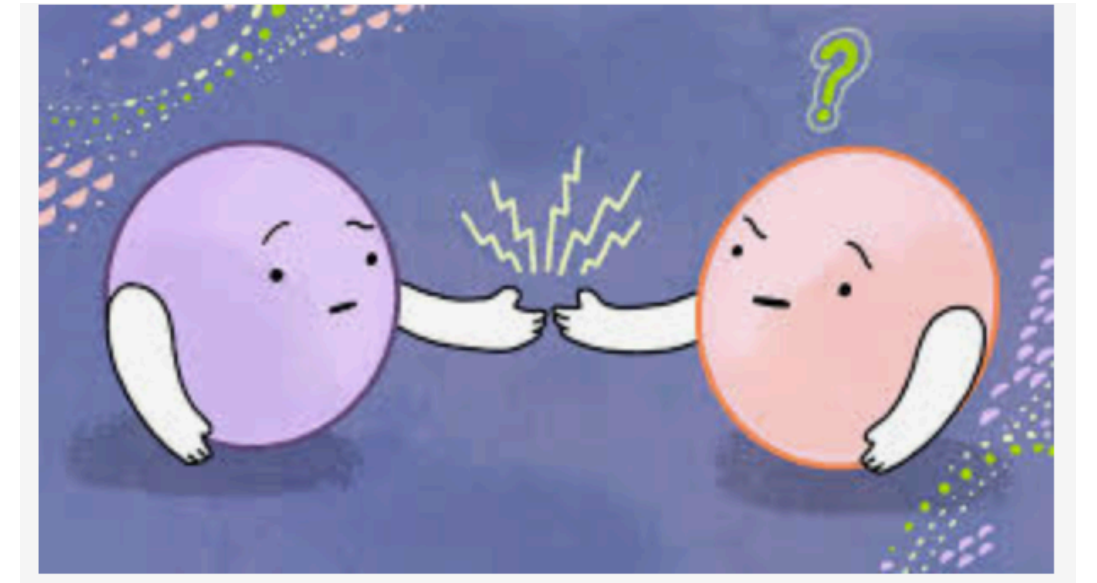
Constraints from: EWPO, collider
phenomenology, meson decays,
cLFV

Additional particles

Neutrino portal

Couple right-handed neutrino to SM

→ Neutrino **portal**



$$\mathcal{L} \supset y \bar{L}_L \tilde{H} N_R$$

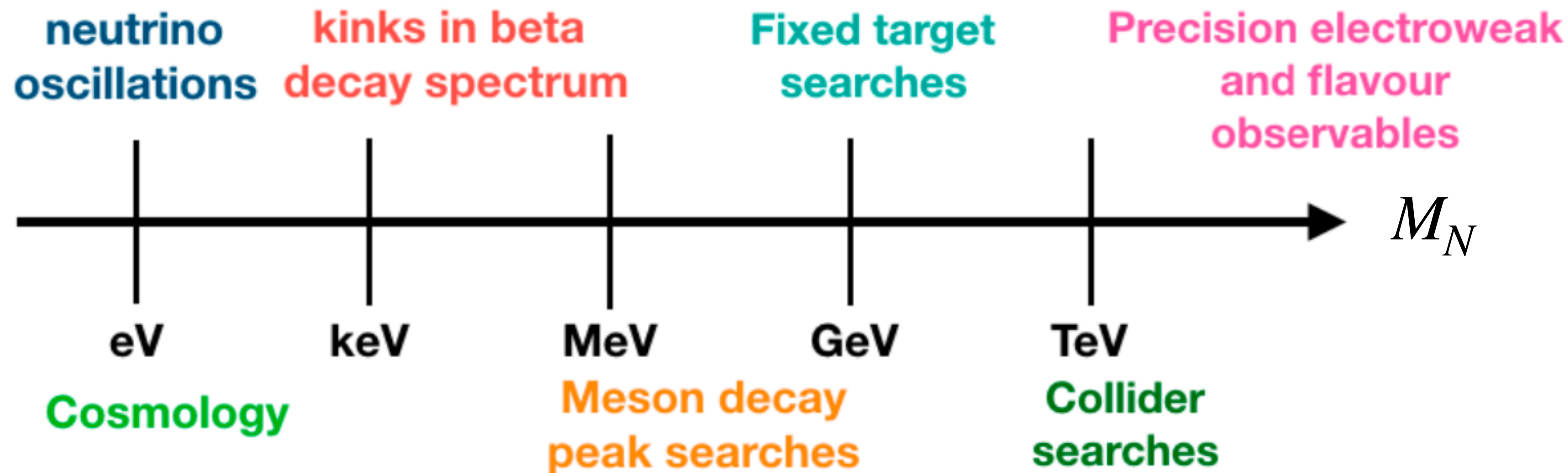
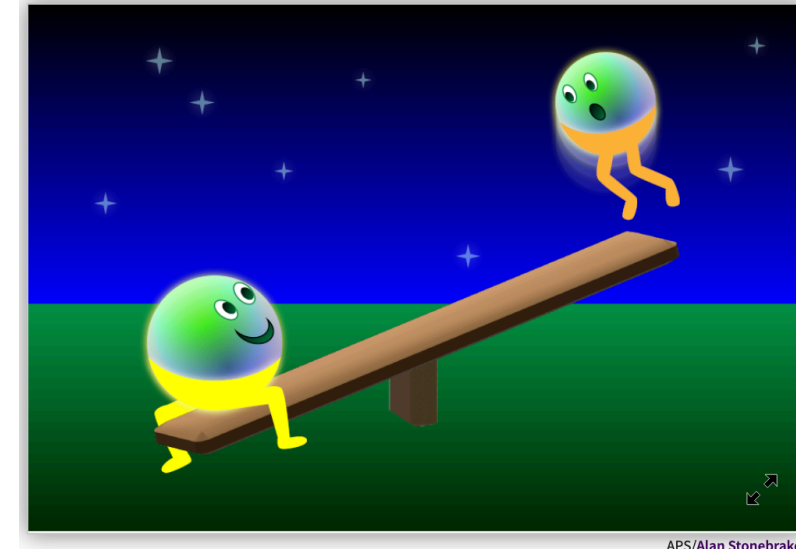
Sterile neutrino does not need to be involved
in neutrino mass generation

Additional particles

Neutrino portal

Interaction of sterile with SM governed by **active-sterile mixing angle**

$$\theta \propto \frac{y\nu}{M_N}, \text{ neutrino mass } m_\nu \propto \theta \times (y\nu)$$



Constraints are around $\theta^2 \lesssim 10^{-5}$ but depend on M_N

[Dev, Deppisch, Bolton '19]

[Fernandez-Martinez et al '23]

Additional particles

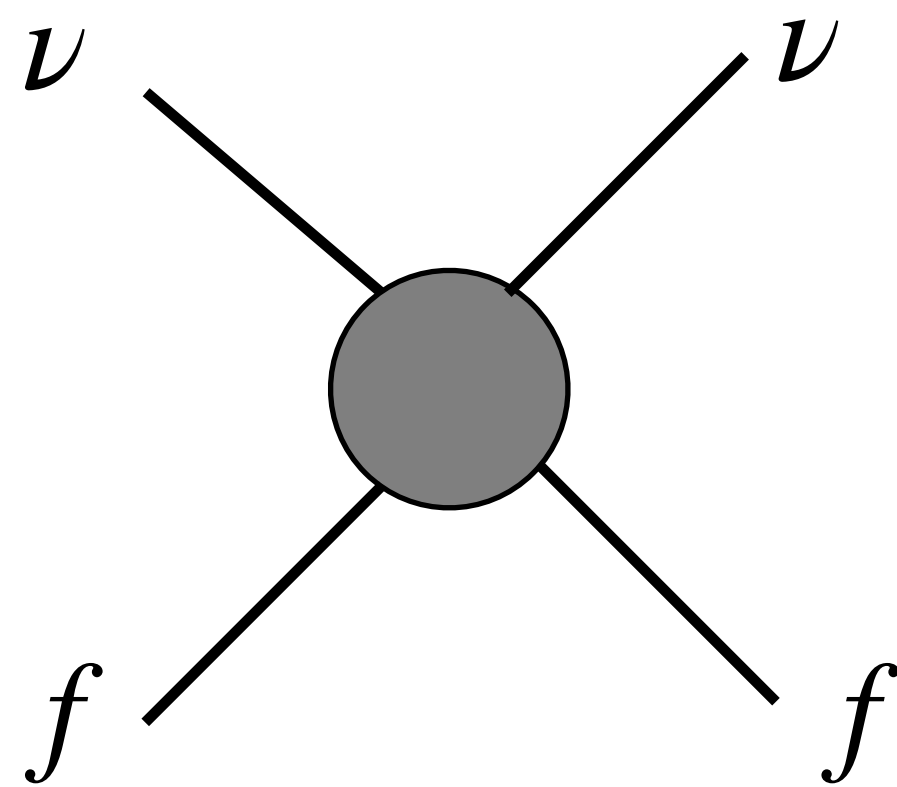
New neutrino interactions

4-fermion operators

→ neutrino non-standard interactions (NSI)

$$\text{NC NSI: } \mathcal{L}_{NSI} = -2\sqrt{2}G_F \sum_{\alpha,\beta,f,P} \epsilon_{\alpha\beta}^f (\bar{\nu}_\alpha P \nu_\beta) (\bar{f} P f)$$

[Wolfenstein '77]



$$P = S, V, A, T$$

$$P = 1, i\gamma_5, \gamma^\mu, \gamma^\mu\gamma_5, \sigma^{\mu\nu}, \sigma^{\mu\nu} = i/2[\gamma^\mu, \gamma^\nu]$$

Lots of possible NSI terms!

$$\epsilon \sim \mathcal{O}(G_X/G_F)$$

Additional particles

Vector NSI effects in oscillations

In oscillations: new forward scattering with matter

→ **independent** of mass of new mediator!

Affect neutrino oscillations as a **new matter effect**

$$H = \frac{1}{2E} \left[U^\dagger M^2 U + a \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{e\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix} \right]$$

Matter potential $a \propto G_F \rho E$

→ experiments with **largest** E and L provide leading constraints

Additional particles

NSI effects in NC scattering

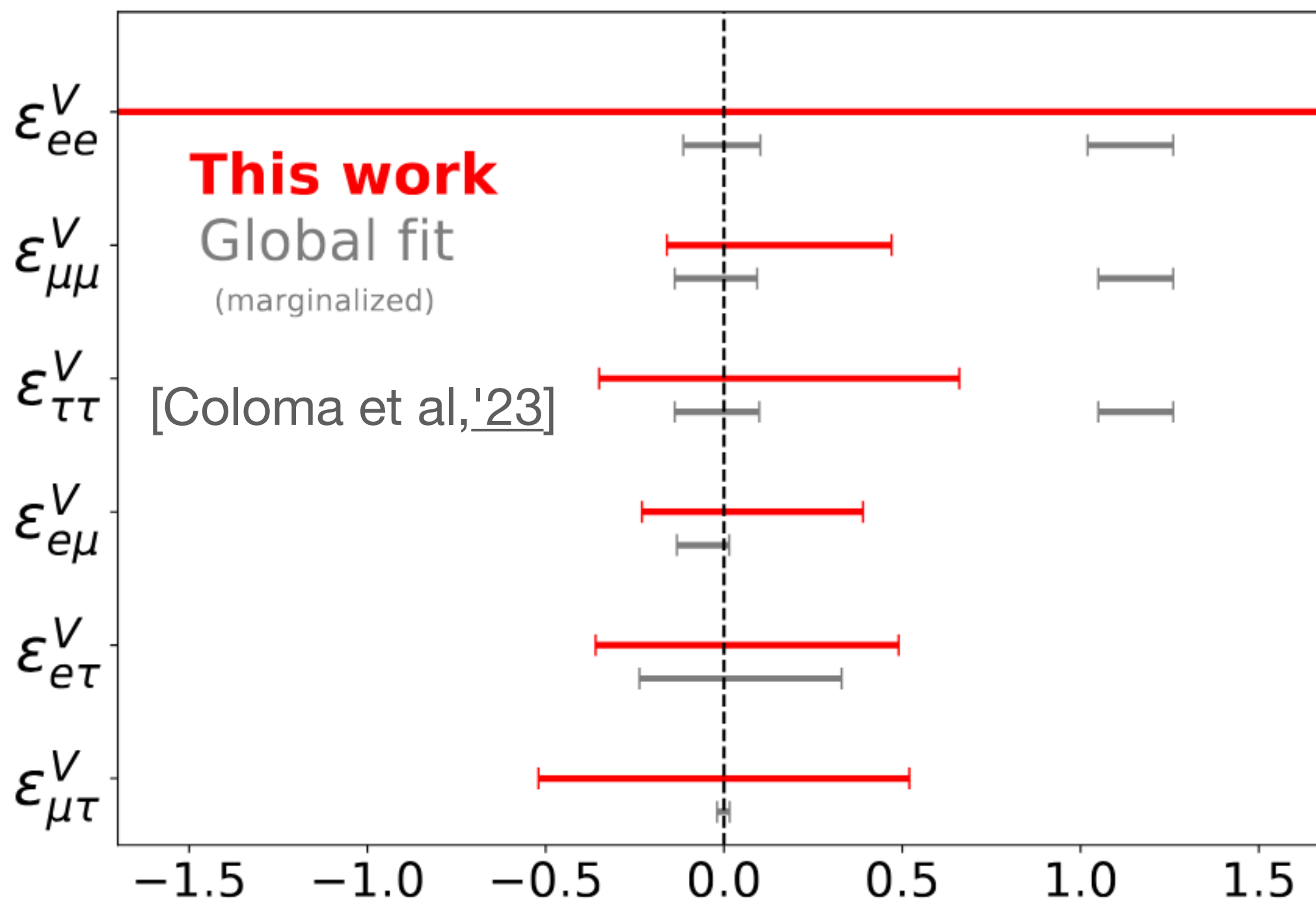
Vector NSI:
Complementarity with
oscillations

dependence on mediator mass

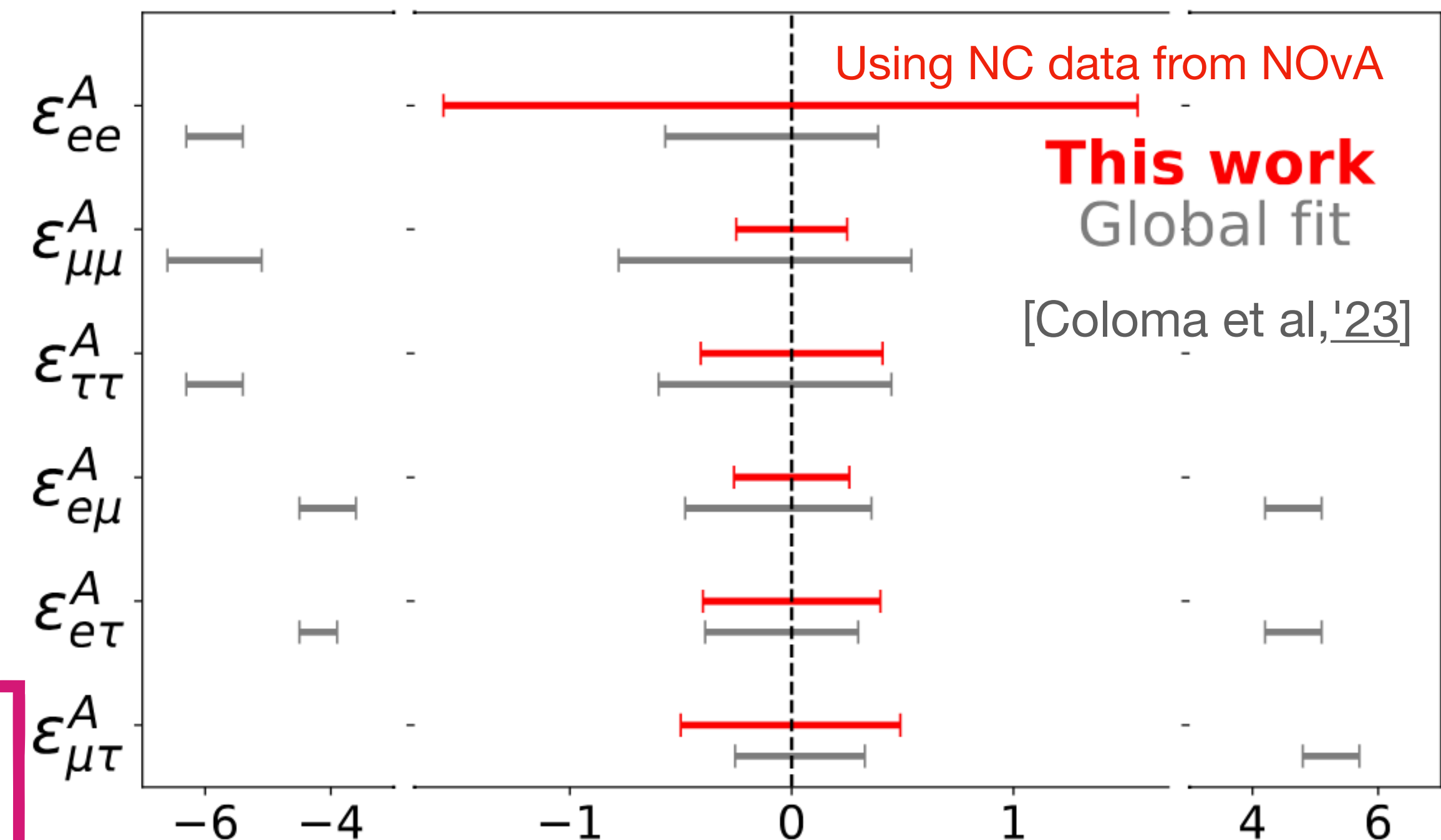
Axial-Vector NSI:
Data from SNO

First analysis of NOvA's NC data set

[JG, Machado, Pinheiro '24]



$\epsilon \sim \mathcal{O}(0.1)$
allowed!



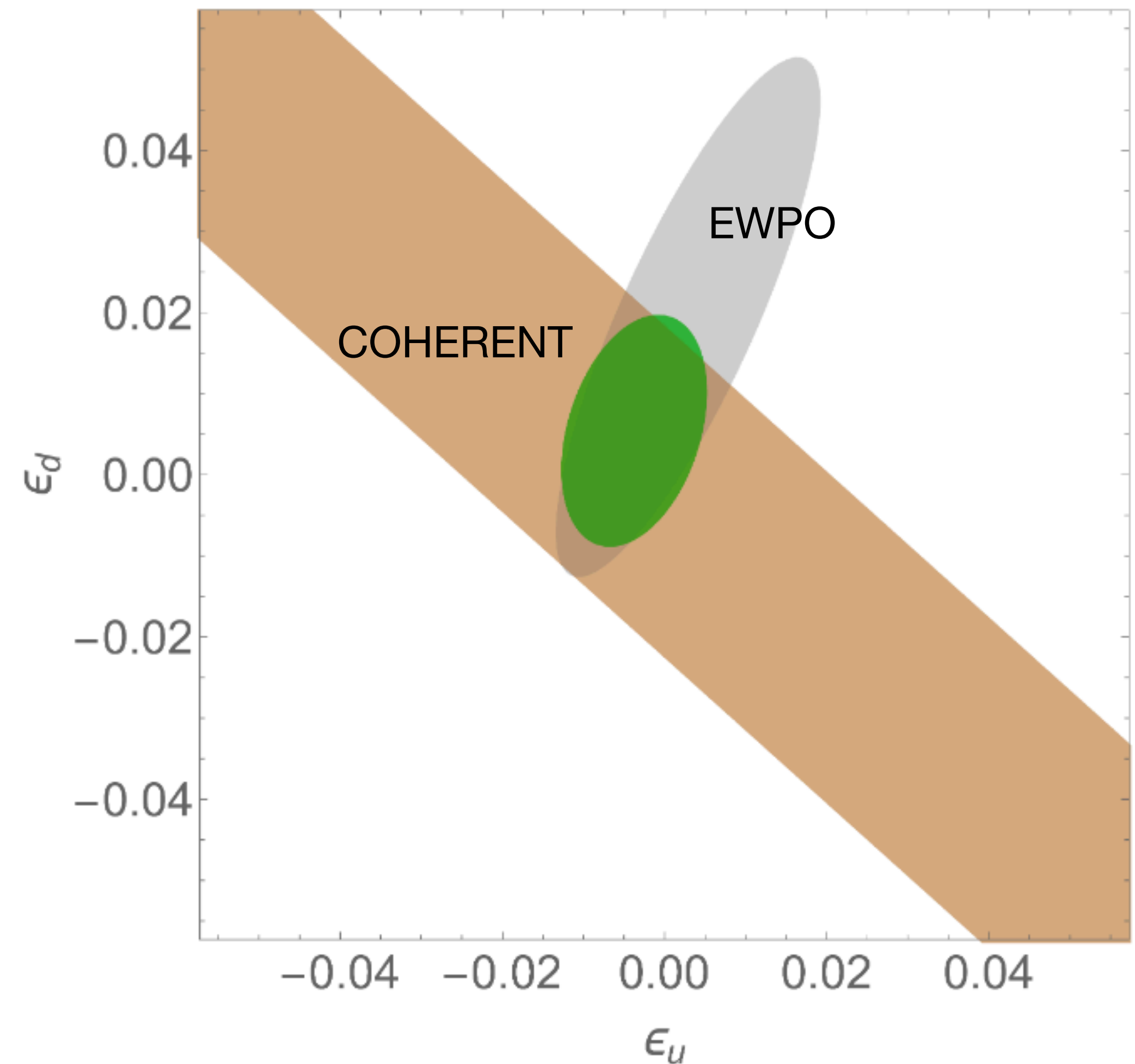
Additional particles

NSI effects in NC scattering

[Breso-Pla et al '23]

In SMEFT framework:
Complementarity with
constraints from charged lepton
observables

[for more on SMEFT see talk by Vryonidou]

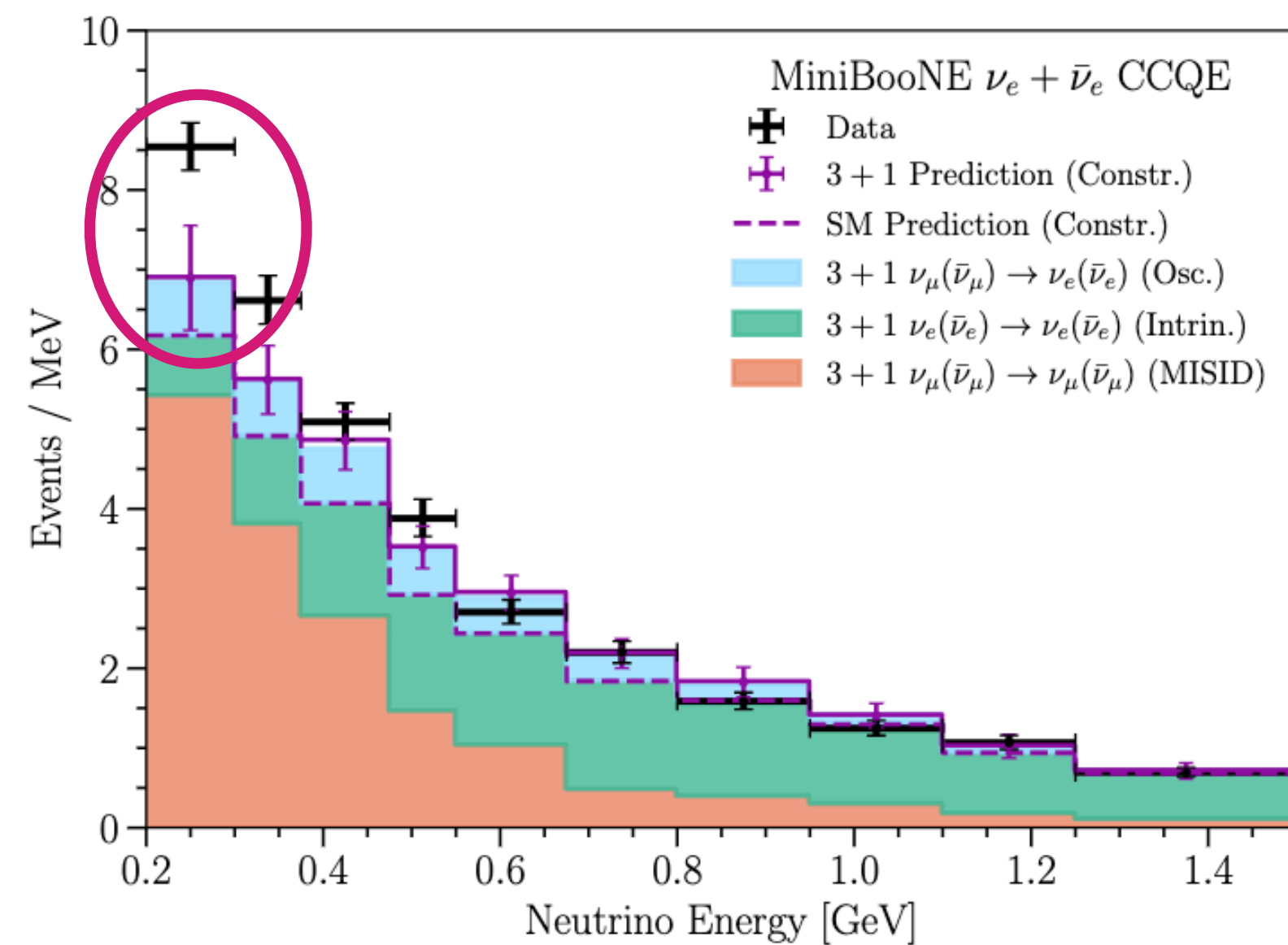


Additional particles

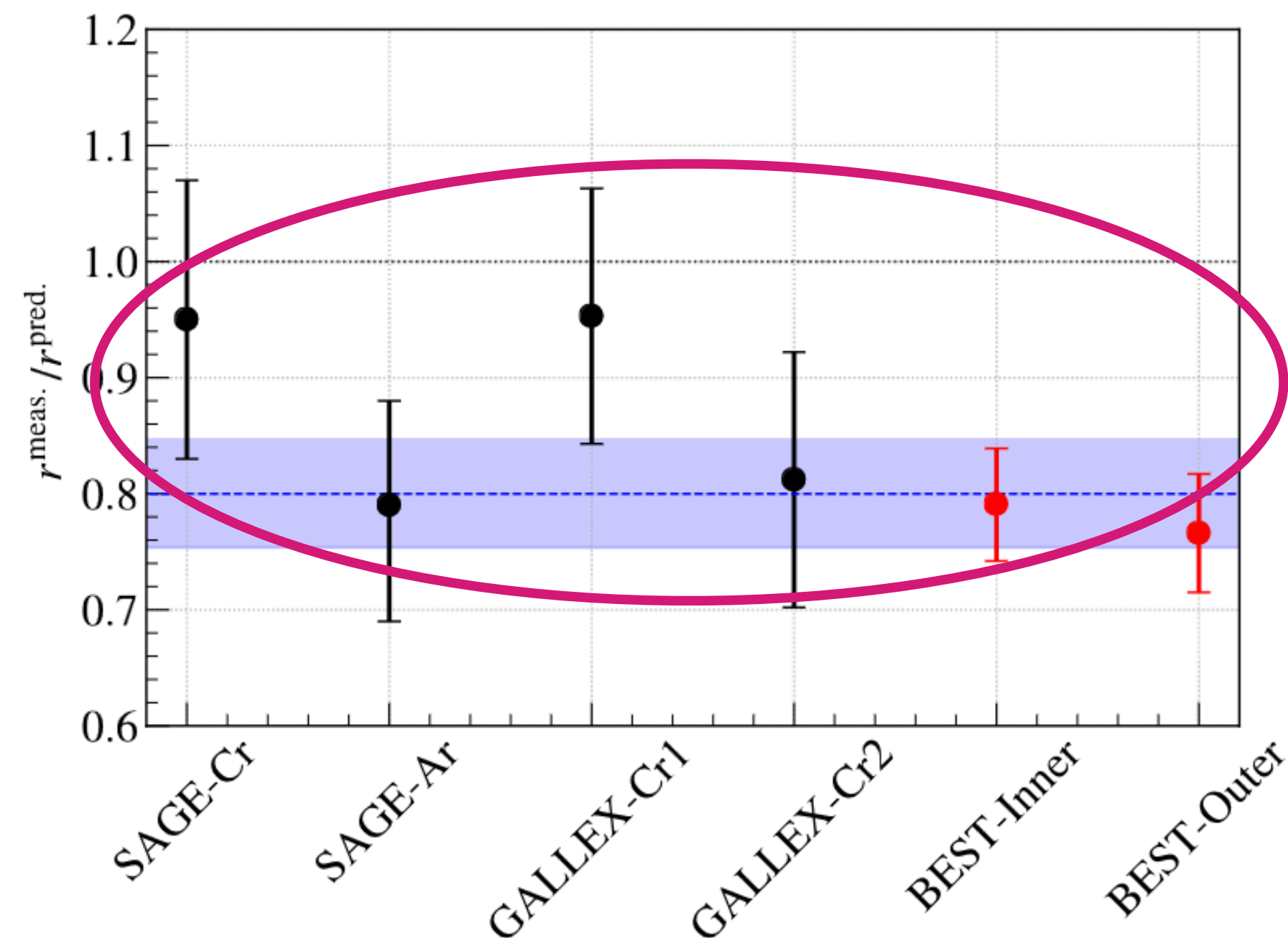
Anomalies

Some (persistent) **anomalies** in neutrino physics
eV-scale vanilla sterile neutrino explanation **not** a good global fit!
So far **no convincing SM explanation** either

[Brdar, JG, Kopp '23]



[MiniBooNE+MicroBooNE 2201.01724]



[BEST 2201.07364]

New symmetries

Neutrino sector

```
graph TD; A([Neutrino sector]) --- B([Leptonic CP violation]); A --- C([Lepton number violation]); A --- D([Flavor symmetry])
```

Leptonic CP violation

Lepton number violation

Flavor symmetry

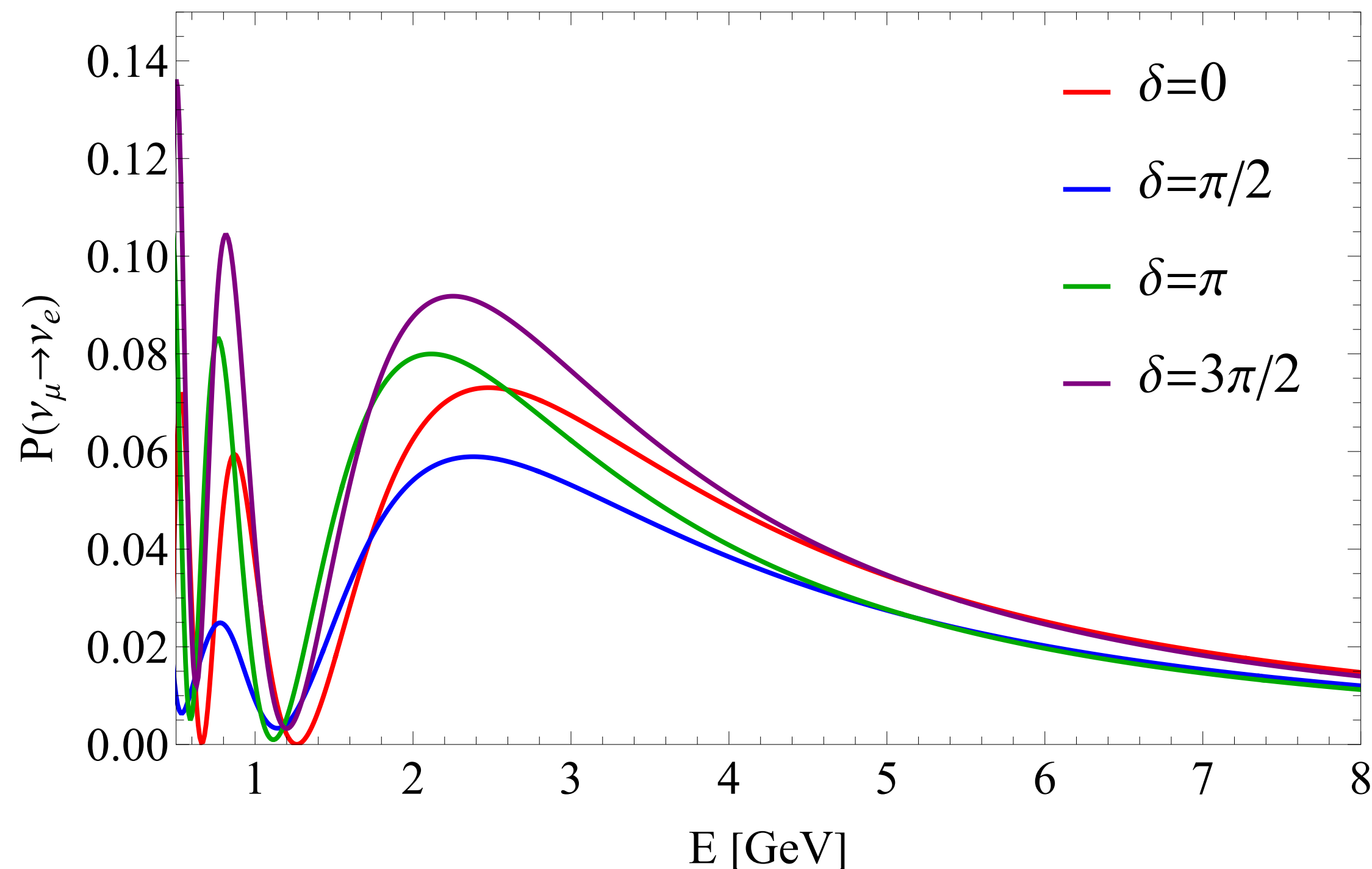
New symmetries

Leptonic CP violation

Use long-baseline experiments:

$\bar{\nu}_e$ appearance probability in $\bar{\nu}_\mu$ beam depends on δ

(However also dependence on mass ordering, other oscillation parameters)



To alleviate effects of systematics uncertainties:
Compare neutrino to anti-neutrino channel

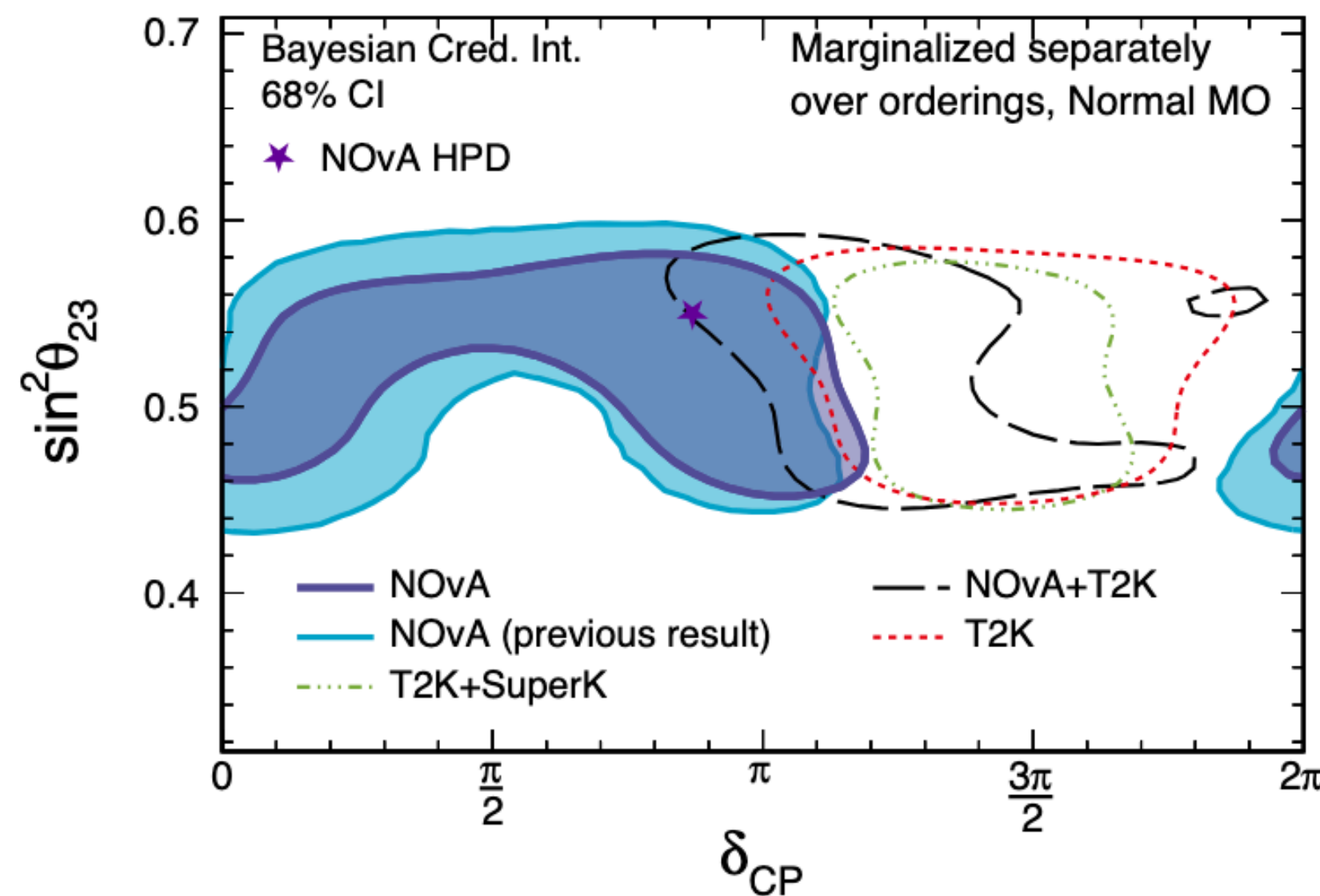
If they behave differently
→ CP violation

New symmetries

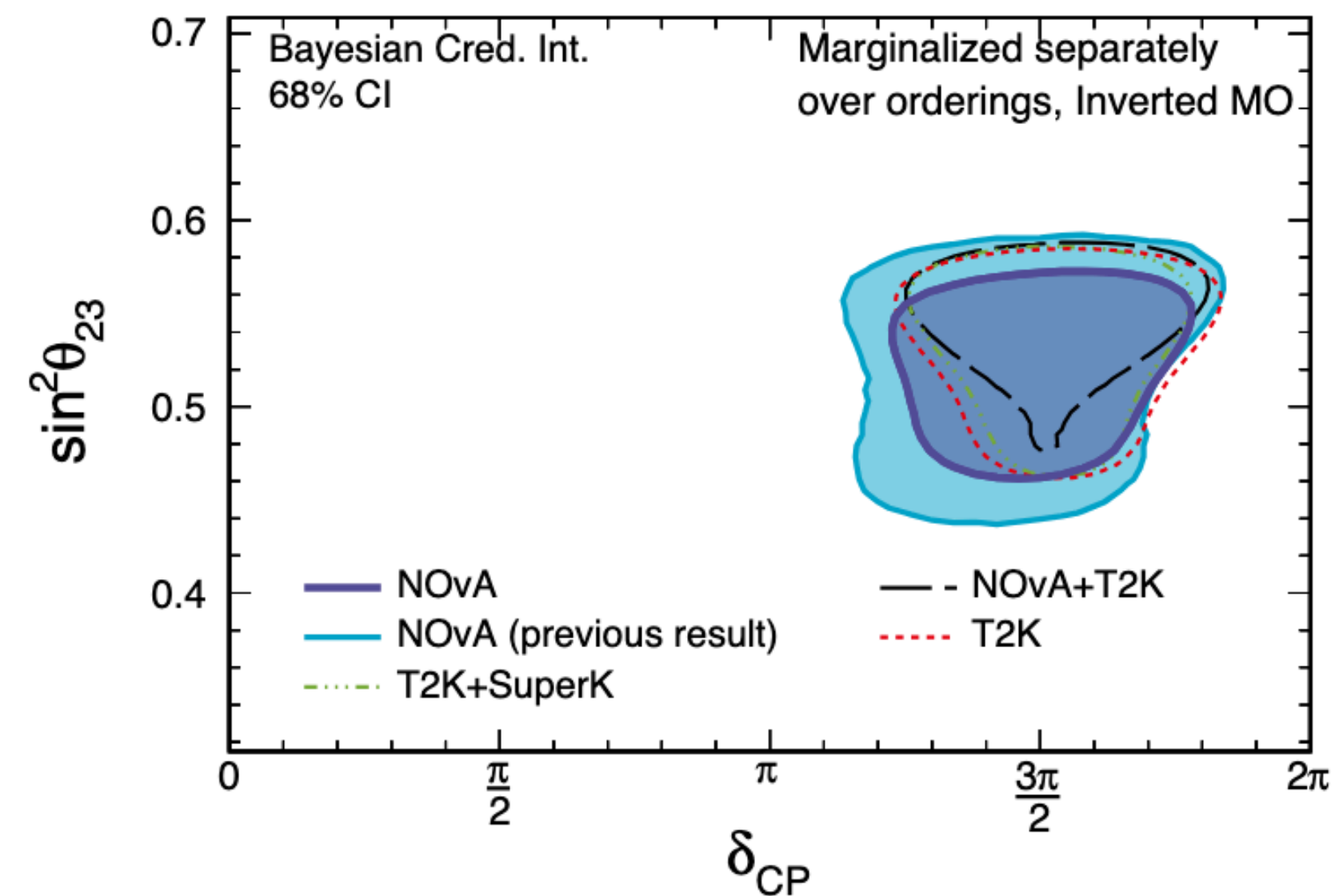
Leptonic CP violation

[see talk by Rusov for the status of CPV in the quark sector]

Current experiments long-baseline experiments:
NOvA (US), T2K (Japan)



In normal ordering:
slight disagreement



In inverted ordering:
improvement on constraints

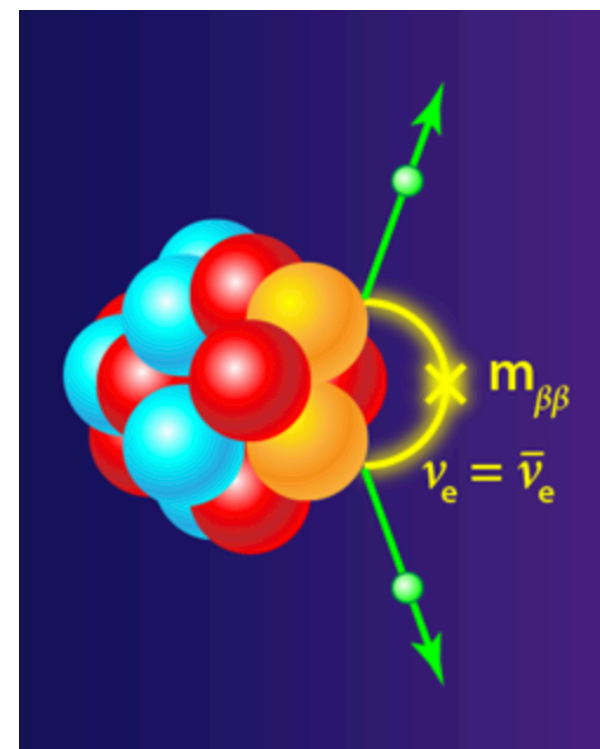
[NOvA '25]

New symmetries

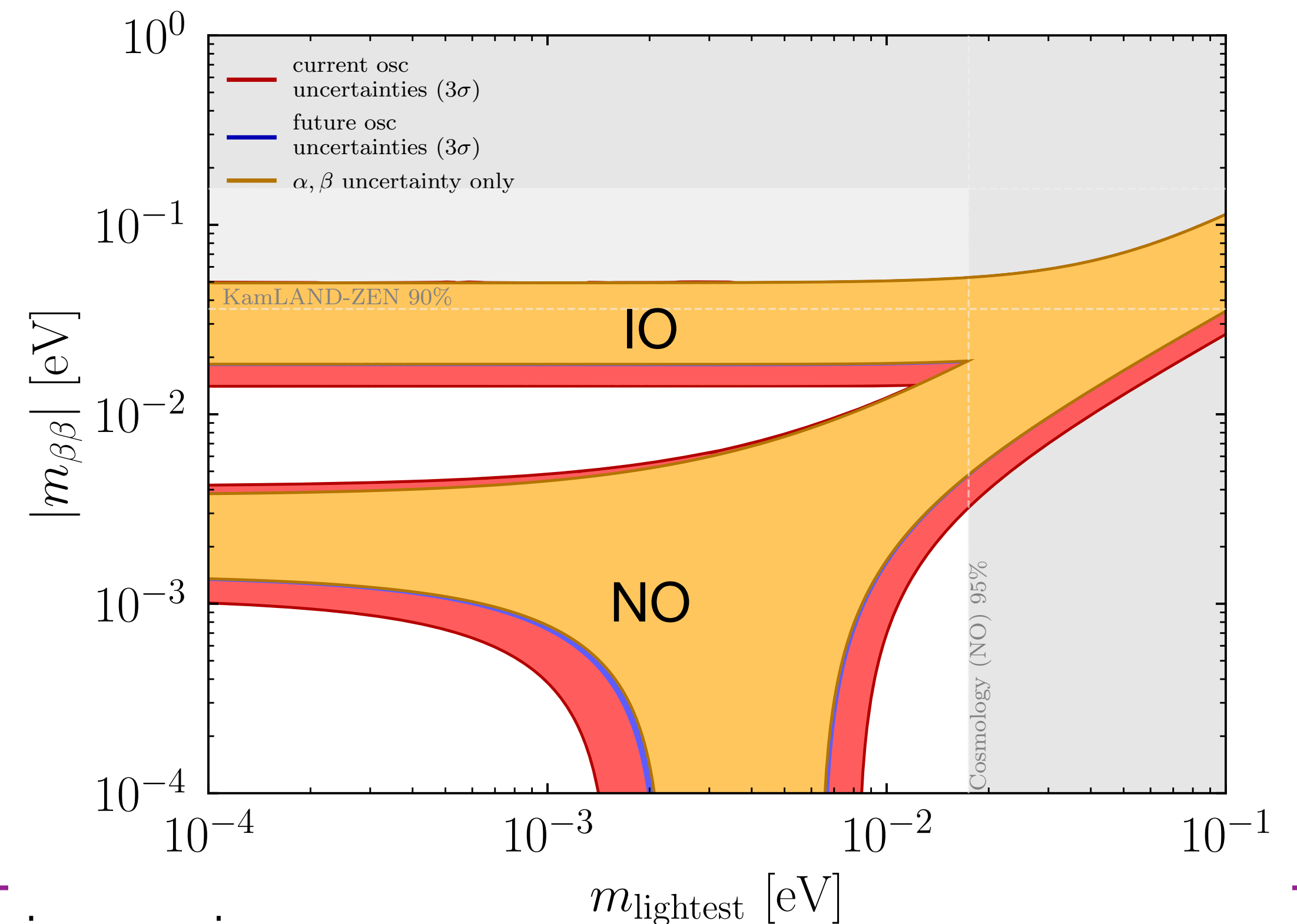
Lepton number symmetry

If neutrinos are Majorana particles $\leftrightarrow$ lepton number is violated

Neutrinoless double beta decay most **promising** probe of Majorana nature of neutrinos



[Denton, JG, '23]



New symmetries

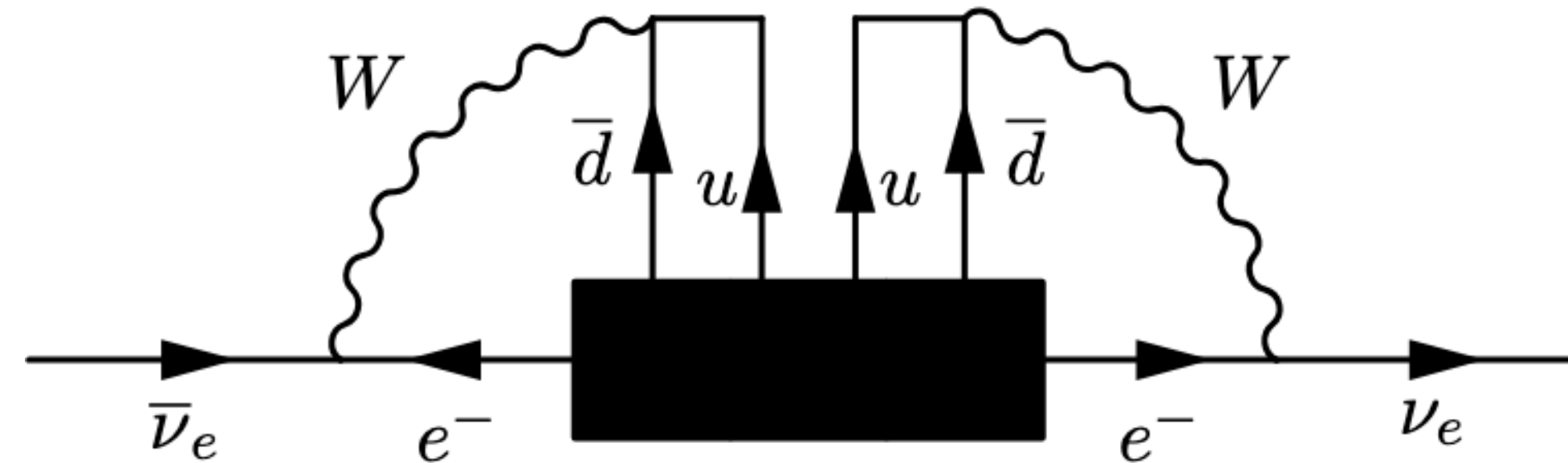
Lepton number symmetry

Schechter-Valle theorem

Any $\Delta L = 2$ operator contributing to $0\nu\beta\beta$ will generate

Majorana neutrino mass contribution

[Schechter, Valle 1982]



Majorana mass induced by this (dim-9) operator tiny: $\lesssim \mathcal{O}(10^{-28} \text{ eV})$

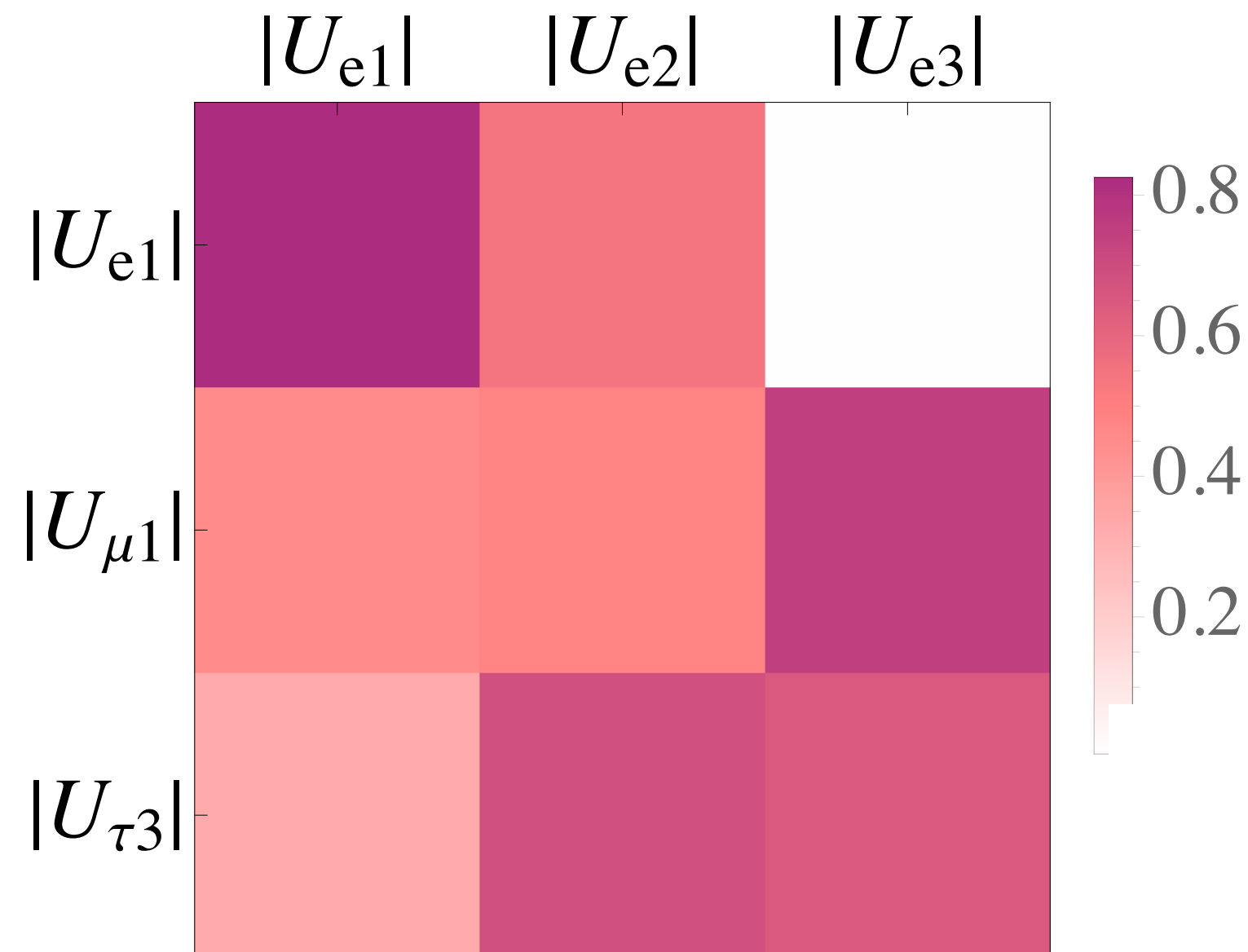
Even if lepton number violation is detected: neutrino mass might be dominantly of Dirac type (=pseudo-Dirac neutrino) [Duerr, Lindner, Merle '11]

New symmetries

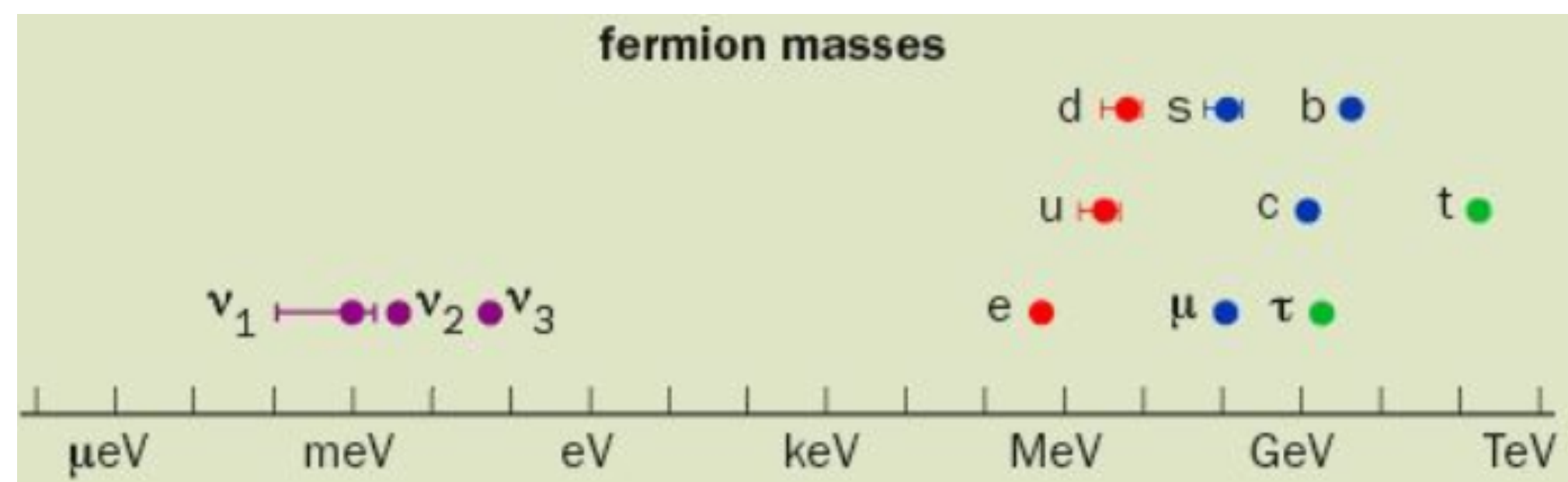
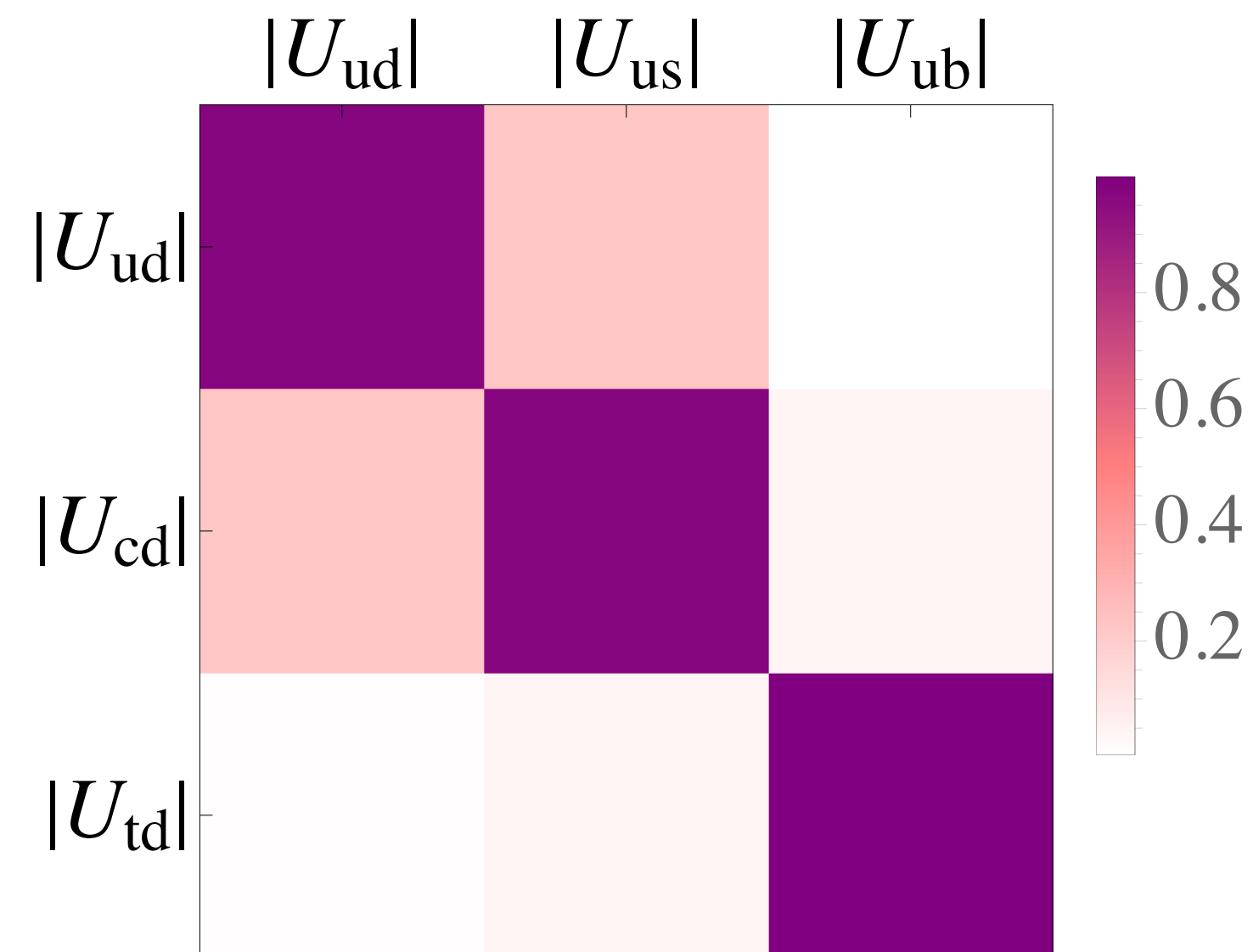
Flavor symmetry

[see talk by Cornella]

Leptons



Quarks



could provide rationale between
difference in mixings in quark sector
and lepton sector

New symmetries

Flavor symmetry

Most successful neutrino flavor models use **discrete** symmetries

Most predictive flavor models predict **relations** between mixing

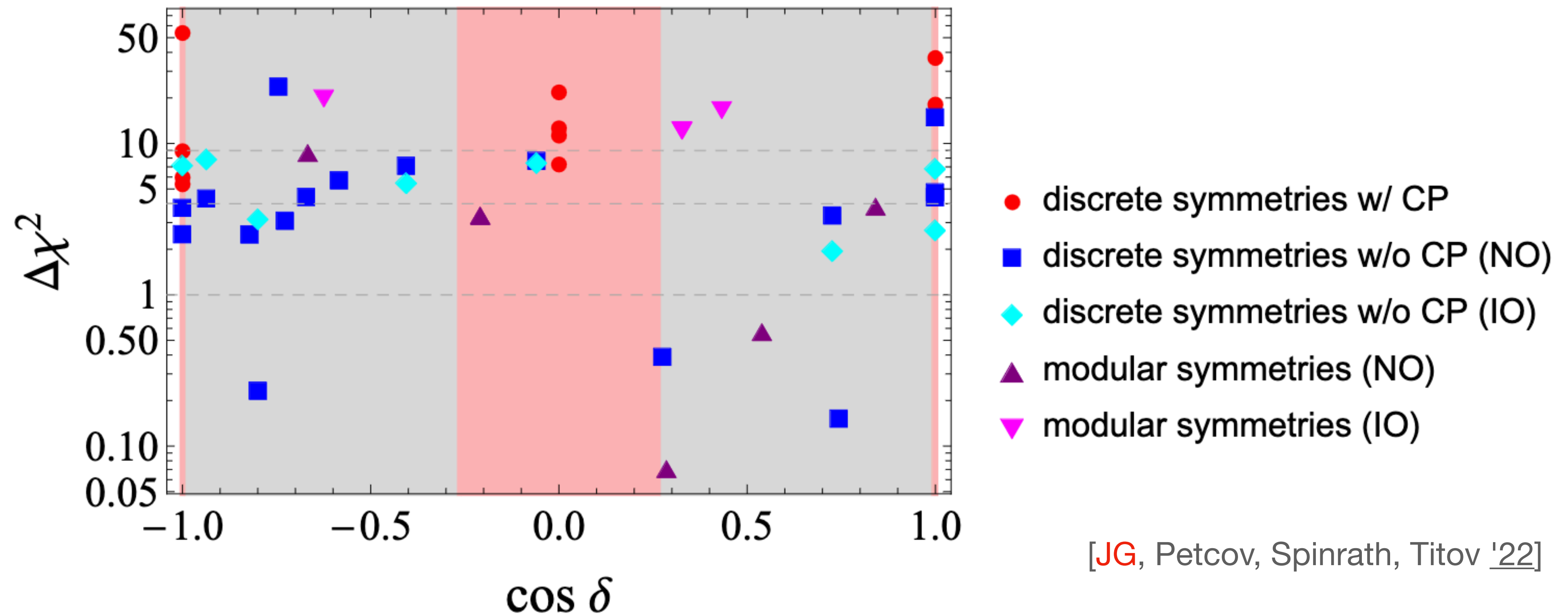
parameter like $\theta_{12}^{\text{PMNS}} - \theta_{12}^{\nu} \approx \theta_{13}^{\text{PMNS}} \cos \delta$

Also prediction for other mixing parameters and masses

New symmetries

Flavor symmetry

Precise measurements of oscillation parameters can distinguish flavor models



[JG, Petcov, Spinrath, Titov '22]

Neutrino physics

The future

Oscillation (standard+BSM) physics

Long baseline (300 km, 1300 km) **accelerator** neutrino experiments:
Hyper-Kamiokande, DUNE

Atmospheric neutrino experiments:
HK, IceCube-Gen2, KM3NeT-ORCA

Medium baseline (~50 km) **reactor** neutrino experiment:
JUNO

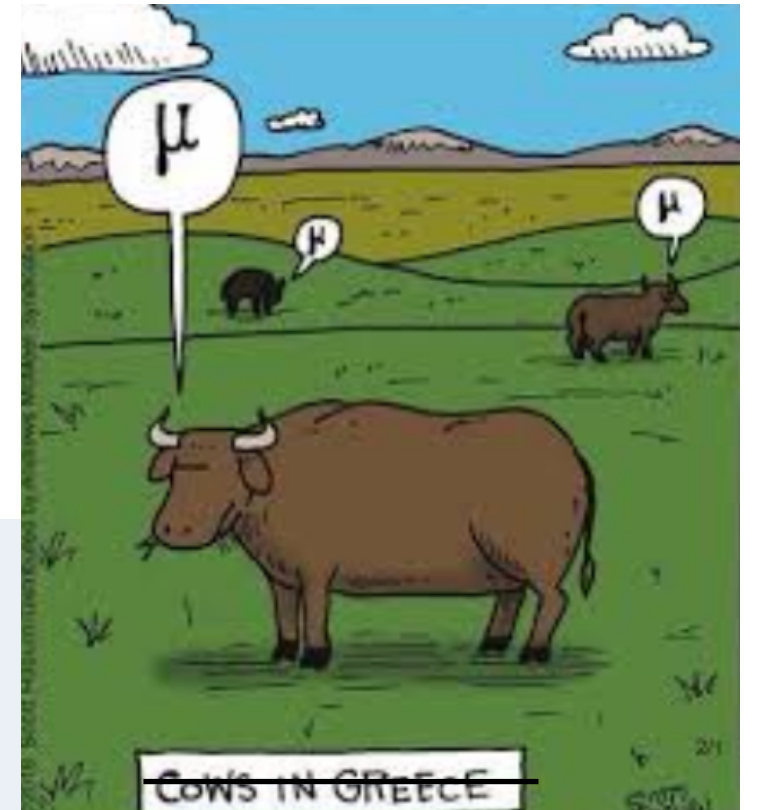
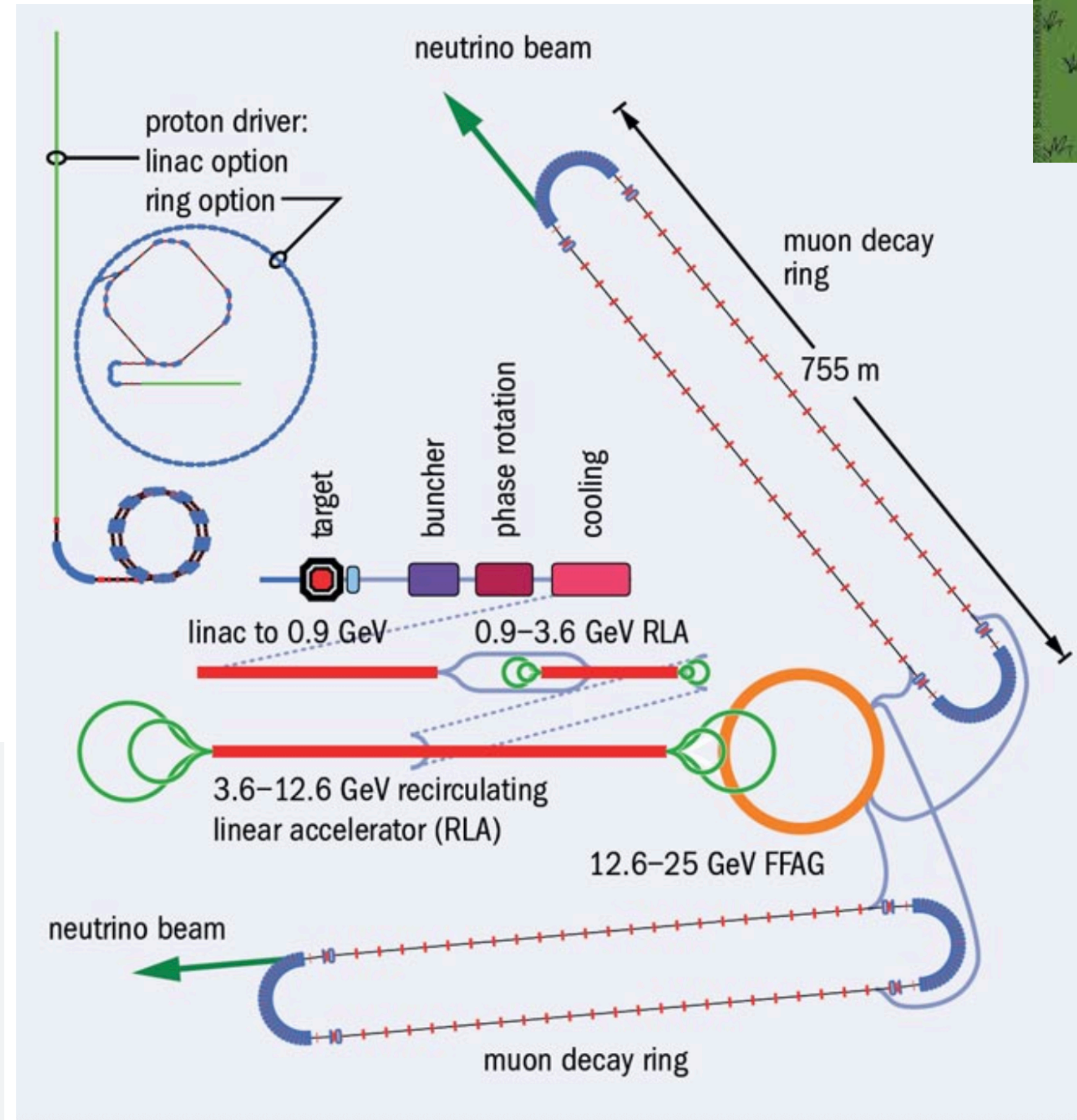
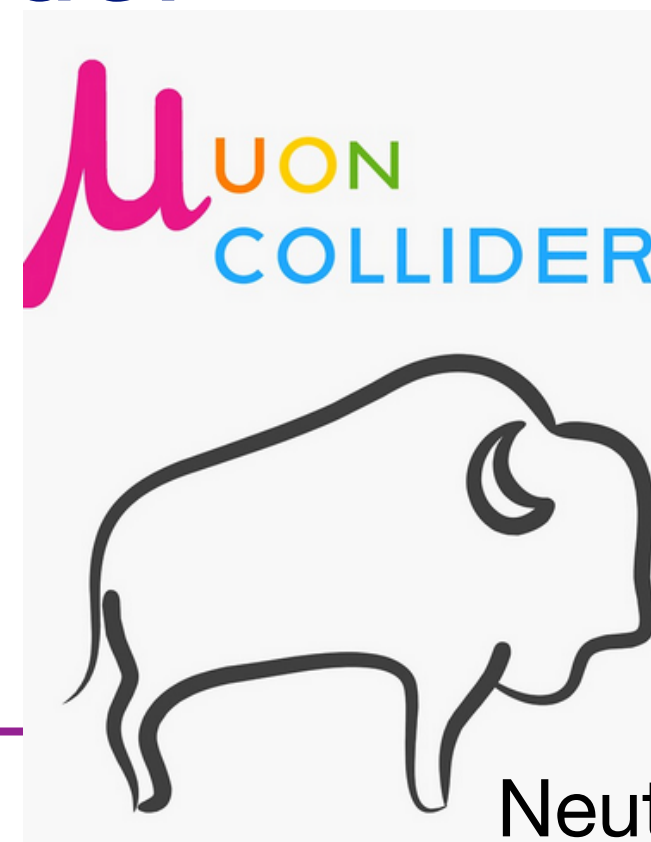
Neutrino physics

The far future (maybe?)

renewed interest of particle physics community in muon colliders

Neutrino factory could be a pre-stage of a high-energy muon collider

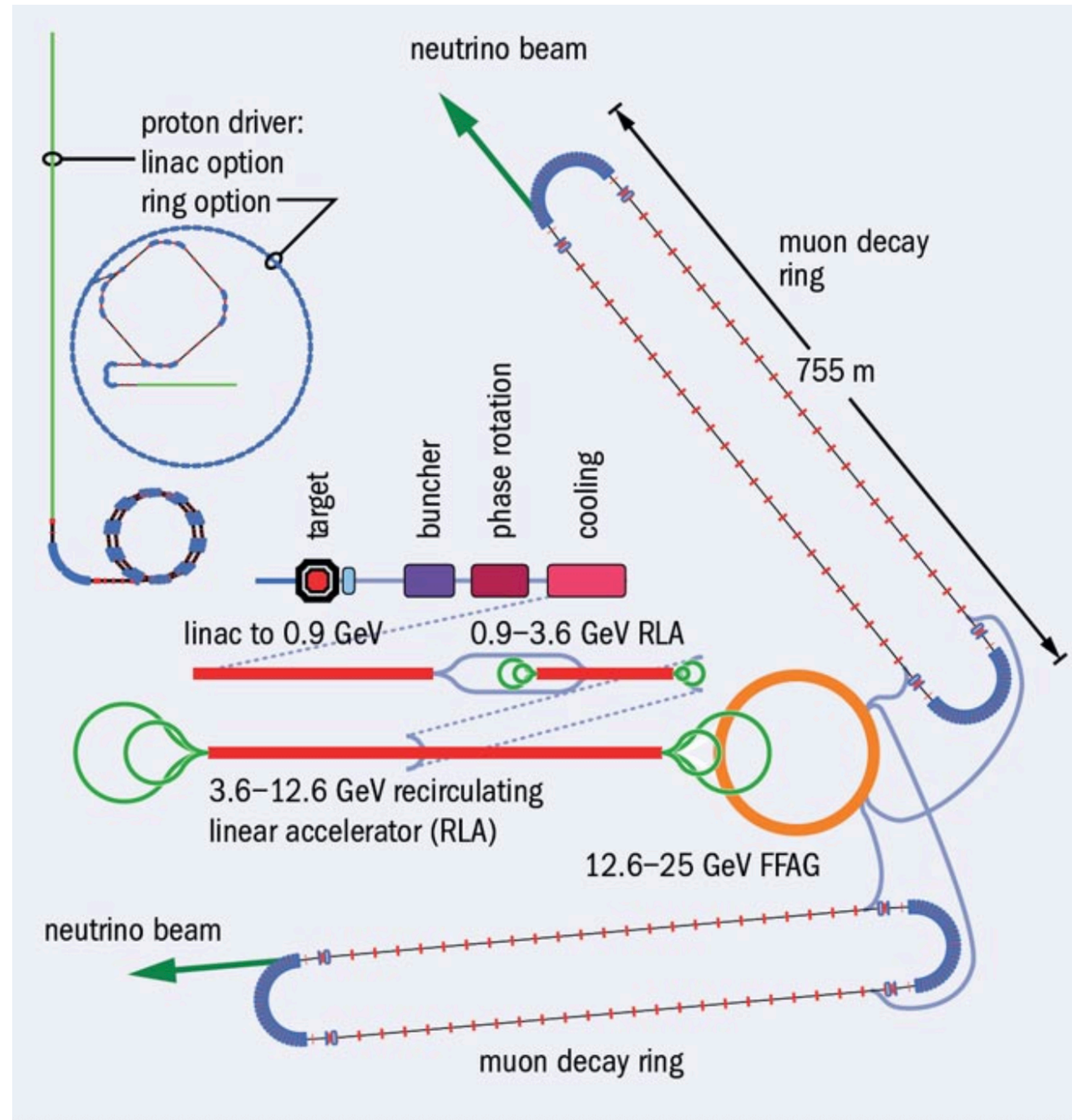
Synergy with future collider program



Bison at Fermilab

Neutrino physics

Neutrino factory



- Neutrino production from GeV (anti-)muon decays
- Muons decay along straight line of race track design → very collimated neutrino beam of **known** composition
- ν_e , ν_μ , ν_τ appearance & ν_e , ν_μ disappearance searches possible
- Will **improve** constraints on SM + BSM scenarios assuming FNAL-SURF baseline and 40 kT – 10^{22} μ decays

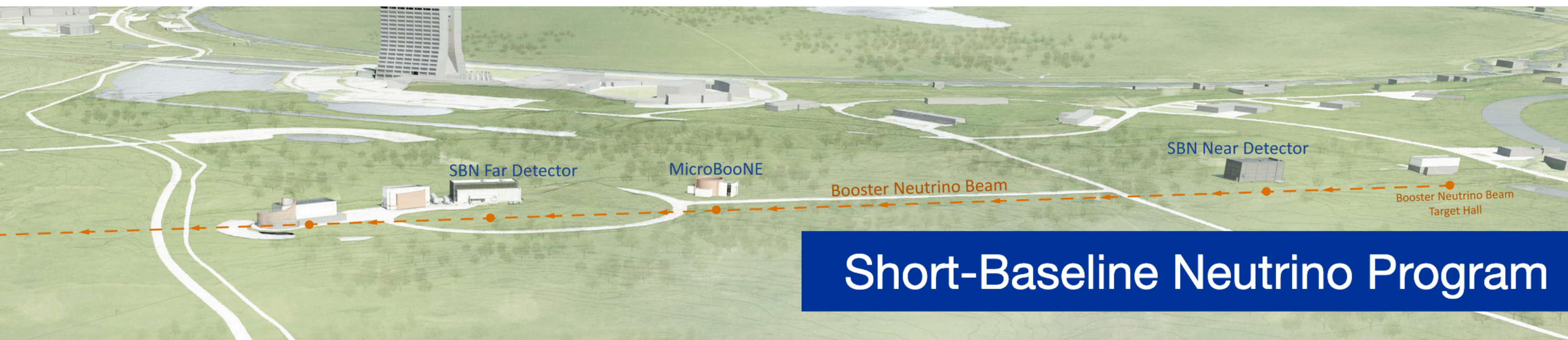
[Denton, JG, '24]

[Denton, JG, Kong, '25]

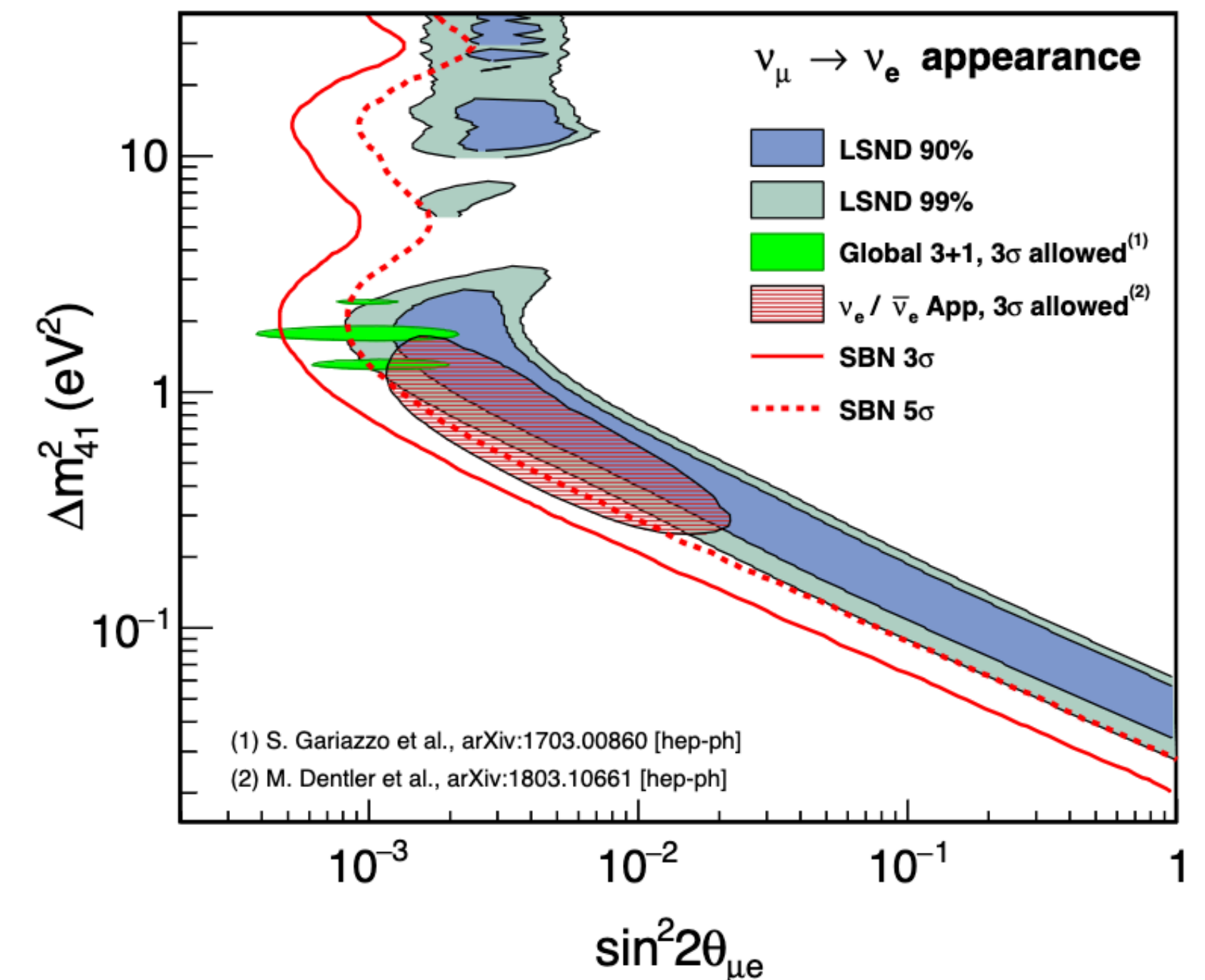
Neutrino physics

The future

Short baseline program:
Testing the MiniBooNE anomaly



3 detectors: MicroBooNE, ICARUS, SBND



[Machado, Palamara, Schmitz
1903.04608]

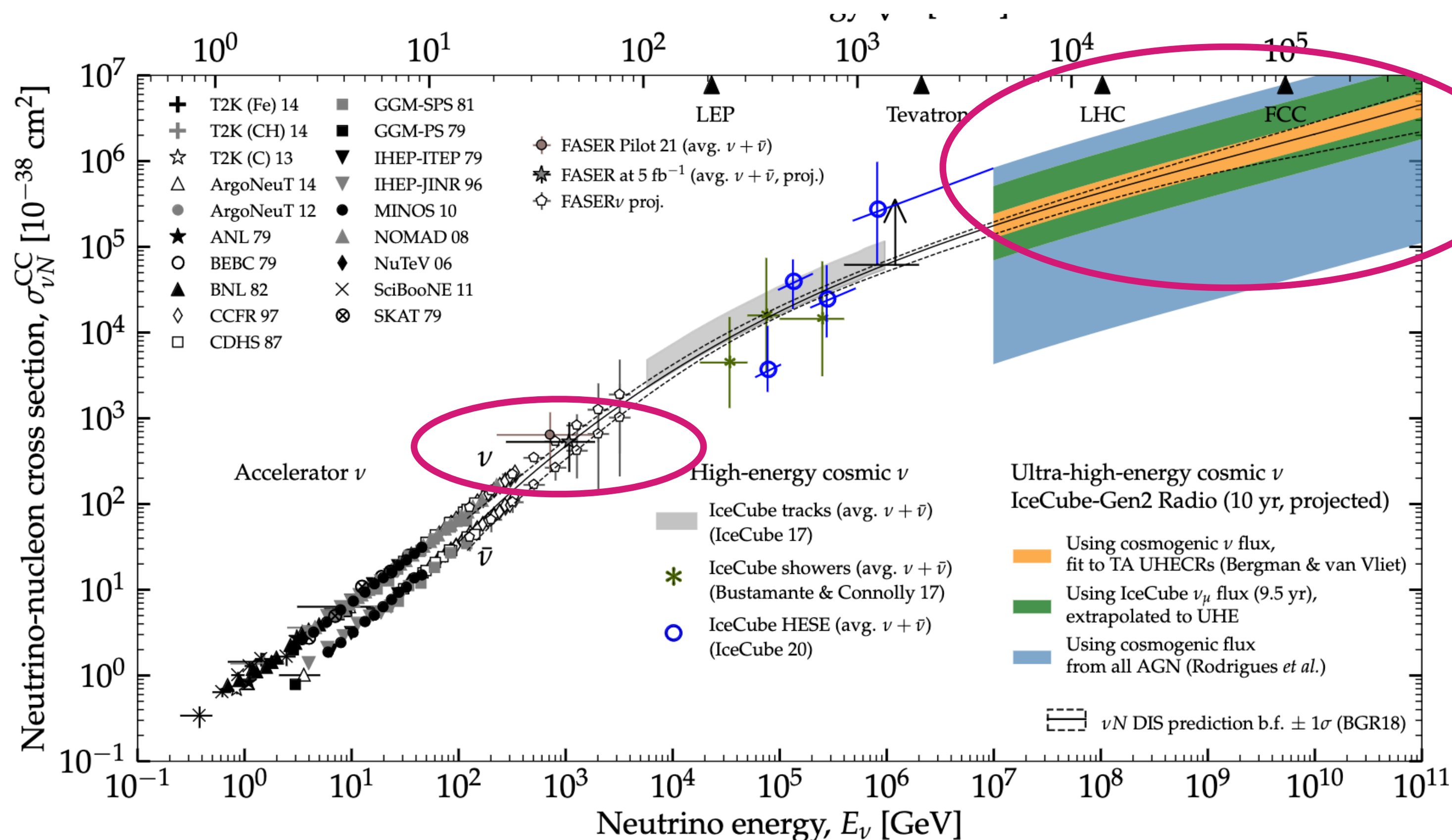
Neutrino physics

The future

FASER ν ,
SBN@LHC:
xsec at TeV
energies

SHIP:
measurement
of ν_τ xsec

Neutrino scattering



IceCube-
Gen2:
UHE cosmic ν s

CEvNS
experiments:
MeV neutrino
scattering

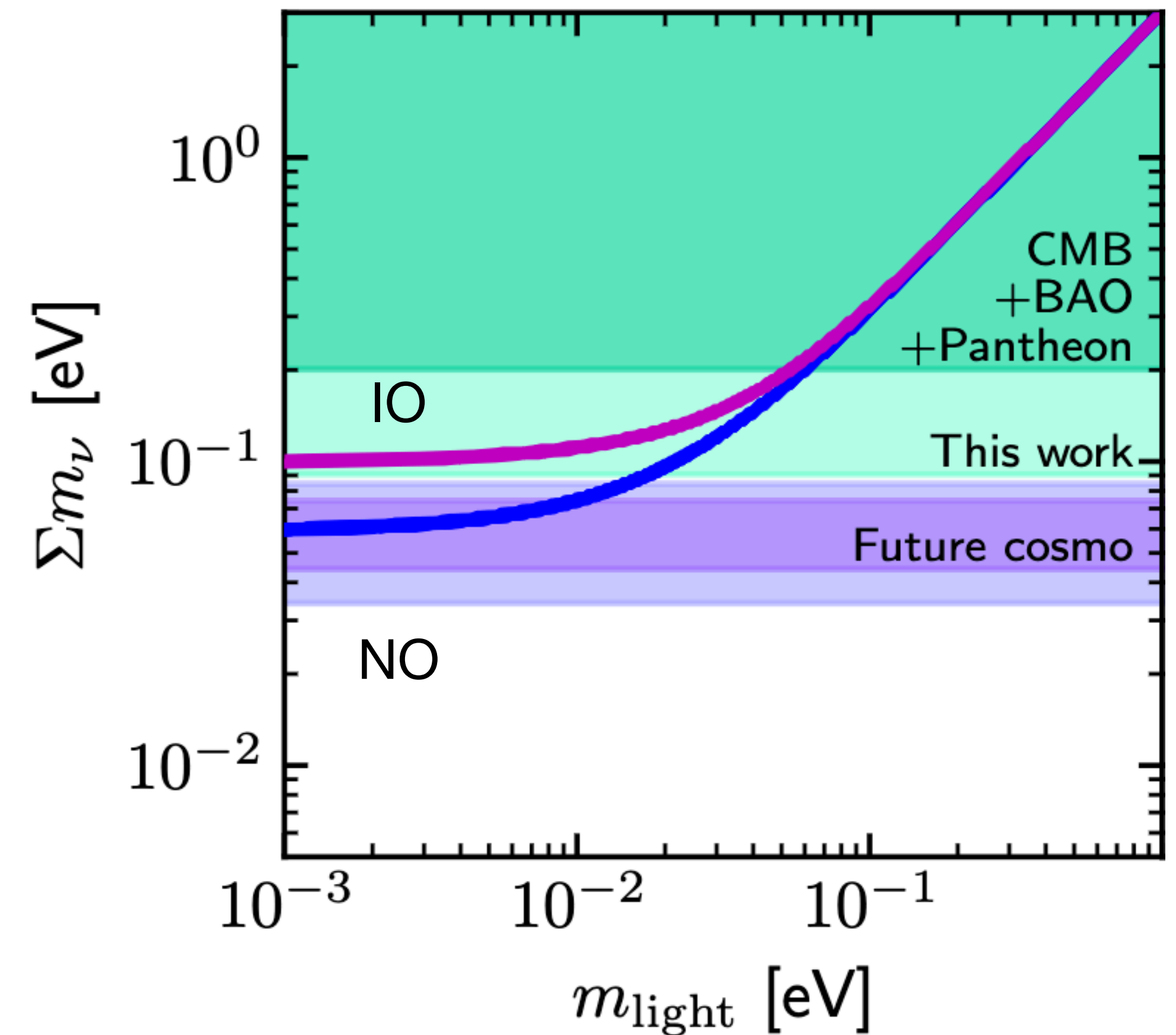
Neutrino physics

The future

[Di Valentino, Gariazzo, Mena, '21]

Absolute neutrino mass scale/
nature of neutrinos:

Future cosmological observatories will
measure sum of neutrino masses (maybe
very soon)
(And potentially mass ordering)



Neutrino physics

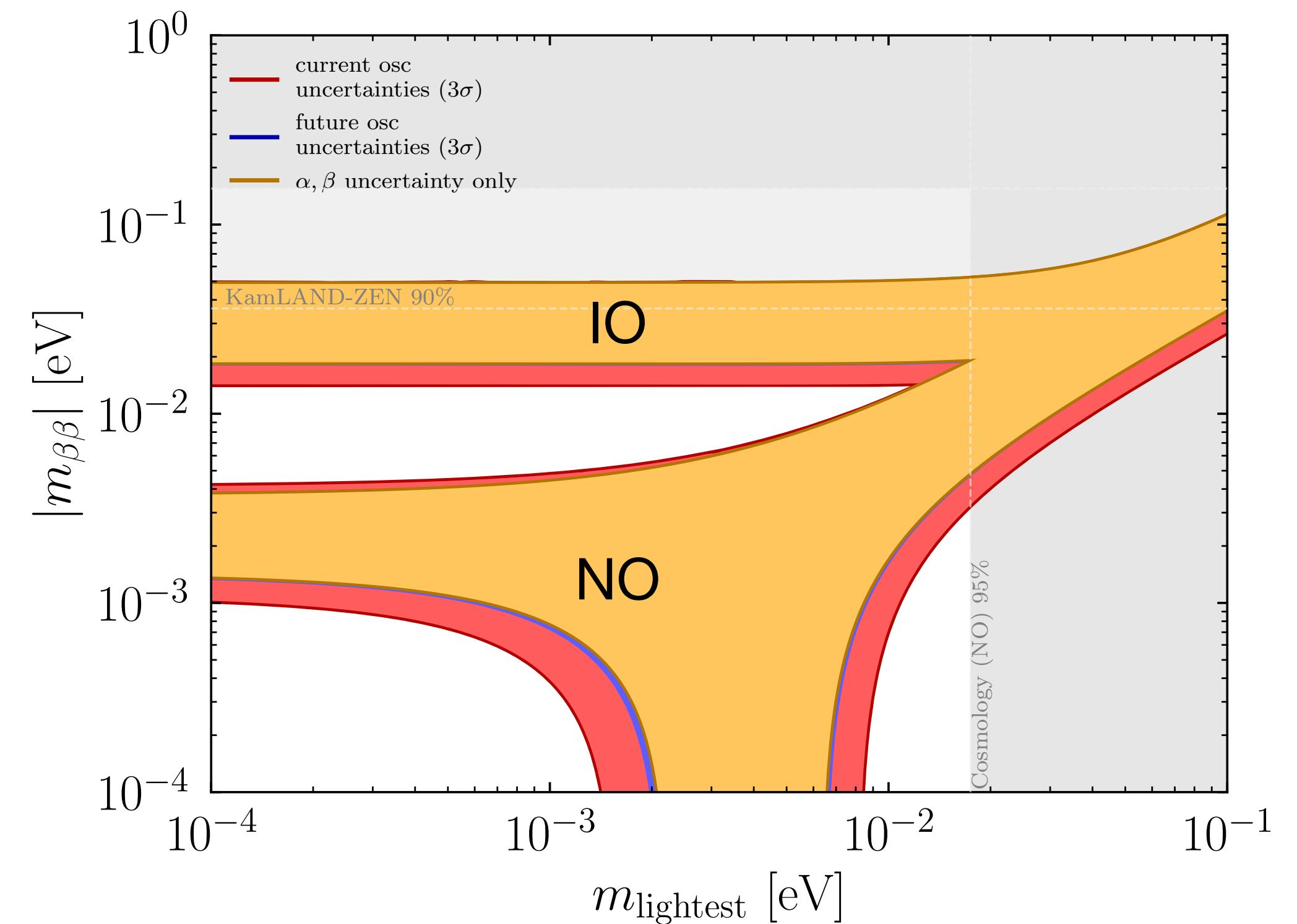
The future

[Denton, JG, '23]

Future neutrinoless double beta experiments
aim to **fully probe** IO region

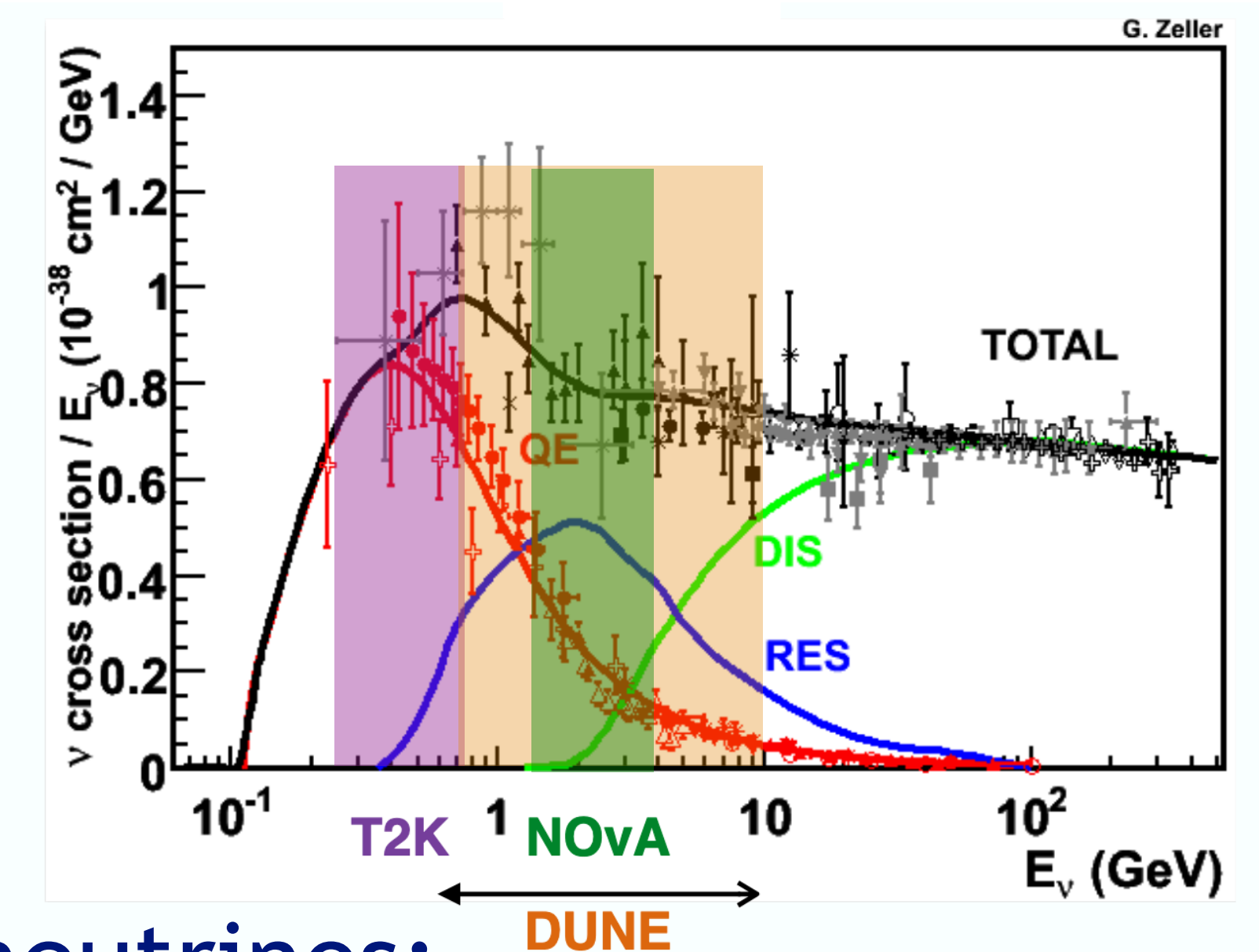
Progress on nuclear matrix element calculation
(new results: matrix elements are smaller
→ bound on $|m_{\beta\beta}|$ weaker)

[Belley et al, '23]



Neutrino physics Challenges

- For neutrinoless double beta decay (MeV):
nuclear matrix element calculation
- For accelerator based neutrino
experiments (GeV):
cross section & flux prediction
- For high energy atmospheric and astrophysical neutrinos:
hadronic interactions for flux prediction
→ forward physics program at the LHC



Neutrino physics

Conclusions

- Neutrinos play an **essential** role in our search for the new SM
- Future experiments will measure **all** neutrino parameters and potentially uncover new symmetries and particles in nature
- Possibilities that the neutrino sector will continue to amaze us
in the future

Thanks for your attention!

